

# Spin-Dependent WIMP-nucleus Elastic Scattering Simplified

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Based on:

**M. C., J.D. Vergados and M. E. Gomez**

***“Scheme for the extraction of WIMP-nucleon scattering cross sections from total event rates”***

**Phys. Rev. D 83, 075010 (2011), arXiv:1011.6108 [hep-ph]**

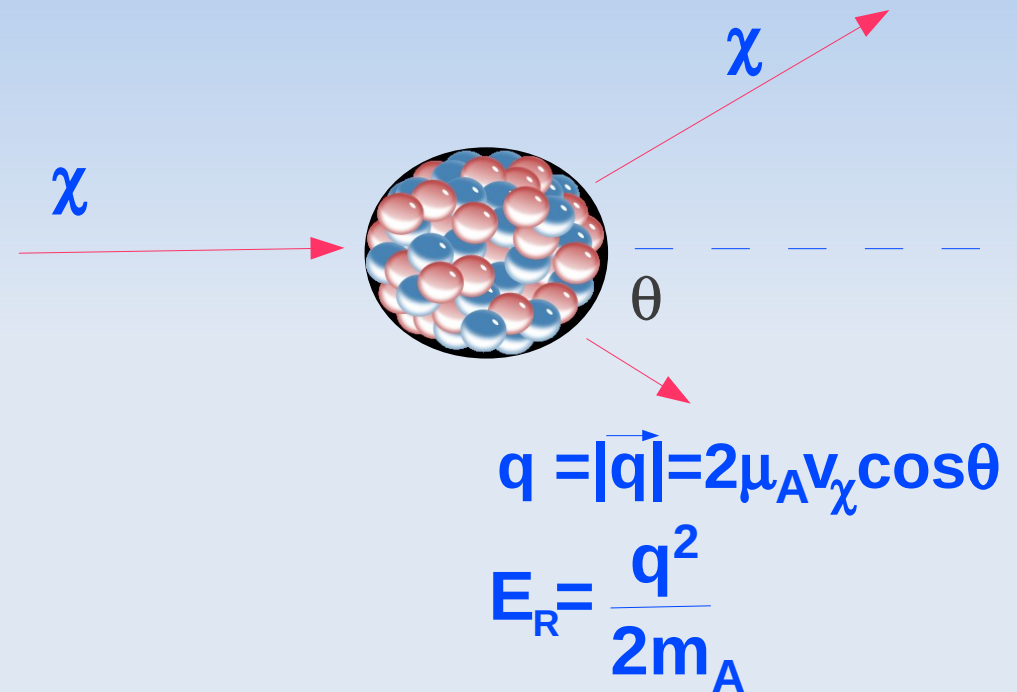
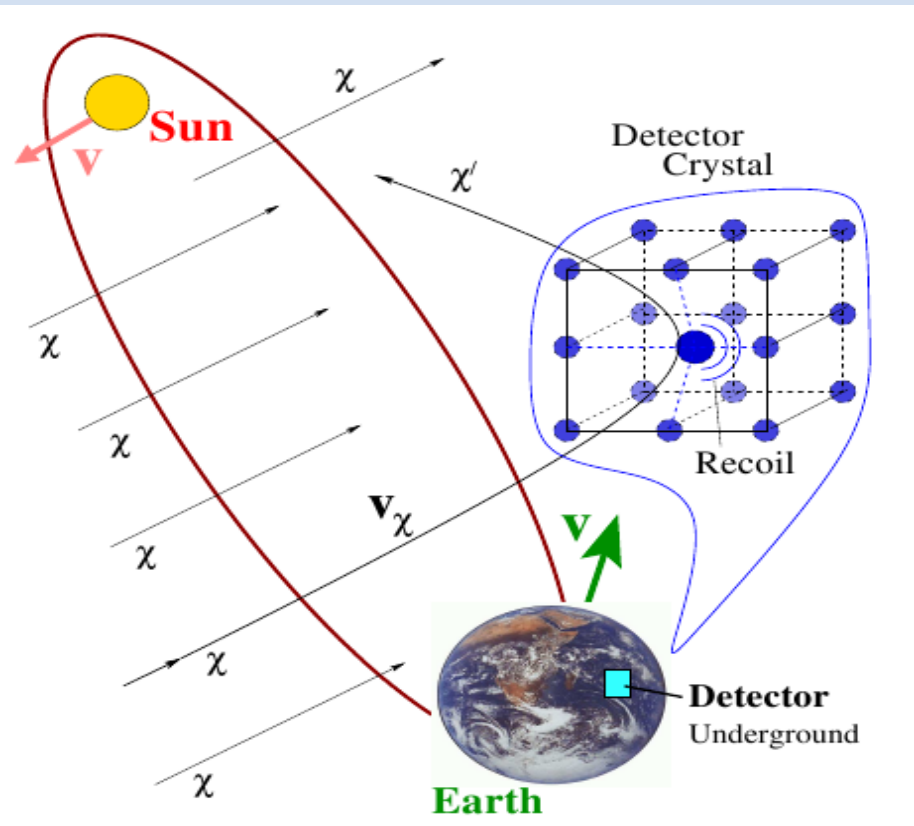
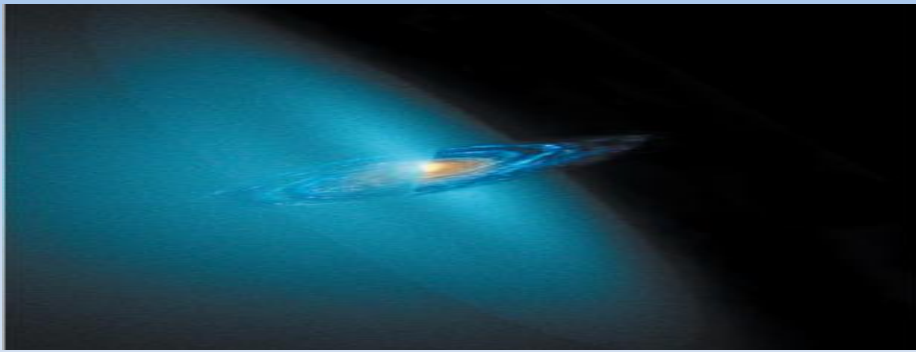
**and**

**M.C.**

***“On upper limits on neutralino-nucleon spin-dependent cross sections”***

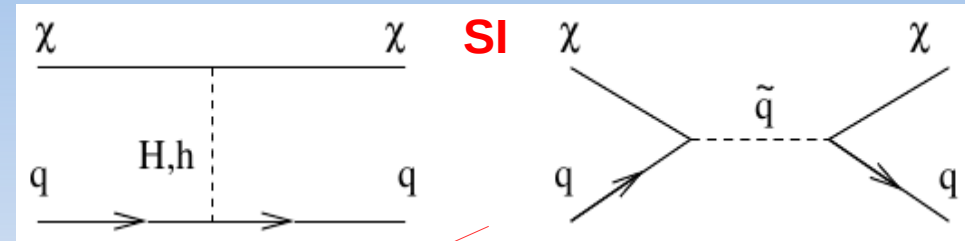
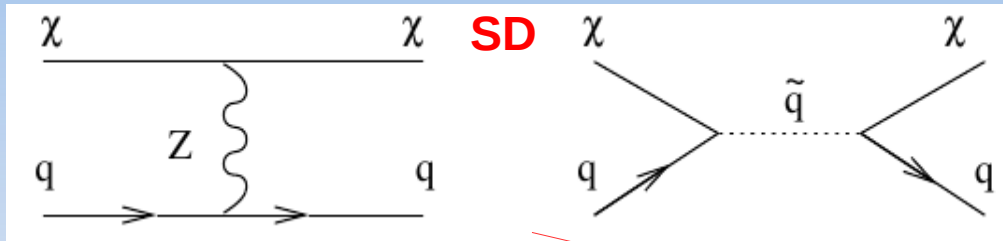
***In preparation***

