

Rotational Effects on Convection

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Conference on Differential Rotation and Magnetism across the HR Diagram

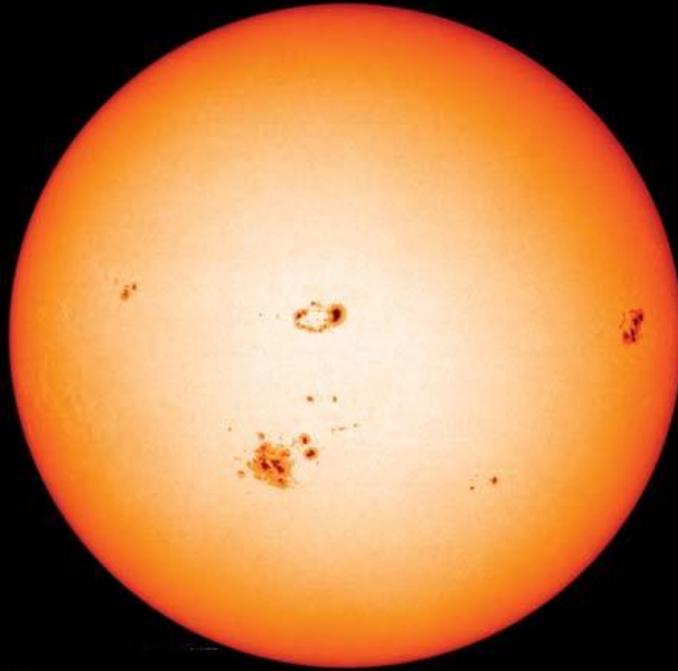
Great Red Spot of Jupiter NASA



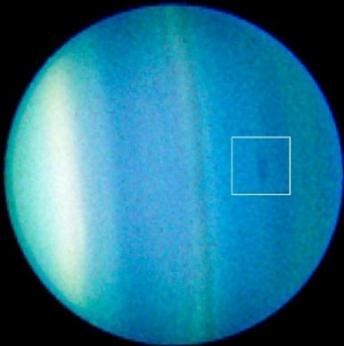
Saturn spots Cassini/NASA



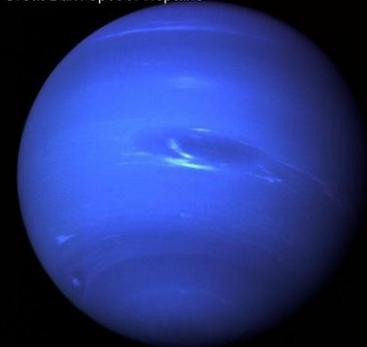
Solar Image by SOHO



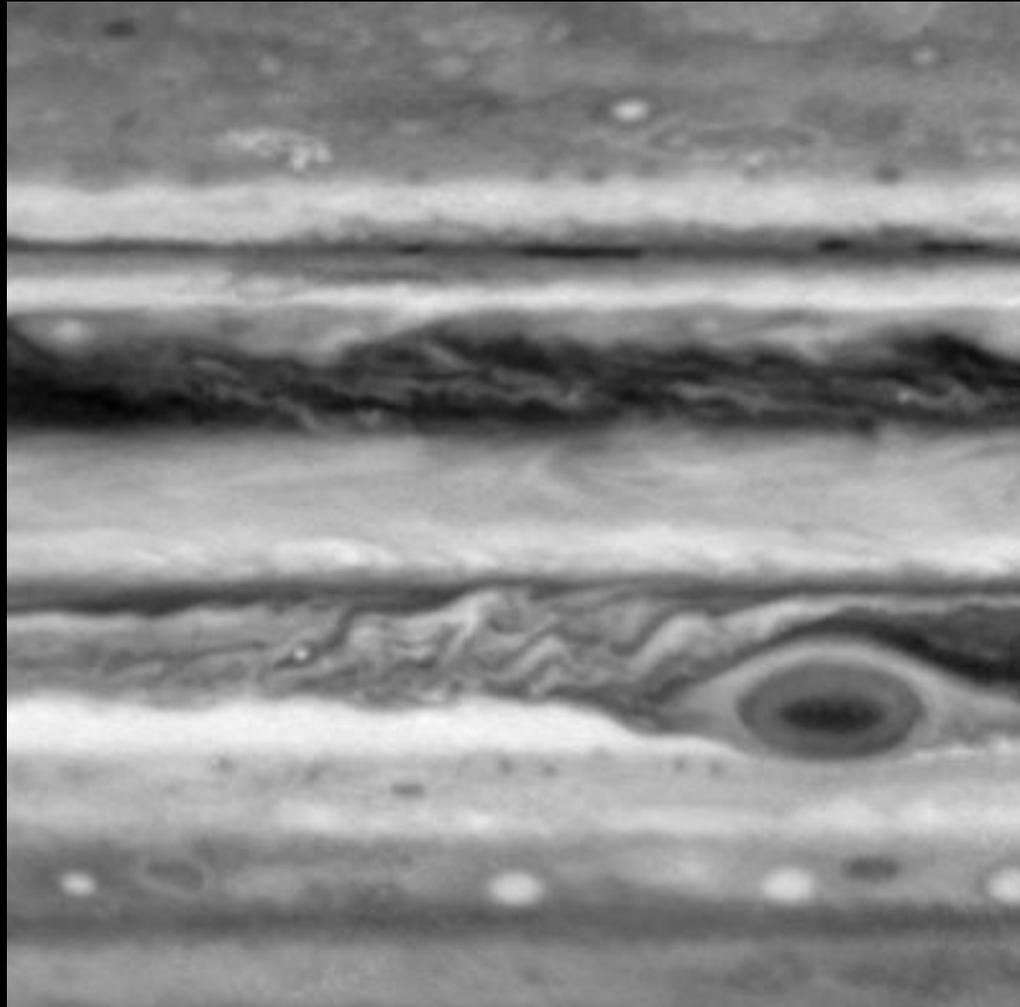
Uranus Dark Spot *Hubble Space Telescope*



