

A NEW MODEL FOR THE EXTRAGALACTIC GAMMA RAY BACKGROUND

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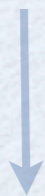
with

Ruben Salvaterra & Francesco Haardt

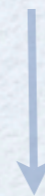
WHAT THE EGB IS

Take all sky detected by Fermi LAT and subtract the Galactic emission

EGB = resolved sources + all the rest

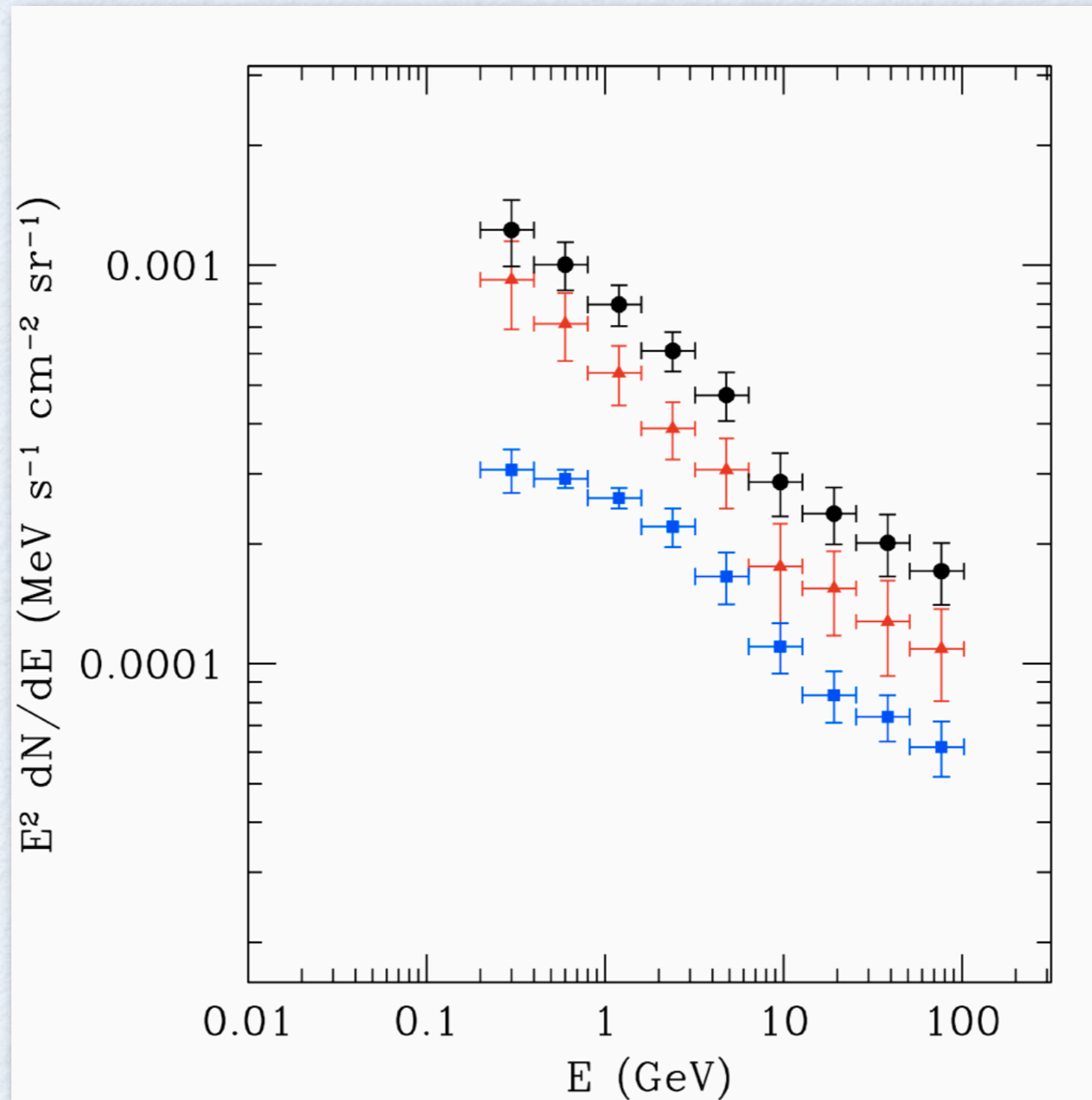


blazars;
star-burst galaxies



unresolved blazars;
star-forming galaxies;
star-burst galaxies;
AGNs;
radio galaxies;
Dark Matter;
Intergalactic shocks
...