# WIMP-nucleus elastic scattering



Non relativistic WIMP scatters elastically with nuclei. The recoil of the nucleus deposits a tiny amount of energy in the detector: recoil energies are from few to 100 keV

# Neutralino-nucleon cross sections



**Axial vector  
interaction:**

**Spin-spin interaction  
in the non relativistic  
limit**

$$\mathcal{L}_{eff} = g_q(\bar{\chi}\gamma^\mu\gamma^5\chi)(\bar{q}\gamma_\mu\gamma^5q) + h_q(\bar{\chi}\chi)(\bar{q}q)$$

**Scalar interaction,  
Spin Independent**

$$a_p = \sum_{q=u,d,s} g_q \Delta q^{(p)}$$

$$a_n = \sum_{q=u,d,s} g_q \Delta q^{(n)}$$

$$\lambda^{(p)} = \sum_q h_q f_q^{(p)} \simeq \lambda^{(n)} = \sum_q h_q f_q^{(n)} \equiv c_0$$

**f's and Δ's factors require  
inputs from non  
perturbative QCD**

$$\sigma_{p,n}^{SD} = \frac{3\mu_p^2}{\pi} |a_{p,n}|^2$$

$$\sigma^{SI} = \frac{\mu_p^2}{\pi} |c_0|^2$$

**3 elementary cross sections  
to be extracted from  
experiments**

# WIMP-nucleus SI cross section

- 1) The structure of the nucleus is important if  $qR \sim 1$ : in this case the Zero Momentum Transfer Limit (ZMTL) cross section is not a good approximation, especially for heavy nuclei
- 2) The scalar SI interaction is sensible to the mass distribution inside the nucleus

$$\frac{d\sigma_{(A)}^{SI}}{dq^2} = \frac{1}{4(\mu_A v)^2} \frac{\mu_A^2}{\pi} |\lambda^p Z F^Z(q^2) + \lambda^n (A - Z) F^N(q^2)|^2$$

$$\lambda^p \simeq \lambda^n \equiv c_0$$

$$F^Z(q^2) \simeq F^N(q^2) \cong F(q^2) \xrightarrow{q=0} 1$$

$$\sigma_{(A)}^{SI}(0) = \frac{\mu_A^2}{\pi} |c_0|^2 A^2$$

$$\frac{d\sigma_{(A)}^{SI}}{dq^2} = \frac{1}{4(\mu_A v)^2} \sigma_{(A)}^{SI}(0) F^2(q^2)$$

**The Particle Physics and the Nuclear Physics degrees of freedom are factorized**

# WIMP-nucleus SD cross section: standard formalism (1)

Engel PLB 264 (1991)    Engel, Pittel and Vogel IJMP E 1 (1992)

$$\begin{aligned} \frac{d\sigma_{(A)}^{SD}}{dq^2} &\propto |\langle J, M_Z = J | \sum_{i=1}^A (a_0 + a_1 \tau_i^3) \vec{\sigma}_i e^{-i\vec{q} \cdot \vec{r}_i} | J, M_Z = J \rangle|^2 \\ &\propto \frac{1}{2J+1} \left| \langle J || \sum_{i=1}^A (a_0 + a_1 \tau_i^3) \vec{\sigma}_i e^{-i\vec{q} \cdot \vec{r}_i} || J \rangle \right|^2 \end{aligned}$$

**Isospin  
representation**

$$a_0 = a_p + a_n$$

$$a_1 = a_p - a_n$$

$$\frac{d\sigma_{(A)}^{SD}}{dq^2} \propto \frac{1}{2J+1} \sum_{L_{\text{odd}}} (|\langle J || \mathcal{T}_L^{el5}(q) || J \rangle|^2 + |\langle J || \mathcal{L}_L^5(q) || J \rangle|^2)$$

$$S(q) = a_0^2 S_{00}(q) + a_0 a_1 S_{01}(q) + a_1^2 S_{11}(q)$$

Donnelly and  
Walecka  
NPA 274 (1976)  
Donnelly and  
Peccei, Phys.  
Rep. 50 (1979)  
Donnelly and  
Haxton At.Data  
Nucl. Data Tables  
23 (1979)

# WIMP-nucleus SD cross section: standard formalism (2)

$$S(q) = a_0^2 S_{00}(q) + a_0 a_1 S_{01}(q) + a_1^2 S_{11}(q)$$

$q=0$

$$S(0) = \frac{2J+1}{\pi} J(J+1) \left[ \frac{a_0(\langle \vec{S}_p \rangle + \langle \vec{S}_n \rangle) + a_1(\langle \vec{S}_p \rangle - \langle \vec{S}_n \rangle)}{2J} \right]^2$$

$$\sigma_{(A)}^{SD}(0) = \frac{4\mu_A^2}{2J+1} S(0)$$

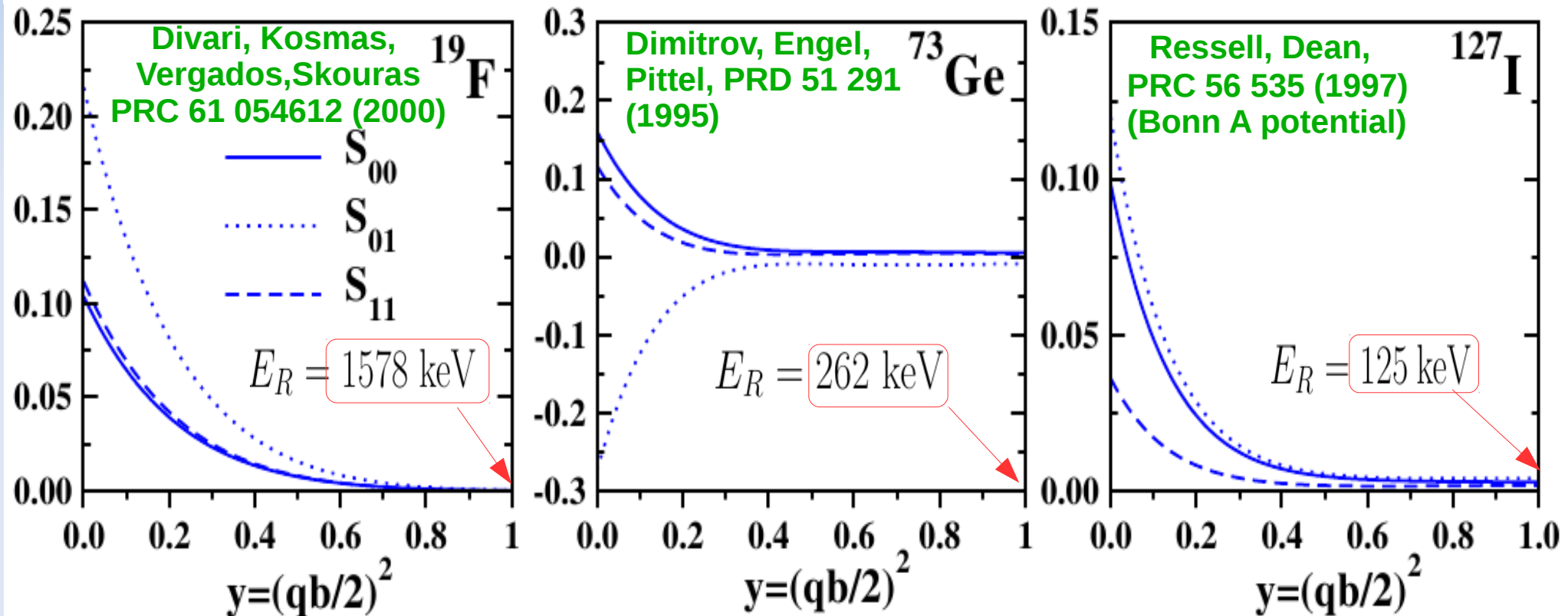
$$\frac{d\sigma_{(A)}^{SD}}{dq^2} = \frac{1}{4(\mu_A v)^2} \sigma_{(A)}^{SD}(0) \frac{S(q)}{S(0)}$$

