

JUNE '09

STOCKHOLM

BARYOGENESIS IN VARIANTS OF THE ELECTROWEAK PHASE TRANSITION

MICHAEL G. SCHMIDT
HEIDELBERG

(Almost)
at the beginning
of this workshop
!

- SOME GENERAL PERSPECTIVES
- OVERVIEW OF SOME OLD AND MORE RECENT MODELS

PHASE TRANSITIONS IN THE EARLY UNIVERSE
ARE MOST INTERESTING BECAUSE THEY MIGHT HAVE
LEFT TRACES EXPLAINING THE PRESENT UNIVERSE.

CANDIDATES:

- QCD PHASE TRANS.

BUT: CROSS OVER BEHAVIOR UNLESS LARGE μ (BARYON DENSITY)
 \leadsto NEUTRON STARS...?

- ELECTROWEAK PHASE TRANS.

THE SUBJECT OF THIS WORKSHOP AND OF THIS TALK \rightarrow

- INFLATION / GUT-PHASE TRANS. (...HYBRID INFLATION)

MOST IMPORTANT FOR STRUCTURE FORMATION

BUT: AT 10^{15} GEV ? HARD TO CONTROL BY LABORATORY EXP.

\leadsto LOW SCALE INFLATION?

? THE ELECTROWEAK PHASE TRANSITION : $\langle H \rangle = 0 \Rightarrow \langle H \rangle \neq 0$
MIGHT PREDICT

- GRAVITATIONAL WAVES
- SEEDS FOR LARGE SCALE MAGNETIC FIELDS
- PRODUCTION OF A BARYON ASYMMETRY

} HOPE TO
LEARN MORE
ABOUT IN THIS
WORKSHOP

THE LATTER IS JUST ONE NUMBER, BUT MOST IMPORTANT
FOR OUR UNIVERSE

$$\eta = \frac{N_B - N_{\bar{B}}}{N_\gamma} = (6.1 \pm \frac{0.4}{0.3}) \cdot 10^{-10}$$

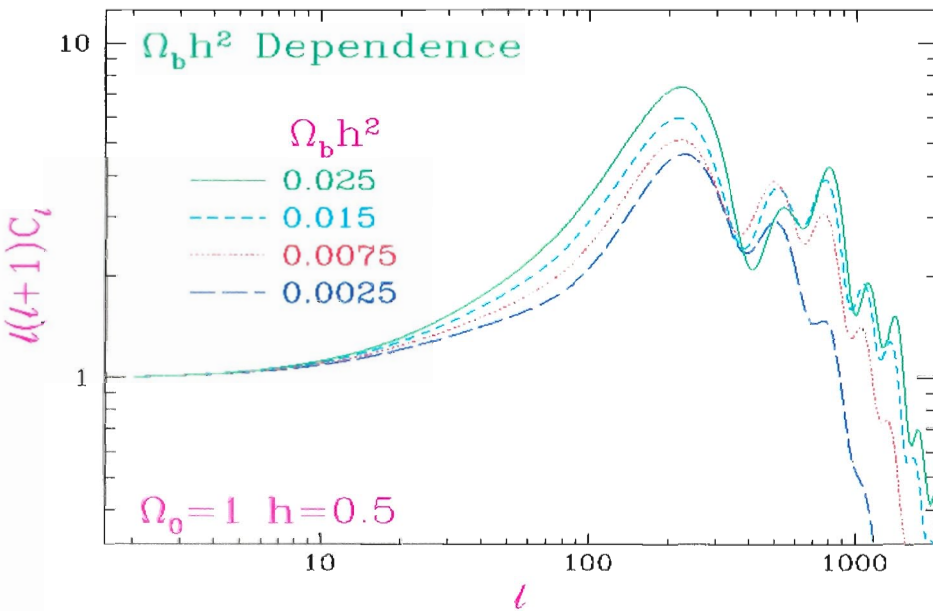
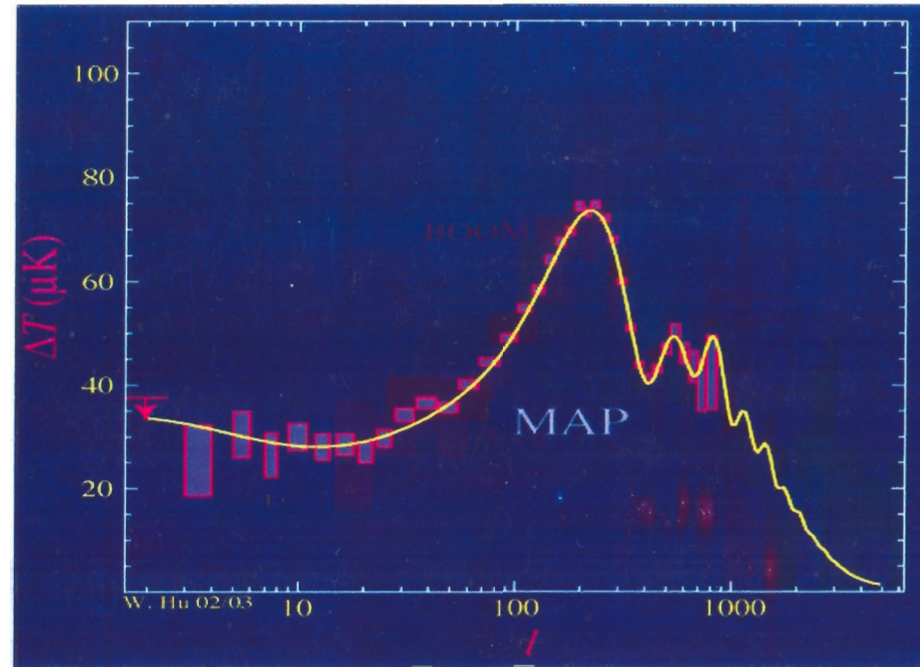
FROM WMAP

IN AGREEMENT WITH
PRIMORDIAL NUCLEOSYNTHESIS

ITS EXPLANATION IS STRONGLY RELATED TO ELEMENTARY PARTICLE PH.
(SM AND "BEYOND"!))

- IT HAS TO BE GENERATED IN THE EARLY COSMOS AFTER INFLATION

Baryonic matter and cmb



baryons: increase compression (odd) peaks, decrease rarefaction peaks

INFLATION

- SOLVES CAUSALITY PROBLEM IN BIG BANG THEORY
- $\Omega = 1$ ($\rho = \rho_c$) FLAT UNIVERSE
- CREATES FLUCTUATIONS LEADING TO STRUCTURE FORM.

NO PARTICLES LEFT AFTER EXPON. GROWTH

→ HAVE TO CREATE BARYONASYMMETRY

WELLKNOWN SAKHAROV NECESSARY CRITERIA

- B - VIOLATION ✓ ;
- C, CP - VIOLATION :

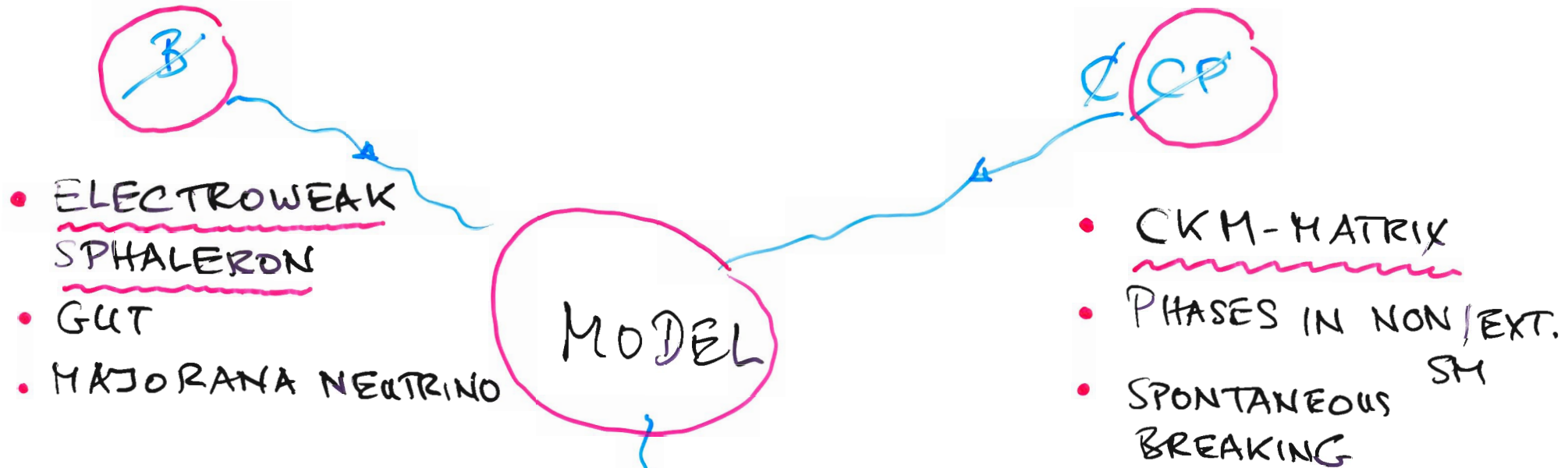
IN THE ABSENCE OF PREFERENCE FOR MATTER/ANTIMATTER:

B NONCONS. REACTIONS WILL PRODUCE B AND \bar{B} AT THE SAME RATE ;

- NON EQUILIBRIUM : CHEMICAL POT. FOR NONCONSERVED QUANTUM-N. VANISHES ; $\mu_B = \mu_{\bar{B}}$ (CPT) → SAME THERMAL DISTRIBUTION

• BARYOGENESIS

SAKHAROV '67



POSSIBLE IN
SM - ELECTROWEAK
PHASE TRANSITION?

SHAPOSHNIKOV '87

- EXPANDING UNIVERSE
- OUT OF EQUILIBRIUM DECAY
- PHASE TRANSITION

• SPHALERON TRANSITION

$SU(2)_L$ GAUGE THEORY IN SM VIOLATES $(B+L)$

($B+L$ CONSERVED!)

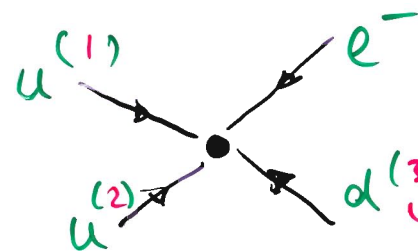
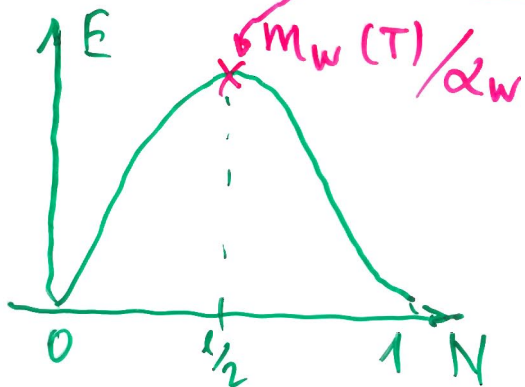
$$\partial_\mu j^\mu B = \frac{g^2 N}{4\pi^2} \epsilon_{\alpha\beta\gamma\delta} \mathbf{F}_{\alpha\beta} \cdot \mathbf{F}_{\gamma\delta}$$

$$\frac{d}{dt} \int d^3x j_0^B = \frac{d}{dt} \int d^3x K_0$$

CHERN-SIMONS NUMBER N_{CS}

⇒ INSTANTON INDUCED TUNNELING IN N_{CS} ($N \rightarrow N+1$)

⇒ THERMAL SPHALERON TRANS.



EFF. ACTION (1 GENER.)

Color
 $\sim e^{-8\pi^2/g_w^2}$
 SHALL!

$$\Gamma_B \sim \dots \alpha_w (\alpha_w T)^4 \exp\left(-\dots \frac{V(T)}{T}\right)$$

UN SUPPRESSED IN SYMMETRIC (HOT) PHASE
 EQUILIBRATES INT-EQUILIBRIUM ($\rightarrow B+L=0$)

BÖDEKER, MOORE, RUMMANNEN

IN THERMAL TRANSITION FROM "HOT" **SYMMETRIC PHASE**: $\langle H \rangle = 0$
TO **HIGGS PHASE**: $\langle H \rangle \neq 0$
?
 $V(T)$

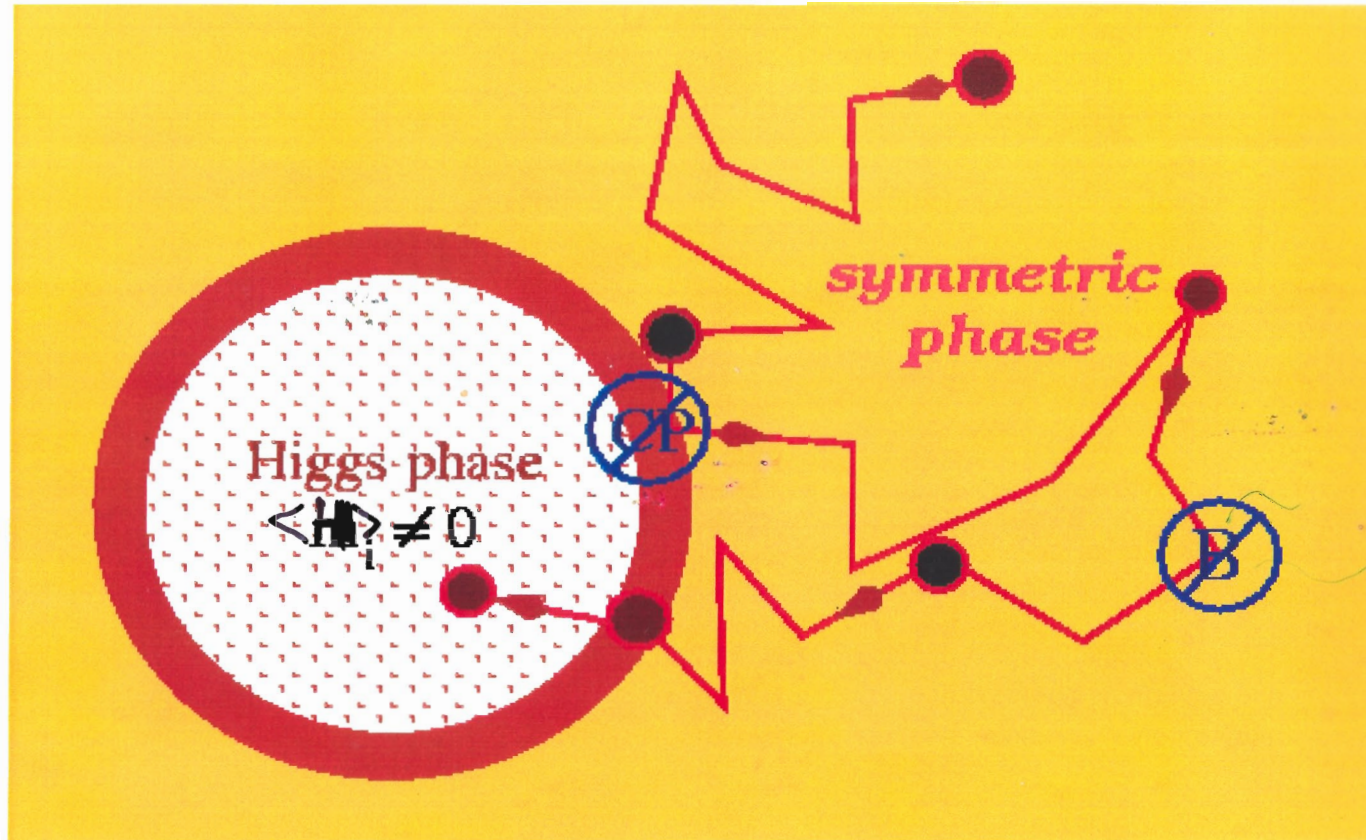
- BARYON ASYMMETRY GENERATED BY (UNSUPPRESSED!)
HOT SPHALERON TRANSITION ($V \equiv 0$)
- WASH OUT BY SPHALERON ($V(T) \neq 0$) IN HIGGS PHASE
STOPPED IF $\frac{V(T)}{T} \gtrsim 1$ (DETAILED LATER) $\Gamma_{\text{SPH.}} < H$

⇒ NEED STRONG FIRST ORDER PHASE TRANSITION

EARLY HOPE: PHASE TRANSITION IN SM?

