

# An update on scalar singlet dark matter

Pat Scott

Department of Physics, McGill University

With: Jim Cline, Kimmo Kainulainen & Christoph Weniger  
(arXiv:1306.4710)

Slides available from <http://www.physics.mcgill.ca/~patscott>



# What is the scalar singlet model?

- Ultra-minimal dark matter model (Silveira & Zee, PLB 1985)
- 1 new scalar particle  $S$ , SM gauge singlet
  - 4 new (renormalizable) Lagrangian terms not forbidden by any symmetries:

$$\mathcal{L}_S = -\frac{\mu_S^2}{2} S^2 - \frac{\lambda_{hs}}{2} S^2 H^\dagger H - \frac{1}{2} \partial_\mu S \partial^\mu S - \frac{\lambda_S}{4} \lambda_S S^4 \quad (1)$$

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- $\partial_\mu S \partial^\mu S = S$  kinetic term,  $\lambda_S S^4 = S$  self-interaction
  - So long as  $\lambda_S \lesssim 1$  (to remain perturbative), only  $\mu_S^2 S^2$  and  $\lambda_{hs} S^2 H^\dagger H$  matter for phenomenology



# Consequences

After  $H$  gets a VEV  $v_0$  during electroweak symmetry-breaking,  $\lambda_{hs} S^2 H^\dagger H$  induces  $hSS$  and  $hhSS$  terms:

$$\mathcal{L}_{S^2|H|^2} \rightarrow O_{hSS} + O_{hhSS} = -\frac{\lambda_{hs} v_0}{2} hSS - \frac{\lambda_{hs}}{4} hhSS \quad (2)$$

Introduces two new interaction vertices:



$S$  gets mass contributions from bare term  $\mu_S^2 S^2$  and new  $SS$  term also induced by  $\lambda_{hs} S^2 H^\dagger H$ :

$$m_S = \sqrt{\mu_S^2 + \frac{\lambda_{hs} v_0^2}{2}} \quad (3)$$



# Advantages

## Question

Why is the scalar singlet model attractive?



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Why is the scalar singlet model attractive?

## Answers

- Occam's razor: super-simple, just one singlet scalar and a  $Z_2$  symmetry
- 2 parameters only:  $\lambda_{hs}$ ,  $m_S$  (or trade for  $\mu_S$  if you prefer)
- All phenomenology fully calculable, some extremely straightforwardly
- Predictive and very testable





# Advantages

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Can it solve the hierarchy problem / give me 130 GeV lines / low-mass DM / make my lunch?



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## Answers

Not really. . .

- any hierarchy 'solution' would require fine-tuning and probably not hold at higher orders anyway
- 130 GeV line requires additional charged scalar
- Low-mass direct detection is conceivable, but parameters are ruled out (**more later**)
- Electroweak baryogenesis works, but not with  $S$  as all DM (**more later**)



# Phenomenology

Plenty of phenomenological studies over the years. . .

- Pre-LHC investigations of collider and dark matter detection prospects

(e.g. McDonald hep-ph/0106249, Burgess et al. hep-ph/0011335, Patt & Wilczek hep-ph/0605188, Barger et al. 0706.4311, 1008.1796)

- As an explanation for DAMA/CoGeNT/CRESST

(Andreas et al. 0808.0255, 1003.2595, Tytgat 1012.0576)

- Indirect detection prospects

(Yaguna 0810.4267, Goudelis et al. 0909.2799, Arina & Tytgat 1007.2765)

- As a way to achieve baryogenesis

(Profumo et al. 0705.2425, Barger et al. 0811.0393, Cline & Kainulainen 1210.4196)

- Impacts of early LHC searches, XENON-100

(e.g. Mambrini 1108.0671, Djouadi et al. 1205.3169)



# Phenomenology

## Question

Why the update?



# Phenomenology

## Question

Why the update?