- 1) To have a spin-dependent “form factor” normalized to one it is necessary to divide by  $S(0)$
- 2) The nuclear physics degrees of freedom are not decoupled from the particle physics ones as in the SI case with the form factor

# WIMP-nucleus SD cross section: standard formalism (2)

The structure functions are furnished as polynomial fits to the results of the shell model calculations in terms of the variable  $y = (qb/2)^2$

$$b = 1 \text{ fm } A^{1/6}$$



- 1) The three momentum-dependent structure functions  $S_{ij}$  look different
- 2) The interference term can be negative

# WIMP-nucleus cross section: Vergados formalism (1)

$$\frac{d\sigma_{(A)}^{SD}}{dq^2} \propto |\langle J, M_Z = J | \sum_{i=1}^A (a_0 + a_1 \tau_i^3) \vec{\sigma}_i e^{-i\vec{q} \cdot \vec{r}_i} | J, M_Z = J \rangle|^2$$

$$\propto \frac{1}{2J+1} |\langle J || \sum_{i=1}^A (a_0 + a_1 \tau_i^3) \vec{\sigma}_i || J \rangle|^2$$

$$\frac{d\sigma_{(A)}^{SD}}{dq^2} \propto (a_0 \Omega_0(q) + a_1 \Omega_1(q))^2$$

Kosmas and Vergados  
PRD 55 (1997)  
Divari, Kosmas, Vergados,  
Skouras PRC 61 (2000)

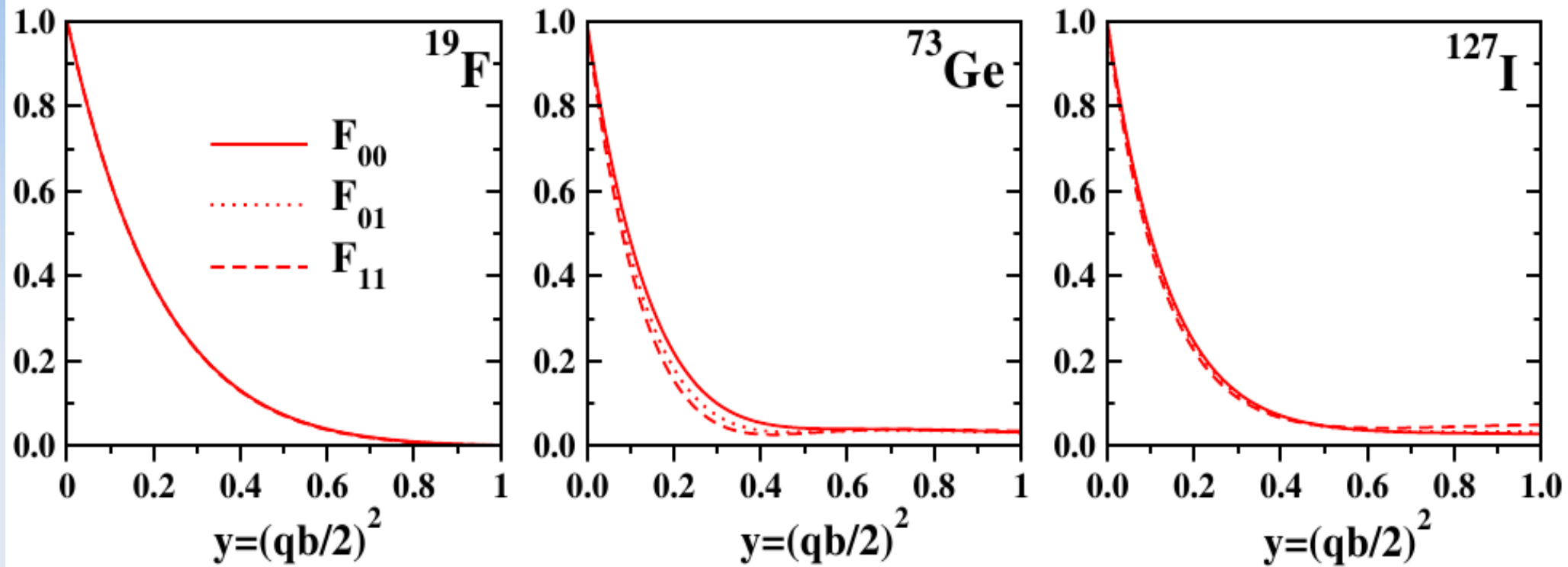
$$a_0^2 \Omega_0^2(0) \left( \frac{\Omega_0(q)}{\Omega_0(0)} \right)^2 + 2a_0 a_1 \Omega_0(0) \Omega_1(0) \left( \frac{\Omega_0(q) \Omega_1(q)}{\Omega_0(0) \Omega_1(0)} \right) + a_1^2 \Omega_1^2(0) \left( \frac{\Omega_1(q)}{\Omega_1(0)} \right)^2$$

$\mathbf{F_{00}(q)}$

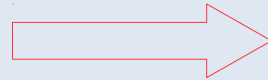
$\mathbf{F_{01}(q)}$

$\mathbf{F_{11}(q)}$

# WIMP-nucleus SD cross section: connection between the two formalisms



$$\begin{aligned}
 S_{00}(q) &= \frac{2J+1}{16\pi} \Omega_0^2 F_{00}(u) \\
 S_{01}(q) &= \frac{2J+1}{8\pi} \Omega_0 \Omega_1 F_{01}(u) \\
 S_{11}(q) &= \frac{2J+1}{16\pi} \Omega_1^2 F_{11}(u)
 \end{aligned}$$

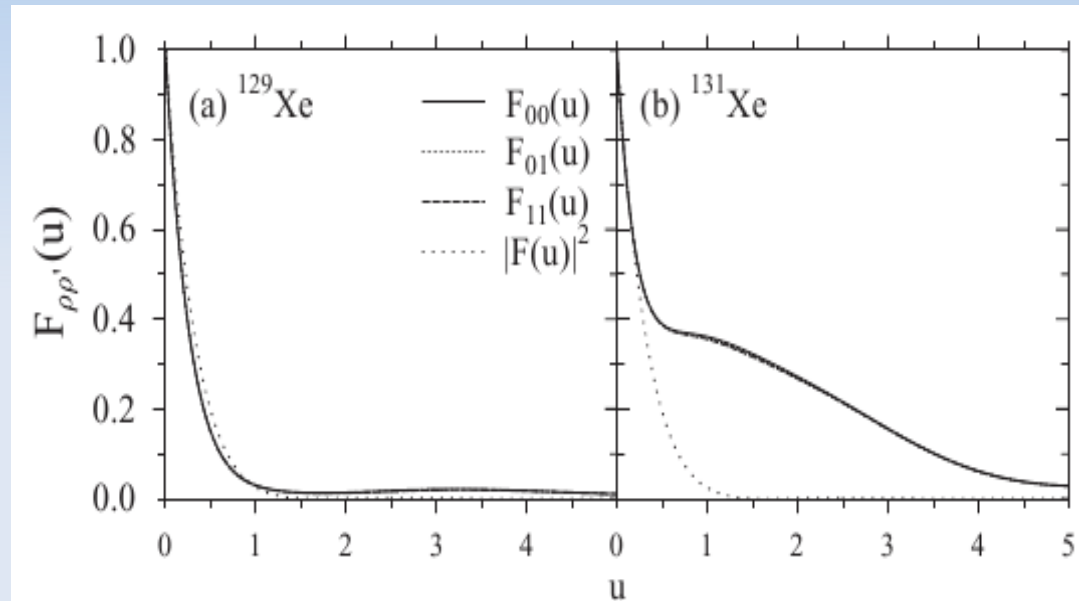
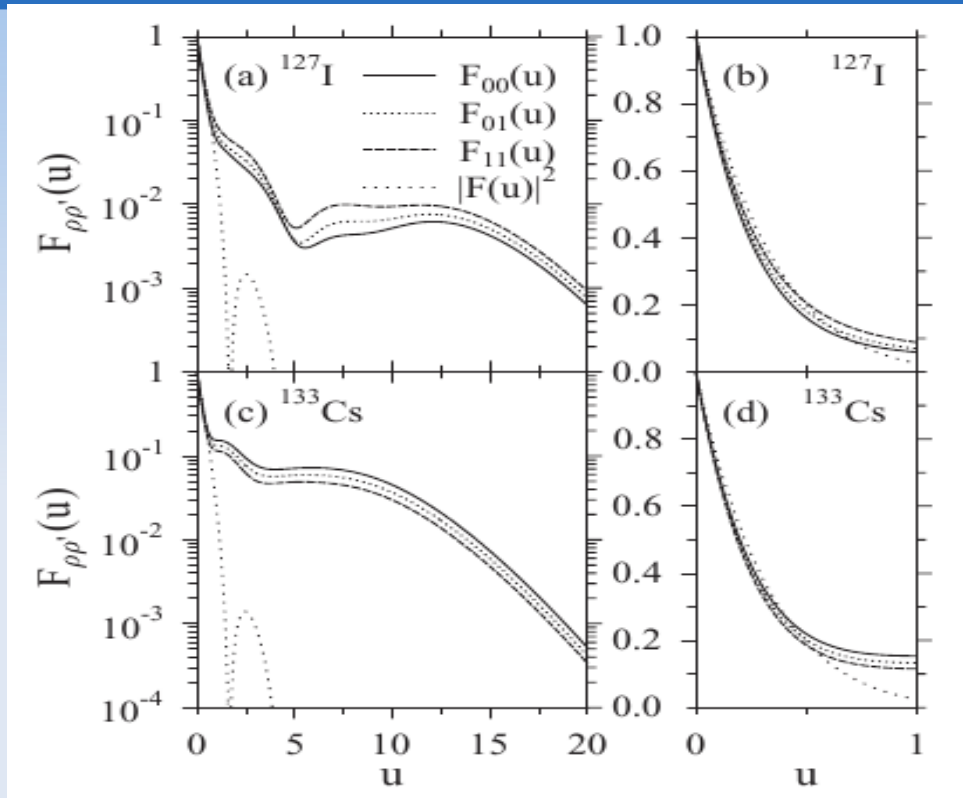


