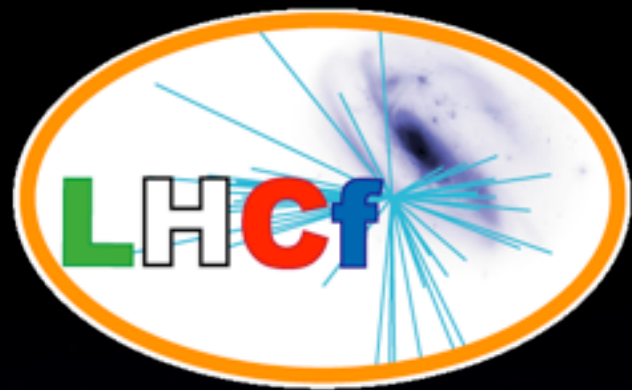




The first result of the CERN LHCf experiment

K.Kasahara

for the LHCf collaboration.

Waseda Univ.

★ For details: P.L to be published

TeVPA @Stockholm: Aug.1, 2011

Purpose of LHCf (=LHC forward experiment)

- ★ To select better nuclear interaction models
- or
- ★ To afford basic information for constructing a better nuclear interaction model

by observing forward energetic neutral particles at LHC for cosmic-ray physics

Excerpt from CERN Courier/Bulletin 2006

LHCf: a **tiny** new experiment joins the LHC

While most of the LHC experiments are on a grand scale, LHC forward(LHCf) is quite different. Unlike the massive detectors that are used by ATLAS or CMS, LHCf's largest detector is a **mere 30 cm**.



Most of the LCHf collaborators in one photo!



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Waseda University, Japan

K.Fukatsu, Y.Itow, K.Kawade, T.Mase, K.Masuda, Y.Matsubara, H.Menjo()*

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Solar-Terrestrial Environment Laboratory and KIT(), Nagoya University, Japan*

K.Yoshida

Shibaura Institute of Technology, Japan

T.Tamura

Kanagawa University, Japan

Y.Muraki

Konan University

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LBNL, Berkeley, USA

O.Adriani(1,2), L.Bonechi(1), M.Bongi(1), G.Castellini(1,3), R.D'Alessandro(1,2),

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J.Velasco, A.Faus

IFTIC, Universitat de Val`encia, Valencia, Spain

D.Macina, A-L.Perrot

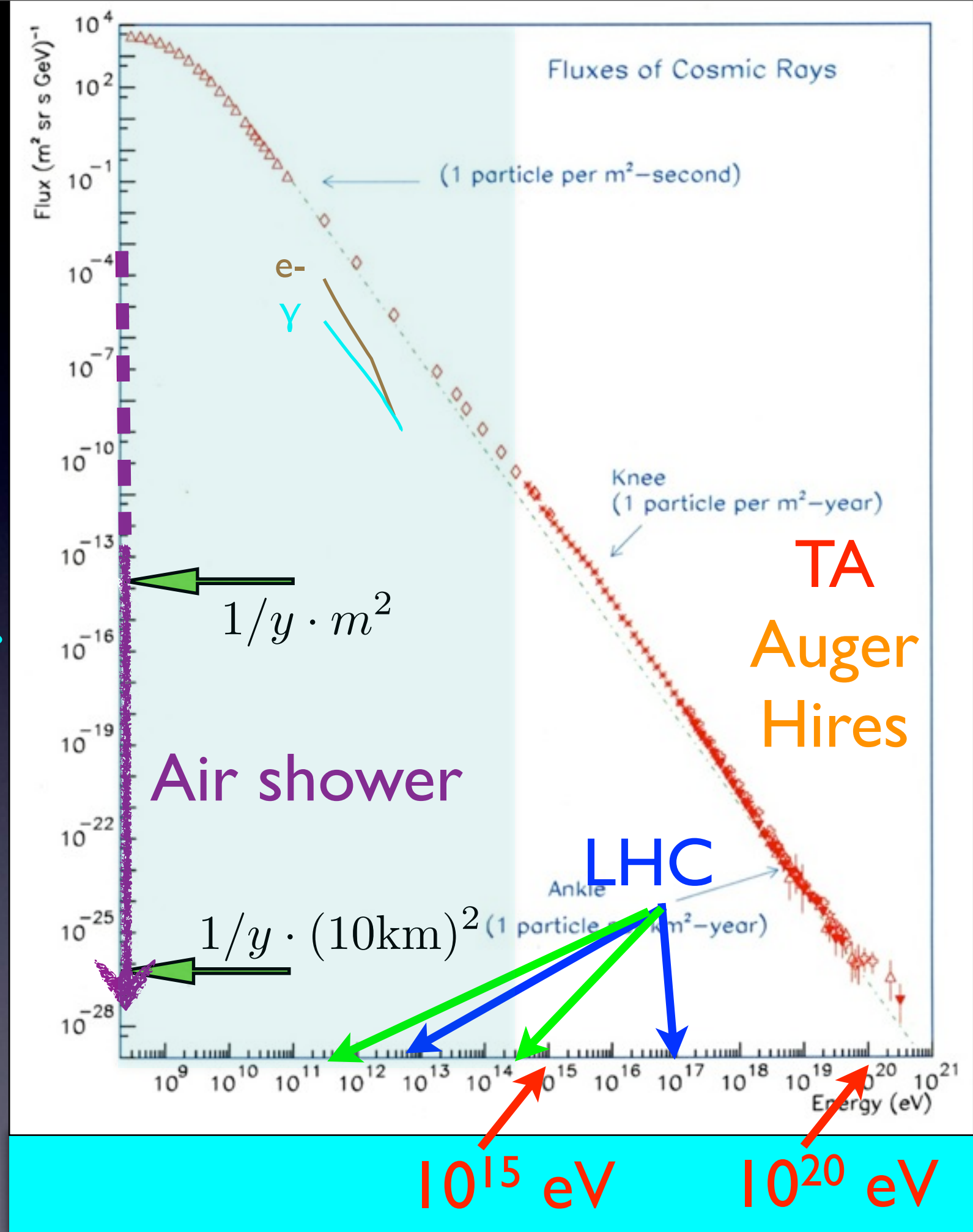
CERN, Switzerland

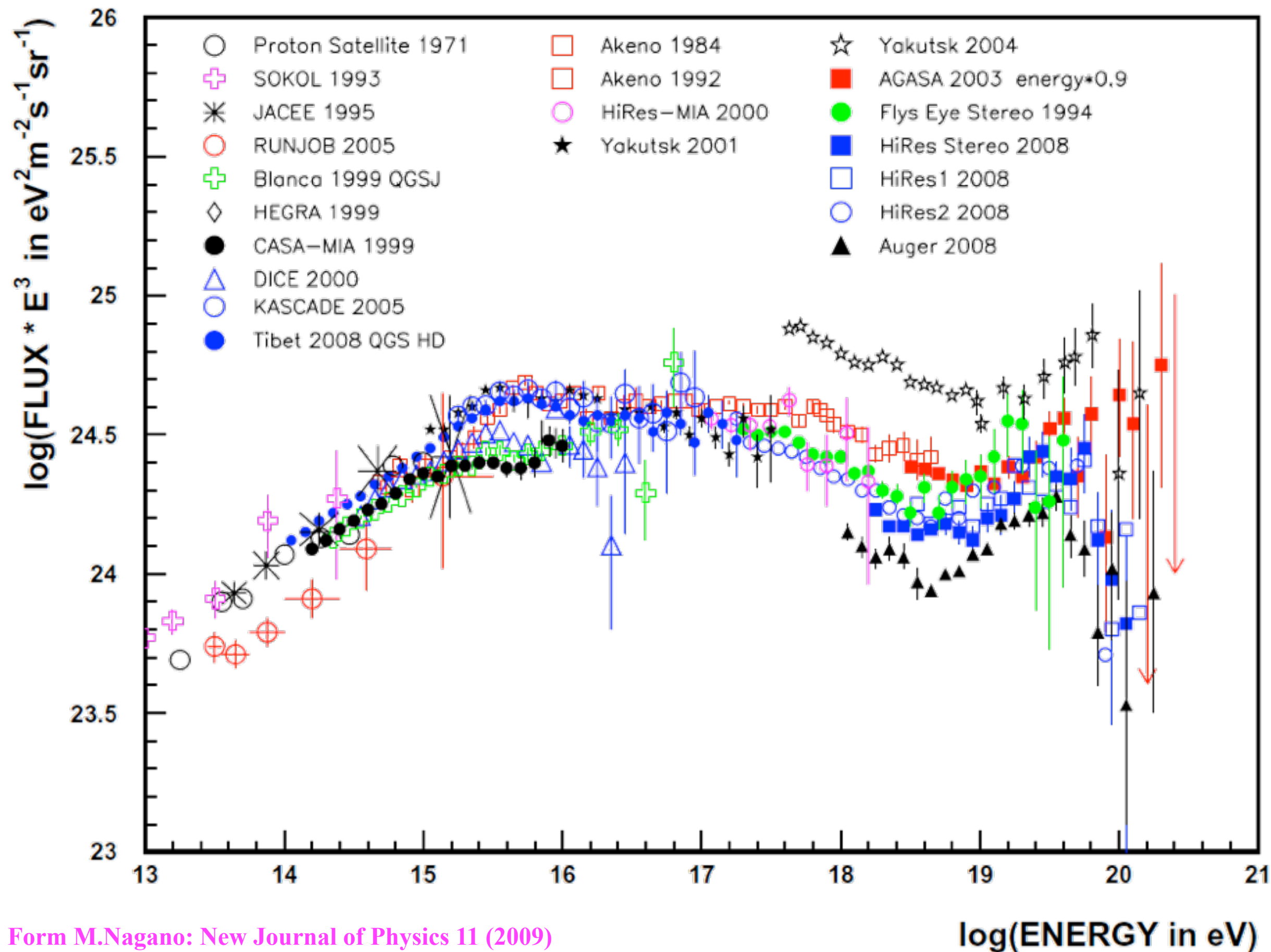


Motivation

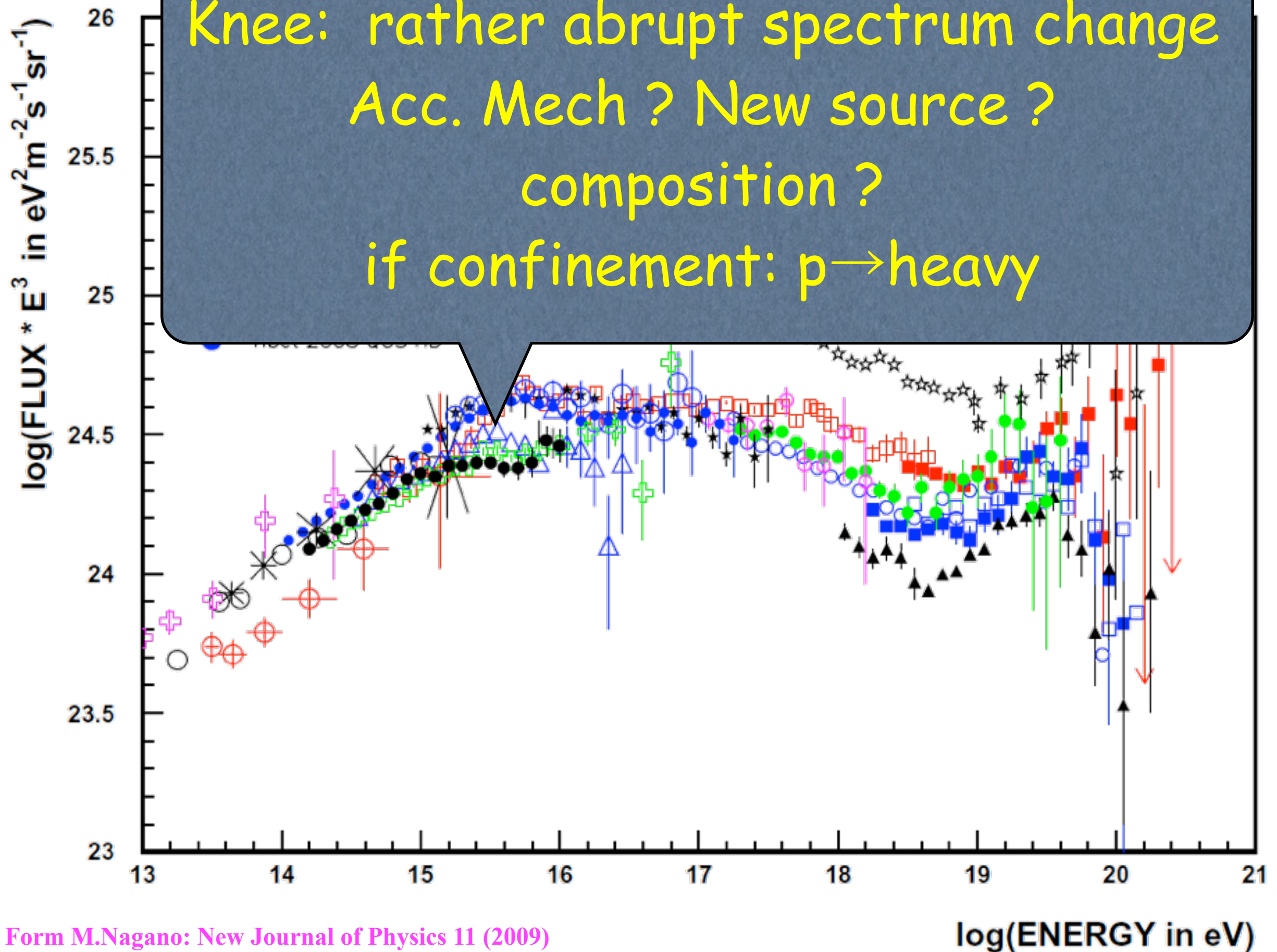
Cosmic-Ray Energy Spectrum

- $< 10^{14} \text{ eV}$: accel. SNR
- near-by e^- source?
- dark matter ?
- B/C ratio ...

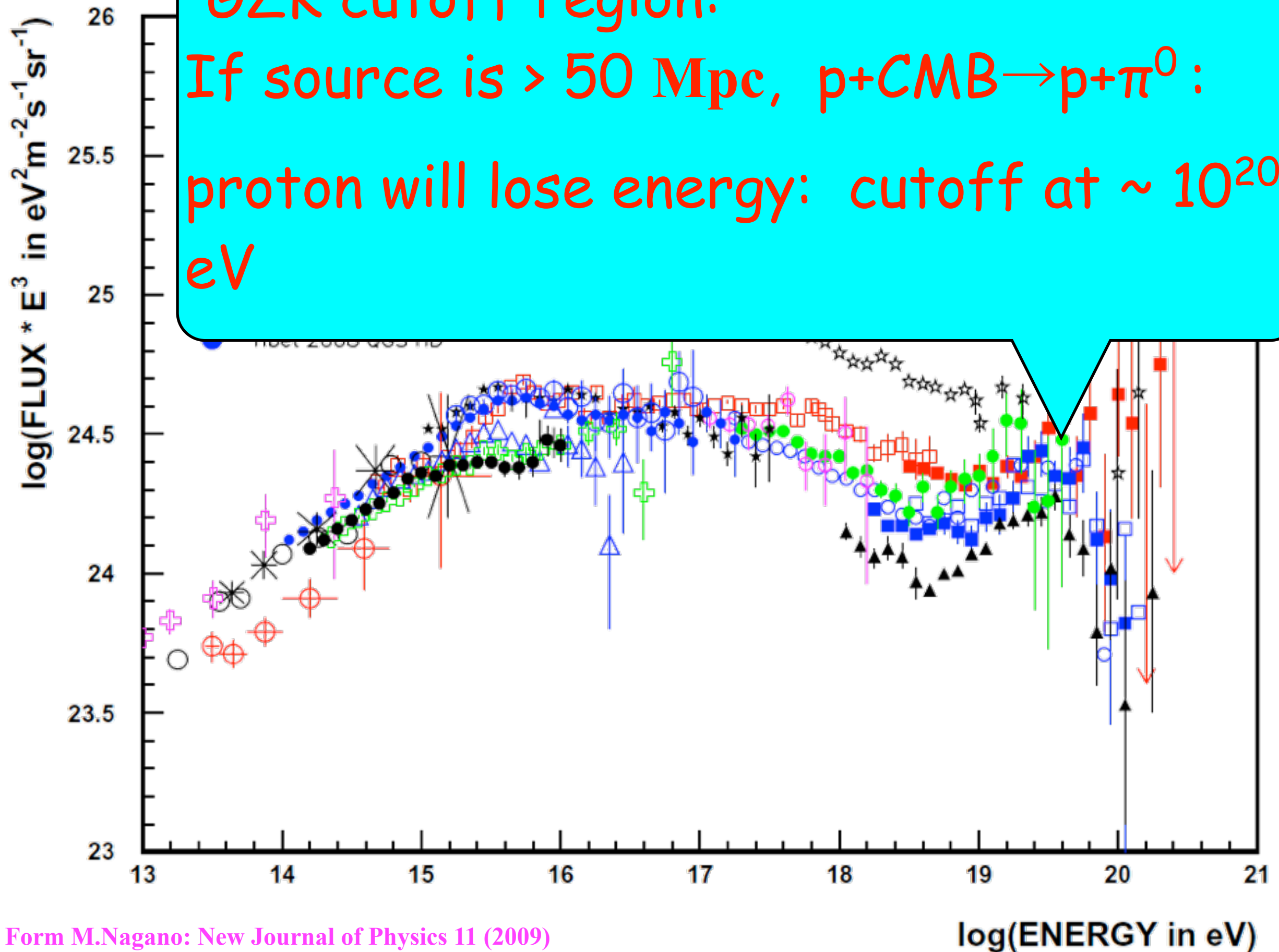




Knee: rather abrupt spectrum change
Acc. Mech ? New source ?
composition ?
if confinement: $p \rightarrow \text{heavy}$