THE FERMI EGB



EGB MODEL: INGREDIENTS

- i) blazar (FSRQ + BL-Lac) component;
- ii) star-forming galaxy component.

EGB MODEL: IDEA

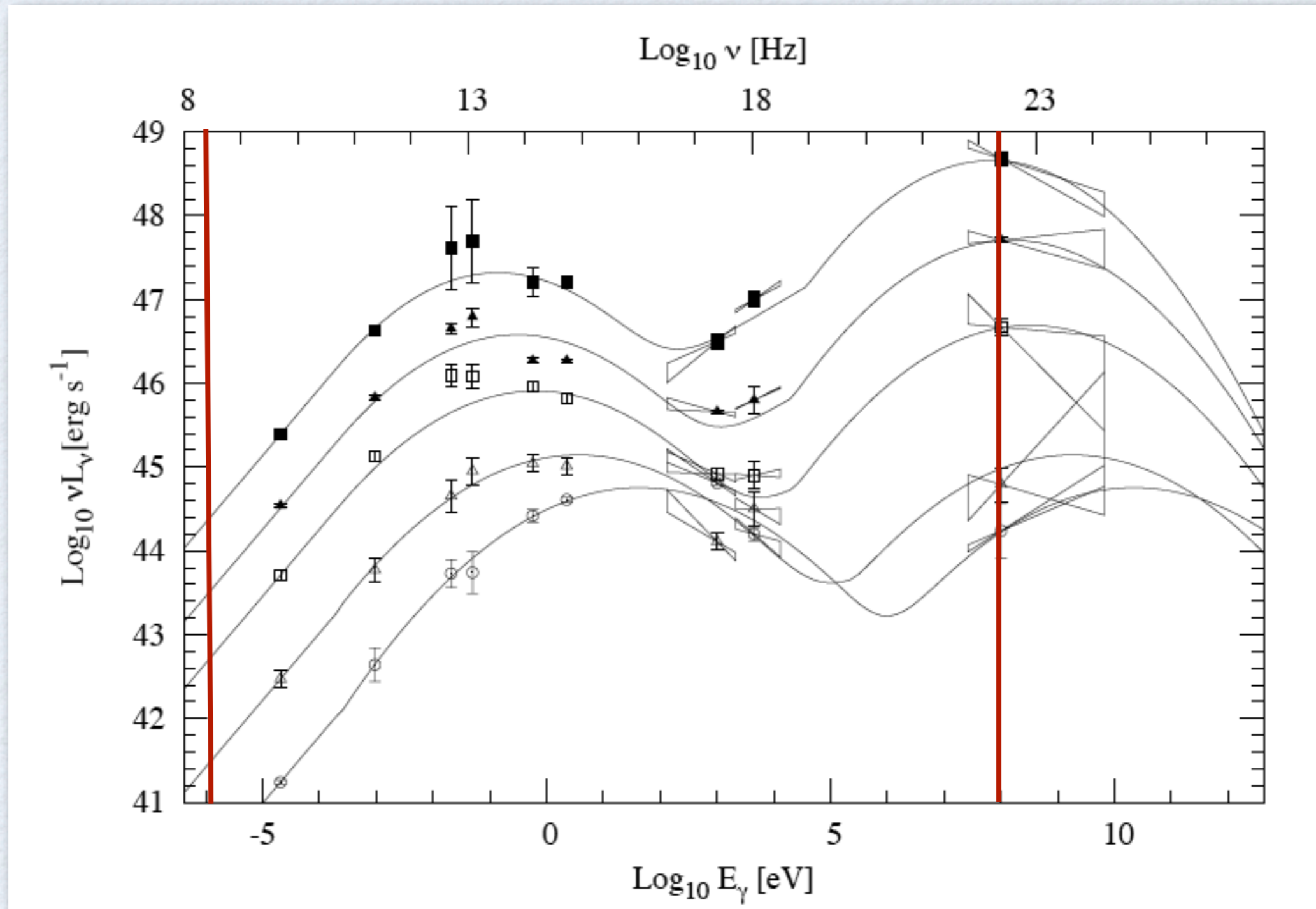
We simultaneously fit the Log N- Log S of Fermi sources and the total EGB with a two component model with two free parameters.