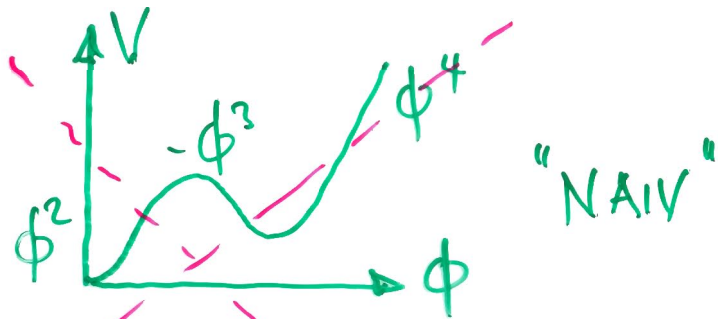
Electroweak baryogenesis at a strong 1st order transition

CHARGE TRANSPORT



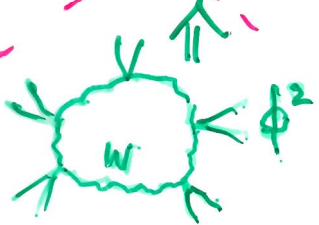
- expanding bubbles of higgs phase
- CP violation on bubble walls \Rightarrow CREATE CHIRAL ASYMM.
- B violation in symmetric phase (SPHALERON)

NEED STRONG FIRST ORDER PHASE TRANSITION



1/2 LOOP PERT. TH. ✓

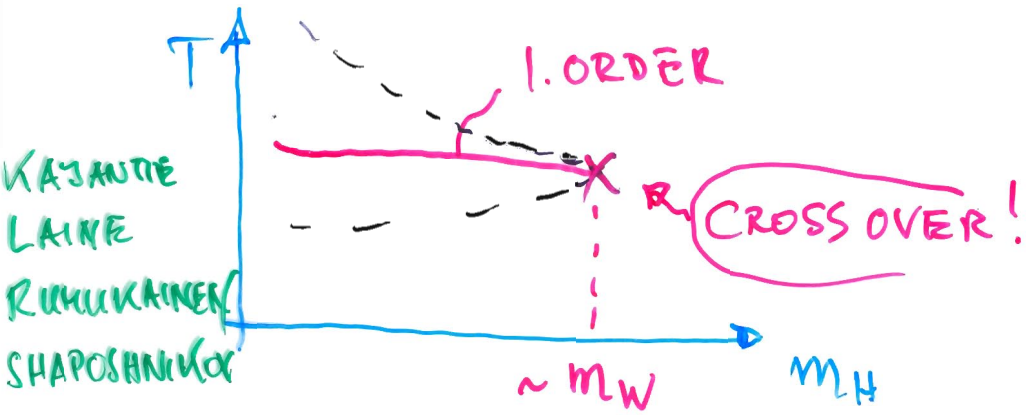
→ MORE CONCISE: "INTEGRATE OUT"
 $M \neq 0$ HIGGS BOSON MODES
 OBTAIN EFF. 3-DIM. THEORY
 (AGAIN GAUGE THEORY WITH HIGGS FIELD,
 NO FERMIONS)



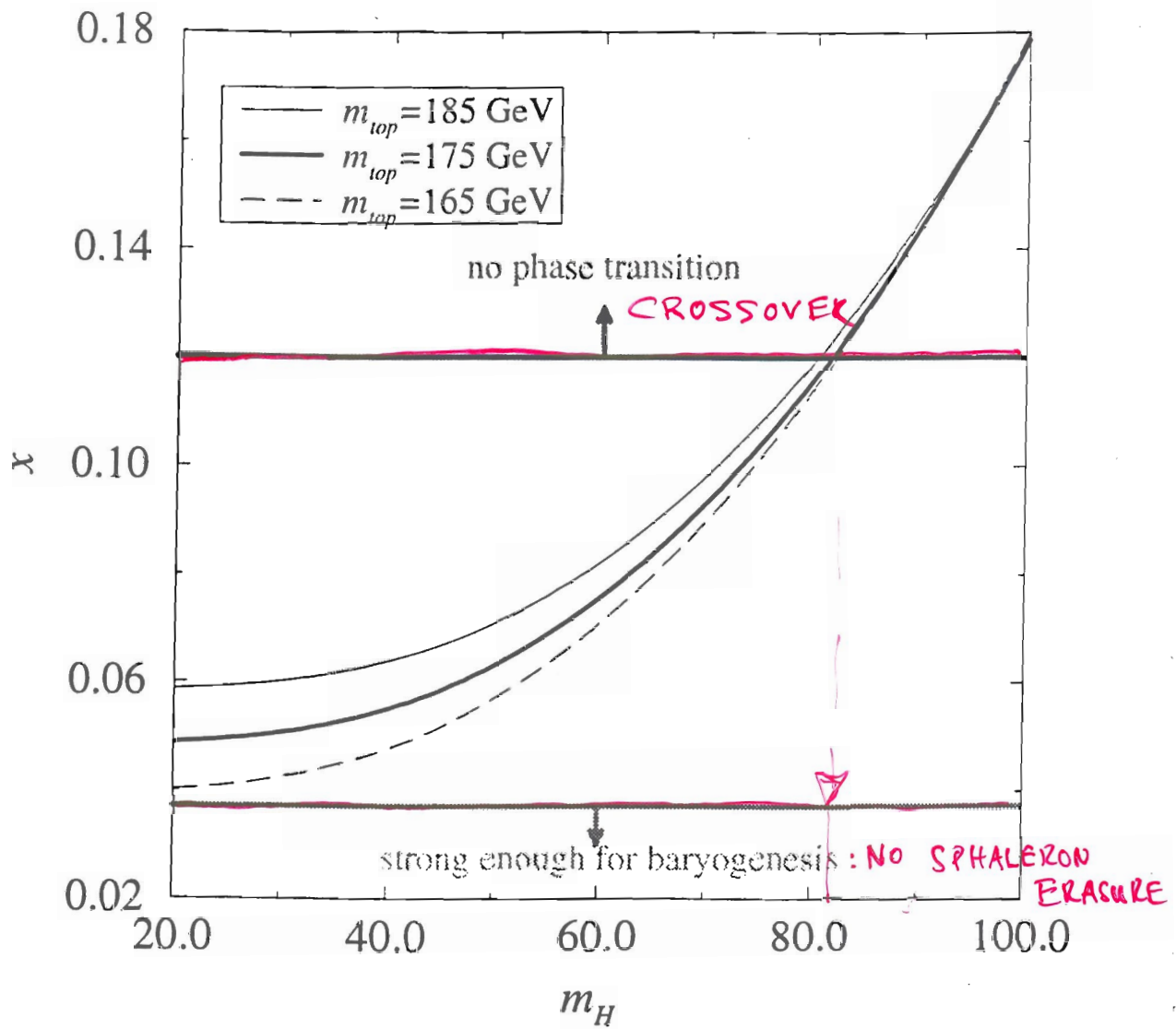
MASS SCALES

M_H, T \Rightarrow $g_W T$ \Rightarrow $g_W^2 T$ CAN BE STRONGLY IR SENSITIVE
MATSUBARA DEBYE 3D THEORY (TRANSV. GAUGE B.)

→ LATTICE GAUGE TH. CALC.



⇒ **NO** FIRST ORDER PHASE TRANS. IN SM!



KAJANTIE
LAINE
KUMKATINEN
SHPAOSTHYKOVA

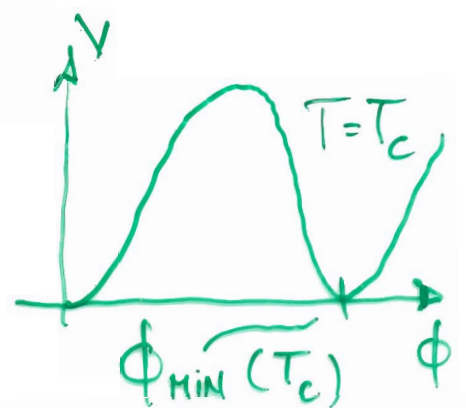
$$x = \frac{\lambda_3(T)}{g_3^2} \sim \frac{\lambda T}{g_{\text{WEAK}}^2} \iff m_{\text{HIGGS}}$$

3D
4D

$$(g_3^2 = g_{\text{WEAK}}^2 T (1 + \dots))$$

$x \lesssim 0.04$
? STRONG 1. ORDER
PT

$$\frac{\phi_{\text{MIN}}(T_c)}{T_c} \gtrsim 1$$



ALSO ~~CP~~ VERY SMALL !?? (ARGUMENT IN LIT.!) (T¹²)
↑
??

JARLSKOG DETERMINANT:

$$\Delta_{CP} = \mathcal{J} (m_u^2 - m_c^2)(m_c^2 - m_t^2)(m_t^2 - m_u^2)(m_d^2 - m_s^2)(m_s^2 - m_b^2)(m_b^2 - m_d^2) / \mathcal{J}^{12}$$

WITH $\mathcal{J} = \mathcal{J}_m (V_{fg} V_{hi} V_{fi}^* V_{hg}^*) = s_1^2 s_2 s_3 c_1 c_2 c_3 \sin \delta$ $\approx 10^{-13}$

↑
CKM-MATRIX

$\approx (3.0 \pm 0.3) 10^{-5}$

NOT SO SMALL

INTEGRATING OUT FERMIONS AND EXPANDING IN m_{DIAGONAL}^2
('PERTURBATIV')

MISSLEADING AT SMALL TEMPERATURES IN A TACHYONIC
ELECTROWEAK TRANSITION WHERE $\langle \phi^+ \phi \rangle$ IS INCREASING FROM
0 TO $v/2$? \leadsto 'COLD BARYOGENESIS'

\leadsto (ANDERS TRANBERG'S TALK!) (NONPERTURBATIV IN m^2)

(i) "FINITE TIME QUENCH" CHANGE SIGN OF ϕ^2 -TERM!

MODEL: INTRODUCE INFLATON FIELD IN ELECTROWEAK SM

→ LOW SCALE HYBRID INFLATION

$$V(\sigma, \phi) = \frac{\lambda}{4} (\phi^2 - \mu^2)^2 + \frac{1}{2} g^2 \phi^2 \sigma^2 + \frac{1}{2} m_\sigma^2 \sigma^2 + \dots$$

$\langle \phi^2 \rangle \neq 0$ FOR $\sigma^2 < \frac{\lambda}{g^2} \mu^2$

INVERTED HYBRID INFLATION

σ DECREASING

$$V(\sigma, \phi) = \frac{\lambda}{4} \phi^4 + \frac{1}{2} m^2 \phi^2 - \frac{1}{2} g^2 \phi^2 \sigma^2 + \frac{1}{q} k_q \sigma^q + \frac{1}{p} k_p \sigma^p$$

($q > p > 4$)

$\langle \phi^2 \rangle \neq 0$ FOR $\sigma^2 > \dots$
 σ INCREASING

HAS BETTER CONTROL OF RADIATIVE CORRECTIONS

(KEEP INFLAT. POT 'FLAT')

FIT TO COSMOLOGICAL DATA POSSIBLE (FINE TUNING?)

ref. in
COPELAND
LYTH
RAGANIE
TRODDEN

HYBRID INFLATION

1

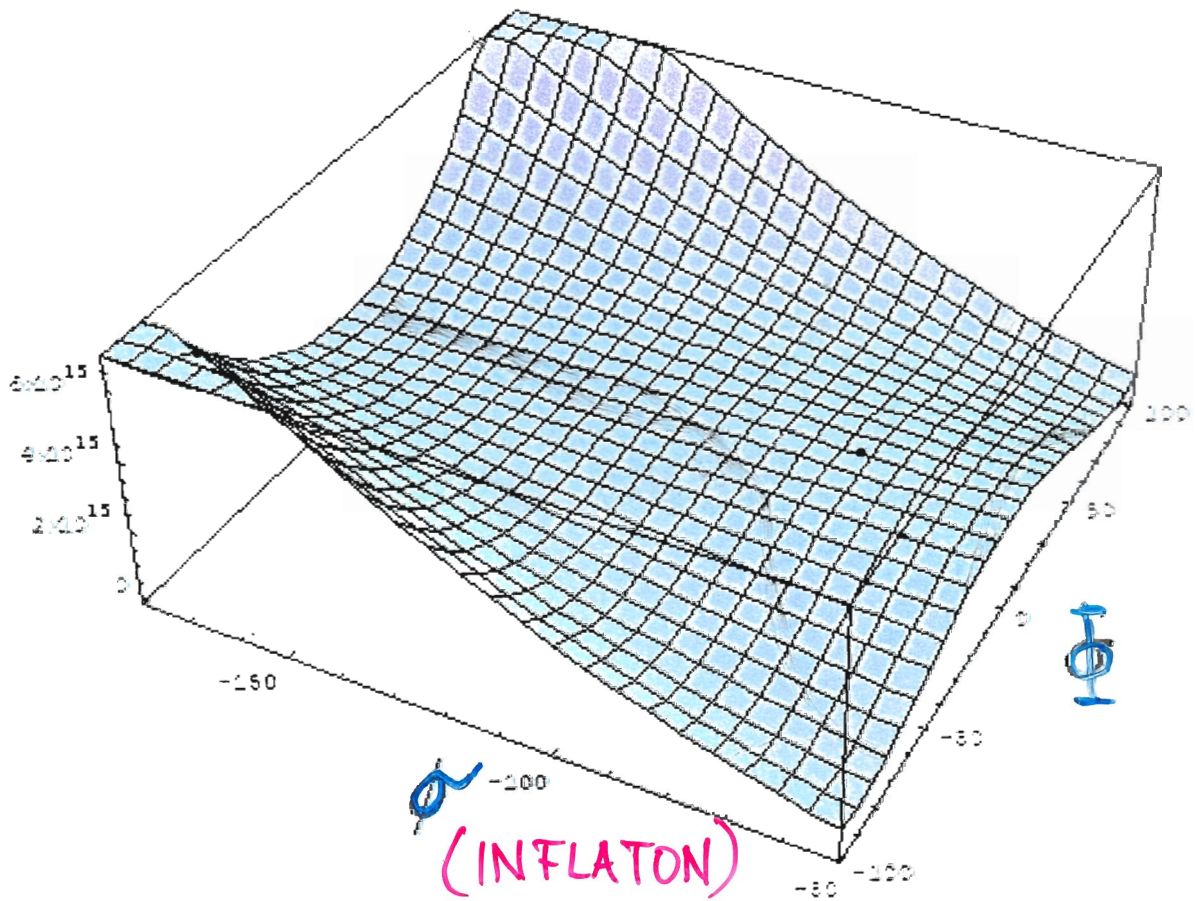


Figure 1: Hybrid Potential, using $m_{pl} = 10^9$, $\lambda = 10^4$, $g = 8 \cdot 10^3$, $m = 1.5 \cdot 10^{-5} m_{pl}$, and $M = 10^{-3} \cdot m_{pl}$.

