



UiO : **Fysisk institutt**

Det matematisk-naturvitenskapelige fakultet

Antideuteron as signature for dark matter

Based on Master's Thesis at NTNU supervised by Michael Kachelrieß

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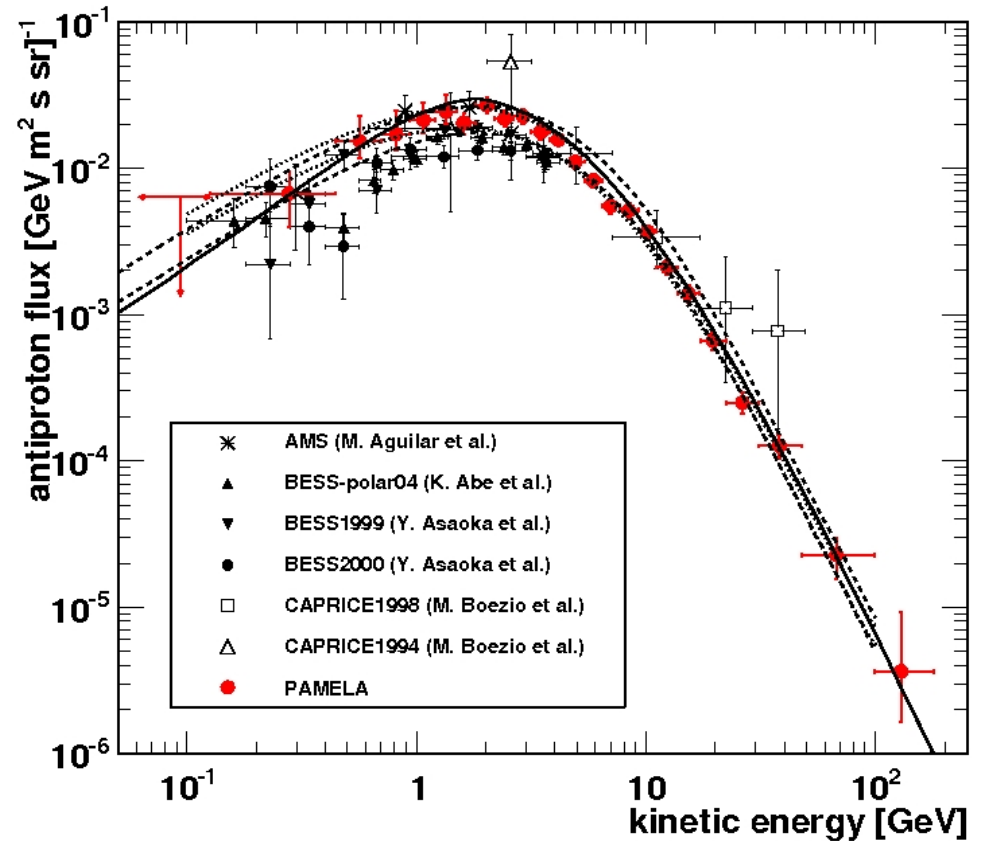
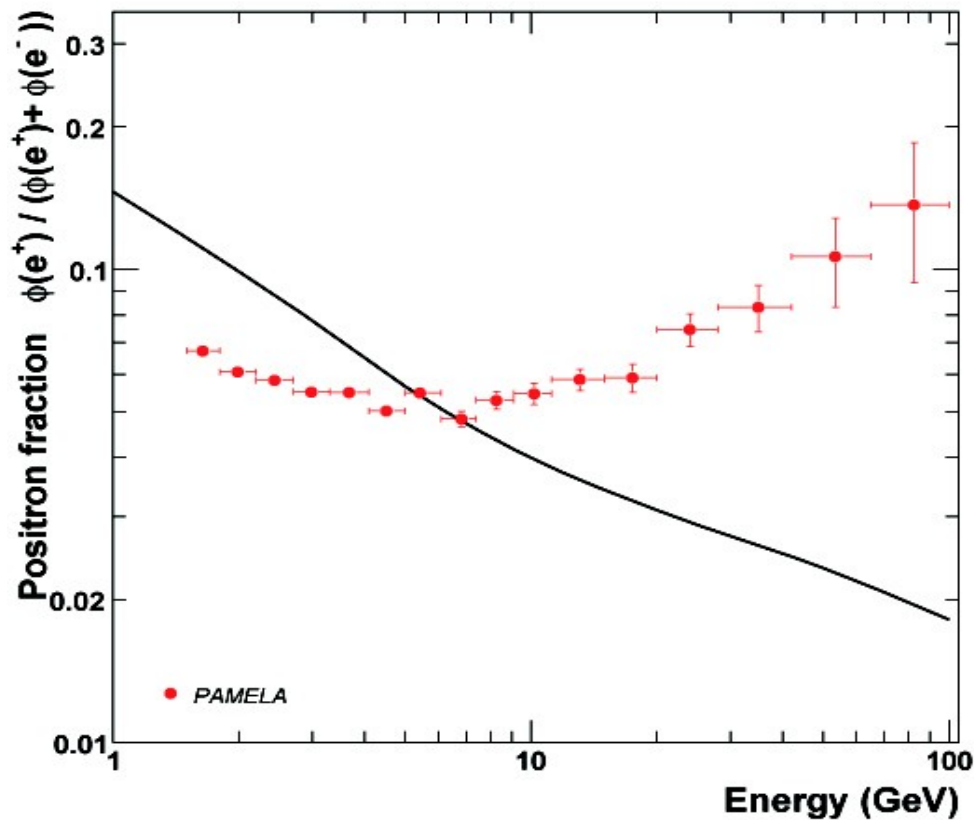
Detection of Dark Matter

