Incorporating Event Rates Into MSSM Parameter Determinations

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PROSPECTS, AlbaNova, September 15th, 2010







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Technical Issues And Implementation Obstacles Our implementation

Some Results With Fittino mSUGRA Results Non-Universal Gaugino Mass Results

Summary and Outlook

Cascade decays



 $\begin{array}{l} \text{Many } \tilde{g}, \tilde{q} \text{ expected.} \\ \tilde{g} \text{ must decay via } \tilde{q}, \tilde{q} \text{ decays} \\ \text{often to } \tilde{\chi}_2^0, \text{ typically } \tilde{\chi}_2^0 \\ \text{decays to } \tilde{\ell} \text{ which then must} \\ \text{decay to } \tilde{\chi}_1^0 \end{array}$

 Most effort for determining SUSY parameters so far has gone into cascade decay kinematics [*e.g.* Gjelsten, Miller and Osland, (2004); Gjelsten, Miller, Raklev, 2005)]

Cascade decays

Motivation for utilizing cross-sections

- \blacktriangleright Unknown center of momentum $+ \, \tilde{\chi}^0_1$ escaping detector = reconstruction difficult
- Event-by-event reconstruction impossible, but various kinematic quantities have distributions with well-defined endpoints

• e.g.
$$m_{q\ell\ell}^2 = (p_q + p_n + p_f)^2$$
 has endpoint $(m_{\tilde{q}}^2 - m_{\tilde{\chi}_2^0}^2)(m_{\tilde{\chi}_2^0}^2 - m_{\tilde{\chi}_1^0}^2)/m_{\tilde{\chi}_2^0}^2$
if $\frac{m_{\tilde{q}}}{m_{\tilde{\chi}_2^0}} > \frac{m_{\tilde{\chi}_2^0}}{m_{\tilde{\chi}_1^0}}$ (or other eqn.s for other cases)

Cascade decays Motivation for utilizing cross-sections

Motivation for utilizing cross-sections

- ▶ More observables = better determination of parameters! (in general)
- LHC cross-sections very sensitive to colored sparticle masses



Further motivation for utilizing cross-sections

- Cascade decay endpoint mimic points exist cross-sections can provide discrimination
- Cascade decay endpoints can be underconstraining for certain parameter regions and for certain hierarchies
- Non-supersymmetric models can have very similar spectra but with differing spins:
 - ► invariant mass distributions need to be well-measured to determine spin
 - cross-sections typically are an order of magnitude more for UED models than SUSY models with similar spectra

Obstacles Our implementation

Technical Issues

- ► Even LO SUSY-QCD calculations are not that fast
 - (2.5 minutes for Prospino 2 to calculate $\sigma(\tilde{g}\tilde{g} + \tilde{q}\tilde{g} + \tilde{q}\tilde{q} + \tilde{q}\tilde{q}^*)$ on my laptop at LO, 6 minutes at NLO)
- Dependent on unknown masses
 - $\blacktriangleright \Rightarrow$ Published studies restricted to single points in many-parameter space
- Not easily invertible
- Not easy to account for experimental cuts, such as on transverse momentum (though possible with a lot of computing power)
- ► Lester, Parker, White (hep-ph/0508143):
 - Markov chain exploration with full Monte Carlo simulation of SUSY events at each point
 - ► Supercomputer ran ISAJET, HERWIG then ATLFAST for LHC
 - ▶ restricted to 1000 events simulated per point at leading order!
 - proof of principle, not intended to be repeated often

Obstacles Our implementation

Our implementation - overview

Herbi Dreiner, Michael Krämer, Jonas Lindert, B. O'L.: self-contained code which

- ► takes LHC-scale SUSY spectrum (*e.g.* from SLHA-format file)
- looks up table of (NLO) cross-sections for colored sparticle production
- works out relevant cascade decays and multiplies with relevant branching ratios (BRs taken from SLHA file, such as produced by SPheno)
- applies approximations for cut acceptances depending on sparticle masses
- returns event rates for particular signals

Error estimated to be 15% on cross-section (typical NLO error including PDF uncertainty *etc.*), 5% on cut acceptances (from comparison with full parton-level MC simulation with Herwig++), 10% to account for jet showering (work in progress).