Simpl. FINITE TIME QUENCHES

A. TRANBERG
J. SMIT
M. HINDMARSH

VAN DER MEULEN
SEXTY

$$\mu_{\text{eff}}^2(t) \phi^\dagger \phi \quad \text{WITH} \quad \mu_{\text{eff}}^2 = \begin{cases} \mu^2 \left(1 - \frac{2t}{t_Q}\right) & 0 < t < t_Q \\ -\mu^2 & t > t_Q \end{cases}$$

TENTATIVELY CP VIOLATING TERM

$$\kappa \phi^\dagger \phi F^{\mu\nu} \tilde{F}_{\mu\nu} \quad \kappa = \frac{3\delta_{CP}}{16\pi^2 m_W^2} \quad ??$$

- SOON AFTER μ_{eff}^2 IS GETTING NEGATIVE: "TACHYONIC PREHEATING"

IR MODES ARE GETTING UNSTABLE \rightarrow GROWING \rightarrow CLASSICAL EQS.

NONLINEAR TERMS IN EOM UNIMPORTANT

$$\ddot{\phi}_k + k^2 \phi_k + \underbrace{\mu_{\text{eff}}^2(t)}_{\text{neg.}} \phi_k = 0$$

$$\phi_k \sim e^{(-\mu^2 - k^2)t}$$

$k^2 < |\mu^2|$
"IR"

- STUDY CLASSICAL EOM'S ON LATTICE WITH INITIAL CONDITIONS:

\rightarrow TIME DEPENDENCE OF $\langle \phi^\dagger \phi \rangle$

$$\rightarrow N_{CS}(t) - N_{CS}(0) = \frac{1}{16\pi^2} \int_0^t dt' \int d^3x F^{\mu\nu} \tilde{F}_{\mu\nu} \Leftrightarrow \text{BARYON NUMBER}$$

• ALSO 'HIGGS WINDING'

$$N_W = \frac{1}{24\pi^2} \int d^3x \epsilon_{ijk} \text{Tr} [u^\dagger (\partial_i u) u^\dagger (\partial_j u) u^\dagger (\partial_k u)]$$

UNITARY GAUGE

START WITH $N_{CS} = N_W = 0$

FINALLY $N_{CS} = N_W \neq 0$

$$U = \frac{\Phi}{(\frac{1}{2} \text{Tr} \Phi^\dagger \Phi)^{1/2}}$$

POSITIONS WITH $\Phi^\dagger \Phi = 0$
IMPORTANT FOR CHANGE!

$$\left(\Gamma_{\text{SPHALERON}} = \frac{d \Delta_{CS}}{dt} \text{ in THERMAL EQUIL.} \right)$$

• $\frac{n_B}{n_\gamma} = (0.20 \pm 0.04) \times 10^{-3} \delta_{CP}$ FOR $m_H = 2m_W, m_H(t_Q) = 18$
OPTIMAL CHOICE

BUT: WHAT IS THE REAL ~~CP~~ TERM IN THE (EXT.?) SM?

EARLIER WORK BY GARCIA-BELLIDO, GRIGORIEV, SHAROSHNIKOV, KRUSENKO ...
COPELAND, LYTH, RAJANTIE, TRODDEN

(ii) DERIVE CP VIOLATING TERM(S) IN BOSONIZED SM :

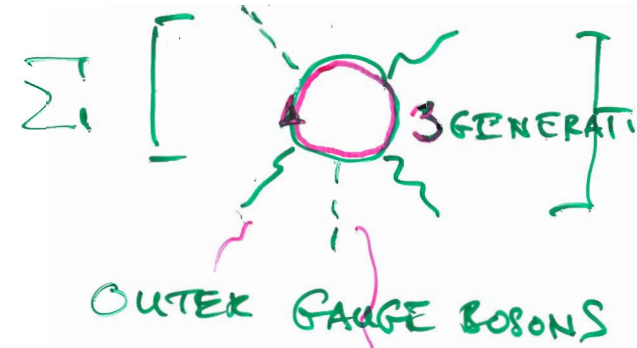
EFFECTIVE ACTION INTEGRATING OUT FERMIONS

J. SHIT, METH. SALCEDO

A. HERNANDES
T. KONSTANDIN
H.G. SCH.

NEED SECOND ORDER (NEXT TO LEADING)
DERIVATIVE EXPANSION!

WORLDLINE METHOD OPTIMAL



KONDRASHIN, NELLEN, PER, SCHUBERT + (PSEUDO) SCALARS
D'HOOKER, GAGNIER

$$S_{CP} = \frac{1}{8(4\pi)^2} \frac{3}{16} \int K^{CP} \epsilon^{\mu\nu\lambda\sigma} \frac{1}{\tilde{m}_c^2} \times$$

$$\int d^4x \left[z_\mu W_\nu^\dagger W_\alpha^- (W_\sigma^+ W_\alpha^- + W_\alpha^+ W_\sigma^-) + c.c. \right]$$

$$K^{CP} \approx 8.87 \quad \int = s_1^2 s_2 s_3 c_1 c_2 c_3 \sin\Gamma \approx (3.0 \pm 0.3) 10^{-5}$$

$$W_{\mu\nu}^+ = \frac{\phi^\dagger W_{\mu\nu} \tilde{\phi}}{\phi^\dagger \phi} \dots W_\mu^+ = \frac{\phi^\dagger \mathcal{D}_\mu \tilde{\phi}}{\phi^\dagger \phi} \quad z_\mu = W_\mu^3 - B_\mu = \frac{\phi^\dagger \mathcal{D}_\mu \phi - \tilde{\phi} \mathcal{D}_\mu \tilde{\phi}}{2\phi^\dagger \phi}$$

($\tilde{\phi} = \epsilon \phi^*$)

$$\lim \frac{\kappa^{CP}}{\tilde{m}_c^2} = \frac{32}{9\tilde{m}_c^2 (\tilde{m}_t^2 - \tilde{m}_s^2)^3 (\tilde{m}_c^2 - \tilde{m}_t^2)^3 (\tilde{m}_s^2 - \tilde{m}_t^2)^2} \times$$

$$\left(\tilde{m}_s^6 \tilde{m}_t^6 (\tilde{m}_s^2 - \tilde{m}_t^2)^2 + 3\tilde{m}_c^{14} (\tilde{m}_s^2 + \tilde{m}_t^2) \right.$$

$$- 5\tilde{m}_c^2 \tilde{m}_s^4 \tilde{m}_t^4 (\tilde{m}_s^2 - \tilde{m}_t^2)^2 (\tilde{m}_s^2 + \tilde{m}_t^2) - 12\tilde{m}_c^{12} (\tilde{m}_s^4 + \tilde{m}_t^4)$$

$$+ \tilde{m}_c^4 \tilde{m}_s^2 \tilde{m}_t^2 (\tilde{m}_s^2 - \tilde{m}_t^2)^2 (13\tilde{m}_s^4 + 28\tilde{m}_s^2 \tilde{m}_t^2 + 13\tilde{m}_t^4) + 18\tilde{m}_c^{10} (\tilde{m}_s^6 + \tilde{m}_t^6)$$

$$+ \tilde{m}_c^8 (-12\tilde{m}_s^8 + 37\tilde{m}_s^6 \tilde{m}_t^2 - 74\tilde{m}_s^4 \tilde{m}_t^4 + 37\tilde{m}_s^2 \tilde{m}_t^6 - 12\tilde{m}_t^8)$$

$$\left. - \tilde{m}_c^6 (3\tilde{m}_s^{10} - 41\tilde{m}_s^8 \tilde{m}_t^2 + 41\tilde{m}_s^6 \tilde{m}_t^4 + 41\tilde{m}_s^4 \tilde{m}_t^6 - 41\tilde{m}_s^2 \tilde{m}_t^8 + 3\tilde{m}_t^{10}) \right)$$

$$\frac{64 \tilde{m}_c^4 \tilde{m}_s^2 \tilde{m}_t^2 (\tilde{m}_c^2 - \tilde{m}_t^2) (\tilde{m}_c^2 - 3\tilde{m}_s^2 + 2\tilde{m}_t^2) \log \left[\frac{\tilde{m}_s^2}{\tilde{m}_c^2} \right]}{3 (\tilde{m}_c^2 - \tilde{m}_s^2)^4 (\tilde{m}_s^2 - \tilde{m}_t^2)^3}$$

$$\frac{64 \tilde{m}_c^4 \tilde{m}_s^2 (\tilde{m}_c^2 - \tilde{m}_s^2) \tilde{m}_t^2 (\tilde{m}_c^2 + 2\tilde{m}_s^2 - 3\tilde{m}_t^2) \log \left[\frac{\tilde{m}_t^2}{\tilde{m}_c^2} \right]}{3 (\tilde{m}_c^2 - \tilde{m}_t^2)^4 (\tilde{m}_s^2 - \tilde{m}_t^2)^3}$$

ALREADY SIMPLIFIED!

FOR $\tilde{m}_u \rightarrow \tilde{m}_d \rightarrow 0$
 $\tilde{m}_b \rightarrow \tilde{m}_c$

?? SALCEDO: • ~~CP~~ - PART FROM W^+ {P CONSERVING?}

• PROPOSAL FOR A DIRECT CALCULATION OF W^-

- WE JUST CHECK THIS

VERY RECENT

IN PROGRESS

WORK WITH A. TRANBERG: "COLD BARYOGENESIS" WITH OUR ~~CP~~-TERM

REMARKS:

- IF AS A RESULT OF TIME VARIATION IN THE YUKAWA COUPLINGS QUARK MASSES WERE LARGE AT THE TIME OF THE ELWK. TRANS. THE KM MECHANISM IS STRONGLY ENHANCED

BERKOOZ
NIR
VOLANSKY
(WITH FN MECHAN.)

- THE SM ENLARGED BY R.H. SINGLET NEUTRINOS OFFERS VERY INTERESTING NEW POSSIBILITIES FOR CP VIOLATION, DARK MATTER ...

"νMSM"

SHAPOSHNIKOV
ASAKA

BARYOGENESIS: 3 SINGLET NEUTRINOS WITH $M_2 \sim M_3 \sim \text{GeV}$

ASAKA, SHAPOSHNIKOV

$$|M_2 - M_3| \sim M_1 \sim \text{keV!}$$

GORBUNOV
KHEBNIKOV
KUSENKO

AKHMEDOV, RUBAKOV, SMIRNOV

- SCILLATE AND MIX WITH L.H. NEUTRINOS

LEPTON NUMBER IN 2,3 GENERATED VIA THERMALIZATION

HOT \rightarrow SPHALERON BARYON NUMBER

- S_1 NOT THERMALIZED ABOVE $T_{PT} \rightarrow$ DARK MATTER CAND., (NO BARYON NUMBER GEN.!) \rightarrow DARK MATTER CAND.,

• STILL CLOSE TO THE SM : ELECTROWEAK PHASE TRANS.

NARDINI
QUIROS
WULZER

IN THE RANDALL-SUNDBRAH MODEL

IR - BRANE SUBSTITUTED BY BLACK HOLE HORIZON AT LARGE TEMPERATURES, CONDENSATES TO USUAL RS-IR-BRANE.

⇒ CREMINELLI
NICOLIS
RATTAZZI

RADION IS STABILIZED BY GOLDBERGER-WISE MECHANISM. THE CRITICAL TEMPERATURE T_c IN SUB-TeV-RANGE, BUT ACTUAL NUCLEATION T_N

IS ORDERS BELOW DUE TO THE FLATNESS OF THE RADION STABILIZING POTENTIAL (\rightarrow LOW SCALE INFLATION), i.e. STRONG SUPERCOOLING,

HIGGS ON THE IR BRANE ($\phi \neq 0$ IN EFF. TH.) STRONG FIRST ORDER PT

ADD CP-VIOLATION \rightarrow BARYOGENESIS

N/QW
'07

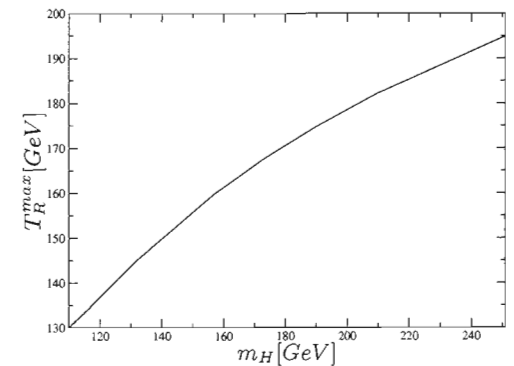
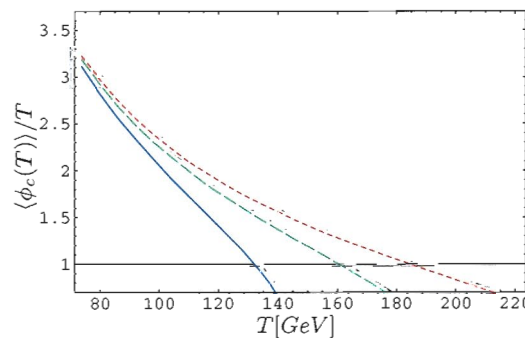


Figure 6: Left panel: Plot of $\langle \phi_c(T) \rangle / T$ as a function of the temperature in GeV for a Higgs mass of 115 GeV (solid line), 165 GeV (dashed) and 225 GeV (short dashed). Right panel: Plot of the temperature at which the SM potential is minimized at $\langle \phi_c(T) \rangle / T = 1$ as a function of the Higgs mass in GeV.

HIGHER DIMENSION OPERATORS IN THE EFFECTIVE

ZHANG THEORY INCLUDING THE SM (RESP. MSSM...)

BLUM, NIR
("BEYOND MSSM"
+ ref.!

LOW SCALE CUT OFF, MODIFY SM - HIGGS SELF INTERACT.

D. BÖDEKER
L. FROMME
S. HUBER!
M. SENIUCH

$$V(\phi) = -\frac{\mu^2}{2} \phi^2 + \frac{\lambda}{4} \phi^4 + \frac{1}{8M^2} \phi^6$$

$$W_{DST} = \frac{\lambda_1}{M} (H_u H_d)^2 + \frac{\lambda_2}{M} \Theta^2 \mu_{SUSY} (H_u H_d)^2$$

DINE, PEIERBERG, THOMAS

GROJEAN
SERVANT
WELLS

HELPS TO OBTAIN STRONG FIRST ORDER PT

BEYOND THE SM MODELS

- TWO-HIGGS DOUBLET MODEL(S)

FROMME
HUBER
SENIUCH

- SUPER SYMMETRIC MODELS

- MOST POPULAR **MSSM**

- INCLUDE GAUGE SINGLET CHIRAL SUPERFIELD S

ELLWANDER
ET. ...