- Assume DM in the form of unknown particles
- Most DM candidates have interactions which allow them to scatter, decay and/or annihilate
- 3 Ways to explore:
 - Direct searches: Particle Colliders
 - Direct detection: Direct Scattering
 - Indirect detection
- Indirect detection: Detection of DM annihilation/decay products (cosmic rays)

Suitable particle channels?

- What channels (particle species) are possible?
- DM is still around:
 - Long lifetime, low annihilation/decay rate
 - Weak signal
- Need channels where a signal is not “drowned”
- Charged VS. neutral particles: γ and ν
- Antimatter: Little around, seldom produced
- Positrons, antiprotons, antideuterons

The PAMELA data



- Antideuterons? Ongoing experiment: AMS-02

The Coalescence Model

- Need a model for nucleus formation
- Simple model: Coalescence
 - Nucleons with $\Delta p < p_0$ form a nucleon
 - p_0 calibrated against known processes
 - Developed in the 70's for heavy ion collisions
- Two approaches:
 - Direct approach: Applied per-event
 - Traditional approach: Applied to average energy spectra from multiple events

- Lab frame:

$$\underbrace{\left(\frac{dN_{\bar{d}}}{d^3k}\right)_{\vec{k}_{\bar{d}}}}_{\bar{d}\text{-density in momentum space}} = \frac{1}{8}\gamma \underbrace{\frac{4\pi p_0^3}{3} \left(\frac{dN_{\bar{p}}}{d^3k}\right)_{\vec{k}_{\bar{p}}=\vec{k}_{\bar{d}}/2}}_{\text{probability of finding } \bar{p} \text{ within sphere of radius } p_0 \text{ around } \vec{k}_{\bar{n}} \text{ in momentum space}} \cdot \underbrace{\left(\frac{dN_{\bar{n}}}{d^3k}\right)_{\vec{k}_{\bar{n}}=\vec{k}_{\bar{d}}/2}}_{\bar{n}\text{-density in momentum space}}$$

$$\left(\frac{dN_{\bar{d}}}{dx_{\bar{d}}}\right) = \frac{p_0^3}{6} \frac{m_{\bar{d}}}{M_{DM}^2 m_{\bar{n}} m_{\bar{p}}} \frac{1}{\sqrt{x_{\bar{d}}^2 + 2m_{\bar{d}}x_{\bar{d}}/M_{DM}}} \left(\frac{dN_{\bar{n}}}{dx_{\bar{n}}}\right) \left(\frac{dN_{\bar{p}}}{dx_{\bar{p}}}\right)$$

– $x_i = T_i / M_{DM}$, $x_{\bar{p}} = x_{\bar{n}} = x_{\bar{d}} / 2$

– Assumptions: Uncorrelated, isotropic distributions

- Why use the traditional approach?