Great Dark Spot of Neptune



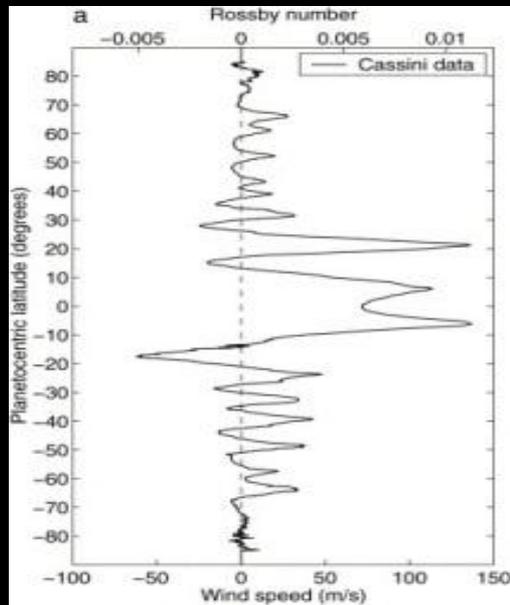
Two Big Puzzles: Jets and Spots



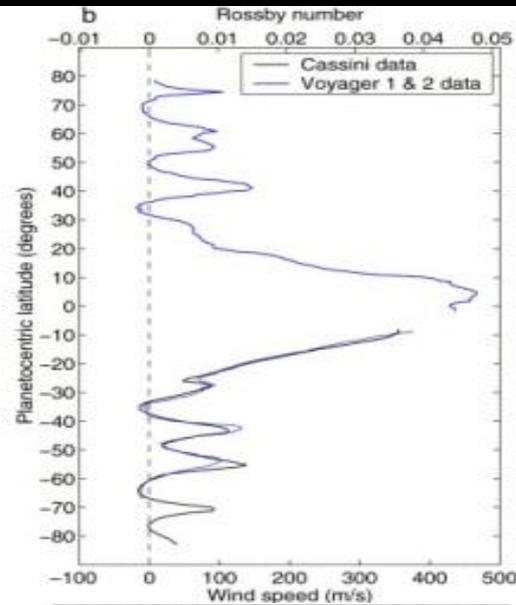
Credit: Cassini/NASA

Mean Zonal Winds

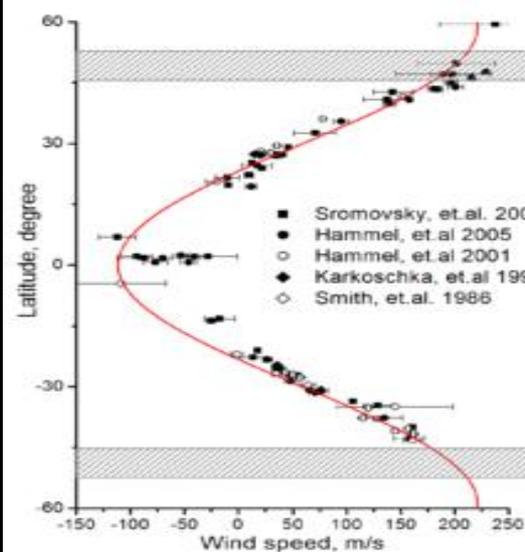
Jupiter



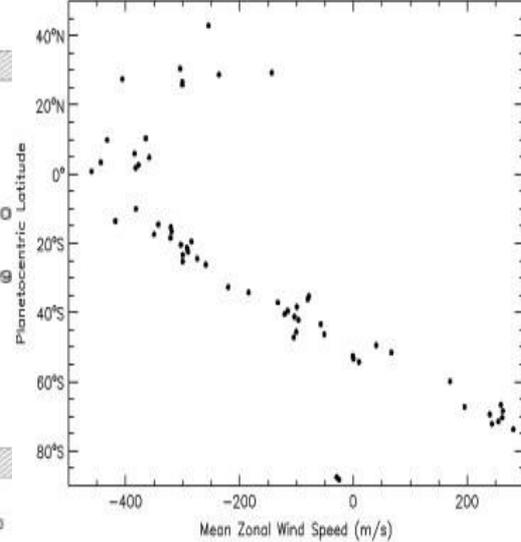
Saturn



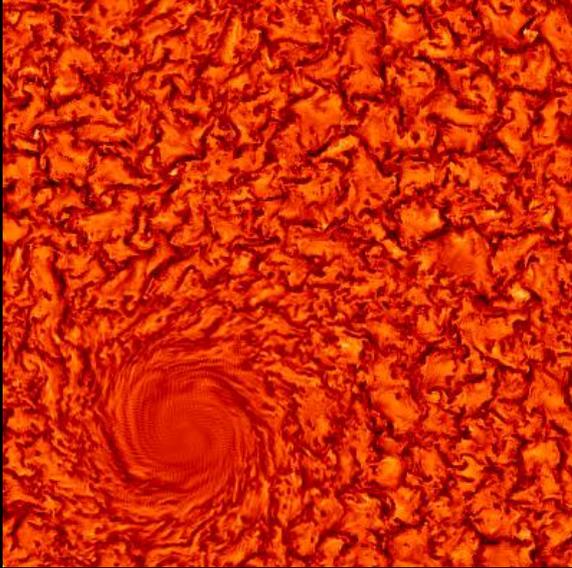
Uranus



Neptune



Long-lived Vortices



- Chan 2003, *Numerical study of convection*, in ASP Conf. Ser. 293, 3D Stellar Evolution, eds. S. Turcotte, S.C. Keller, & R.M. Cavallo (San Francisco)
- Chan 2005, *Flow patterns and transitions in rotating convection*, in Proceedings of the Workshop on Interdisciplinary Aspects of Turbulence, eds. F. Kupka and W. Hillebrandt, p.159-160 (Munich).
- Chan 2007, *Rotating convection in f-boxes: Faster rotation*, *Astron. Nachr.*, 328: 1059-1061.
- Kapyla et al. 2011, *Starspots due to large-scale vortices in rotating turbulent convection*, *Ap. J.*, 742: 34-41.
- Mantere et al. 2011, *Dependence of the large-scale vortex instability on latitude, stratification, and domain size*, *Astron. Nachr.*, 332: 876-882
- Chan & Mayr 2013, *Numerical simulation of convectively generated vortices: application to the Jovian planets*, *Earth Planetary Science Lett.*, in press.

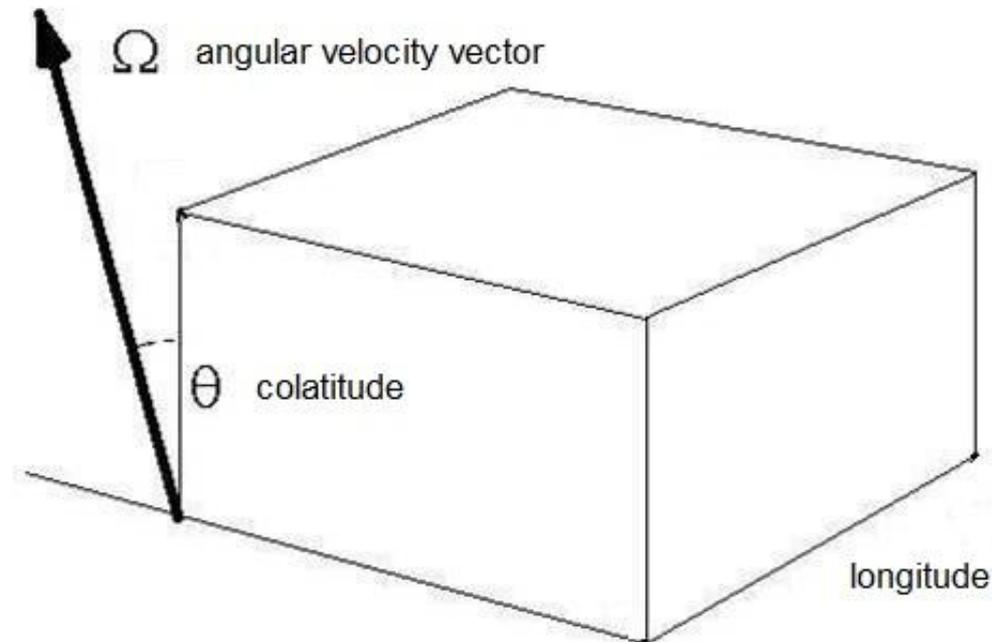


3D LES of Turbulent Compressible Convection in a f-Box

$$\partial_t \rho = -\nabla \cdot (\rho \mathbf{v})$$

$$\partial_t (\rho \mathbf{v}) = -\nabla \cdot (\rho \mathbf{v} \mathbf{v}) - \nabla p + \nabla \cdot \Sigma + \rho \mathbf{g} + 2\rho \mathbf{v} \times \Omega$$

$$\partial_t E = -\nabla \cdot \{ (E + p) \mathbf{v} - \mathbf{v} \cdot \Sigma + \mathbf{f} \} + \rho \mathbf{v} \cdot \mathbf{g}$$



Model Description

- Constant angular velocity vector
- Depth ~ 4 pressure scale heights; aspect ratio = 4; $\gamma = 5/3$
- Heat flux = 0.03125 pressure \times C_s ; C_s = isothermal sound speed at top
- Units: length = d , time = d/C_s , thermal quantities = values at top
- Latitude = 67.5, 22.5; $\Omega = 0.375, 0.75, 1.5, 3$.
- Conservative finite-difference scheme; 540x540x80 grids
- $Re \sim 2 \times 10^3 - 10^4$; $Ta \sim 2 \times 10^8 - 10^{10}$