S. HUBER
H.G. SCH.

FUNAKURO ET AL.

NMSSM

$$W = W_{\text{MSSM}} + \lambda S H_1 \epsilon H_2 - m^2 S + \frac{k}{3!} S^3 + \mu (H_1 \epsilon H_2)$$

n MSSM

$$W = W_{\text{MSSM}} + \lambda S H_1 \epsilon H_2 - m^2 S$$

+ SUSY (SOFT)
BREAKING TERMS

PANAGIOTKOPOULOS
ET AL. PILAFIS

MENON, MORRISSEY, WAGNER

HUBER, KONSTANDIN, PROKOPEE, SCH.

MSSM

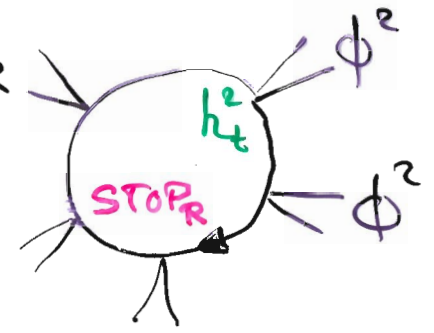
• INCREASE " ϕ^3 "-TERM WITH A LIGHT S-TOP_R IN THE LOOP

IN 3D EFFECTIVE THEORY: COLORED U SCALAR

$$m_{\text{Loop}}^2 = \underbrace{\tilde{m}_u^2}_{\text{SUSY BREAKING}} + \dots T^2 + \log \dots$$

SHALL!

$$(m_{\text{STOP}_R}^2 = \tilde{m}_u^2 + \underbrace{m_{\text{TOP}}^2}_{h_t^2 \phi^2} \text{ FOR } T=0)$$



CARENA
QUIROS
WAGNER

\tilde{m}_u^2 NEGATIVE

• GET $L_3^{\text{EFF.}}$ IN U AND H \Rightarrow LATTICE CALC. 2 LOOP PERT. TH. } ROUGHLY AGREE!

\Rightarrow FOR A RANGE OF \tilde{m}_u^2 THERE IS AN (INTERMEDIATE?!) COLORED U PHASE

• CAN GET STRONG FIRST ORDER P.T. TO HIGGS PHASE ($\frac{V(T)}{T} \gtrsim 1$)

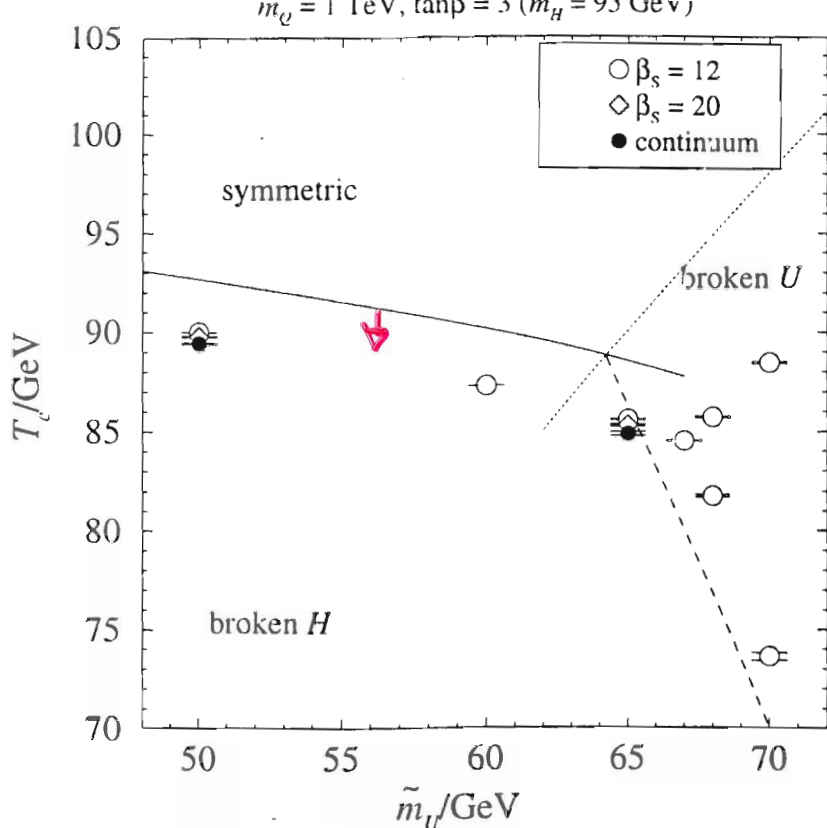


FIG. 1. The phase diagram and the critical temperatures. The continuous lines are from the 2-loop perturbative effective potential in the Landau gauge. Open symbols correspond to infinite volume extrapolations, and filled symbols to continuum limit extrapolations.

LATTICE

LAINÉ

RUMMUKAINEN

hep-ph/9804255

vs.

2-LOOP
PERT. TH.

BÖDEKER
JOHN
LAINÉ
SCH.

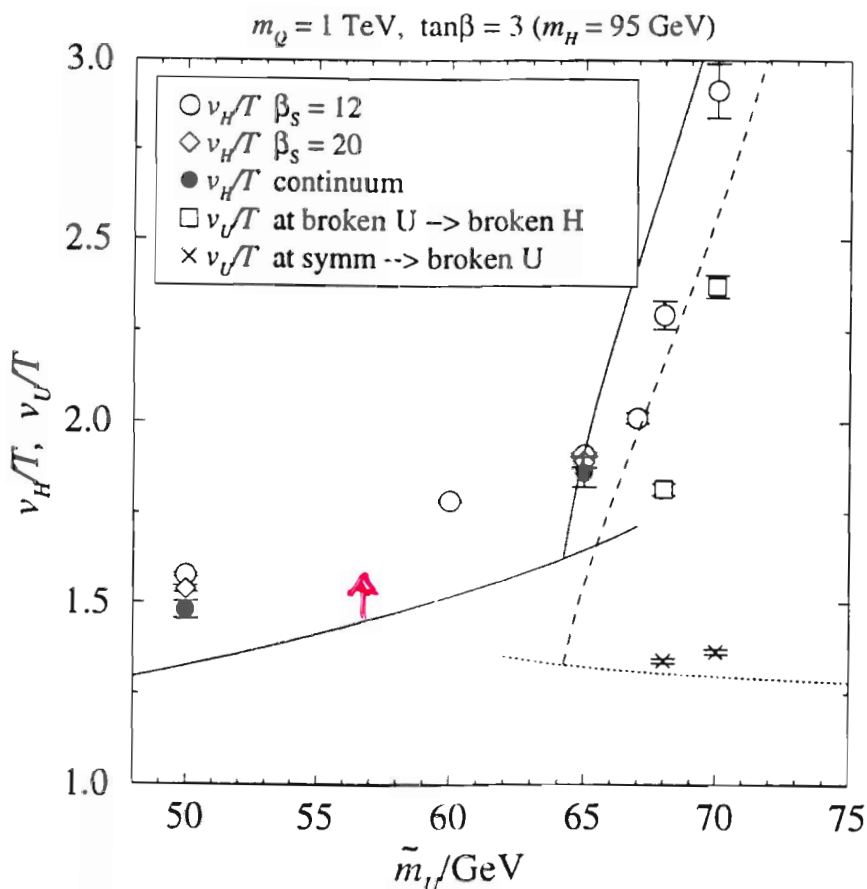


FIG. 3. The scalar field expectation values in the broken phases at T_c .

NONPERT.
EFFECTS:

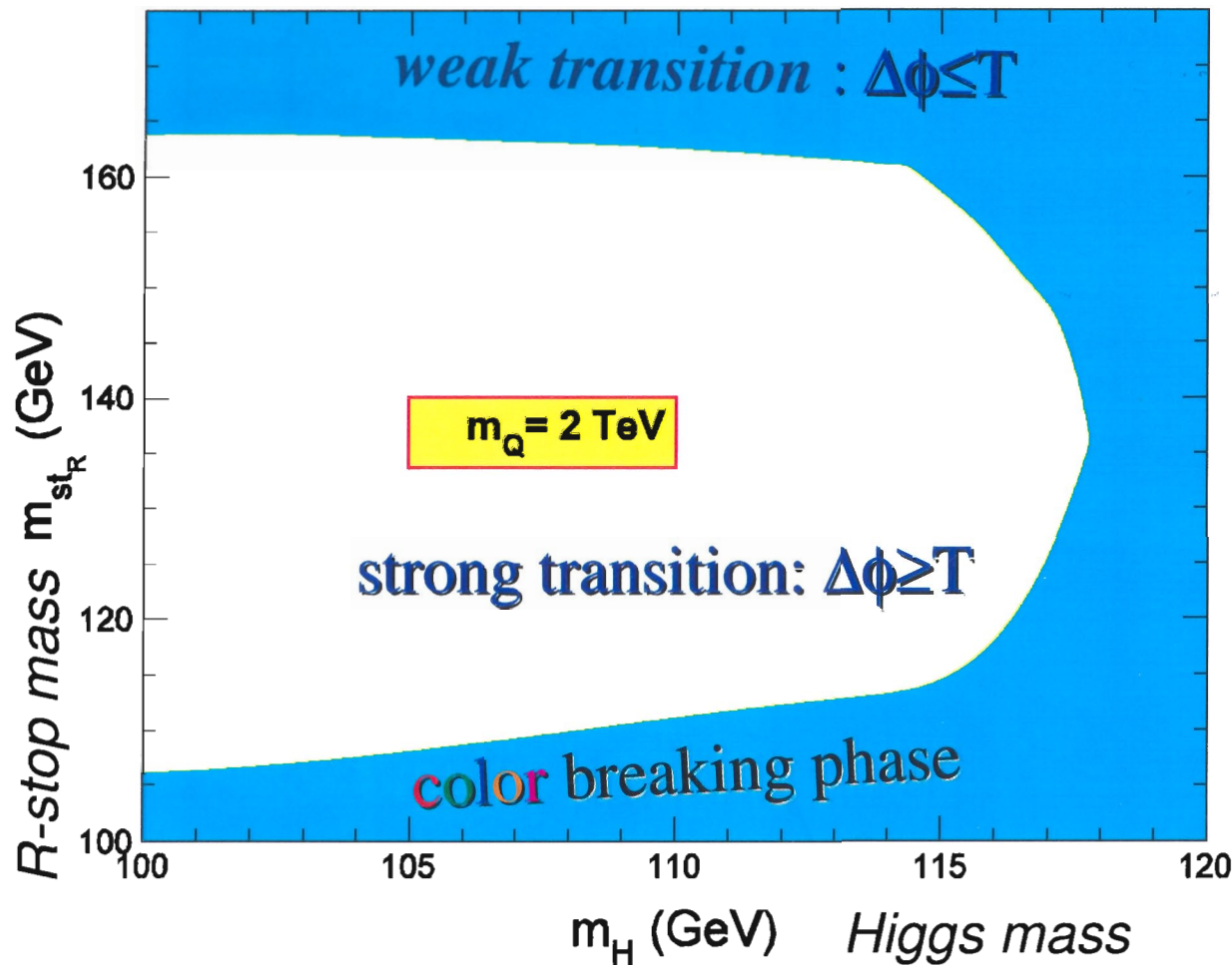
\Rightarrow STRONGER
P.T.

! ~ 2
 m_a NEGATIV

Strong first order transition in MSSM

allowed “triangle” for MSSM:

Carena, Quiros, Seco, Wagner, 2000



CARENA, NARDINI, QUIRÓS, WAGNER

'08 ANALYSE IN 1-LOOP IMPROVED (DAISY, "2-LOOP" LOG'S) THEORY
THE PHASE DIAGRAM. FIND EXTENSION OF (NOT IN EFF. 3D THEORY)
"ALLOWED" REGION IN $(m_{\text{stop}_2} - m_H)$ ADMITTING
QUASI STABLE H-VACUA FOR LARGE SUSY MASS SCALE \tilde{m}
(THUS ALSO SUPPRESSING EDM)

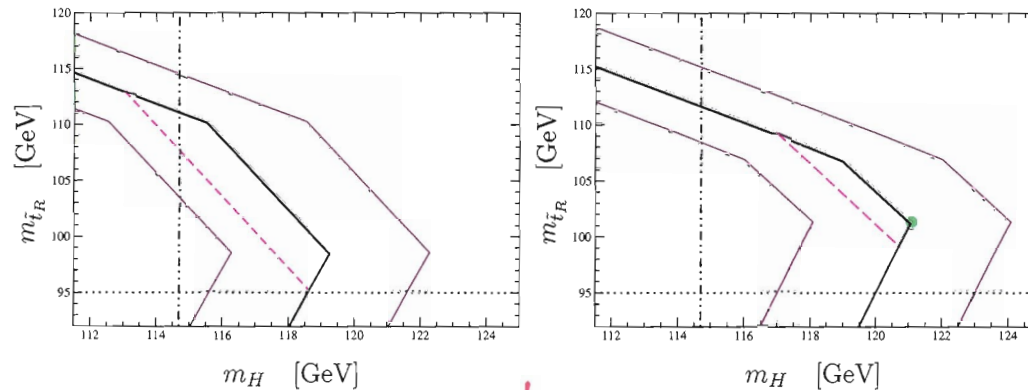


Figure 3: Window where $\phi(T_H^c)/T_H^c \geq 0.9$ and $T_H^c \geq T_U^c + 1.6$ GeV in the $m_H - m_{\tilde{t}}$ plane for $\tilde{m} = 500$ TeV (left panel) and $\tilde{m} = 8000$ TeV (right panel). The allowed region is below the solid lines and dashed lines for $\tan \beta \leq 15$ and $\tan \beta \leq 5$, respectively. The thick solid line is obtained by ignoring the Higgs mass uncertainty, while the solid thin lines is obtained by including an uncertainty of 3 GeV in the Higgs mass computation. The Higgs (stop) mass lower bound is marked by a dotted-dashed (dotted) straight line. In green (right panel) the point that will be numerically analyzed in the tunneling analysis.

MSSM - BARYOGENESIS ANALYZED IN THE "LIGHT HIGGS BOSON SCENARIO" (LHS)

• $\frac{g_{hZZ}}{g_{h^{SM}ZZ}} = \sin(\beta - \alpha)$ $\frac{g_{HZZ}}{g_{h^{SM}ZZ}} = \cos(\beta - \alpha) \dots$

SMALL

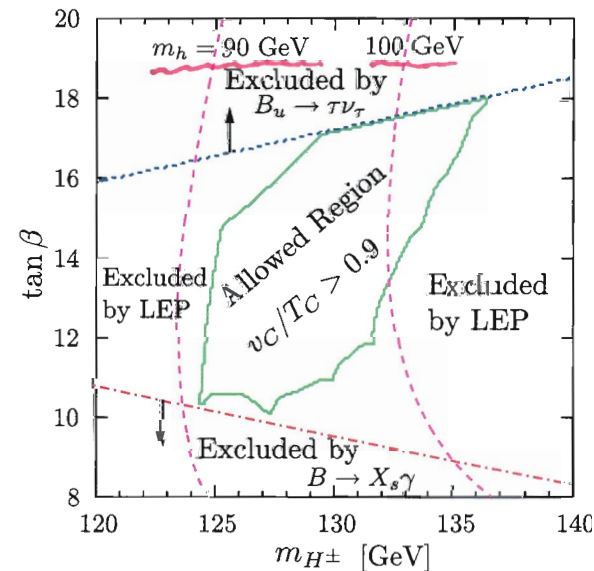
($\approx 0.5 \rightarrow m_h \sim 100 \text{ GeV}$)

• CHOICE $m_{\tilde{F}_R}^2(T) = m_{\tilde{F}_R}^2 + cT^2 \approx 0$.
 NEGLECTED?! NO "DAISIES"

• REANALYSIS OF $v(T_{\text{NUCLEATION}})/T_N$ NEEDED FOR SPHALERON DECOUPLING
 $> 1.38!$ (\rightarrow "BARYOGENESIS INFEASIBLE"?!)

