

Universality or non-universality in applications of gauge/gravity duality

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Based on joint work with

M. Ammon, V. Grass, M. Kaminski, P. Kerner, S. Müller, A. O'Bannon, J. Shock, M. Strydom, H. Zeller

Motivation:

Gauge/gravity duality: New tools for strongly coupled systems

Famous result: Shear viscosity/Entropy density

Kovtun, Son, Starinets

$$\frac{\eta}{s} = \frac{1}{4\pi} \frac{\hbar}{k_B T}$$

From 'Planckian time' $\tau_P = \frac{\hbar}{k_B T}$, **Universal result**

This talk:

Deviations from this result at leading order in λ and N

Search for similar result in condensed matter context

Holographic proof of universality relies on space-time isotropy

Key ingredient for changes to the
universal result: Spacetime anisotropy

Rotational invariance broken

Holographic p-wave superfluids/superconductors

Nematic phase (Condensate breaks rotational symmetry)

Outline

- Holographic superconductors
- Transport coefficients in anisotropic systems
- Universality in AdS/CMT
- Condensates at finite magnetic field

Reminder: Holographic Superfluids/Superconductors

- Holographic Superconductors from charged scalar in Einstein-Maxwell gravity
(Gubser; Hartnoll, Herzog, Horowitz)
- p-wave superconductor
current dual to gauge field condensing
(Gubser, Pufu)
SU(2) Einstein-Yang-Mills model

s-wave superconductor:

$$\mathcal{L} = R + \frac{6}{L^2} - \frac{1}{4} F^{ab} F_{ab} - V(|\psi|) - |\nabla\psi - iqA\psi|^2$$

Operator \mathcal{O} dual to scalar ψ condensing

Herzog, Hartnoll, Horowitz 2008

p-wave superconductor:

$$S = \frac{1}{2\kappa^2} \int d^4x \left[R - \frac{1}{4} (F_{\mu\nu}^a)^2 + \frac{6}{L^2} \right]$$

Current J_x^1 dual to gauge field component A^{1x} condensing

Gubser, Pufu 2008

P-wave superconductor from probe branes

Ammon, J.E., Kaminski, Kerner 0810.2316, 0903.1864

- A holographic superconductor with field theory in $3+1$ dimensions for which
- the dual field theory is explicitly known
- there is a qualitative ten-dimensional string theory picture of condensation

D7 probe branes

On gravity side:

Probe brane fluctuations described by **Dirac-Born-Infeld action**

$$S_{\text{DBI}} = -T_{D7} \int d^8\xi \text{Str} \sqrt{|\det(G + 2\pi\alpha'F)|}$$

On field theory side: Lagrangian explicitly known

$$\mathcal{L} = \mathcal{L}_{\mathcal{N}=4} + \mathcal{L}(\psi_q^i, \phi_q^i)$$

Turn on finite temperature and isospin chemical potential:

Finite temperature: Embed D7 brane in black hole background

Isospin chemical potential: Probe of two coincident D7 branes

Additional symmetry $U(2) = SU(2)_I \times U(1)_B$

$$A_0^3 = \mu - \frac{\tilde{d}_0^3}{2\pi\alpha'} \frac{\rho_H}{\rho^2} + \dots, \quad A_3^1 = -\frac{\tilde{d}_1^3}{2\pi\alpha'} \frac{\rho_H}{\rho^2} + \dots$$

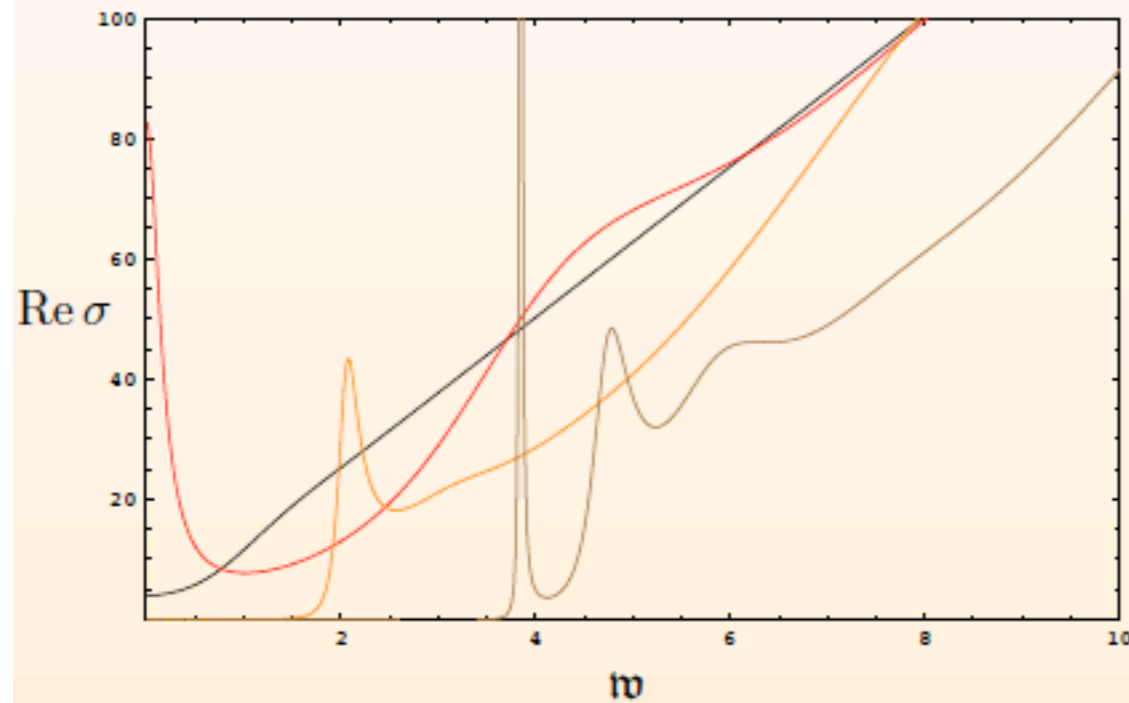
Condensate $\langle J_3 \rangle$, $J_3 = \bar{\psi}_d \gamma_3 \psi_u + \text{bosons}$

Calculate correlators from fluctuations

Conductivity

Frequency-dependent conductivity $\sigma(\omega) = \frac{i}{\omega} G^R(\omega)$

G^R retarded Green function for fluctuation a_2^3



Ammon, J.E., Kaminski, Kerner '08

$$\nu = \omega / (2\pi T)$$

T/T_c : Black: ∞ , Red: 1, Orange: 0.5, Brown: 0.28.

(Vanishing quark mass)

Interpretation: Frictionless motion of mesons through plasma

Effective 5d model \rightarrow anisotropic shear viscosity

Bottom-up: Including the backreaction

Ammon, J.E., Graß, Kerner, O'Bannon 0912.3515

Einstein-Yang-Mills-Theory with $SU(2)$ gauge group

$$S = \int d^5x \sqrt{-g} \left[\frac{1}{2\kappa^2} (R - \Lambda) - \frac{1}{4\hat{g}^2} F_{\mu\nu}^a F^{a\mu\nu} \right]$$

$$\alpha = \frac{\kappa_5}{\hat{g}}$$

$\alpha^2 \propto$ number of charged d.o.f./all d.o.f.

