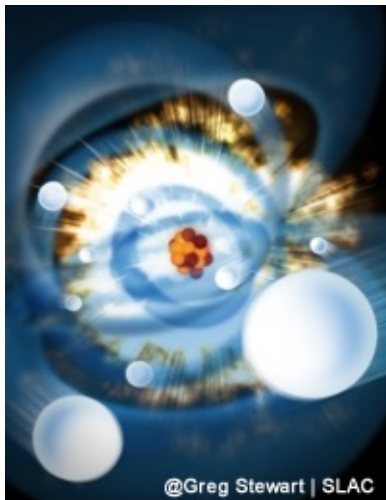


Harnessing ultraintense x-rays: from atomic response to applications



Linda Young
Nobel Symposium 158
Sigtunahöjden, Sigtuna, Sweden
14 Jun 2015

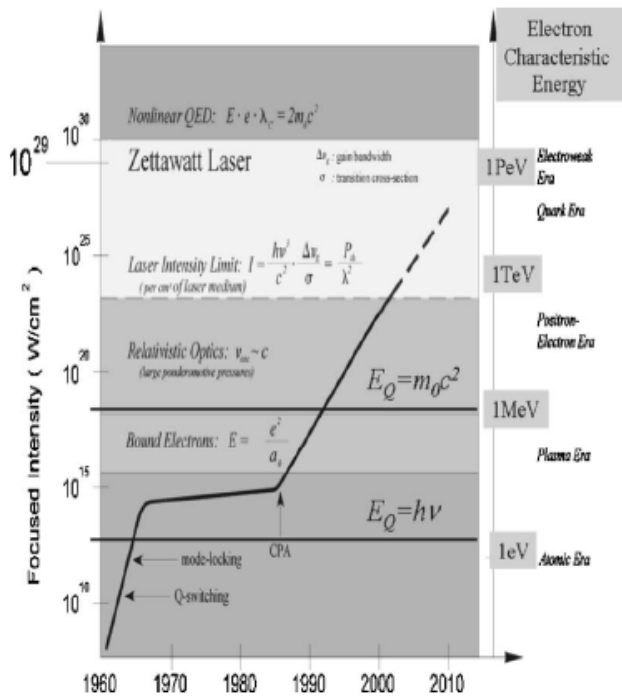
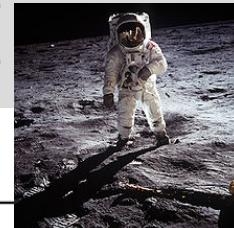
Compare ultraintense optical and x-ray sources

High-intensity at optical wavelengths

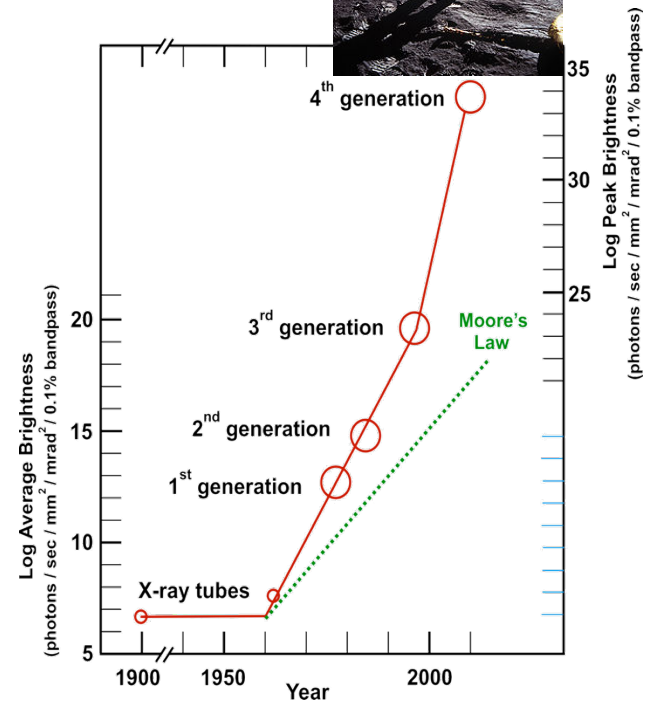
- high harmonic generation
- tabletop coherent x-ray radiation
- attosecond pulses

High-intensity at x-ray wavelengths

?
?
?



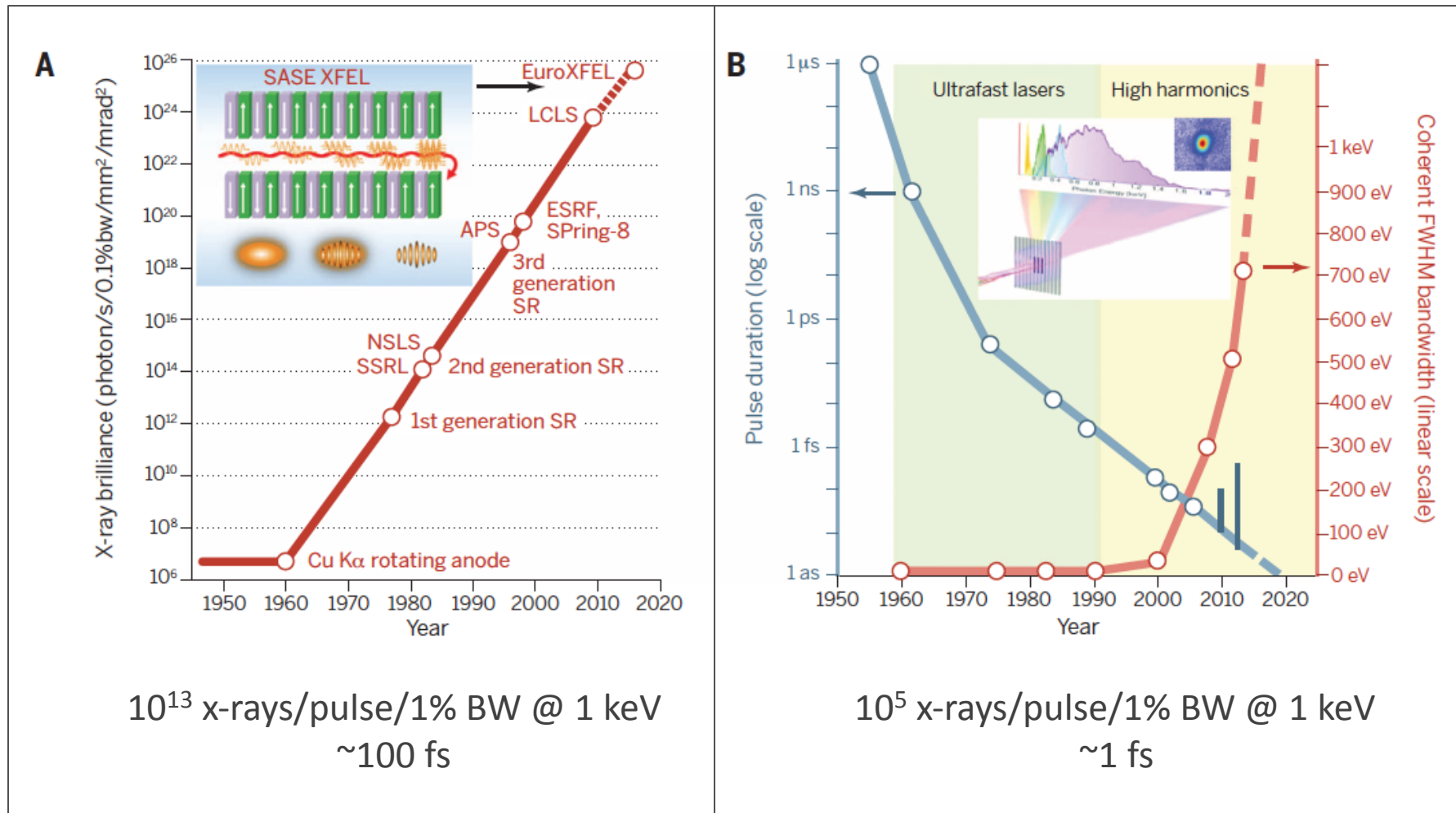
G. Mourou RMP 2006



D. Moncton, George Brown



X-ray sources: accelerator-based vs laser-based HHG



XFELs 10^8 "brighter" than HHG sources



Outline

- Birth of world's first hard x-ray FEL - LCLS
- Non-resonant high intensity x-ray phenomena
LCLS Experiment 1: Oct 1 - 6, 2009
L. Young *et al.*, Nature 466, 56 (2010)
- Resonant high intensity x-ray processes
LCLS Experiment 5: Oct 1 - 6, 2009
E.P. Kanter *et al.*, PRL 107, 233001 (2011)
- Towards single particle imaging



Linac Coherent Light Source at SLAC

X-FEL based on last 1-km of existing 3-km linac

Proposed by C. Pellegrini in 1992

1.5-15 Å
(14-4.3 GeV)

Injector (35°
at 2-km point

Existing 1/3 Linac (1 km)

New e⁻ Transfer Line (340 m)

X-ray Transport
Line (200 m)

Undulator (130 m)

Near Experiment Hall

Far Experiment
Hall



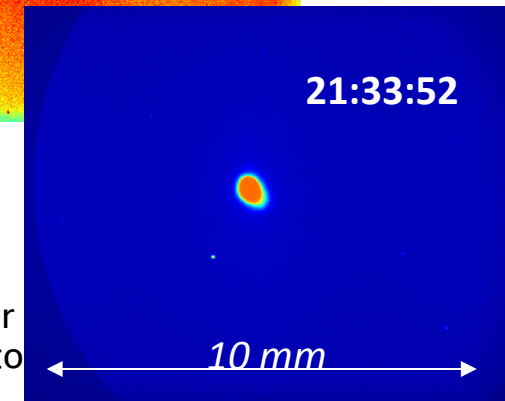
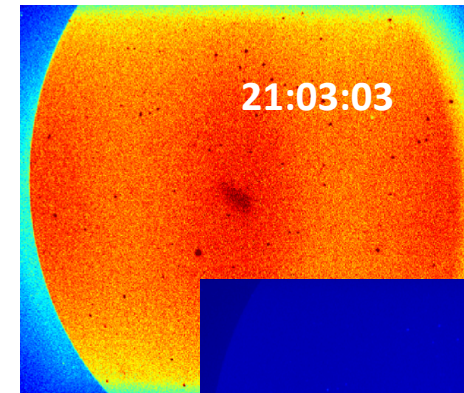
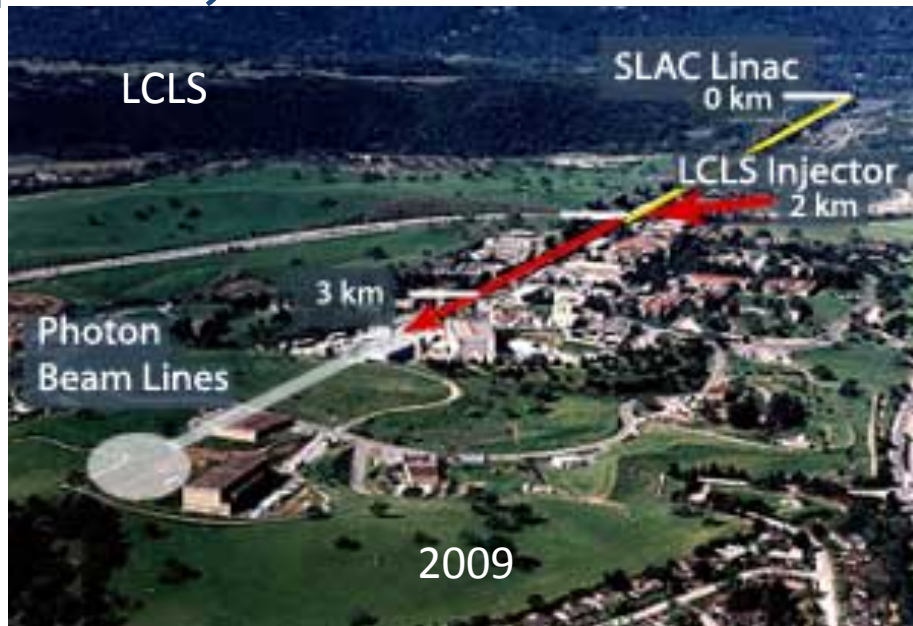
UCLA



LLNL



April 10, 2009: LCLS lases at 1.5 Angstroms



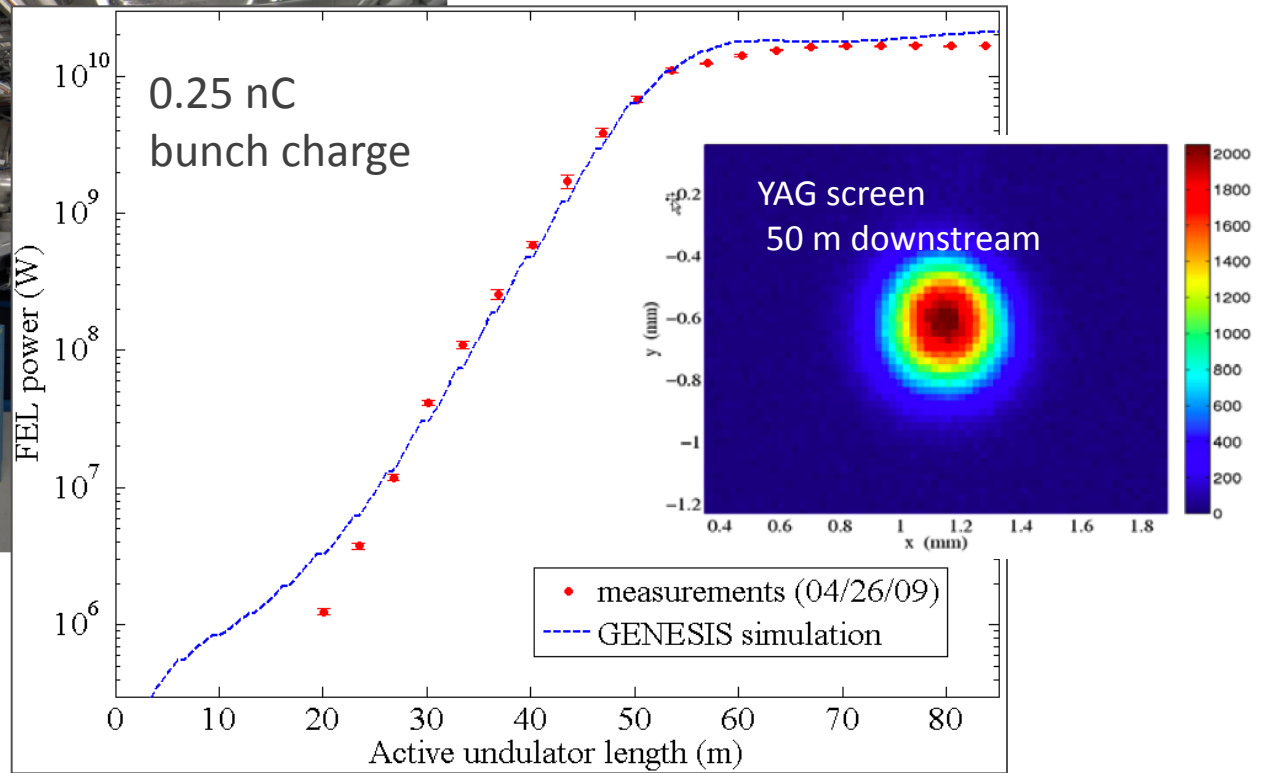
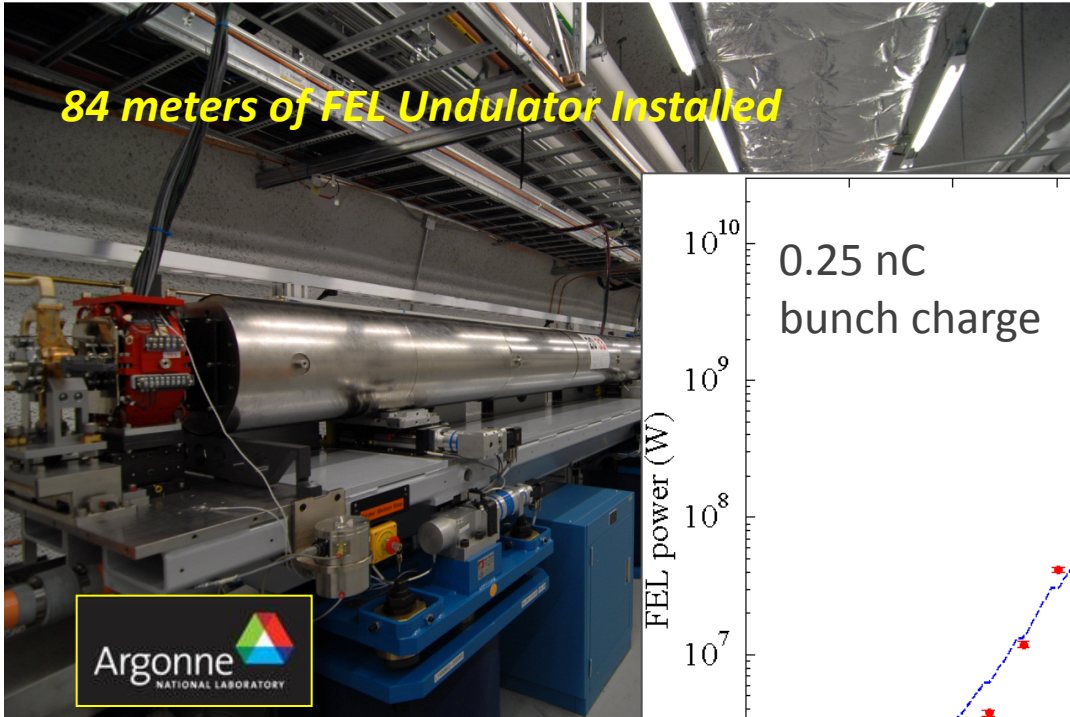
In my life before SLAC, I had the privilege to participate (in various capacities) in the design, construction and commissioning of two linacs, two synchrotrons, four storage rings and three FELs (free-electron lasers). Now I have had the privilege to be in SLAC's Main Control Center on April 10, when the Linac Coherent Light Source became a 1.5 Ångstrom laser. I don't expect I will ever, as long as I live, see such a beautiful, smooth turn-on of any light source. With each undulator placed on the beam path, the FEL power increased by a factor of about 2.3; two hours into the first attempt at lasing, the pinpoint of FEL light from twelve undulators was nearly 2,000-fold more intense than plain old undulator radiation. The team called it quits at 11:30 p.m. that night. When they returned at 8:00 a.m. the next morning, the FEL light came back as soon as the shutter was opened.