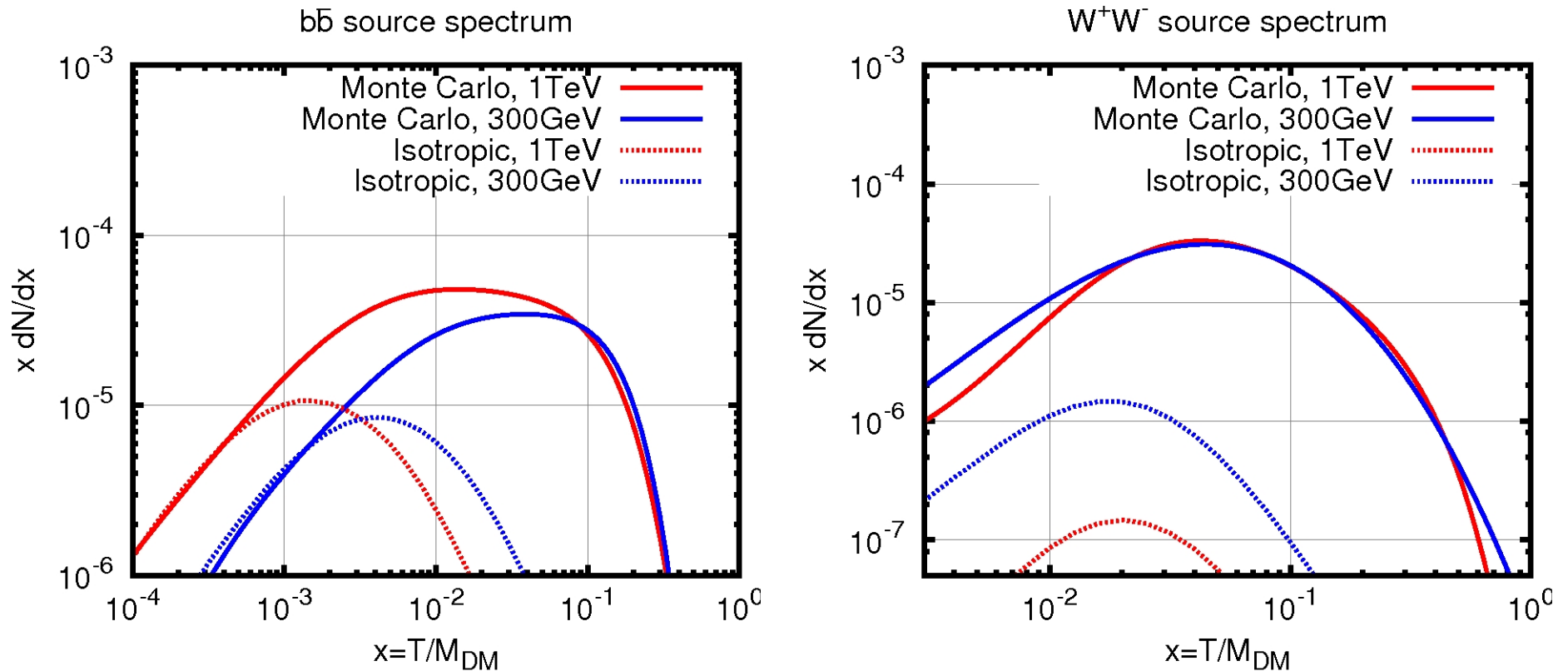
– \bar{d} production rate: $\sim 10^{-4}$