$$\begin{aligned}
 \frac{S_{00}(q)}{S_{00}(0)} &= F_{00}(u) \\
 \frac{S_{01}(q)}{S_{01}(0)} &= F_{01}(u) \\
 \frac{S_{11}(q)}{S_{11}(0)} &= F_{11}(u)
 \end{aligned}$$

$$u = q^2 b^2 / 2$$

# A recent calculation for heavy nuclei

P. Toivanen, Kortelainen, Suohnen,  
J. Toivanen PRC 79 (2009)



- 1) At high momentum transfer there can be larger differences depending on the nucleus
- 2) There are uncertainties due to the nucleon-nucleon potential used in the shell model calculation
- 3) These regions are less interesting for experiments because cross sections are suppressed and with bigger background

# Advantages of the normalized functions

$$F_{00}(u) \simeq F_{01}(u) \simeq F_{11}(u)$$

One only needs one structure function, that by definition is normalized to one at zero momentum transfer

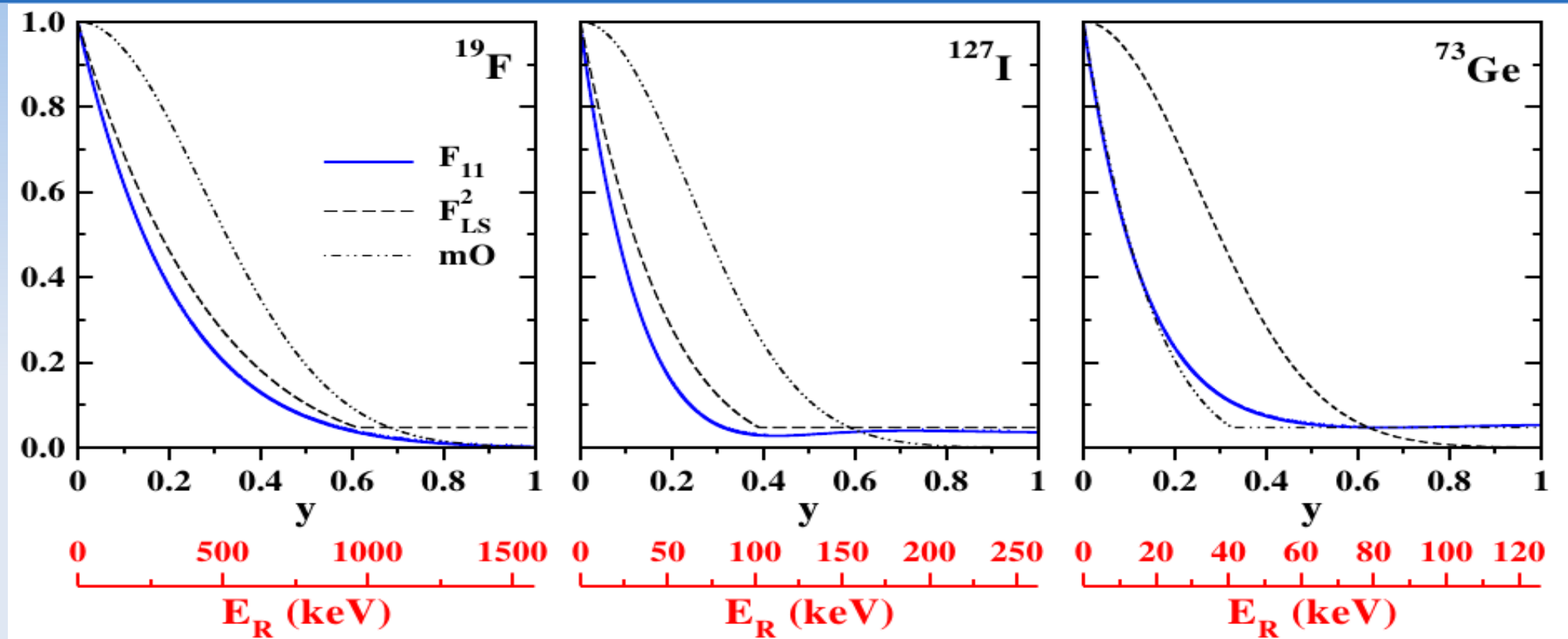
$$\frac{d\sigma_{(A)}^{\text{SD}}}{dE_R} = \frac{1}{2\mu_A^2 v^2} \sigma_{(A)}^{\text{SD}}(0) F_{11}(E_R)$$

Particle physics and nuclear physics are factorized. The cross section has the same form as in the SI case

$$F_{11}(E_R) = S_{11}(E_R) / S_{11}(0)$$

is all that you need for the  
“spin-dependent form factor”

# No need of other parametrizations...



$$F_{LS}^2(qr_n) = \begin{cases} \left( \frac{\sin(qr_n)}{qr_n} \right)^2 & qr_n < 2.55, qr_n > 4.5 \\ 0.047 & 2.55 \leq qr_n \leq 4.5 \end{cases}$$

$$\frac{S_{ij}(q)}{S_{ij}(0)} = \exp\left(-\frac{q^2 R_A^2}{4}\right)$$

J. D. Lewin, P. F. Smith, *Astropart. Phys.* **6** (1996)

MicrOMEGAs, G.~Belanger et al. *Comput. Phys. Commun.* **180** (2009)

# Notation for the total rate

**Nuclear cross sections at zero momentum transfer**

$$\sigma_{(A)}^{SI}(0) = \left( \frac{\mu_A}{\mu_p} \right)^2 A^2 \sigma^{SI}$$

$$\sigma_{(A)}^{SD}(0) = \left( \frac{\mu_A}{\mu_p} \right)^2 \frac{1}{3} \left( \Omega_p^A \sqrt{\sigma_p^{SD}} + \varrho \Omega_n^A \sqrt{\sigma_n^{SD}} \right)^2$$