## Answers

- Include 125 GeV Higgs
- Update to new LHC limits
- Update to new XENON100 limits
- Add *Fermi* dwarf and CMB limits (best ID limits available)
- Treat models with sub-dominant relic densities consistently
- Higgs-nucleon coupling now much better understood from lattice calculations
- Use full thermal-average  $\langle\sigma v\rangle$  for relic density



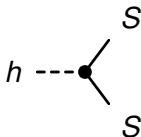
# Outline

- 1 Background
- 2 Updated constraints
  - LHC: Invisible Higgs width
  - Relic density and indirect detection
  - Direct detection



# Dude, where's my Higgs?

If  $S$  is light enough, the Higgs decay  $h \rightarrow SS$  is kinematically allowed:



with rate:

$$\Gamma_{h \rightarrow SS} = \frac{\lambda_{hs}^2 v_0^2}{32\pi m_h^2} \left( m_h^2 - 4m_S^2 \right)^{1/2} \quad (4)$$

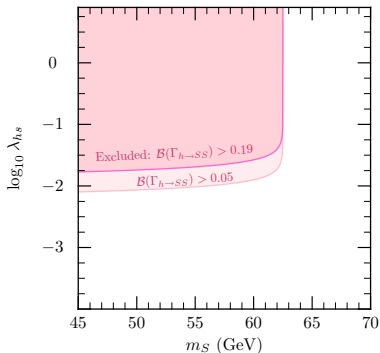
- Higgs production rates (e.g. via  $gg$  fusion, vector boson fusion) can be compared to observed  $h$  decay rates (e.g.  $\gamma\gamma$ ,  $b\bar{b}$ , etc) to test for missing decays
- Latest 95% CL combined LHC+Tevatron limit across all production and decay channels is  $BF_{SS} < 19\%$  (Belanger et al.

1306.2941)



# Constraints

Resulting LHC+Tevatron limits:



300 fb<sup>-1</sup> 14 TeV LHC will manage BF limit of 5% (Peskin, arXiv:1207.2516)





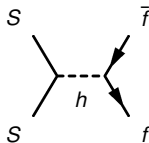
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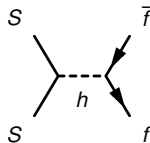
# Annihilation Channels

$$SS \rightarrow f\bar{f}:$$

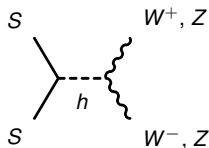


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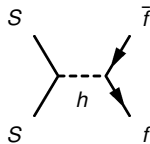


$$SS \rightarrow V\bar{V}:$$

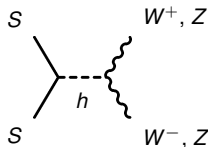


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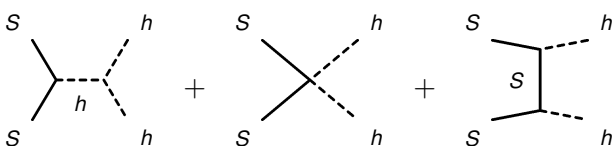
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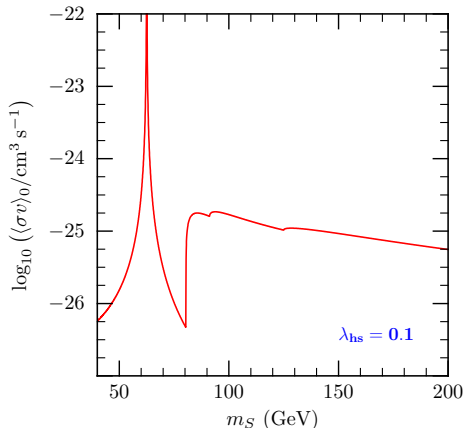
$$SS \rightarrow V\bar{V}:$$



$$SS \rightarrow hh:$$



## Annihilation cross-section

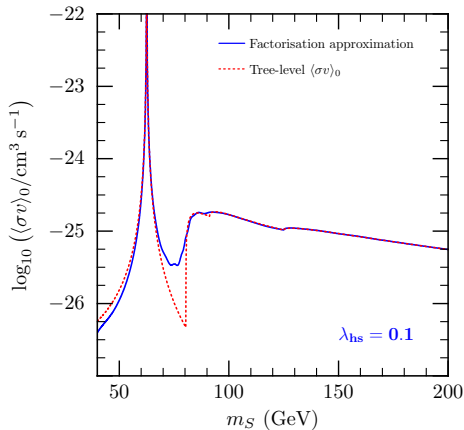
Resulting  $\langle\sigma v\rangle_0$ :