EGB MODEL: BLAZAR COMPONENT

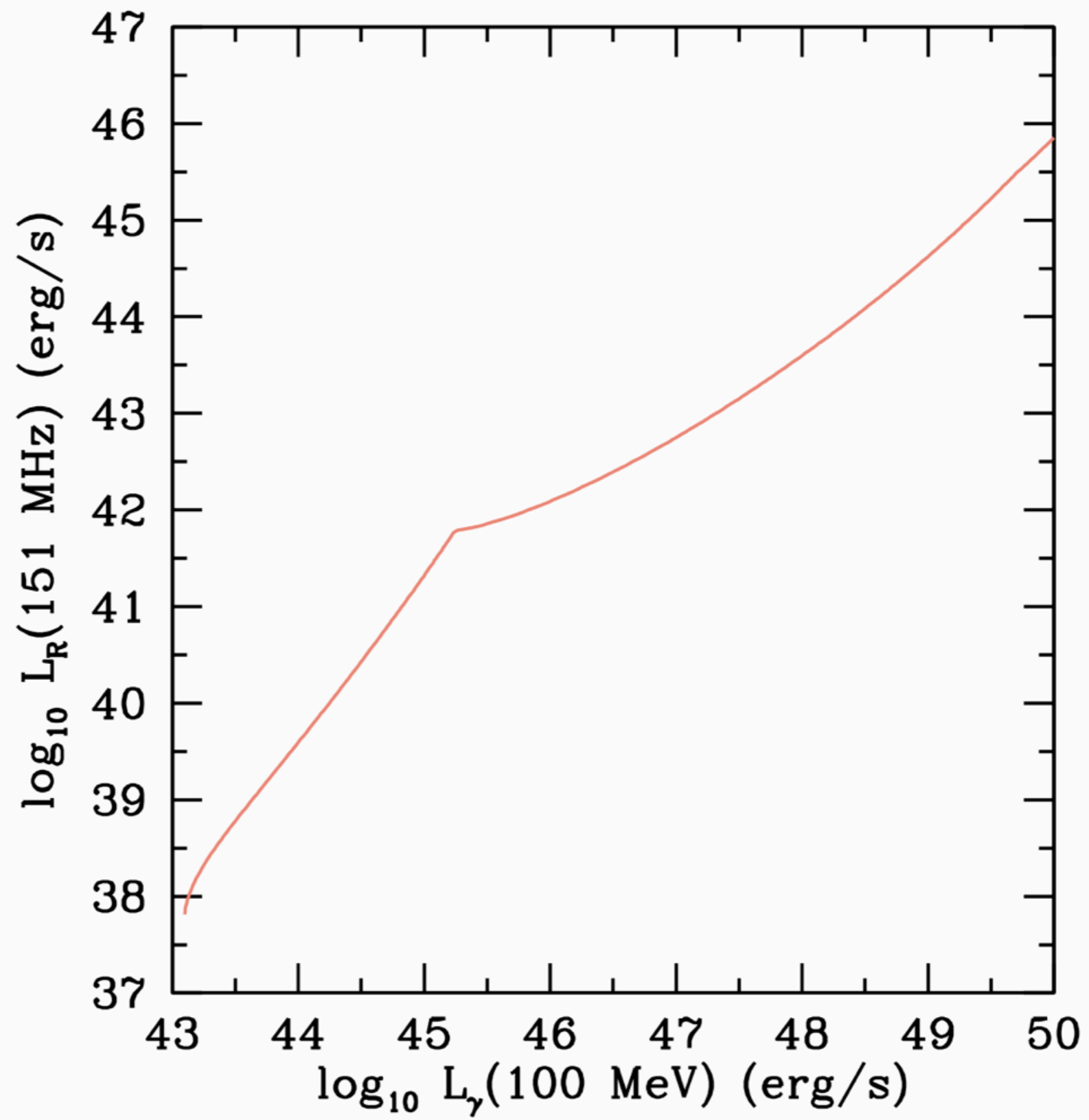
- Two basic elements:
- Radio LF (@ 151 MHz; FR I and FR II) (Willott et al. 2001);
 - Blazar Flux (using the Blazar SED) (Fossati et al. 1998).