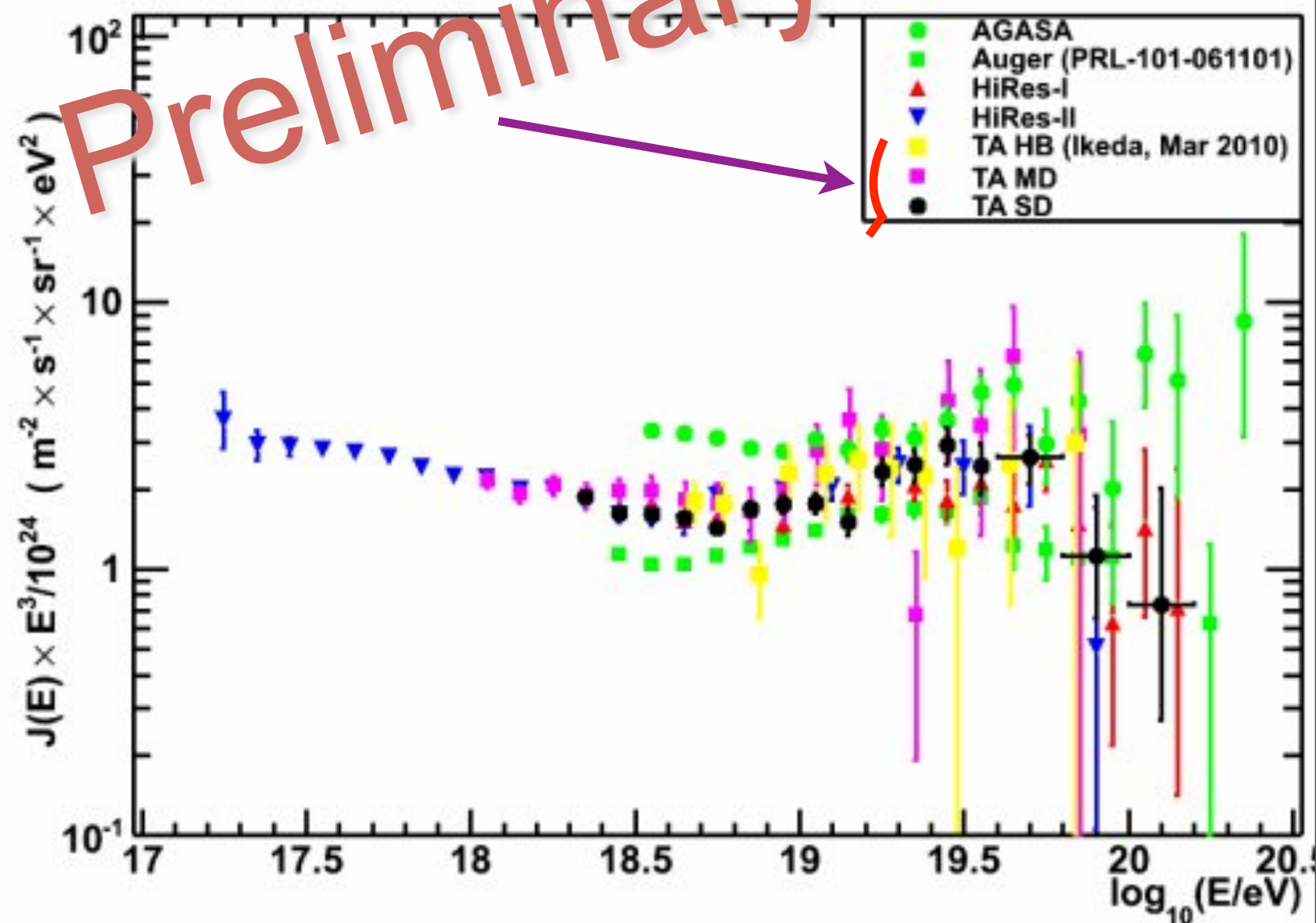
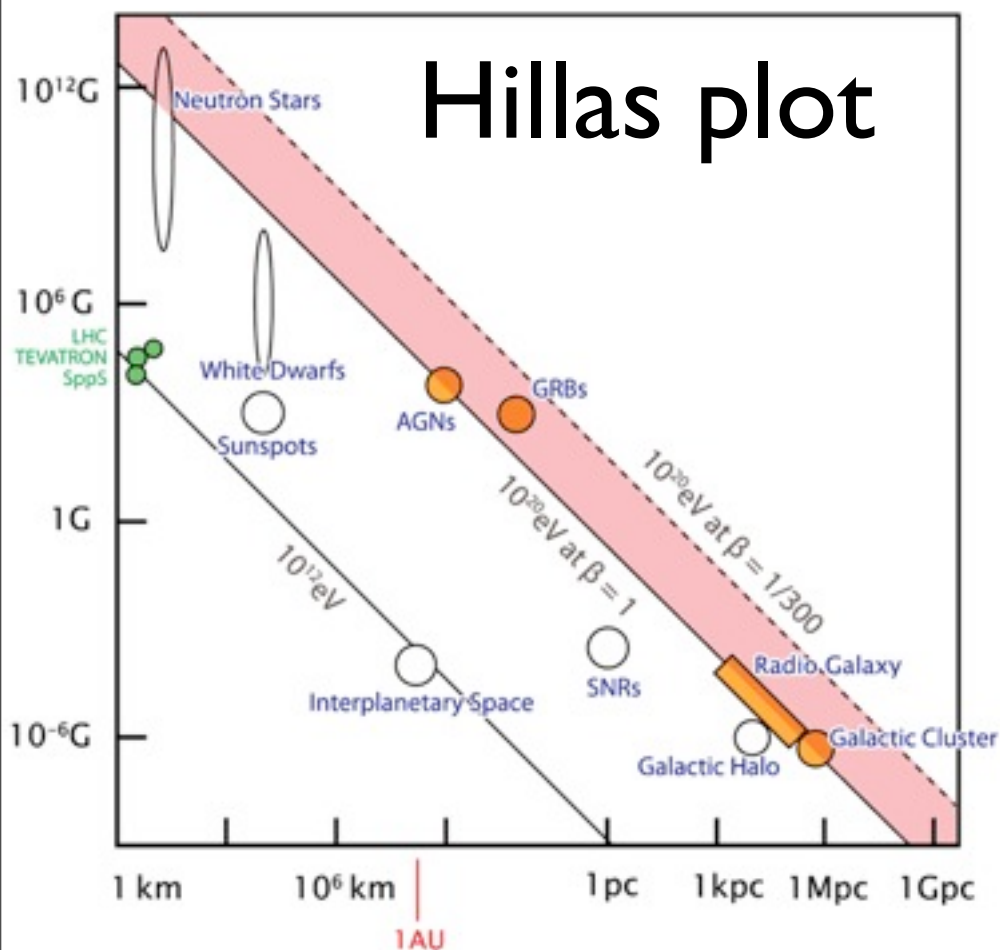
GZK cutoff region:
If source is > 50 Mpc, $p + CMB \rightarrow p + \pi^0$:
proton will lose energy: cutoff at $\sim 10^{20}$ eV



Recent TA result

If Super GZK:
various interesting
scenarios:

New results: GZK
cutoff

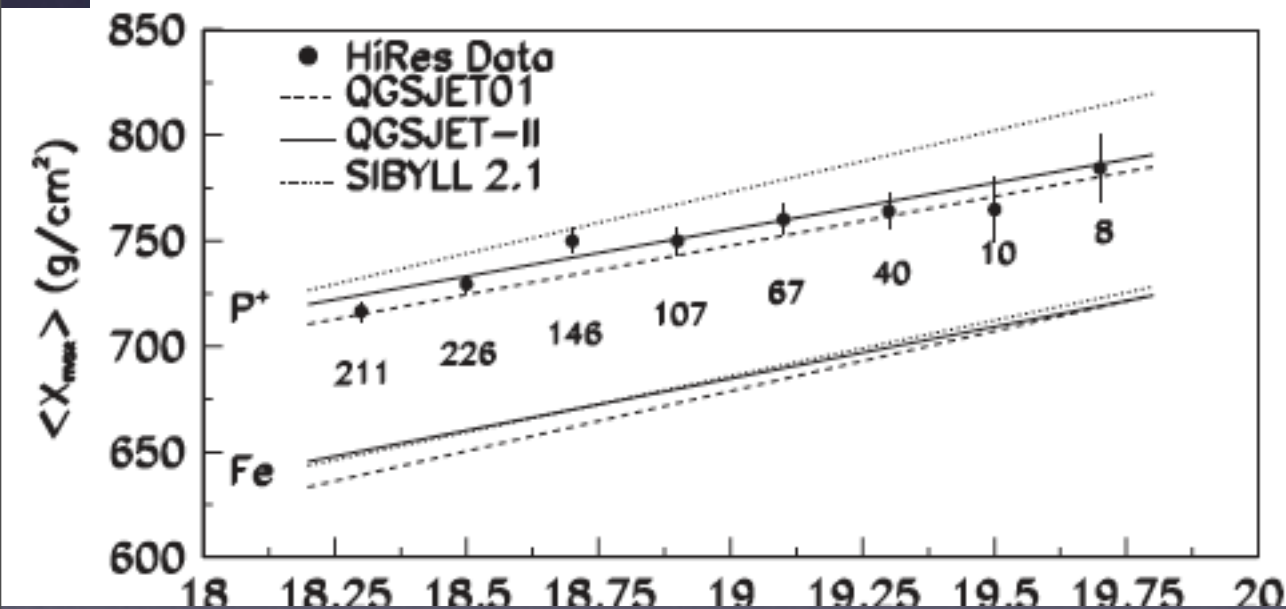
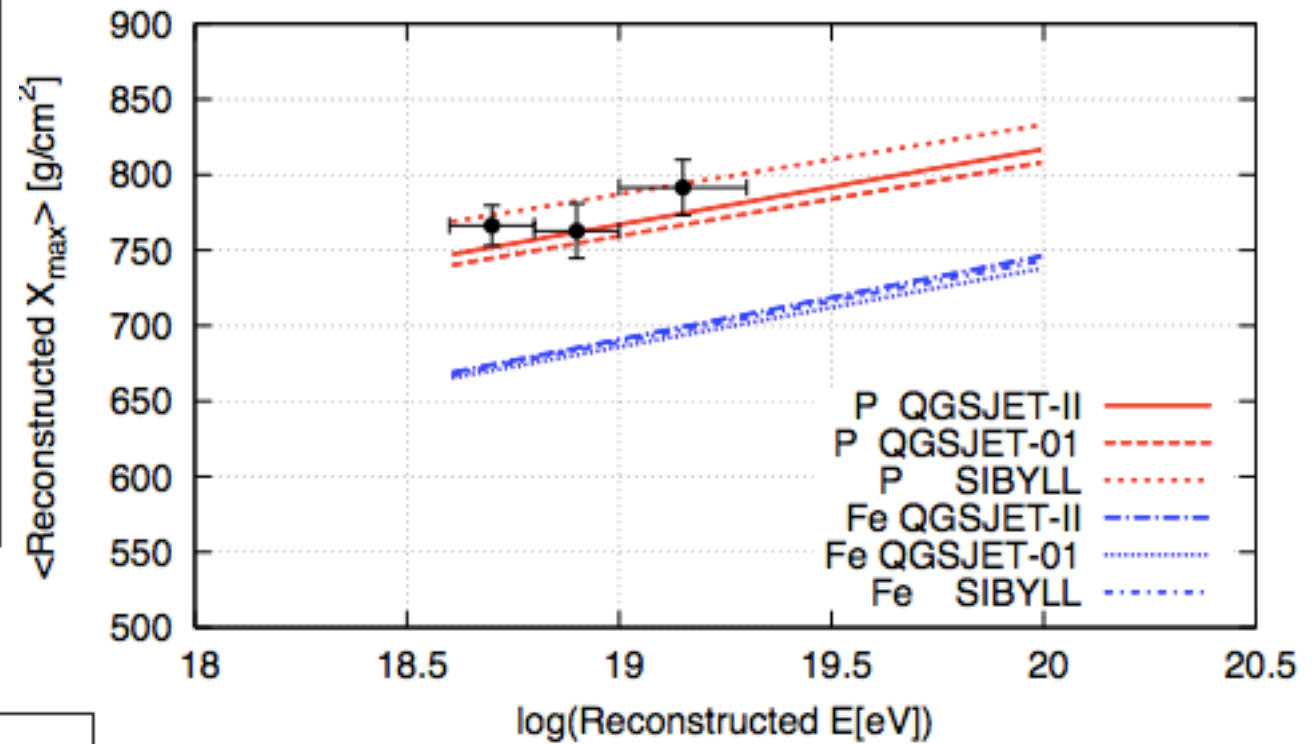
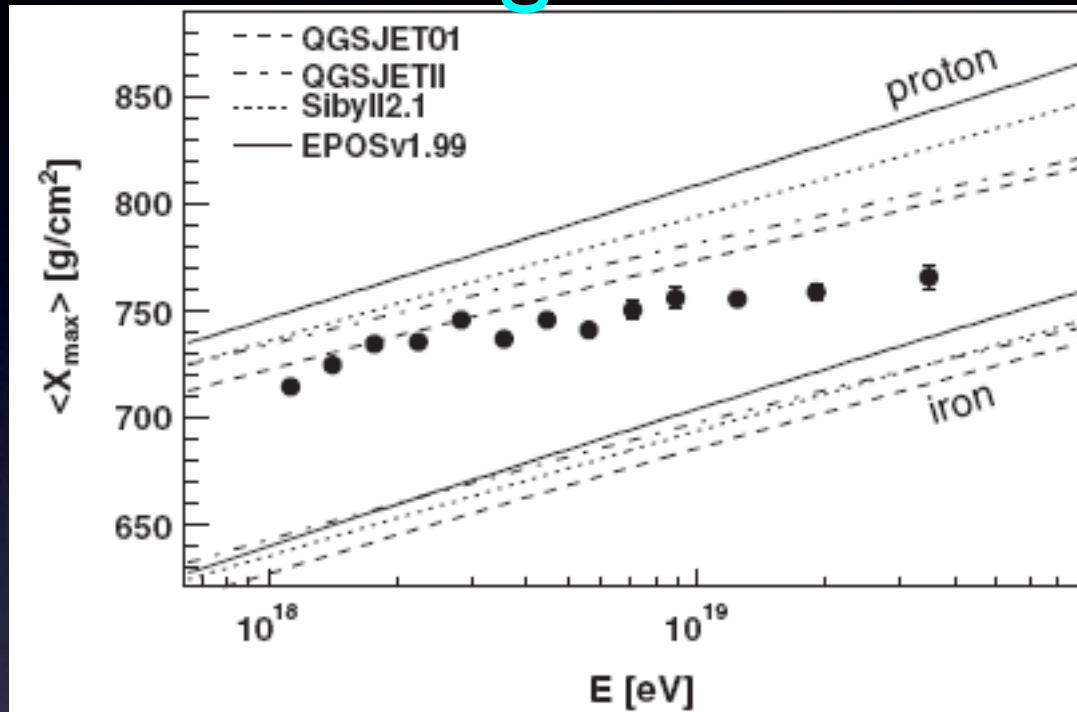


There still remain lots of problems:
source: AGN ? GRB ? .Composition ?
near-by source $\rightarrow$ super GZK recovery