Three ways to define the Coriolis number:

$$Co1 = f_0 d/V$$

$$V = \text{rms velocity, } f_0 = 2\Omega$$

$$Co2 = f_0 d/V_0$$

$$V_0 = V \text{ of the non-rotating case}$$

$$Co3 = f d/V_0$$

$$f = 2\Omega \cos(\text{colatitude})$$

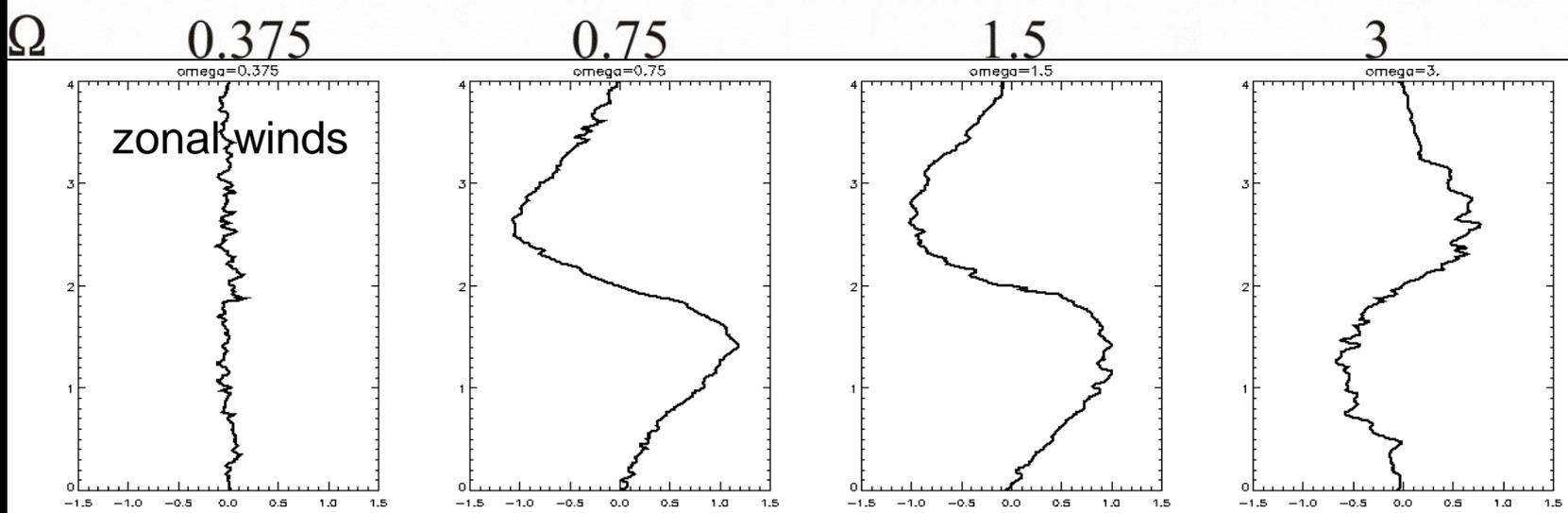
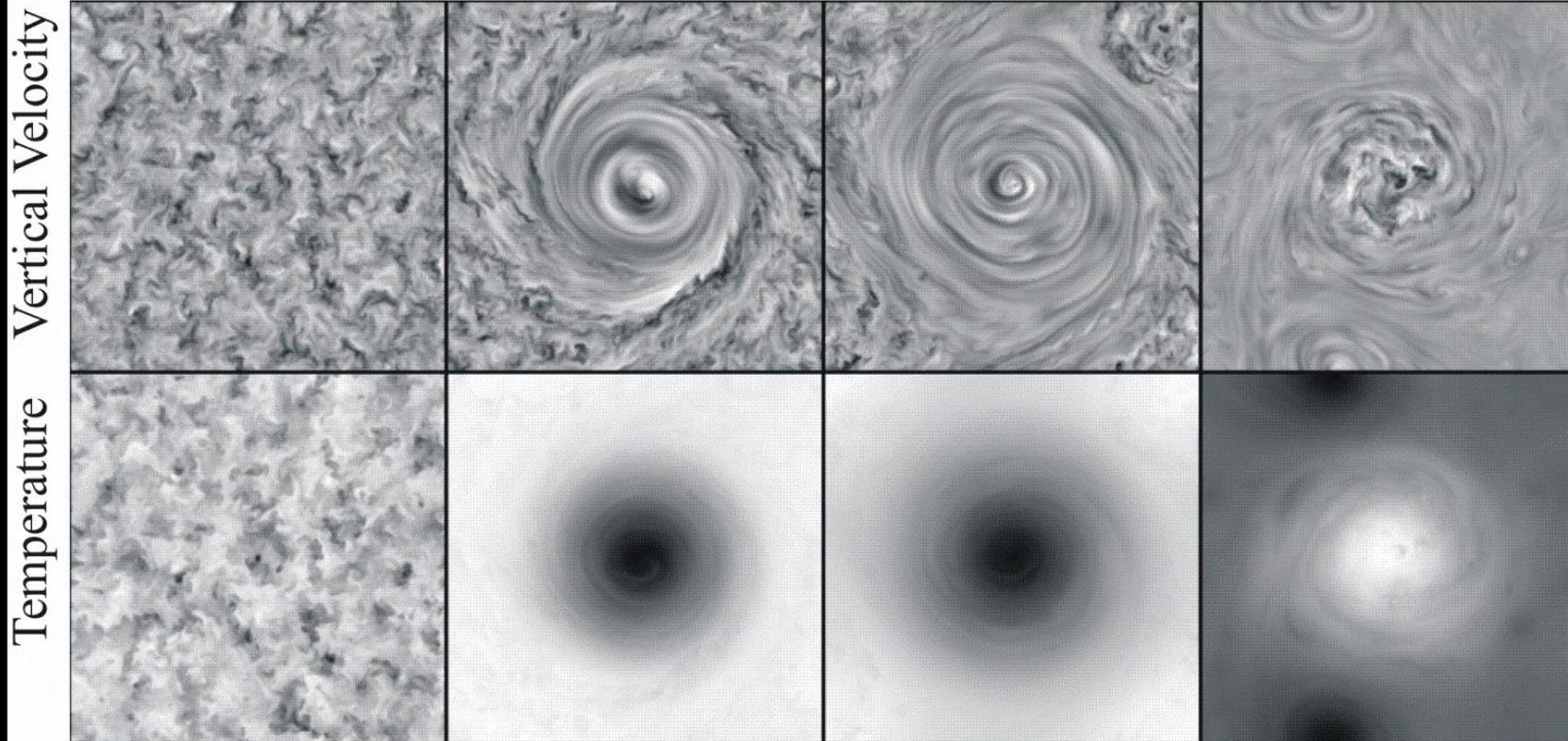
Case Set 1

Case	λ	Ω	V	Co1 f_0d/V	Co2 f_0d/V_0	Co3 fd/V_0
0			0.136			
1	67.5	0.375	0.103	6.02	4.31	3.98
2	67.5	0.75	0.526	2.28	8.62	7.94
3	67.5	1.5	0.582	4.17	17.2	15.9
4	67.5	3.0	0.354	13.7	34.5	32.3
5	22.5	0.375	0.118	5.03	4.31	1.65
6	22.5	0.75	0.230	5.24	8.62	3.30
7	22.5	1.5	0.198	12.0	17.2	6.58
8	22.5	3.0	0.285	17.2	34.5	13.2