- John Galayda





LCLS saturation at 1.5 Å



- Saturation after ~65 meters of undulator!

Science Drivers for LCLS



AMO: Atomic Molecular and Optical

SXR: Soft X-ray Materials Science

XPP: X-ray Pump-Probe

XCS: X-ray Correlation Spectroscopy

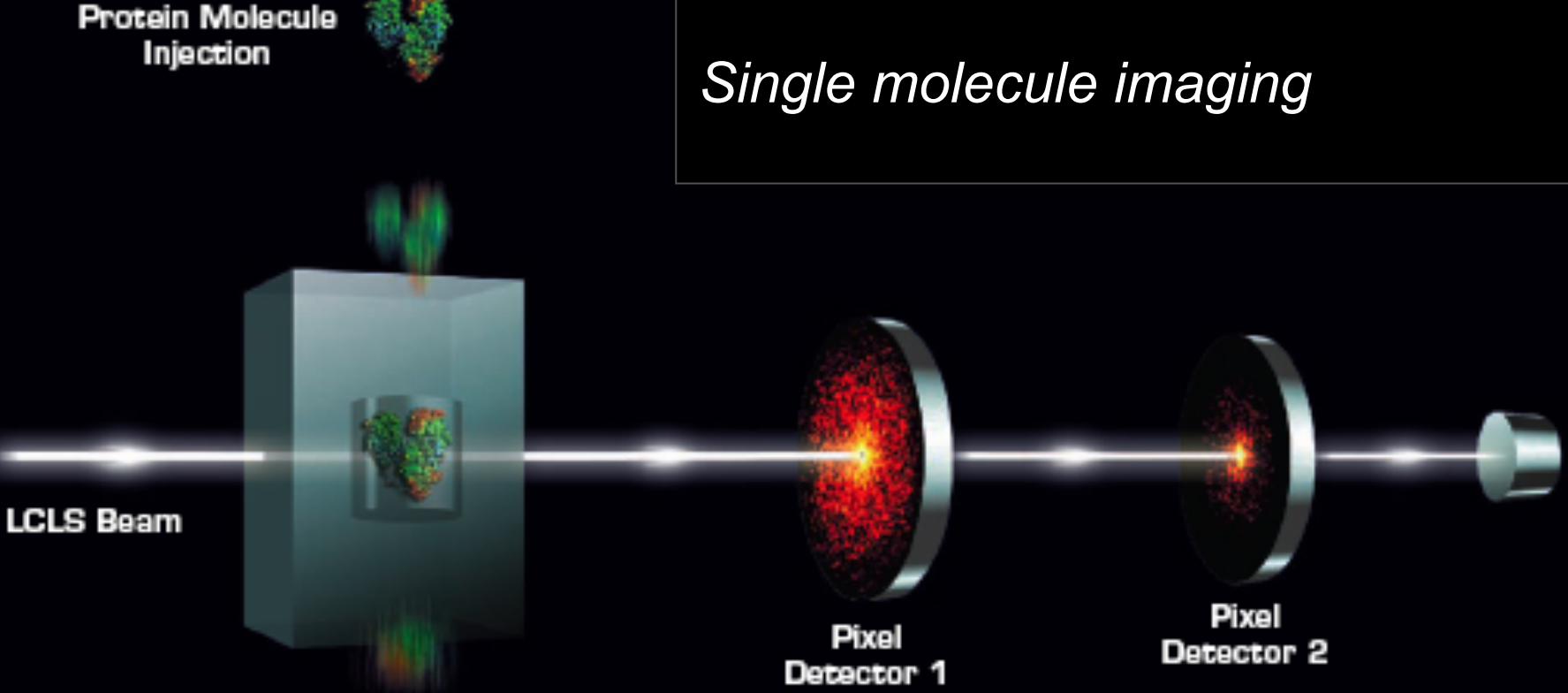
CXI: Coherent X-ray Imaging

MEC: Materials in Extreme Conditions

AMO

- Understand and control x-ray atom/molecule interactions at ultrahigh x-ray intensity as a foundation for other applications.
- Provide diagnostics of the LCLS radiation

Single molecule imaging



Protein Molecule Injection

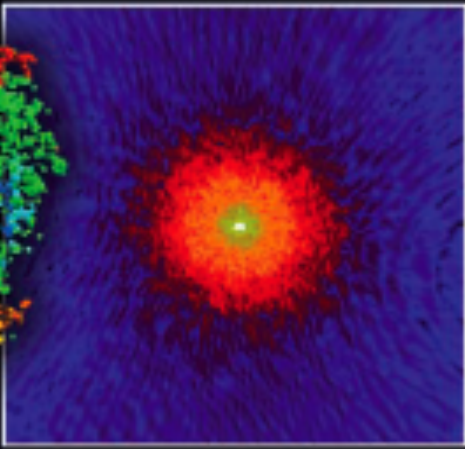
LCLS Beam

Pixel Detector 1

Pixel Detector 2

To Mass Spectrometer

Protein Molecule

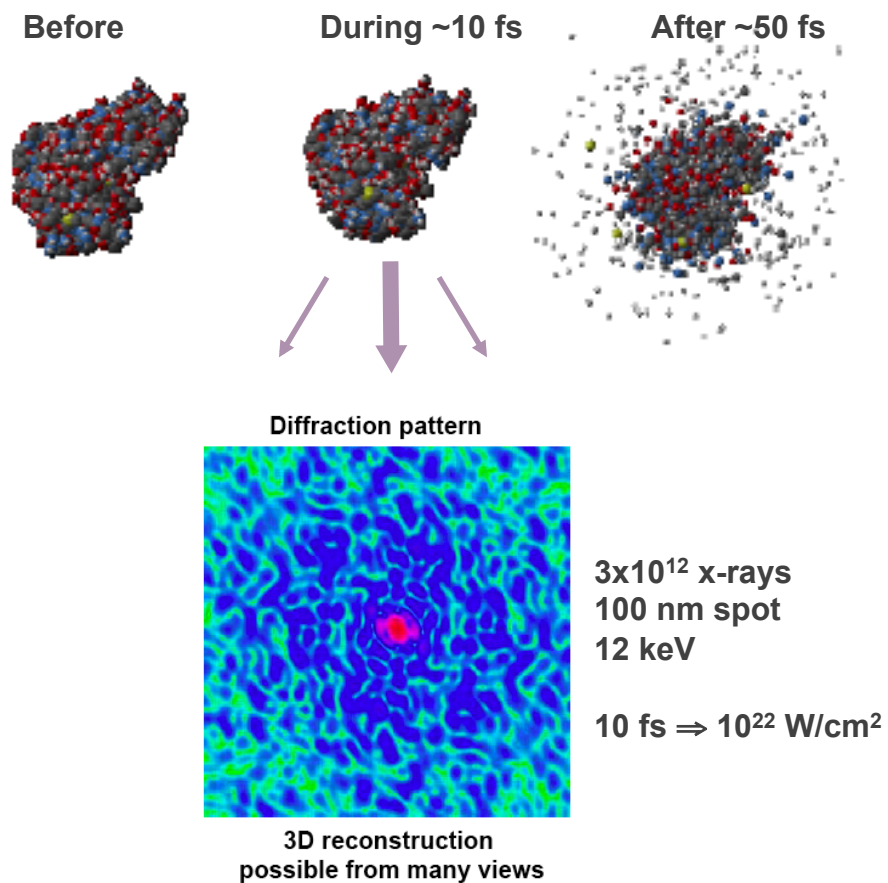


X-ray Diffraction Pattern

Operating with ultrafast pulses, LCLS will take images of molecules dropped into the x-ray beam. Scientists will merge the series of diffraction patterns of the molecules in many different positions. The resulting three-dimensional reconstruction will reveal the structures of proteins that cannot be crystallized and thus studied any other way.

AMO questions at the ultraintense x-ray frontier

- fundamental nature of x-ray damage at high intensity
 - Coulomb explosion
 - electronic damage
 - behavior at 10^{22} W/cm² - 1Å
- nonlinear x-ray processes
 - role of coherence
- quantum control of inner-shell processes



Neutze, Wouts, van der Spoel, Weckert, Hajdu Nature 406, 752 (2000)



LCLS Experiment 1 - Oct 1, 2009

Nature of the electronic response to

10^5 x-rays/Å²
80 - 340 fs
800 - 2000 eV

$\sim 10^{18}$ W/cm²

Original single molecule imaging parameters, Neutze et al. Nature (2000)

3×10^{12} x-rays/(100 nm)² = 3×10^6 x-rays/Å²

10 fs

$\sim 10^{22}$ W/cm²



Our approach to understanding ultraintense x-ray interactions

- Start with a well-characterized target

Binding energies in neutral neon

2p : ~21 eV

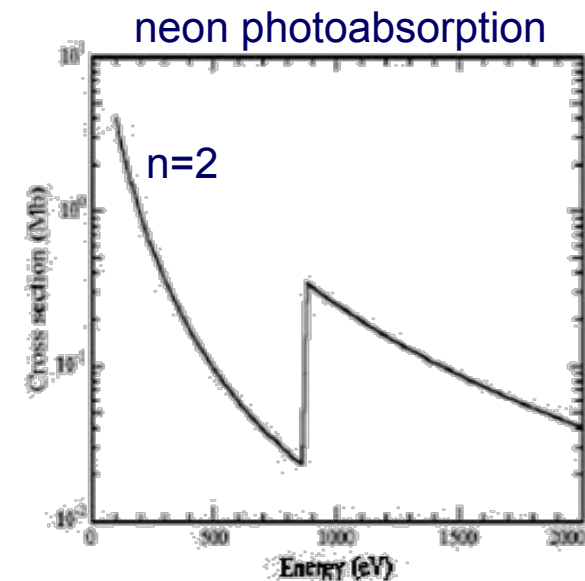
2s : ~48 eV

1s : ~870 eV

Inner-shell excitation

Auger yield 98%

Auger clock - τ_{1s} : 2.4 fs

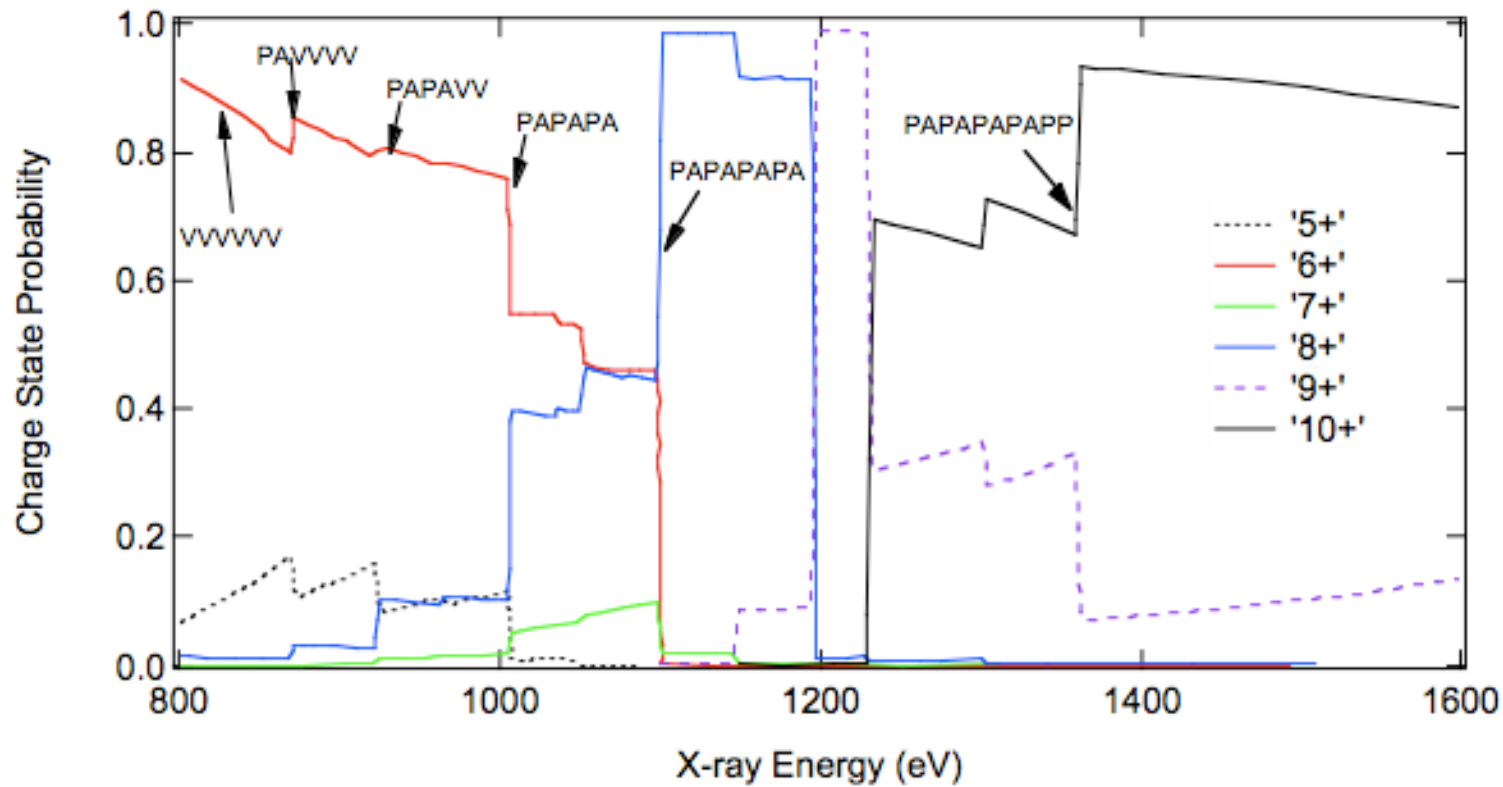


- Probe changes in interaction from outer- to inner-shell between 800-2000 eV



Guided by theory

Theory: Rohringer & Santra, PRA 76, 033416 (2007)