In presence of $SU(2)$ chemical potential, same condensation process as before

Hairy black hole solution

- metric ansatz

$$ds^2 = -N(r)\sigma(r)^2 dt^2 + \frac{1}{N(r)} dr^2 + \frac{r^2}{f(r)^4} dx^2 + r^2 f(r)^2 (dy^2 + dz^2)$$

with

$$N(r) = -\frac{2m(r)}{r^2} + r^2$$

AdS boundary $r = r_{\text{bdy}} \rightarrow \infty$ & black hole horizon $r = r_h$

- gauge field ansatz

$$A = \phi(r)\tau^3 dt + w(r)\tau^1 dx$$

Field Theory	\Leftrightarrow	Gravity
chemical potential μ $SU(2) \rightarrow U(1)_3$		$A_t^3 = \phi(r) \neq 0$ $SU(2) \rightarrow U(1)_3$
$\langle \mathcal{J}_1^x \rangle \neq 0$ $U(1)_3 \rightarrow \mathbb{Z}_2, SO(3) \rightarrow SO(2)$		$A_x^1 = w(r) \neq 0$ $U(1)_3 \rightarrow \mathbb{Z}_2, SO(3) \rightarrow SO(2)$

- $w(r_{\text{bdy}}) = 0 \Rightarrow$ **SSB** $U(1)_3 \rightarrow \mathbb{Z}_2$ & $SO(3) \rightarrow SO(2)$

\Rightarrow holographic p-wave superfluid with backreaction

- 5 fields: $\{\phi(r), w(r), \sigma(r), f(r), m(r)\}$

Variation of on-shell action at AdS boundary gives

- energy-momentum tensor

$$\langle \mathcal{T}_{\mu\mu} \rangle \propto T^4 \cdot \text{Func}(m_0^b, f_2^b), \text{ with: } \langle \mathcal{T}_{yy} \rangle = \langle \mathcal{T}_{zz} \rangle \neq \langle \mathcal{T}_{xx} \rangle$$
$$\langle \mathcal{T}_{\mu\nu} \rangle = 0 \text{ for } \mu \neq \nu$$

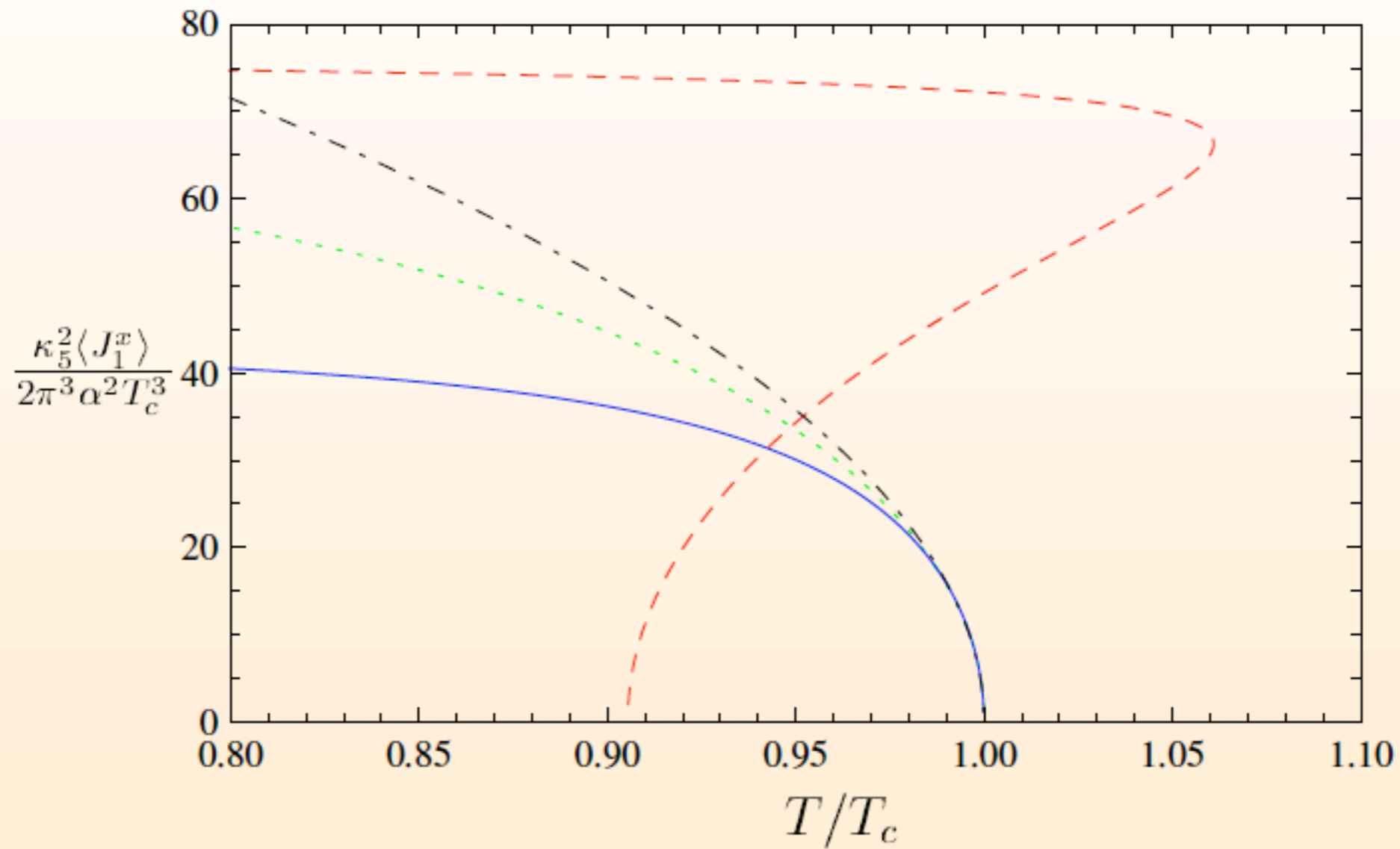
- density current

$$\langle \mathcal{J}_3^t \rangle \propto T^3 \phi_1^b$$

- condensate

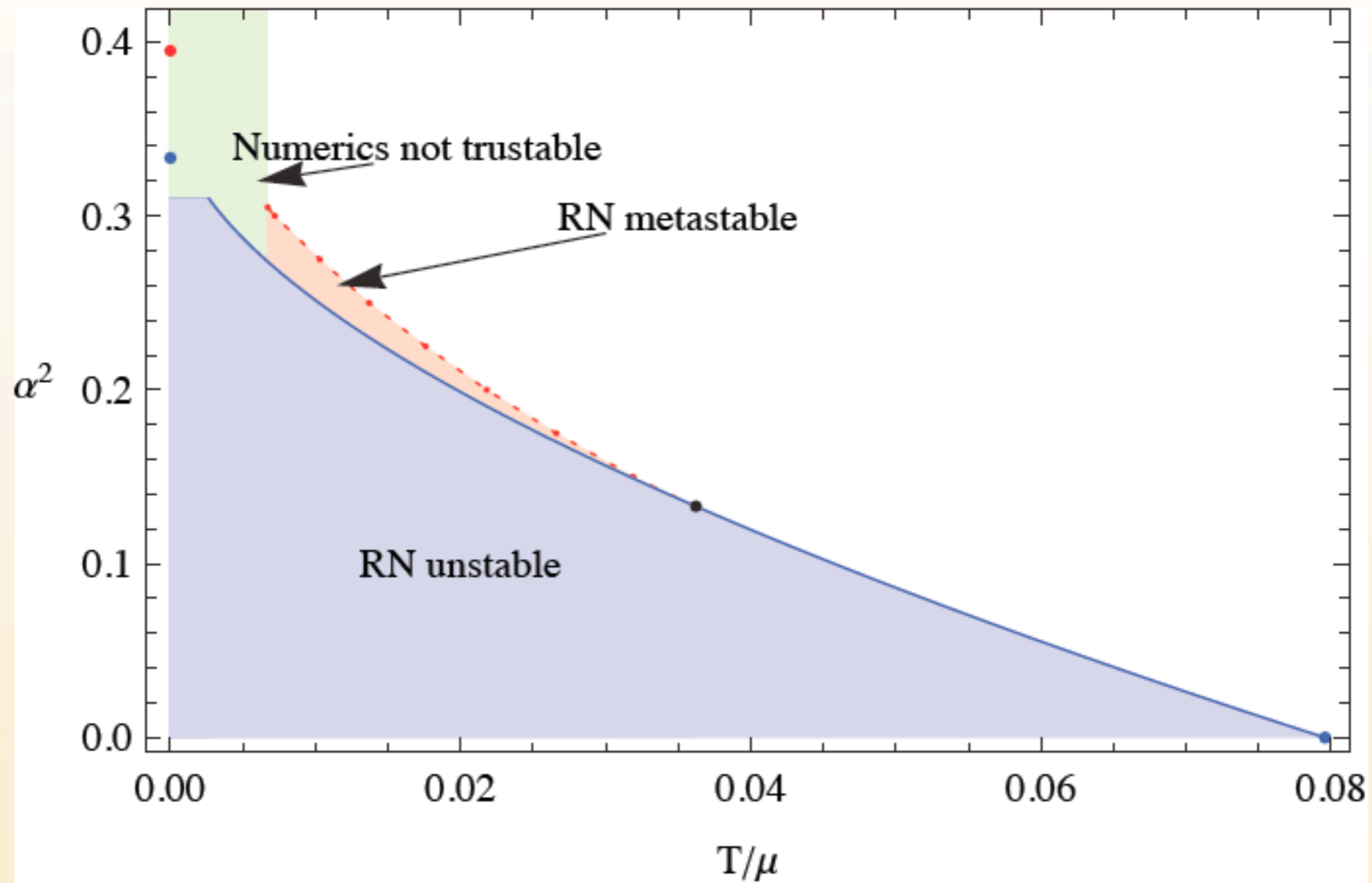
$$\langle \mathcal{J}_1^x \rangle \propto T^3 w_1^b$$