67.5° latitude

← Cyclones →

Anticyclone

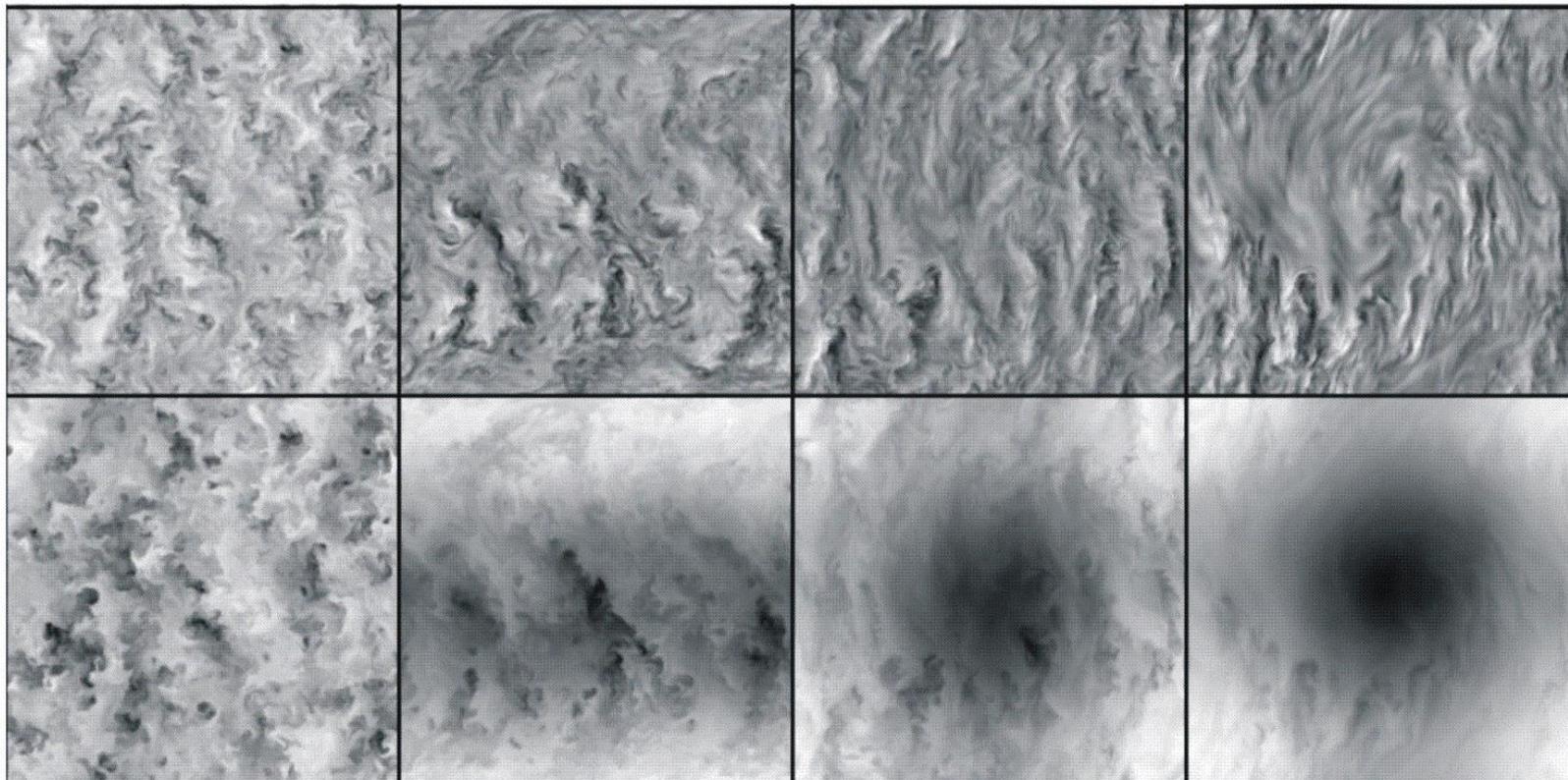


22.5° latitude

← Cyclones →

Vertical Velocity

Temperature



Ω

0.375

0.75

1.5

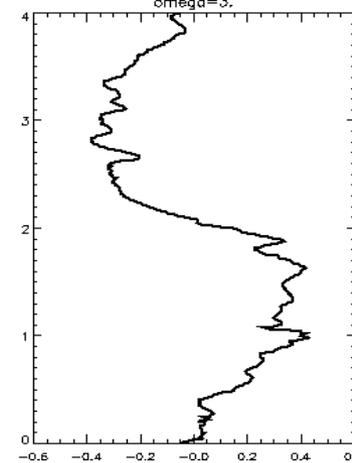
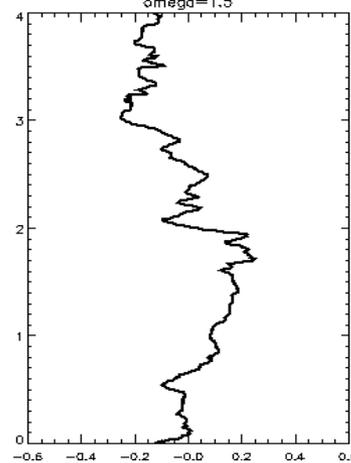
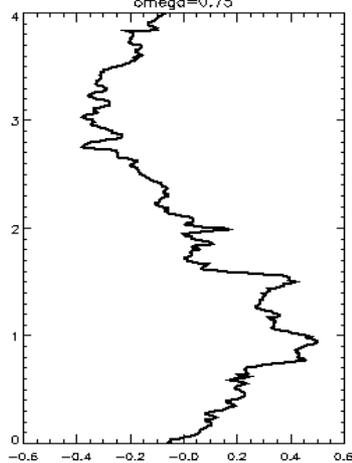
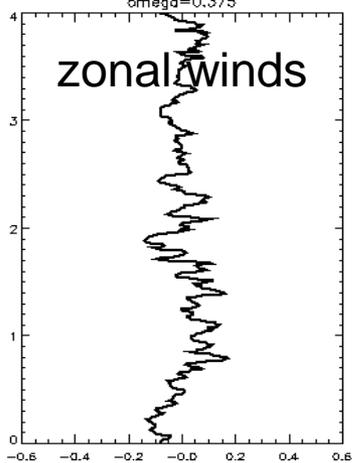
3