$$I_{\text{blaz}}(E_0) = \frac{1}{4\pi} \int_0^\infty dz \frac{dV}{dz} \int_{\log L_\gamma^{\min}}^{\log L_\gamma^{\max}} d \log L_\gamma \frac{d\Phi_\gamma(L_\gamma, z)}{d \log L_\gamma} \frac{dn(L_\gamma, z)}{dE} e^{-\tau_{\gamma\gamma}(E_0, z)}$$

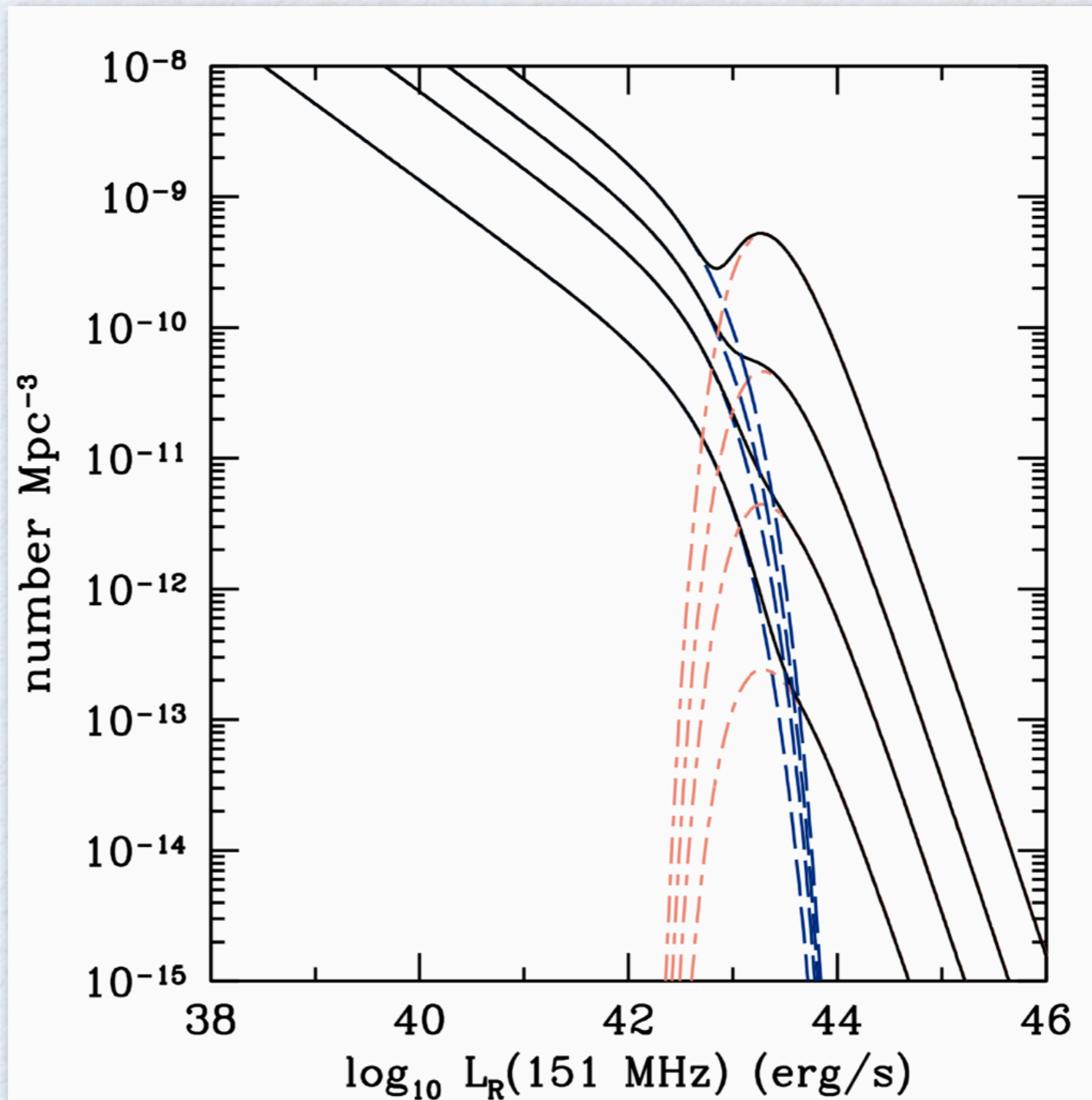
THE BLAZAR SEQUENCE



See Fossati et al 1998; Inoue & Totani 2010