Phase transition



Phase transition becomes first order above α_{crit}

Phase diagram



Fluctuations about equilibrium

small perturbations:

- metric $\hat{g}_{MN} = g_{MN}(r) + h_{MN}(x^\mu, r)$
- gauge field $\hat{A}_M^a = A_M^a(r) + a_M^a(x^\mu, r)$
- x^μ -spacetime translational invariance still unbroken

⇒ Fourier decomposition of fluctuations possible:

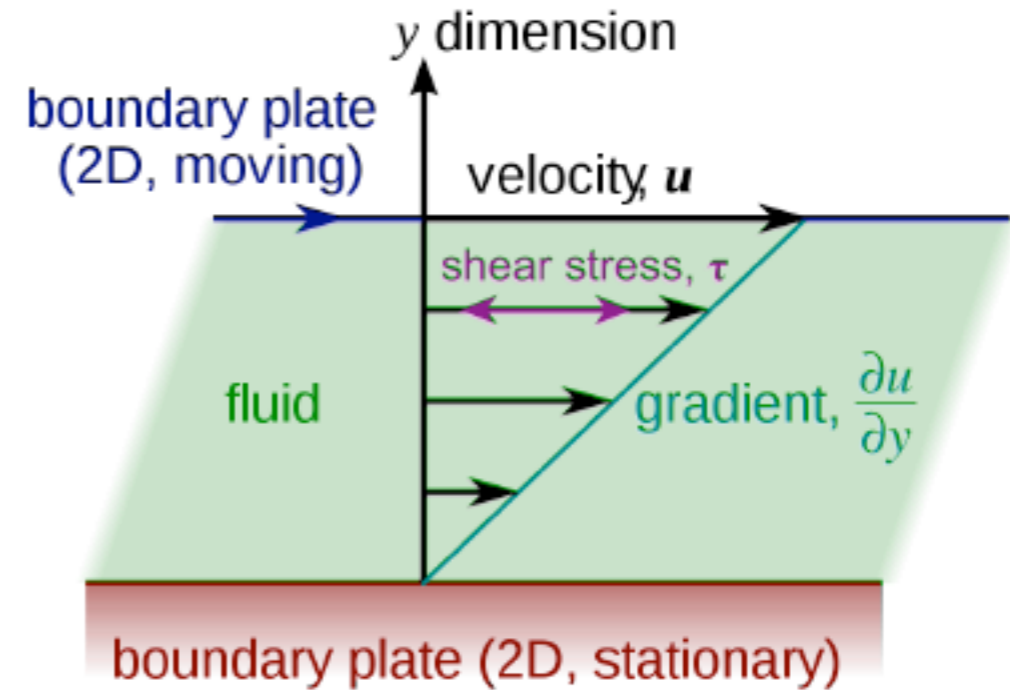
$$h_{MN}(x^\mu, r) = \int \frac{d^4 k}{(2\pi)^4} e^{ik_\mu x^\mu} h_{MN}(k^\mu, r)$$

$$a_M^a(x^\mu, r) = \int \frac{d^4 k}{(2\pi)^4} e^{ik_\mu x^\mu} a_M^a(k^\mu, r)$$

- $SO(2)$ symmetry ⇒ two distinct momenta needed: k_{\parallel} and k_{\perp}

Anisotropic shear viscosity

- viscosity tensor $\eta_{\alpha\beta\gamma\delta}$
- anisotropic systems:
21 components
- isotropic systems:
1 shear viscosity
- transversely isotropic systems:
2 shear viscosities



Holographic calculation: J.E., Kerner, Zeller 1011.5912; 1110.0007

Classification of Fluctuations

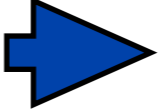
- set $k_{\perp} = 0$

⇒ classification under $SO(2)$ rotational symmetry around x-axis possible:

	dynamical fields	constraints	# physical modes
helicity 2	$h_{yz}, h_{yy} - h_{zz}$	none	2
helicity 1	$h_{ty}, h_{xy}; a_y^a$	h_{yr}	4
	$h_{tz}, h_{xz}; a_z^a$	h_{zr}	4
helicity 0	$h_{tt}, h_{xx}, h_{yy} + h_{zz}, h_{xt};$ a_t^a, a_x^a	$h_{tr}, h_{xr}, h_{rr}; a_r^a$	4

gauge choice $h_{Mr} = 0$ and $a_r^a = 0 \Rightarrow 14$ physical modes

Transport coefficients from Green functions

One non-trivial helicity 2 mode 
gives well-known result $\eta/s = 1/4\pi$

Helicity 1 modes:

- in $\vec{k} \rightarrow 0$ limit additional symmetry:
 $\mathbb{Z}_2: x \rightarrow -x, w \rightarrow -w$

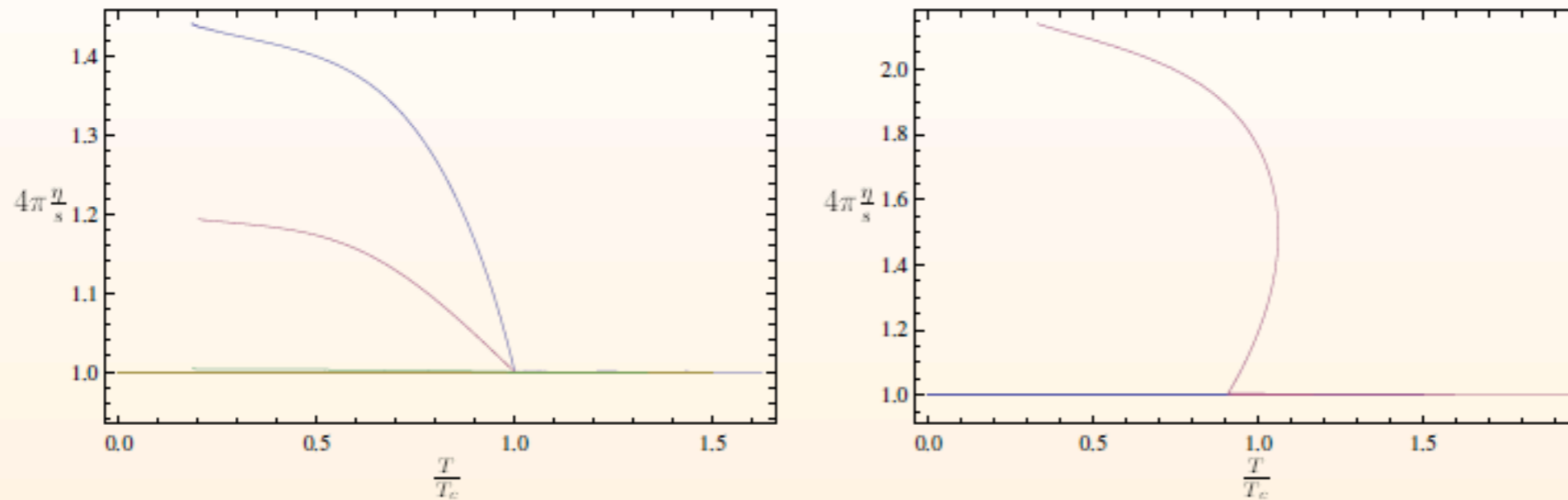
\Rightarrow helicity 1 modes decouple in 2 blocks:

even parity: $\{\Psi_t = g^{yy} h_{t\perp}, a_{\perp}^3, h_{r\perp}\}$

odd parity: $\{\Psi_x = g^{yy} h_{x\perp}, a_{\perp}^1, a_{\perp}^2\} \Rightarrow$ 3 independent fields: $\Psi_x, a_{\perp}^1, a_{\perp}^2$

\Rightarrow Green's function: 3×3 matrix

Anisotropic shear viscosity



$\eta_{yz}/s = 1/4\pi$; η_{xy}/s dependent on T and on α

Critical behaviour: $1 - 4\pi \frac{\eta_{xy}}{s} \propto \left(1 - \frac{T}{T_c}\right)^\beta$ with $\beta = 1.00 \pm 3\%$, α -independent

Non-universal behaviour at leading order in λ and N

Critical exponent confirmed analytically in Basu, Oh 1109.4592

Is there a similar universal result as $\frac{\eta}{s} = \frac{1}{4\pi} \frac{\hbar}{k_B}$ within condensed matter applications of holography?