• RESTRICTIONS BY B-PHYSICS DATA

FIG. 3: The allowed region in the LHS. We take $A_t = A_b = -300 \text{ GeV}$, $\mu = 100 \text{ GeV}$, $m_{\tilde{q}} = 1200 \text{ GeV}$, $m_{\tilde{t}_R} = 10^{-4} \text{ GeV}$ and other input parameters are presented in the text.



CONDITIONS FOR A STRONG FIRST ORDER ELWK. PT. MUCH EASIER TO FULFILL IN THE NMSSM, nMSSM → FIG.

GIVEN SUCH A STRONG PHASE TRANSITION THE PROCEDURE TO OBTAIN A BARYON ASYMMETRY HAS QUITE A FEW STEPS, BUT ALL OF THEM VERY CONCRETE AND FEASIBLE

- CRITICAL BUBBLE (MULTIDIM. IN HIGGS FIELDS) ("DET"!)

→ FIG.

- TRANSITION PROBABILITY (LANGER FORM.)

$$\sim e^{-S_{\text{EFF}}}$$

- SUPERCOOLING ("1 BUBBLE/UNIVERSE")
NUCLEATION TEMPERATURE

- SPHALERON RATE (MULTIDIM. IN HIGGS F.)

(FLUCTUATIONS!)

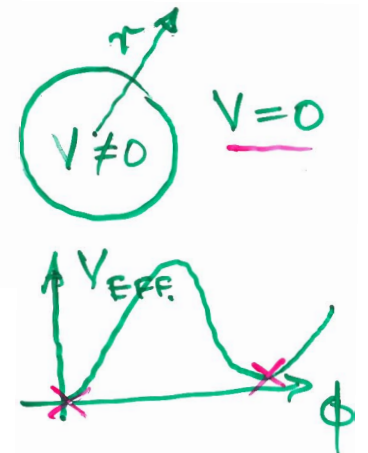
- STATIONARY EXPANSION OF BUBBLE

→ $V_W = \rho_1$ WALL PROFILE

- DIFFUSION IN PRESENCE OF MOVING WALL

→ WITH CP-VIOLATING WALL OR EXPLICIT ~~CP~~ INTERACTION

(QUANTUM) BOLTZMANN EQS. , GENERATE CHIRAL ASYMMETRY $n_{q_L} - n_{\bar{q}_L}$



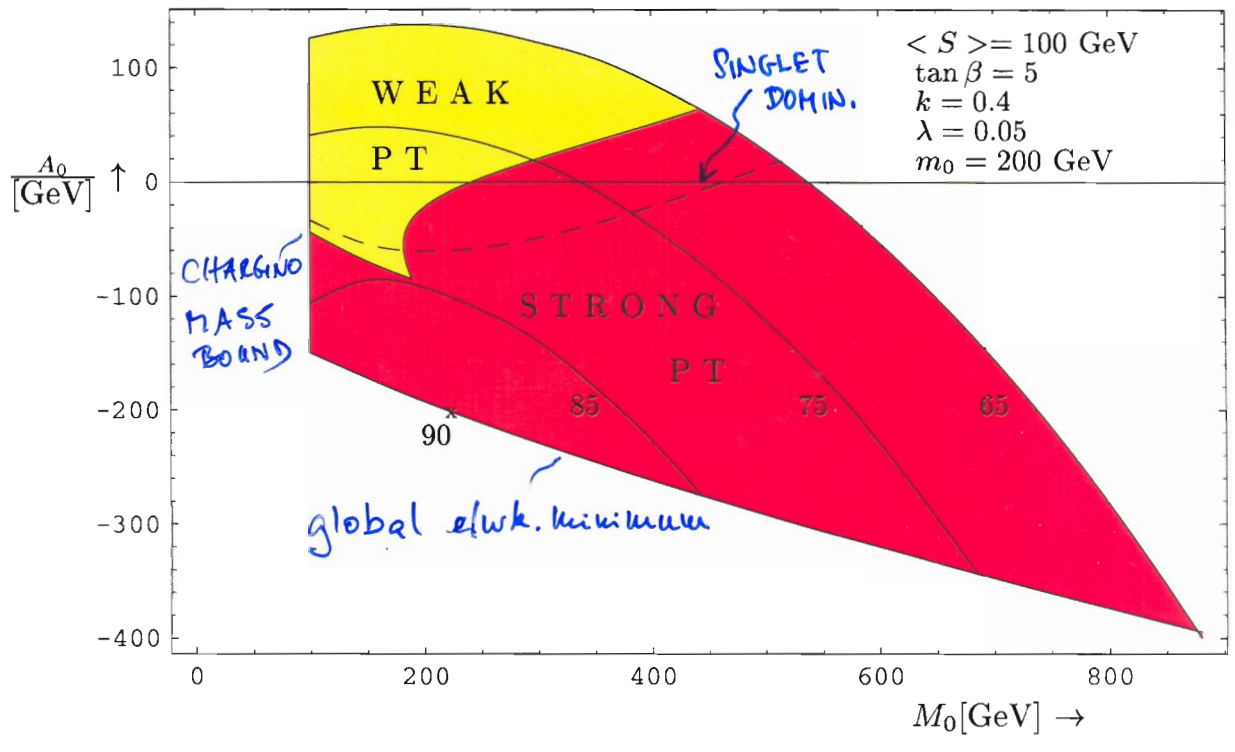
HIGGS

SYMM.

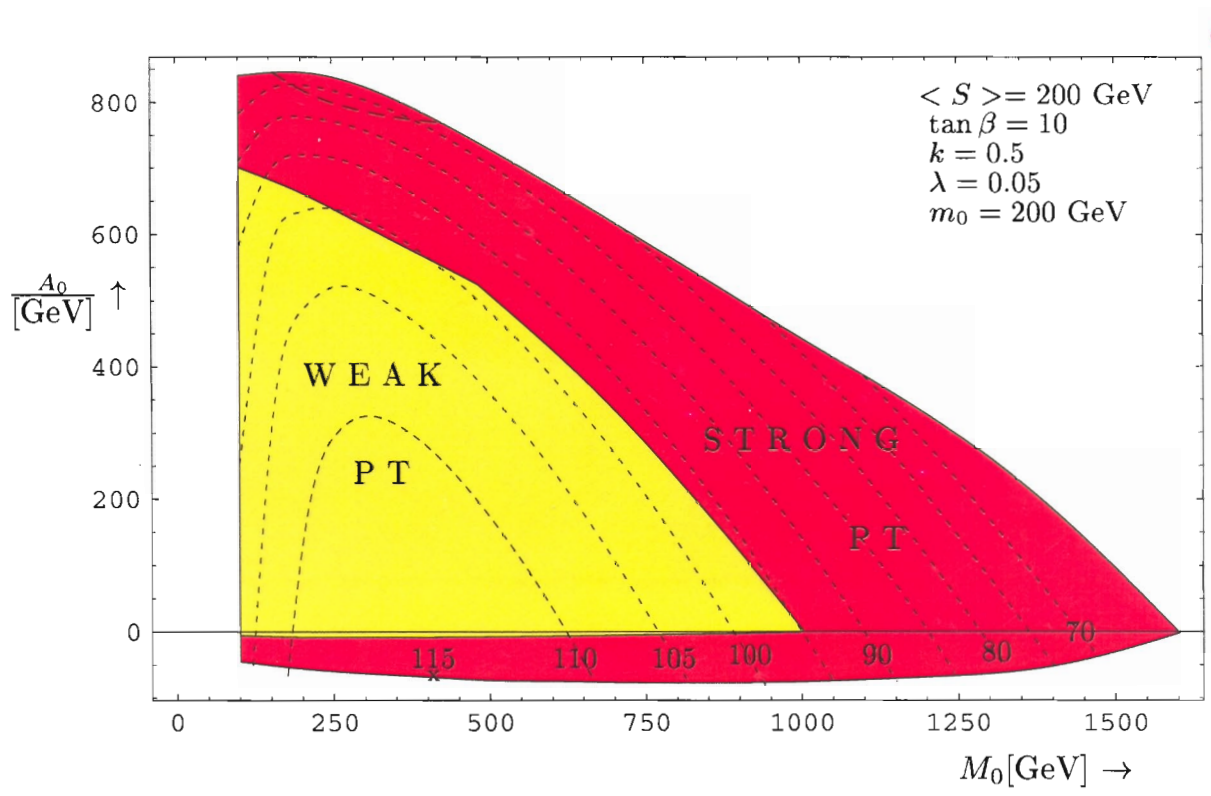
→ V_W DEFLAGRATION

NMSSM RESULTS for strength of the PT :

S. Huber
M.G. Schmidt



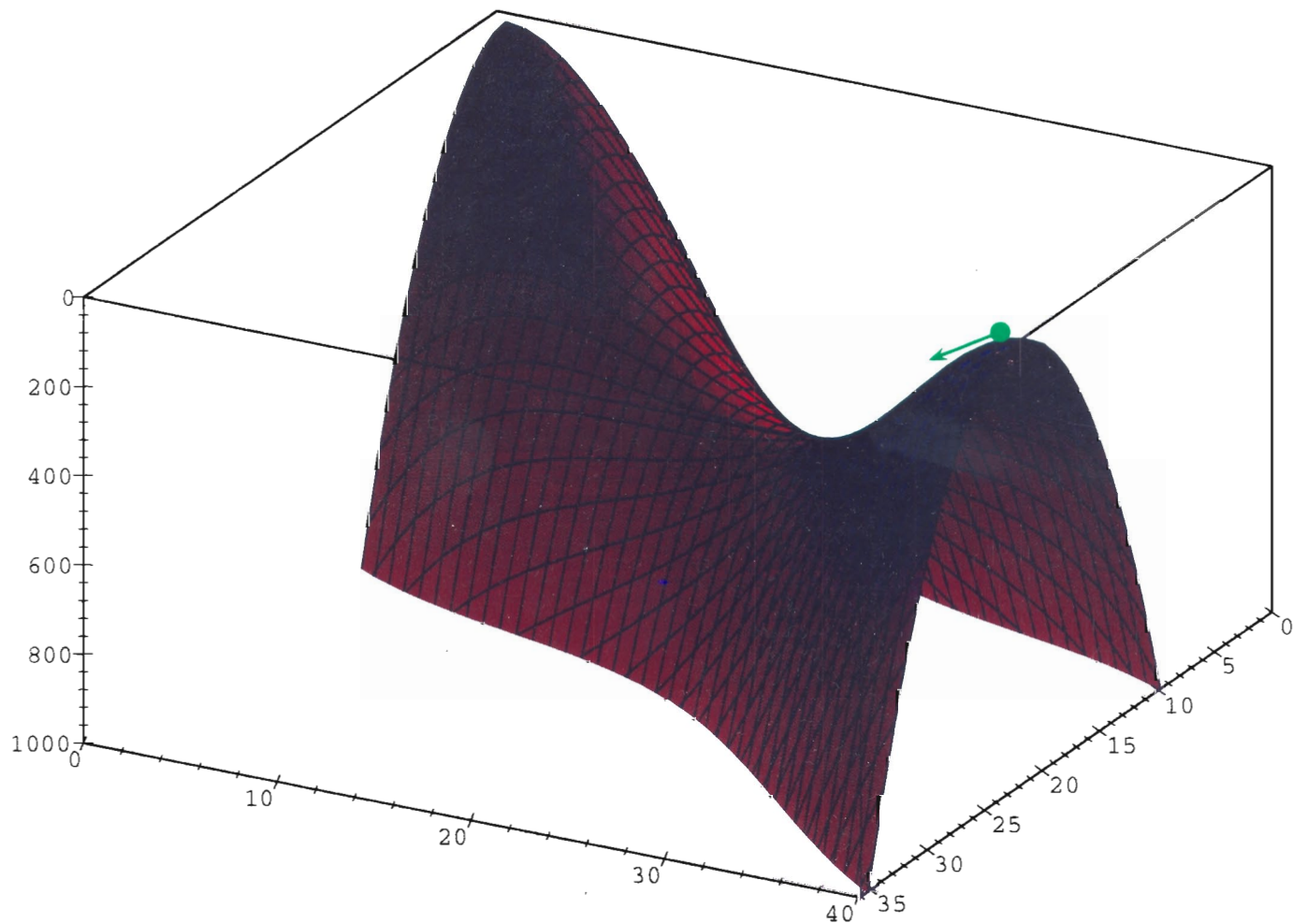
$\langle S \rangle = 100 \text{ GeV}$



$\langle S \rangle = 200 \text{ GeV}$

$M_H^{\text{max}} = 115 \text{ GeV}$ and $\frac{v_c}{v_s} > 1$

Thermisches Tunneln
mit mehr als einem Skalarfeld



MORE RECENT WORK BY S. HUBER + T. KONSTANDIN

- PRODUCTION OF LH QUARKS VIA DIFFUSION
~~B~~ BY "HOT" SPHALERON OF ELWK. THEORY IN
 FRONT OF BUBBLE WALL
- FAST FREEZING OUT OF SPHALERON TRANSITION IN
 HIGGS PHASE CONSERVES BARYON ASYMMETRY

RECENT WORK

CHUNG
 GARBRECHT
 RAMSEY-MUSOLF
 CIRIGLIANO
 PROFUMO

$$\frac{V(T_N)}{T_N} > \dots$$

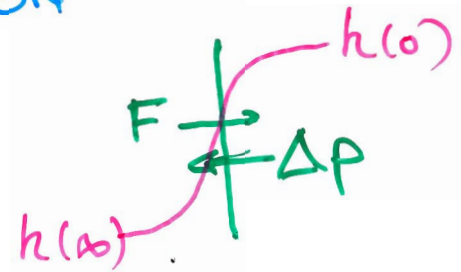
⇒ WALL VELOCITY V_w

! SMALL FOR BARYOGENESIS: $n_B \sim \frac{\Gamma_{\text{HOT}}}{V_w} \int_{-\infty}^0 dz n_L(z) \exp\left(-z \frac{\Gamma_{\text{HOT}}}{V_w}\right)$

STATIONARY CASE:

BUBBLE PRESSURE-DIFFERENCE = - BUBBLE FRICTION

$$\Delta V_T = \int_{-r}^{+r} h' dx \left(h'' - \frac{\partial V}{\partial h} \right)$$



WALL FRAME $x = z + V_w t$
ORDER V_w CALC.

$$\partial_t \delta = \underline{V_w} \delta'$$

G. MOORE
T. PROKOPEC

$$\underbrace{\square h + \frac{\partial V_T}{\partial h}}_{\text{CONTRIB. TO PRESSURE}} + \underbrace{\sum_i \left(\frac{\partial m_i^2}{\partial h} \right) \int \frac{d^3 p}{2E_i(z)} \delta f_i(p, z)}_{\substack{\text{TO FRICTION} \\ \text{"HEAVY PART."}}} = 0$$

LINEARIZED IN δ

$$f_i = \left(\exp\left(\frac{E_i + \delta_i}{T}\right) + 1 \right)^{-1}$$

FLUID APPROXIMATION

$$\delta_i = - \left[\delta \mu_i + \frac{E}{T} (\delta T_i + \delta T_{\text{bg}}) + \frac{p x}{T} (T \delta v_i + T \delta v_{\text{bg}}) \right]$$

BOLTZMANN EQS.

$$d_t f_i = \partial_t f_i + \dot{z} \frac{\partial}{\partial z} f_i + \dot{p}_z \frac{\partial}{\partial p_z} f_i = C(f_i)$$

INTEGRATE $C(f_i)$ WITH $\left(\frac{1}{T_1}, \frac{E_i}{T_2}, \frac{p_z}{T_3} \right)$
 \vec{w}

$$\int d^3p \vec{w} C(f) = \left(\Gamma \right) \begin{pmatrix} \delta T \\ \delta U + \delta U_{bg} \end{pmatrix}$$

LINEARIZED!
 DRIVING TERM $-f_0' \frac{\partial_t (u^2)}{2E}$

$$f_i = f_{0i} + \delta f_i$$

LIN. IN δ_i

FOR "HEAVY" + "LIGHT" PART.
 (BACKGROUND)

CALCULATE
 IN VARIOUS MODELS!