Composition

Auger

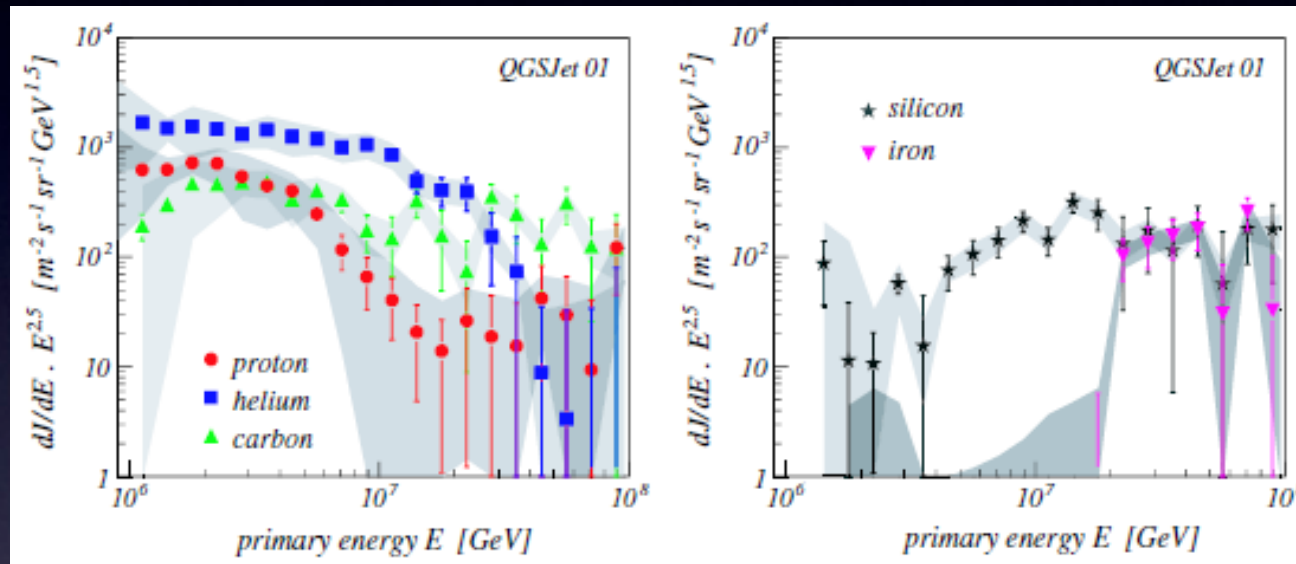
TA



HiRes

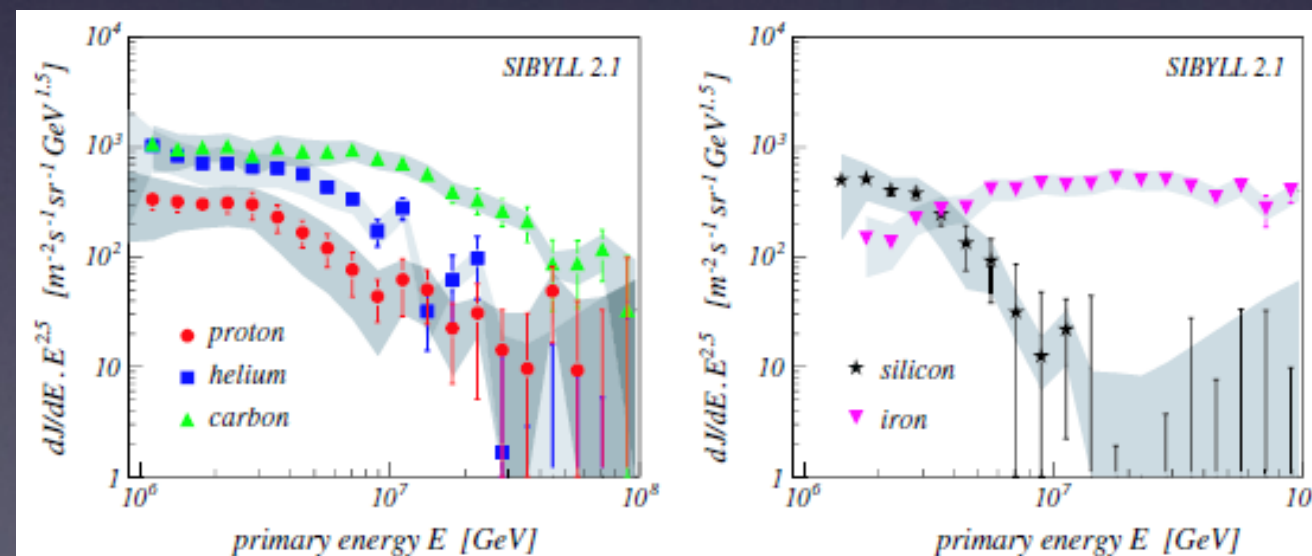
Not only UHECRs

composition (@ $\sim 10^{16}$ eV) by KASCADE



qgsjet I

Sibyll2.1



Air shower observation



Air shower observation

- SD: Surface array Detectors
- FD: Fluorescence Detectors
- Cherenkov
- Radio



A dark, blurry image of a landscape, possibly a field or forest, with a small, bright, circular light source visible in the lower center.

M.C: Indispensable tool

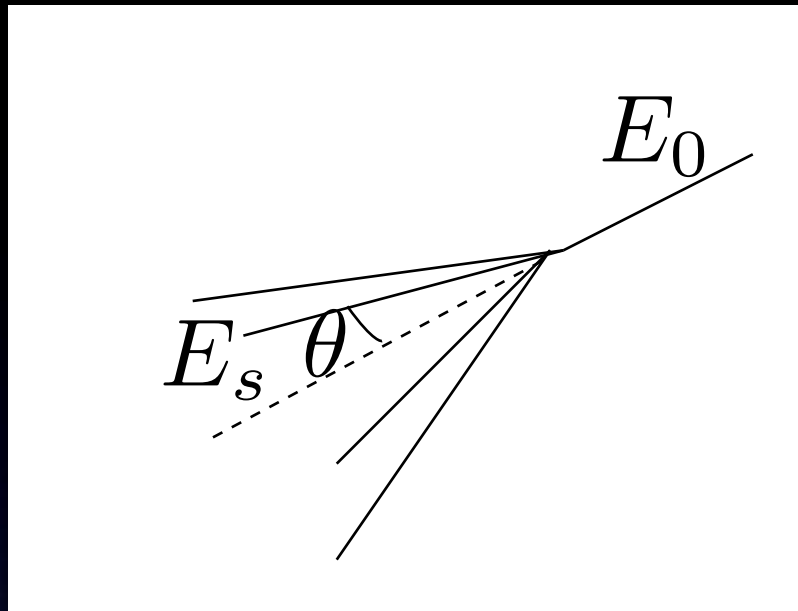
- Energy scale
- Composition
- Trigger efficiency, $S\Omega$

* Problems:

- * Hadronic interaction model
- * Computation time at $>10^{16}\text{eV}$

- Hadronic Interaction model
- Several interaction models in cosmic ray field
 - qgsjet1
 - qgsjet2
 - dpmjet3
 - sibyll
 - EPOS

Important variables for AS development



$$\frac{1}{\sigma} \frac{d\sigma}{dx}$$

$$x = \frac{E_s}{E_0}$$

$$\frac{1}{\sigma} \frac{d\sigma}{d\eta}$$

$$\int_{0.05}^{\infty} \eta f(\eta) d\eta \log\left(\tan\left(\frac{\theta}{2}\right)\right)$$

- For N_e and N_γ

- large x is important

- For N_μ

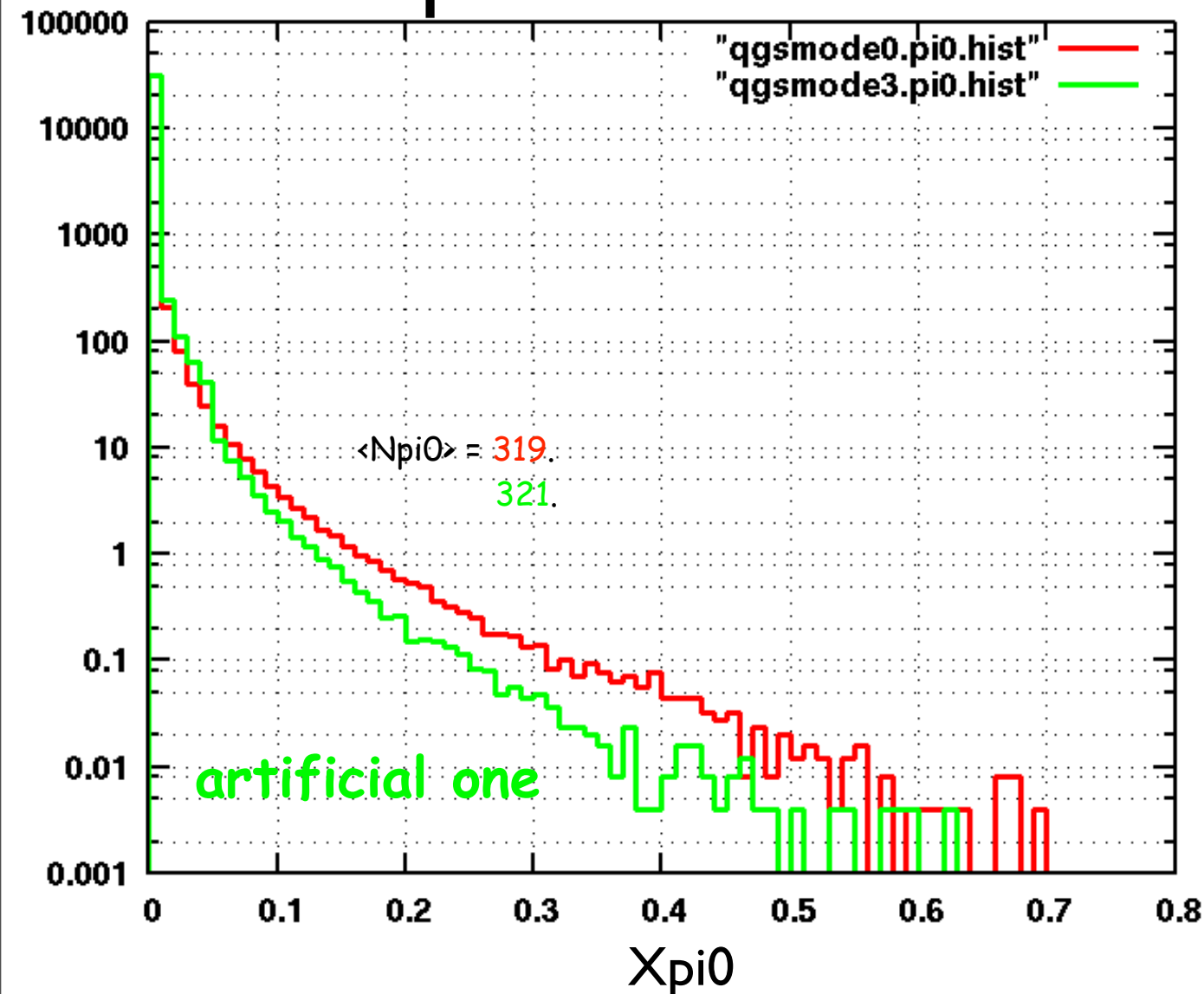
- large x is important at high energies but

- small x (or $\eta_{\text{cms}} \sim 0$) becomes important at lower energies

electron/gamma in AS
 50~60 % are from pi, K with $x > 0.1$
 (~40 % from γ with $x > 0.05$)