Candidate: Homes' Law

Homes' Law

Homes' Law $\rho_s = C\sigma(T_c)T_c$

Shown to hold experimentally to great accuracy (Homes et al, Nature 2004)

Zaanen (Nature, 2004): $\tau(T_c) = \frac{\hbar}{k_B T_c}$

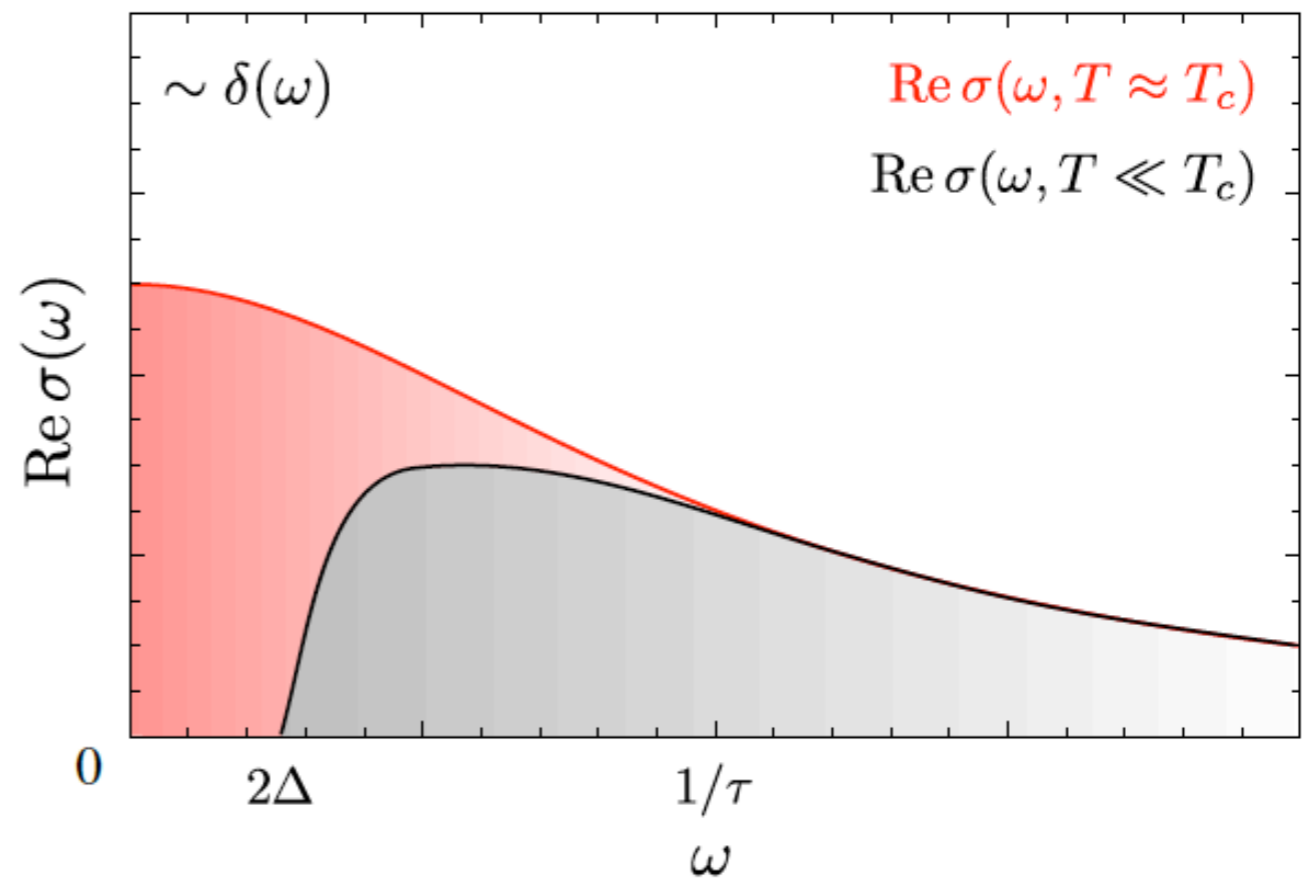
Planckian dissipation: Shortest possible dissipation timescale

Holographic version:

Preliminary results in J.E., Kerner, Müller 1206.5305

Not possible to calculate superconducting density ρ_s holographically

$$\text{Re } \sigma(\omega) = \rho_s \delta(\omega)$$



Idea: Rewrite Homes' Law using sum rules

Homes' Law

Homes' Law: $\rho_s = C\sigma(T_c)T_c$

Sum rule: $\omega_P^2(T = 0) = \omega_P^2(T = T_c)$

$$\rho_s \propto \omega_P^2(T = 0)$$

Drude law: $\sigma = \frac{ne^2\tau}{m}$, $\omega_P^2 = \frac{4\pi ne^2}{m}$

$$\Rightarrow 4\pi\sigma(T_c) = \omega_P^2(T_c)\tau(T_c)$$

\Rightarrow Homes' Law equivalent to

$$\tau(T_c)T_c = \text{const}$$

Homes' Law

Assume diffusion can be used to determine the timescale

$$\Rightarrow D(T_c)T_c = \text{const}$$

Holography in the probe limit without backreaction
(Einstein-Maxwell theory):

$$D = \frac{1}{4\pi} \frac{d}{d-2} \frac{1}{T}$$

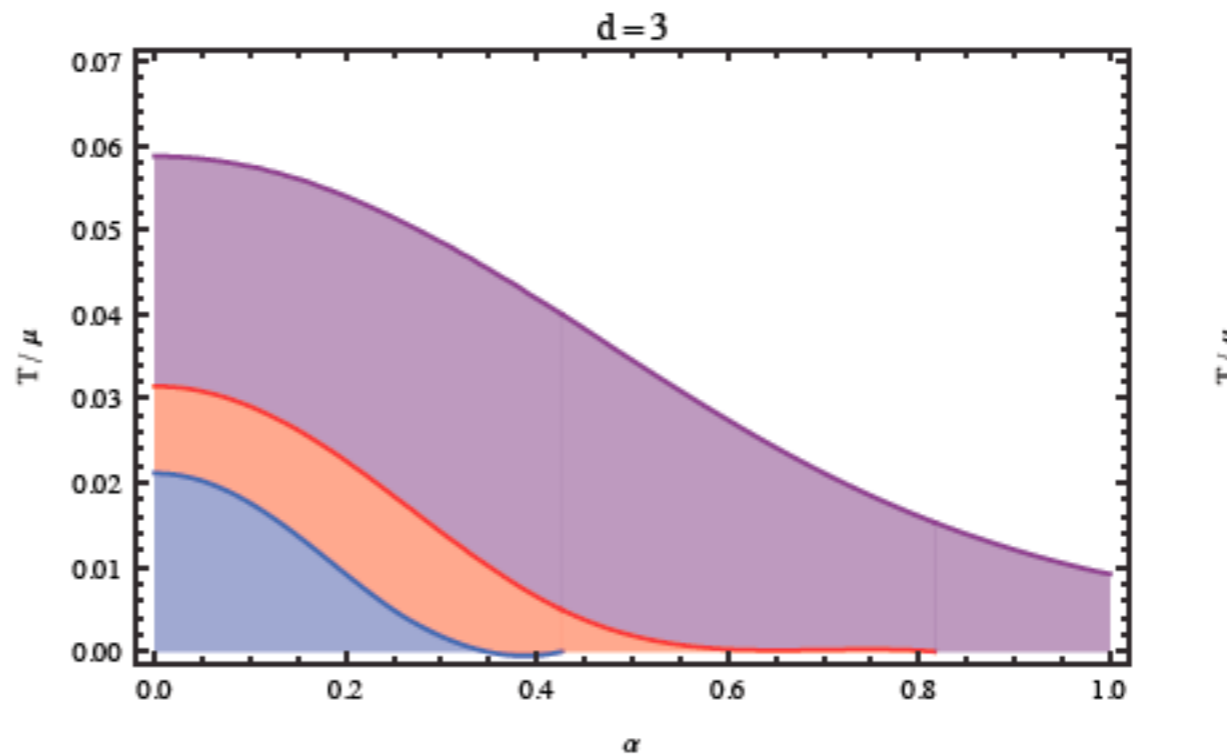
Including the backreaction we expect $D = \frac{1}{4\pi} \frac{d}{d-2} \frac{1}{T} f\left(\frac{T}{\mu}\right)$

