

UNITARITY, NEAR-LOCALITY, AND HILBERT SPACE NETWORKS

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THE HOLOGRAPHIC WAY
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Refs: SBG 1108.2015, 1201.1037;
SBG and Y. Shi, 1205.4732; WIP

This talk:

1. Introductory/motivational comments
 - comments on AdS/CFT
2. Review Hawking/nice slice description
3. Restoring unitary
 - “nonviolent” nonlocality vs. complementarity,
massive remnants (fuzzballs, firewalls)
4. Comments on Hilbert space networks

The information “paradox,” in a nutshell:

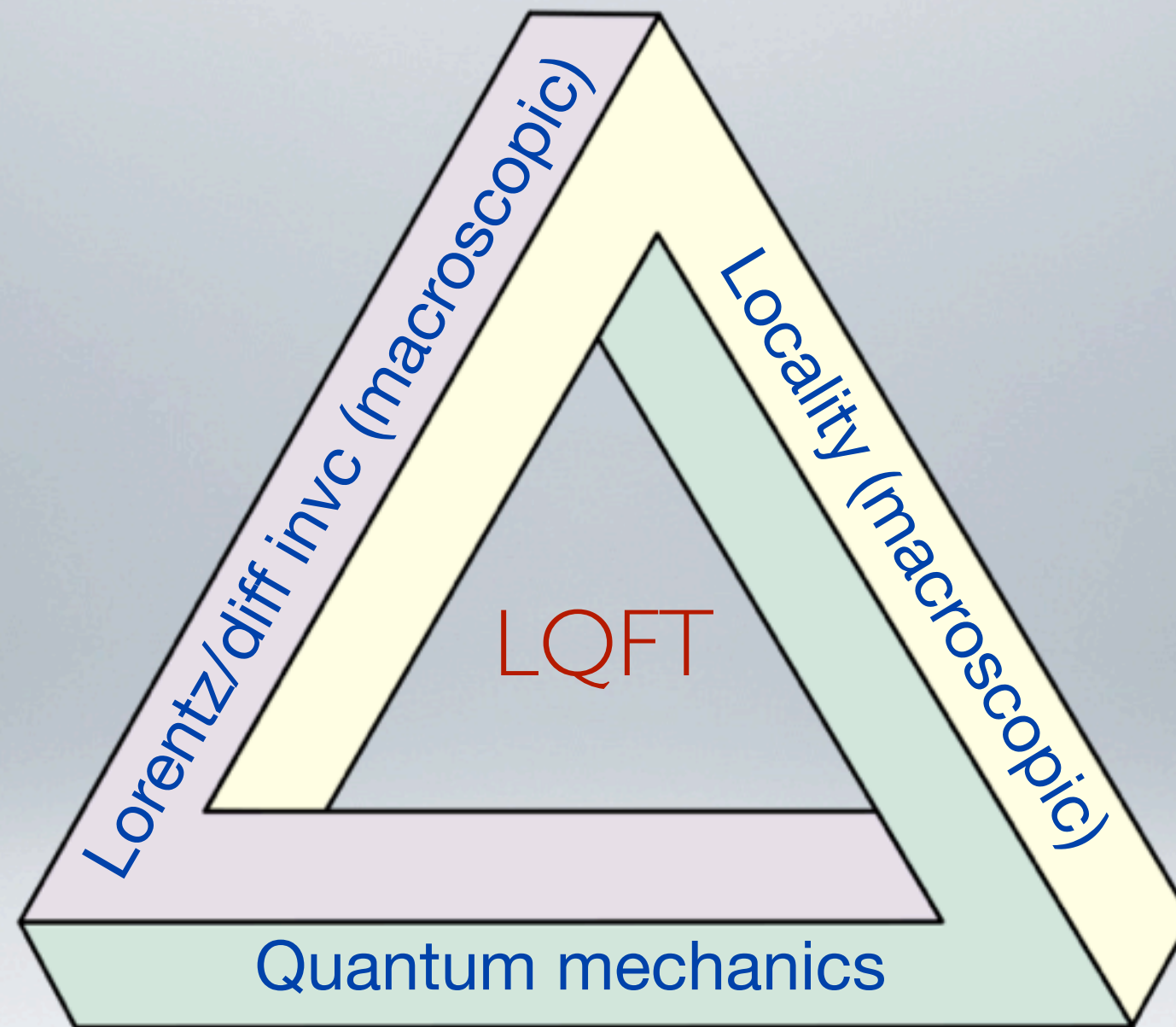
Information cast into a black hole

- can't get out locality
- can't be destroyed QM; energy conserv.
- can't be left in remnant catastrophic instabilities

Stupid mistake?

Guide to new physics?

APPARENTLY A FUNDAMENTAL CONFLICT:

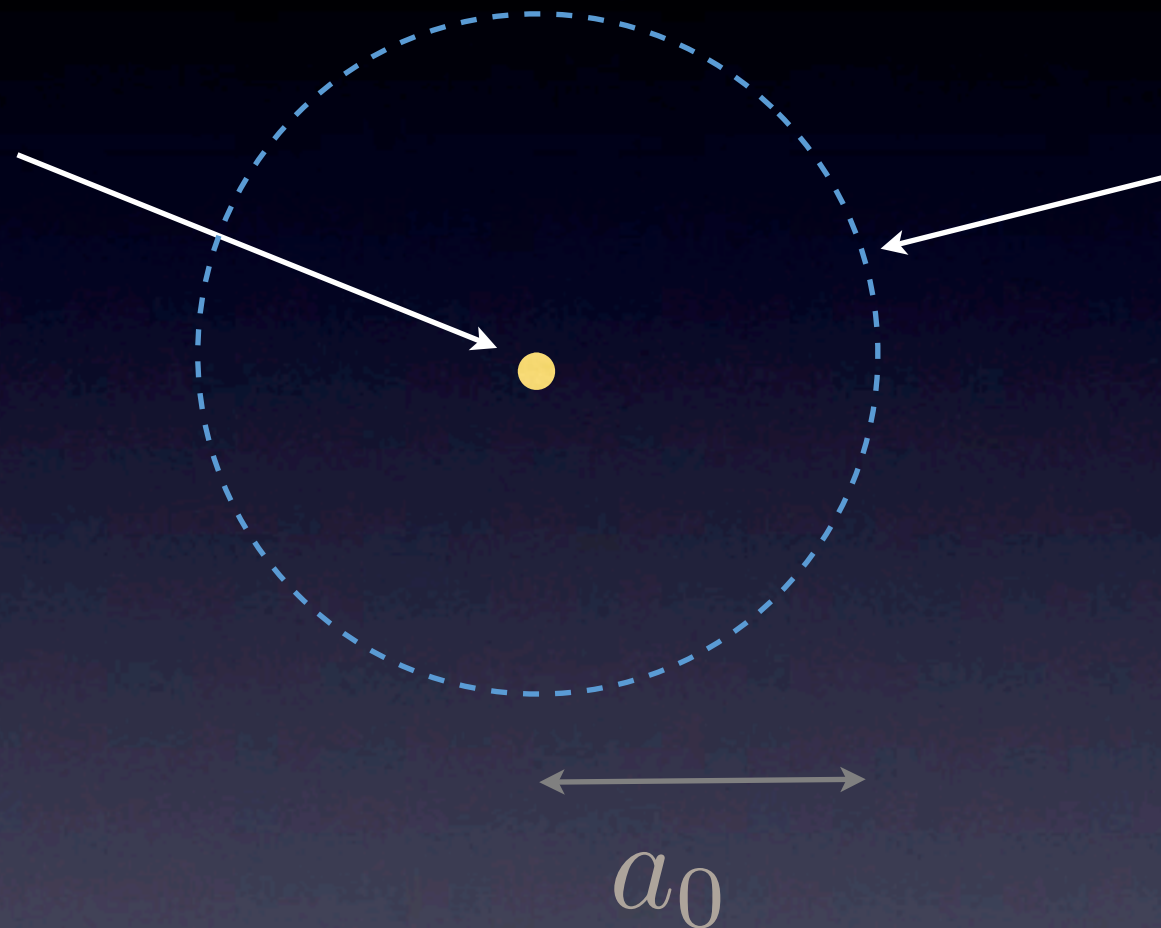


- QM, LI: hard to modify (consistency, observation). Locality?
- It's not about singularities, renormalizability? Long distance.