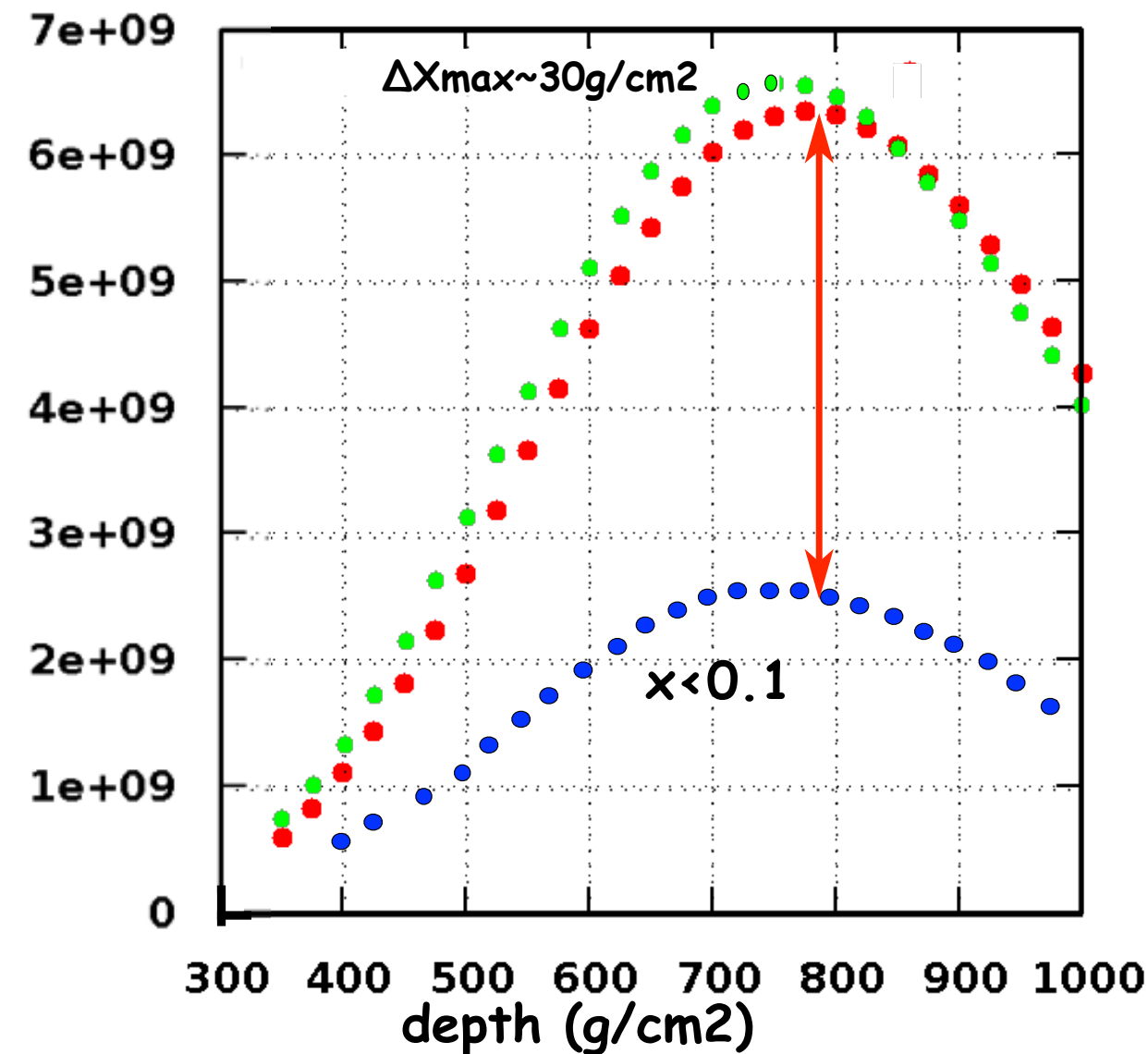
$p \ 10^{19} \text{eV}$

$dn/dx/ev: \pi^0$

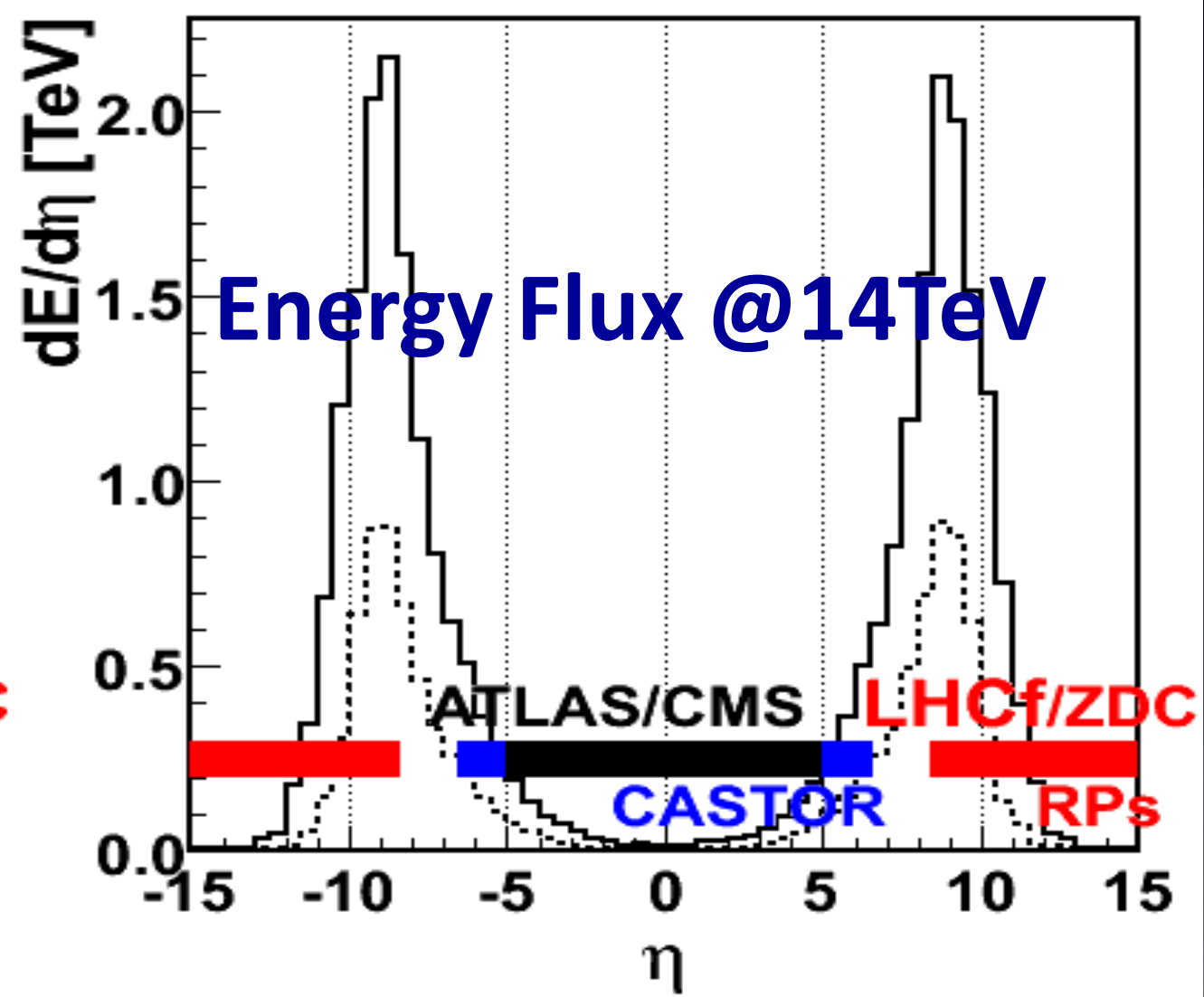
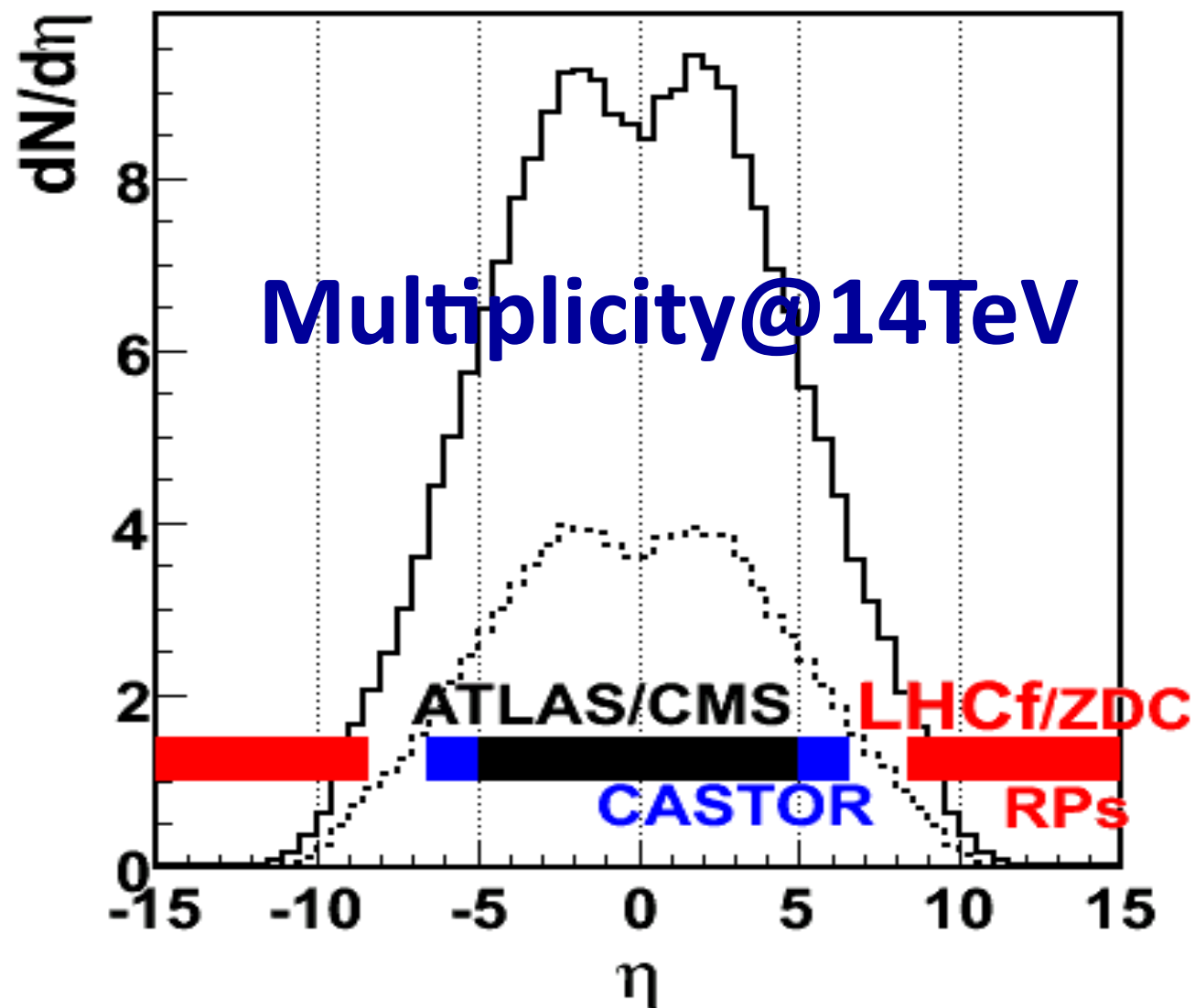