**Truncated Maxwellian velocity distribution function**

$$f(v) = \frac{1}{\kappa v_E} \left( e^{-\frac{(v-v_E)^2}{v_0^2}} - e^{-\frac{(v+v_E)^2}{v_0^2}} \right)$$

**Tipycal recoil energy**

$$\epsilon_0 = 2\mu_A v_0^2 \left( \frac{\mu_A}{m_A} \right)$$

**The integral are dimensionless**

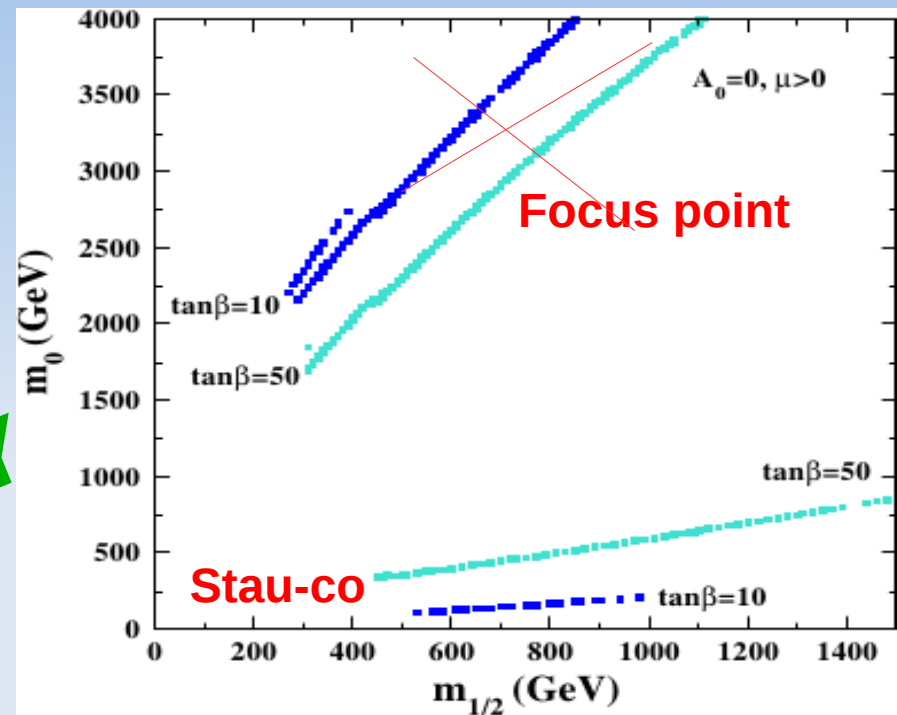
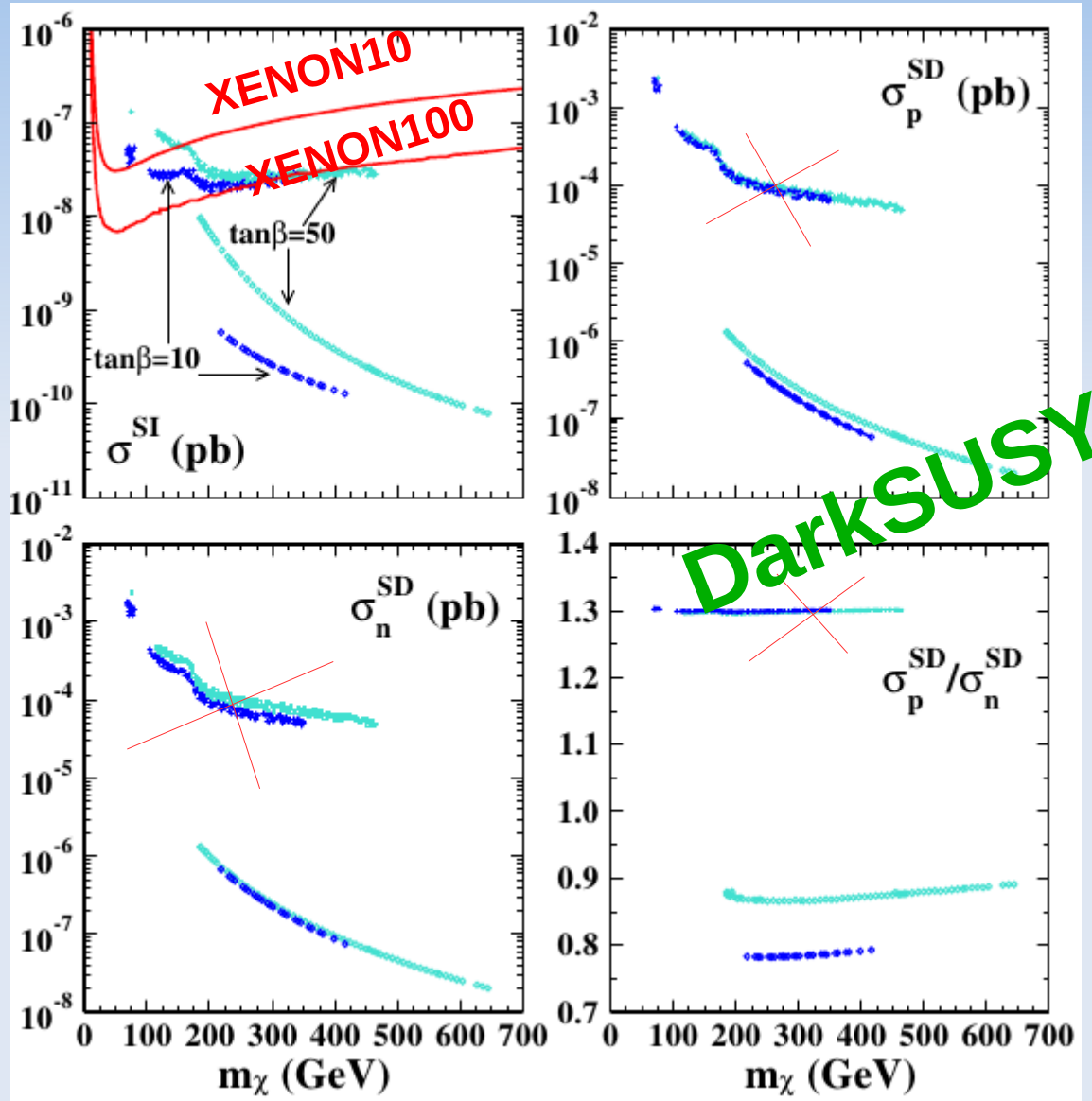
$$t_A^{SD} = \int_{E_1}^{E_2} \frac{dE_R}{\epsilon_0} F_{11}(E_R) \int_{v_{min}(E_R)}^{v_{max}} dv f(v)$$

$$t_A^{SI} = \int_{E_1}^{E_2} \frac{dE_R}{\epsilon_0} F^2(E_R) \int_{v_{min}(E_R)}^{v_{max}} dv f(v)$$

**Total event rate**

$$R = \frac{\rho_0 v_0}{m_\chi m_A} (\sigma_{(A)}^{SI}(0) t_A^{SI} + \sigma_{(A)}^{SD}(0) t_A^{SD})$$

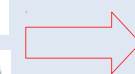
# Constrained MSSM: the stau co-annihilation region



WMAP( $3\sigma$ ):  $0.094 < \Omega h^2 < 0.128$

$$a_n > 0$$

$$a_p < 0$$



$$\varrho = \text{sgn}(a_p/a_n) = \pm 1$$

Can the relative sign between the SD amplitudes and the three cross sections be determined experimentally?