omega=0.375

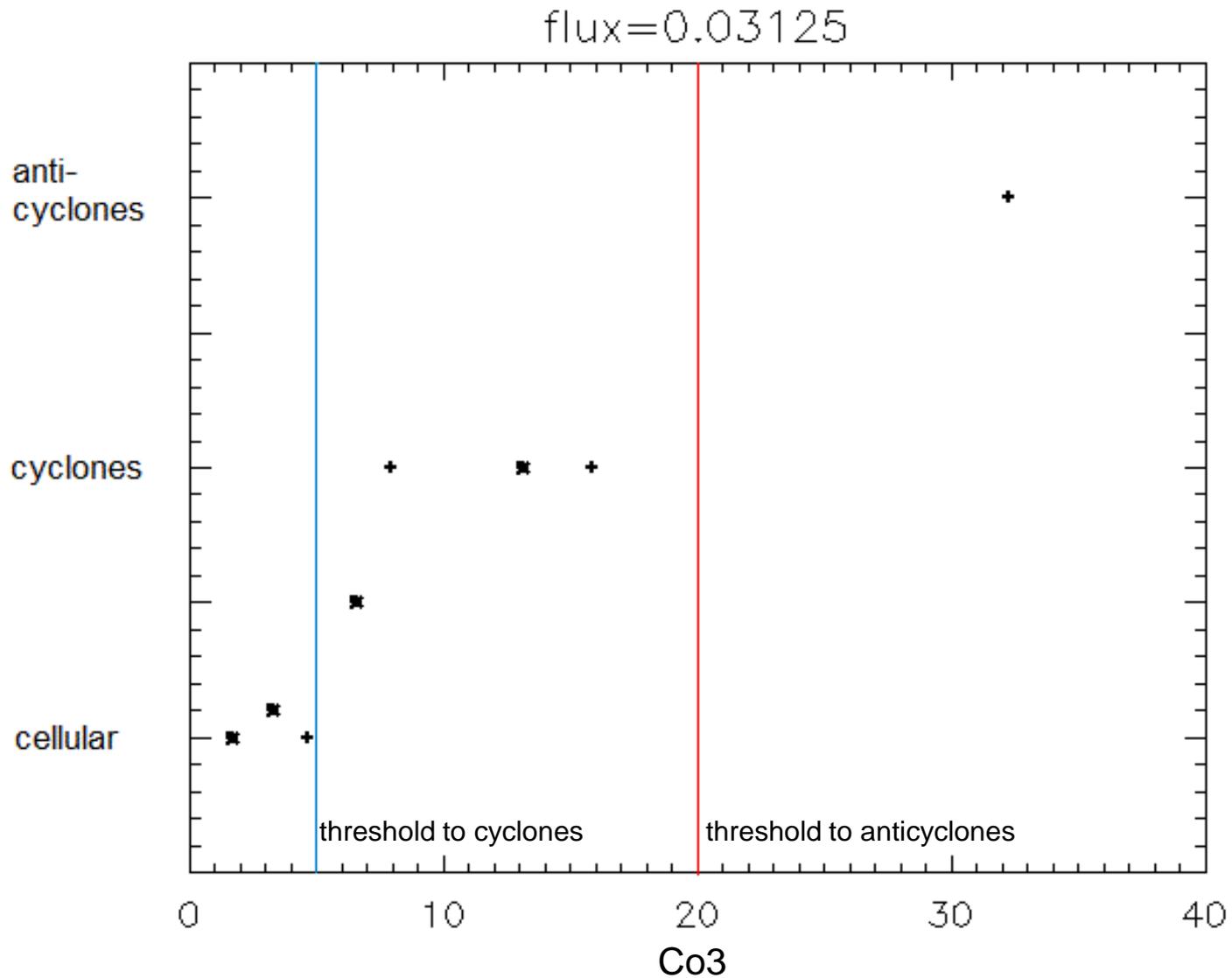
omega=0.75

omega=1.5

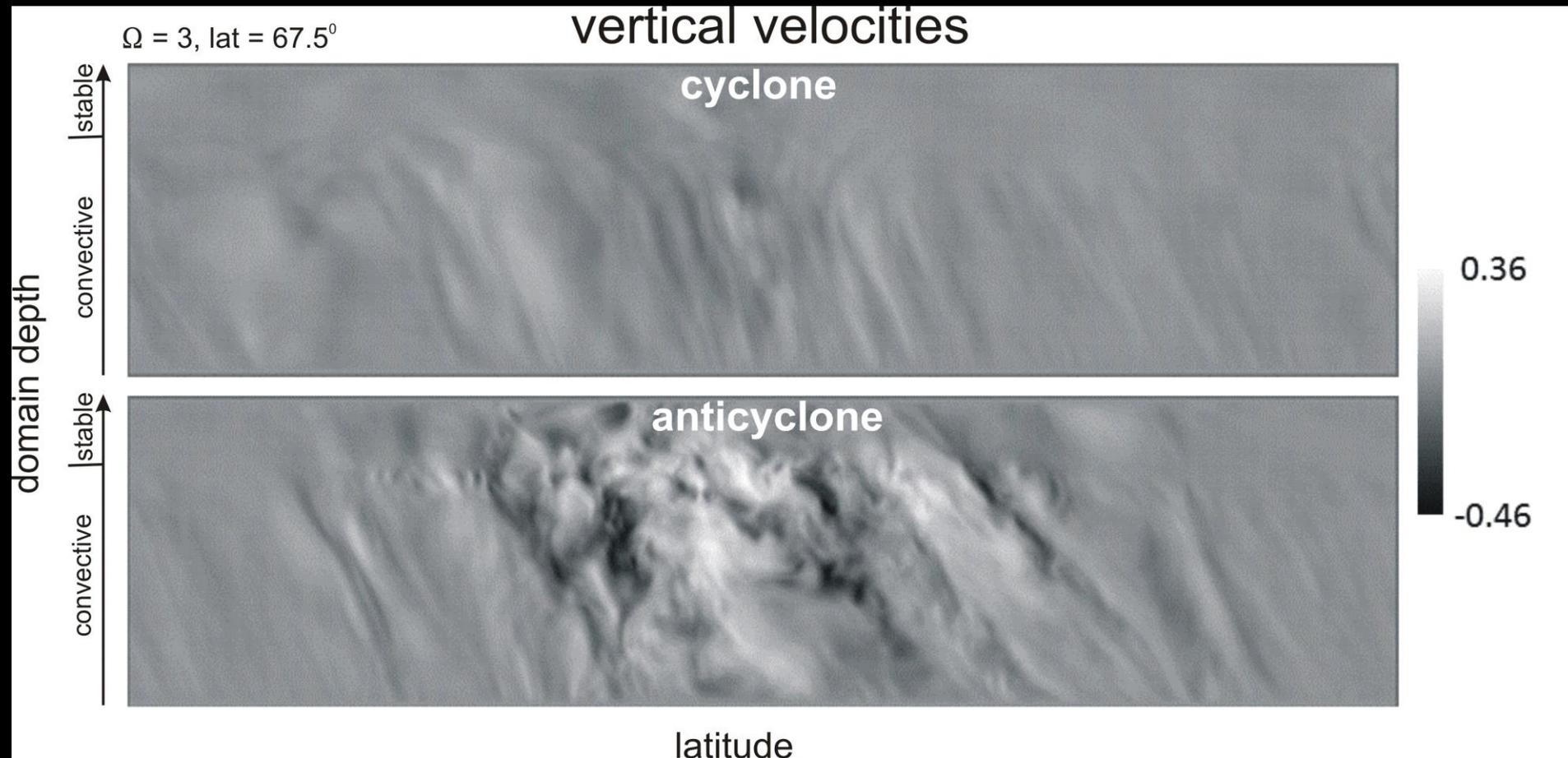
omega=3.