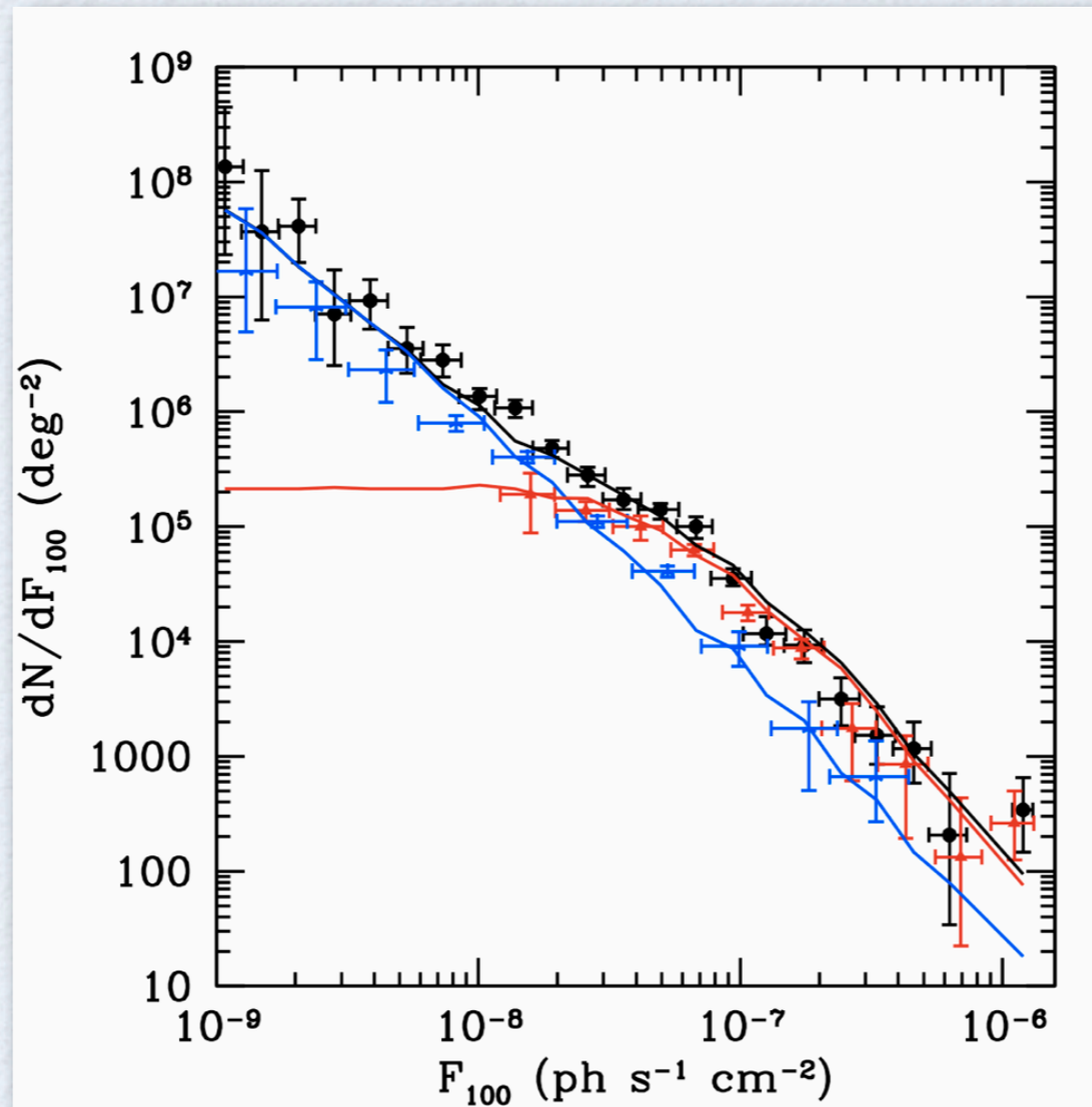


THE BLAZAR LF



(Willott et al 2001)

BLAZAR DIFFERENTIAL logN-logS



$$N(> F_{\text{ph}}) = \frac{1}{4\pi} \int_0^\infty dz \frac{dV}{dz} \int_{\log L_\gamma^{\min}}^{\log L_\gamma^{\max}} d \log L_\gamma \frac{d\Phi_\gamma(L_\gamma, z)}{d \log L_\gamma}$$

$$\frac{\Phi_\gamma(L_\gamma, z)}{d \log L_\gamma} = \kappa \frac{\Phi_{151\text{MHz}}(L_{151\text{MHz}}, z)}{d \log L_{151\text{MHz}}}$$

$$k = 3.93 \pm 0.01 \times 10^{-4}$$

EGB MODEL: STAR-FORMING GALAXY COMPONENT

Basic idea:



$$I_{\text{gal}}(E_0) = \frac{1}{4\pi} \int_0^\infty dz \frac{dl}{dz} \frac{(1 + \mathcal{R})}{m_{\text{H}} \xi(\text{H}_2)} \langle q_{\text{H}}(E) \rangle \dot{\rho}_{\text{SFR}}(z) e^{-\tau_{\gamma\gamma}(E_0, z)}$$

