Status and Future Plans of the Cryogenic Dark Matter Search Experiment

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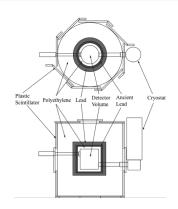
S. Arrenberg, T. Bruch, L. Baudis, M. Tarka



The Cryogenic Dark Matter Search Experiment

- Search for Dark Matter in the form of Weakly Interacting Massive Particles (WIMPs)
- Located in Soudan Mine, Minnesota, USA (2090 mwe)
- ullet Ge & Si crystals cooled to pprox 40 mK
- Passive Pb & polyethylene shielding



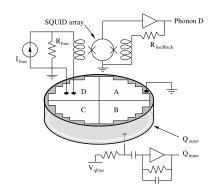


- Muon scintillator veto
- Collect phonon & ionisation signals for event-by-event discrimination

CDMS-II Detector: ZIP

- Z-sensitive Ionisation & Phonon detector (ZIP)
- 3'' diameter \times 1 cm thick
- Mass:
 - 19 × 250 g (Ge)
 - 11 × 100 g (Si)





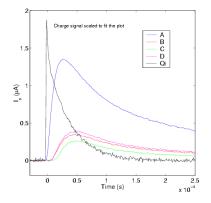
- 2 charge electrodes on one surface:
 - inner disc & outer ring
- 4 phonon sensors on opposite surface:
 - each covers one quadrant



Ionisation & Phonon Signals

Simultaneous measurement of charge-carriers (ionisation) and phonons generated by an event

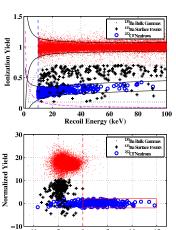
- Ionisation signal:
 - Fast: 1 μ s rise- & 40 μ s fall-times
 - Provides event time information
- Phonon signal:
 - Pulse shape depends on event position
- Recoil energy from ionisation & phonon signal amplitudes
- Event position reconstruction from timing and amplitude of phonon signals



Electron-Nuclear Recoil Discrimination

Nuclear recoils generate less ionisation than electron recoils of the same deposited energy

Therefore, ionisation-phonon recoil energy ratio is a discrimination parameter \Longrightarrow Ionisation Yield



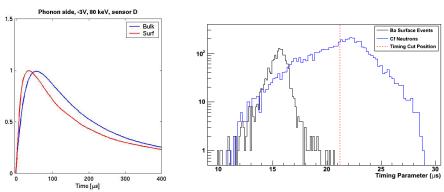
Normalized Timing Parameter

- γ -rays from ¹³³Ba source define electron recoil band
- neutrons from ²⁵²Cf source define nuclear recoil band
- Bands are well separated down to 10 keV
- Surface electron-recoil events:
 - Incomplete charge collection
 - Reduced yield ⇒ background



Surface Event Rejection

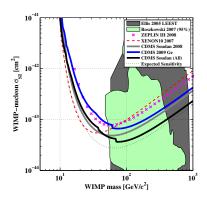
Shape of phonon pulses indicates interaction depth



Surface events have faster rise time than bulk events, enabling definition of a timing cut

Results from Final Exposure of CDMS-II

- 612 kg-days exposure
- 2 candidate events:
 - **1** $E_R = 12.3 \text{ keV}$, Tower 1, ZIP 5, 02:48 GMT 28 Oct 07
 - ② $E_R = 15.5 \text{ keV}$, Tower 3, ZIP 4, 19:41 GMT 5 Aug 07
- Expected irredicible backgrounds:
 - 0.8 ± 0.1 (stat) ± 0.2 (sys) misclassified surface electron recoils
 - ullet pprox 0.1 background neutrons from cosmic-rays & radioactivity
- 23% probability of observing 2 or more background events over this exposure time

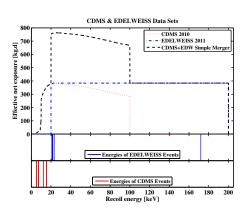


Z. Ahmed et al., Science 327, 1619 (2010)

- Cannot be interpreted as significant evidence for WIMP interactions, nor can either event be rejected as a WIMP
- In process of reanalysing results to reduce background estimates

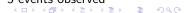
CDMS-EDELWEISS Combined Limit

 EDELWEISS also uses cryogenic Ge detectors to search for WIMPs at Modane Underground Laboratory (4800 mwe)