# Spin matrix elements

$$\Omega_{p,n} = 2\sqrt{\frac{J+1}{J}} \langle \vec{S}_{p,n} \rangle$$

**DAMA, KIMS  
COUPP, ANAIS**

$$^{127}\text{I} \Longleftrightarrow A_1: \Omega_p^{127} = 0.731 \quad \Omega_n^{127} = 0.177$$

**Ressell, Dean,  
PRC 56 535 (1997)  
(Bonn A potential)**

**The only nucleus with large protons and neutrons matrix elements**

**CDMS, COGENT  
EDELWEISS**

$$^{73}\text{Ge} \Longleftrightarrow A_3: \Omega_p^{73} = 0.066 \quad \Omega_n^{73} = 0.836$$

**Dimitrov, Engel, Pittel,  
PRD 51 291 (1995)**

**Protons contribution is negligible**

**COUPP, PICASSO  
SIMPLE, DRIFT**

$$^{19}\text{F} \Longleftrightarrow A_2: \Omega_p^{19} = 1.646 \quad \Omega_n^{19} = -0.030$$

**Divari, Kosmas,  
Vergados, Skouras  
PRC 61 054612  
(2000)**

**Neutrons contribution negligible**

$^{19}\text{F} (L_J = S_{1/2})$	$\langle \mathbf{S}_p \rangle$	$\langle \mathbf{S}_n \rangle$	$\mu$ (in $\mu_N$ )
ISPSM, Ellis-Flores [28, 39]	1/2	0	2.793
OGM, Engel-Vogel [29]	0.46	0	(2.629) <sub>exp</sub>
EOGM ( $g_A/g_V = 1$ ), Engel-Vogel [29]	0.415	-0.047	(2.629) <sub>exp</sub>
EOGM ( $g_A/g_V = 1.25$ ), Engel-Vogel [29]	0.368	-0.001	(2.629) <sub>exp</sub>
SM, Pacheco-Strottman [14]	0.441	-0.109	
SM, Divari et al. [5]	0.4751	-0.0087	2.91

**WARNING!**

**In some papers it is stated the opposite using the Pacheco-Strottman matrix elements.**

**The calculations of Divari, Kosmas, Vergados, Skouras are more accurate and do not support this conclusion.**

# Application (1): Cross sections from rates

$$\begin{cases} \text{I} \\ \text{F} \\ \text{Ge} \end{cases} \begin{cases} \sigma^{SI} + \mathcal{R}_{A_1} \left( \Omega_p^{A_1} \sqrt{\sigma_p^{SD}} + \varrho \Omega_n^{A_1} \sqrt{\sigma_n^{SD}} \right)^2 - \mathcal{S}_{A_1} = 0 \\ \sigma^{SI} + \mathcal{R}_{A_2} (\Omega_p^{A_2})^2 \sigma_p^{SD} - \mathcal{S}_{A_2} = 0 \\ \sigma^{SI} + \mathcal{R}_{A_3} (\Omega_n^{A_3})^2 \sigma_n^{SD} - \mathcal{S}_{A_3} = 0 \end{cases}$$

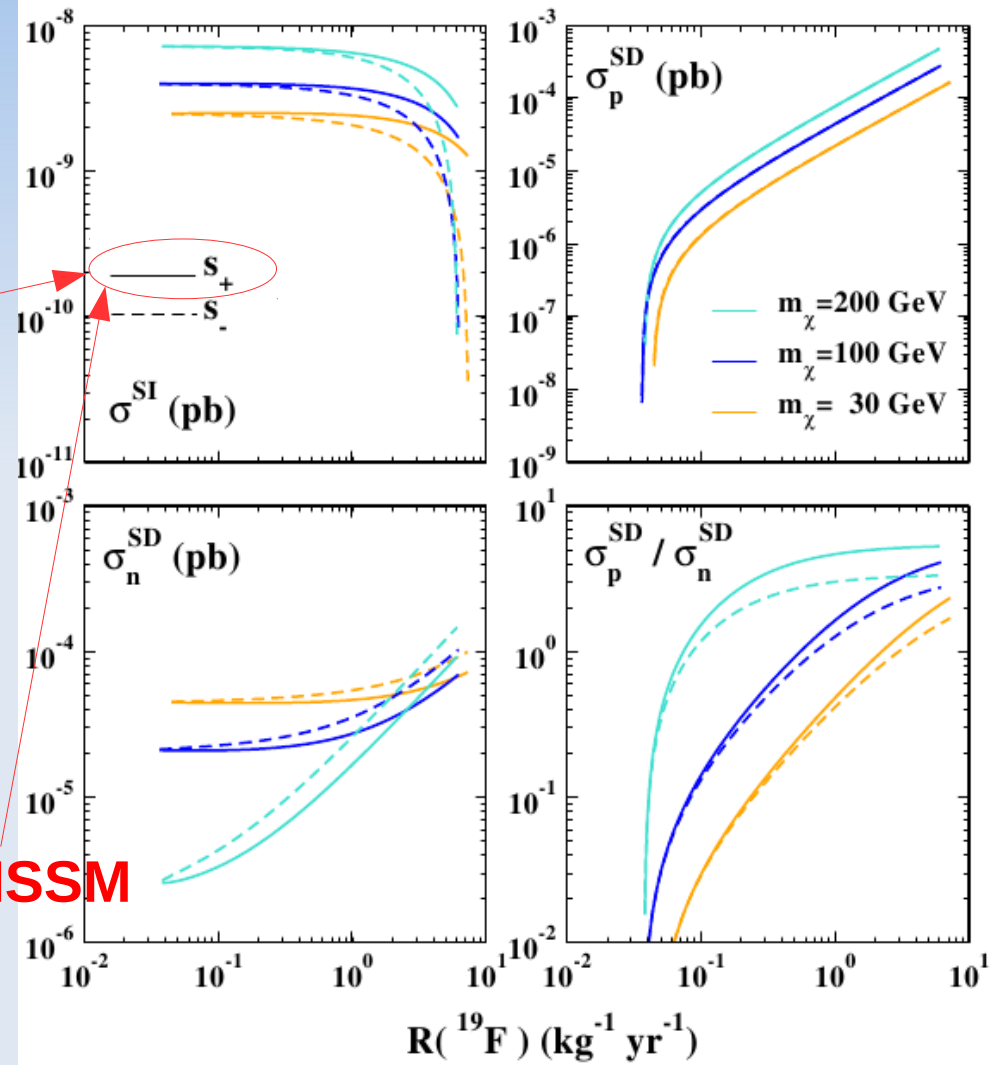
$$s_+ : \begin{cases} \sigma_+^{SI} = \frac{\mathcal{B} + \sqrt{\mathcal{B}^2 - \mathcal{A}\mathcal{C}}}{\mathcal{A}} \\ \sigma_{p,+}^{SD} = \frac{\mathcal{S}_{A_2} - \sigma_+^{SI}}{\mathcal{R}_{A_2} (\Omega_p^{A_2})^2} \\ \sigma_{n,+}^{SD} = \frac{\mathcal{S}_{A_3} - \sigma_+^{SI}}{\mathcal{R}_{A_3} (\Omega_n^{A_3})^2} \end{cases} \quad s_- : \begin{cases} \sigma_-^{SI} = \frac{\mathcal{B} - \sqrt{\mathcal{B}^2 - \mathcal{A}\mathcal{C}}}{\mathcal{A}} \\ \sigma_{p,-}^{SD} = \frac{\mathcal{S}_{A_2} - \sigma_-^{SI}}{\mathcal{R}_{A_2} (\Omega_p^{A_2})^2} \\ \sigma_{n,-}^{SD} = \frac{\mathcal{S}_{A_3} - \sigma_-^{SI}}{\mathcal{R}_{A_3} (\Omega_n^{A_3})^2} \end{cases}$$