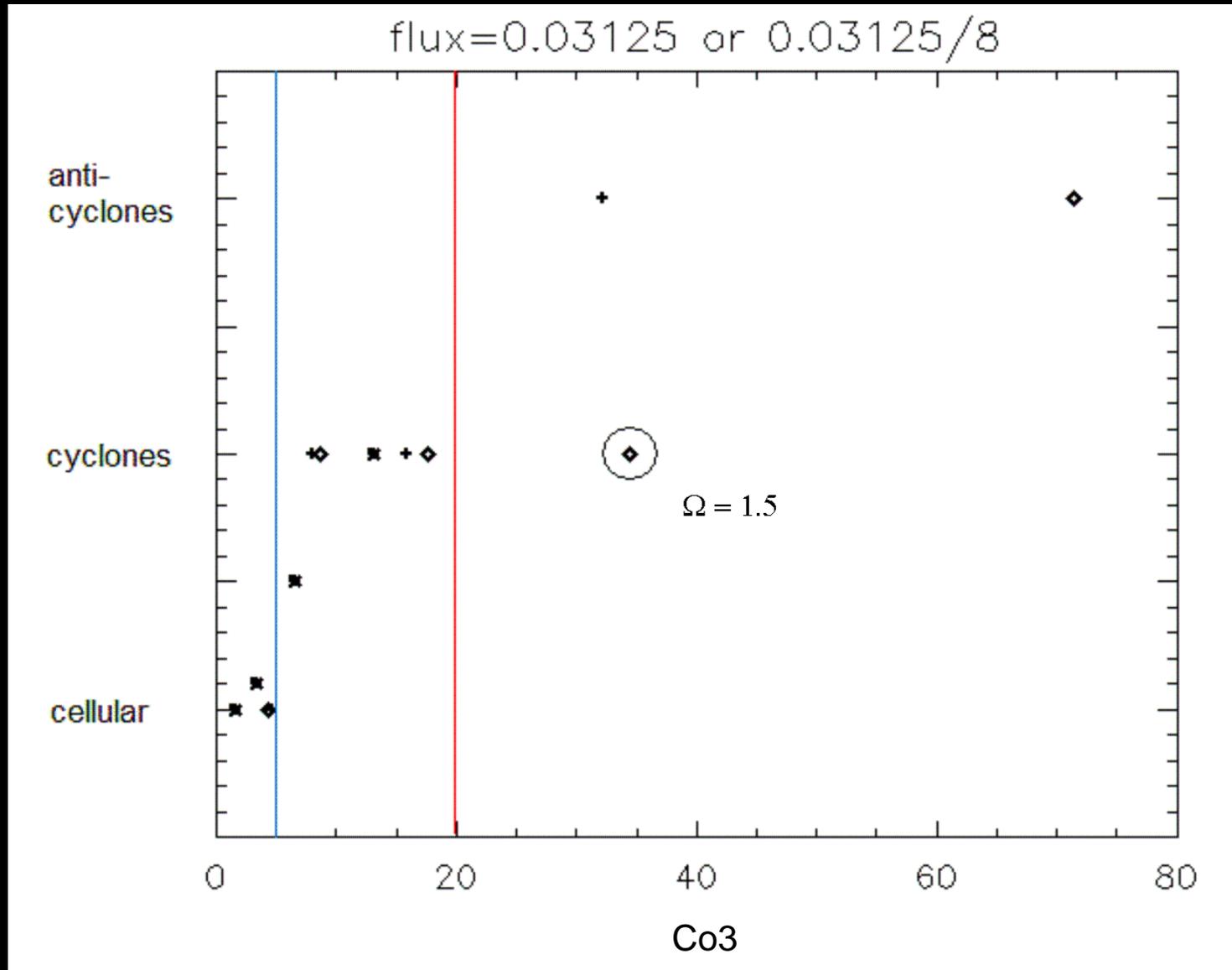
Regime Diagram



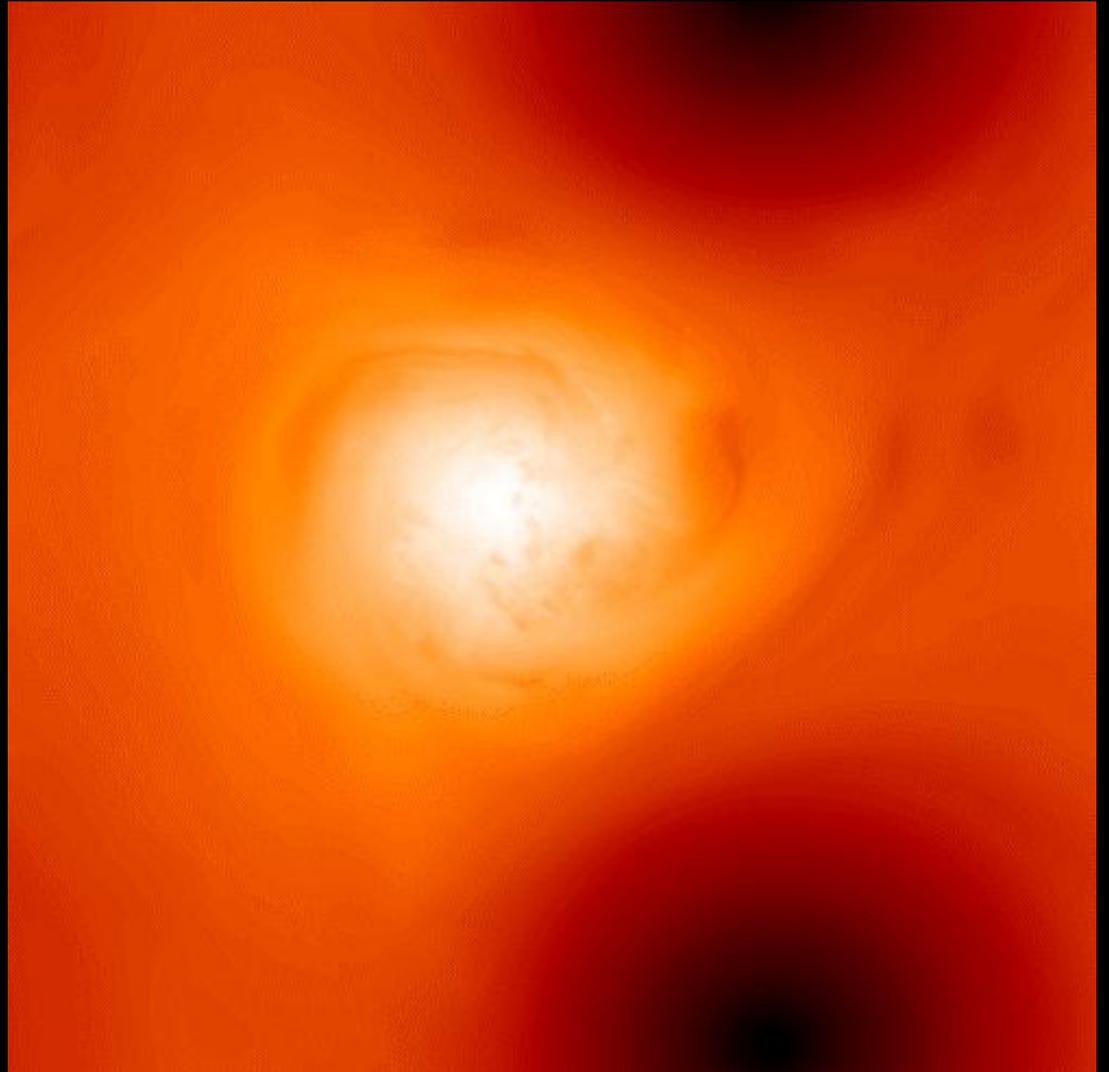
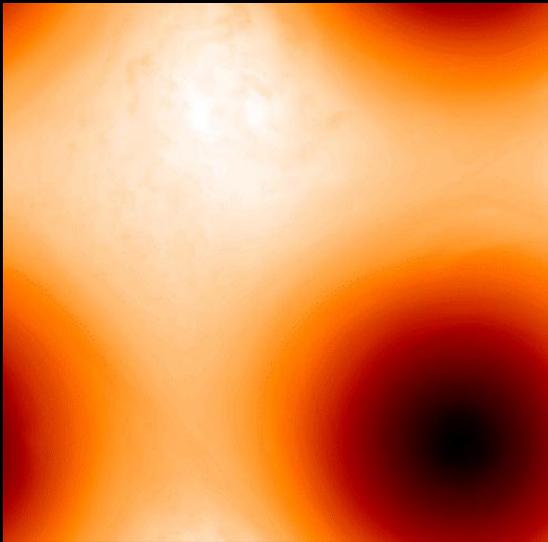
Vertical Cuts of Velocity Fields through Centers of Cyclone and Anticyclone



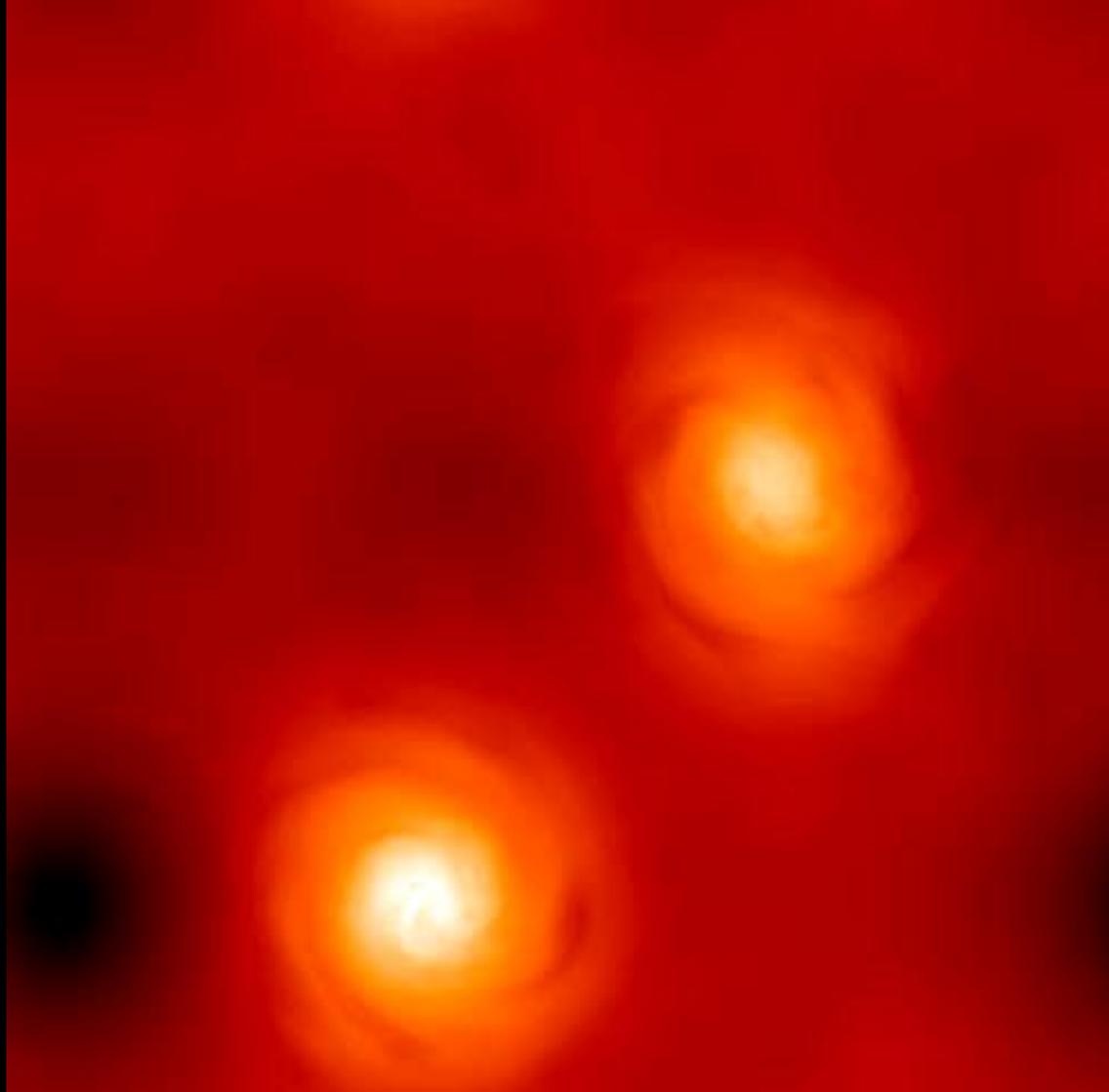
Regime Diagram (more cases)