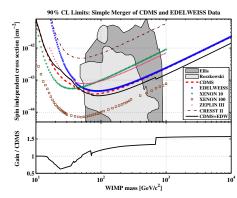
Z. Ahmed *et al.*, Phys. Rev. D **84**, 011102 (2011)

- Same target nucleus allows results to be combined without introducing additional model dependence
- Technology not identical: 400 g crystals & patterned electrodes
- CDMS data:
 - Threshold pushed down to 5 keV
 - 4 events observed
- EDELWEISS data:
 - Threshold set at 20 keV
 - 5 events observed



CDMS-EDELWEISS Combined Limit

- Sum exposure-weighted efficiencies
- Treat all events on equal footing regardless of experiment of origin
- Biggest improvement at high WIMP masses due to large eventless window between 23.2 & 172 keV
 - ullet Up to a factor of ~ 1.6 for $M_{\gamma} > \sim 50 \text{ GeV}$
- Low energy events in EDELWEISS lead to degredation in limit for low mass WIMPs

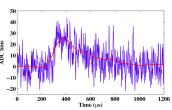


Z. Ahmed et al., Phys. Rev. D 84, 011102 (2011)

 Alternative procedures for combining data are explored, and could provide stronger constraints at certain masses

CDMS-II Low Energy Analysis

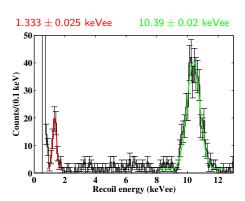
- \bullet Recent results from DAMA/LIBRA & CoGENT have been interpreted in terms of elastic scatters from a WIMP with mass ~ 10 GeV and cross-section $\sim 10^{-40}~\rm cm^2$
- \bullet Previous CDMS results not sensitive to such WIMPs since thresholds were ~ 10 keV to maintain sufficient rejection of electron recoils
- Can lower threshold significantly at cost of higher backgrounds
- Repeat analysis with lower 2 keV threshold:
 - Use 8 Ge detectors with lowest thresholds
 - Results driven by best detector (Tower 1, ZIP 5)



2 keVnr phonon pulse (Tower 1 ZIP 5)

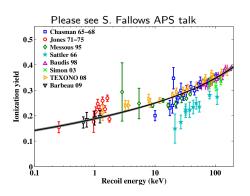
Energy Calibration

- Ionisation energy scale initially calibrated with 356 keV line (¹³³Ba)
- Phonon energy calibrated by normalising phonon-based recoil energy for electron recoils to their mean ionisation energy
- \bullet From 1.3 and 10.4 keV lines, small rescaling (\sim 4%) applied to ensure electron recoil energy scale was not underestimated



Energy Calibration

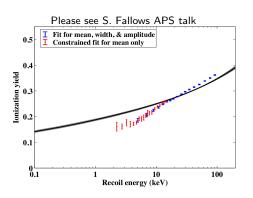
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 Ionisation yield measured for nuclear recoils from ²⁵²Cf source

Energy Calibration

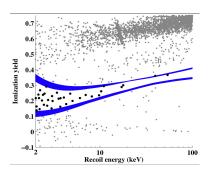
- \bullet lonisation energy scale initially calibrated with 356 keV line (133 Ba)
- Phonon energy calibrated by normalising phonon-based recoil energy for electron recoils to their mean ionisation energy
- \bullet From 1.3 and 10.4 keV lines, small rescaling (\sim 4%) applied to ensure electron recoil energy scale was not underestimated



- Ionisation yield measured for nuclear recoils from ²⁵²Cf source
 - Measured to \sim 4 keV, power-law extrapolation to lower energies
 - Conservative when compared with literature

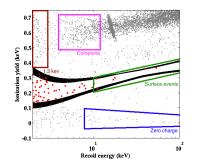
Nuclear Recoil Definition

- ullet lonisation energy within $(+1.25, -0.5)\sigma$ of mean from $^{252}{
 m Cf}$ calibration
- Maximise sensitivity while limiting leakage from electron recoils & zero-charge events