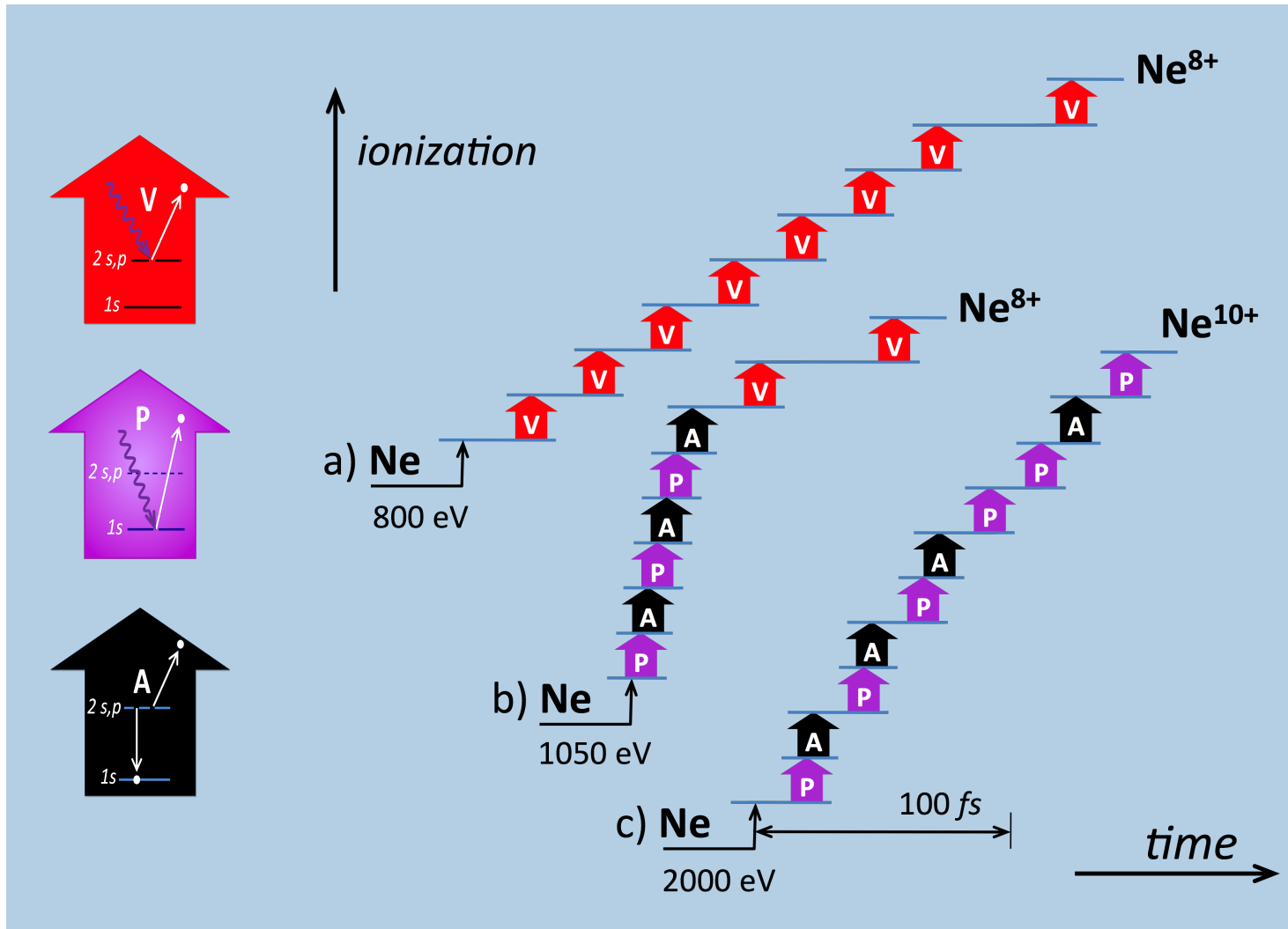


LCLS specs
10¹³ x-rays
230 fs
1 μm spot

Three target energies: 800 eV, 1050 eV, 2000 eV



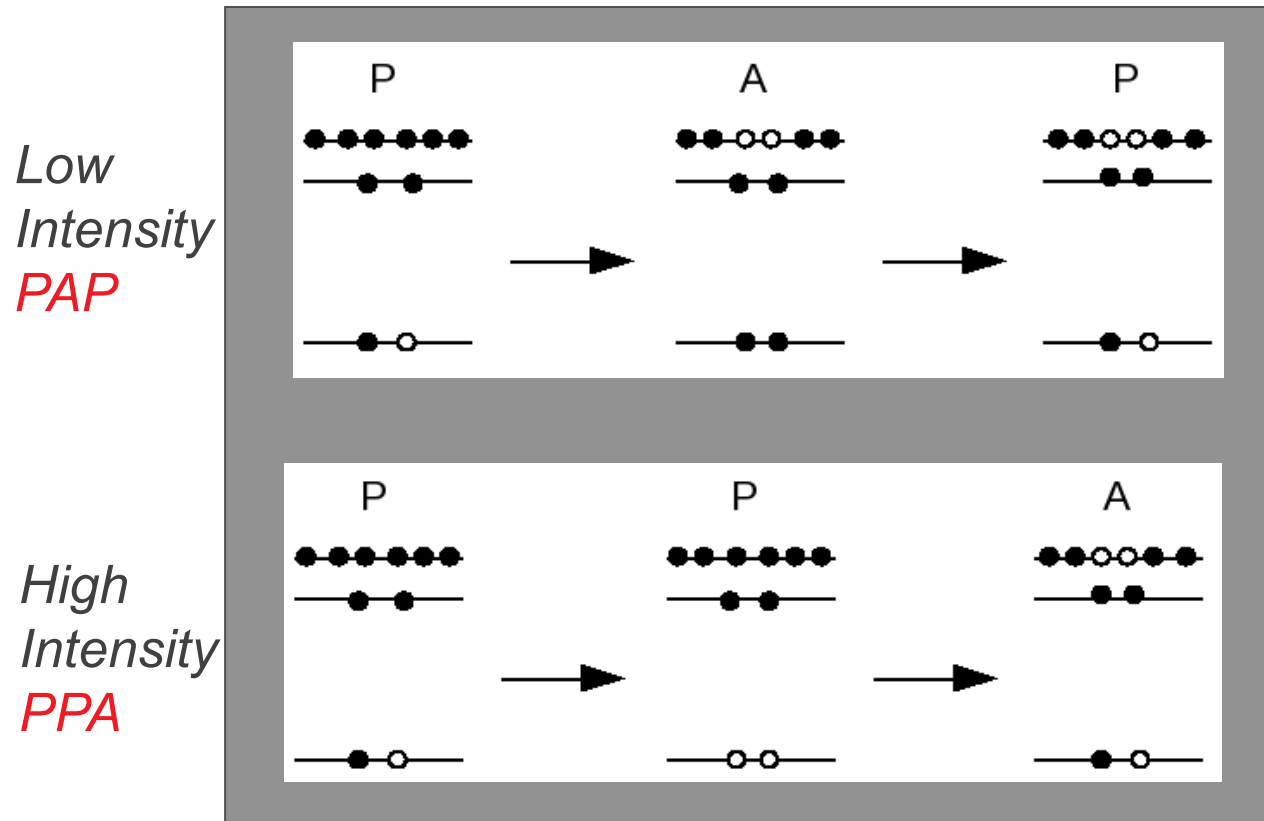
Valence ionization, core ionization and Auger decay



Sequential single photon processes dominate the interaction



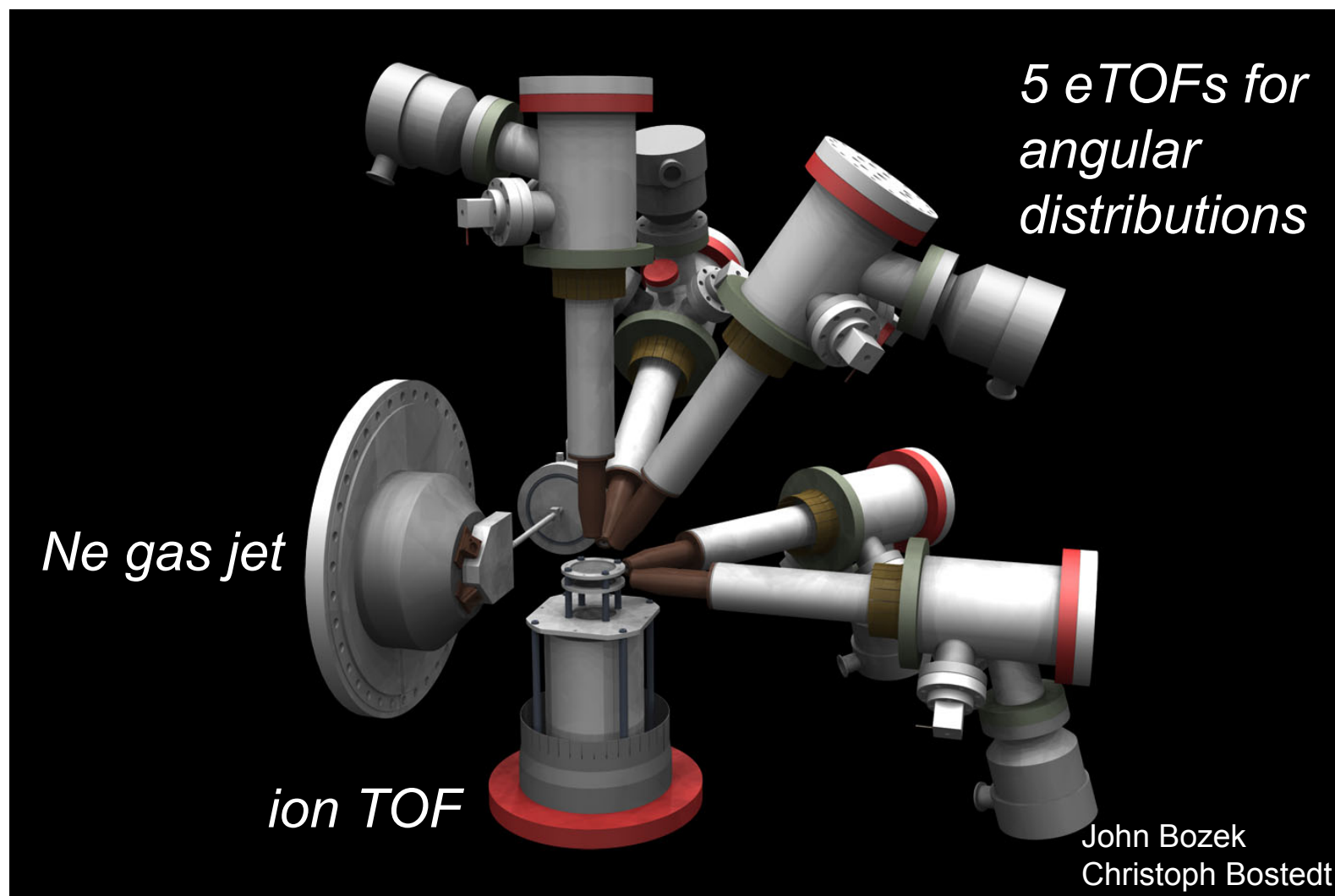
How does one arrive at a particular charge state?



- Hollow atoms produced at high x-ray intensity
- Electron spectroscopy can define the mechanism

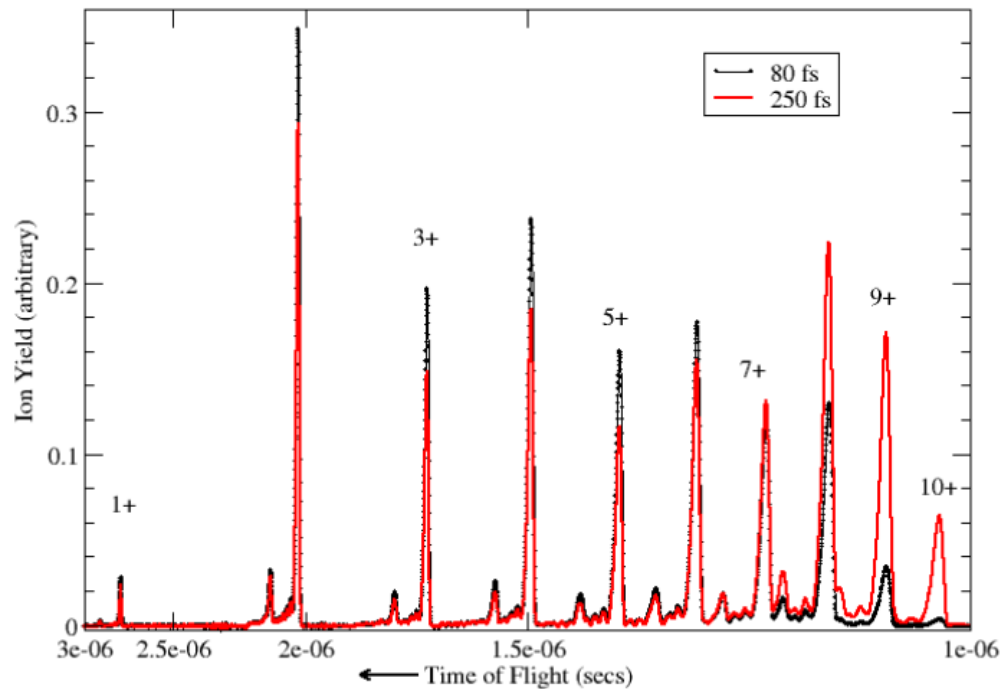


High field physics chamber



Day 1 - two interesting observations

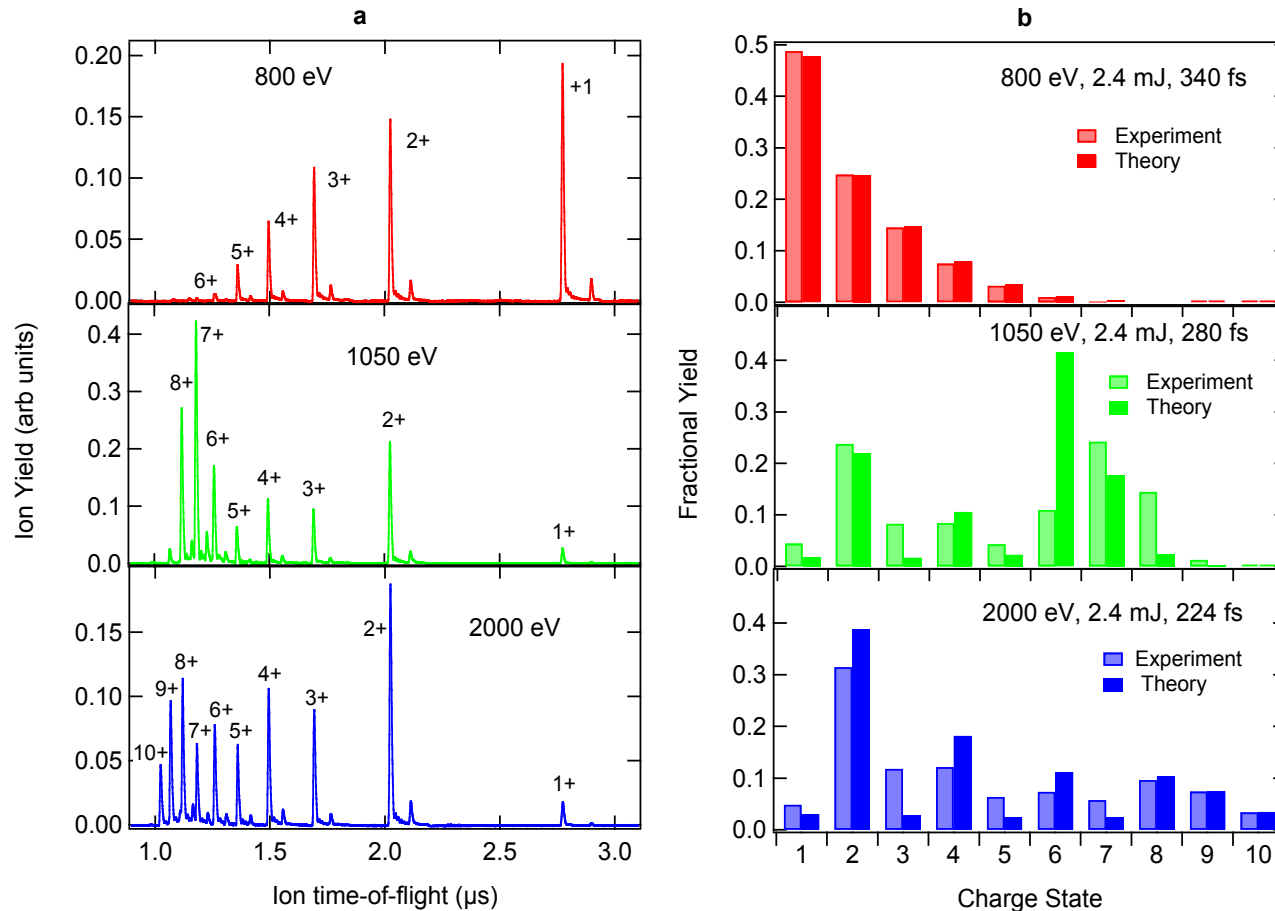
- Single ~ 100 fs pulse at 2000 eV fully strips neon
6-photon, 10-electron process



- Shorter pulses with equal pulse energy & fluence suppress absorption & damage.