Including the backreaction

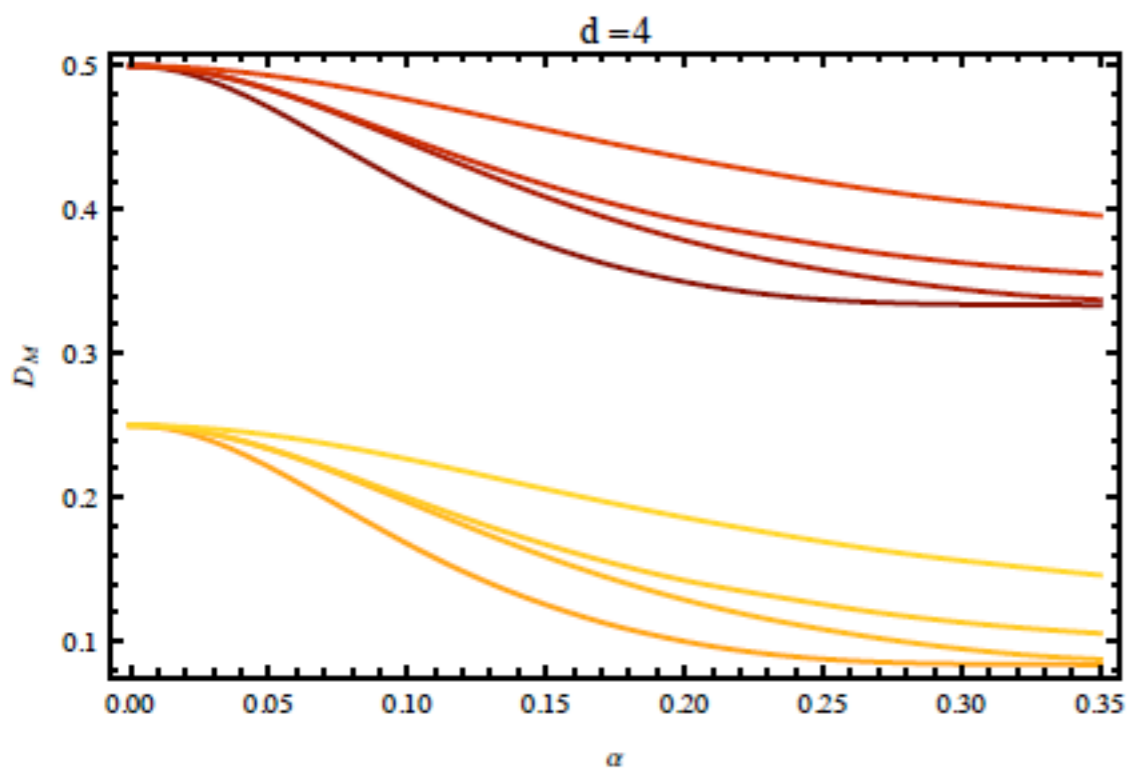
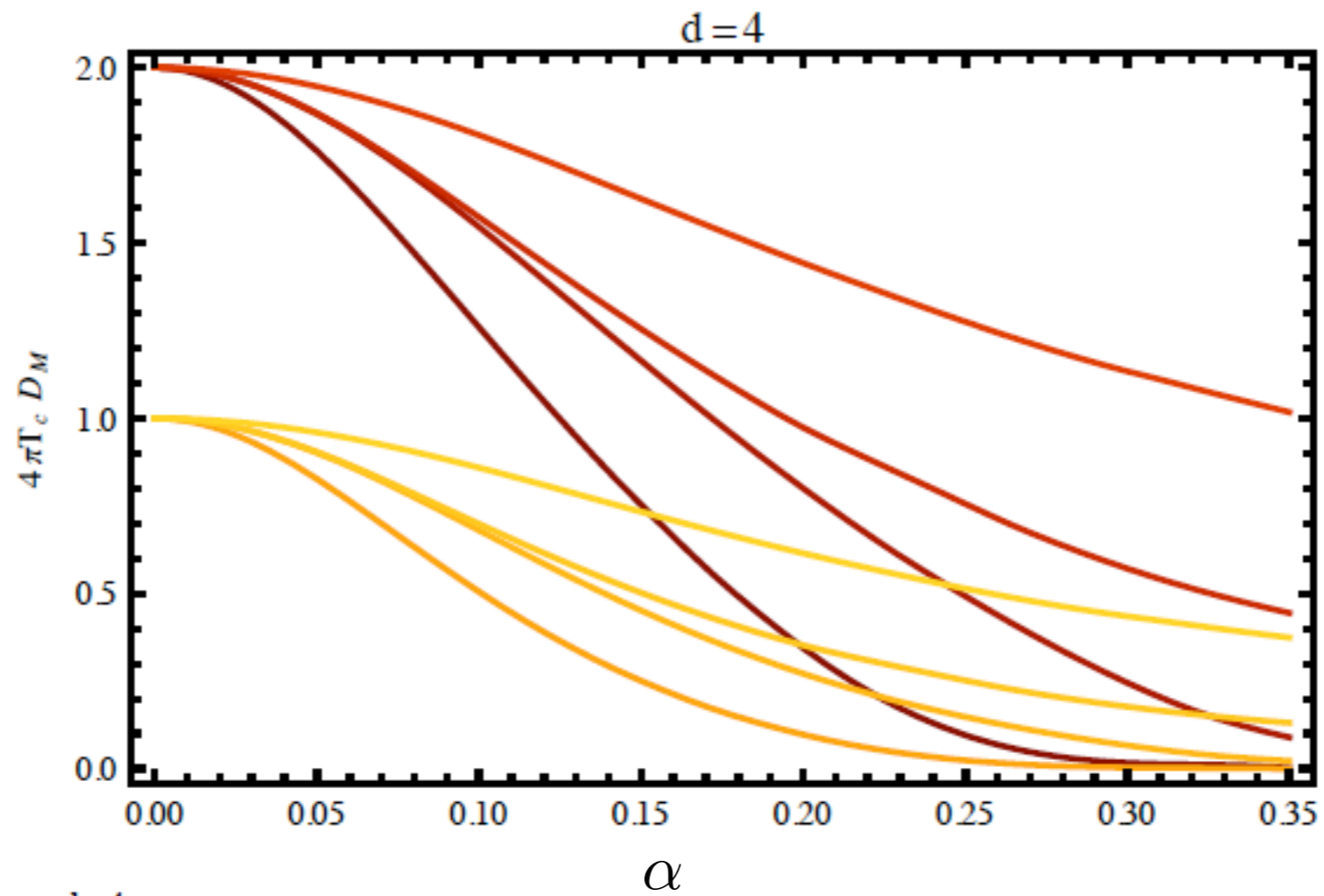
$$S = \frac{1}{2\kappa^2} \int d^{d+1}x \sqrt{-g} \left[R - 2\Lambda - \frac{2\kappa^2}{e^2} \left(\frac{1}{4} F_{ab} F^{ab} - |\nabla\Phi - iA\Phi|^2 - V(|\Phi|) \right) \right]$$