- Higgs resonance:  
 $m_S \sim m_h/2 = 62.5$  GeV
- $W, Z$  thresholds: 4-body final states from  
 $SS \rightarrow V^* \bar{V}^* \rightarrow f\bar{f} f\bar{f}$   
become significant just below  $m_{W,Z}$



# Annihilation cross-section

Resulting  $\langle\sigma v\rangle_0$ :



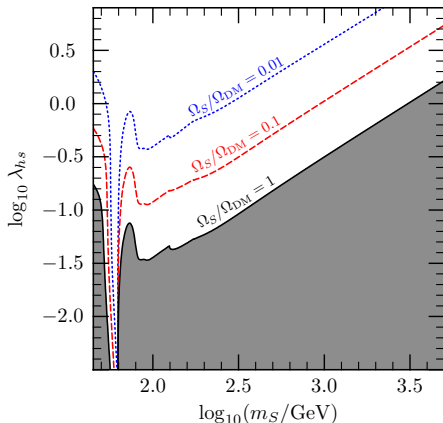
→ better to factor non- $hh$  channels into  $SSh$  fusion part  $\times$  full Higgs decay width  $\Gamma_h$  to all SM particles:

$$\begin{aligned} \langle\sigma v\rangle_0 &= \langle\sigma v\rangle_{0,hh} + \langle\sigma v\rangle_{0,\text{others}} \\ &= \langle\sigma v\rangle_{0,hh} + \\ &\quad \frac{\lambda_{hs}^2 v_0^2 \Gamma_h}{m_S [(4m_S^2 - m_h^2)^2 + m_h^2 \Gamma_h^2]} \end{aligned} \quad (5)$$



# Relic density

Resulting relic densities:



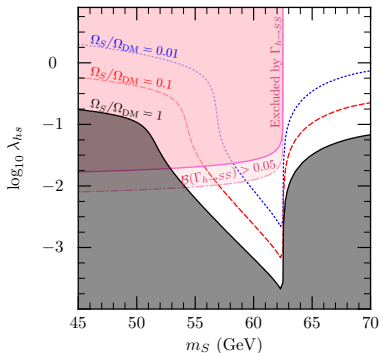
- When  $m_S \gtrsim m_h/2$ : OK to just get relic density estimate direct from  $\langle \sigma v \rangle_0$   
(a la Steigman et al. 1204.3622)
- When  $m_S < h/2$ : must include thermal effects  
→ relic density in this region from explicit freeze out calculation (just assuming instantaneous thermalisation – gives a conservative limit)



# Relic density

Together with Higgs invisible width, rules out all models with  $m_S < 53$  GeV

$\Rightarrow$  DAMA/CoGeNT/CRESST not seeing scalar singlet DM





# Indirect Detection Limits

## 1 *Fermi*-LAT 3-year combined dwarf gamma-ray limits

(Geringer-Sameth & Koushiappas 1108.2914, *Fermi*-LAT Collab 1108.3546)

- Implemented from *full set of CLs* on gamma-ray flux particle physics parameter  $\Phi_{PP}$  (1108.2914; kindly provided by Alex Geringer-Sameth)

## 2 CMB constraints on energy injection (WMAP7 / Planck polarisation)

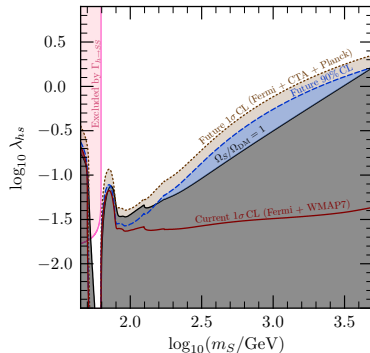
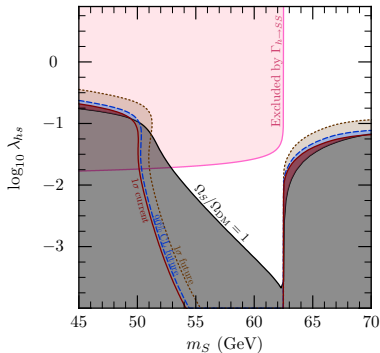
- Implemented with tabulated likelihoods for all SM final states (Cline & PS 1301.5908)

→ Model-by-model *Fermi* + CMB combination of limits from all final states, according to actual branching fractions as each  $\{m_S, \lambda_{hs}\}$ .