SOLVE FLUID EQ. TO ELIMINATE $\delta T_{bg}, \delta U_{bg}$

⇒ FINAL FORM $\boxed{A \delta' + \Gamma \delta = F}$

MATRIX IN C_i, V_w $\rightarrow \frac{V_w}{2T} [c_1 (m_i^2)', c_2 (m_i^2)', 0]$

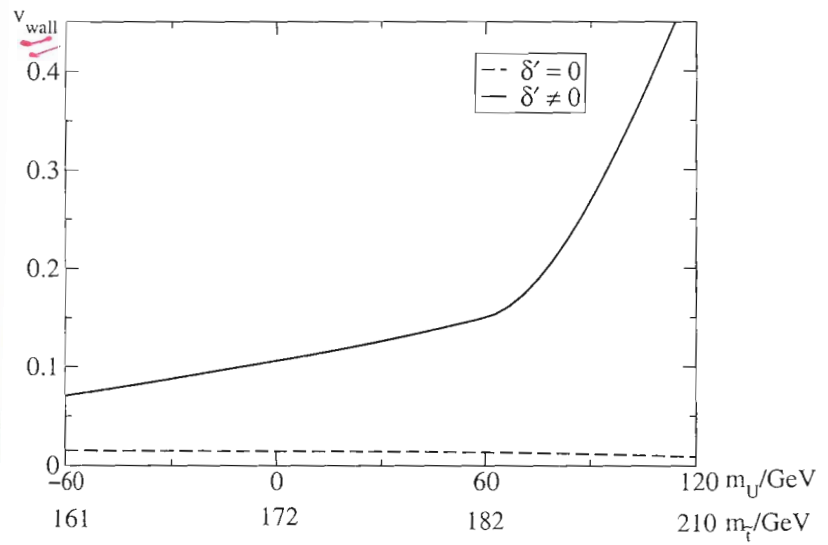
δ : INSERT IN FRICTION TERM(S); SOLVE FOR V_w
 IN (*)

AND WALL PROFILE $h(x) = \frac{h_{cut}}{2} (1 + \tanh \frac{x}{L})$

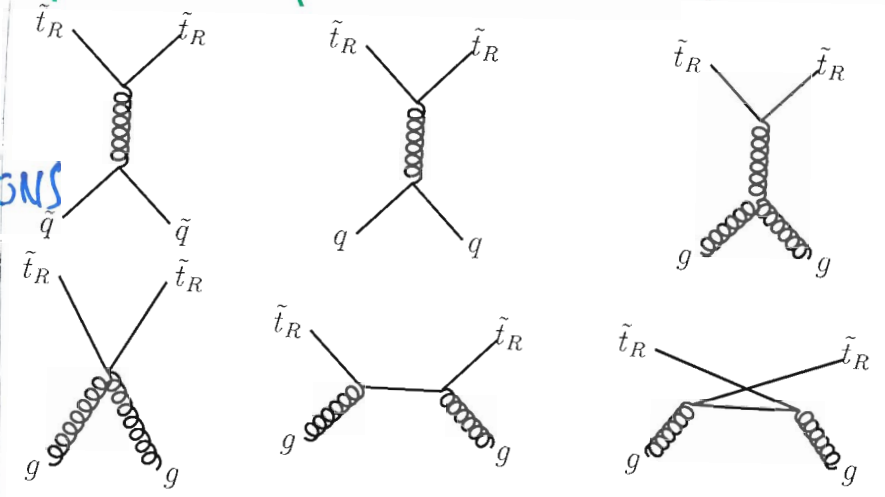
MSSM: (i) "heavy": T, \tilde{t}_R, W

P. JOHN
M.B. SEAR

TWO HIGGSSES \Rightarrow APPROX.: ONE HIGGS WITH FIXED $\tan\beta$ ✓ (! BUT CP-Viol.)..



COLLISIONS



Wall velocity for increasing stop mass parameter m_U .

The diagram is calculated for $\tan\beta = 3$, $A_t = \mu = 0$, and $m_A = 400\text{GeV}$.

COMBINE (i) AND (ii) !?

(ii) GAUGE BOSON PLASMA IN HOT PHASE: OVERDAMPED

$$\pi m_D^2 / 8k \frac{df}{dt} = - \underbrace{(k^2 + m^2)}_{E^2} f + \text{NOISE} \approx E^2 \delta f \quad (f \sim A^2)$$

G. MOORE
(\rightarrow ARNOLD SON, YAFFE)

\Rightarrow FRICTION $\frac{3m_D^2 T}{16\pi L} \int_0^1 \frac{(1-x) dx}{x} \quad (x = \phi/\phi_0)$

IR DIVERGENT

\rightarrow CUTOFF AT $\phi/\phi_0 = \sqrt{\Lambda/g}$

INFRARED DOMINATED!

ESTIMATE: $v_w \approx 0.1$

FOR $\lambda/g \approx 0.04$

⇒ TRANSPORT EQS. IN PRESENCE OF MOVING WALL

"THICK WALL" $L \gg$ MEAN FREE PATH ; $p \sim T \gg L^{-1}$

RELAT. THERMAL PART.

↪ QUASICLASSICAL, DERIVATIVE EXPANSION

NEED ORDER \hbar - ~~CP~~ DIRAC PARTICLES! (↪ WKB?!)
exp. \hbar^2

SAFE FRAMEWORK: KADANOFF BAYM EQS.

$$\left(\cancel{k} + \frac{i}{2} \cancel{\gamma} - P_L m e^{-\frac{1}{2} \vec{\partial} \cdot \partial_k} - P_R m^+ e^{-\frac{1}{2} \vec{\partial} \cdot \partial_k} \right) g^<(k, X) = \text{COLLISION TERM}$$

↑ EXPAND! (Order \hbar^2)

KANUKLINEN
 PROKOPEC
 SCH.
 WEINSTOCK

SPIN CONSERVED → DECOUPLING

$$m = |m| e^{i\theta(z)}$$

(ONE FLAVOR)

$$g_s^< = 2\pi f_s \delta(k_0 - \omega_s)$$

$$\omega_s = \frac{(k_\perp^2 + k_\parallel + |m|^2)^{1/2} - s |m|^2 \theta'}{2\omega_0 (\omega_0^2 - k_\parallel^2)^{1/2}}$$

⇒ KINETIC EQ.

$$\left| \frac{k_x}{\omega_s} \partial_x f_s + \boxed{F_s} \partial_{k_x} f_s = \text{COLL. T.} \right|$$

WITH SEMICLASSICAL FORCE

$$F_s = - \frac{|m|^2}{2\omega_s} + s \frac{(|m|^2 \theta')'}{2\omega_0 (\omega_0^2 - k_\parallel^2)^{1/2}}$$

DRIVING TERM

CP-VIOLATING

WKB ✓
 S. HUBER

MSSM

$$\Psi_R = \begin{pmatrix} \tilde{W}_L^+ \\ \tilde{h}_{1,R} \end{pmatrix} \quad \Psi_L = \begin{pmatrix} \tilde{W}_R^+ \\ \tilde{h}_{2,L} \end{pmatrix} \quad m(z) = \begin{pmatrix} M_2 & g H_2^*(z) \\ g H_1^*(z) & \mu_c \end{pmatrix}$$

$$g^< = g_{eq}^< + \delta g \quad \text{LINEAR RESPONSE}$$

→ KINETIC EQS. WITHOUT USE OF IN GENERAL NONALGEBRAIC CONSTRAINTS

T. KONSTANTIN
T. PROKOPEC
H.G. SCH.

$$k_z \partial_z \delta g + \frac{i}{2} [m^2, \delta g] + k_0 \Gamma \delta g = S(g_{eq}^<)$$

(PHEN. DAMPING BY QLT.)

FIRST ORDER IN GRAD. EXPANSION

(IN MASS EIGENBASIS)

OSCILLATIONS IN OFF-DIAGONAL DENSITIES!

CP VIOLATING SOURCES $S_{\mu}^{a,b,c}$

$$(*) S_{\mu}^c = -2g^2 T_c^{-2} \text{Im}(M_2 \mu_c) (H_2 \partial_z H_1 - H_1 \partial_z H_2) (m_{(1)}^0 + 4m_{(2)}^3)$$

SECOND ORDER IN GRAD. EXPANSION:

$$[S_0^d = 2V_w g^2 T_c^{-4} \text{Im}(M_2 \mu_c) (H_2 \partial_z^2 H_1 + H_1 \partial_z^2 H_2) \varphi_{(1)}^3]$$

MOM. INTEGRALS OF EQUIL. DISK.

SEMICLASSICAL FORCE

IN (*) FOR MAXIMUM: ALMOST MASS DEGENERATE CHARGINOS!
OTHERWISE STRONG DAMPING!!!

DIFFERS FROM!
CARENA, FORENO, QUIROS
SECO, WAGNER

$$\eta_{10} = \frac{n_B}{S_\gamma} \times 10^{10}$$

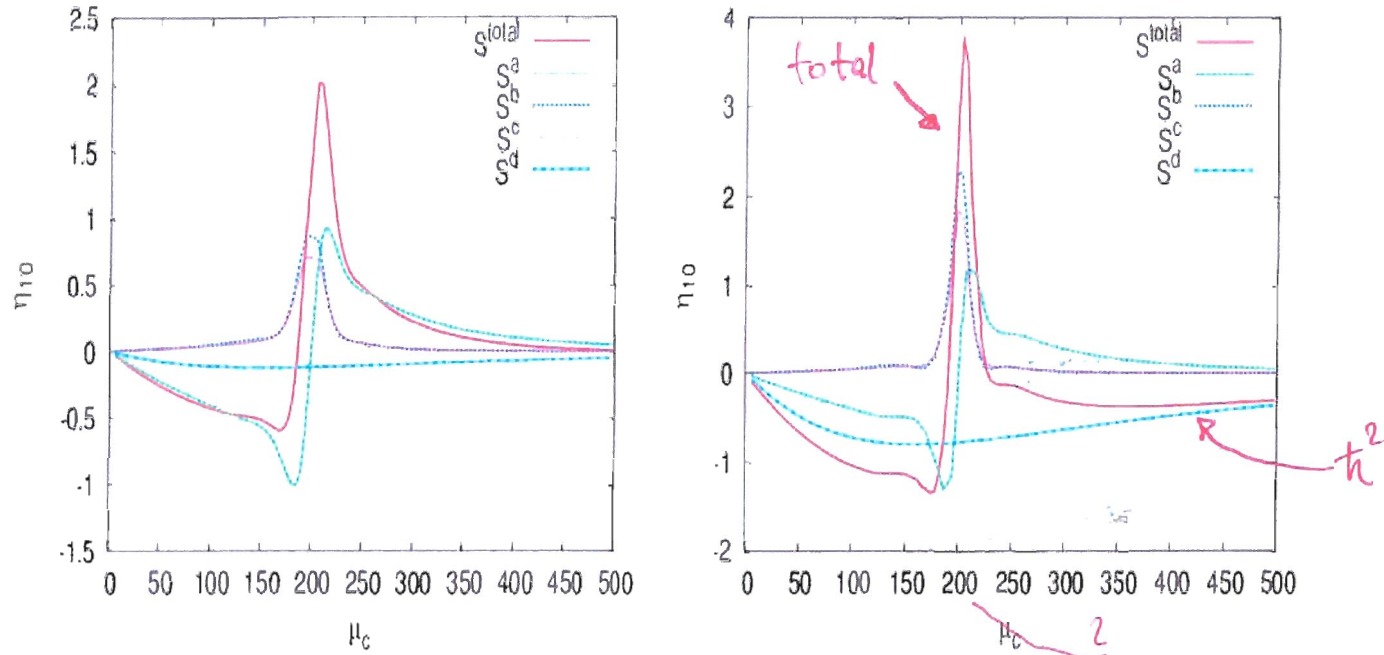


FIG. 2: This plot shows the first and second order sources as a function of μ_c with $M_2 = 200$ GeV. The plot on the left are the sources with the damping, $\Gamma = \alpha_w T_c$, while on the right plot, $\Gamma = 0.25\alpha_w T_c$.