Theory can model ultraintense x-ray-induced electronic damage in neon



Theory

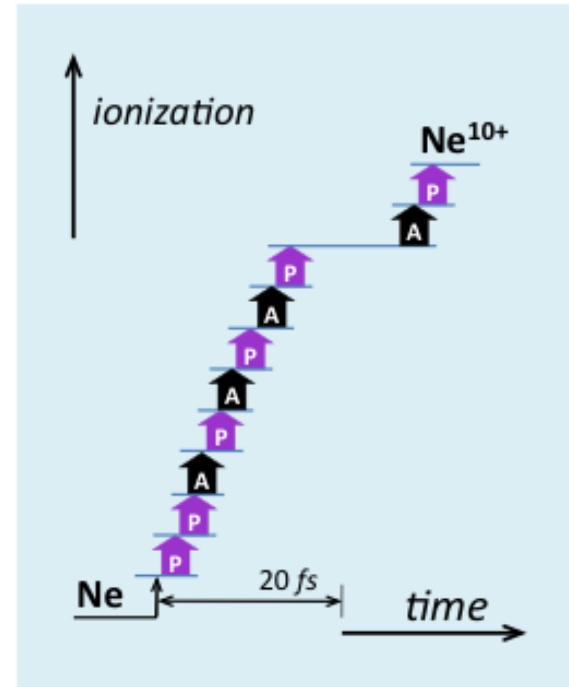
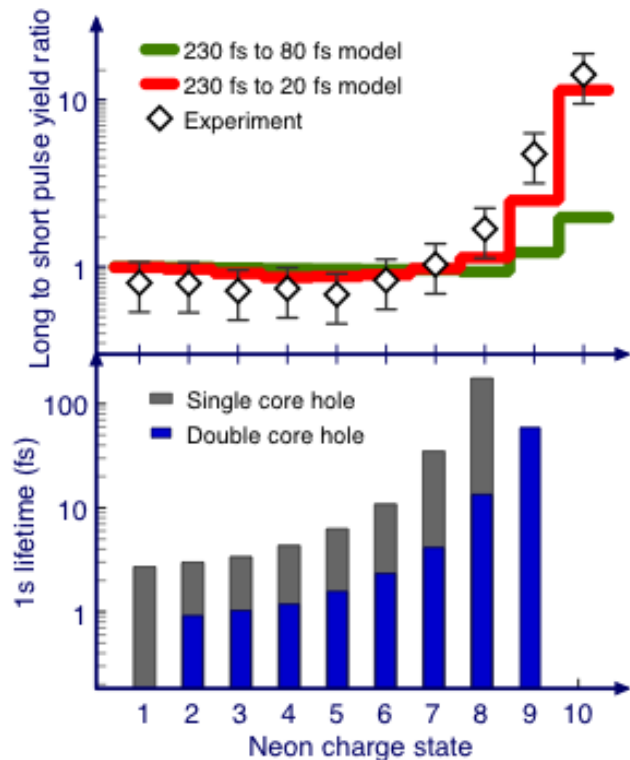
- Intensity averaged
- Fluence determined by experiment

Consistent with “measured” pulse energy and focus.

Sang-Kil Son, Robin Santra – refined calcs include shakeoff – G. Doumy et al, PRL 2011)



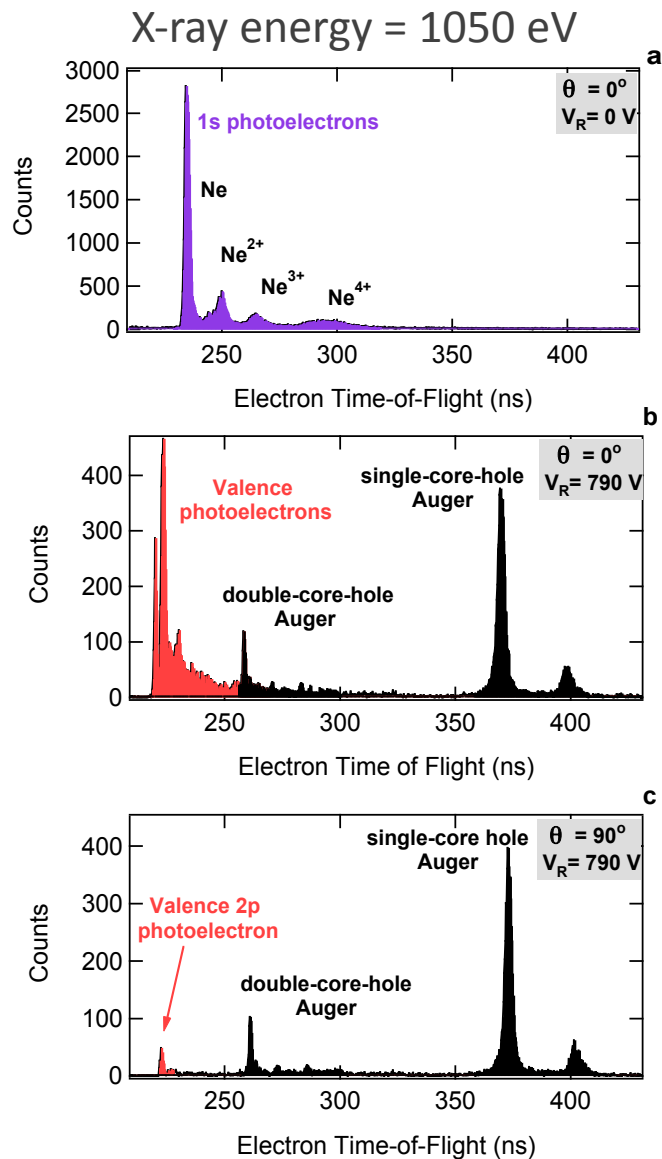
Atoms become transparent at high x-ray intensity !



- x-ray absorption is due to the presence of 1s electrons
- high x-ray intensities eject both 1s electrons rendering the atom transiently transparent
- slowing atomic clocks create transparency at surprisingly long timescales



Electron spectrometers track ionization mechanism



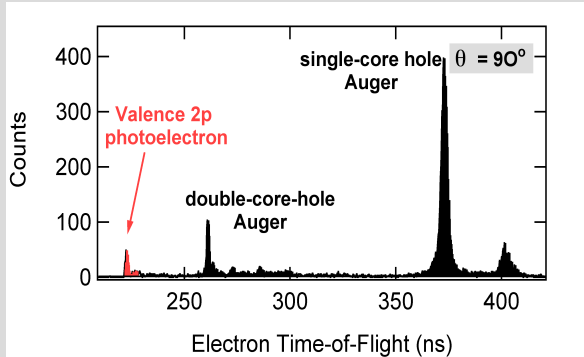
“Slow” 1s photoelectrons along x-ray polarization axis

“Fast” valence photoelectrons and Augers along polarization axis

Clean hollow atom signature
double-core-hole Auger
 $\theta = 90^\circ$

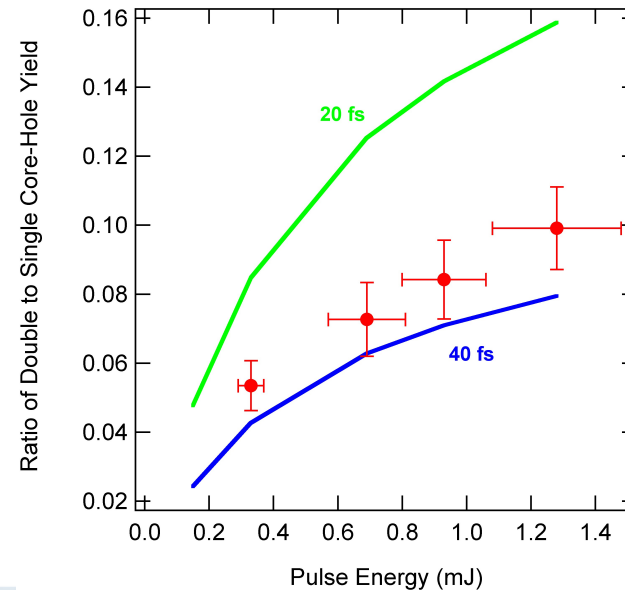


Hollow atom production: deliberate, huge and an an indicator of x-ray pulse duration

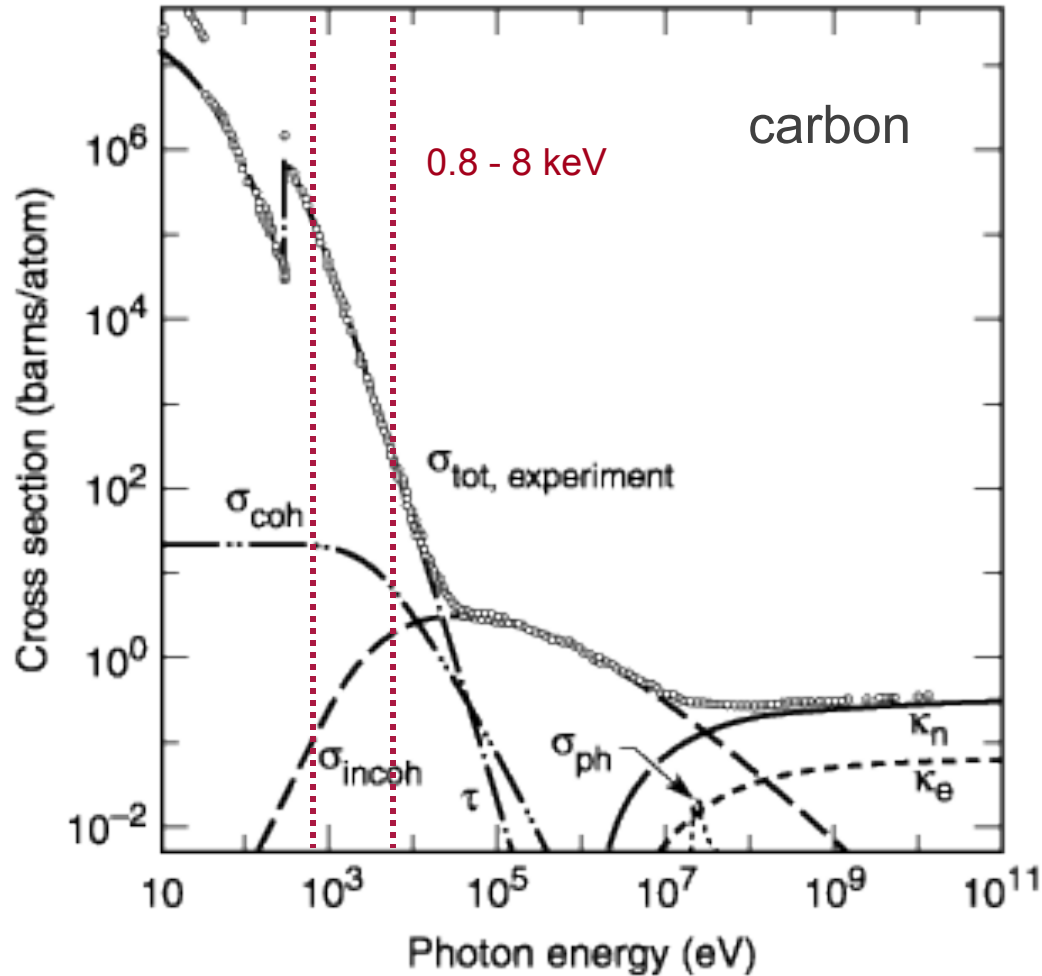


Hollow atom yield
@ LCLS ~10%
@ synchrotron ~0.3%
due to electron correlation

1050 eV,
nominal electron bunch
duration ~80 fs



Absorption vs scattering: normal and hollow atoms



	$\frac{\sigma_{\text{photo}}}{\sigma_{\text{elastic}}}$	$\frac{\sigma_{\text{Compton}}}{\sigma_{\text{elastic}}}$
2 keV	360	0.05
8 keV	20	0.60
8 keV hollow	2	

Impact of hollow atom formation on coherent x-ray scattering

Sang-Kil Son, L Young, R Santra
Phys. Rev A. **83**, 033402 (2011)



Prescient prediction

“Furthermore, during short intense pulses numerous K-holes may be present at any one time, reducing the photoelectric cross-sections of atoms in which they were produced and thus lowering the total number of primary ionization events in the sample (see trend in Fig. 1). **This effect makes the system radiation hardened to photo-ionization during very short exposures, and is more pronounced at higher radiation intensities.**”

Neutze, Wouts, van der Spoel, Weckert, Hajdu Nature (2000)



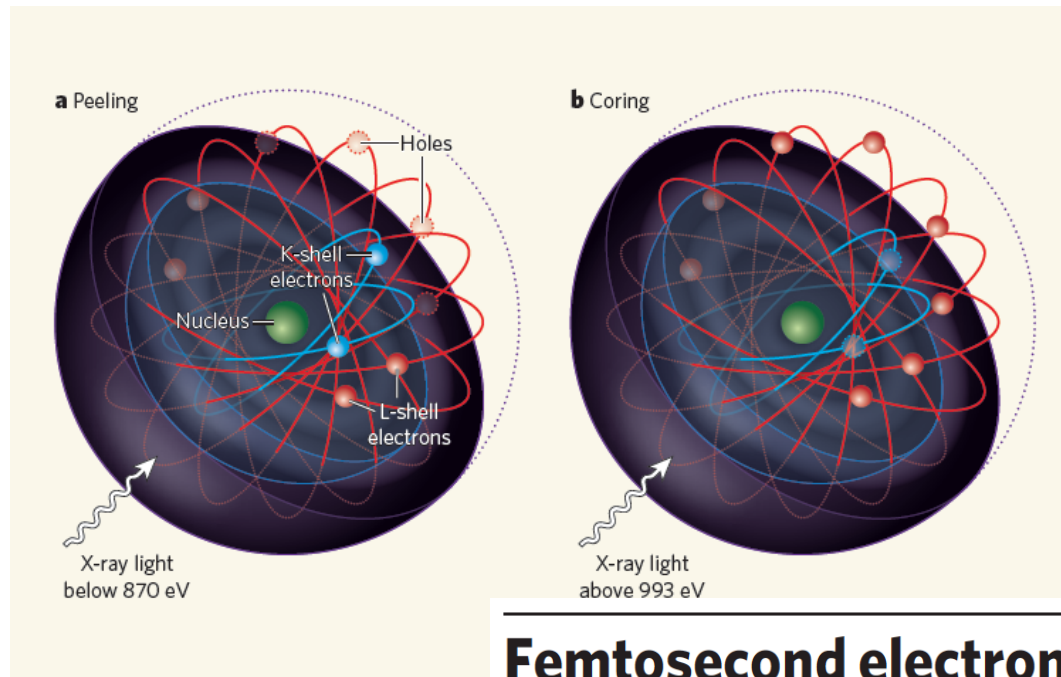
NEWS & VIEWS

ATOMIC PHYSICS

X-ray laser peels and cores atoms

Justin Wark

The world's first kiloelectronvolt X-ray laser produces such a high flux of photons that atoms can be 'cored'. In other words, the light source can knock out both the electrons of an atom's innermost shell.



Femtosecond electronic response of atoms to ultra-intense X-rays

Nature 466, 56 (2010).

L. Young¹, E. P. Kanter¹, B. Krässig¹, Y. Li¹, A. M. March¹, S. T. Pratt¹, R. Santra^{1,2}, S. H. Southworth¹, N. Rohringer³, L. F. DiMauro⁴, G. Doumy⁴, C. A. Roedig⁴, N. Berrah⁵, L. Fang⁵, M. Hoener^{5,6}, P. H. Bucksbaum⁷, J. P. Cryan⁷, S. Ghimire⁷, J. M. Glowia⁷, D. A. Reis⁷, J. D. Bozek⁸, C. Bostedt⁸ & M. Messerschmidt⁸

Summary: non-resonant ultraintense x-ray interactions

- Ultraintense x-ray interactions – multiphoton processes rule!
 - establish sequential single photon absorption as dominant ionization mechanism
 - fully stripped neon: six-photon, ten-electron ($\sim 10^{12}/\mu\text{m}^2$)
 - multiple photon absorption probability high when fluence $> 1/\sigma$
 - controlled electron stripping (outer v inner shells)
- X-ray induced transparency – a general phenomena
 - transient x-ray transparency caused by ejection of inner-shell electrons
 - induced transparency = frustrated absorption = core-level bleaching
 - molecules: Hoener *et al.*, PRL **104**, 253002 (2010)
 - clusters: Schorb *et al.*, PRL **108**, 233401 (2012)
 - solids: Yoneda *et al.*, Nat. Comm. (2014), Rackstraw *et al.*, PRL (2015)
 - implications for imaging: $\sigma_{\text{scatt}}/\sigma_{\text{abs}}$ is increased
- Femtosecond time-scale atomic processes provide FEL diagnostics





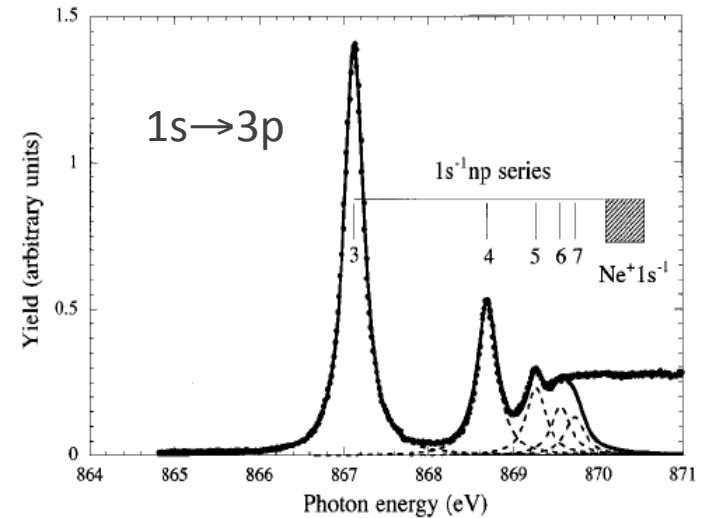
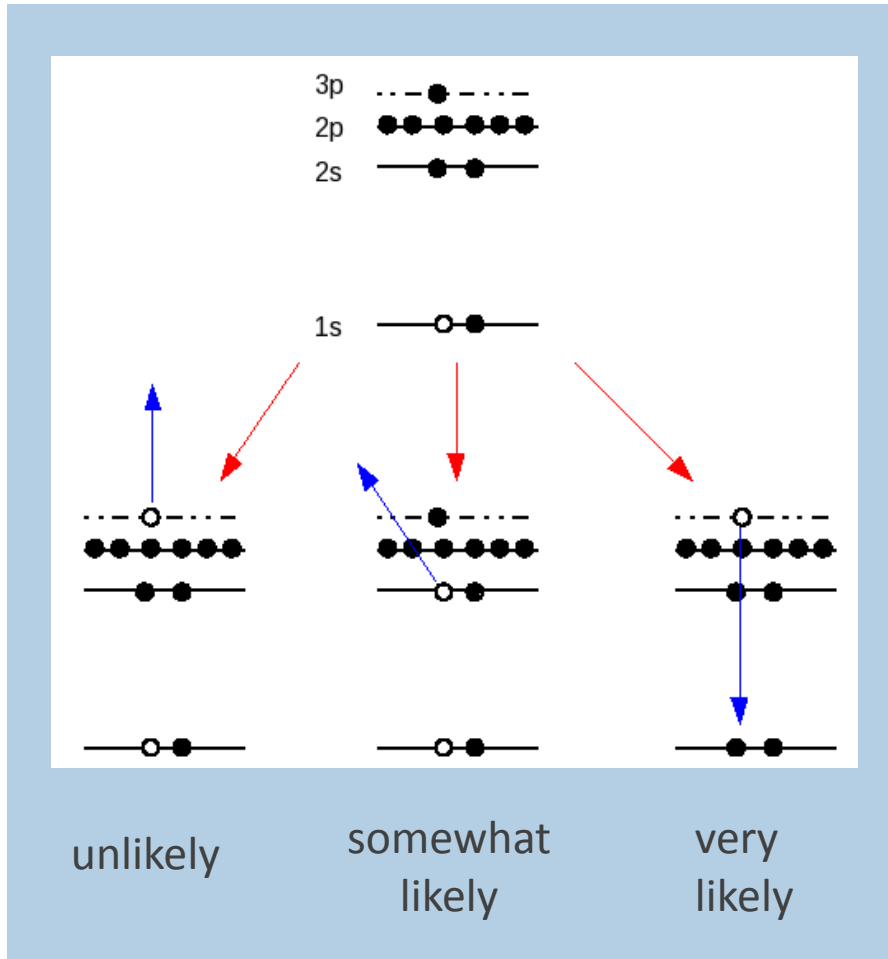
LCLS Experiment 5

Resonant x-ray processes at high intensity



Can we control inner-shell electron dynamics?