A SEEMINGLY SIMILAR CRISES:

Classical atom

CM breaks
down here



QM takes over
here
(CM irrelevant)

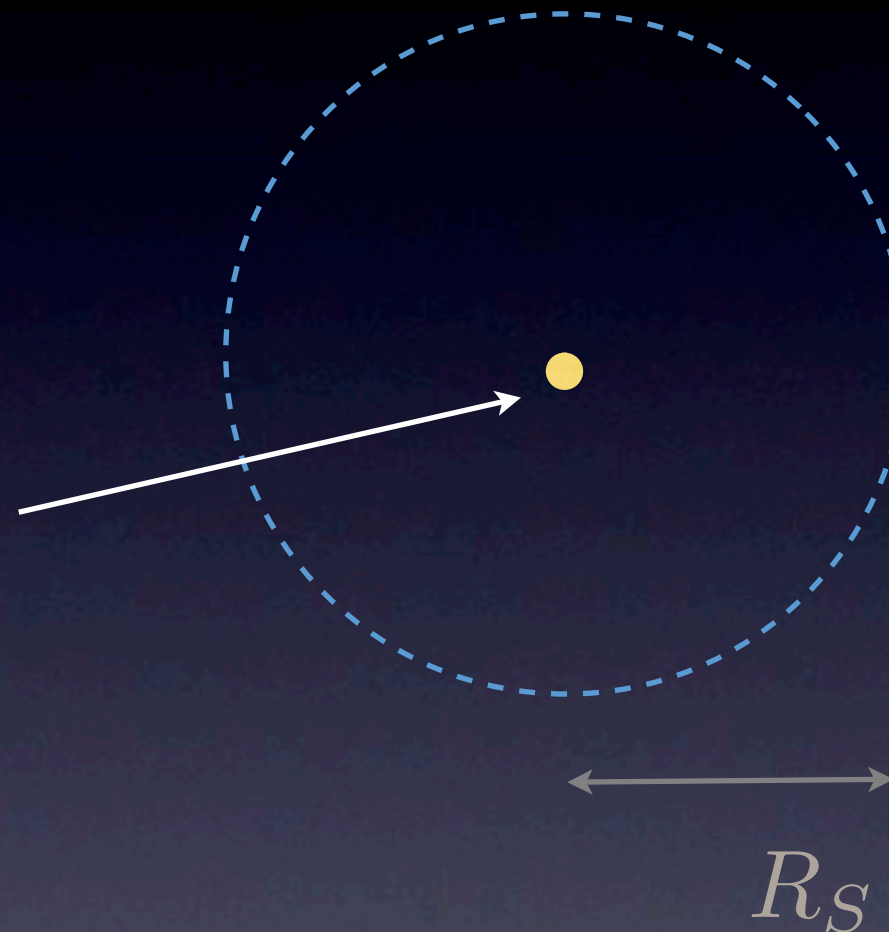
New physics was needed:

Uncertainty principle
Wave mechanics...

A SEEMINGLY SIMILAR CRISES:

Black hole

LQFT breaks
down here



QG relevant
here?
(subtle breakdown
LQFT?)

New physics ~~was~~ needed:
is

Uncertainty principle
~~Wave mechanics...~~

???

“CLASSICAL INSTABILITY PARADOX”



“BLACK HOLE INFORMATION PARADOX”

This unitarity crisis is likely an important guide to understanding new principles/mechanisms.

(As was the stability problem of the atom)

What other guides do we have?

AdS/CFT??

Microstates??

Cosmological comparisons; tests?

Locality/ local observables

S-matrix

Amplitude magic??

Correspondence

Quantum info. transfer from BHs

SBG and Sloth;
many others

e.g. hep-th/0512200,
w/ Marolf and Hartle

Erice lectures: 1105.2036

nonperturbative?

Comments on AdS/CFT (More detail: 1105.6359)

- many regard as resolution -

A question: can we recover from the boundary theory a sufficiently fine-grained bulk description, e.g., of evolution of a small ($\ll R$) BH, and of infalling observers?

Need:

$$\begin{array}{ccc} & M : \mathcal{H}_{\text{bulk}} \rightarrow \mathcal{H}_{\text{bdy}} & \text{interacting} \\ & \nearrow \text{I-I, unitary} \nwarrow & \\ U_{\text{bulk}} = M^\dagger U M & & U \end{array}$$

Unitary bulk evolution?

Approaches:

(~) local bulk observables

- challenge in QG
- much discussion at recent KITP workshop
- general “relational” approach: e.g SBG, Marolf, Hartle hep-th/0512200; used in inflation
- no clear and general story

S-matrix (flat space limit)

- Problem: construct scattering states from boundary data; extract fine-grained S-matrix
 - there are obstacles
- (more discussion:
1106.3553 w/ M. Gary)

More general approaches (e.g. beyond AdS/CFT):

1. Investigate correspondence boundary
2. Quantum information transfer
3. The gravitational S-matrix

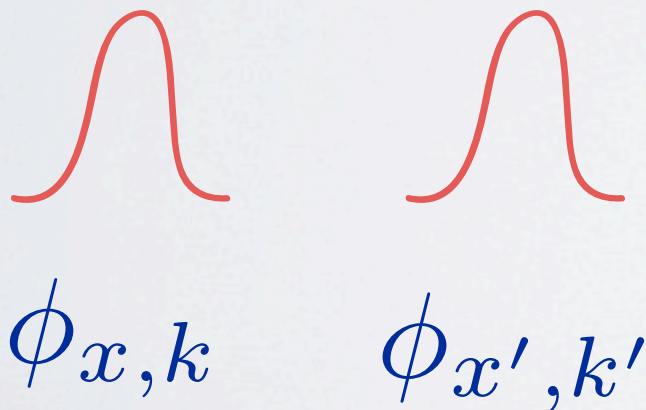
Correspondence

Existing theory: LQFT, semiclassical background

correspondence point - various proposals: planckian curvature, modified/string uncertainty, modified dispersion, holographic bounds ...

Configurations: $\phi_{x,k} \phi_{x',k'} |0\rangle$
(min uncertainty wavepackets)

Where is description untrustworthy?



$\phi_{x,k}$

$\phi_{x',k'}$

Correspondence

Existing theory:

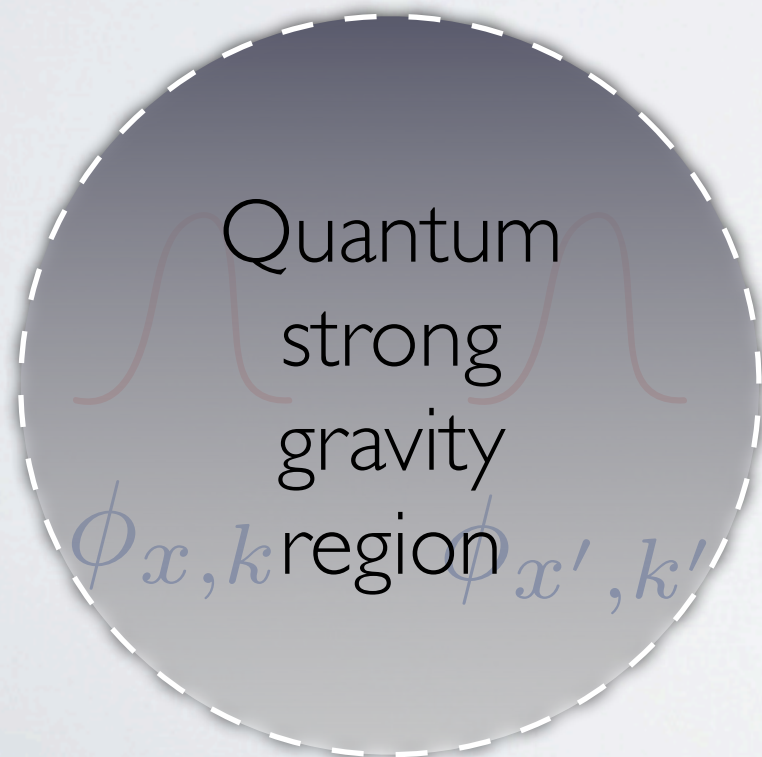
LQFT, semiclassical background

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Configurations:

$$\phi_{x,k} \phi_{x',k'} |0\rangle$$

(min uncertainty wavepackets)



Where is description untrustworthy?

$$|x - x'|^{D-3} < (\hbar G) |k - k'|$$

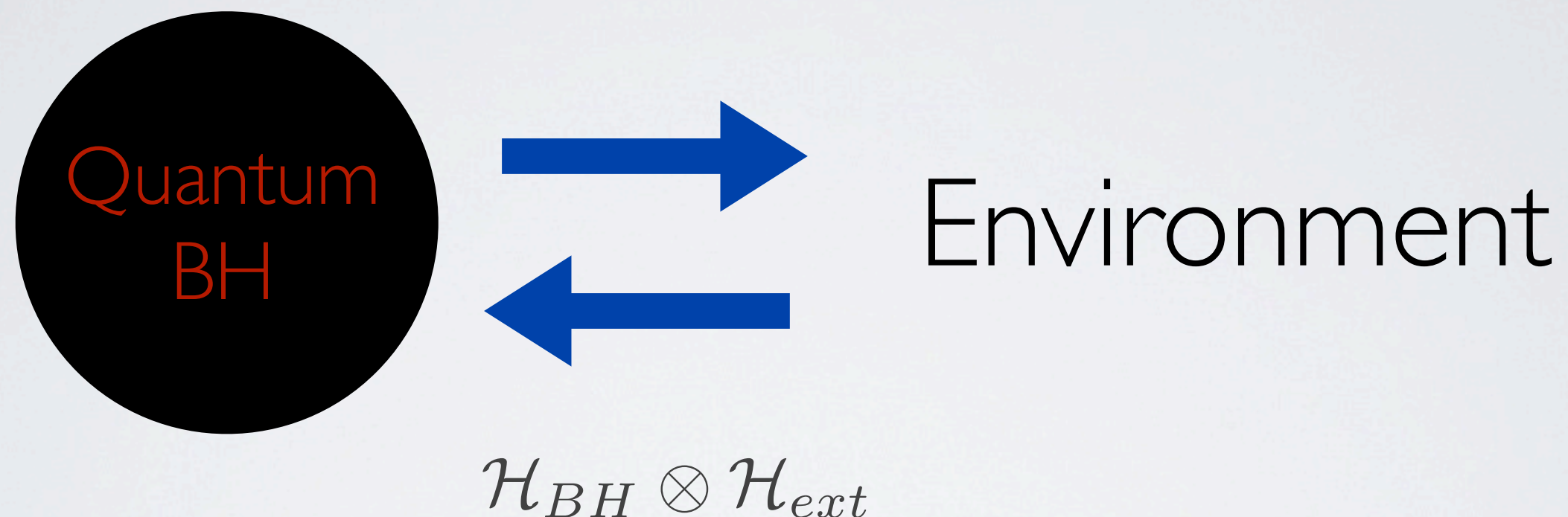
“locality bound”

~ Heisenberg microscope
(multiparticle generalizations)

How do we describe physics in this regime?