Obstacles Our implementation

Some details of our implementation

- currently 2 signals:
 - ▶ 2 or more jets passing p_T, η cuts with $\not\!\!\!E_T$ cut
 - $e\bar{e} + \mu\bar{\mu} \mu\bar{e} e\bar{\mu}$, with p_T, η cuts
- ▶ looks up table by $m_{\tilde{g}}, m_{\tilde{q}}$, for this point we have:
 - NLO cross-sections for g̃g, q̃g, q̃g, q̃q combinations, b̃b, t̃t also accounted for separately
 - ► numbers parameterizing cut acceptances for massless particles (e, µ, j assumed so) for given energies in q̃ rest frame
- works out relevant cascade decays and multiplies with relevant branching ratios
- ► applies approximations for cut acceptances depending on sparticle masses
 - \blacktriangleright distribution of lepton energies in \tilde{q} rest frame calculated from sparticle masses
 - ▶ this distribution is convoluted with acceptances for given energies
- returns event rates for particular signals

Obstacles Our implementation

Implementation into Fittino

- ▶ Thus far, only used in (private at the moment) version of Fittino
- Used as normal observables in Fittino

Fittino: a program by Philip Bechtle, Klaus Desch and Peter Wienemann (http://www-flc.desy.de/fittino/)

• See Klaus Desch's talk on Thursday

Some results from this implementation have been published in JHEP **4** (2010).

mSUGRA Results Non-Universal Gaugino Mass Results

I + rates

Plots for 7 TeV, M_0 against $M_{1/2}$

I, rates



 $(\text{``I''} = m_{\ell\ell}^{\max}, m_{q\ell\ell}^{\max}, m_{q\ell}^{\log}, m_{q\ell}^{\operatorname{high}}, \text{``II''} = m_{q\ell\ell}^{\operatorname{thr.}}, m_{T2}^{\tilde{q}}, m_{\tau\tau}^{\max}, m_{tb}^{w})$

Data input was SPS1a: $M_0 = 100$ GeV, $M_{1/2} = 250$ GeV, $A_0 = -100$ GeV, $tan(\beta) = 10$.

mSUGRA Results Non-Universal Gaugino Mass Results

Plots for 14 TeV, M_0 against $M_{1/2}$

I + II, rates





Some improvement, but at 14 TeV, we expect that there will be plenty of endpoint observables anyway.

mSUGRA Results Non-Universal Gaugino Mass Results

Plots for 7 TeV, M_0 against M_3

I, rates





Two million MC steps.

Ten thousand MC steps.

Adding rates may allow a fit to 6 parameters when endpoints alone cannot constrain the system enough with early data.

mSUGRA Results Non-Universal Gaugino Mass Results

Plots for 14 TeV, M_0 against M_3

I + II, rates



I + II + rates

Summary and Outlook

Summary:

- Cross-sections can be calculated within 20 30% quickly
- \blacktriangleright Rates can make a big difference to reducing errors on $M_{1/2}$ and $\tan\beta$ in Fittino

Outlook:

- ► Further signals being added (*e.g.* multilepton signals without OSSF-OSDF subtraction)
- Correlations to be investigated

Fit at 7 TeV with 100% uncertainties SPSIa numbers Invariant mass distribution shapes Motivation for NLO Other implementations



B. O'Leary



Fit at 7 TeV with 100% uncertainties SPS1a numbers Invariant mass distribution shapes Motivation for NLO Other implementations

Results for a fit to SPS1a data at 7 TeV, 1 fb⁻¹, 100% uncertainties



I + rates, 100%



Fit at 7 TeV with 100% uncertainties **SPS1a numbers** Invariant mass distribution shapes Motivation for NLO Other implementations

SPS1a spectrum



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PROSPECTS, 15/09/2010

Fit at 7 TeV with 100% uncertainties **SPS1a numbers** Invariant mass distribution shapes Motivation for NLO Other implementations