$f = 0.0375/8$, $\Omega = 1.5$; aspect ratio = 4 & 8

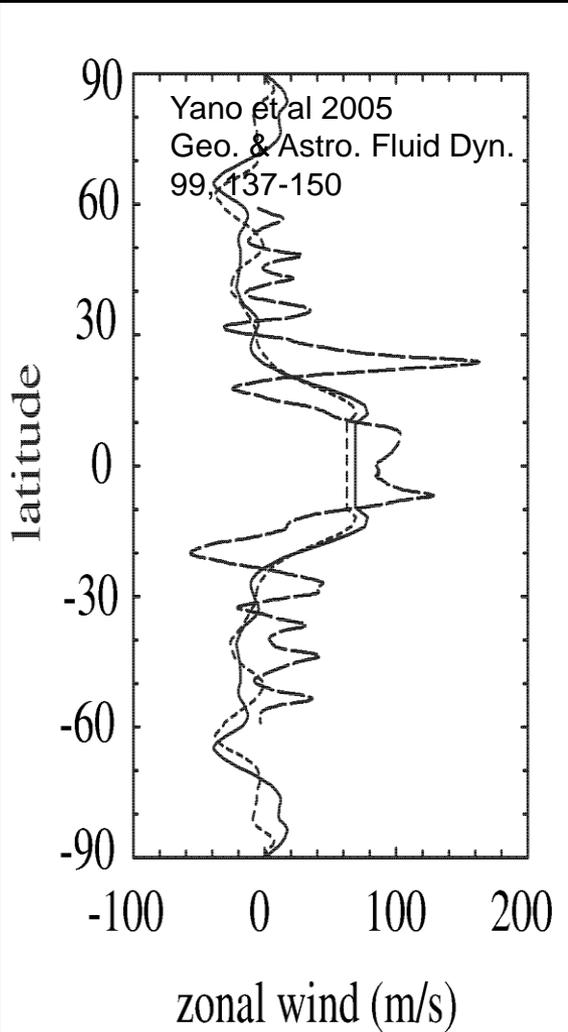


$f = 0.0375/8$, $\Omega = 3$; aspect ratio = 8

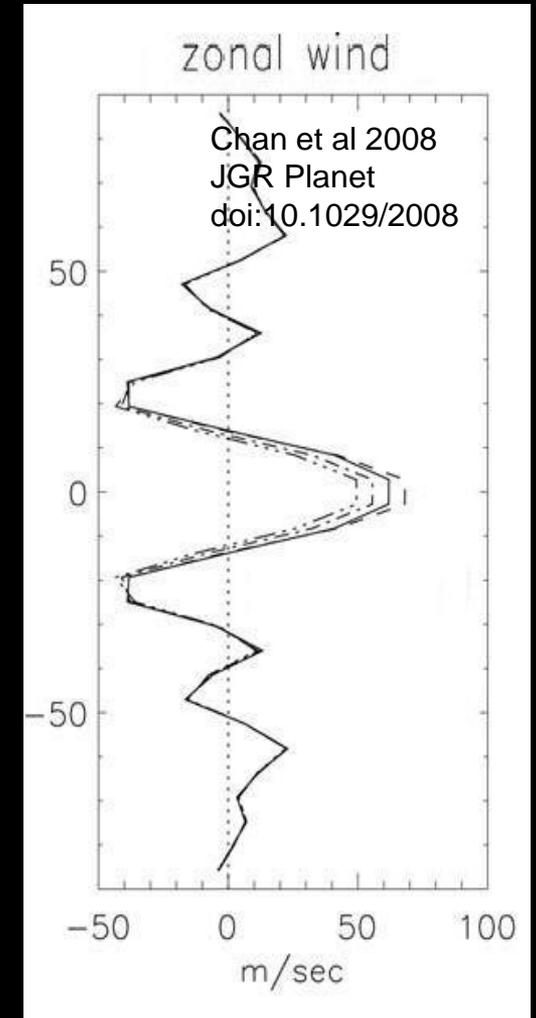
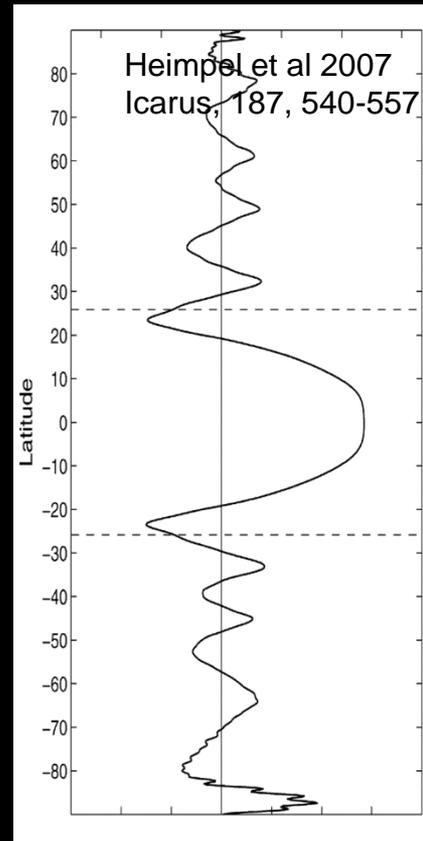