Assume: **quantum mechanics** (take seriously)

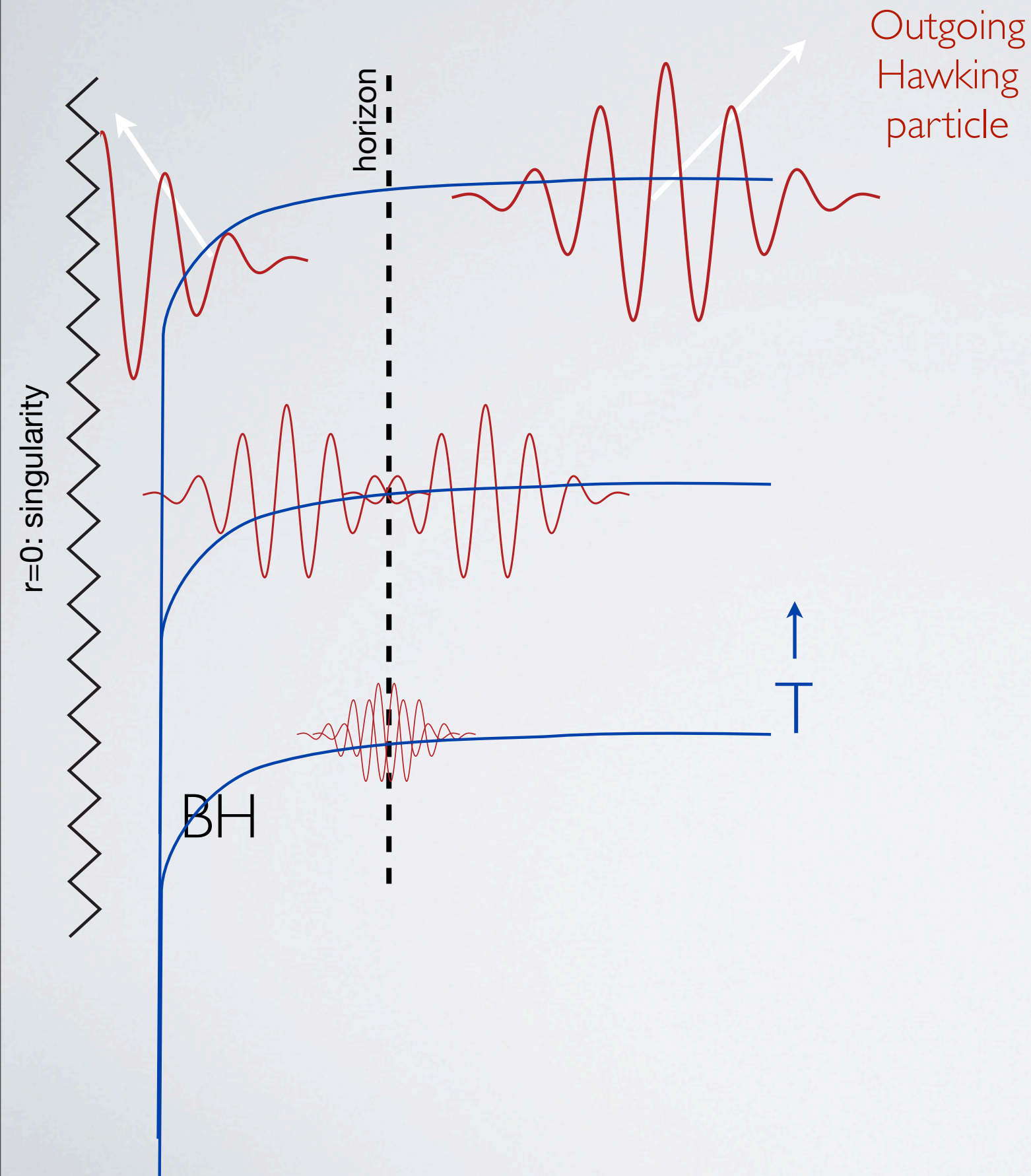
Also assume: quantum subsystems (\sim localization)



Information exchange; **unitary**

“Effective quantum theory” -- but more?
(recall: “QM+locality+Poincare \Rightarrow QFT”)

Another guide: “as close as possible” to LQFT. Review



nice-slice description
-- sharpens tension

X^+ , X^- : null, Kruskal

$$X^+(X^- + e^{-2T} X^+) = R_c^2$$

“Hawking state”

(explicit in 2d:
SBG and Nelson, 1992)

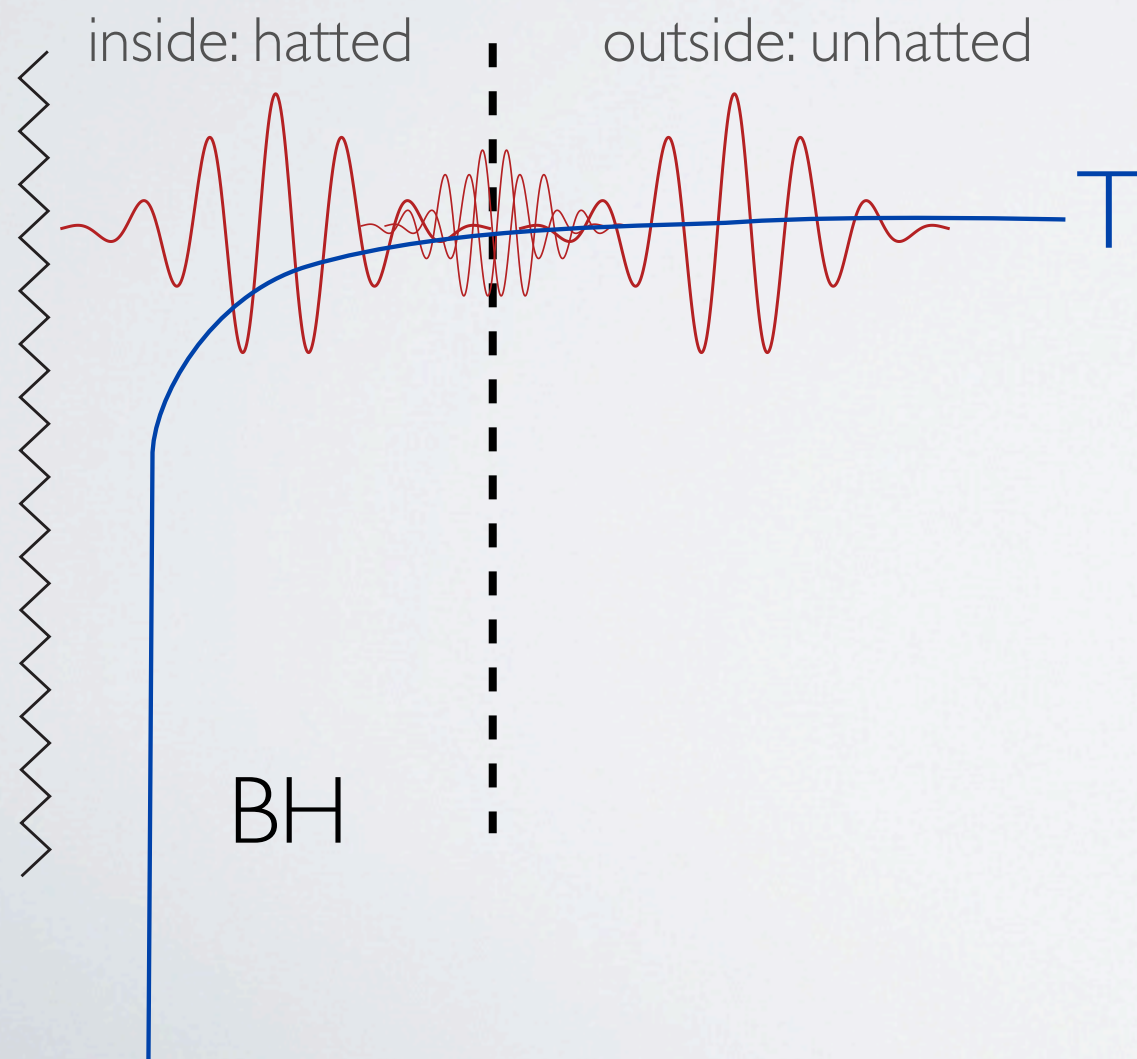
$$|\psi\rangle_{\text{Hawk}} = \prod_{jn} S_{jn} |\hat{0}\rangle |0\rangle$$

$j \sim$ asymp. frequency

$n \sim$ position along slice

$$\exp \left\{ \tanh^{-1} [e^{-\omega_j/2T}] \left(\hat{b}_{jn}^\dagger \hat{b}_{jn}^\dagger - \text{h.c.} \right) \right\}$$

$$= \frac{1}{Z} \sum_{\{n_{jn}\}} e^{-\beta H/2} |\widehat{\{n_{jn}\}}\rangle |\{n_{jn}\}\rangle$$



Regularization:

For given T :

$$n \rightarrow \infty \leftrightarrow \rightarrow \text{horizon}$$

$$\text{cutoff: } n < N(T) \sim \lambda > L$$


Shorter modes “look like” vacuum

“Hawking state” (cont’d):

So rewrite:

$$|\psi\rangle_{\text{Hawk}} = \prod_{jn} S_{jn} |\hat{0}\rangle |0\rangle = |0\rangle_N \prod_j \prod_{n=1}^{N-1} (S |\hat{0}\rangle |0\rangle)_{jn}$$

UV modes:
“in vacuum”



$\in \mathcal{H}_{BH} \otimes \mathcal{H}_{ext}$
As advocated

- $|0\rangle_N$ can either go with BH, or “ancillary”
- Hilbert spaces effectively finite dim. (if finite time)

Evolution: e.g 2d: $U = 1$

$$|\psi\rangle_{\text{Haw}} = |0\rangle_N \prod_j \prod_{n=0}^{N-1} (S|\hat{0}\rangle|0\rangle)_{jn}$$

$$= \frac{|0\rangle_N}{Z} \sum_{\{n_{jn}\}, n < N} e^{-\beta H/2} |\widehat{\{n_{jn}\}}\rangle |\{n_{jn}\}\rangle$$

more generally: U_{LQFT}
 (also, can include infalling matter)

and: $N \rightarrow N + 1$

Note: “generalized” unitary transform: dimensions of
 $\mathcal{H}_{BH}, \mathcal{H}_{ext}$ change (“isometry”)

Cartoon: timestep $\sim R$ (See also Mathur)

$$|0\rangle_N \rightarrow |0\rangle_{N+1} (|\hat{0}0\rangle + |\hat{1}1\rangle)_{n=N}$$

“qubit model”

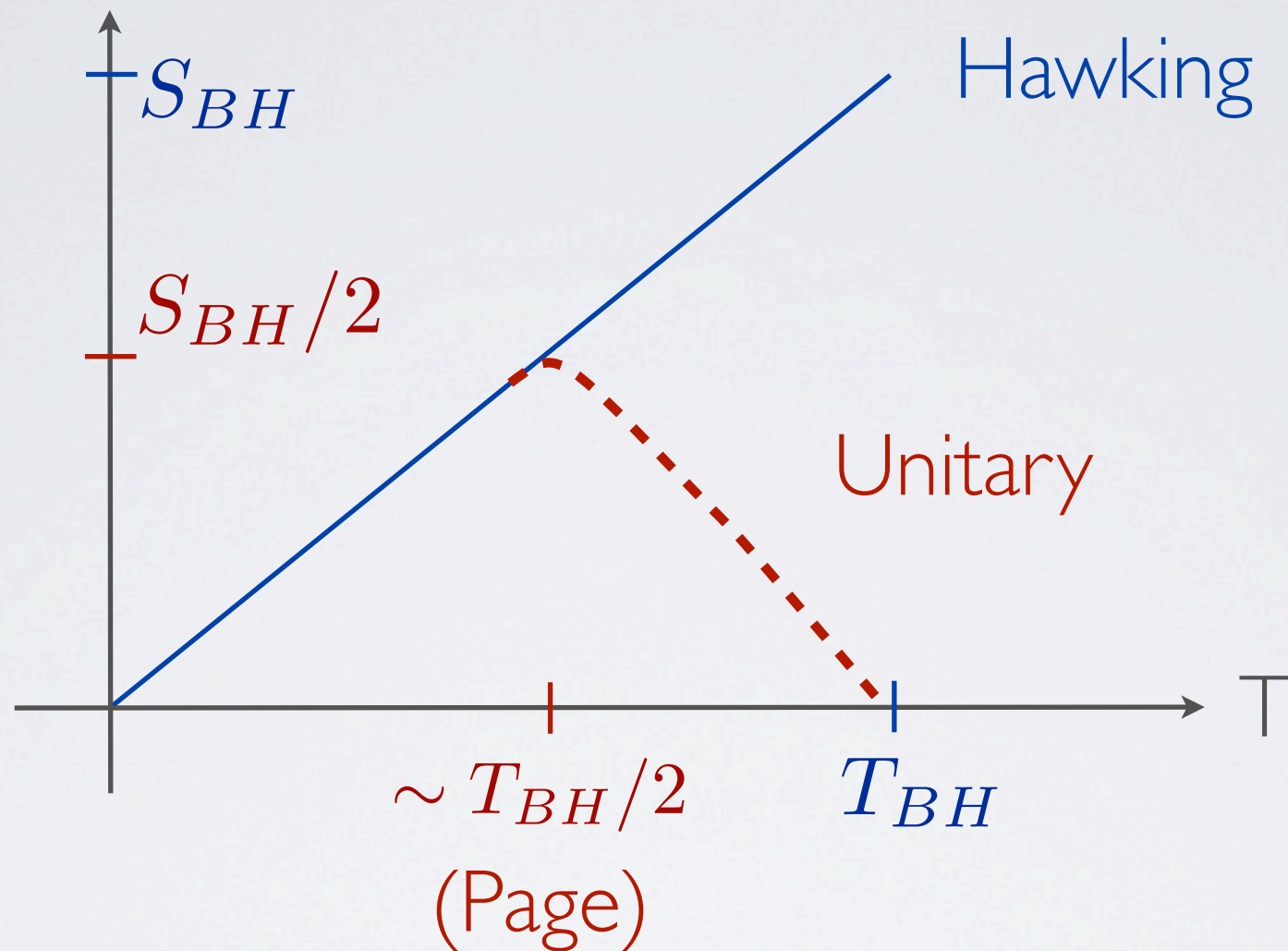
↑
pairs produced

$(\omega_j \simeq 1/\beta)$

Nonunitarity

$$\rho_{ext} = \text{Tr}_{BH} (|\psi\rangle_{\text{Hawk}} \langle\psi|) \propto \sum_{\{n_{jn}\}, n < N} e^{-\beta H} |\{n_{jn}\}\rangle \langle\{n_{jn}\}|$$

$$S_{ext} = -\text{Tr} \rho_{ext} \ln \rho_{ext}$$



arguments for failure of LQFT/nice-slice description
(hep-th/0703116, 0911.3395, etc.)

- problems making gauge invariant
- problems w/ perturbative quant.

How is unitarity restored?

General scenarios:

Massive remnants

includes fuzzballs, firewalls

Complementarity (holography)

“Nonviolent” nonlocality

All approaches nonlocal. How to describe physics?

(No fundamental spacetime?)

General framework -- not necessarily based on spacetime



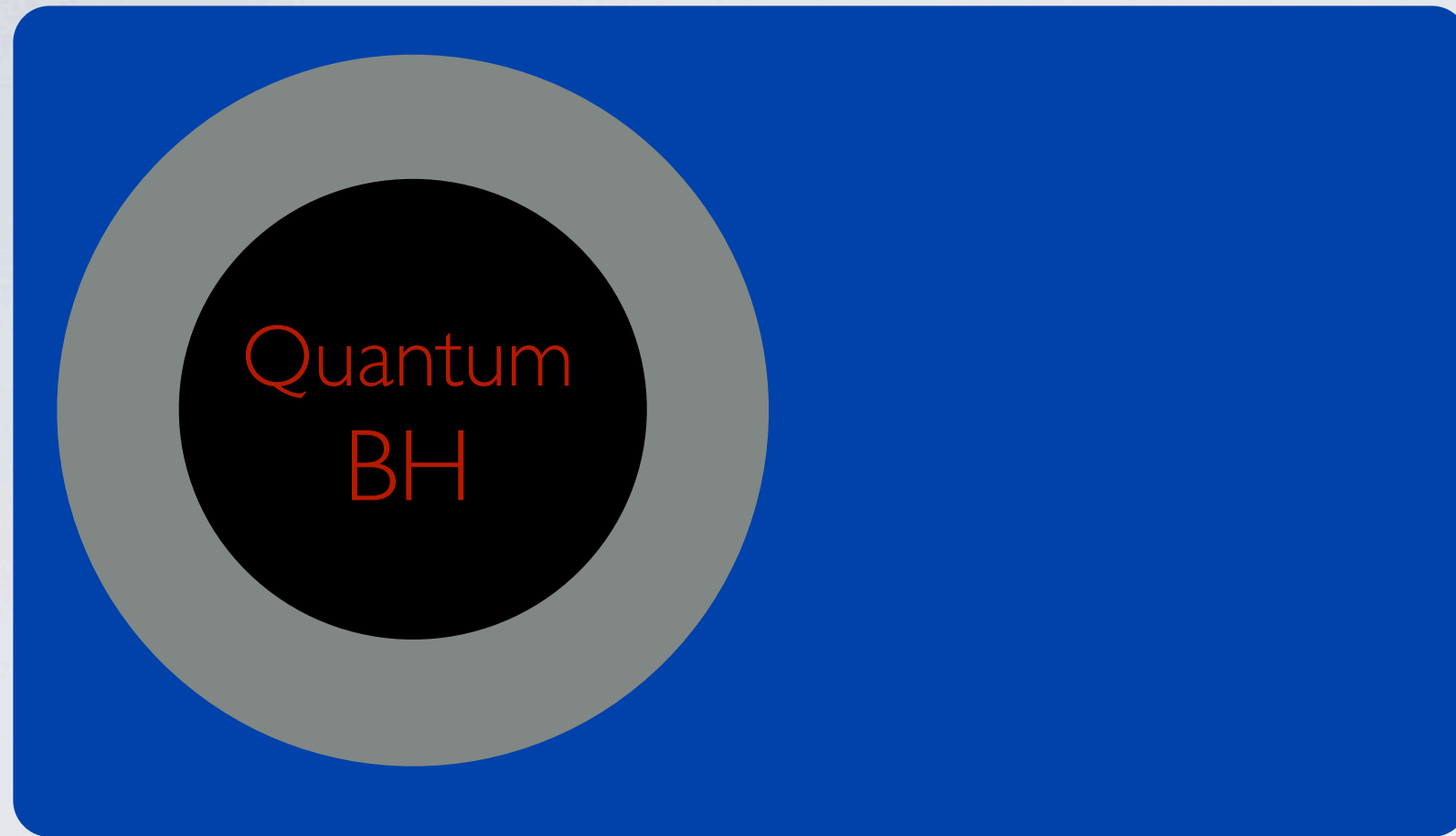
Quantum
BH

“Hilbert space
network”

$$\mathcal{H}_{BH} \otimes \mathcal{H}_{ext}$$

(nb: AdS/CFT could
work this way)