MAXIMAL CP-VIOLATION ASSUMED



MSSM

T. KONSTANTIN
T. PROKOPEC
M.G. SCH.
M. SECO

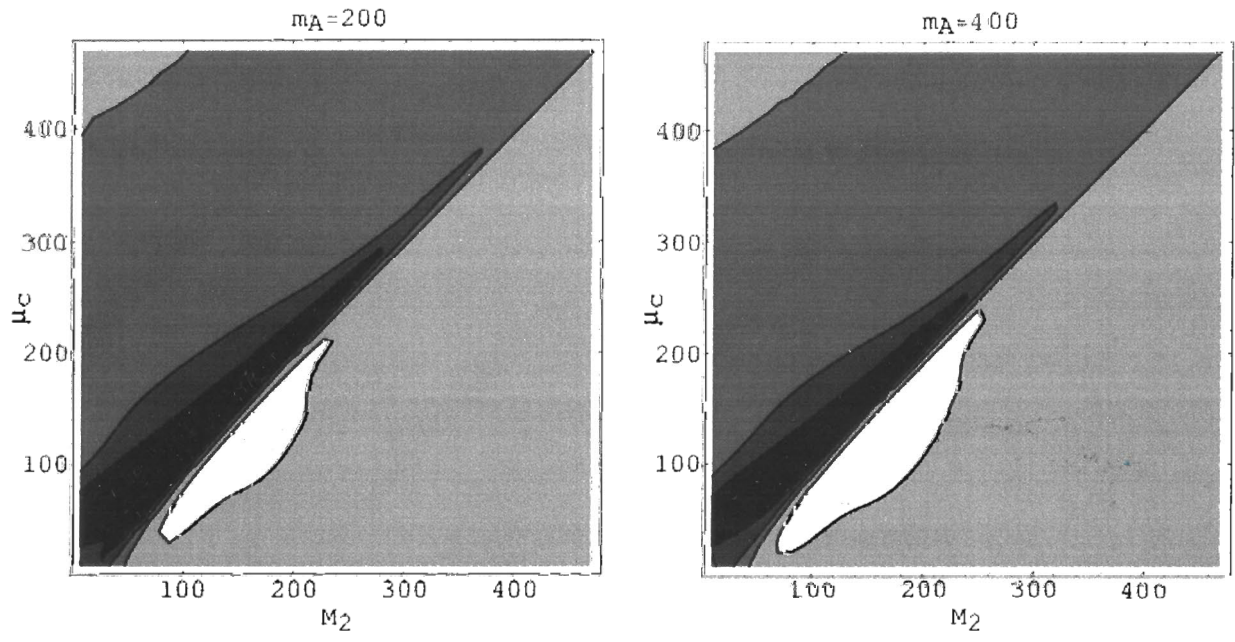


FIG. 5: The baryon-to-entropy ratio $\eta_{10} = 10^{10} \times \eta$ in the (M_2, μ_c) parameter space from $(0 \text{ GeV}, 0 \text{ GeV})$ to $(400 \text{ GeV}, 400 \text{ GeV})$. For the left plot the value $m_A = 200 \text{ GeV}$ is used, for the right plot $m_A = 400 \text{ GeV}$. The black region denotes $\eta_{10} > 1$, where baryogenesis is viable. The other four regions are bordered by the values of η_{10} , $\{-0.5, 0, 0.5, 1\}$, beginning with the lightest color.

MAXIMAL CP-VIOLATION

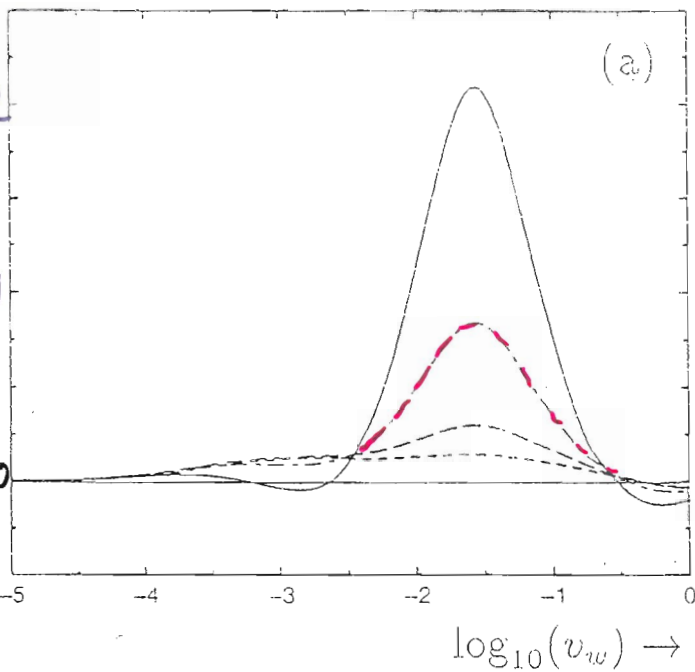
! RESTRICTIONS By exp. n/e - ELECTRIC DIPOLE LIMITS
 \sim CP-VIOL. PHASE < 0.1

$|d_e| \lesssim 1.6 \cdot 10^{-27} \text{ ecm}$
 Regan et al PRL 88
 071805, 2002

$$\eta^* = \eta_B / \eta_B^{\text{observed}}$$

NMSSM

S. HUBER
H. POH.



η^*

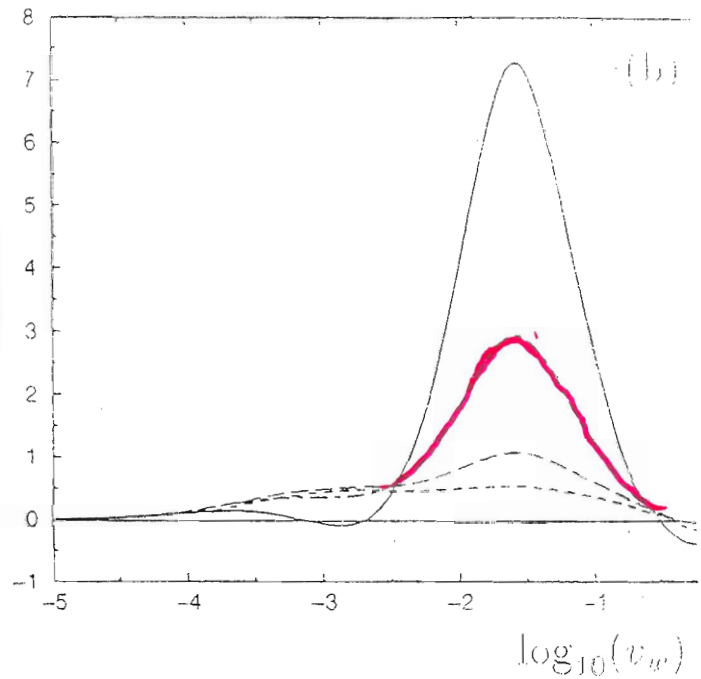


Figure 10: The chargino contribution to the baryon asymmetry in units of 2×10^{-11} as a function of the wall velocity for different values of the wall thickness $L/T, 10/T, 5/T, 3/T$ (from below). We use the squark spectrum C and the explicit CP-violation considered in the context of fig. 6. (b) The same quantity for a transitionally CP-violating bubble wall of fig. 7 and the squark spectrum D.

$$\eta_B = \frac{n_B}{n_\gamma} = \frac{135 \Gamma_{\text{SPHALETON}}}{g^* 2\pi^2 v_w T} \int_0^\infty d\bar{z} \mu_{B_L}(\bar{z})$$

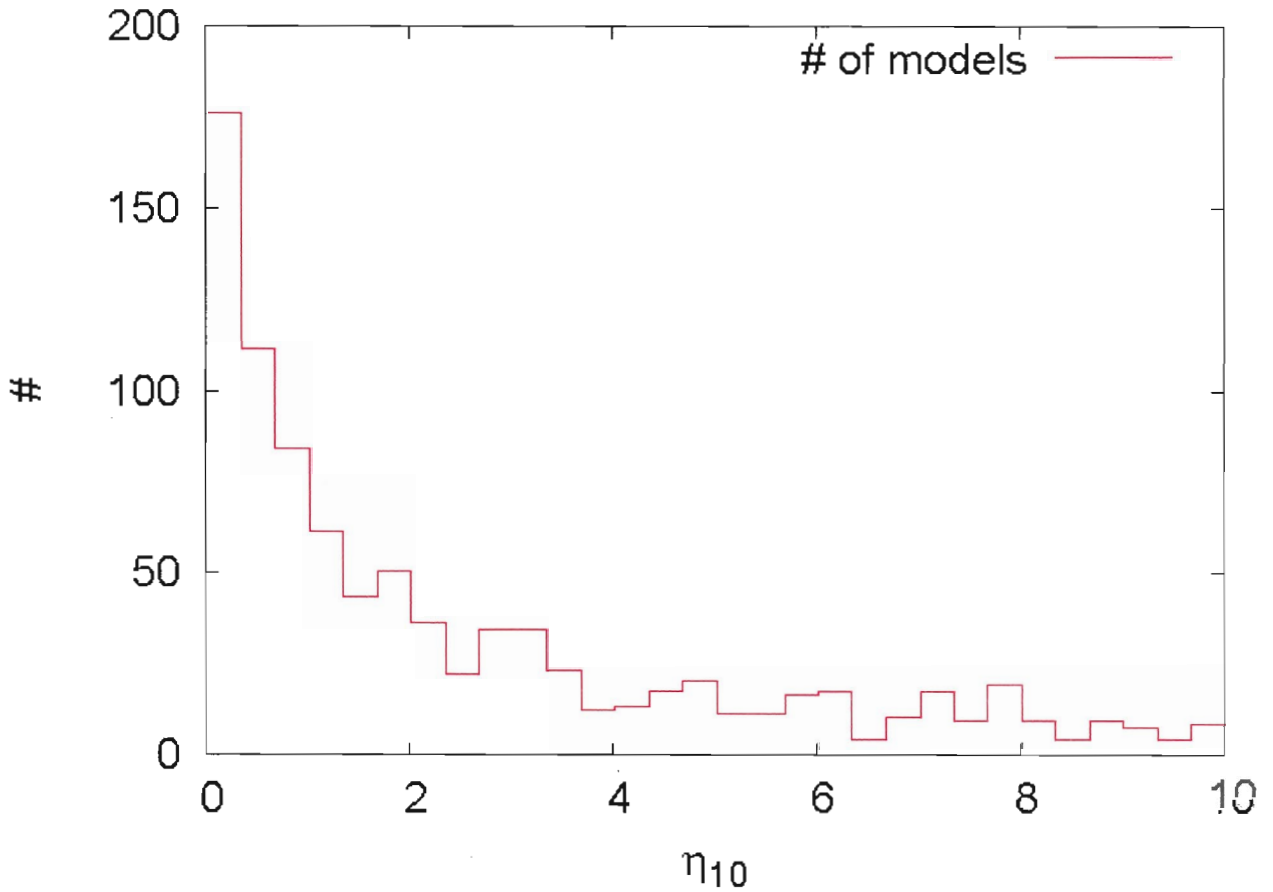
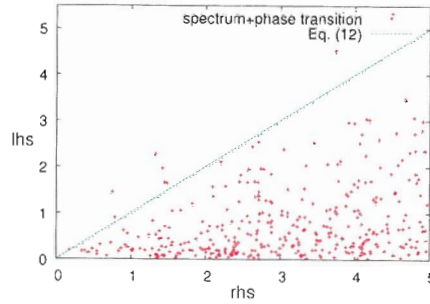
IN FRONT OF WALL

(η^* IN UNITS OF 2×10^{-11} !)

The nMSSM allows for a strong first order phase transition (PT) with only tree level dynamics. The criterion for a first order PT is on tree level given by

$$m_s^2 < \frac{1}{\tilde{\lambda}} \left| \frac{\lambda^2 t_s}{m_s} - m_s \tilde{a} \right|. \quad (12)$$

what tends to persist even if one includes the one-loop potential.



Produced baryon asymmetry in random nMSSM models.

HUBER '06
 KONSTANDIN
 PROKOPEC
 SCHMIDT

$\eta_{10} \approx 1$ exp.

CONCLUSIONS + FINAL REMARKS

- INGREDIENTS OF MODELS FOR AN ELECTROWEAK PHASE TRANSITION CAN BE TESTED IN FORTH COMING EXPERIMENTS
- IN MOST ELWK. BARYOGENESIS MODELS A VERY STRONG PHASE TRANSITION IS REQUIRED - "BEYOND THE SM" OR "EXTENDED" SM IS MANDATORY.
MSSM RULED OUT?! NMSSM, μ MSSM MUCH MORE FLEXIBLE (FOR ELWK. BARYOGENESIS)
(NONTHERMAL)?!
LEPTOGENESIS?
(PROMISING?)
- CP VIOLATION ALSO BY DERIVATIVE TERMS IN EFF. ACTION
TRANSITIONAL ~~CP~~; edm BOUNDS
- ELECTROWEAK BARYOGENESIS A WELL FOUNDED PROCEDURE IN A GIVEN MODEL
- USEFUL FOR PRODUCTION OF GRAVITY WAVES, PRIMORDIAL MAGN. FIELDS?
NEED SMALL V_W FOR BARYOGENESIS (CHARGE TRANSPORT MECH.)
LARGE V_W ? ('RUN AWAY' BODEKER, MOOR '09) ← WISHFUL THINKING?

The current measurement bound of the electron electric dipole moment (EDM)

Regan et al, Phys. Rev. Lett. 88:071805, 2002

$$|d_e| \quad 1.6 \times 10^{-27} \text{ ecm}$$

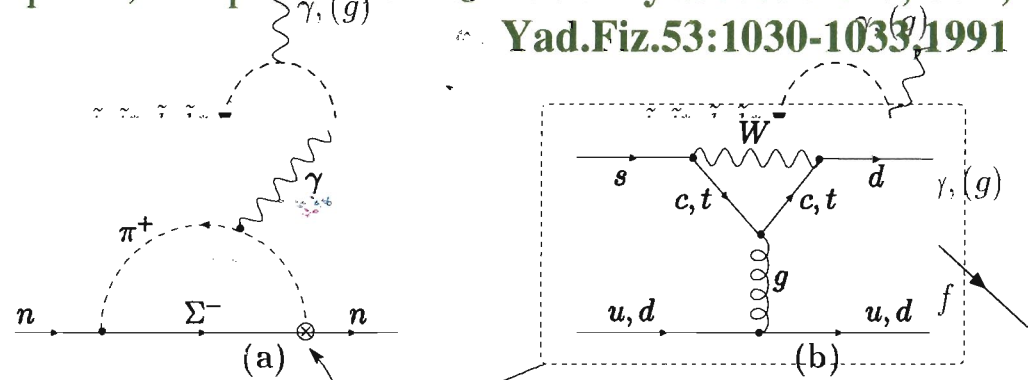
The standard model (MSM) value for eEDM (4 loop)

$$d_e^{\text{CKM}} \quad 1 \times 10^{-38} \text{ ecm}$$

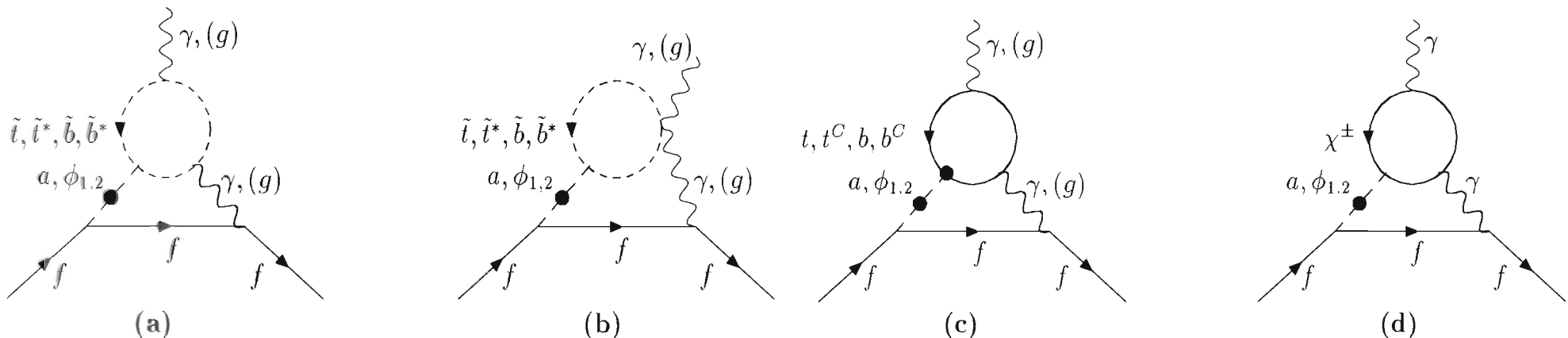
Pospelov, Khriplovich, Sov.J.Nucl.Phys.53:638-640,1991,
Yad.Fiz.53:1030-1039,1991

The standard model (MSM) value for neutron EDM (2 loop penguin)

$$d_n^{\text{CKM}} \sim 1 \times 10^{-32} \text{ ecm}$$



The MSSM 2 loop Higgs contribution for electron EDM



ALTERNATIVE :

STRENGTHEN $-\phi^3$ -TERM; FURTHER CP-VIOL.