In the case of I, F, Ge, the solutions are connected with the relative sign between the proton and neutron amplitudes appearing in the equation for I

$$\varrho = +1 \Leftrightarrow s_-$$

$$\varrho = -1 \Leftrightarrow s_+$$

CMSSM



$$R(^{127}\text{I}) = R(^{73}\text{Ge}) = 1 \text{ kg}^{-1} \text{ y}^{-1}$$

# Application (2): Model independent UL are...naturally model independent!

Suppose we can neglect the  
SI rate

$$\phi_A = \frac{\rho_0 v_0}{m_\chi m_A}$$

$$R^{\text{SD}} = \phi_A \left( \mathcal{C}_A^p \sqrt{\sigma_p^{\text{SD}}} \pm \mathcal{C}_A^n \sqrt{\sigma_n^{\text{SD}}} \right)^2 t_A^{\text{SD}}$$

$$\mathcal{C}_A^{p,n} = \frac{\mu_A}{\mu_p} \frac{\Omega_{p,n}^A}{\sqrt{3}}$$

If an experiment with exposure  $\mathcal{E}$  sets an  
upper limit on the number of events then:

$$R^{\text{SD}} \times \mathcal{E}_A < N^{\text{UL}}$$

$$\left( \mathcal{C}_A^p \sqrt{\sigma_p^{\text{SD}}} \pm \mathcal{C}_A^n \sqrt{\sigma_n^{\text{SD}}} \right)^2 < \frac{N^{\text{UL}}}{\phi_A t_A^{\text{SD}} \mathcal{E}_A}$$

$$= \sigma_A^{\text{lim}} \equiv \frac{N^{\text{UL}}}{\phi_A t_A^{\text{SD}} \mathcal{E}_A}$$

**Define  
also**

$$\sigma_{p,n}^{\text{lim(A)}} \equiv \frac{\sigma_A^{\text{lim}}}{(\mathcal{C}_A^{p,n})^2}$$

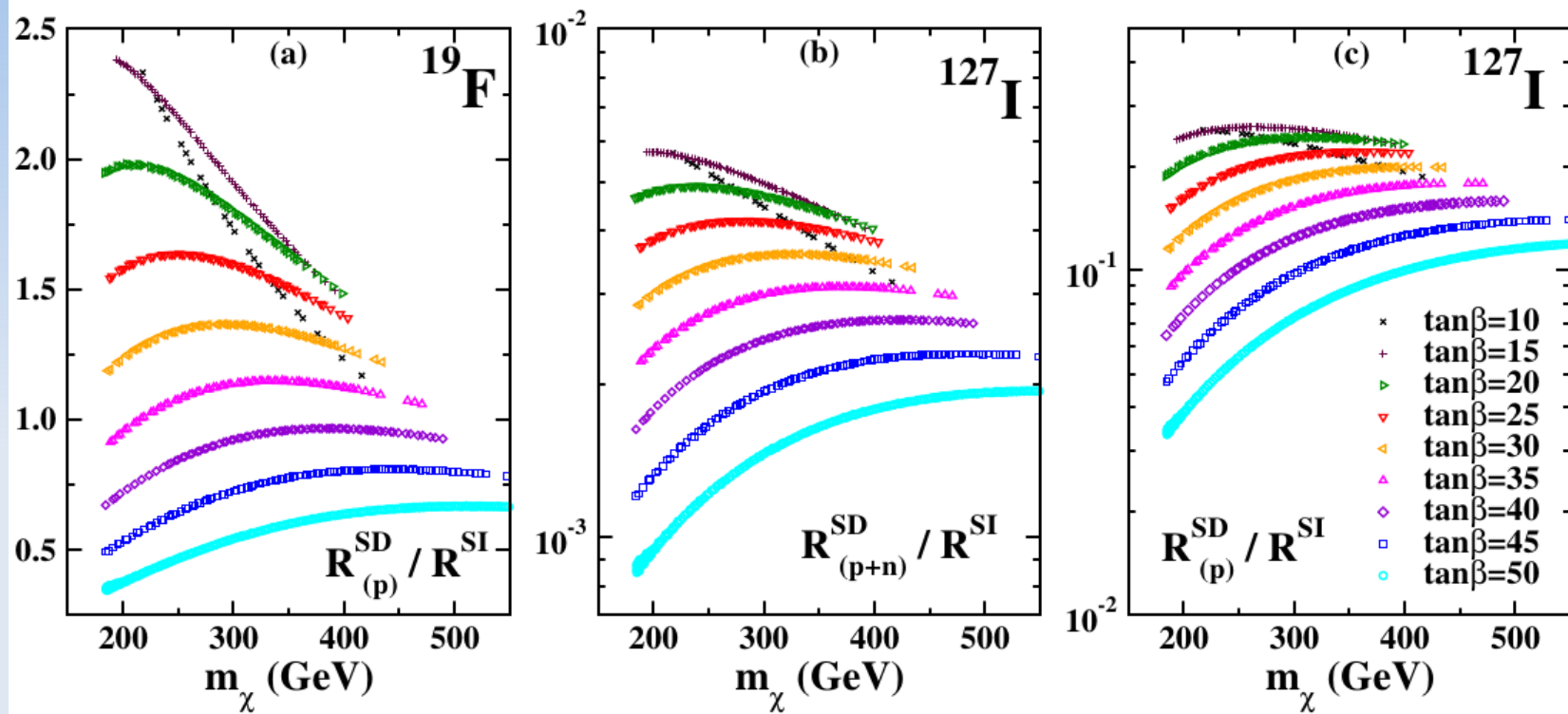
$$\left( \frac{\sqrt{\sigma_p^{\text{SD}}}}{\sqrt{\sigma_p^{\text{lim(A)}}}} \pm \frac{\sqrt{\sigma_n^{\text{SD}}}}{\sqrt{\sigma_n^{\text{lim(A)}}}} \right)^2 < 1$$

- 1) The formula of **Tovey et al. PLB 488 (2000)** follows naturally from the factorization.  
The model independent method... is not a method, after all.
- 2) Not other assumptions are needed to derive it. The correct nuclear physics is taken into account

# Fluorine, Iodine and the stau co-annihilation region

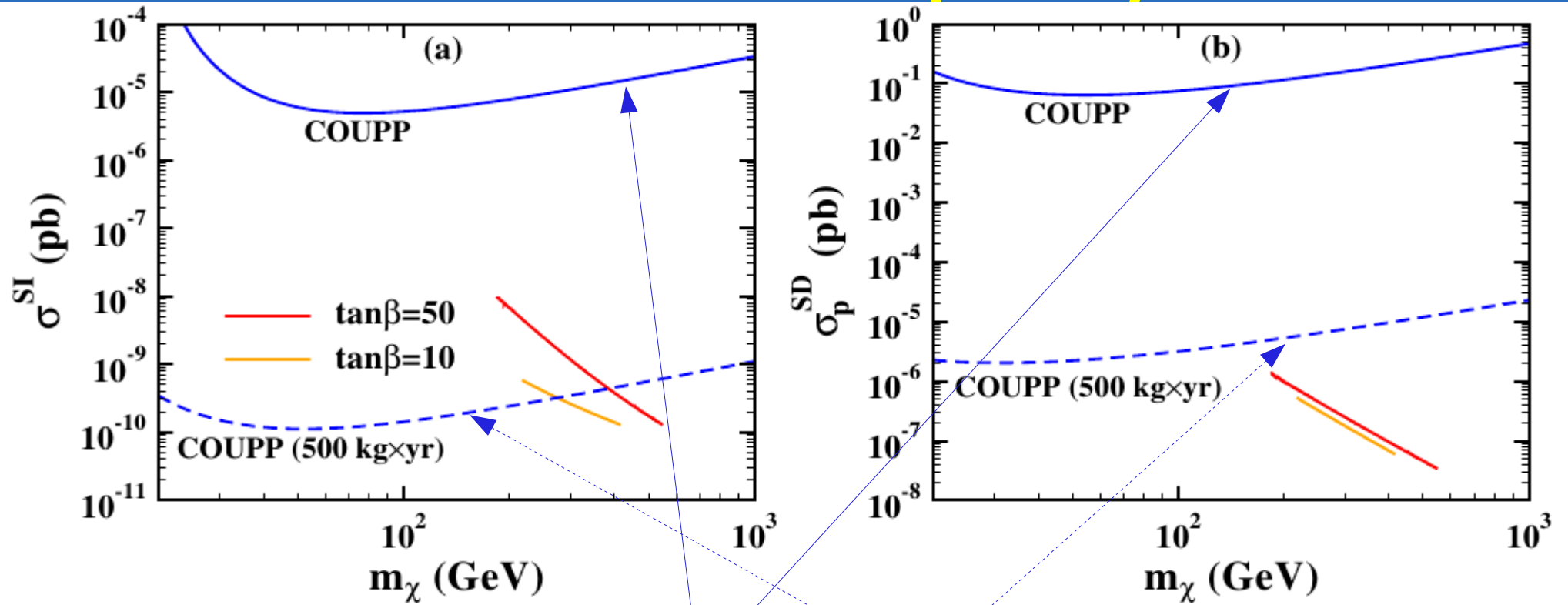
$$\sigma_{p,n}^{\text{lim}(A)} \equiv \frac{\sigma_A^{\text{lim}}}{(C_A^{p,n})^2}$$

Are these  
“universal”  
upper limits?