“Rabi flopping” may inhibit Auger decay & x-ray damage.



- Strong 1s to 3p resonance

$$\mu_{Ne\ 1s-3p} = 0.01\ ea_0$$

$$\tau_{Ne\ 1s^{-1}} = 2.4\ fs = 100\ a.u.$$

- Rabi flopping possible

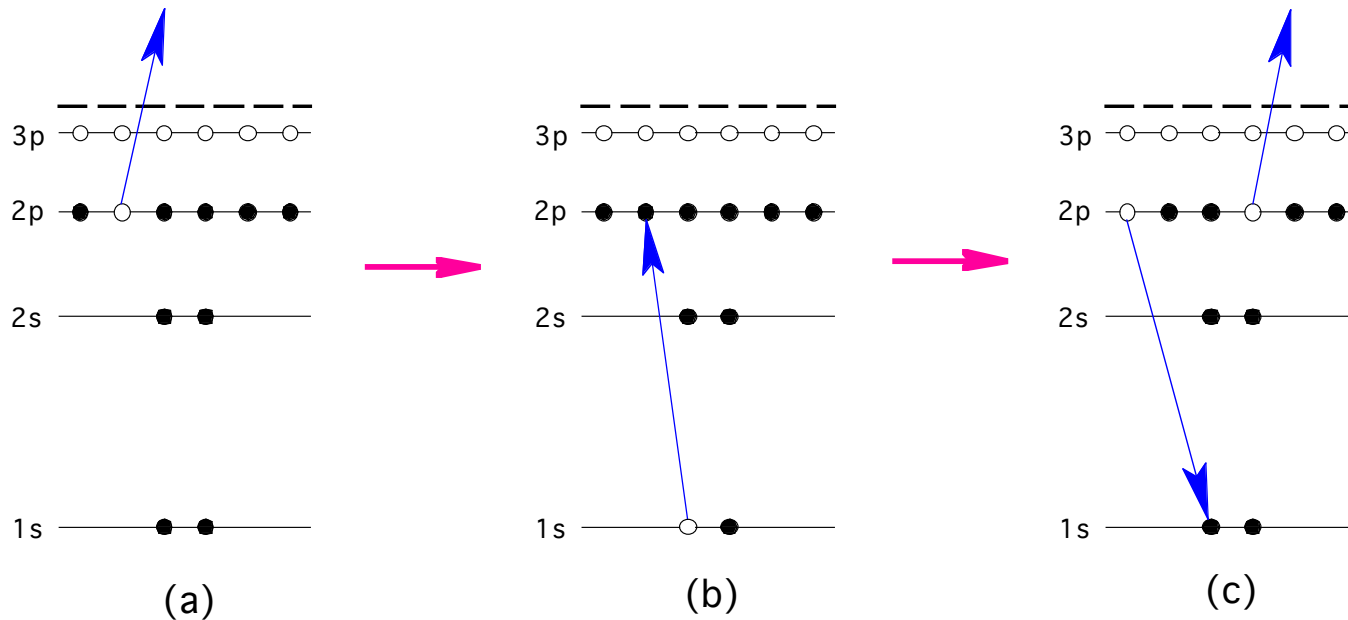
$$E_{Ne} \sim 6.3\ a.u.$$

$$I_{Ne} \sim 1.4 \times 10^{18}\ W/cm^2$$

But LCLS linewidth $\sim 8\ eV!$



Rabi-flopping on 1s - 2p resonance more feasible

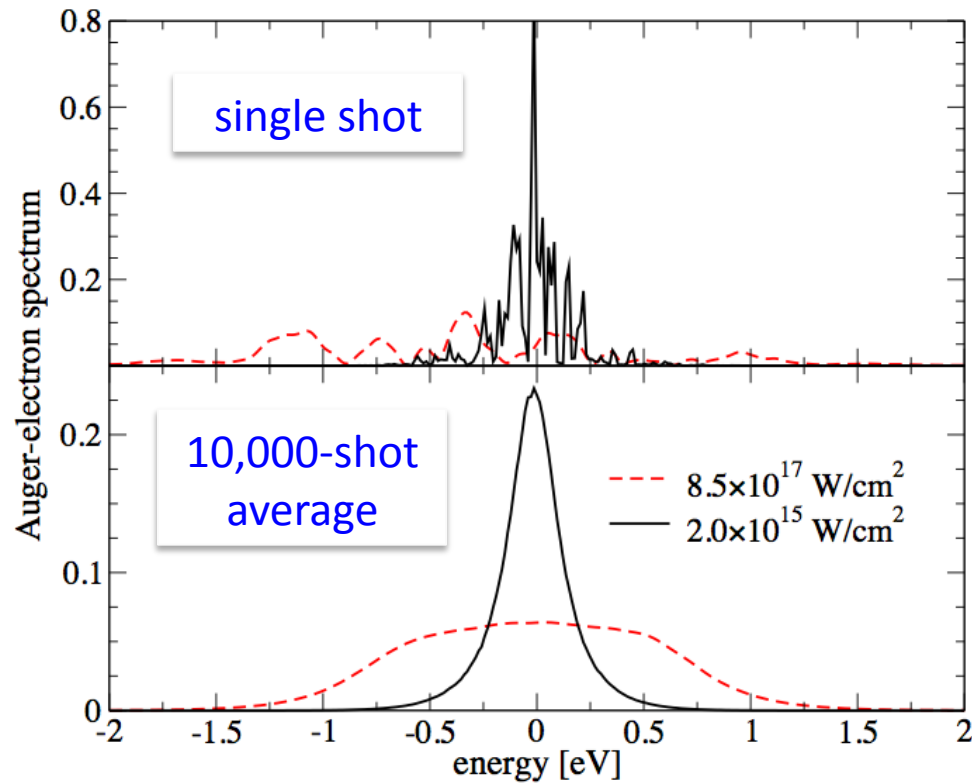


$$E_{x\text{-ray}} = 848.6 \text{ eV}$$

$$\sigma_{1s-2p} = 500\sigma_{2p-\infty} = 30 \sigma_{1s-3p}$$

Observe Auger yield when x-rays scanned over 1s - 2p resonance.
 Observe broadening at resonance to indicate Rabi flopping
 Theory: Rohringer & Santra PRA (2008).

Calculated “Resonant Auger effect at high x-ray intensity”

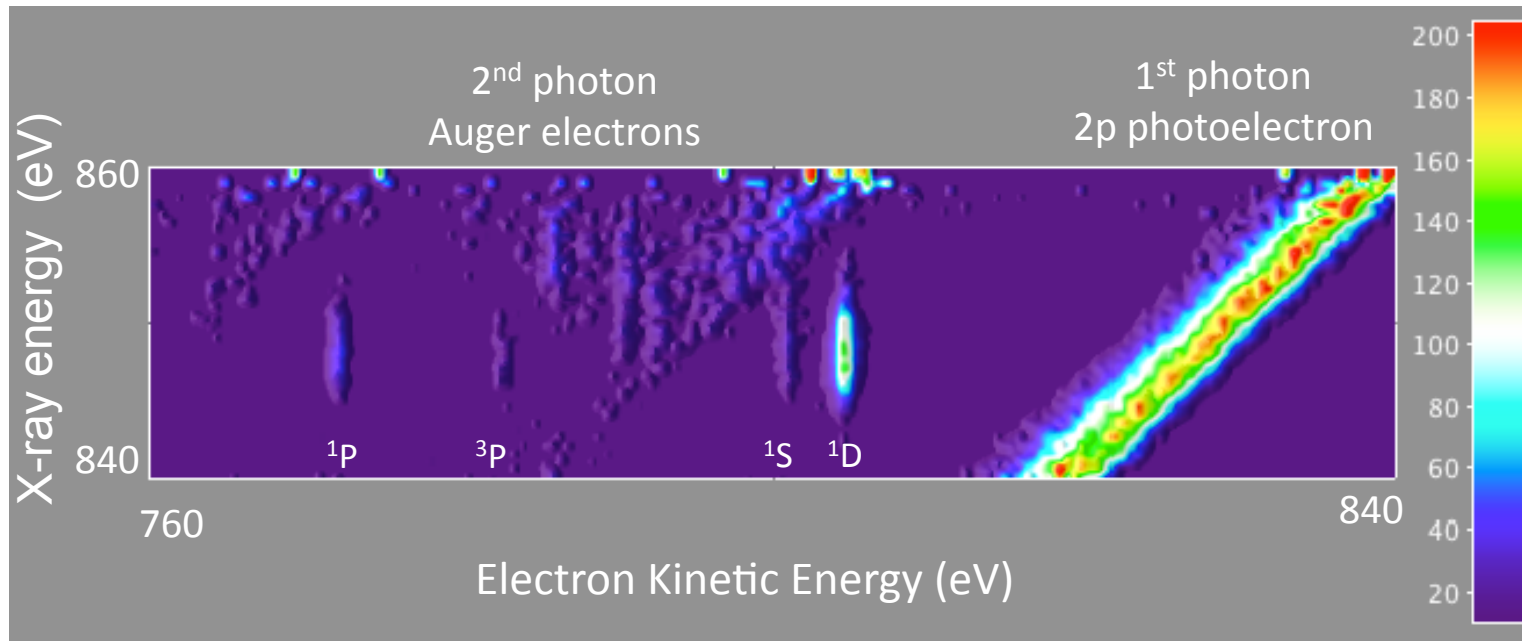


-> Look for Auger line broadening on resonance

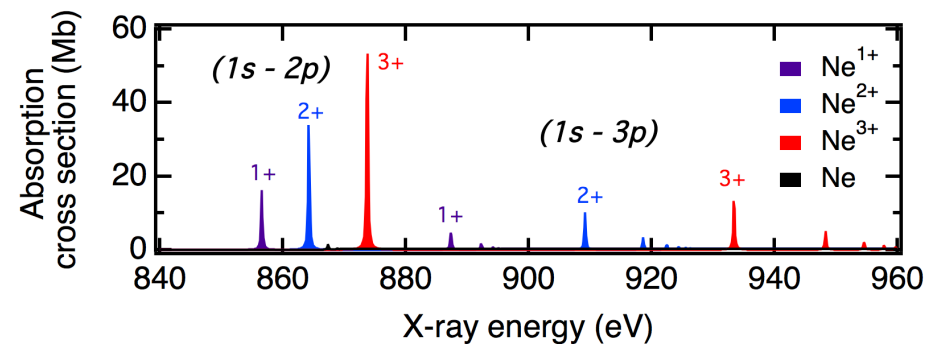
N. Rohringer & R. Santra, PRA 77, 053404 (2008)



Looking for Rabi flopping: unveiling and driving hidden resonances with LCLS pulses



- High fluence pulse alters target to reveal enormous “hidden” resonances ~1000x larger than background
- X-ray absorption spectrum changes during the fs duration pulse



Characterization of SASE x-ray pulse required

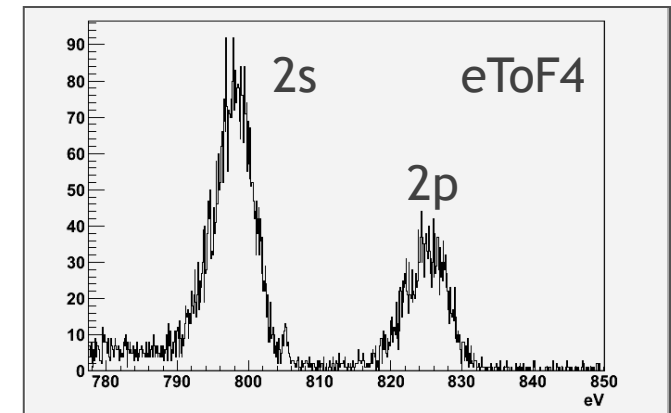
- Shot-to-shot photon energy jitter

- Photoelectron energy spectra give x-ray energy centroid

- Jitter derived from shot-to-shot GeV electron beam energy measurements

$$E_x \text{ (eV)} = 44.25 E_e \text{ (GeV)}^2$$

Conditions	FWHM photon energy jitter (eV)
40 pC (<10 fs) 850 eV 0.3 mJ 4500 A	4.25
250 pC (100 fs) 787 eV 1.5 mJ 2500 A	4.79
250 pC (100 fs) 769 eV 1.5 mJ 2500 A	5.24



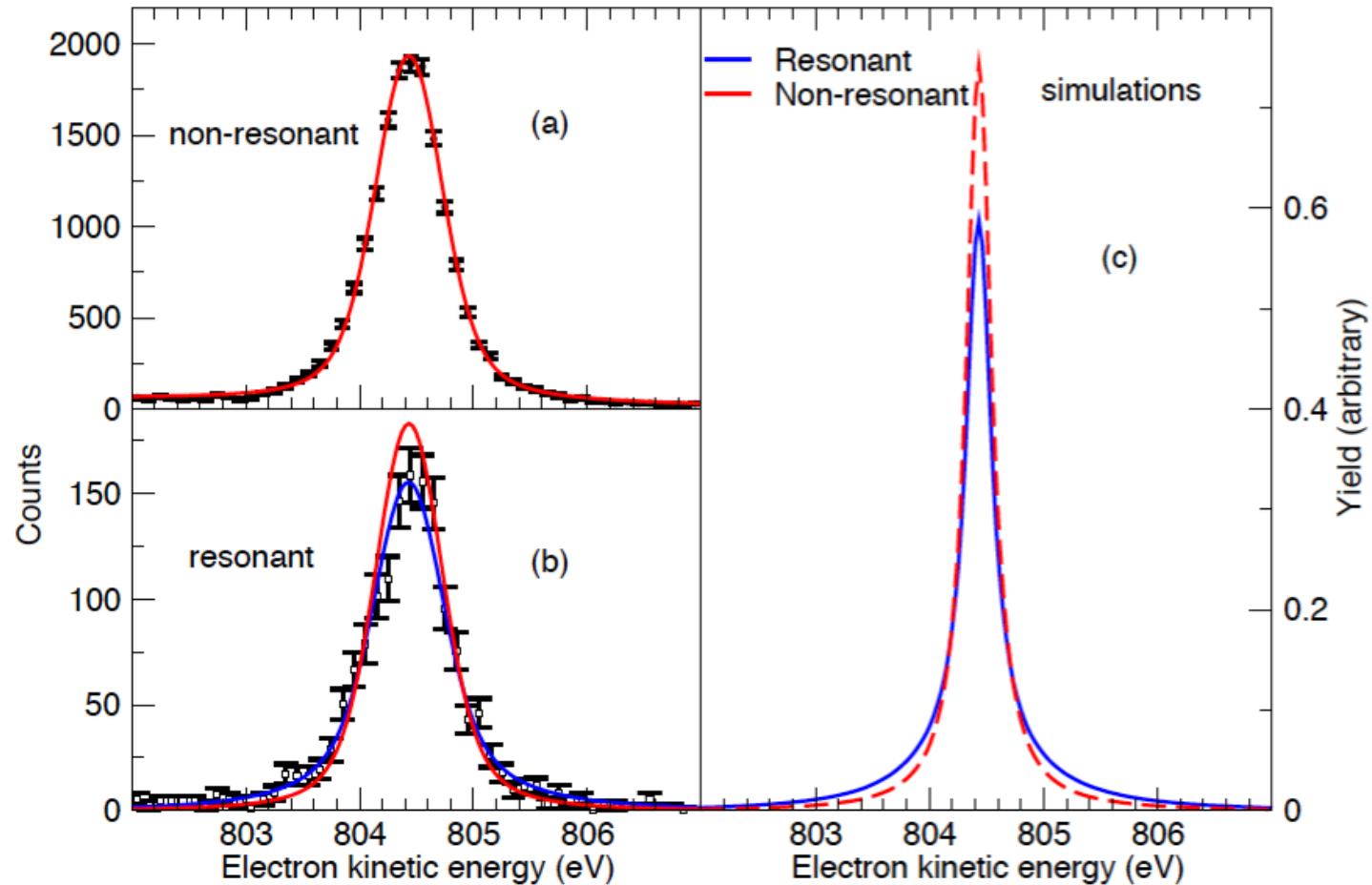
Electron kinetic energy (eV)

- Intrinsic x-ray bandwidth

Conditions	Intrinsic x-ray pulse bandwidth (from 2s photopeak) (eV) (FWHM)	Intrinsic x-ray pulse bandwidth (from 2p photopeak) (eV) (FWHM)	Average bandwidth (eV) (FWHM)	%
40 pC (<10 fs) 850 eV 0.3 mJ 4500 A	4.3	4.5	4.4	0.5 %
250 pC (100 fs) 787 eV 1.5 mJ 2500 A	7.1	7.8	7.45	0.9%
250 pC (100 fs) 769 eV 1.5 mJ 2500 A	7.7	7.8	7.77	1%



Is the 1D Auger line broadened on 1s-2p resonance?