General framework -- not necessarily based on spacetime



“Hilbert space network”

$$\mathcal{H}_{BH} \otimes \mathcal{H}_{near} \otimes \mathcal{H}_{far}$$

(nb: AdS/CFT could work this way)

Some expectations:

$$\mathcal{H}_{far} \approx LQFT$$

$$\mathcal{H}_{near} \sim LQFT$$

$$\mathcal{H}_{BH} \neq LQFT$$



$$U \approx LQFT$$

$$U \neq LQFT$$

$$\mathcal{H}_{BH}, \mathcal{H}_{near} \text{ finite dimensional; } \log \dim \mathcal{H}_{BH} = S_{BH}?$$

Physical constraints on evolution:

- A) $S \rightarrow 0$
- B) Innocuous to infalling observers (?) (uneventful horizon)
- C) $dS/dt \sim 1/R$ (?)
- D) Near-Hawking; \sim thermodynamic (?)
- E) Correspondence limit: LQFT + GR
- F) Energy conservation
- G) Complete, consistent

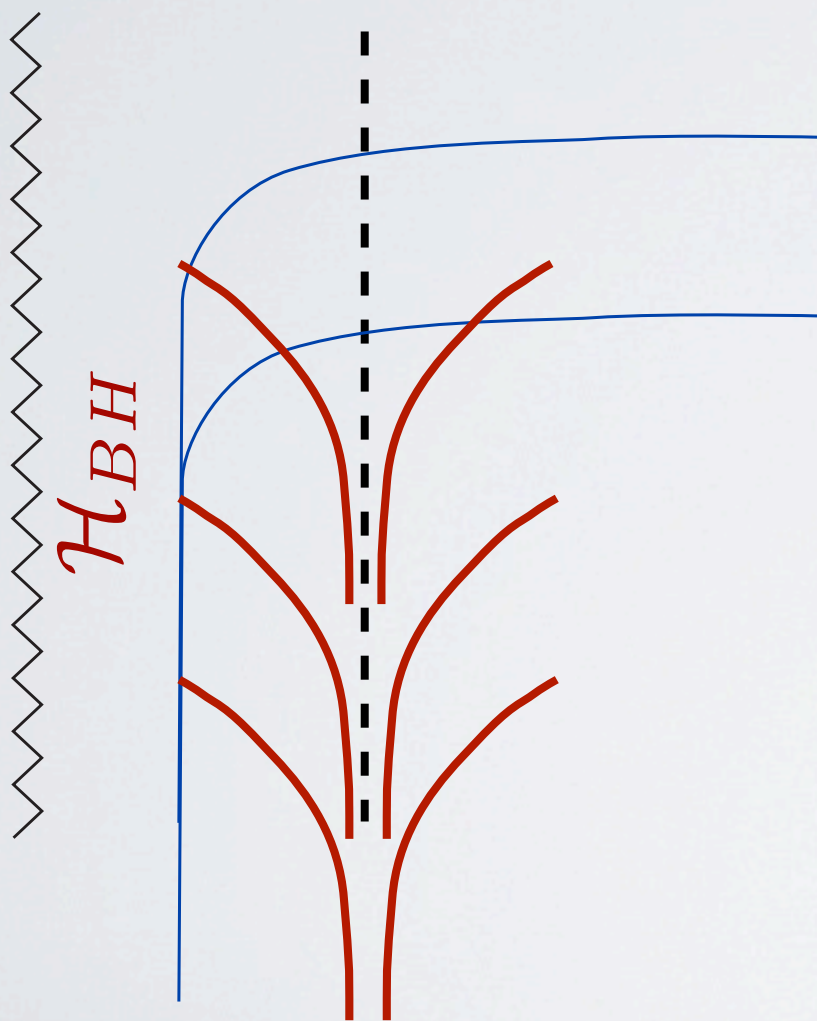
...

Basic guideline: be as conservative as possible!

What is “as close as possible to LQFT”?

Explore examples (e.g. qubit models)

Recap:



Explore examples
(e.g. qubit models)

Not most conservative?

I A) “Fast scrambling”
(complementarity)

e.g. qubit “ $q=0,1$ ” in:

$$|\hat{a}\rangle|\psi\rangle \rightarrow |\hat{a}\rangle \underbrace{|\hat{q}\rangle}_{\text{scrambled}}|\psi\rangle$$

$$\hat{U}(|\hat{a}\rangle|\hat{q}\rangle)$$

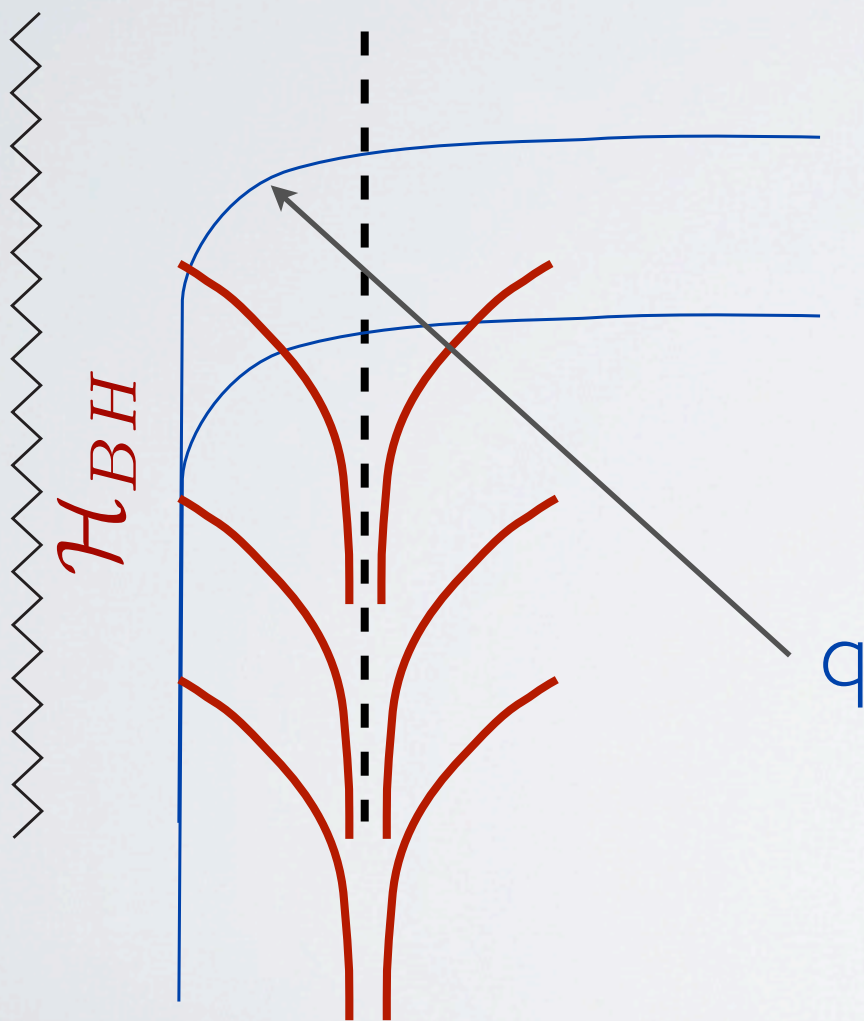
$$\Delta t \sim R$$

$$(\text{or, w/ } |\hat{00}\rangle + |\hat{11}\rangle)$$

then:

$$|\hat{q}'\hat{a}'\rangle|\psi\rangle \rightarrow |\hat{a}'\rangle|\hat{q}'\rangle|\psi\rangle$$

(separate scrambling/transfer)



Not most conservative?

Big departure from LQFT evolution

Indeed, Hayden/Preskill:

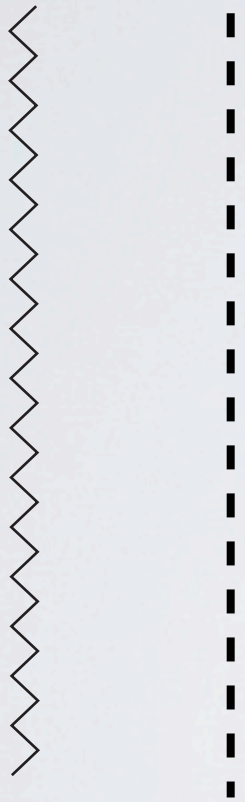
After sufficient evolution, a BH behaves as an information mirror on the scrambling/transfer time!