SPS1a inputs for 1 $\rm fb^{-1}$ at 7 and 14 TeV

observable	nominal	statistical uncertainty			1	
	value	for 7 Te	$V/1 \text{ fb}^{-1}$	for 14 TeV/1 fb	_1	
group I						
$m_{\ell\ell}^{ m max}$	80.4	4	.4	1.5		
$m_{q\ell\ell}^{\max}$	452.1	36	.0	12.0		
$m_{q\ell}^{\rm low}$	318.6	19	.7	6.5		
$m_{q\ell}^{ m high}$	396.0	13	.5	4.5		
group II						
$m_{q\ell\ell}^{\rm thr.}$	215.6		-	22.8		
$m_{T2}^{\ddot{q}}$	531.0		-	16.9		
$m_{ au au}^{ m max}$	83.4		-	10.8		
m_{tb}^{w}	359.5		-	37.0		
$r_{ ilde{\ell} ilde{ au}\mathrm{BR}}$	0.076		-	0.008		
Event rate [fb]		7 TeV		1	14 TeV	
	nomi	nal value	uncertain	ty nominal val	ue uncertainty	
R _{jj} ∉ _⊤	4.6	5×10^3	9.1×10^{-1}	4.8×10^4	9.5×10^{3}	
$R_{\ell\ell j j \not \in_T}$	1.6	5×10^2	3.2 ×10	1.5×10^{3}	3.0×10^2	

Table of results

SPS1a	M ₀ [GeV] 100	M _{1/2} [GeV] 250	tan eta 10	A ₀ [GeV] -100
7 TeV and 1 fb ^{-1} I + rates	99.0 ^{+9.9} 9.1	$250.0 \ ^{+8.7}_{-6.5}$	$10.7 \ ^{+4.0}_{-8.8}$	$55.2 \ ^{+1048}_{-254}$
14 TeV and 1 fb⁻¹ I + rates I + II, rates I + II + rates	$\begin{array}{r} 99.7 \begin{array}{c} +4.3 \\ -5.7 \\ 99.8 \begin{array}{c} +3.3 \\ -4.4 \\ 99.8 \begin{array}{c} +3.9 \\ -4.2 \end{array}$	$251.1 \stackrel{+7.5}{_{-5.8}} \\ 249.7 \stackrel{+6.6}{_{-5.2}} \\ 251.3 \stackrel{+5.0}{_{-5.0}} \\$	$11.2 \begin{array}{c} +3.5 \\ -5.1 \\ 10.1 \begin{array}{c} +3.8 \\ -3.2 \\ 10.7 \begin{array}{c} +3.1 \\ -3.1 \end{array}$	$\begin{array}{r} -50.9 \begin{array}{c} ^{+1233}_{-350} \\ -94.1 \begin{array}{c} ^{+1610}_{-216} \\ -55.7 \begin{array}{c} ^{+263}_{-233} \end{array}$

SPS1a numbers Invariant mass distribution shapes Motivation for NLO Other implementations

Table of results

SPS1a	<i>M</i> 0 [GeV] 100	<i>M</i> 1 [GeV] 250	M ₂ [GeV] 250	<i>M</i> ₃ [GeV] 250
7 TeV I + rates	$91.1 \ ^{+27.3}_{-36.1}$	$236.5 \ ^{+67.1}_{-57.9}$	$242.6^{+51.6}_{-33.7}$	$251.0^{+9.5}_{-8.5}$
14 TeV I + rates I + II, rates I + II + rates	$\begin{array}{r} 98.5 \begin{array}{c} ^{+16.5}_{-18.4} \\ 102.7 \begin{array}{c} ^{+9.4}_{-21.4} \\ 98.6 \begin{array}{c} ^{+12.6}_{-11.2} \end{array}$	$\begin{array}{c} 245.8 \begin{array}{c} +55.7 \\ -40.7 \\ 258.0 \end{array} \\ \begin{array}{c} +32.5 \\ -51.1 \\ 249.6 \end{array} \\ \begin{array}{c} +31.7 \\ -24.7 \end{array}$	$\begin{array}{r} 244.2 \begin{array}{c} +42.1 \\ -19.4 \\ 255.4 \begin{array}{c} +43.6 \\ -41.7 \\ 248.7 \begin{array}{c} +24.9 \\ -15.5 \end{array}$	$250.3 \begin{array}{c} ^{+11.1}_{-7.0} \\ 251.4 \begin{array}{c} ^{+9.9}_{-12.2} \\ 252.1 \begin{array}{c} ^{+6.0}_{-7.1} \end{array}$

Fit at 7 TeV with 100% uncertainties **SPS1a numbers** Invariant mass distribution shapes Motivation for NLO Other implementations