– \bar{n} , \bar{p} production rate: $\sim 10^0$

Simulations

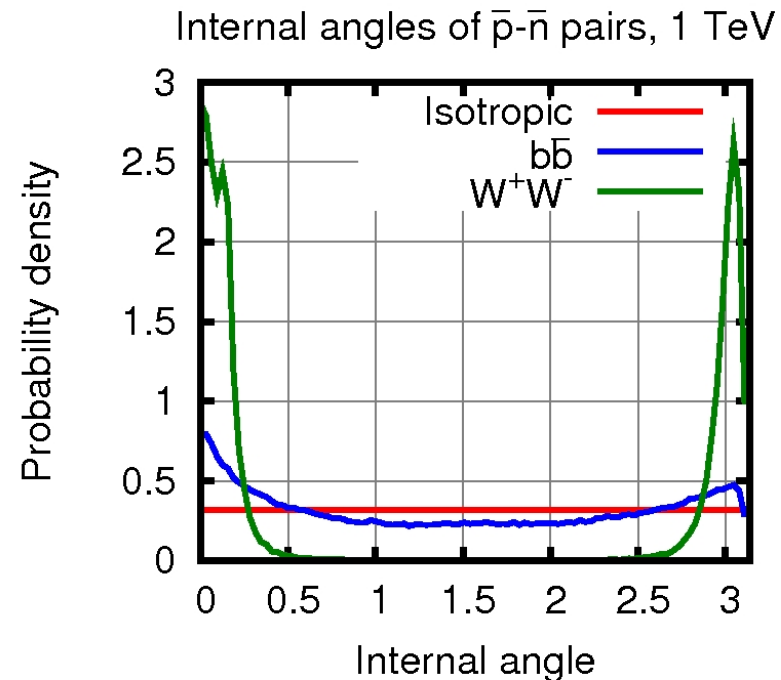
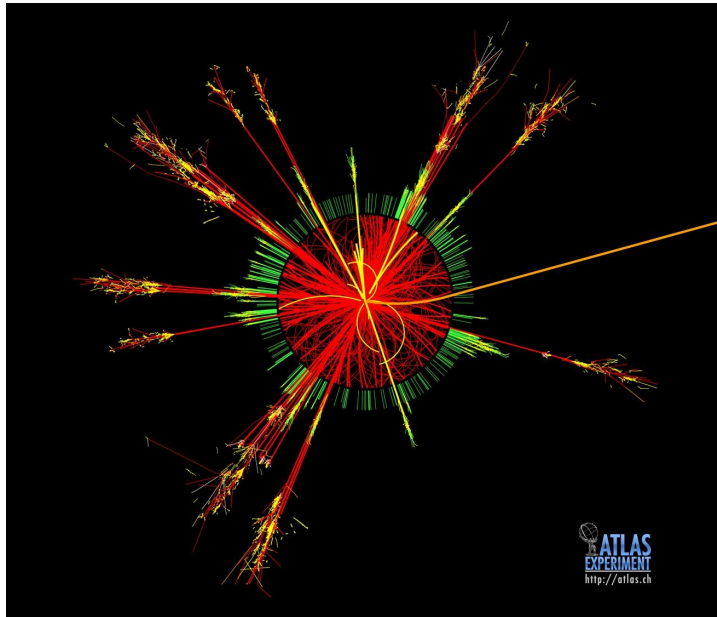
- Tree level annihilation processes
 - $X \bar{X} \rightarrow W^+ W^-$
 - $X \bar{X} \rightarrow b \bar{b}$
 - Spectra only depend on M_{DM} , not on the model
 - Total spectrum from branching ratios
 - Lightest MSSM neutralino
- Programs (Monte Carlo Generators)
 - MadGraph: Tree level process events
 - Herwig++: Showering, hadronization, (coalescence)