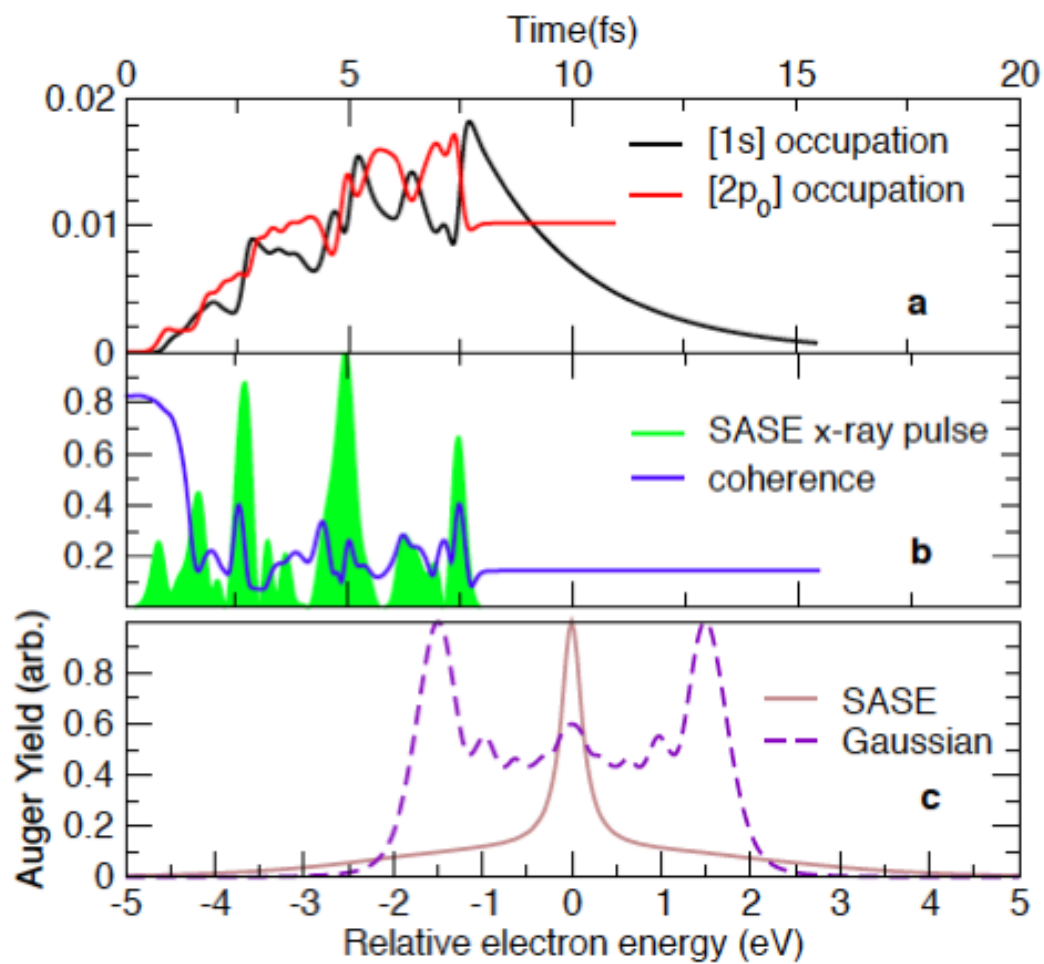


E.P. Kanter *et al.*, PRL (2011)

Theory from N. Rohringer and R. Santra



SASE vs Gaussian pulse for Rabi flopping



Summary: resonant x-ray processes at high intensity

- First hint of Rabi cycling for inner-shell electrons: Ne 1s – 2p resonance
- Need XFEL with improved longitudinal coherence – SEEDING
 - Quantum control – multidimensional spectroscopies
 - Single particle imaging (reduced radiation damage & increased x-ray intensity)
- “Hidden” resonances critical in atomic response to ultraintense x-rays
 - Enhanced two-photon absorption probability
 - Doumy *et al.*, PRL **106**, 083002 (2011)
 - Ionization beyond sequential single photon model
 - Schorb *et al.*, PRL **108**, 233401 (2012) - Ar
 - Rudek *et al.*, Nat. Phot. **6**, 858 (2012) – Xe
 - Rudek *et al.*, Phys. Rev. A **87**, 023413 (2013) – Kr





Improving the X-ray Laser



SASE pulses, baseline mode of operation: poor longitudinal coherence

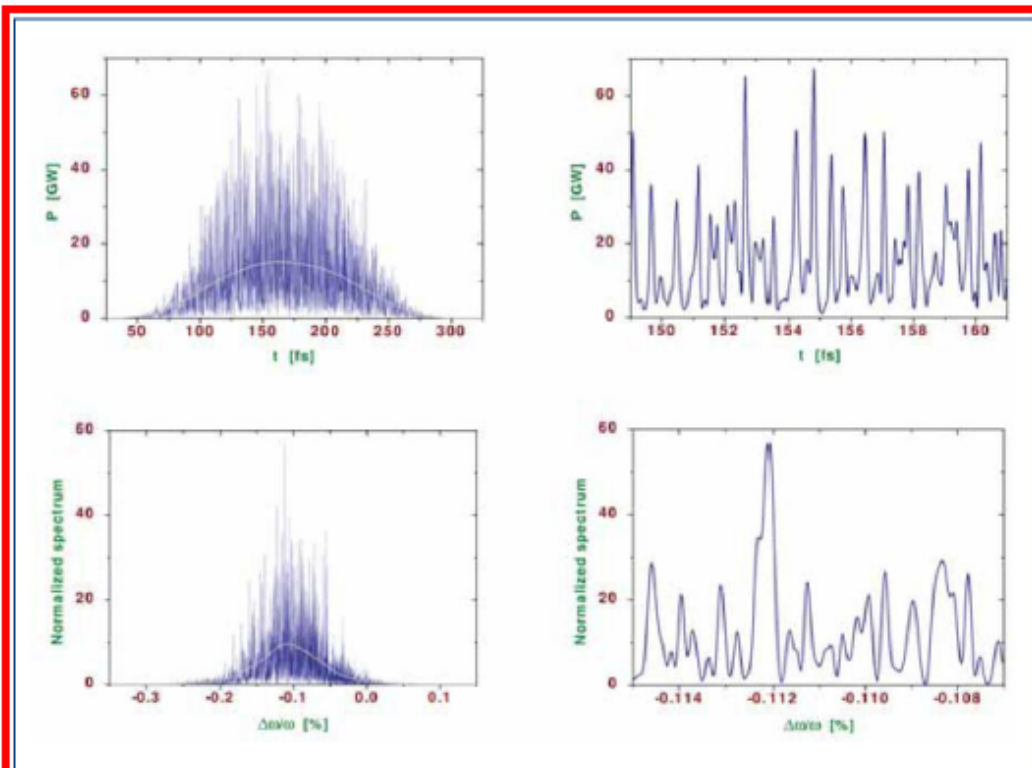


Figure 5.2.4 Temporal (top) and spectral (bottom) structure for 12.4 keV XFEL radiation from SASE 1. Smooth lines indicate averaged profiles. Right side plots show enlarged view of the left plots. The magnetic undulator length is 130 m.

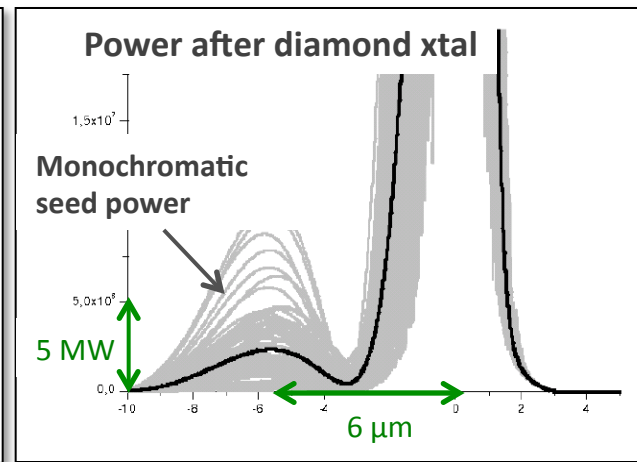
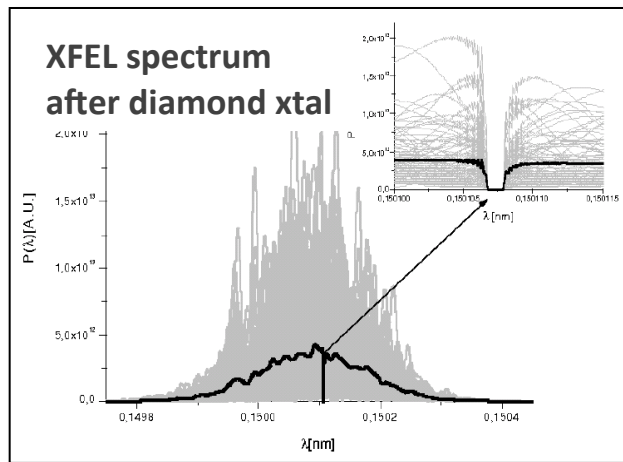
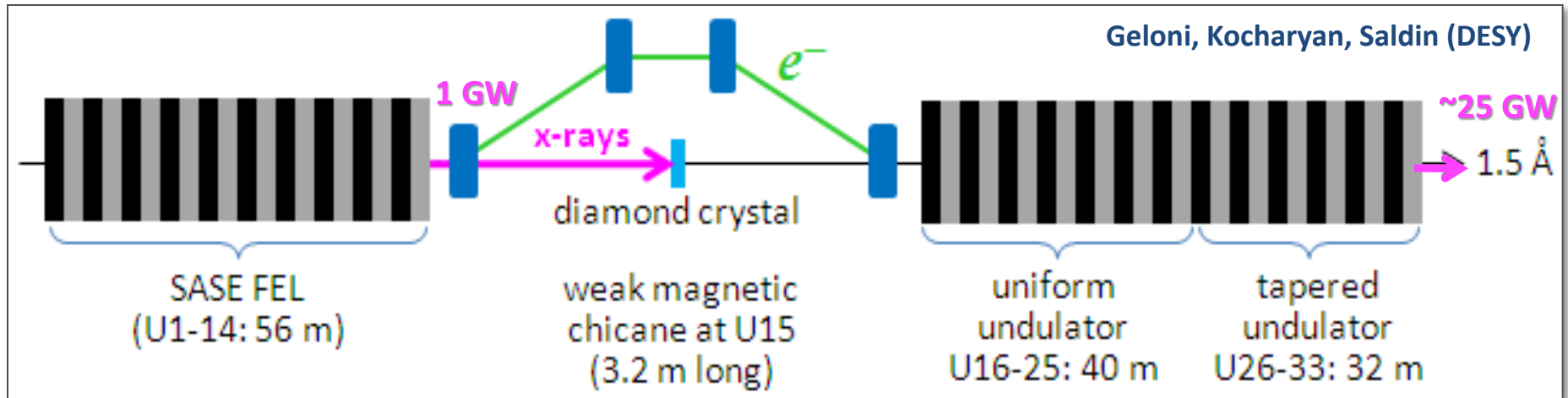
Source: The European XFEL TDR – DESY 2006-097 (2006)

$$\frac{\Delta\omega}{\omega} \sim 2\rho \sim 10^{-3}$$

$$\left(\frac{\Delta\omega}{\omega}\right)_{\text{spike}} \sim \frac{1}{\sigma_T \omega} \sim 10^{-5}$$

- Hundreds of longitudinal modes
- A lot of room for improvement
- Self-seeding schemes answer the call for increasing longitudinal coherence

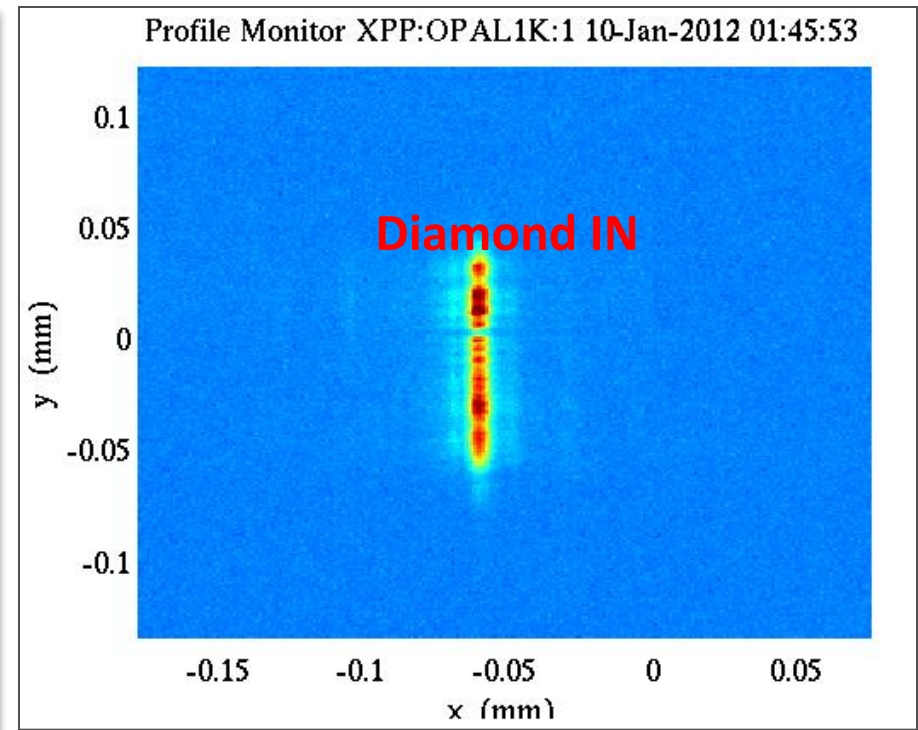
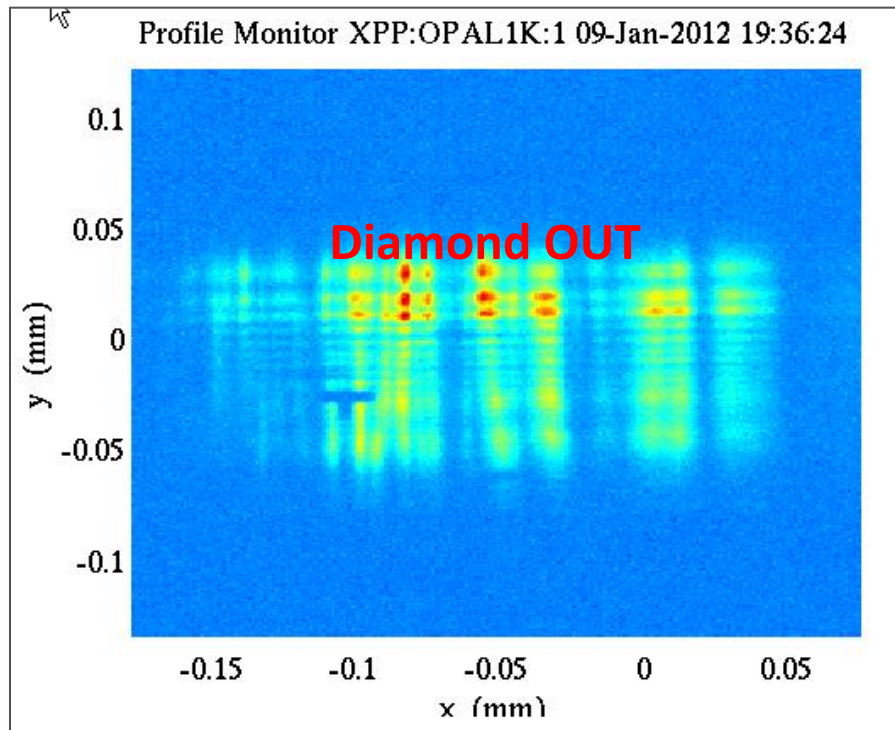
Hard x-ray self-seeding proposed 2010



Diamond C(004): 100 μm
 $\lambda = 0.15 \text{ nm}$, $\theta_B = 57^\circ$



Hard x-ray self seeding realized Jan 2012 - P. Emma et al.