Free parameter:
SF efficiency of molecular
Hydrogen

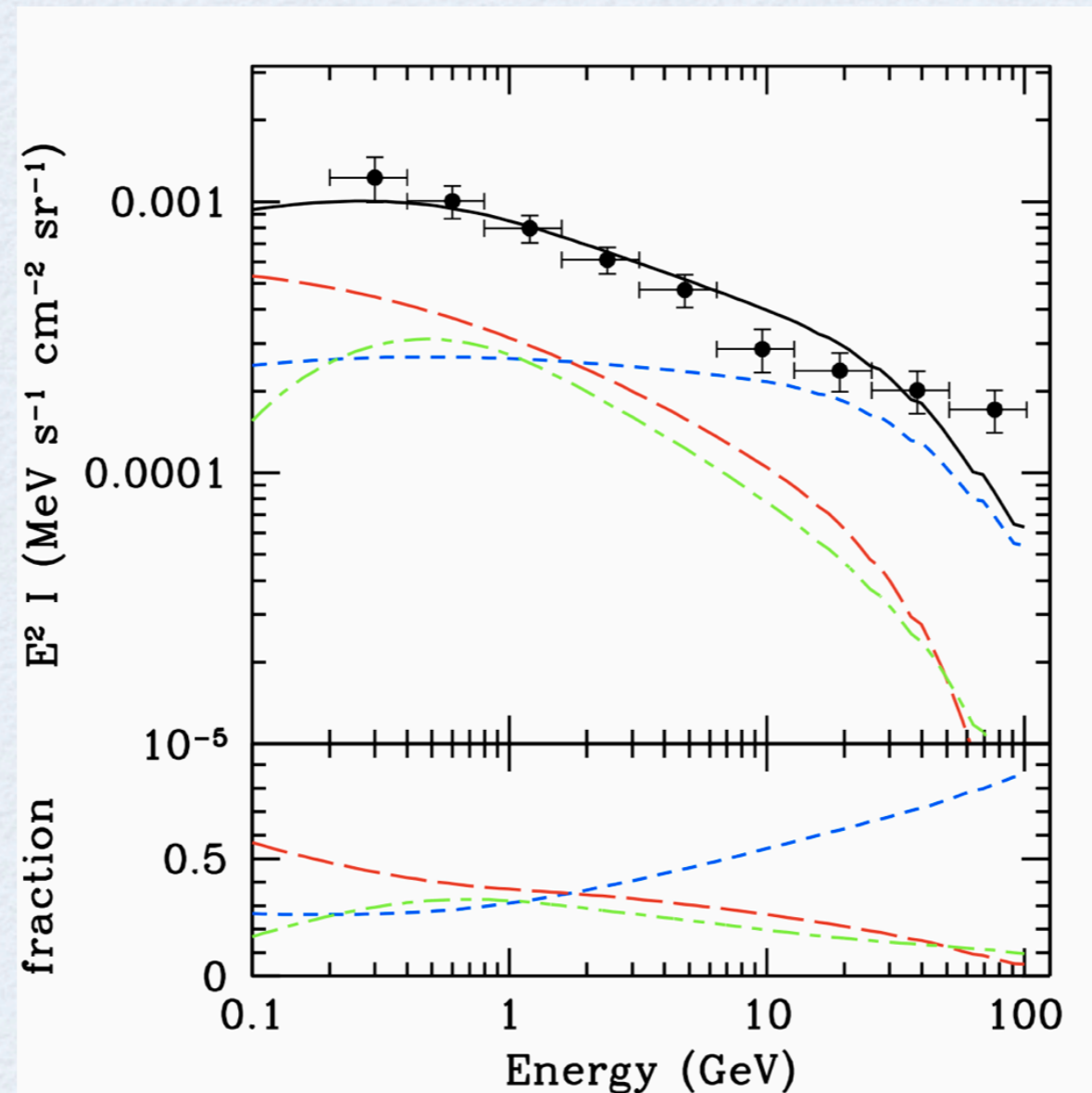
Pionic emissivity

Cosmic SFR

$$\dot{\rho}_{\text{SFR}} \sim \xi(\text{H}_2) \rho_{\text{H}_2}$$

(see Stecker & Venters 2010 for details)

EGB MODEL: RESULTS



$$k = 3.93 \pm 0.01 \times 10^{-4}$$

$$\xi = 4.07 \pm 0.4 \times 10^{-10} \text{ yr}^{-1}$$

$$\chi^2/\text{d.o.f} = 1.15 \text{ where d.o.f} = 29$$

$$\xi \sim (5.25 \pm 2.5) \times 10^{-10} \text{ yr}^{-1}$$

CONCLUSIONS

- We proposed a new model for EGB with two components: star forming galaxies and blazars (FSRQs and BL-Lacs);
- We explain all the EGB Fermi points but the last at 75 GeV. This could be due to a galactic DM or to a blazar LF not precise at low redshift.

Thank you!