Ne

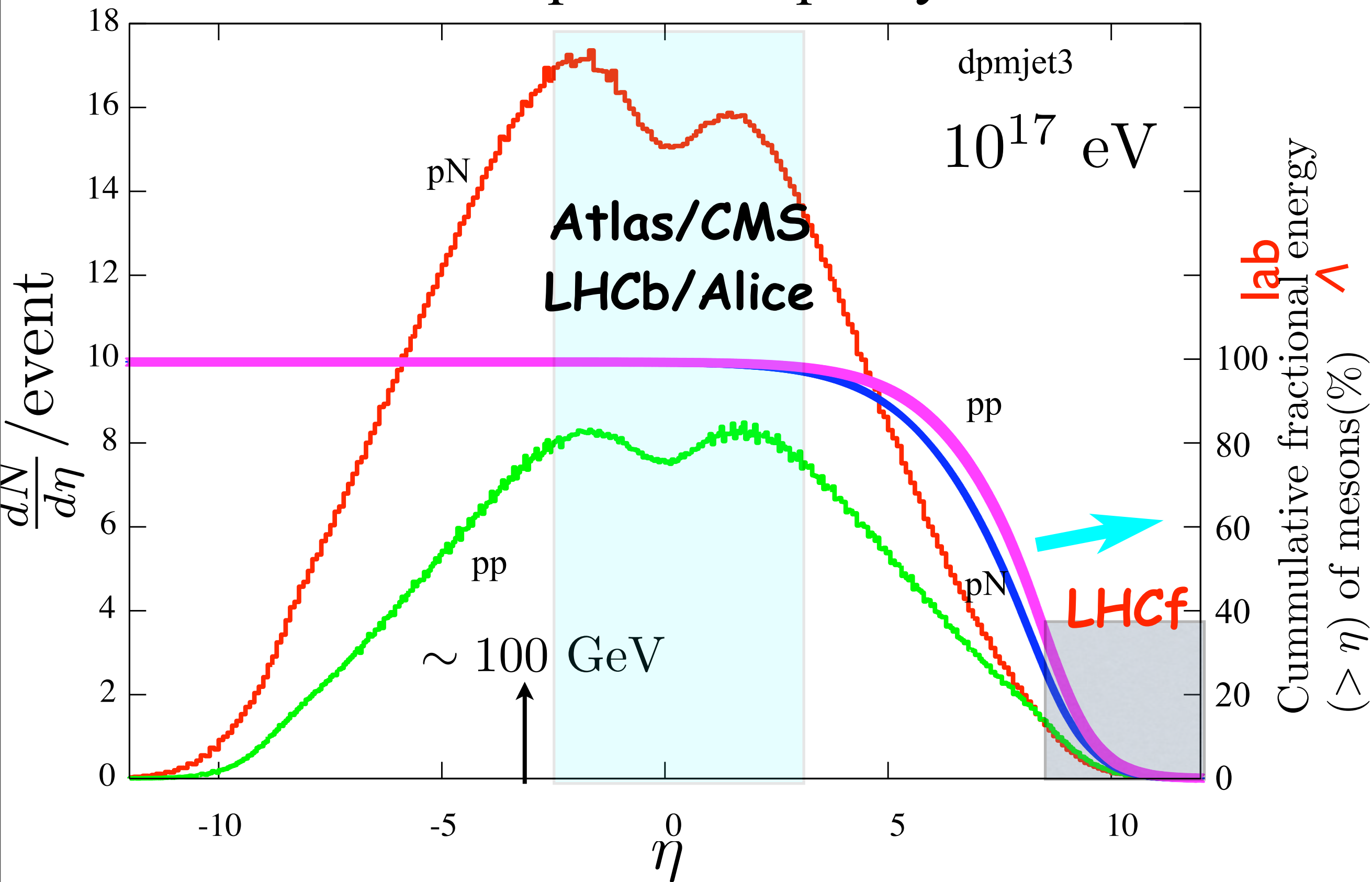
As transtion curve



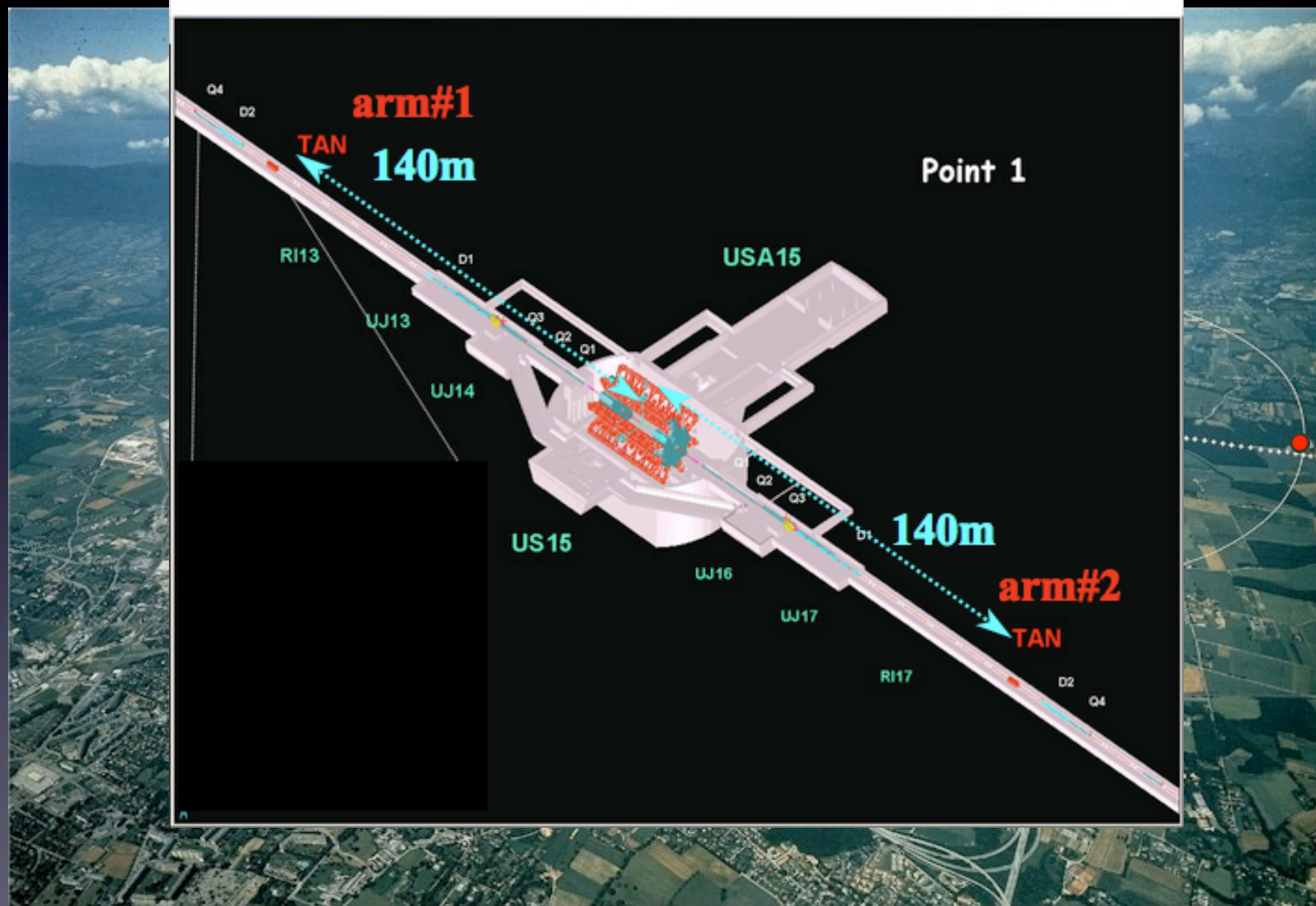
In terms of number and energy in CMS



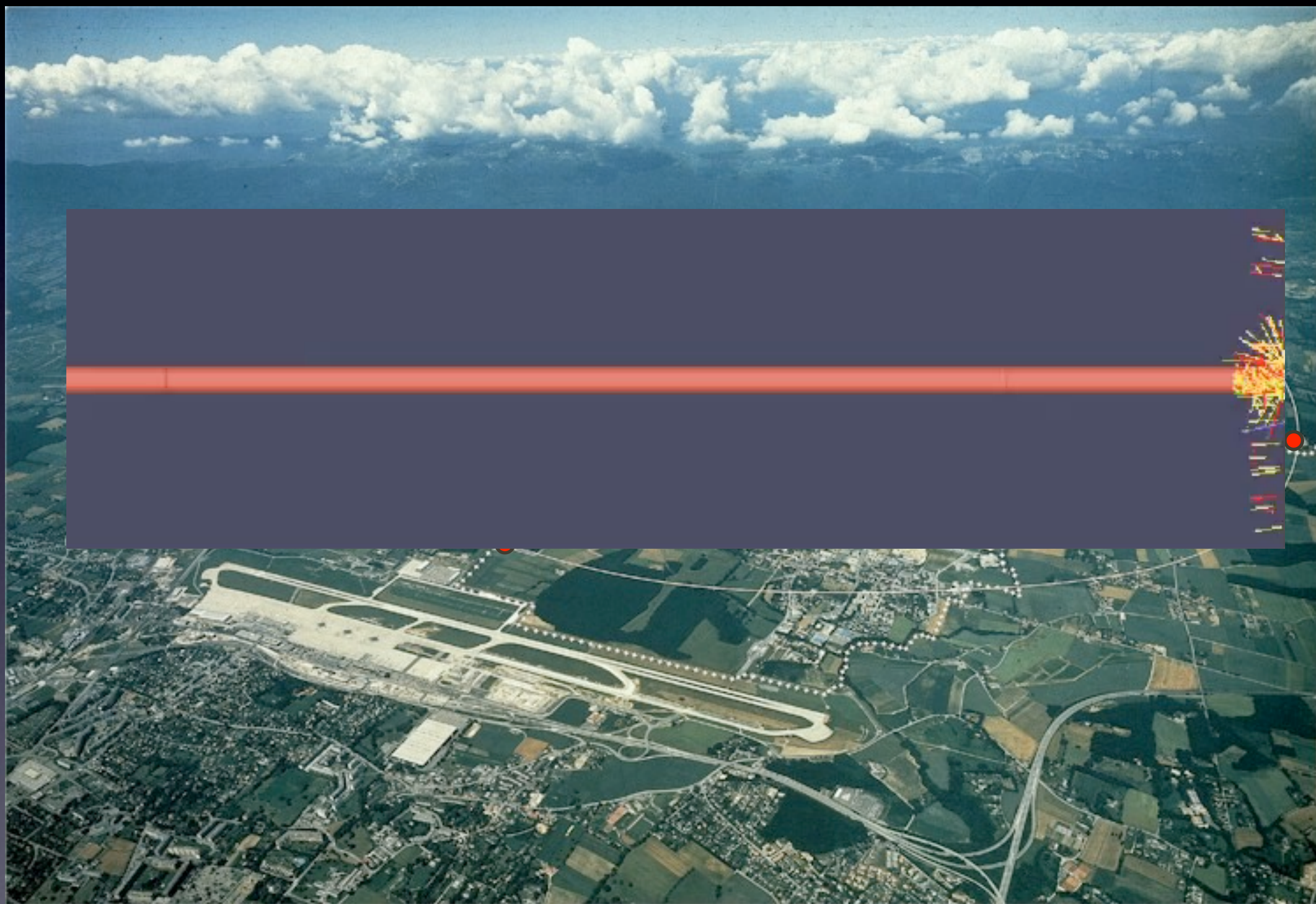
Meson pseudo rapidity

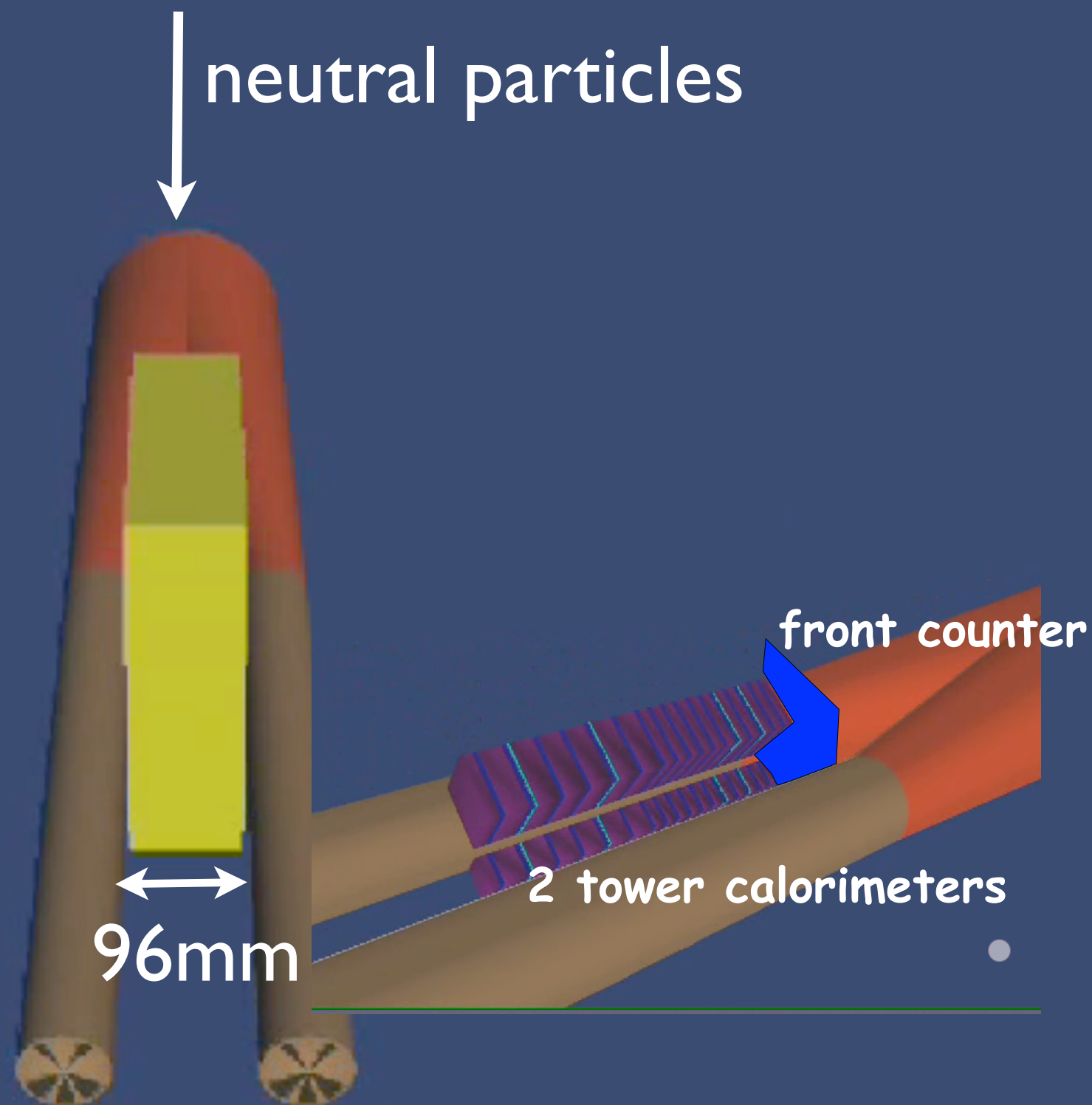
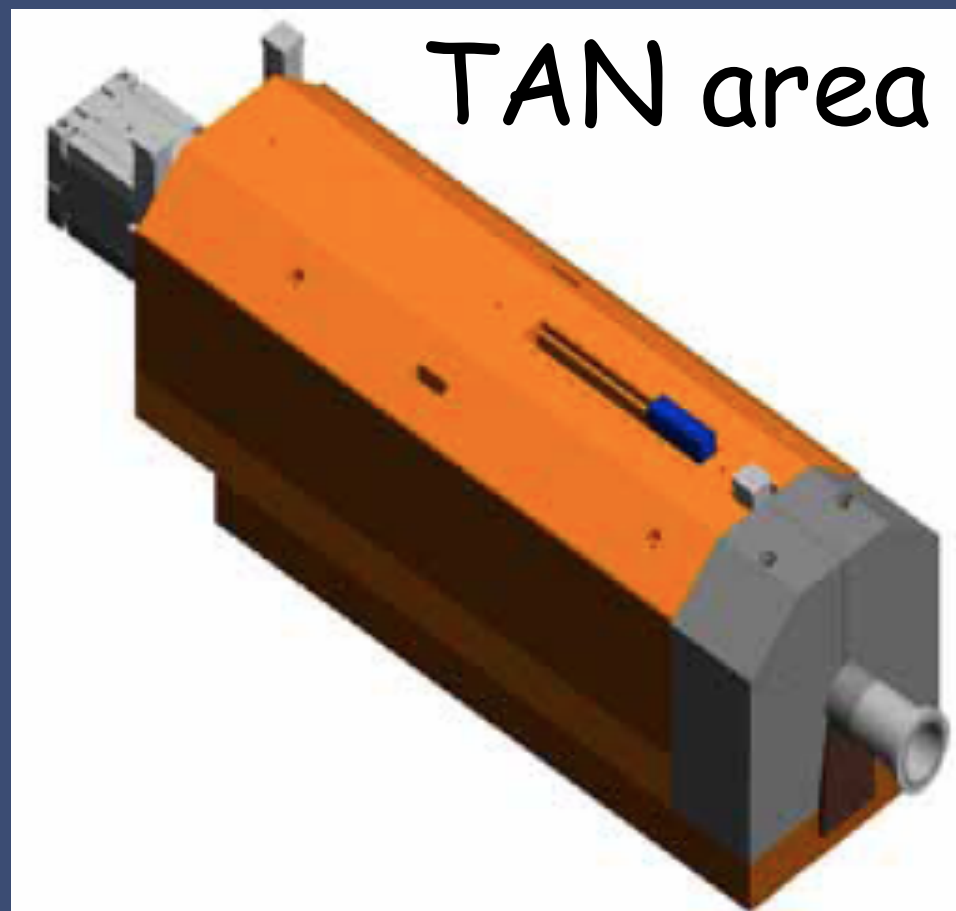