# Indirect Detection Limits

Resulting ID limits:



- Current ID reveals slight tension at  $\geq 1\sigma$  level in small parts of parameter space, but not much better than that
- Future searches (*Fermi* 10 yr, 20 dwarfs + Planck pol. + CTA 500 hr GC) will exclude more at 90% CL; yet more will be in  $\geq 1\sigma$  tension.

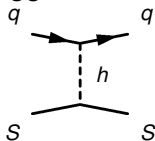


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Higgs portal interactions give spin-independent nuclear scattering via  $t$ -channel Higgs exchange:



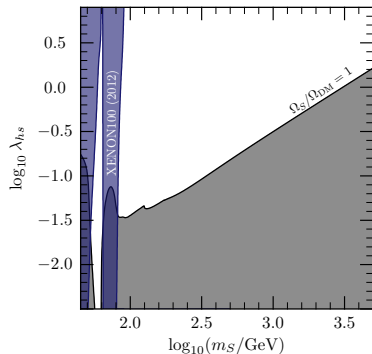
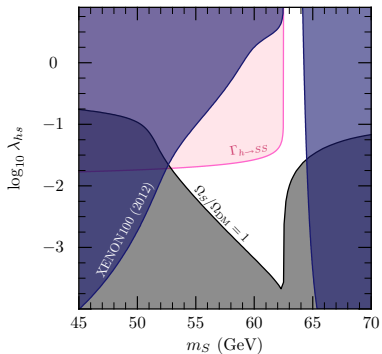
Cross-section in terms of  $S$ - $N$  reduced mass  $\mu_N$  is quite simple:

$$\sigma_{\text{SI},N} = \frac{\lambda_{hs}^2 f_N^2 \mu_N^2 m_N^2}{4\pi m_h^4 m_S^2}. \quad (6)$$

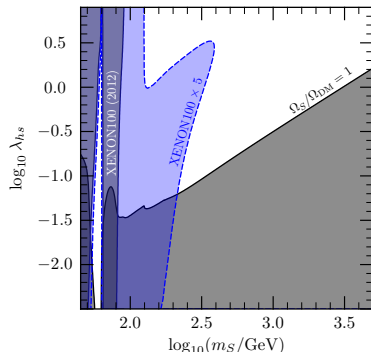
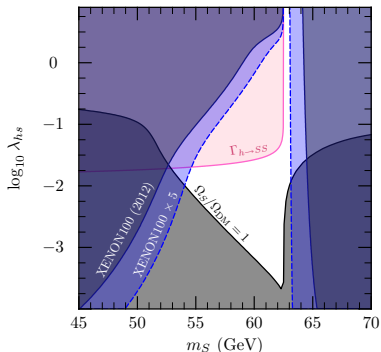
- Pion nuclear sigma uncertainty recently reduced (ChPT, lattice vals  $\uparrow$  to  $\sim 60$  MeV – matches experiment)
  - strange quark content of nucleons now not as important
  - Higgs- $N$  effective coupling now  $f_N \sim 0.345 \pm 0.02$
- Following results just use standard Maxwellian halo.



## Resulting direct limits:



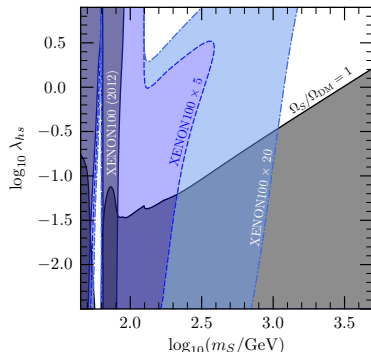
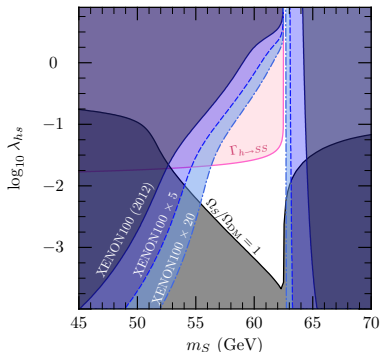
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- ...yet a small window at  $m_S \sim 60$  GeV will remain viable.