Nuclear Recoil Definition

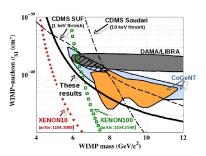
- ullet lonisation energy within $(+1.25, -0.5)\sigma$ of mean from $^{252}{
 m Cf}$ calibration
- Maximise sensitivity while limiting leakage from electron recoils & zero-charge events
- Events in nuclear recoil band can be identified as possible background
- Zero-charge events have ionisation signals consistent with noise
 - Charge carriers collected on cylindrical wall
 - ullet At $E_R < 10$ keV, phonon-based reconstruction unreliable



- Signal-to-noise too low at E_R < 5 keV to reject surface events with phonon timing
- At E_R < 5 keV ionisation-based discrimination breaks down
- 1.3 keV line leaks above 2 keV threshold as recoil energy estimate assumes ionisation signal is consistent with nuclear recoil

90% Confidence Upper Limits

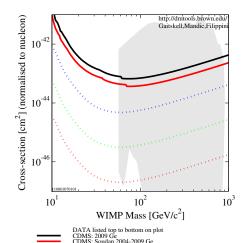
- Can extrapolate observed spectra for each potential background to lower energies to estimate leakage, but systematic errors are potentially large and difficult to quantify
- Therefore, be conservative and assume all candidates are WIMPs
- No background subtraction
- For spin-independent, elastic scattering, 90% confidence level limits incompatible with DAMA/LIBRA & CoGENT excess



Z. Ahmed *et al.*, Phys. Rev. Lett. **106**, 131302 (2011)
D. S. Akerib *et al.*, Phys. Rev. D **82**, 122004 (2010)

$CDMS-II \Longrightarrow SuperCDMS \Longrightarrow GEODM$

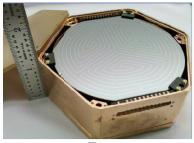
- CDMS-II @ Soudan
 - $3''\phi \times 1 \text{ cm} \rightarrow 250 \text{ g/det}$
 - Target mass 4 kg
 - ullet \sim 2 years ightarrow 612 kg-days
- SuperCDMS Soudan
 - $3'' \phi \times 1'' \rightarrow 600 \text{ g/det}$
 - Target mass 15 kg
 - ullet 2 years $\to \sim$ 8000 kg-days
- SuperCDMS SNOLAB
 - 100 mm $\phi \times$ 33.3 mm \rightarrow 1.37 kg/det
 - Target mass 100 kg
 - ullet 3 years $\to \sim$ 100000 kg-days
- Ge-Observatory for Dark Matter (GEODM) @ SNOLAB/DUSEL
 - $6'' \phi \times 2'' \rightarrow 5 \text{ kg/det}$
 - Target mass 1.5 T
 - ullet 4 years $o\sim$ 1500000 kg-days

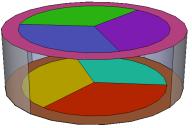


Ellis et. al 2005 LEEST (mu>0, pion Sigma=64 MeV)

SuperCDMS - 15 kg at Soudan

SuperCDMS Soudan Detector: iZIP

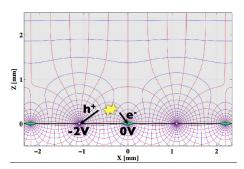


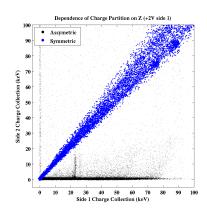


- interleaved Z-sensitive Ionisation & Phonon detector (iZIP)
- New detector design to improve surface event rejection efficiency
- Larger 3" diameter, 1" thick, 600 g crystals
- 8 phonon sensors, 4 on each side, provides new information on interaction depth
- Increased phonon sensor surface coverage improves collection efficiency and signal-to-noise ratio
- Better timing resolution from outer phonon sensor ring, which also breaks degeneracies in position reconstruction

Surface Event Rejection: Ionisation Signals

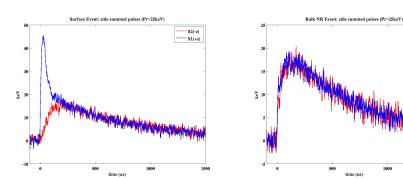
 Two ionisation electrodes on each surface: central disc & outer ring





 Interleaved electrodes, so surface events show up on one detector side only

Surface Event Rejection: Phonon Signals



Two phonon-based discrimination parameters:

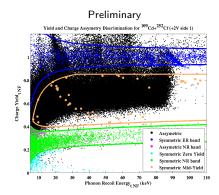
- Pulse shape differences as with ZIPs
- Side asymmetry provides new surface event rejection parameter



S2(-z)