One classification of scenarios:

	T_{scramble}	T_{transfer}
Susskind	$R \ln R$	$R \ln R$
Page		$< \mathcal{O}(RS)$
HR, nat. slice	$\sim R?$	∞
HR, nice slice	∞	∞

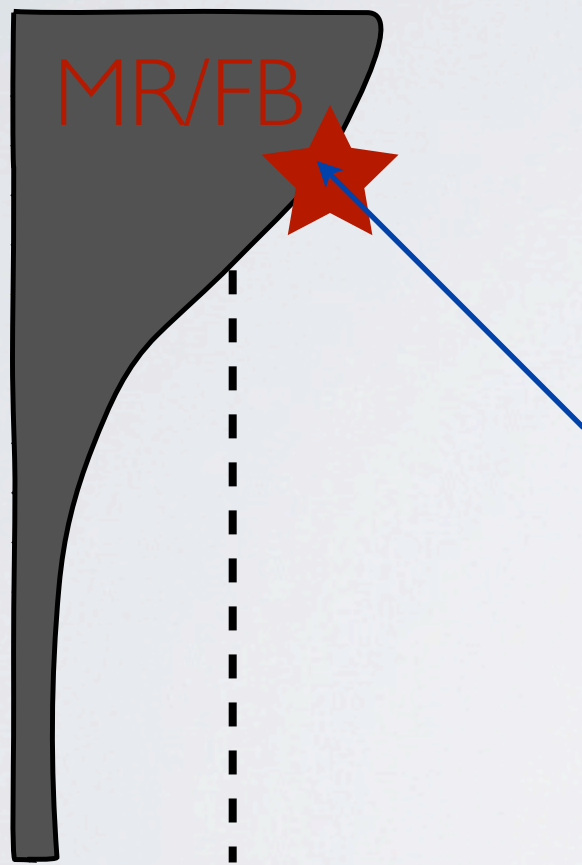
Not most conservative? (cont'd)

I B) Massive remnant; fuzzball; firewall



Not most conservative? (cont'd)

I B) Massive remnant; fuzzball; firewall



expect:

big mods. to \mathcal{H}_{BH}

rapidly varying microstructure
outside horizon


rapid, more limited(?) scrambling
(\sim “neutron star”)

Unitary models: “more conservative”

$$2) \quad |\hat{0}\hat{0}\hat{a}\rangle|a\rangle \rightarrow |\hat{a}\rangle \frac{|\hat{0}0\rangle + |\hat{1}1\rangle}{\sqrt{2}}|a\rangle$$

$$|\hat{0}\hat{1}\hat{a}\rangle|a\rangle \rightarrow |\hat{a}\rangle|\hat{0}1\rangle|a\rangle$$

$$|\hat{1}\hat{0}\hat{a}\rangle|a\rangle \rightarrow |\hat{a}\rangle|\hat{1}0\rangle|a\rangle$$

e.g. “leftmost”
qubits transfer 

$$|\hat{1}\hat{1}\hat{a}\rangle|a\rangle \rightarrow |\hat{a}\rangle \frac{|\hat{0}0\rangle - |\hat{1}1\rangle}{\sqrt{2}}|a\rangle$$

(toy model)

Oversimplified, but:

- can include unitary evol. acting inside, outside

$$|\hat{0}\hat{0}\hat{a}\rangle|a\rangle \rightarrow \hat{U}|\hat{a}\rangle \frac{|\hat{0}0\rangle + |\hat{1}1\rangle}{\sqrt{2}} U|a\rangle$$

$$|\hat{0}\hat{1}\hat{a}\rangle|a\rangle \rightarrow \hat{U}|\hat{a}\rangle |\hat{0}1\rangle U|a\rangle$$

$$|\hat{1}\hat{0}\hat{a}\rangle|a\rangle \rightarrow \hat{U}|\hat{a}\rangle |\hat{1}0\rangle U|a\rangle$$

$$|\hat{1}\hat{1}\hat{a}\rangle|a\rangle \rightarrow \hat{U}|\hat{a}\rangle \frac{|\hat{0}0\rangle - |\hat{1}1\rangle}{\sqrt{2}} U|a\rangle$$

- can generalize to more realistic modes (arXiv: 1201.1037)
(not just qubits)

Also:

“Hawking-like”

“minimal” mods to evolution

- info imprinted in typical Hawking modes
- $\langle T_r^0 \rangle = \langle T_r^0 \rangle_{Hawking}$

3) More general information transferring models:

$$|\hat{q}_1 \hat{q}_2 \hat{a}\rangle |a\rangle \rightarrow |\hat{a}\rangle \frac{|\hat{0}\rangle|0\rangle + |\hat{1}\rangle|1\rangle}{\sqrt{2}} |\hat{0}'\hat{0}''\rangle |q'_1 q''_2\rangle |a\rangle$$

Usual Hawking
particles

Not typically
occupied

(again, modulo unitaries)

... again, representative of more general (e.g. multimode) models:

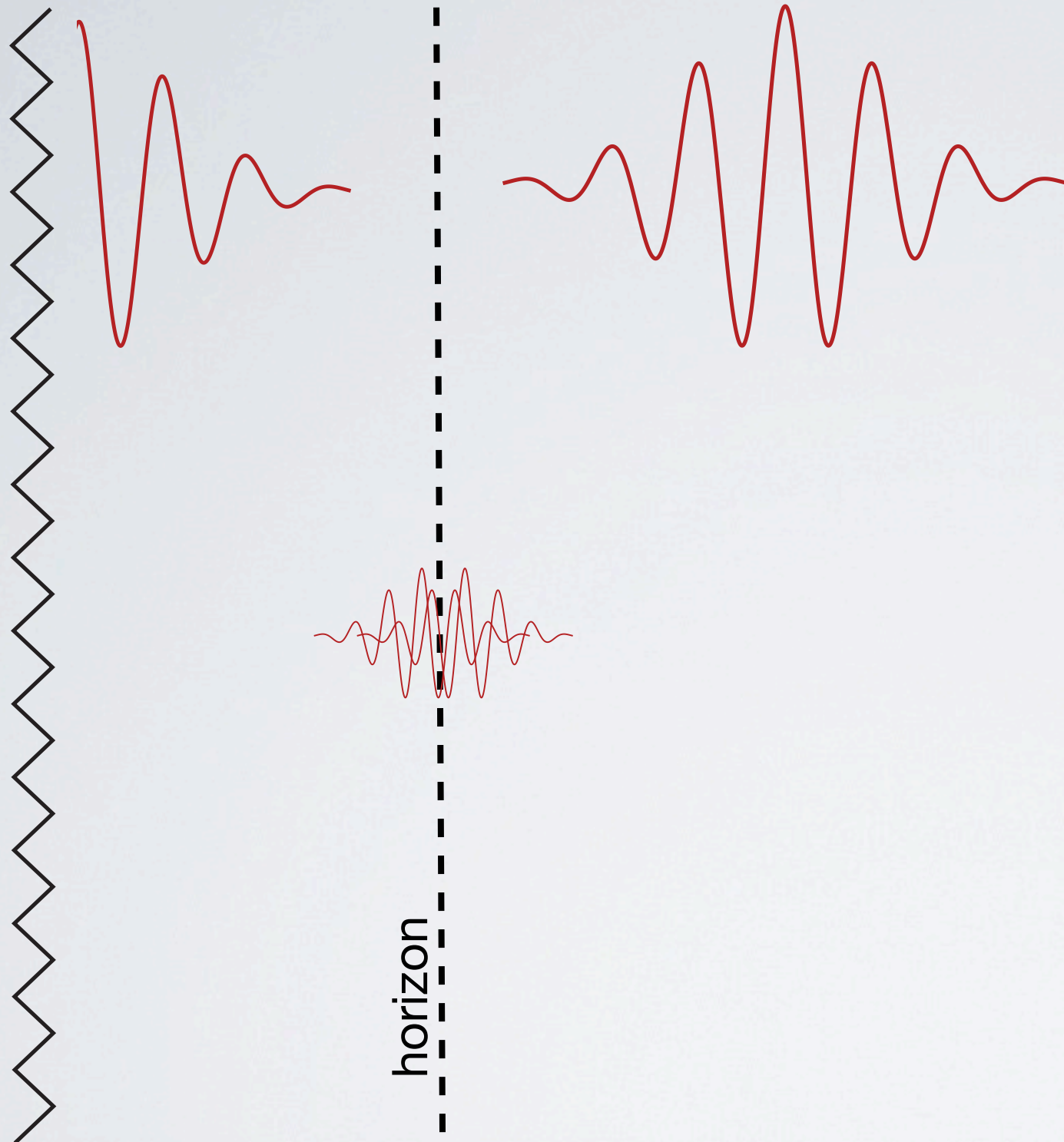
- more generic
- yield extra energy flux

(But: $dE/dI > 1/R$?)