- MSSM, NMSSM, n MSSM, 2-HIGGS...
- LEPTOGENESIS, COHERENT BARYOGENESIS, AFFLECK-DINE BARYOGENESIS, ...

MY USUAL DIRECTION
UP TO NOW

HOW TO GET (i) STRONG NONEQUILIBRIUM AND (ii) ENOUGH CP-VIOLATION IN (ALMOST) THE STANDARD MODEL?

(i) SWITCH SIGN OF $\phi^+\phi$ -TERM (HIGGS MASS)

(ii) INTEGRATE OUT FERMIONS PROPERLY

A. HERNANDEZ
T. KONSTANTIN
M.G. SCH.

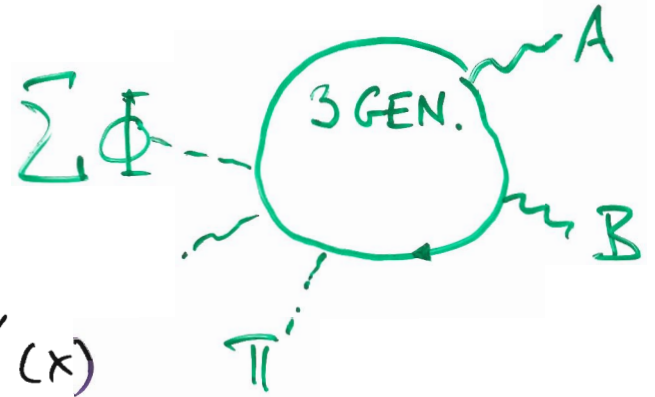
DERIVE CP VIOLATING TERM(S) IN BOSONIZED SM

- INTEGRATE OUT FERMIONS
- OBTAIN EFFECTIVE ACTION

$$-W = \log \text{Det} [\mathcal{O}]$$

$$\mathcal{O} = \not{\partial} - i \not{\Phi}(x) - \gamma_5 \not{\Pi}(x) - \not{A}(x) - \gamma_5 \not{B}(x)$$

(EUCLIDEAN $t = -t_E$, $\not{\partial} = i \not{\partial}_E$)



$$-W = \underbrace{\log |\text{Det} [\mathcal{O}]|}_{-W^+} + i \underbrace{\arg (\text{Det} [\mathcal{O}])}_{-iW^-}$$

EVEN / ODD NUMBER OF γ_5
IN $W^+/-$, REAL / IMAG.

WORLD LINE METHOD OPTIMAL

(PROPAGATING FERM.
IN THE BACKGROUND OF X-DEP.
SCALAR AND GAUGE FIELDS)

GET DERIVATIVE EXPANSION

MONDRAGON
NELLEN
SCH.
SCHUBERT

→ D'HOOVER
GAGNIER
LATER: VERY ELEGANT!

MAIN POINT FOR FERMIONS IN THE LOOP: TRANSLATE γ -MATRICES TO GRASSMANN VARIABLES IN ("FIRST QUANTIZED") WORLDLINE PATH INTEGRAL. NEED EVEN NUMBER OF γ 'S. (\leadsto SECOND ORDER FORM.) ... FINKLER
 (\leadsto COHERENT STATE FORM.)

REAL PART: (i) $W^+ = -\frac{1}{2} \log \text{Det} [\mathbb{O} \mathbb{O}^+] = -\frac{1}{2} \text{Tr} \log [\mathbb{O} \mathbb{O}^+]$

(ii) $\mathbb{O} \rightarrow \tilde{\Sigma} \equiv \begin{pmatrix} 0 & \mathbb{O} \\ \mathbb{O}^+ & 0 \end{pmatrix}; W^+ = -\frac{1}{4} \log \text{Det} [\tilde{\Sigma}^2] = \dots$

(8x8) Γ_A MATRICES ($A=1, \dots, \underline{6}$)

$$\tilde{\Sigma} = \Gamma_\mu (p_\mu - A_\mu) - \Gamma_6 \Phi - \Gamma_5 \Pi - i \Gamma_\mu \Gamma_5 \Gamma_6 B_\mu$$

$$\begin{pmatrix} 0 & \gamma_\mu \\ \gamma_\mu & 0 \end{pmatrix} \quad \begin{pmatrix} 0 & \gamma_5 \\ \gamma_5 & 0 \end{pmatrix} \quad \begin{pmatrix} 0 & i 1_4 \\ -i 1_4 & 0 \end{pmatrix}$$

BASIS WHERE $i \Gamma_5 \Gamma_6$ DIAGONAL \Rightarrow CHIRAL COVARIANT

$$\tilde{\Sigma} = \begin{pmatrix} \gamma_\mu (p_\mu - A_\mu^L) & -i \gamma_5 H \\ +i \gamma_5 H & \gamma_\mu (p_\mu - A_\mu^R) \end{pmatrix}$$

$$A^{L/R} = A \pm \frac{1}{2} B$$

$$H = \Phi - i \Pi$$

$\tilde{\Sigma}^2$ STILL WITH ODD COMB. OF γ ($\gamma_5 = \dots$)

(iii) $\gamma_A \rightarrow \hat{T}_A = T_A \otimes \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \quad A=1, \dots, 5$

$$W^+ = \frac{1}{8} \int_0^{\infty} \frac{dT}{T} \text{Tr}_{T, X} \exp(-T \hat{\Sigma}^2)$$

SCHWINGER PROPER TIME

$$\left[\hat{\Sigma}^2 = (p-A)^2 + \mathcal{H}^2 + \frac{i}{2} \Gamma_\mu \Gamma_\nu \tilde{F}^{\mu\nu} + i \Gamma_\mu \Gamma_5 (\mathcal{D}_\mu \mathcal{H}) \right.$$

$$\left. A_\mu = \begin{pmatrix} A_\mu^L & 0 \\ 0 & A_\mu^R \end{pmatrix}; \mathcal{H} = \begin{pmatrix} 0 & iH \\ -iH & 0 \end{pmatrix} \right.$$

(iv) WORLDLINE INTEGRAL

$$W^+ = \frac{1}{8} \int_0^{\infty} \frac{dT}{T} \int_{X(T)=X(0)} \mathcal{D}X \int_A^T \mathcal{D}\psi_A \text{tr P exp} \left(- \int_0^T d\tau \mathcal{L}(\tau) \right)$$

$\psi(T) = -\psi(0)$

ANTI-PERIODIC GRASSMANN

$$\mathcal{L}(\tau) = \dot{X}^2/4 - i \dot{X}_\mu A^\mu + \frac{1}{2} \psi_A \dot{\psi}_A + \mathcal{H}^2 + i \psi_\mu \psi_\nu \tilde{F}^{\mu\nu}$$

$$+ 2i \psi_\mu \psi_5 (\mathcal{D}_\mu \mathcal{H}) \quad A=1, \dots, 5$$

EVALUATE : $X(\tau) = X_0 + y(\tau)$ WITH $\int_0^T dt y(\tau) = 0$

$$\rightarrow \mathcal{D}X = \mathcal{D}y dK_0$$

$$\langle y(\tau_1) y(\tau_2) \rangle = \frac{(\tau_1 - \tau_2)^2}{T} - |\tau_1 - \tau_2|$$

$$\langle \psi_A(\tau_1) \psi_B(\tau_2) \rangle = \frac{1}{2} \delta_{AB} \text{sign}(\tau_1 - \tau_2)$$

ψ ANTICOMM.

IMAGINARY PART :

CALCULATE CURRENTS (DER. OF ACTION W.R.T. FIELDS)

$$\delta W^- = \frac{1}{2} \delta (\log \text{Det} [0] - \log \text{Det} [0^+])$$

$$= \frac{1}{2} \text{Tr} \left(\delta 0 \frac{1}{0} - \delta 0^+ \frac{1}{0^+} \right)$$

$$= \frac{1}{2} \text{Tr} \begin{pmatrix} 0 & \delta 0 \\ -\delta 0^+ & 0 \end{pmatrix} \begin{pmatrix} 0 & 1/0^+ \\ 1/0 & 0 \end{pmatrix} = \frac{1}{2} \text{Tr} \begin{pmatrix} 1_4 & \\ & 1_4 \end{pmatrix} \delta \Sigma \Sigma^{-1}$$

$$= \frac{1}{4} \text{Tr} \chi [\delta \Sigma, \Sigma] \Sigma^{-2} = \frac{1}{4} \text{Tr} \int_0^{\infty} dt \chi [\delta \Sigma, \Sigma] \exp(-t \Sigma^{-2})$$

$$\Sigma \rightarrow \tilde{\Sigma} \quad (\chi \rightarrow \tilde{\chi} = \chi \gamma_5), \quad \gamma_A \rightarrow \tilde{\Gamma}_A \text{ AS BEFORE}$$

- PREJUDICE CP WITH χ
- WRONG? SALCEDO
- LOWEST ORDER IN DER. EXP \rightarrow J. SHIT ($\neq 0!$)

$$\Rightarrow \delta W^- = \frac{i}{8} \text{Tr} \int_0^\infty dt \left(T_7 \right) T_6 \chi W(T) \exp(-T \hat{\Sigma}^2)$$

↑
CHANGES BOUNDARY COND. FOR GRASSMANN ψ

FOR $\delta \hat{\Sigma}^2 = -\gamma_\mu \delta A_\mu$

$$\delta W^- = \frac{1}{8} \text{tr} \int_0^\infty dt \int \mathcal{D}X \int \mathcal{D}\psi_A \chi W(T) P \exp \exp(-\int_0^T dt \mathcal{L}(\tau))$$

$$W(T) = -\psi_i \psi_c \psi_v \delta \left\{ A_\mu \dot{X}_\nu - 2i \psi_s \mathcal{D}_\mu \delta x_\mu - 2\psi_\mu \left[\delta x_\mu, \mathcal{L} \right] \right\}$$

NOW PERIODIC \rightarrow ZERO MODES

$$\psi_A(t) = \psi_A^0 + \psi_A'(t)$$

STATE CORRELATOR AS \dot{X}

PRODUCE ϵ -TENSOR!

(\therefore EACH ZERO MODE ONCE!)

HAVE TO INTEGRATE δW^- TO OBTAIN EFF. ACTION

CONSISTENT VS. COVARIANT CURRENT

$$W^- = W_c^- + \Gamma_{\text{gauge}} W_{\text{FH}}^- \text{ - ACTION}$$

(\leadsto SALCEDO CONTRIBUTES ONLY LOWEST ORDER IN DER. EXP)

INSATE FOR ACTION \rightarrow DERIV. \rightarrow COMPARE

EVALUATION

A. HERNANDEZ, T. KONSTANTIN, H.G. SCH.

- \mathcal{H} is a 3×3 MASS MATRIX (X-DEP.!) - UNITARY GAUGE

DIAGONALIZATION WITH CKM MATRIX \mathcal{V}

- $SU(2)_L$ GAUGE FIELD STRENGTH IN FLAVOR SPACE $\overline{F}_L = \begin{pmatrix} F_0 & F^+ \mathcal{V} \\ \mathcal{V}^+ F & F_0 \end{pmatrix}$

- USE "LABELED OPERATOR NOTATION" FEYNMAN
...SALCEDO

E.G. $\langle \epsilon^{\mu\nu\lambda\sigma} \mathcal{H} F_{\mu\nu} \mathcal{H}^3 F_{\lambda\sigma} \rangle = \langle \epsilon^{\mu\nu\lambda\sigma} \underbrace{m_1}_{\dots} \underbrace{m_2^3}_{\dots} F_{\mu\nu} F_{\lambda\sigma} \rangle$

$$\nabla f(m_1, m_2) := \frac{f(m_1) - f(m_2)}{m_1 - m_2} ; \quad \mathcal{D}_\mu f(m) = \nabla f(m_1, m_2) \mathcal{D}_\mu \mathcal{H}$$

$$\mathcal{D}_\mu (\mathcal{H}^3) = (m_2^2 + m_1 m_2 + m_1^2) \mathcal{D}_\mu \mathcal{H} = \frac{m_1^3 - m_2^3}{m_1 - m_2} \mathcal{D}_\mu \mathcal{H}$$

$$\mathcal{L}(\tau) = \underbrace{\dot{\chi}^2/4 + \frac{1}{2} \Psi_A \dot{\Psi}_A}_{\mathcal{L}_0(\tau)} + \mathcal{H}^2 \quad \text{GAUSSIAN}$$

$$\begin{aligned} \text{P exp} \left[- \int_0^T dt \mathcal{L}(t) \right] &= \text{exp} \left[- \int_0^T dt \mathcal{L}_0(t) \right] \left(- \int_0^T dt_1 e^{-T m_1^2 - \tau_1 (m_2^2 - m_1^2)} \mathcal{L}_1(t_1) \right. \\ &\quad \left. + \int_0^T dt_1 \int_0^{\tau_1} dt_2 \text{exp} \left[-T m_1^2 - \tau_1 (m_2^2 - m_1^2) - \tau_2 (m_3^2 - m_2^2) \right] \mathcal{L}_1(t_1) \mathcal{L}_1(t_2) + \dots \right) \end{aligned}$$

$\left. \right) = \nabla m^3(m_1, m_2) \mathcal{D}_\mu \mathcal{H}$

- EXPAND AROUND X_0

mostly "AS USUAL"

- DO ψ -INTEGRATION

- DO ψ -INTEGRATION SATURATING ZERO MODES AND

- CONTRACTING WITH FERMIONIC GREENSF.

- DO τ, T INTEGRATION

HEAVY COMPUTER ALGEBRA (A. HERNANDEZ..)

NEXT TO LEADING ORDER: ALMOST ALL ~~CP~~ CONTR. CANCEL
REMAINS:

$$S_{\cancel{CP}} = \frac{1}{8(4\pi)^2} \frac{3}{16} \int K^{CP} \epsilon^{\mu\nu\lambda\sigma} \frac{1}{\tilde{m}_c^2} \times$$

$$\int d^4x \left[z_\mu W_{\nu\lambda}^+ W_\alpha^- \times (W_\sigma^+ W_\alpha^- + W_\alpha^+ W_\sigma^-) + c.c. \right]$$

$K^{CP} \approx 9.87$ IN 'BROKEN' PHASE

$$\int = S_1^2 S_2 S_3 C_1 C_2 C_3 \sin\delta \approx (3.0 \pm 0.3) \times 10^{-5}$$

$$W_{\mu\nu}^+ = \frac{\phi^+ W_{\mu\nu} \tilde{\phi}}{\phi^+ \phi} \dots$$

$$W_\mu^+ = \frac{\phi^+ \partial_\mu \tilde{\phi}}{\phi^+ \phi}$$

$$z_\mu = W_\mu^3 \frac{1}{f} = \frac{\phi^+ \partial_\mu \phi - \tilde{\phi} \partial_\mu \tilde{\phi}}{2\phi^+ \phi} \times 10^{-5}$$

($\tilde{\phi} = \epsilon \phi^*$)