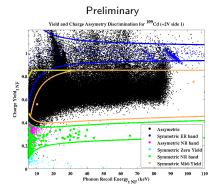
S1(+z)

iZIP Performance



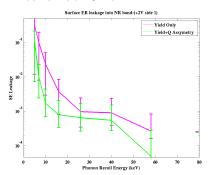
Electron recoils (133 Ba γ -rays) Surface events (109 Cd β -particles) Nuclear recoils (252 Cf neutrons & 7 evt/h background at surface test facility) • Electron & nuclear recoils well-separated above $\approx 6 \text{ keV}$

iZIP Performance

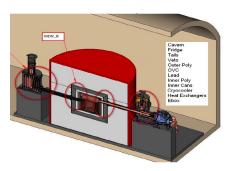


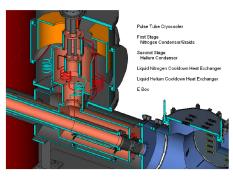
Electron recoils (133 Ba γ -rays) Surface events (109 Cd β -particles) Nuclear recoils (7 evt/h background at surface test facility)

- Electron & nuclear recoils well-separated above $\approx 6 \text{ keV}$
- Yield-only discrimination begins to fail at 10 keV



SuperCDMS SNOLAB





- Target mass 100 kg
- SNOLAB provides deeper facility (> 4000 mwe) than Soudan
- New fridge and shield design work in progress at Fermilab
- Deploy crystals of iZIP detector technology

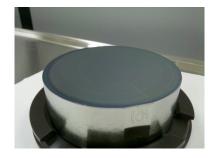
Scaling to 100-1000 kg

- iZIP technology appears to meet requirements for larger target masses
- Scaling difficult: \sim 170 $3''\phi\times 1''$ crystals for 100 kg target mass
- A number of factors make this expensive:
 - Increased manpower: fabrication and testing are labour intensive
 - Increased heat load: additional wiring to room temperature and Field Effect Transistors (FETs) on 4 K stage (5 mW/FET)
 - Increased cold hardware & warm electronics
- Can alleviate these problems through a number of R & D directions we are currently pursuing, to simplify the technology and hence reduce the cost:
 - Multiplexing to reduce number of signal wires to room temperature
 - Alternative SQUID-based ionisation readout
 - Increase the size of individual detectors



100 mm Diameter Germanium Crystals

- Up to 100 mm diameter detector-grade Ge crystals can be grown
- Purchased 2 crystals from Umicore, 33.3 mm thick and of mass 1.37 kg





- Fabricated at SLAC & Stanford, and tested at Minnesota
- New hardware to enable standard 3" readout



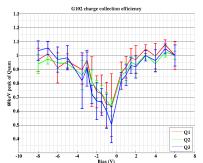
Ionisation Measurements

- Ionisation electrodes patterned as four concentric rings
- ²⁴¹Am source above each electrode



Ionisation Measurements

- Ionisation electrodes patterned as four concentric rings
- ²⁴¹Am source above each electrode



To be presented at LTD14 on Thursday & to appear in J. Low Temp. Phys.

- Charge collection efficiency:
 - Bias voltage varied across detector
 - Gaussian fit to summed charge spectra from all four channels
 - Mean of peak used as a measure of charge collection efficiency
 - Scaling past measurements on 1 cm thick crystals, a bias voltage of 1.7 V required to achieve complete charge collection
- These crystals have necessary charge collection efficiency to be operated as dark matter detectors

Summary

- CDMS-II has completed operation
 - See 2 candidate events, but cannot claim nor reject these as possible WIMPs
 - Reanalysis in progress to reduce background estimates
 - New result published combining CDMS & EDELWEISS data
 - Recent reanalysis with lower 2 keV threshold incompatable with an interpretation of DAMA/LIBRA & CoGENT in terms of spin-independent elastic scatters of low-mass WIMPs
- iZIP technology allows for better rejection of surface events:
 - Essential for scaling to larger target masses
 - 15 kg of 3" diameter, 1" thick Ge iZIP detectors to be installed at Soudan in Autumn 2011 forming SuperCDMS Soudan
- lonisation tests conducted on larger 100 mm diameter,
 33.3 mm thick Ge crystal:
 - Crystals have necessary charge collection efficiency to be operated as dark matter detectors
 - New mask designed with interleaved electrodes to investigate surface event rejection