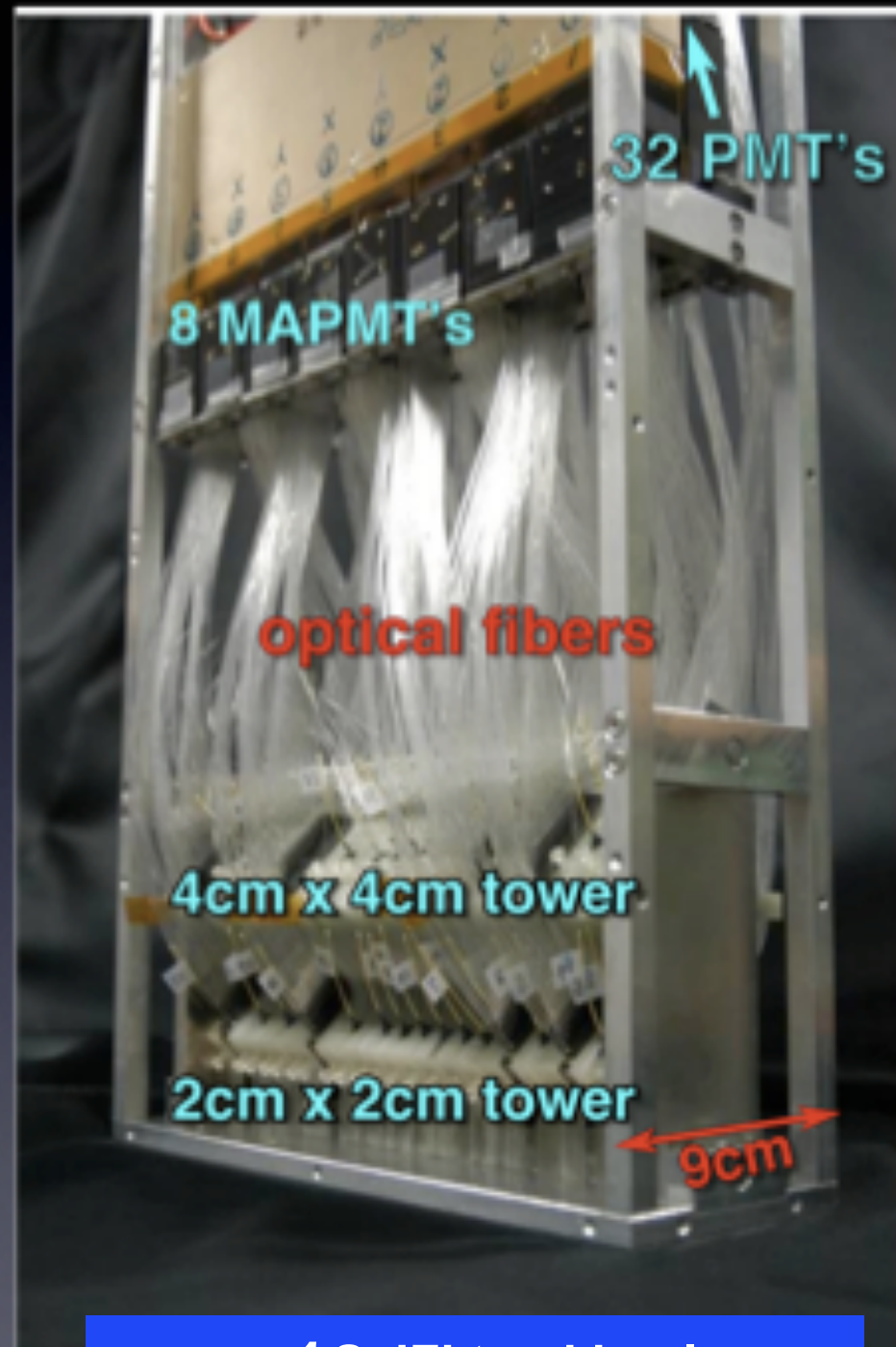








Arm #1



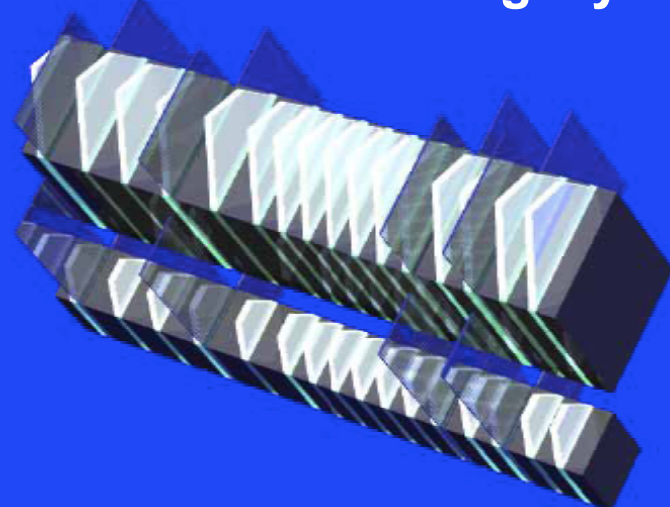
Arm #2



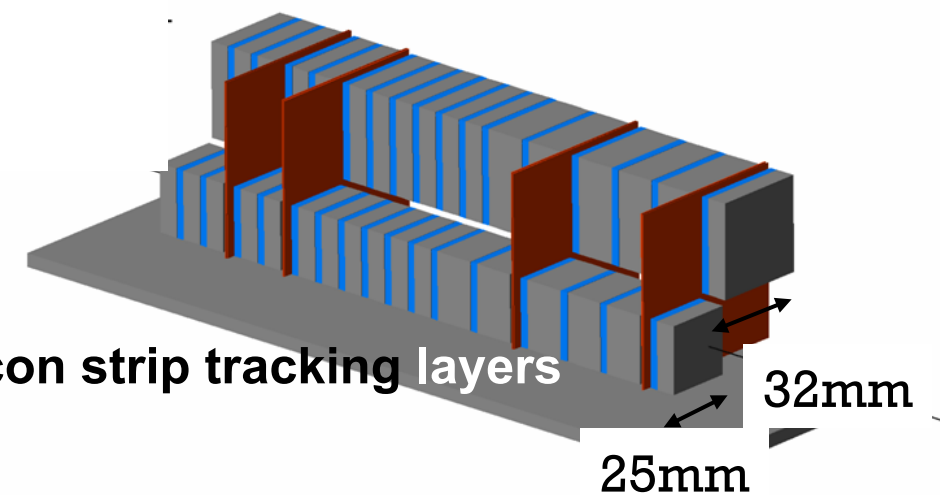
W 44X0

$1.7\lambda_c$

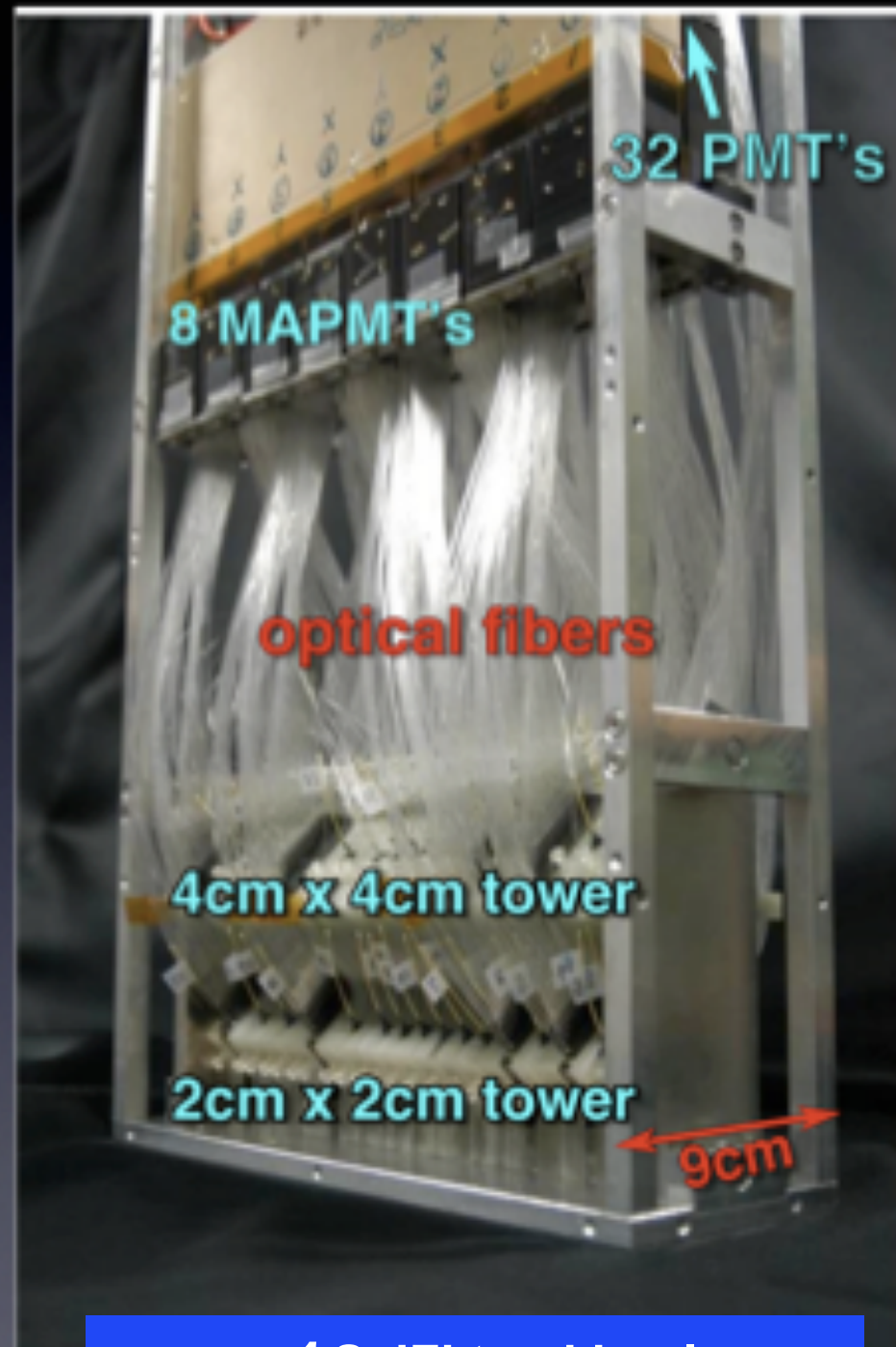
4 SciFi tracking layers



4 Silicon strip tracking layers



Arm #1

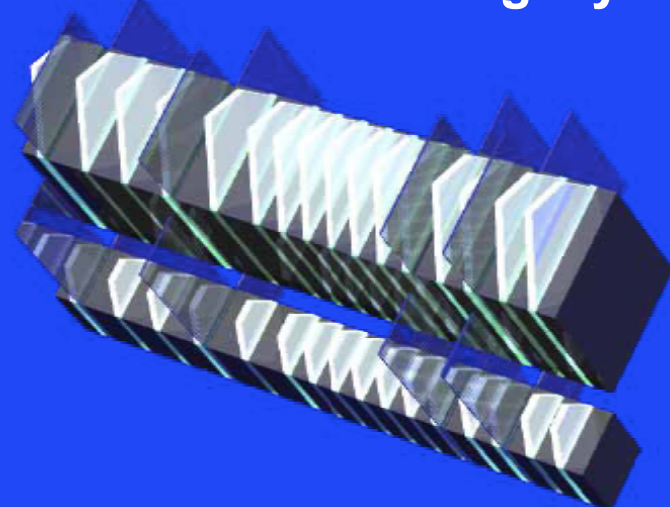


Arm #2

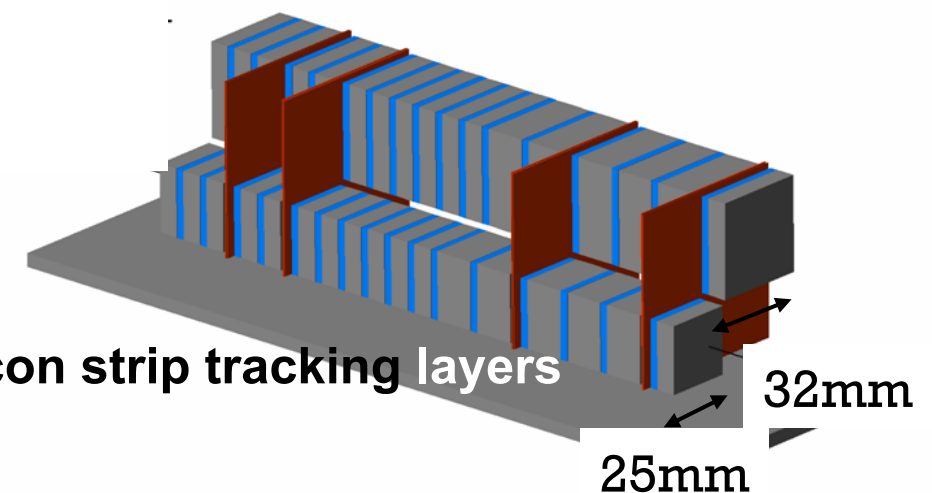
W 44X0
 $1.7\lambda_c$



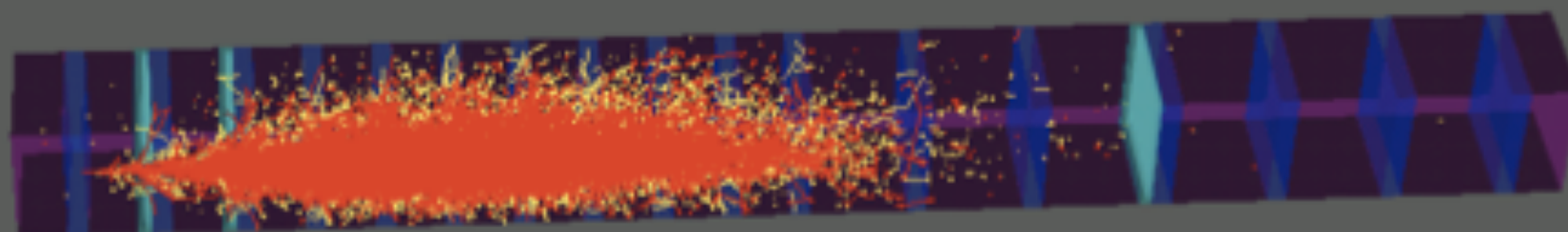
4 SciFi tracking layers



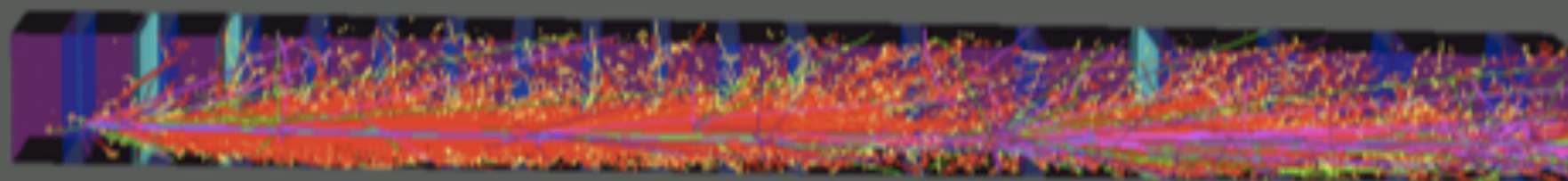
4 Silicon strip tracking layers



Photon vs Neutron



400 GeV photon

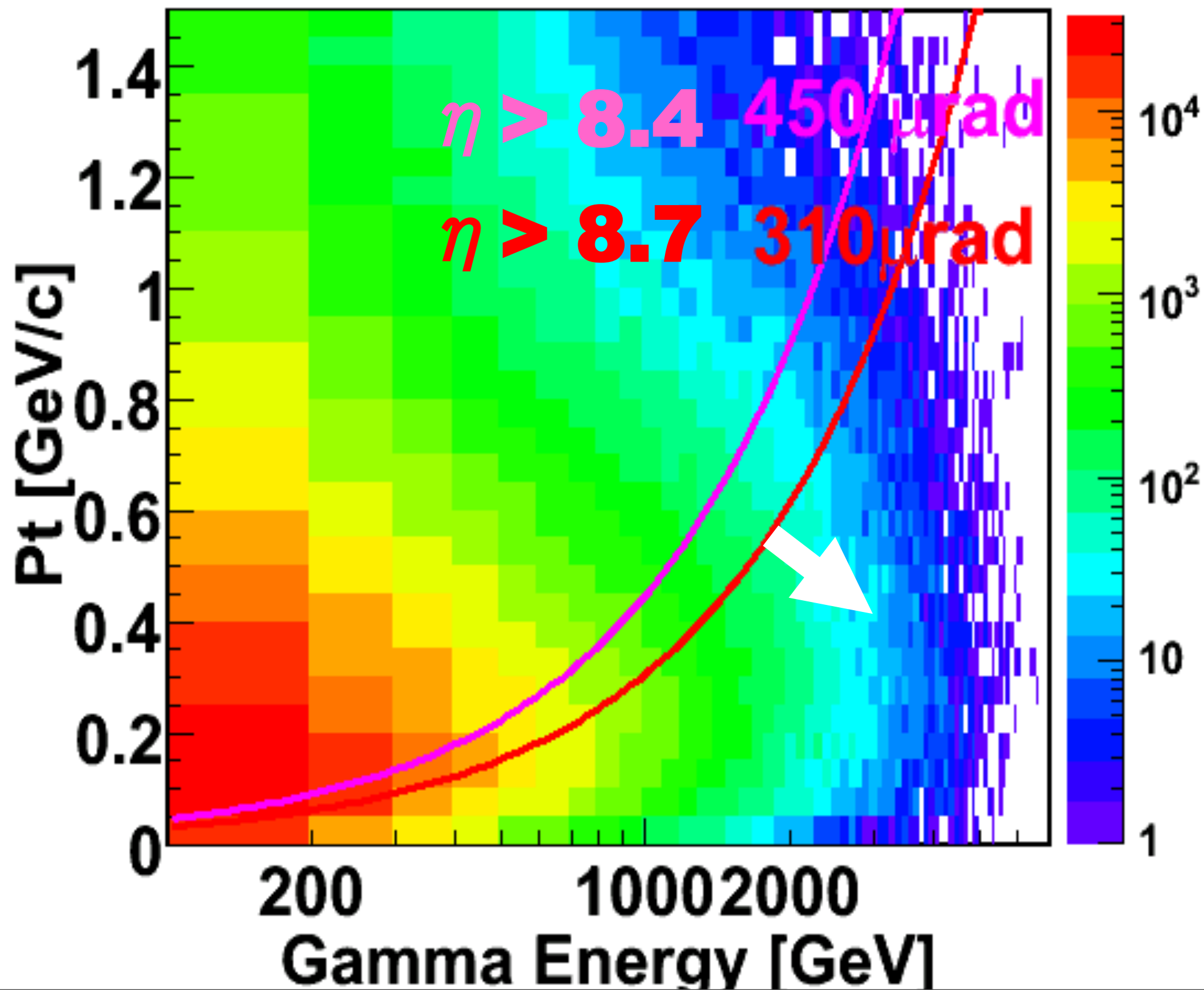


1TeV Neutron

X_F

0.1

1.0



2009-2010 run summary

- $\sqrt{s} = 900 \text{ GeV}$ run (no crossing angle)
 - 06 Dec.- 15 Dec. 2009 (27.7 hrs, 500k collisions)
 - 2.8k/3.7k single showers at Arm1/Arm2
 - 02 May-27 May 2010 (15 hrs. 5.5M collisions)
 - 44k/63k single showers at Arm1/Arm2
- $\sqrt{s} = 7 \text{ TeV}$ run (0 and 100 μrad crossing angle)
 - 30 Mar.- 19 July, 2010 (~150 hrs.)
 - 172M/161M single showers at Arm1/Arm2
 - 345k/676k Pi^0 's at Arm1/Arm2
- Detectors were removed at 20 Jul. 2010

Inclusive photon spectrum

Data

Date : 15 May 2010 17:45-21:23 (Fill Number : 1104)

except runs during the luminosity scan.

Luminosity : $(6.5-6.3) \times 10^{28} \text{cm}^{-2} \text{s}^{-1}$,

DAQ Live Time : 85.7% for Arm1, 67.0% for Arm2

Integrated Luminosity : 0.68 nb^{-1} for Arm1, 0.53 nb^{-1} for Arm2

Number of triggers : 2,916,496 events for Arm1

3,072,691 events for Arm2

Detectors in nominal positions and Normal Gain

Monte Carlo

QGSJET II-03, DPMJET 3.04, SYBILL 2.1, EPOS 1.99 and

PYTHIA8.145: about 10^7 pp inelastic collisions each

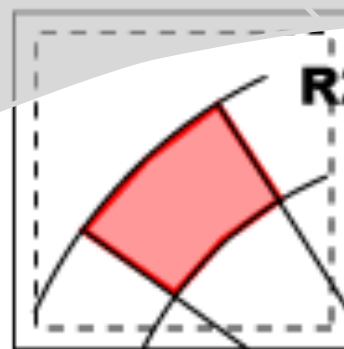
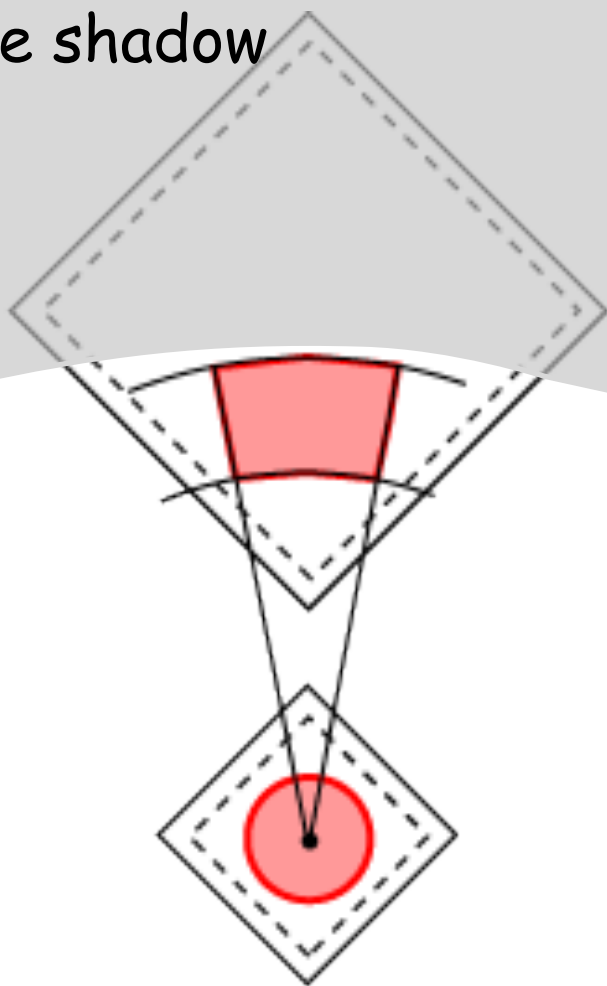
Eta region for photon spectrum

Common region for Arm I and Arm2

Arm I

Arm2

beam pipe shadow



theta

R2-2

R2-1

$$R2-2 = 42\text{mm}$$

$$R2-1 = 35\text{mm}$$

$$\theta = 20^\circ$$

$$8.81 < \eta < 8.99$$

R1

$$R1 = 5\text{mm}$$

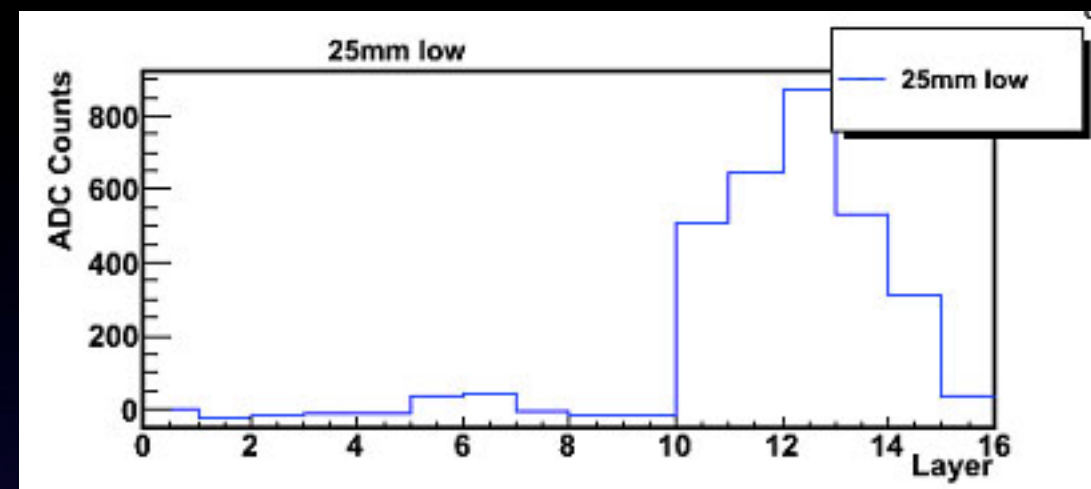
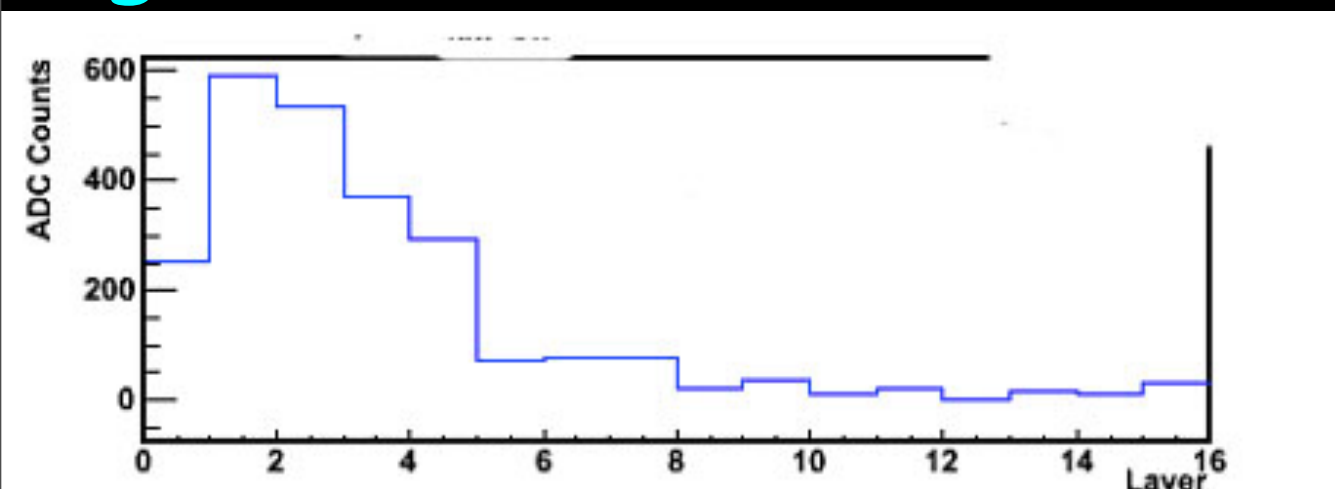
$$\eta > 10.94$$