Antideuteron source spectra



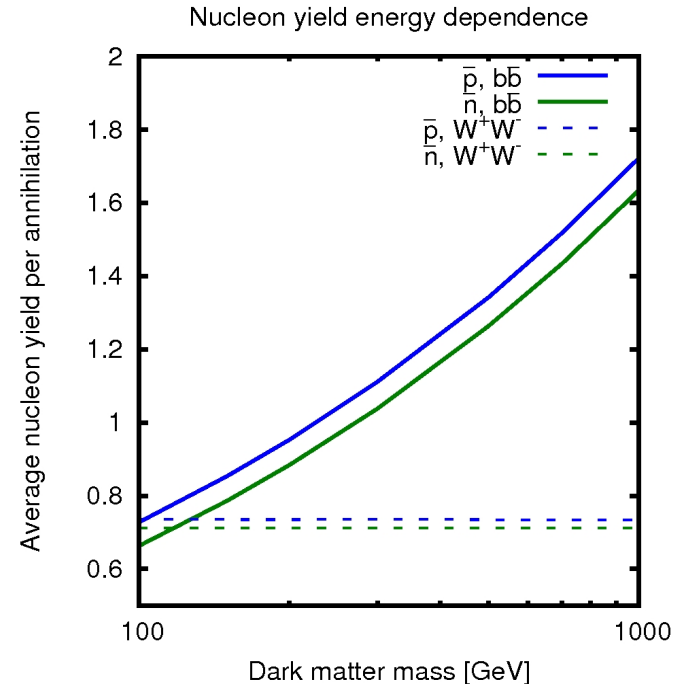
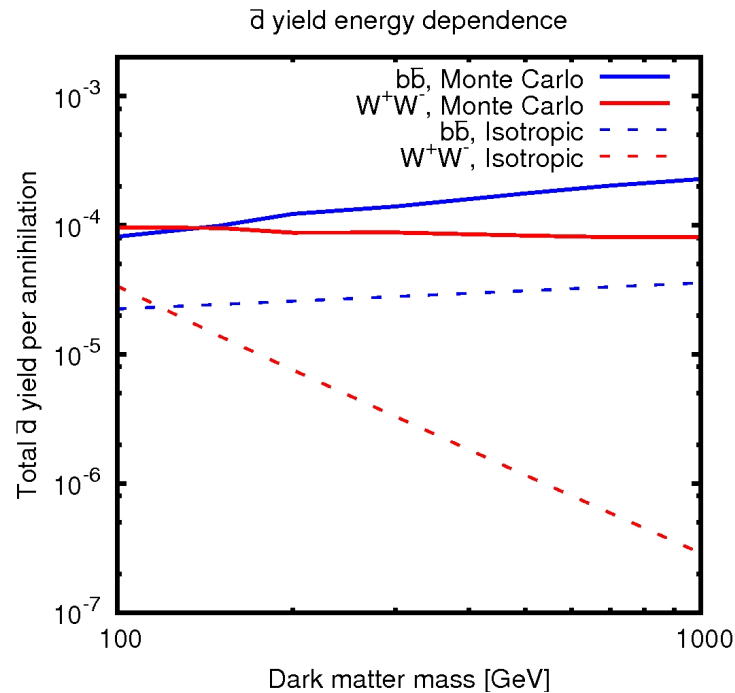
- Differences in the spectra at high energies
- DM mass dependence in the $W^+ W^-$ case

Bad assumptions



- Assumed uncorrelated, isotropic distributions
- QCD effects lead to confined jets
- Heavy ion collisions VS. DM annihilations

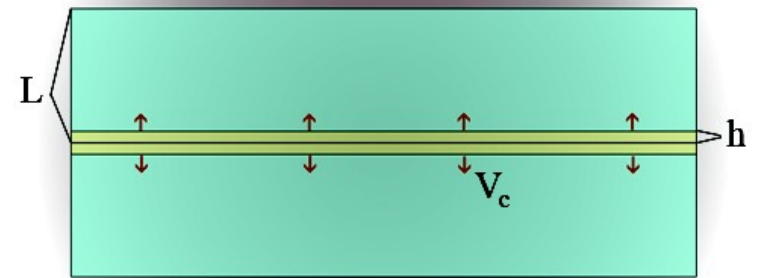
The gauge boson case



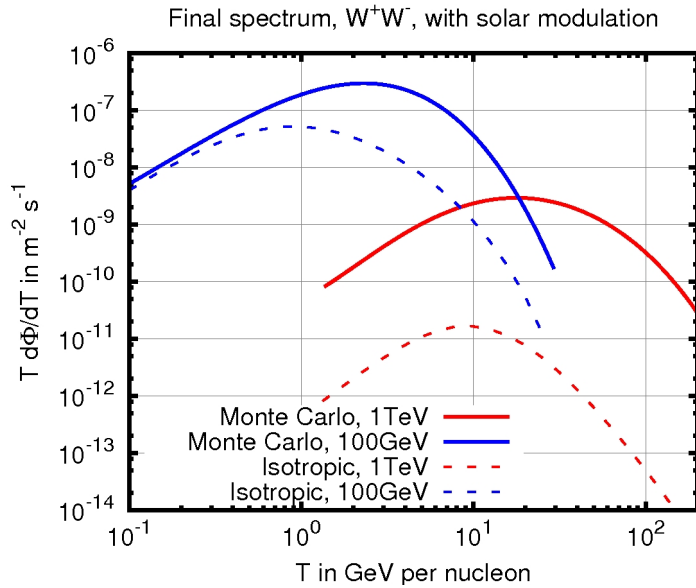
- MadGraph/Herwig++ connection:
 - Quarks always virtual
 - Gauge bosons taken to be real (on-shell)
 - Recurring mistake