Bandwidth $<10^{-4}$ at 8-9 keV and tunable

But ... did not achieve saturation and power jitter still present

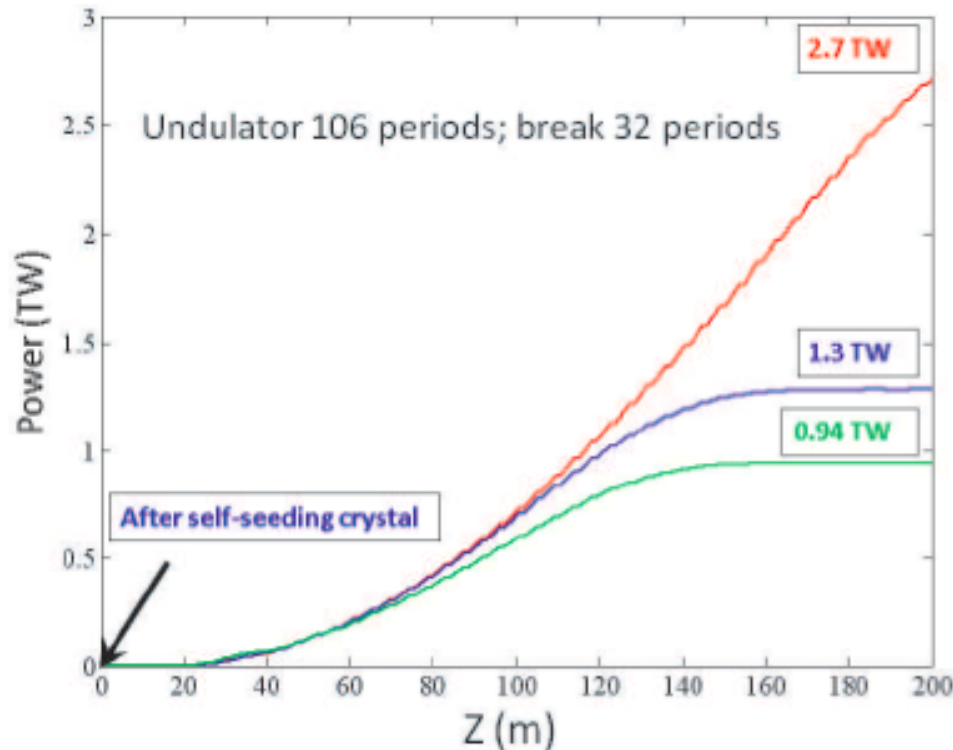
Three lab collaboration:



J. Amann *et al.*, Nature Photonics 6, 693 (2012)

On the road to a TW FEL : LCLS-TN-11-3

GENESIS simulation



200 m undulator tapered

10^{13} x-rays @ 8-10 keV
10 fs

W.M. Fawley¹, J. Frisch¹, Z. Huang¹, Y. Jiao¹, H.-D. Nuhn¹, C. Pellegrini^{1,2}, S. Reiche³, J. Wu^{1†}

¹SLAC National Accelerator Laboratory, Menlo Park, CA 94025, USA

²Department of Physics and Astronomy, UCLA, Los Angeles, CA 90095-1547, USA

³Paul Scherrer Institute, Villigen PSI, 5232, Switzerland

Also arXiv Jun 2013: S. Serkez *et al.* 10 TW FEL, 10^{14} x-rays, 10 fs @ 3.5 keV





Towards single particle imaging



LETTER

doi:10.1038/nature09748

Single mimivirus particles intercepted and imaged with an X-ray laser

Seibert et al., *Nature* 470, 78 (2011)

PRL 114, 098102 (2015)

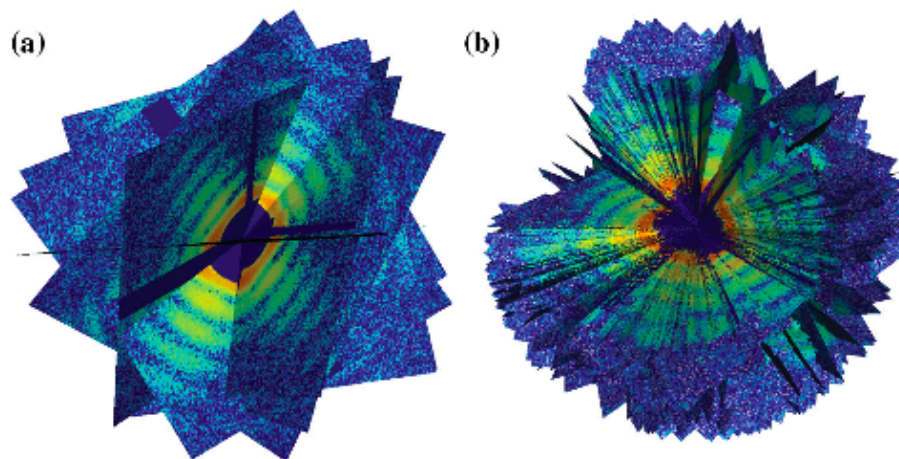
Selected for a *Viewpoint in Physics*
PHYSICAL REVIEW LETTERS

week ending
6 MARCH 2015



Three-Dimensional Reconstruction of the Giant Mimivirus Particle with an X-Ray Free-Electron Laser

Ekeberg et al., *PRL* (2015)



Mimivirus

- Largest known virus – 0.75 μm
 - Does not crystallize
 - Large for 3D cryoelectron microscopy
- Single Shot Scattering Pattern
2D: 32 nm resolution
Set of 198 scattering patterns
3D: reconstruction to 120 nm



Viewpoint

X-Ray Imaging of a Single Virus in 3D

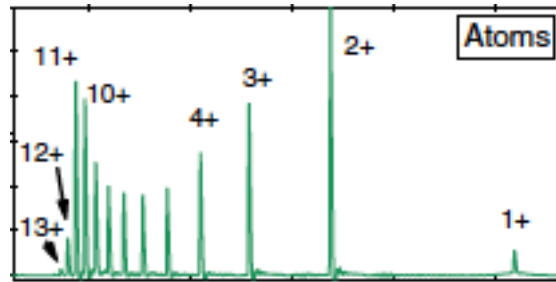
“And there are still open questions on the impact of electronic damage on x-ray scattering on femtosecond time scales: The above-mentioned work by Neutze et al. tracked the movements of the atomic nuclei of the biomolecule, showing they don't move on the few-femtosecond timescale of an x-ray pulse. Electrons, however, move faster than nuclei. Since electrons are what scatters x rays, it is yet to be confirmed that few-femtosecond pulses can probe an unperturbed electronic structure.” - Keith Nugent

March 2, 2015



Beyond the sequential single photon ionization model

Argon

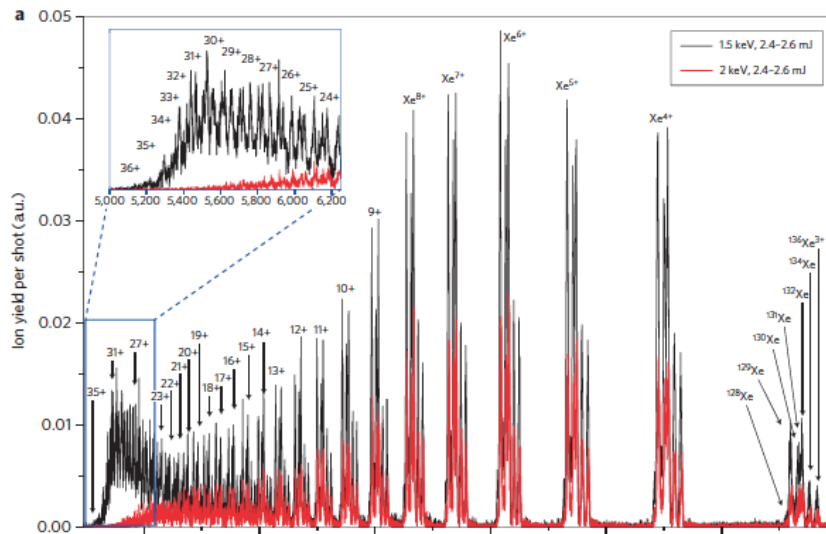


@ 480 eV

sequential single photon limit 10+
observe 13+

Schorb et al., PRL (2012)

Xenon



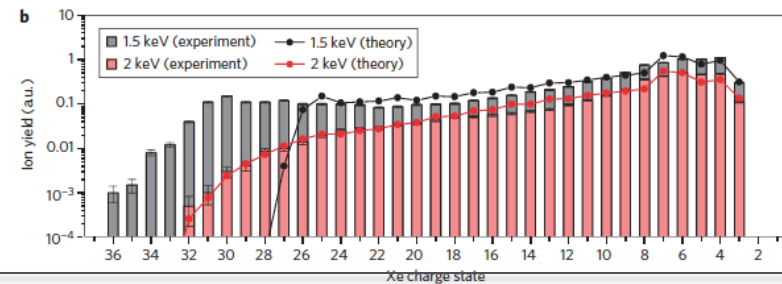
@1500 eV

sequential single photon limit 27+
observe 36+

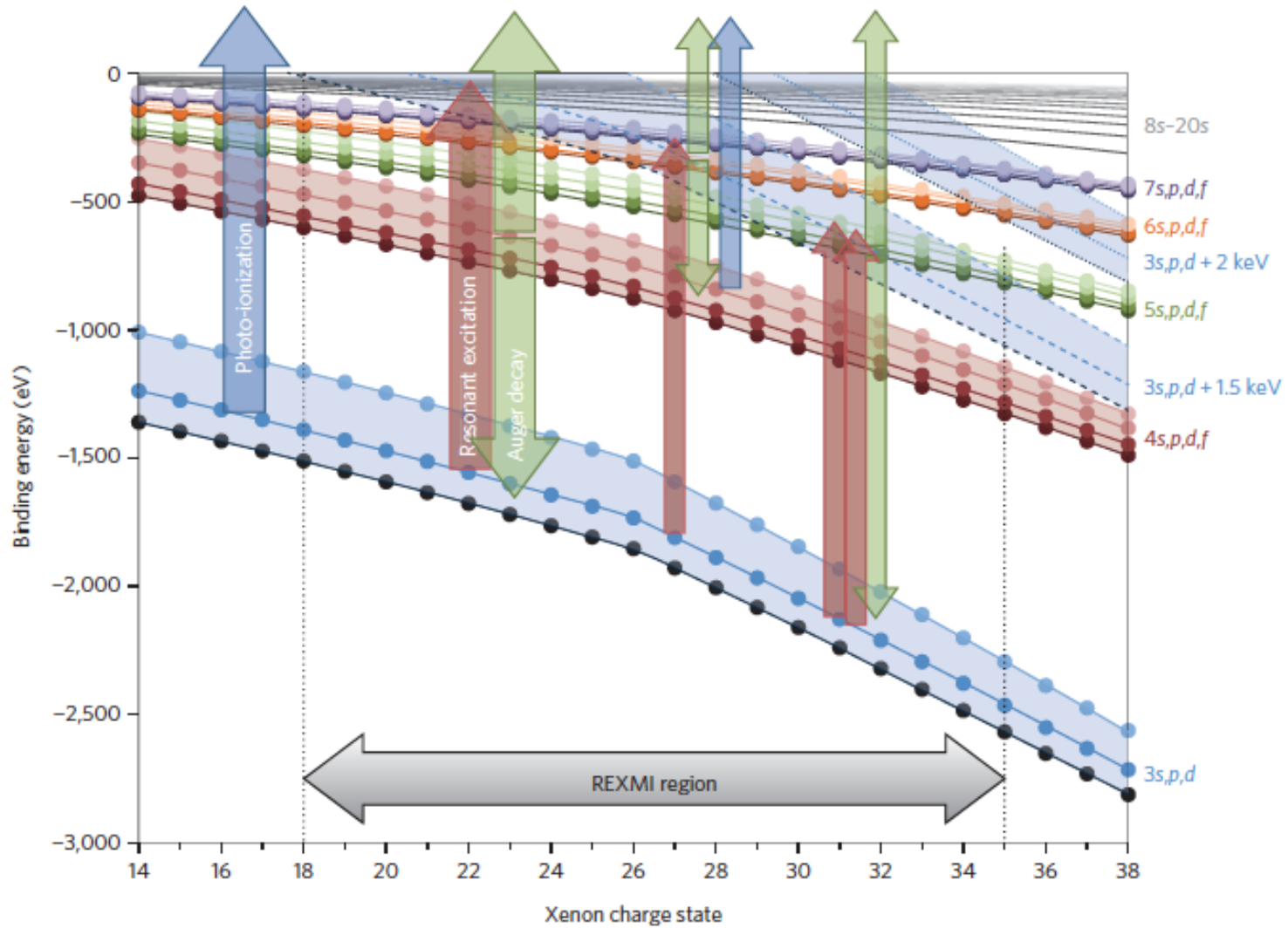
@2000 eV

sequential single photon limit 32+
observe 32+

Rudek et al., Nat. Pho. (2012)



Resonance-enabled x-ray multiple ionization



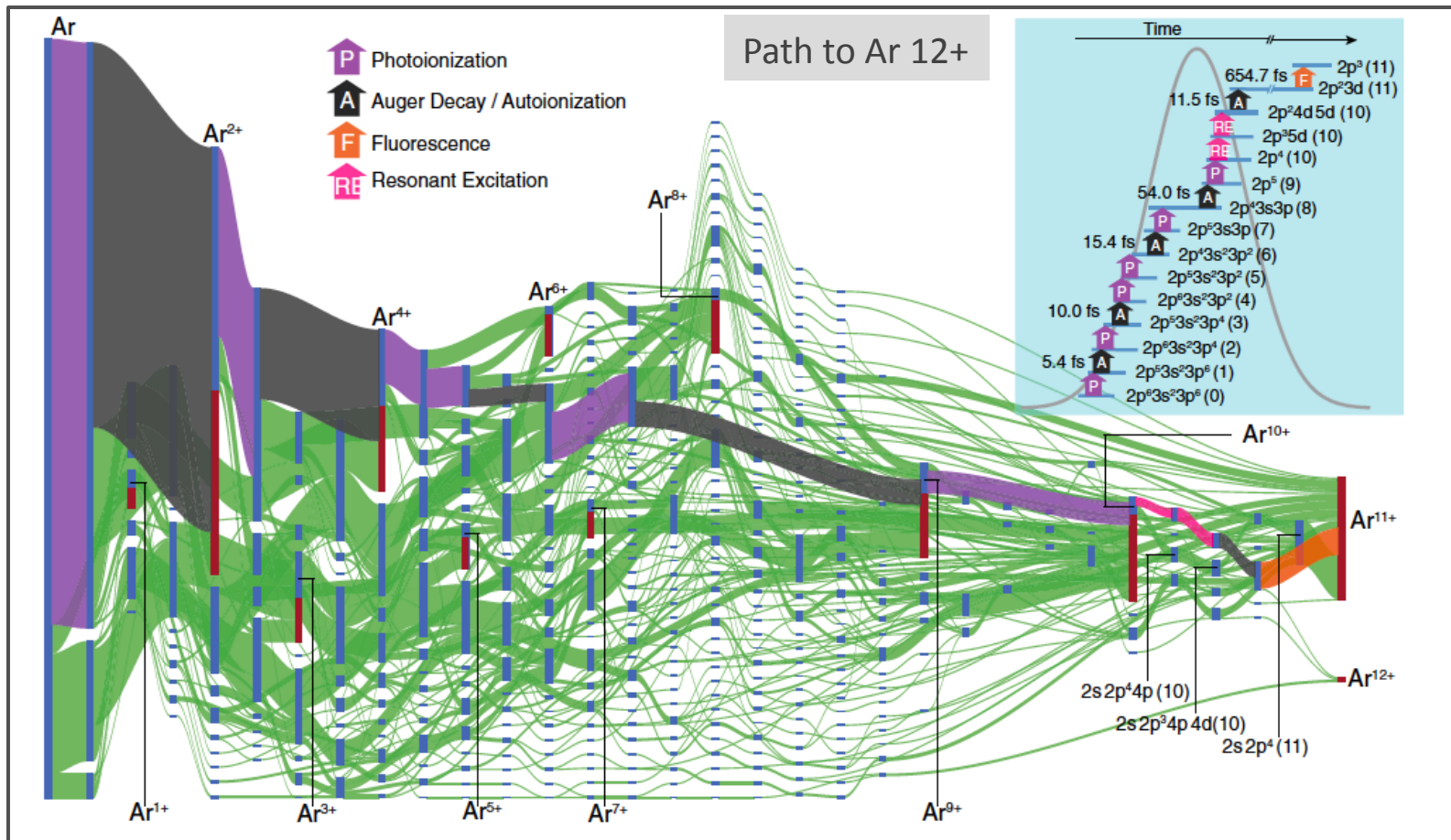
B. Rudek *et al.*, Nat. Phot. (2012)



Tracking electronic configurations during XFEL pulse including resonances!