Particle ID

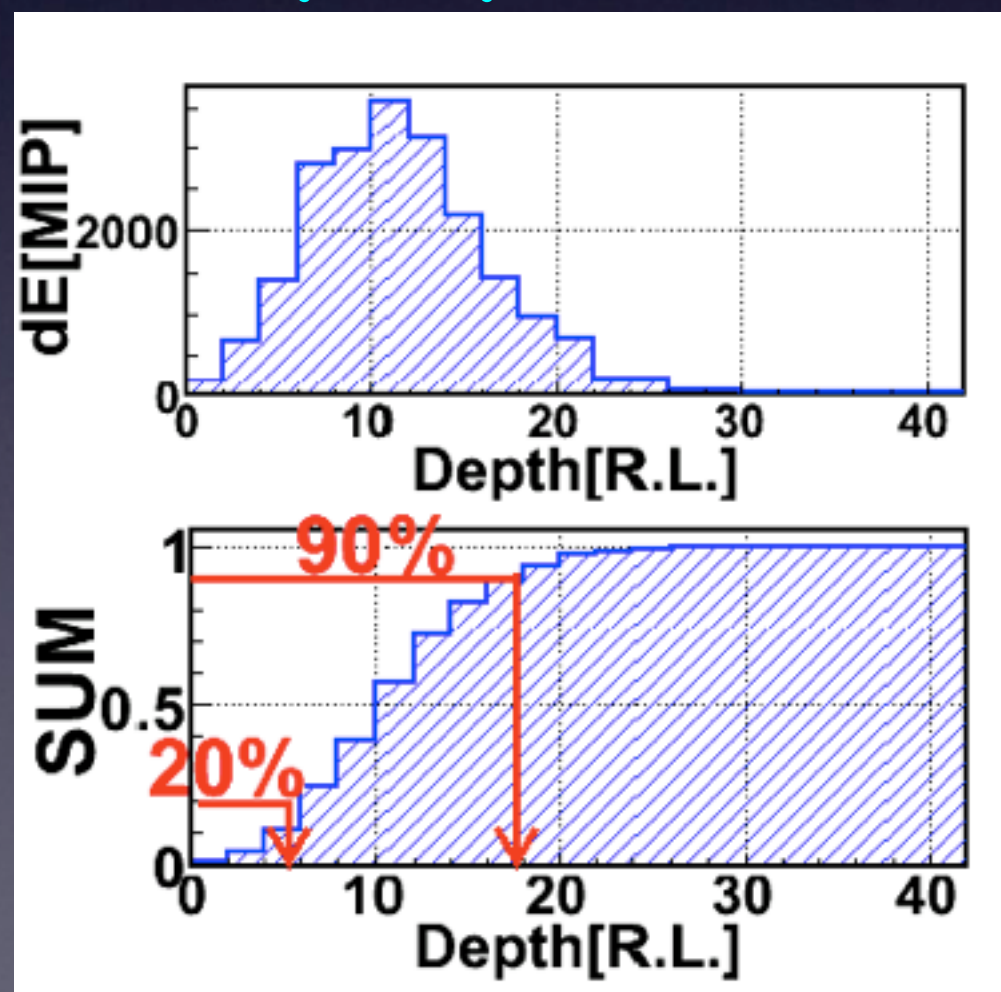
gamma like

shower transition curve

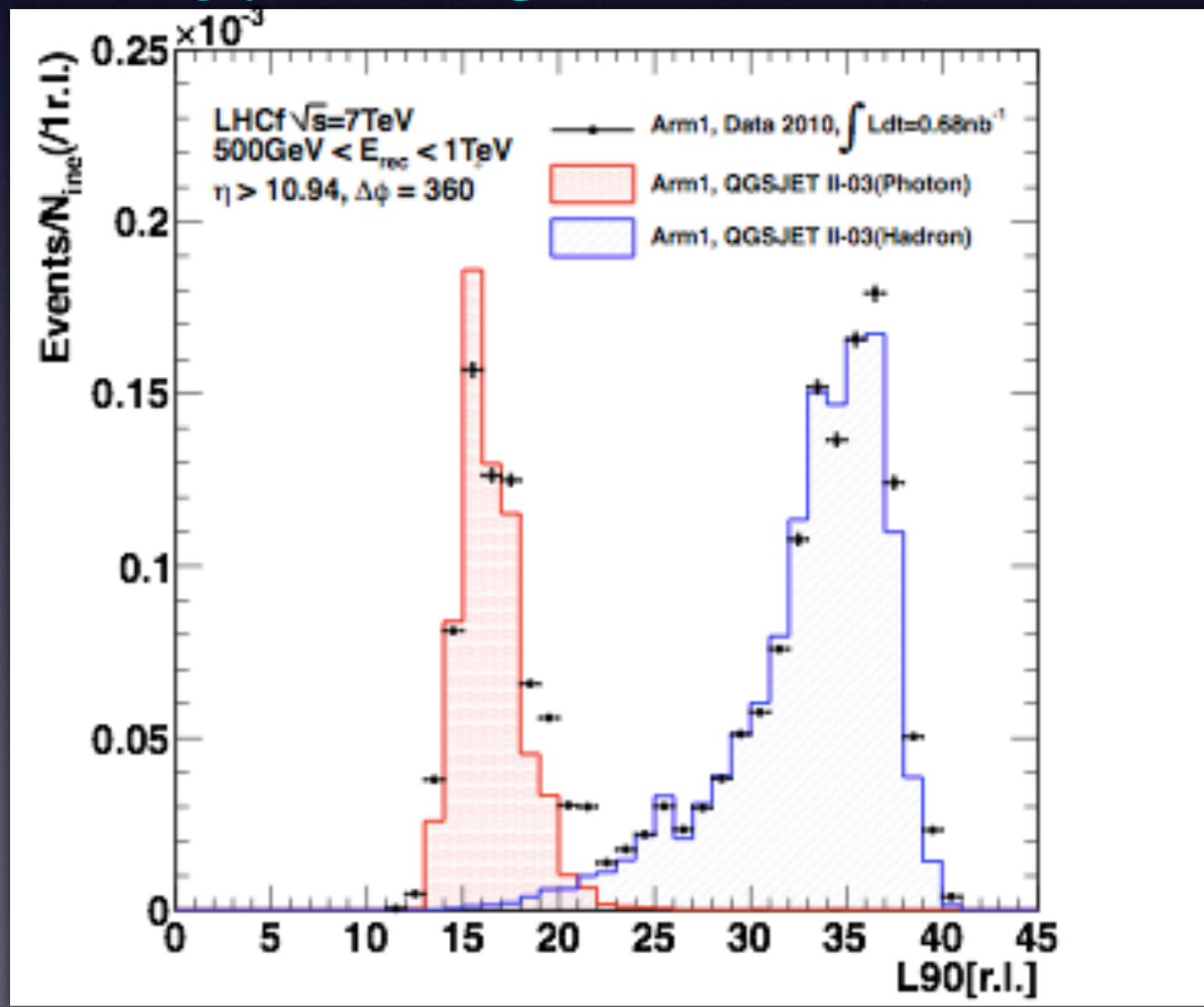
hadron like



Def. of L90%



L90%: Data vs MC

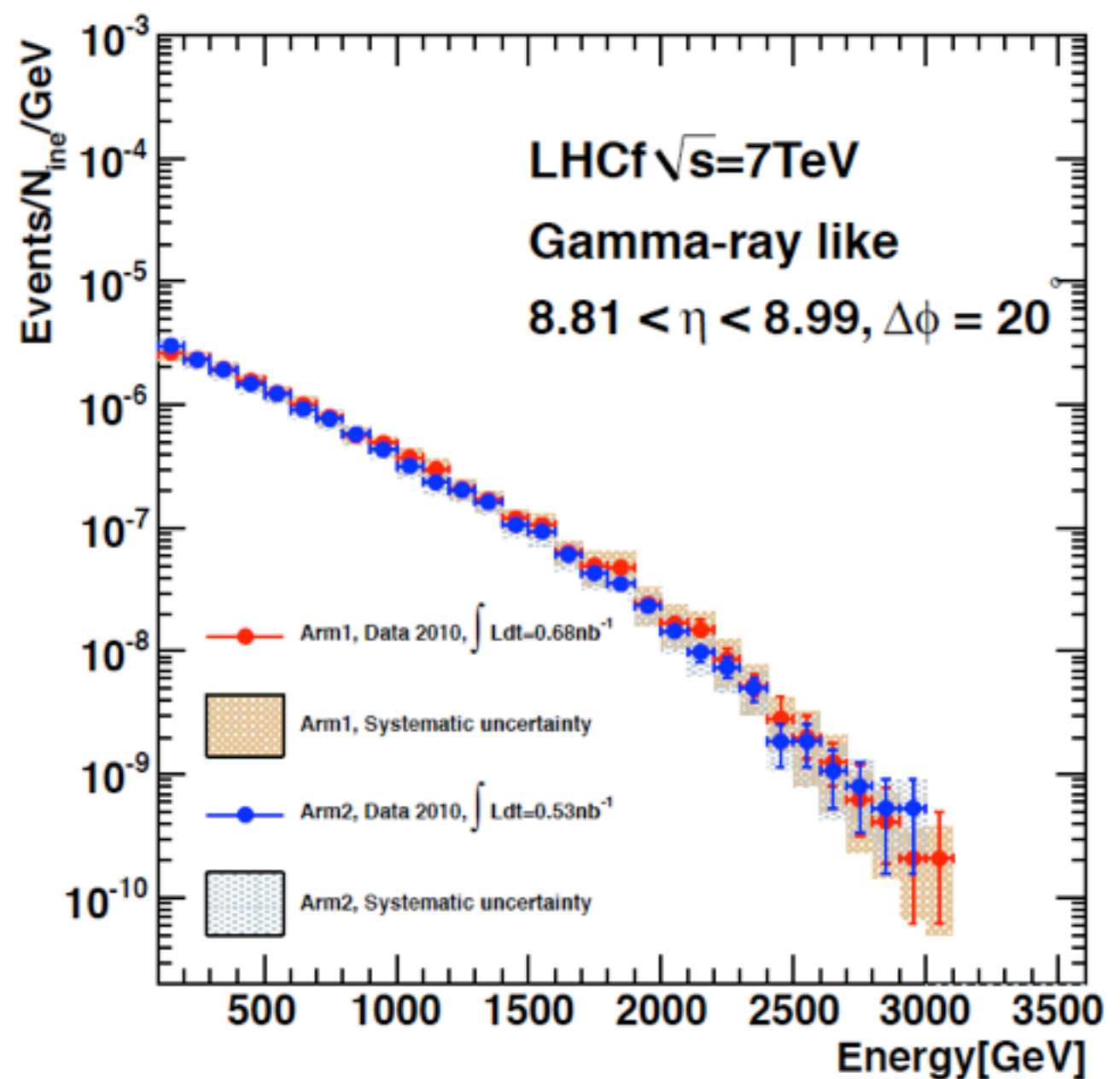
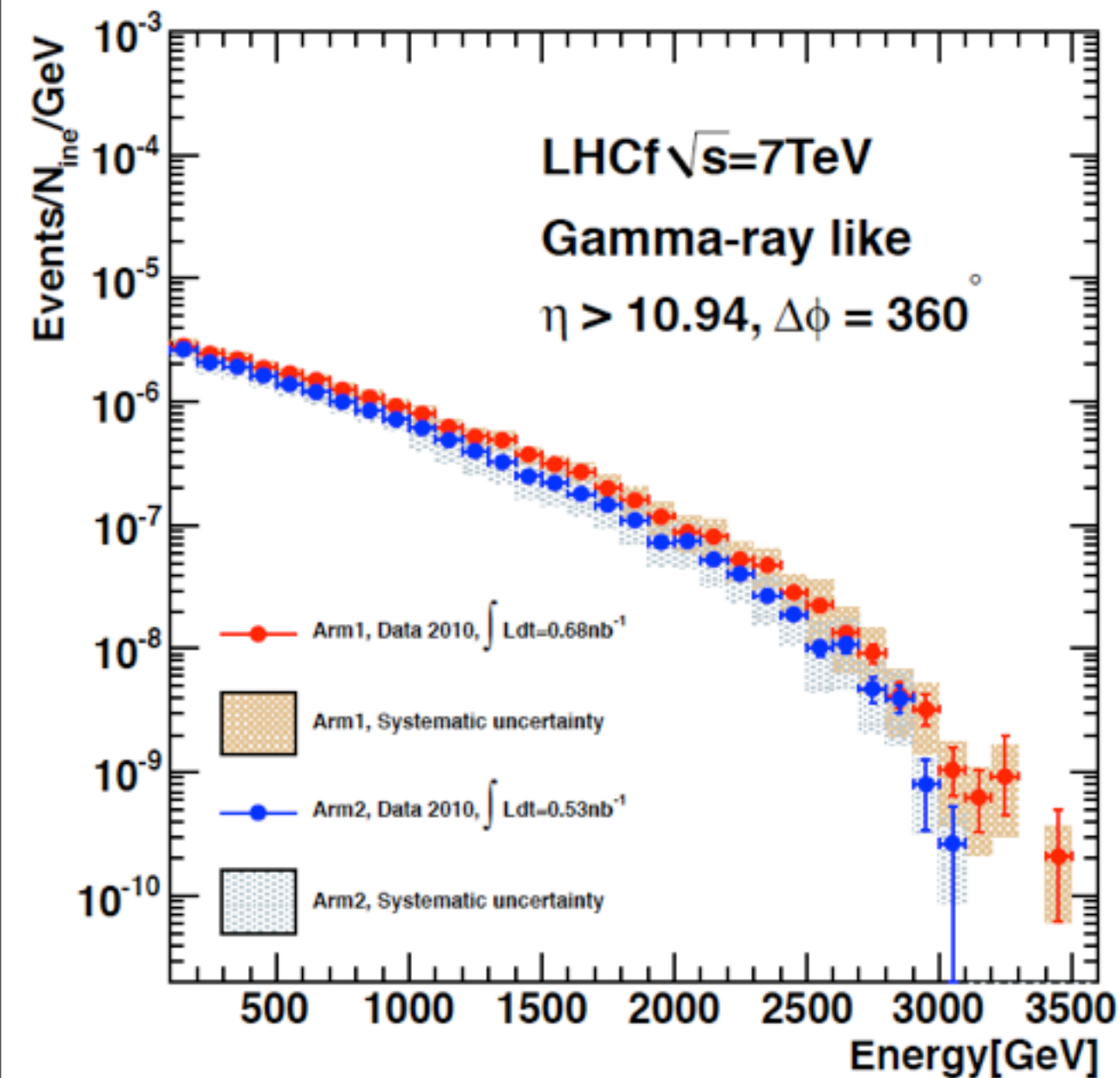