- 1) In Fluorine the SI rate is never negligible: bigger than the SD for  $\tan\beta > 35$  and smaller for  $\tan\beta < 35$ . Limits on the proton SD cross section are obtained assuming no SI rate....
- 2) In Iodine there is a cancellation in the SD rate that is two orders of magnitude smaller than the SI: the limits on the single nucleon SD cross sections derived by Iodine are meaningless for this model

# The stau co-annihilation region and COUPP (CF<sub>3</sub>I)



COUPP Collaboration PRL 106 (2011)

$$\mathcal{E} = 28.1 \text{ kg} \times \text{days} \quad N_{90\% \text{C.L.}}^{UL} = 6.7$$

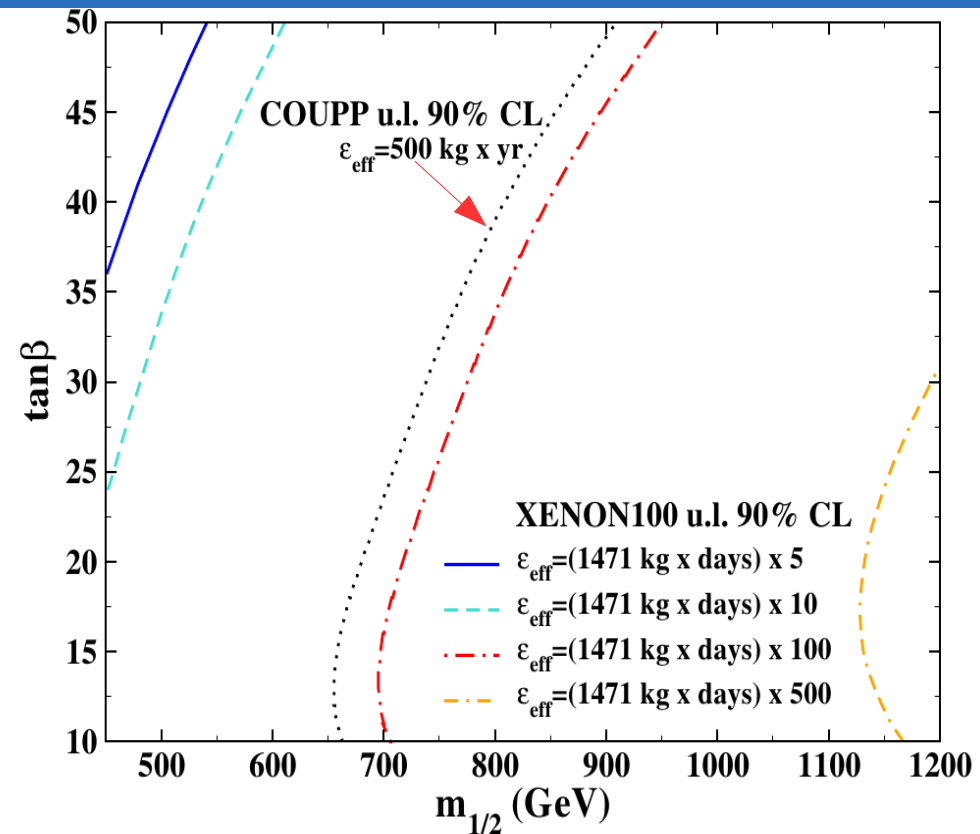
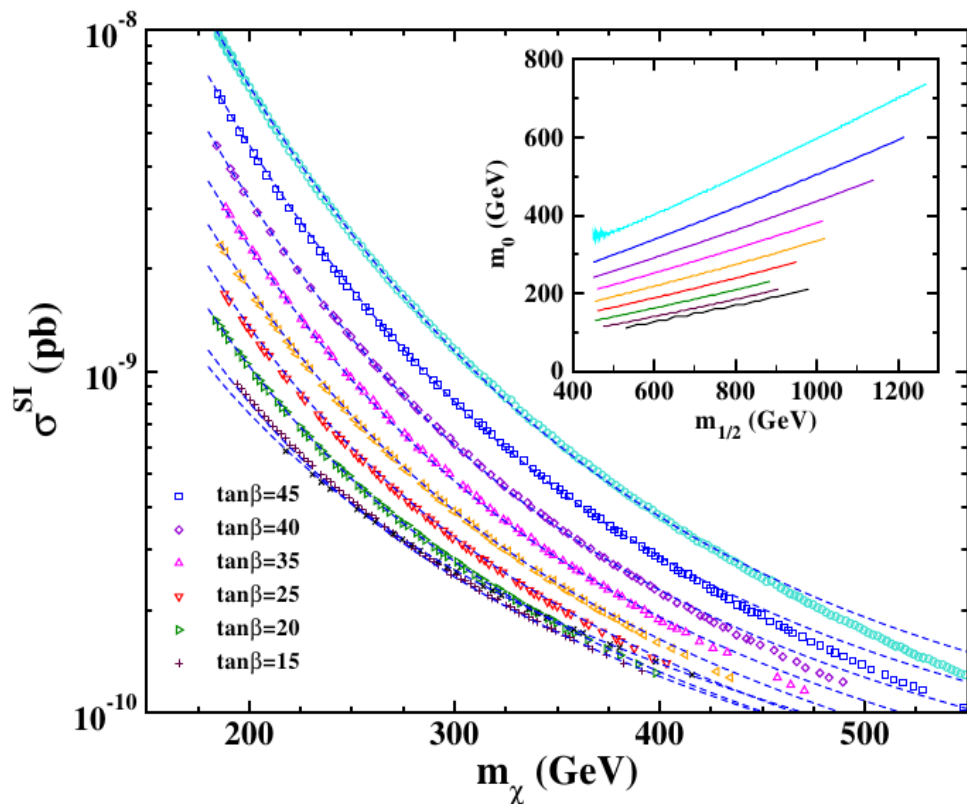
$$50\% \text{ efficiency} \quad E_{th} = 21 \text{ keV}$$

$$\mathcal{E} = 500 \text{ kg} \times \text{year} \quad N_{90\% \text{C.L.}}^{UL} = 6.7$$

$$100\% \text{ efficiency} \quad E_{th} = 7 \text{ keV}$$

**COUPP can probe the model only through SI interaction with Iodine but not through the SD interaction with Fluorine**

# Constraining directly the $(m_{1/2}, \tan\beta)$ plane



Fit as a function of  $\tan\beta$  and  $m_{1/2}$

$$\sigma^{SI}(\tan\beta, m_\chi) = \sum_{k=2}^4 \left[ \sum_{i=0}^3 \sigma_{ki} (\tan\beta)^i \left( \frac{100 \text{ GeV}}{m_\chi} \right)^k \right]$$

$$m_\chi \simeq 0.44 m_{1/2} - 15 \text{ GeV}$$

Translate an upper limit  
 $N_{UL}$  directly  
into a constrain on the  
 $(m_{1/2} - \tan\beta)$  plane