Propagation through the Galaxy

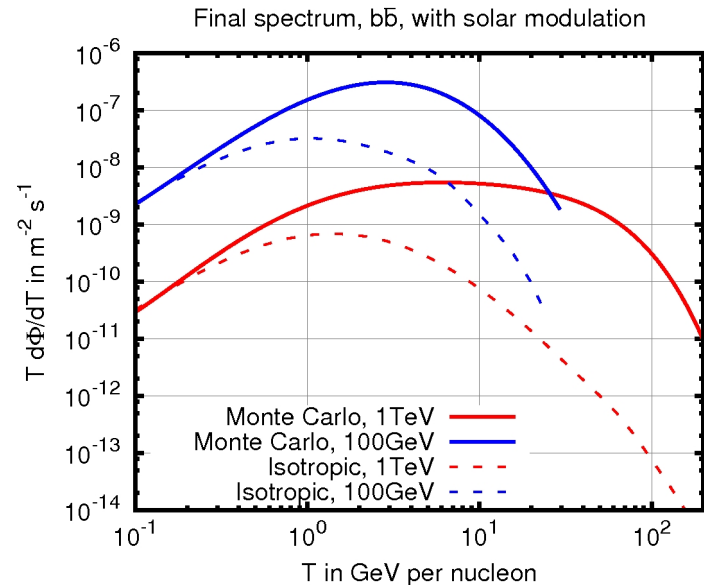
- Two-zone diffusion model
 - Cylindrical symmetry
 - Large cylinder
 - Diffusion of charged particles in magnetic field
 - Small cylinder (disk)
 - Annihilations on interstellar matter
 - Convective wind
 - Diffusion equation solved (mostly) numerically



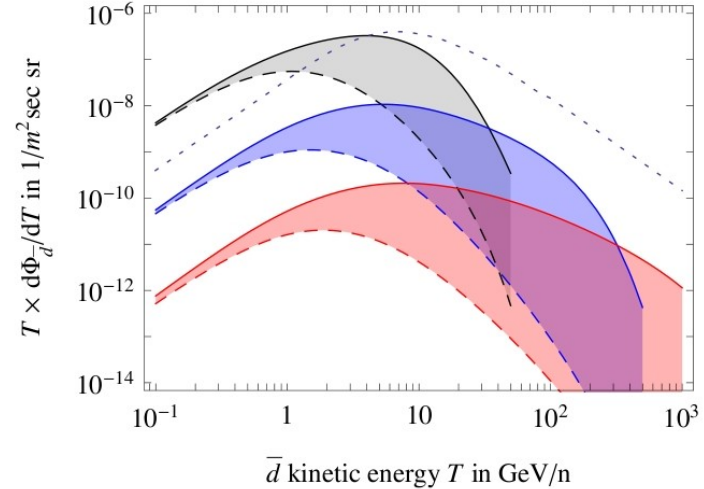
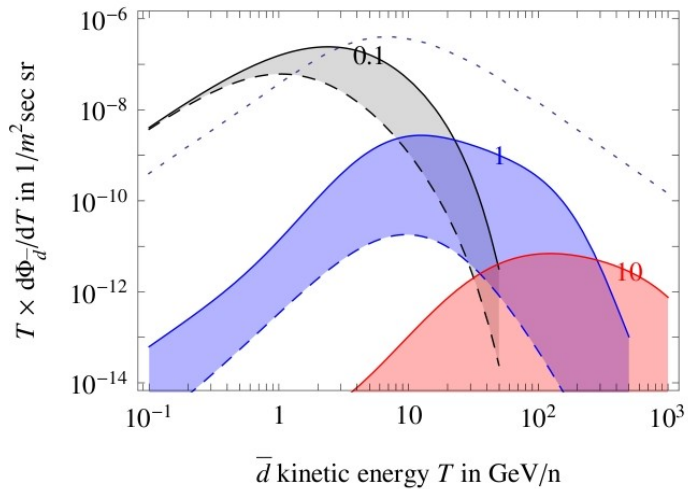
Final spectra



$DM DM \rightarrow W^+W^-$



$DM DM \rightarrow q\bar{q}$



Summary & current project

- Summary
 - The isotropic approximation in coalescence is not valid for DM annihilation processes
 - Recurring error when simulating spectra from annihilations into gauge bosons
 - The antideuteron channel is being measured by the ongoing AMS-02 experiment
- Current project
 - Use antideuteron spectrum from gravitino decays in R-parity violating SUSY models to set limits on R-parity violating couplings