- Changing PID criteria:
 - We get the same energy spectrum

Inclusive photon spectrum: **result**

Small towers

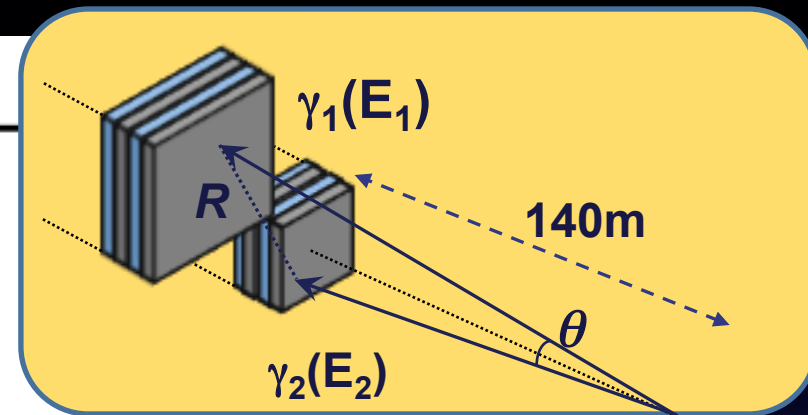
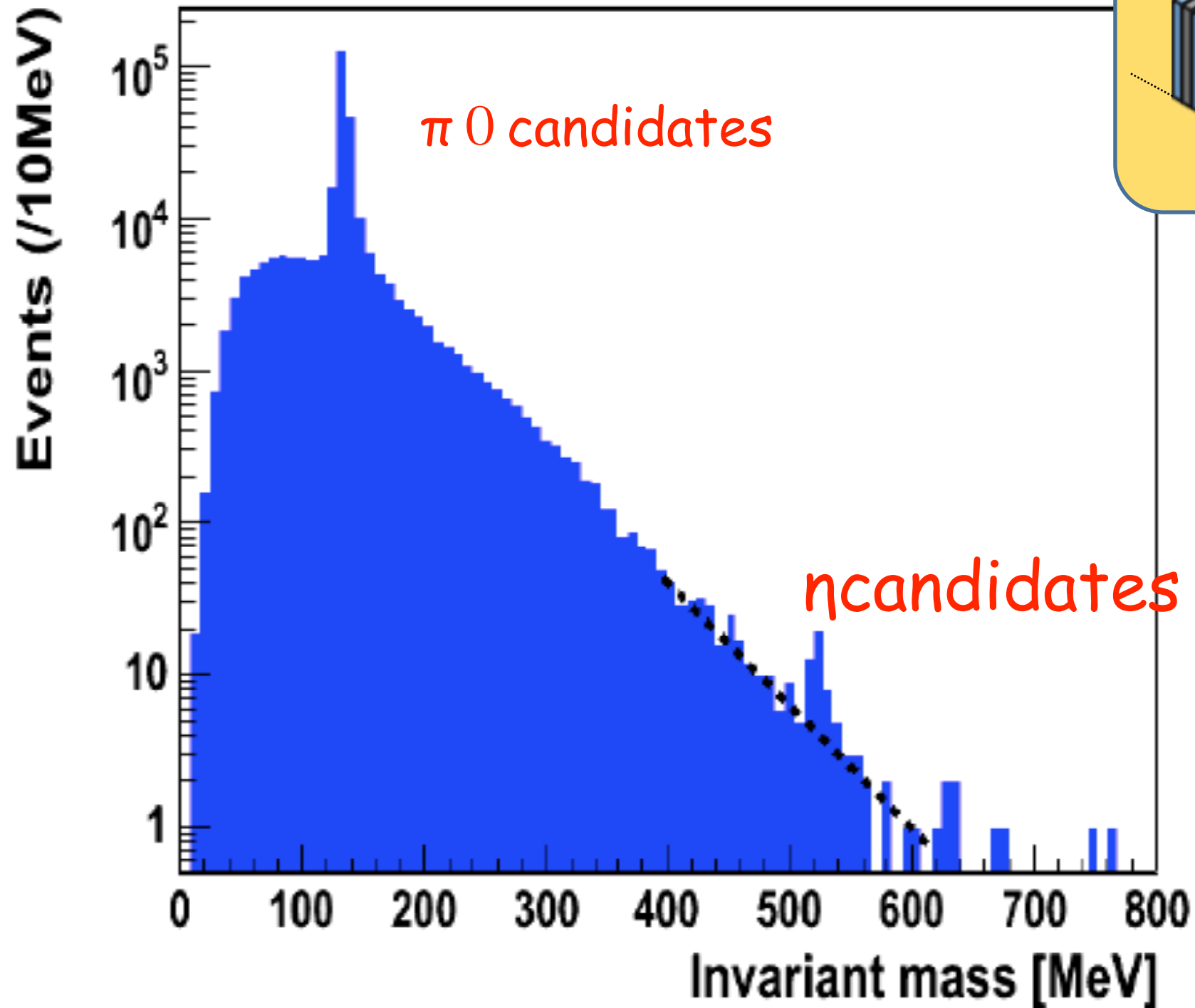
Large towers



Error sources

- * Leak-out/Leak-in
- * Light collec. scin pos. dependeht
- * Incident position estimation
- * Particle ID
- * Multi-hit
- * Pile-up
- * Beam-gas collision
- * B.G from beam pipe
- * Beam center (crossing angle)
- * Inelastic cross-section 71.5 mb

Energy calibration by π^0

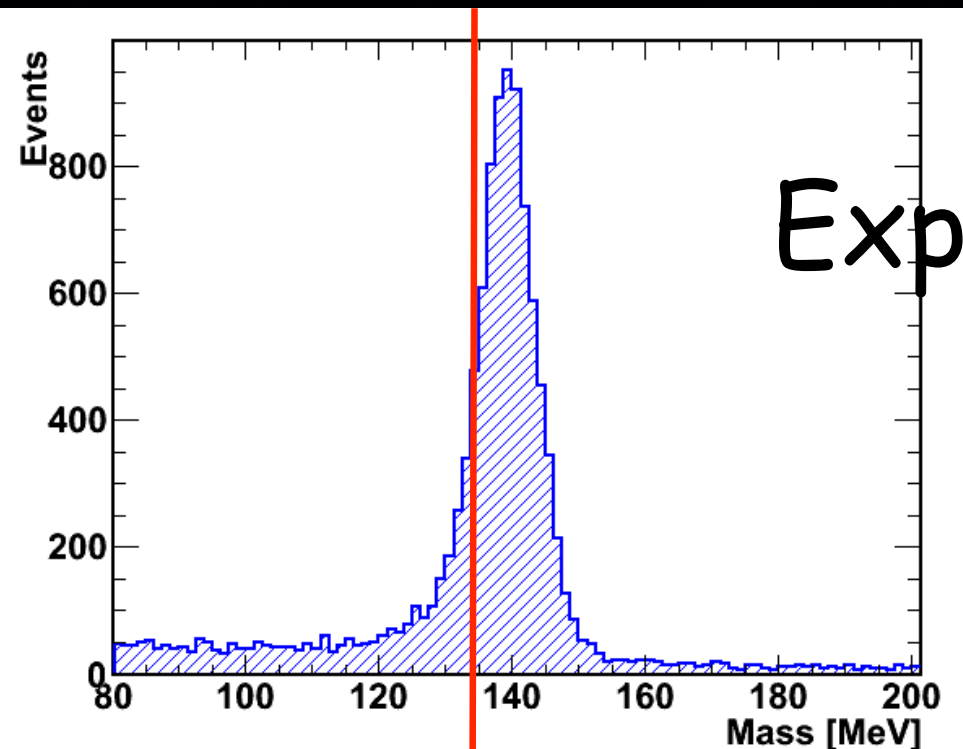


π^0 mass = 135 (MeV)

Arm2

Exp.

M.C



Arm1	145.8 (+7.8%)	135.2
Arm2	140.0 (+3.5%)	135.0

$\pm 0.1 \text{ MeV}$

$\pm 0.2 \text{ MeV}$

Systematic error:

charge to E conversion factor

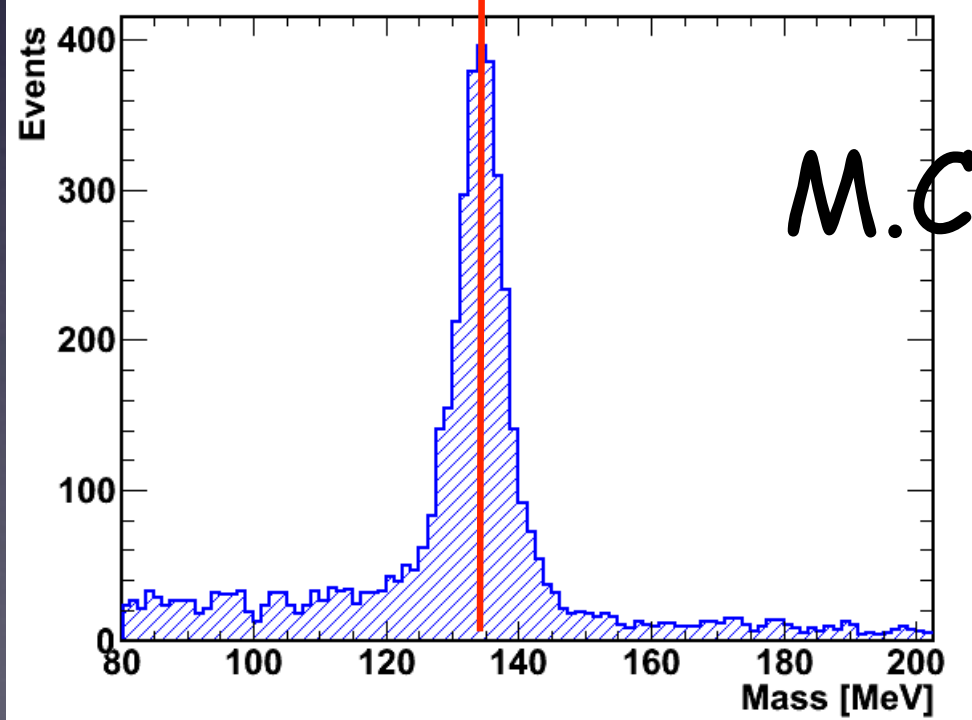
+ non-uniform light yield+.. (3.5 %)

+ leakage-in (1%) + pos. error (2%)

→ 4.2%

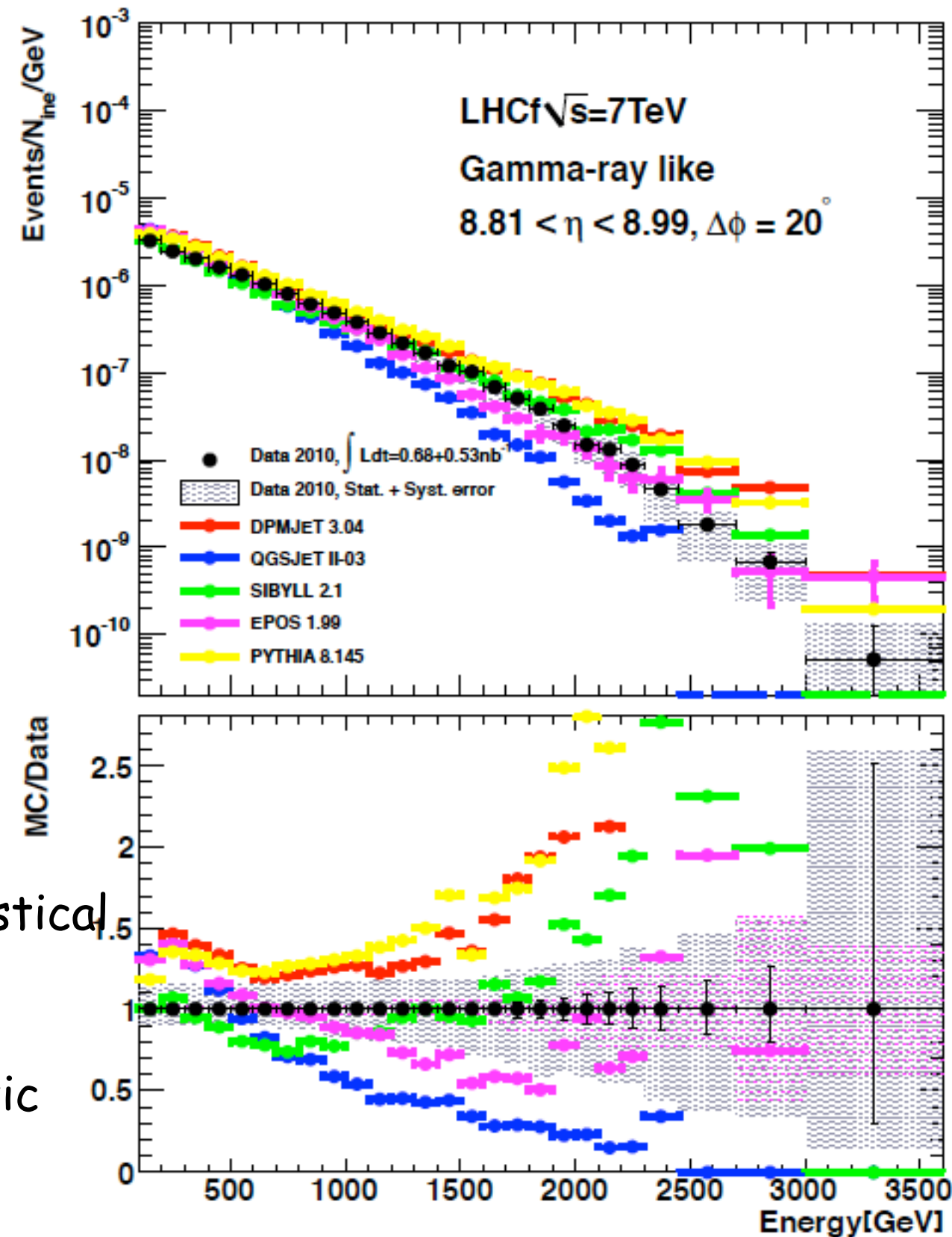
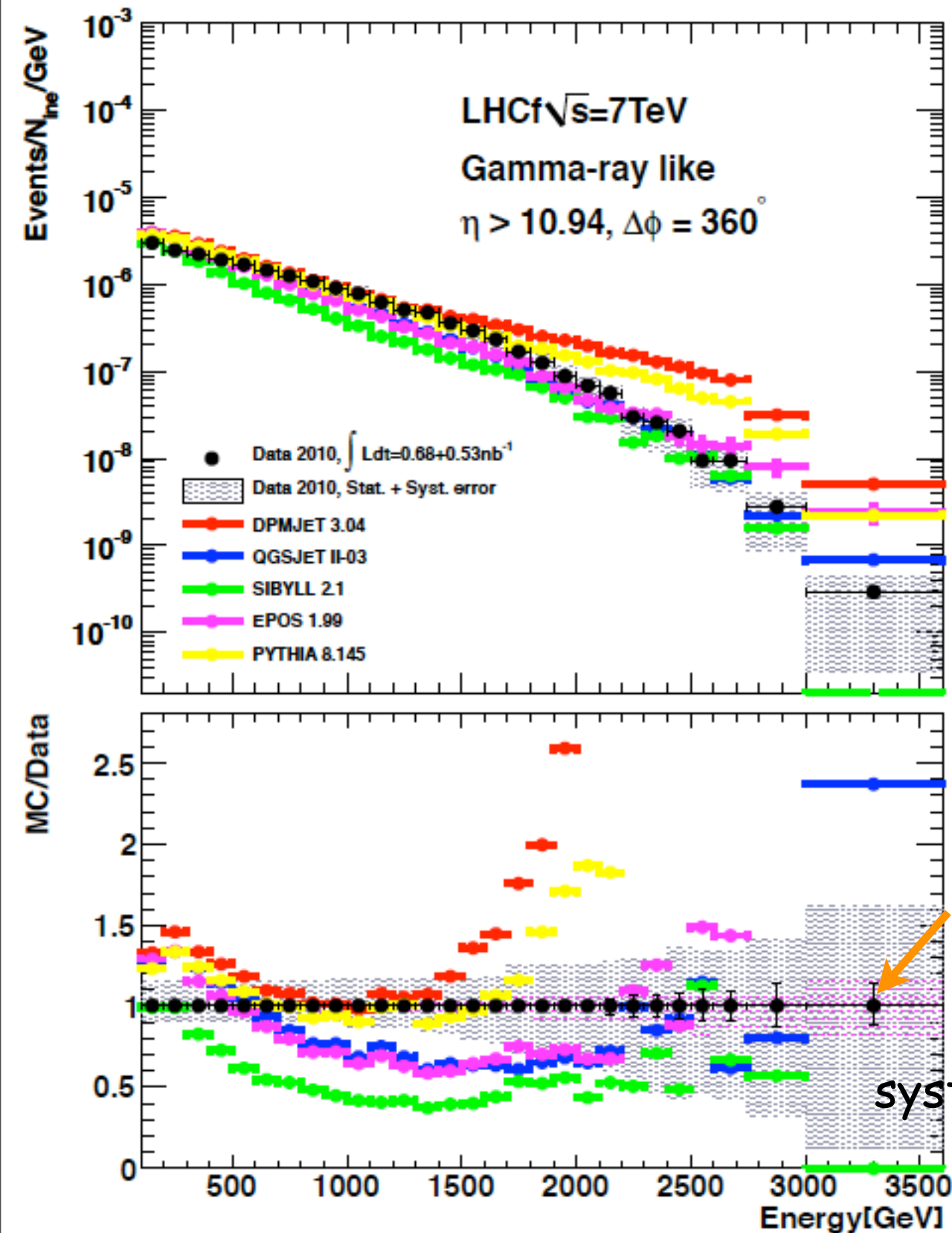
Arm2 is OK. But not Arm1

For the moment, we put the errors into energy scale



Comparison with M.C models (Arm1, 2 combined data)

DPMJET3.04 SIBYLL 2.1 EPOS 1.99 PYTHIA 8.145 QGSJET II-03



Summary



- Photon spectrum in the very forward region was obtained at $\sqrt{s} = 7$ TeV
- Non of the *M.C* models lies within the systematic errors of our data in the entire region
- But the departure is NOT very much surprising/dramatic level
 - For all models, it's not monotonic wrt energy
 - Theoretical implication ? How to update MC ?
- Better model selection / estimate effect on AS.
 - Hadron (neutron) analysis. Inelastic cross-section...