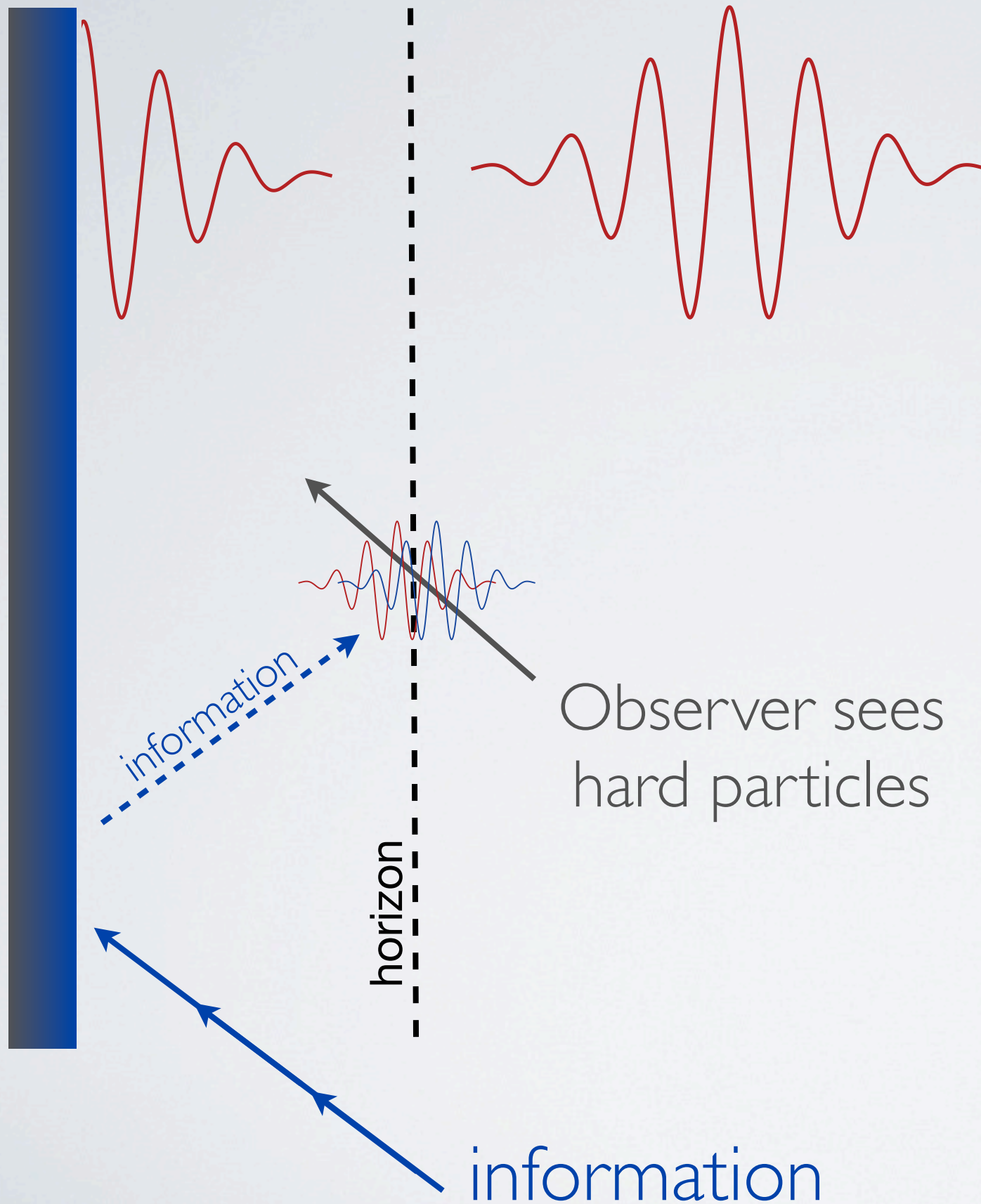
Important points:

- These forms of evolution are **not local** with respect to the semiclassical geometry
- “Nonviolent” or uneventful horizon?

Uneventful horizons:

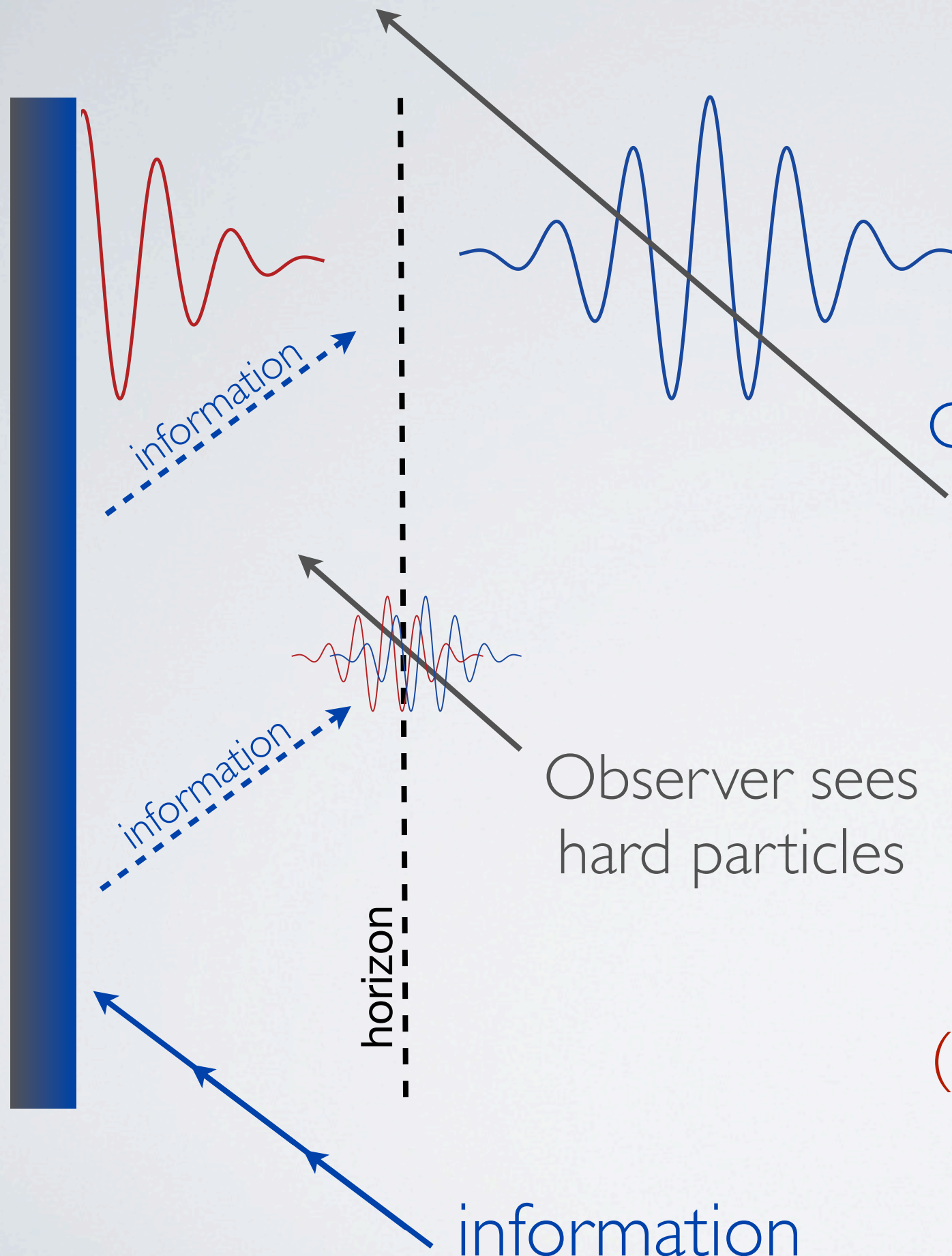


Uneventful horizons:



(recently rebranded:
"firewall")

Uneventful horizons:

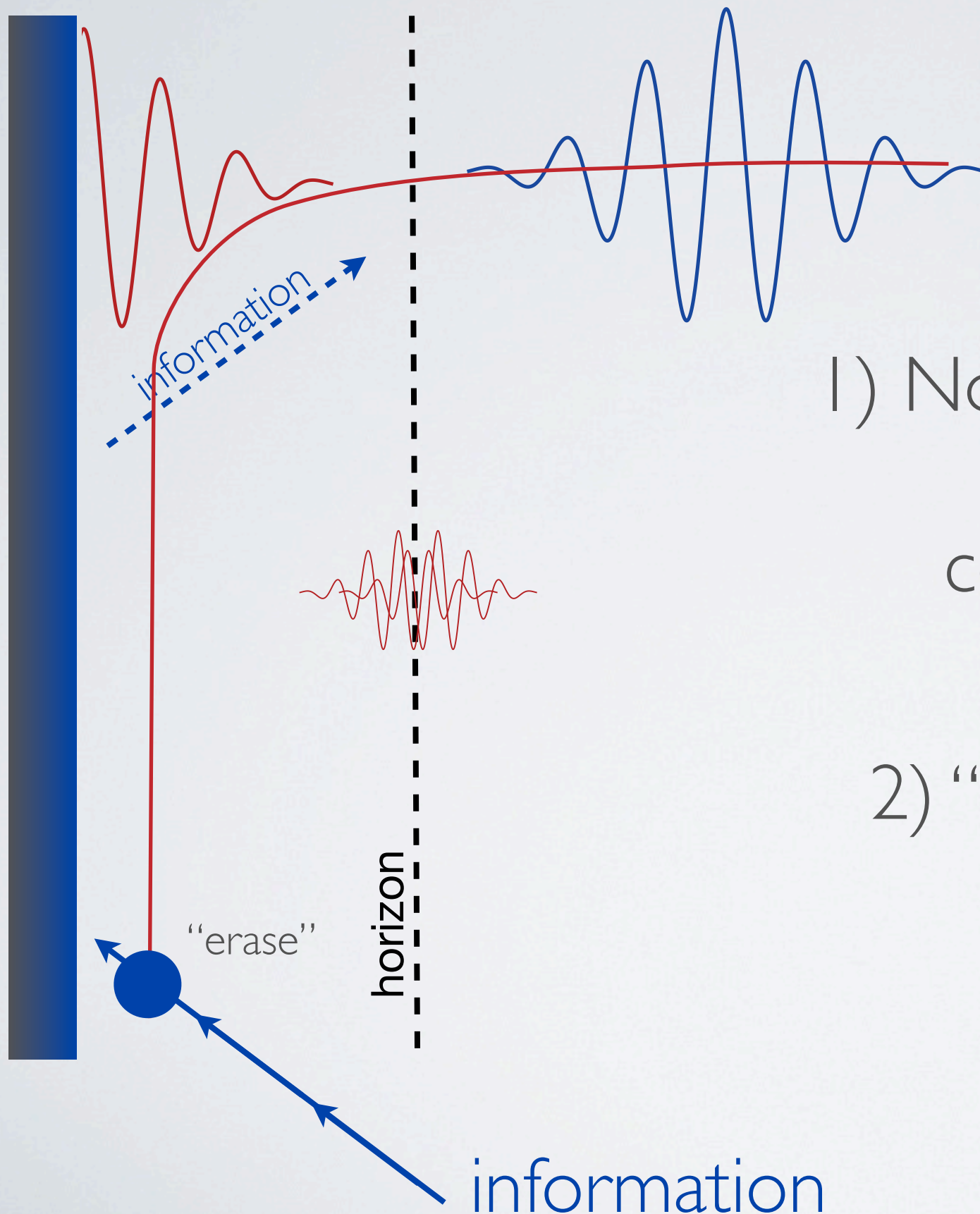


Observer sees ~ 1 extra quantum
of energy $\sim 1/R$ per time R
(~innocuous)

(recently rebranded:
“firewall”)

(n.b.: idea is: geometry not strictly
correct picture of physics -
spacetime=approx. !);

Two further comments:



1) Nonviolent horizon:
small dim of \mathcal{H}_{near}
constraint on evolution

2) "Weak complementarity"
As long as $T \sim RS$
(e.g. nice slice descript.
not good at long times)

- Can explore other restrictions from physical + q. info-theoretic requirements (SBG & Shi 1205.4732 & WIP)

E.g. characterize information transfer

Minimal -- Simplest, most efficient form:



A



B

- Can explore other restrictions from physical + q. info-theoretic requirements (SBG & Shi 1205.4732 & WIP)

E.g. characterize information transfer

Minimal -- Simplest, most efficient form:



$$A = A_1 \otimes A_2$$



B

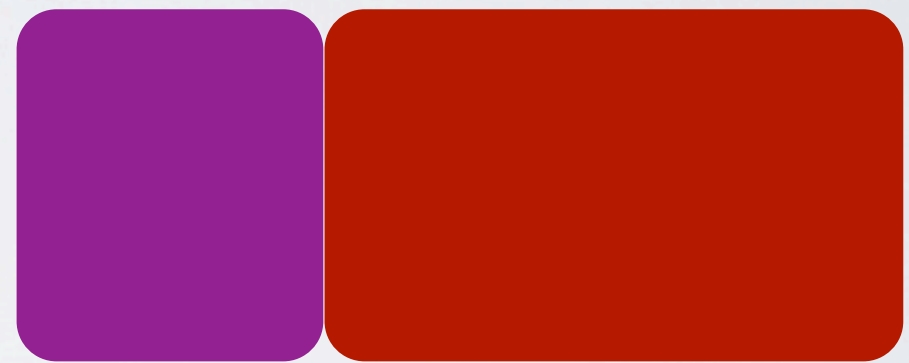
- Can explore other restrictions from physical + q. info-theoretic requirements (SBG & Shi 1205.4732 & WIP)

E.g. characterize information transfer

Minimal -- Simplest, most efficient form:



$$A = A_1$$



$$A_2 \otimes B$$

“Subsystem transfer”

Mod unitaries U_A, U_B etc. Saturates a subadd. condition

Contrast nonminimal:

$$\begin{array}{lcl} |\hat{0}\rangle|0\rangle & \rightarrow & |\hat{0}\rangle|0\rangle \\ |\hat{1}\rangle|0\rangle & \rightarrow & |\hat{1}\rangle|1\rangle \end{array}$$

Might allow you to measure dead or alive
Schrodinger's cat inside a black hole, but doesn't
transfer quantum information ...

Example 2 was minimal; Example 3 was not
(extra flux)

Possible reasons to expect (near-)minimal:

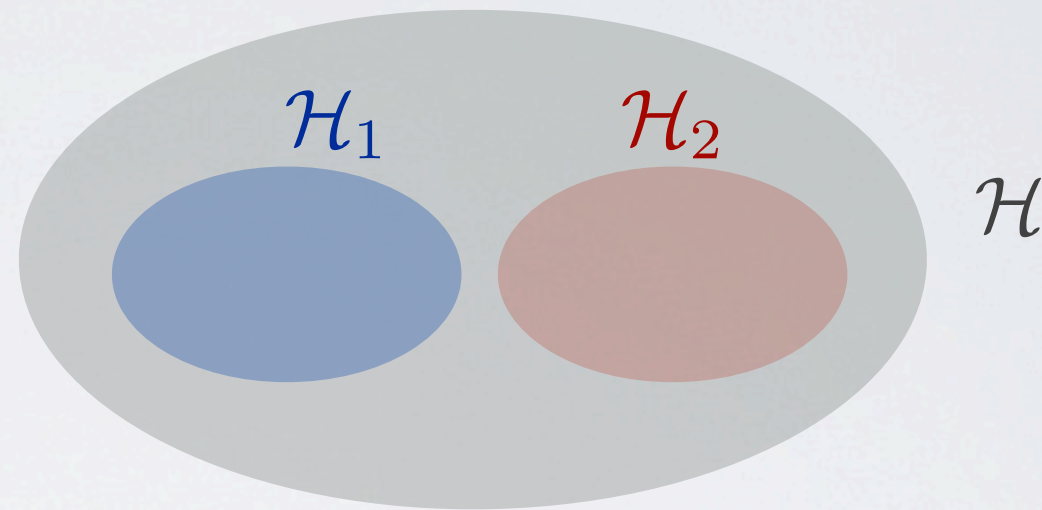
Weak coupling; \sim thermodynamic; small \mathcal{H}_{near}

Comments on Hilbert space networks:

A proposed broad picture -

Quantum states more basic than spacetime

Approximate
“localization”

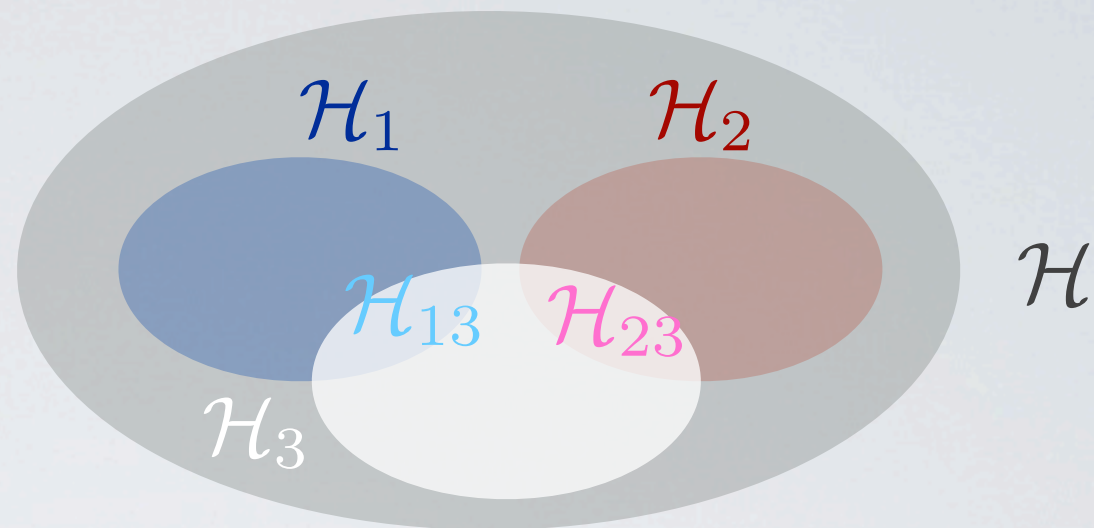


$$\mathcal{H}_1 \otimes \mathcal{H}_2 \in \mathcal{H}$$

Compare LQFT:

$$[\mathcal{O}(x), \mathcal{O}(y)] = 0 \quad , \quad (x - y)^2 > 0$$

~ quantum analog
of manifold:



(some common ideas w/ algebraic
QFT; also Banks “HST” -- though
important differences)

- Unitary evolution; ~local, LQFT

- Symmetries $\mathcal{H} \rightarrow S\mathcal{H}$ global
 $\mathcal{H} \rightarrow S_{loc}\mathcal{H} = S_1\mathcal{H}_1 \otimes S_2\mathcal{H}_2 \cdots$ local

Hilbert space networks: a possible framework for a
unitary theory of quantum gravity

Summary:

The information problem appears foundational.

A “most conservative” approach is to modify
macroscopic locality. (not QM, LI)

Unitarity can be restored through QI transfer from BH
subsystem ... not LQFT (but “nearly local,” nonviolent?)

Proposal: approximate spacetime emerges from a
broader framework - Hilbert space network