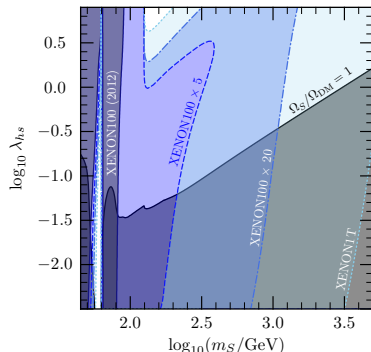
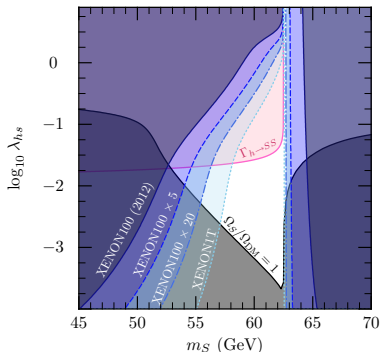
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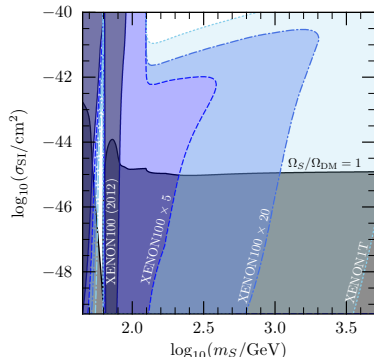
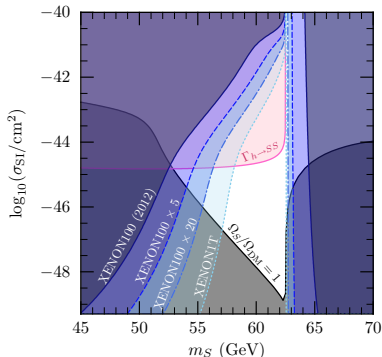


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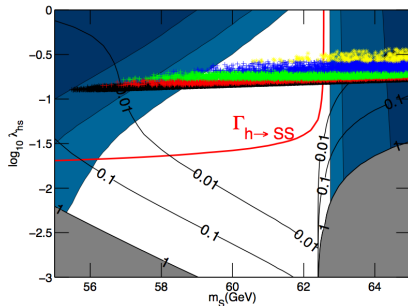
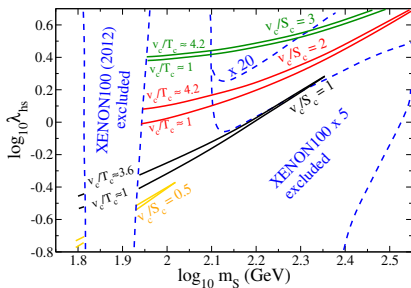


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Electroweak baryogenesis with  $S$ 

- Electroweak baryogenesis (EWBG) is possible with singlet model, but not at the same time as  $\Omega_S/\Omega_{\text{DM}} = 1$
- Future DD will squeeze the EWBG parameter space severely.



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- Signals expected in LHC, direct detection and indirect detection experiments
- Constraints on parameter space from DD, ID and LHC are highly complimentary
- Model is excluded as explanation for low-mass DM hints/anomalies
- Upcoming experiments (DD especially) will access large parts of the remaining parameter space

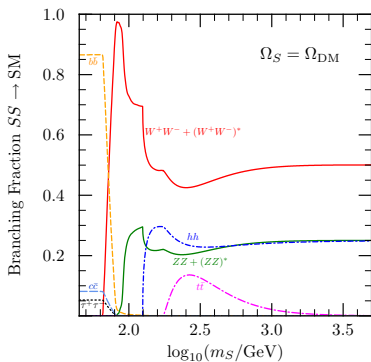


## Backup Slides





# Branching Fractions



## Contributions to ID limits