Backreaction parameter $\alpha^2 L^2 = \frac{\kappa^2}{e^2}$

Phase diagram



R charge and momentum diffusion times T_c vs. α



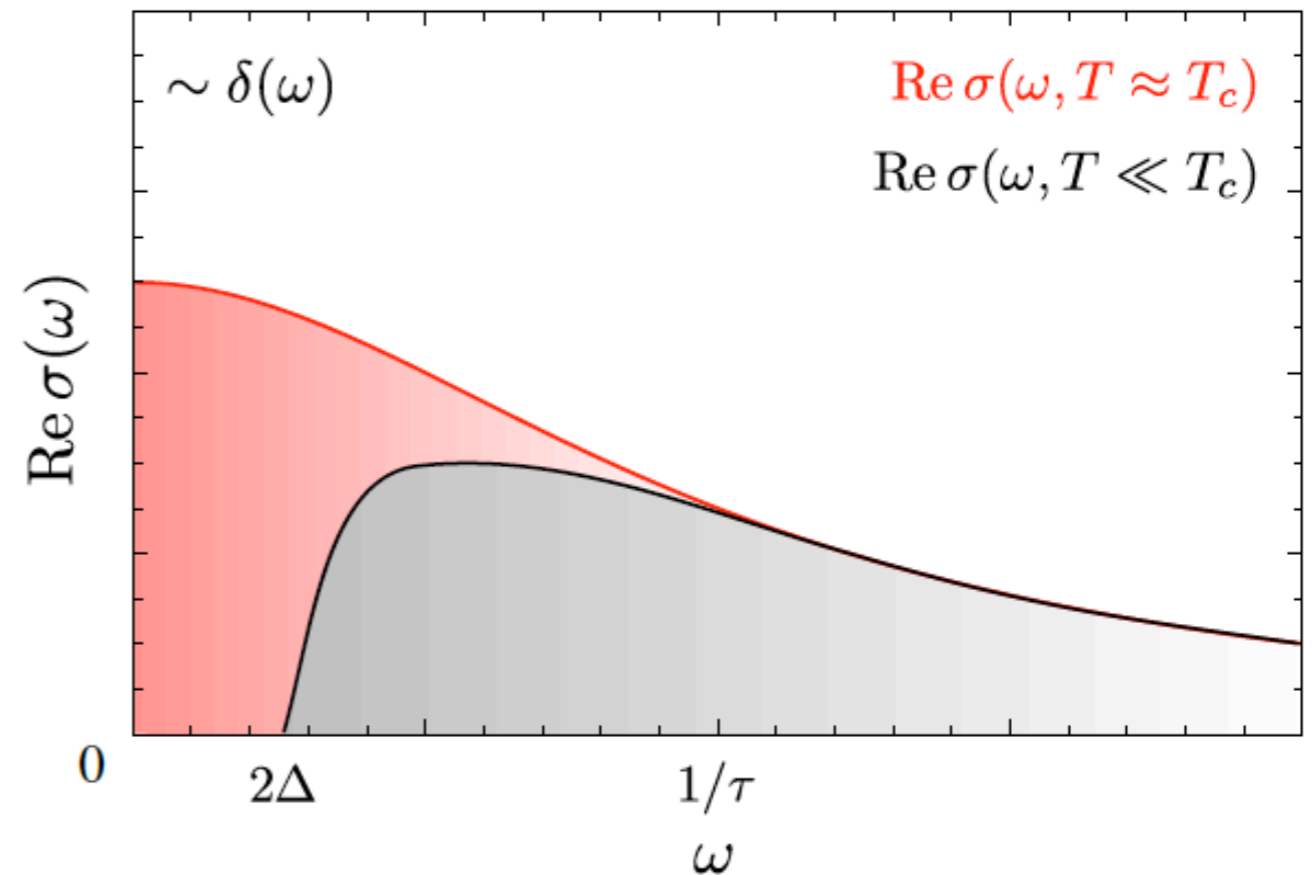
Reasons for decrease

Decrease may originate from pseudogap states whose number increases with backreaction

$$1 - \frac{8N_s}{\omega_{\text{Pn}}^2} = \frac{C}{4\pi} \tau_c T_c$$

Sum rule

$$\frac{\omega_{\text{P}}^2}{8} = \int_0^\infty d\omega \operatorname{Re} \sigma(\omega)$$



$$N_n = \int_0^\infty d\omega \operatorname{Re} \sigma(\omega) \Big|_{T > T_c} = \frac{\omega_{\text{Pn}}^2}{8},$$

$$N_s = \int_{0^+}^\infty d\omega \operatorname{Re} \sigma(\omega) \Big|_{T < T_c}$$

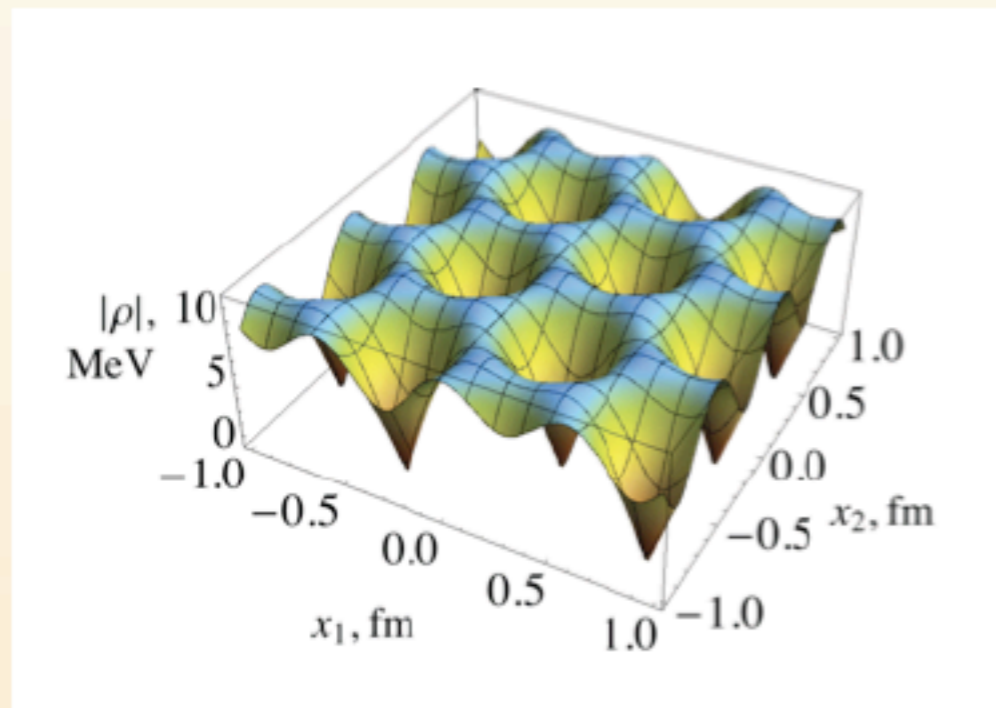
$$\rho_s \equiv \omega_{\text{Ps}}^2 = 8(N_n - N_s) = \omega_{\text{Pn}}^2 - 8N_s$$

External electromagnetic fields

A magnetic field leads to

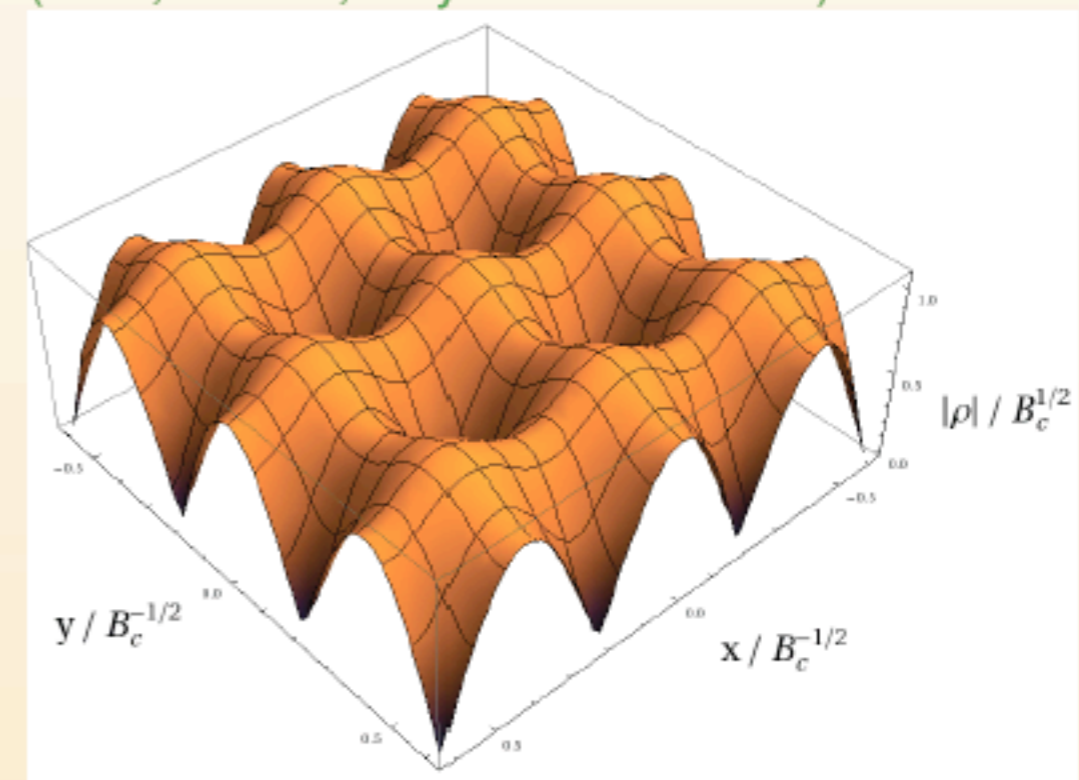
ρ meson condensation and superconductivity in the QCD vacuum