Monte Carlo Rate Equation Approach

	# of ECs with no RE	# of ECs with RE
Ar	1.33×10^3	2.85×10^{13}
Kr	3.05×10^5	2.08×10^{19}
Xe	7.06×10^7	9.05×10^{22}



Ultraintense hard x-ray interactions: seeded v SASE

CXI endstation experiment Apr 2014: $I > 10^{20} \text{ W/cm}^2$

5.5 – 8.3 keV, 2 mJ, 30-40 fs

“100 nm” focus

$10^5 \text{ x-rays/\AA}^2$

...but no shot-by-shot spectral monitor

Observations:

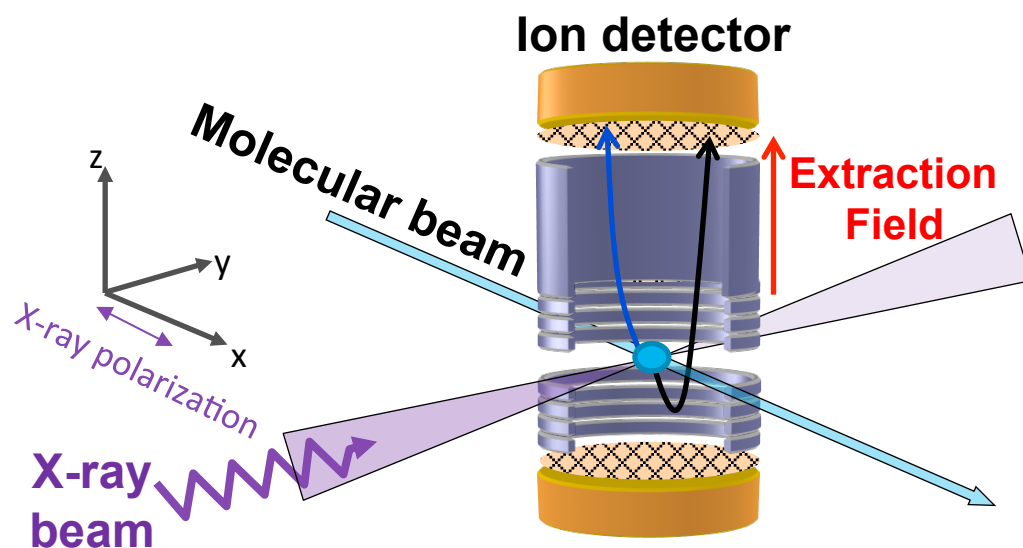
Ar^{18+} : fully stripped

Kr^{34+} : two 1s electrons

Xe^{48+} : 6 electrons

Largely consistent with sequential single photon

CH_3I : sum of charges 55+



Calculations by Phay Ho:

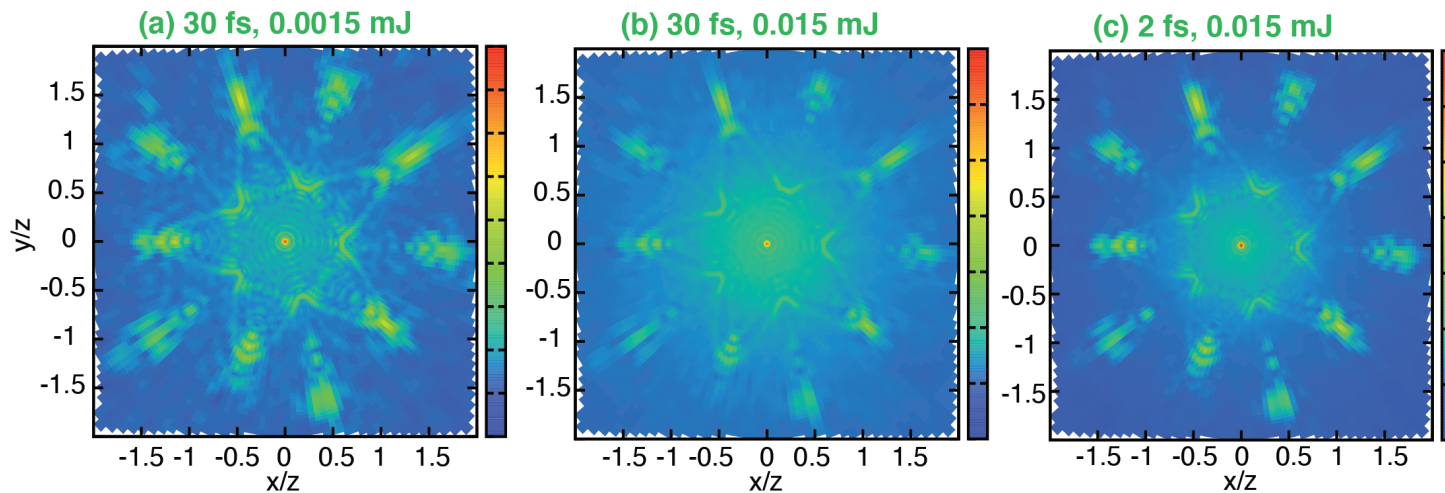
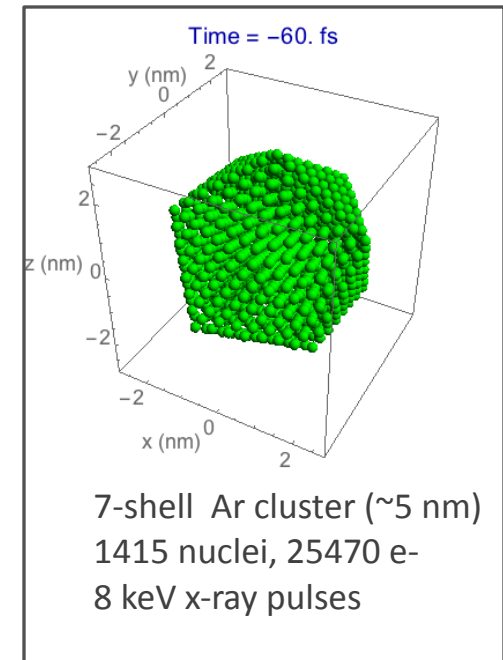
Identified photon energies where seeded pulses reduce absorption & radiation damage relative to SASE pulses.

Joined/Artem Rudenko, Daniel Rolles & team



High Intensity X-ray Imaging of Nanosystems

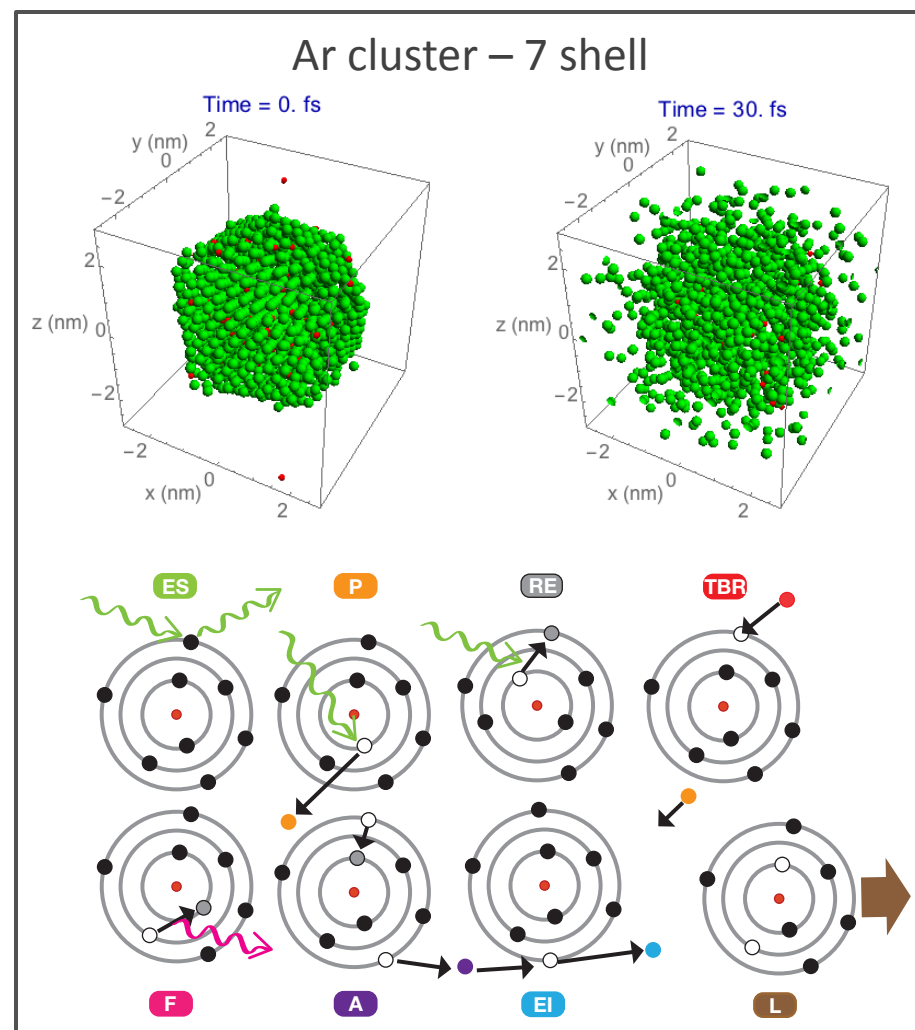
- Intense x-ray pulses lead to efficient stripping of e^- from atoms/ions and significant displacement of e^- , ions/atoms from equilibrium during the pulse – nanoplasma
- Our MC/MD model includes: Elastic Scattering, Photoionization, Auger, Fluorescence, Resonant excitation, Three-body recombination, Electron impact ionization, Lattice dynamics
- Investigate the impact of electronic damage on various observables (ion, photoelectron, Auger and fluorescence spectra, x-ray diffraction)
- Recorded diffraction patterns depend strongly on pulse parameters.



AMO “solutions” at the ultraintense x-ray frontier

- Fundamental nature of x-ray damage at high intensity
 - Coulomb explosion
 - electronic damage
 - behavior at 10^{22} W/cm² – 1Å
 - nanoplasma formation
- Expt’l AMO observables:
 - Ion, photoelectron, Auger, fluorescence and ...
 - x-ray diffraction pattern
- Computational studies on large scale systems

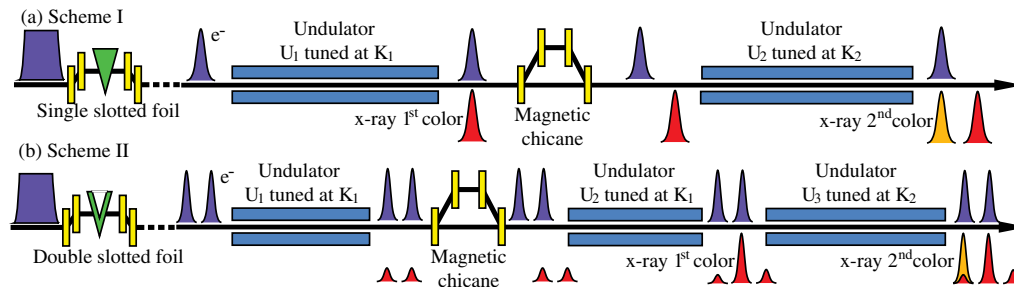
Phay Ho
Chris Knight



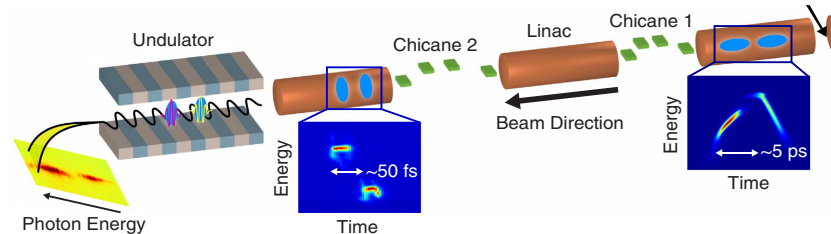
Atomistic computational studies based upon Monte Carlo/Molecular Dynamics

New XFEL capabilities becoming available

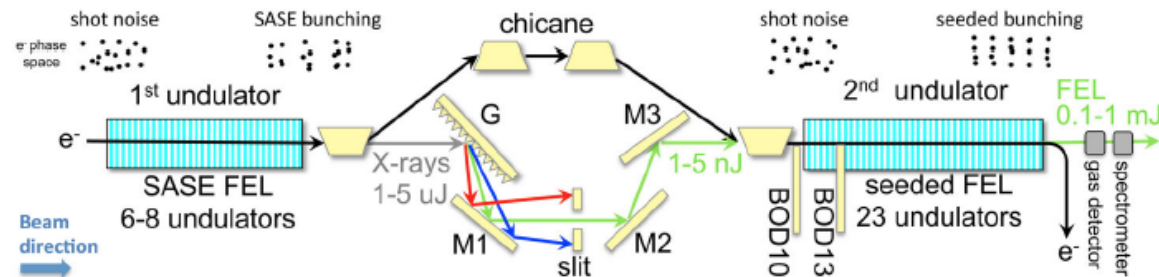
One bunch – two pulse – two color (Lutman *et al.*, PRL **110**, 134801 (2013))



Two bunch – two pulse – two color (Marinelli *et al.*, Nat. Comm. **6**, 6369 (2015))



Seeded soft x-ray beam (Ratner *et al.*, PRL **114**, 054801 (2015))



International hard X-ray FELs here and on the horizon

SACLA
Image Gallery

2011



PAL POHANG ACCELERATOR
LABORATORY

2016



SwissFEL

2017



European
XFEL

2017



Summary

- Our atomic physics expts have established fundamental understanding of the response of matter to ultraintense XFEL irradiation
 - sequential single photon ionization dominates
 - intensity-induced x-ray transparency (frustrated absorption)
 - intense x-rays can “control” inner-shell electron dynamics
 - resonances can be critical in XFEL interactions
- This fundamental understanding will aid in the quest for single molecule imaging and other applications, e.g. high energy density matter
 - AMO methods (ion, electron, photon) in concert with computational studies promise fuller understanding of radiation damage in extended systems
- Future is bright with well-characterized ultraintense x-ray lasers, multiple pulse configurations ... with 1000x higher intensity.

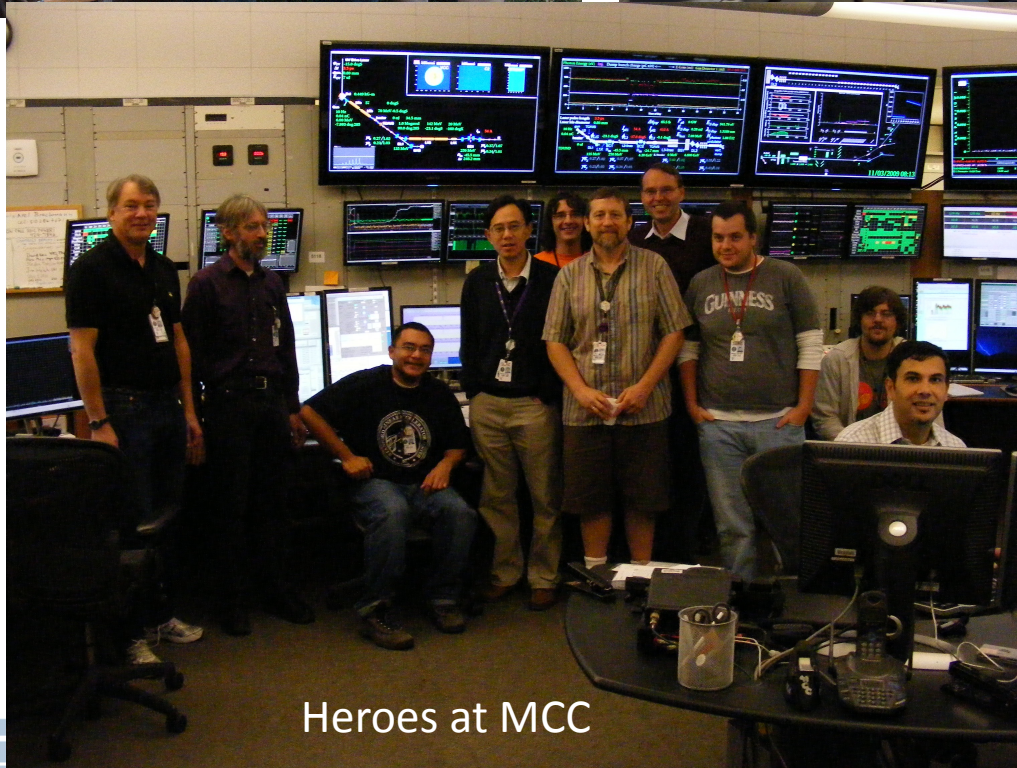




Argonne AMO group Oct 2009



Heroes at AMO Control



Heroes at MCC