NLO SUSY cross-sections

▶ NLO needed since LO cross-sections can vary by 100%



▶ K-factors also different for different sparticle production,

- ▶ e.g. at 7 TeV, $m_{\tilde{g}} \simeq 600$ GeV, $m_{\tilde{q}} \simeq 550$ GeV (close to SPS1a and LM1)
- $K_{\tilde{g}\tilde{g}} = 1.45, K_{\tilde{g}\tilde{q}} = 1.20, K_{\tilde{q}\tilde{q}} = 1.18$
- ► This can lead to quite different amounts of *b*-jets from sbottoms from gluinos, for example

Fit at 7 TeV with 100% uncertainties SPS1a numbers Invariant mass distribution shapes Motivation for NLO Other implementations

Endpoints

 Event-by-event reconstruction impossible, but various kinematic quantities have distributions with well-defined endpoints

$$\begin{array}{ll} \bullet \ e.g. \ m_{q\ell\ell}^2 = (p_q + p_n + p_f)^2 \ \text{has endpoint} \\ \left\{ \begin{array}{ccc} (m_{\tilde{q}}^2 - m_{\tilde{\chi}_2^0}^2)(m_{\tilde{\chi}_2^0}^2 - m_{\tilde{\chi}_1^0}^2)/m_{\tilde{\chi}_2^0}^2 & \text{if} & \frac{m_{\tilde{q}}}{m_{\tilde{\chi}_2^0}} > \frac{m_{\tilde{\chi}_2^0}}{m_{\tilde{\chi}_1^0}} \\ (m_{\tilde{q}}^2 m_{\tilde{\ell}}^2 - m_{\tilde{\chi}_2^0}^2 m_{\tilde{\chi}_1^0}^2)(m_{\tilde{\chi}_2^0}^2 - m_{\tilde{\ell}}^2)/m_{\tilde{\chi}_2^0}^2 m_{\tilde{\ell}}^2 & \text{if} & \frac{m_{\tilde{\chi}_2^0}}{m_{\tilde{\ell}}} > \frac{m_{\tilde{q}}m_{\tilde{\ell}}}{m_{\tilde{\chi}_2^0}} \\ (m_{\tilde{q}}^2 - m_{\tilde{\ell}}^2)(m_{\tilde{\ell}}^2 - m_{\tilde{\chi}_1^0}^2)/m_{\tilde{\chi}_2^0}^2 & \text{if} & \frac{m_{\tilde{\ell}}}{m_{\tilde{\chi}_1^0}} > \frac{m_{\tilde{q}}m_{\tilde{\ell}}}{m_{\tilde{\chi}_1^0}} \\ (m_{\tilde{q}}^2 - m_{\tilde{\chi}_2^0}^2) & \text{otherwise} \end{array} \right.$$

Fit at 7 TeV with 100% uncertainties SPS1a numbers Invariant mass distribution shapes Motivation for NLO Other implementations

Invariant mass distribution shapes

There is also information in the shape of the distribution [e.g. Gjelsten, Miller, Osland (2005, 2006)]...



... but maybe not so much in practice, such as when jet combinatorics are considered.



Fit at 7 TeV with 100% uncertainties SPSIa numbers Invariant mass distribution shapes **Motivation for NLO** Other implementations

Motivation for NLO

- Lot of effort has gone into supersymmetric QCD NLO processes [*e.g.* Beenakker, Höpker, Spira, Zerwas (1996)]
 - necessary to know if signals will be visible
 - may need to know SUSY backgrounds to some processes
 - may learn about sparticle masses
 - limits on sparticle masses if not seen at a collider
- Automated calculation available (e.g. Prospino2)



Figure: Sparticle mass exclusion plot, taken from CDF's website

Fit at 7 TeV with 100% uncertainties SPS1a numbers Invariant mass distribution shapes Motivation for NLO Other implementations

Other implementations

- Cross-section information has been used for SUSY parameter space explorations in the literature
- ► Lester, Parker, White (hep-ph/0508143):
 - Markov chain exploration with full Monte Carlo simulation of SUSY events at each point
 - ► Supercomputer ran ISAJET, HERWIG then ATLFAST for LHC
 - restricted to 1000 events simulated per point at leading order!
 - proof of principle, not intended to be repeated often
- Our implementation is *fast* O(100) floating-point operations
- Our aim, in context of Fittino, is to improve errors on fit, be reproducible, be flexible