Effective field theory:
(Chernodub)



Gauge/gravity duality
magnetic field in black hole supergravity
background

(J.E., Kerner, Strydom PLB 2011)



Condensation in magnetic field

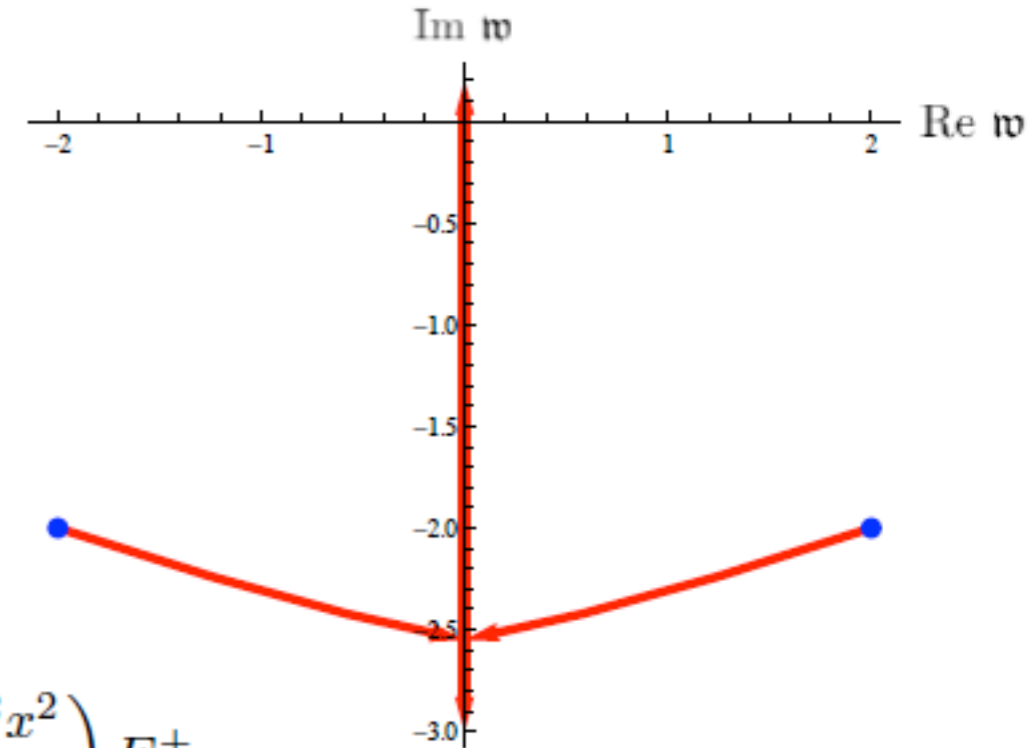
$$S = \int d^5x \sqrt{-g} \left[\frac{1}{2\kappa_5^2} (R - 2\Lambda) - \frac{1}{4\hat{g}^2} F_{\mu\nu}^a F^{a\mu\nu} \right] + S_{\text{bdy}}$$

$$F_{\mu\nu}^a = \partial_\mu A_\nu^a - \partial_\nu A_\mu^a + \epsilon^{abc} A_\mu^b A_\nu^c$$

$$A_y^3 = xB$$

Fluctuations

$$0 = \partial_u^2 E_x^+ + \frac{1}{f} \partial_x^2 E_x^+ + \left(\frac{f'}{f} - \frac{1}{u} \right) \partial_u E_x^+ - \frac{2}{xf} \partial_x E_x^+ + \left(\frac{\omega^2}{f^2} - \frac{B^2 x^2}{f} \right) E_x^+$$