Zonal Wind Bands

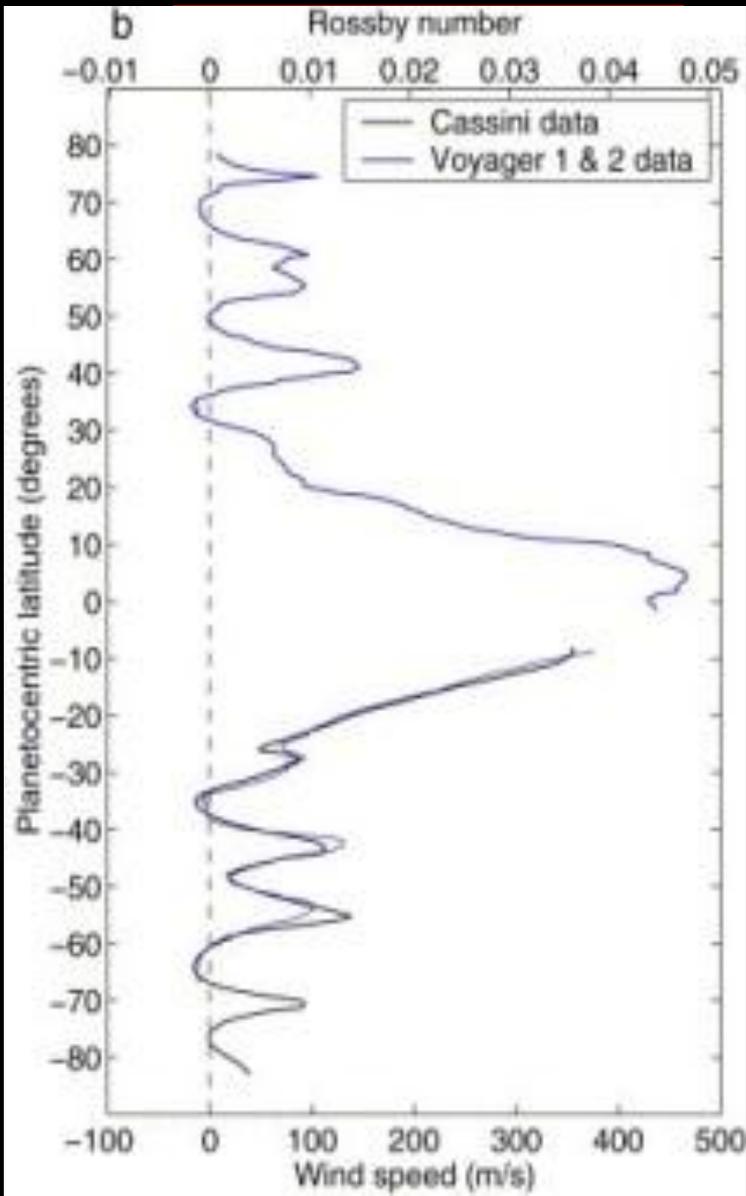
2D turbulence models



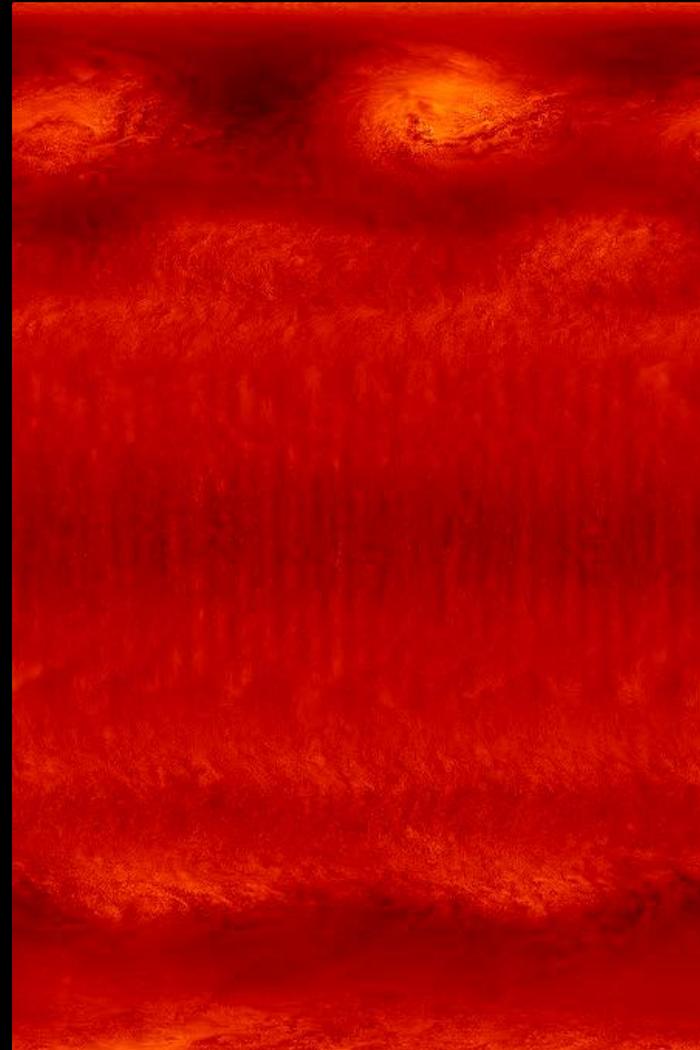
3D convection models



Multiple jets with prograde equatorial wind

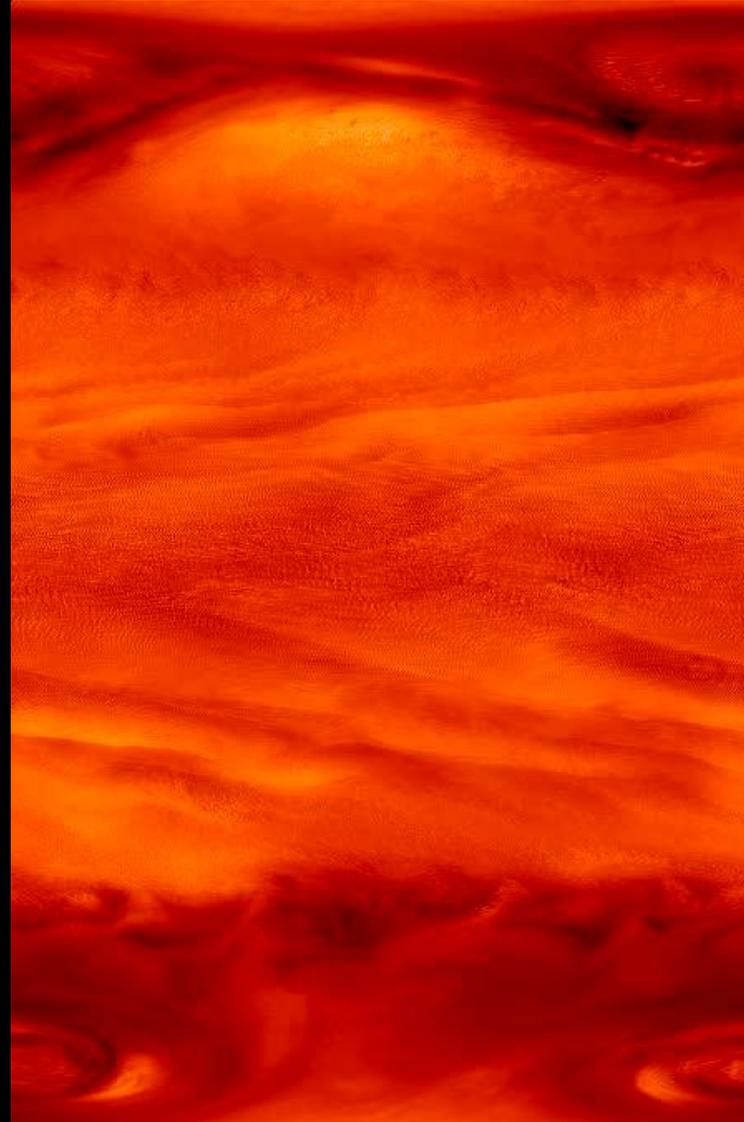
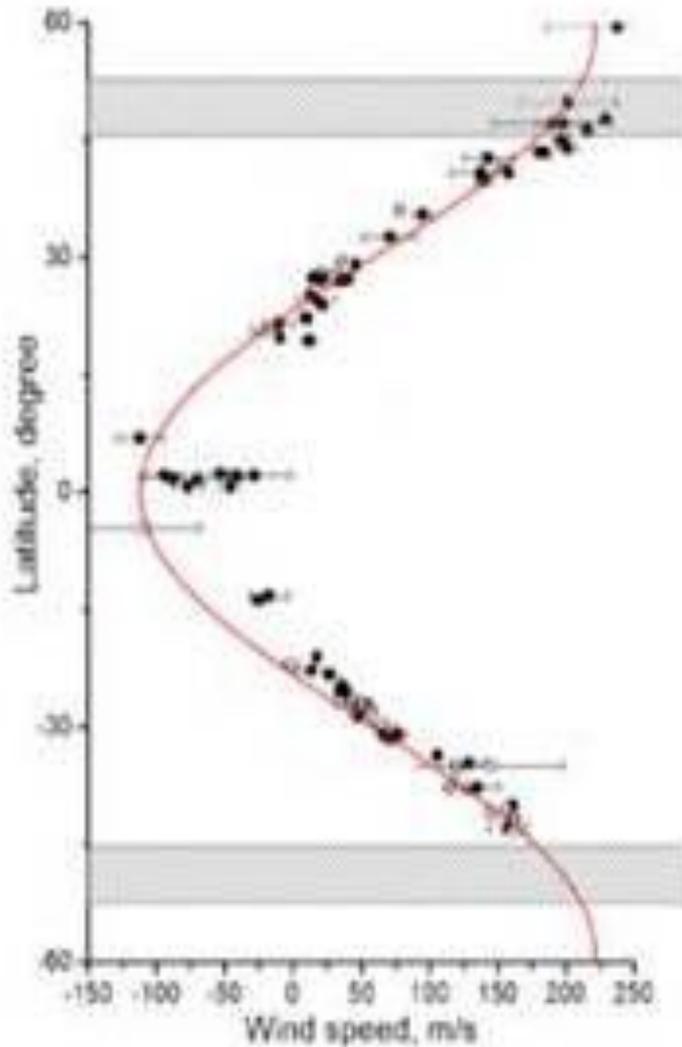


Horizontal cut of temperature



Broad Retrograde Equatorial Wind

Horizontal cut of temperature



Summary

- Long-lived cyclones and anticyclones can be generated in a rotating convection zone
- As Co increases, cyclones appear first
- Multiple zonal wind bands can be generated in a rotating convection shell
- Prograde (or retrograde) equatorial wind depends on Co .
- Long-lived vortices can be generated in co-existence with the zonal wind bands

Thank You!