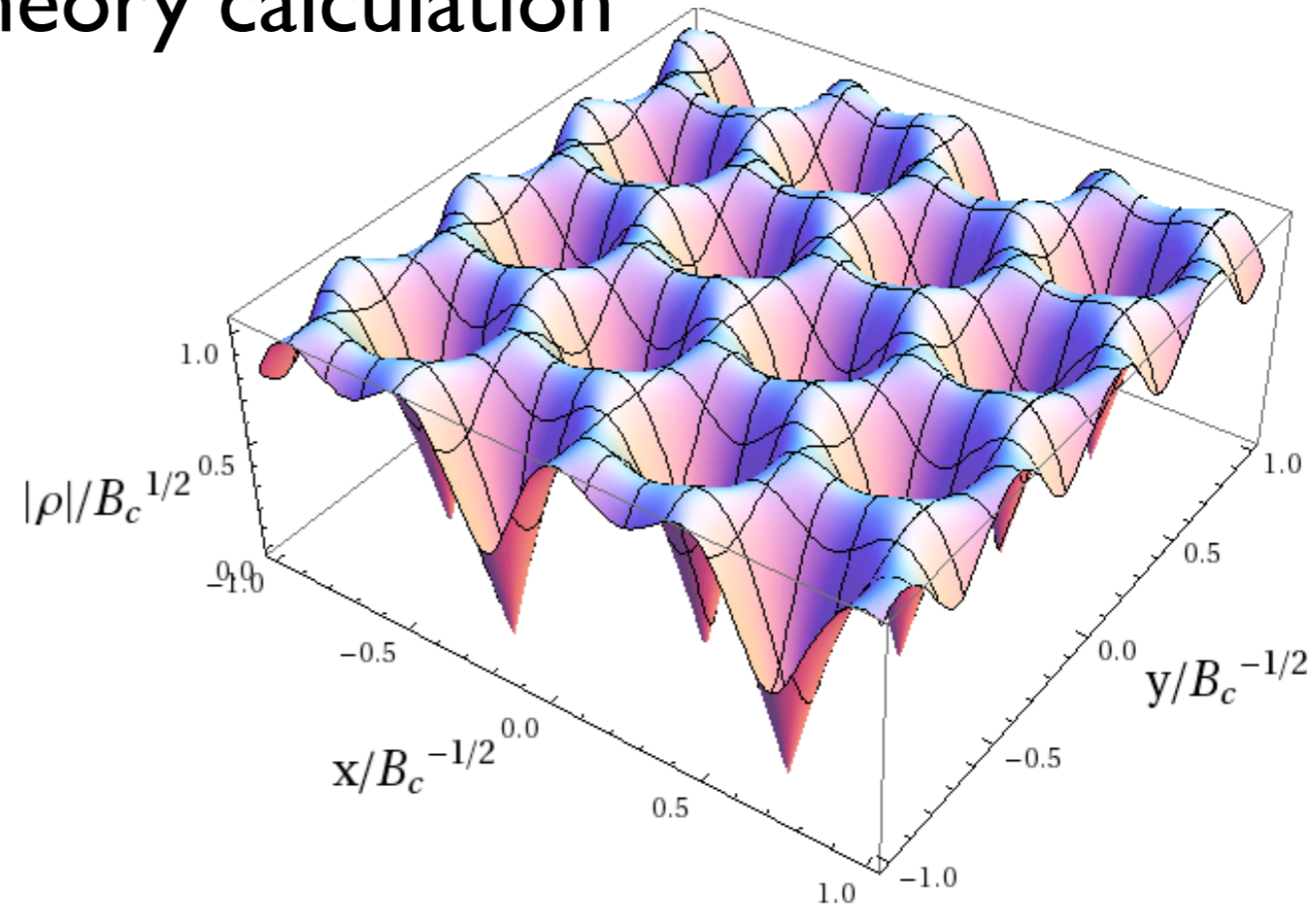
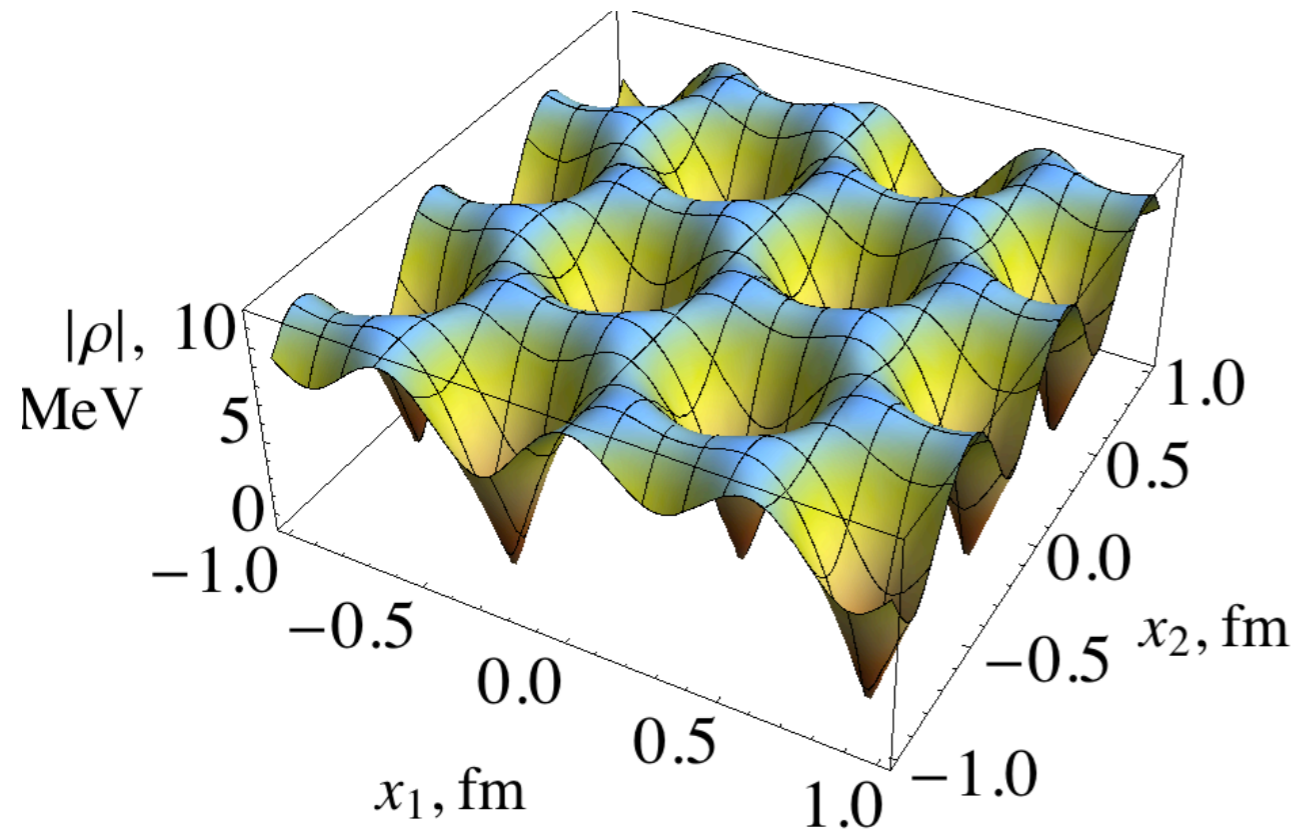


cf. Chernodub;

Callebaut, Dudas, Verschelde;

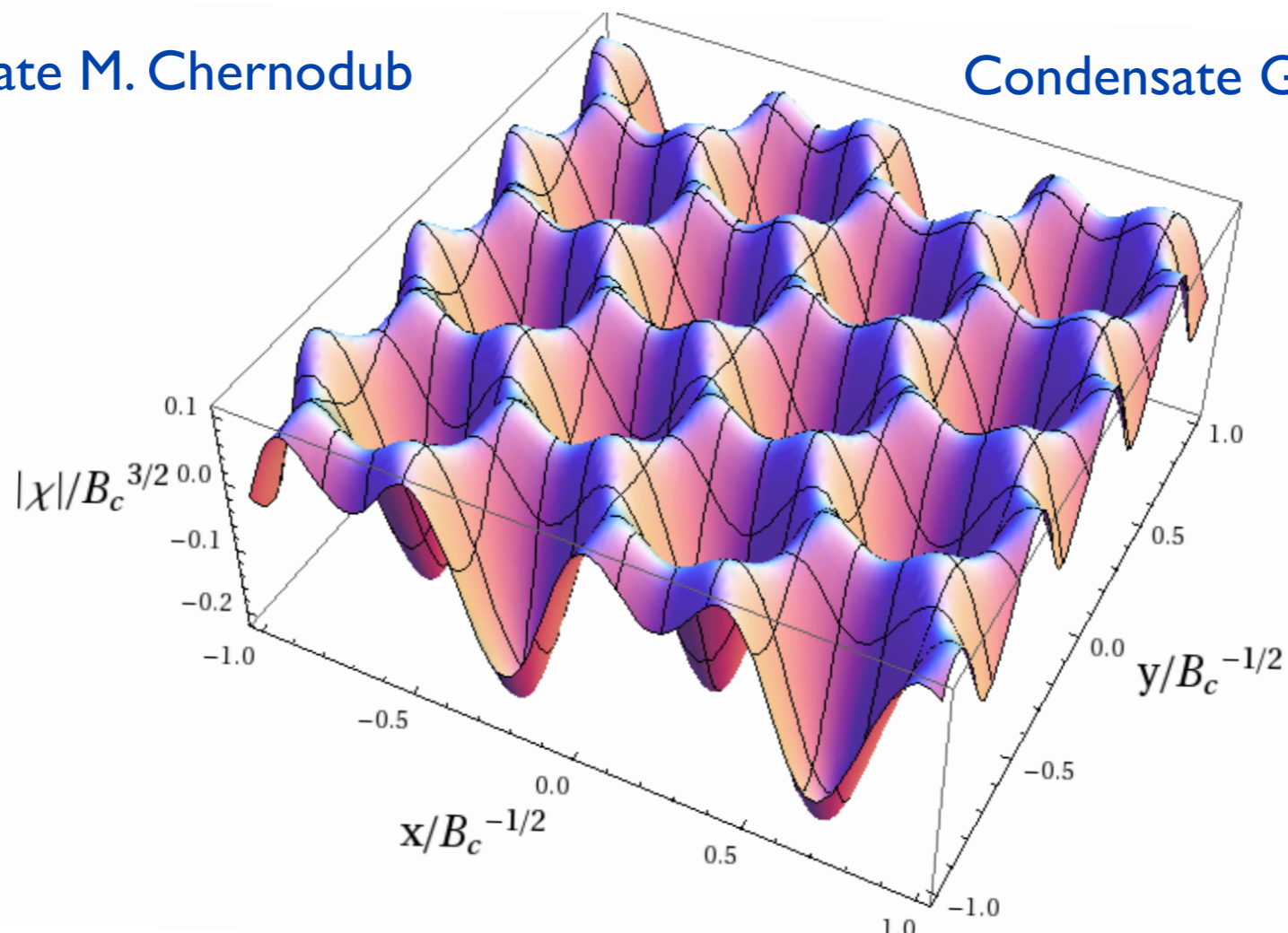
Donos, Gauntlett, Panteleidou

Comparison to field theory calculation



Condensate M. Chernodub

Condensate Gauge/Gravity Duality



Magnetization
Gauge/Gravity Duality

Conclusions

- D3/D7 with finite isospin: Holographic p-wave superconductor with known dual field theory
- Add backreaction in bottom-up model
- Anisotropic shear viscosity:
Non-universal contribution at leading order in N and λ
- Progress towards holographic Homes' law
- Condensation at finite magnetic field