Dark matter searches with ATLAS

SN

SN

ATLAS

David Berge (GRAPPA & Nikhef, Amsterdam) For the ATLAS Collaboration





The Large Hadron Collider



1331	2520		
4 TeV x 4 TeV	6.5 TeV x 6	5 TeV	
50 ns	25 ns		
7.7 x 10 ³³ cm ⁻² s ⁻¹	1.5 x 10 ³⁴ c	m ⁻² s ⁻¹	
~35	~40		
	PS		
SPS			
	4 TeV x 4 TeV 50 ns 7.7 x 10 ³³ cm ⁻² s ⁻¹ ~35 SPS SPS	1331 2520 4 TeV x 4 TeV 6.5 TeV x 6. 50 ns 25 ns 7.7 x 10 ³³ cm ⁻² s ⁻¹ 1.5 x 10 ³⁴ ct ~35 ~40 SPS PS SPS SPS	1331 2520 4 TeV x 4 TeV 6.5 TeV x 6.5 TeV 50 ns 25 ns 7.7 x 10 ³³ cm ⁻² s ⁻¹ 1.5 x 10 ³⁴ cm ⁻² s ⁻¹ ~35 ~40

Task: measure transverse energy



1: "Standard" Dark Matter Searches at Colliders 5



SUSY:

- High- p_T jets from squark & gluino decays
- Leptons from gaugino & slepton decays
- Missing transverse energy from LSPs



ATLAS SUSY Searches* - 95% CL Lower Limits

Status: Moriond 2014

ATLAS Preliminary

 $\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1} \qquad \sqrt{s} = 7, 8 \text{ TeV}$

		Model	e, μ, τ, γ	Jets	$E_{\rm T}^{\rm miss}$	∫ <i>L dt</i> [fb	⁻¹] Mass limit	Reference
Incl. searches	Inclusive Searches	$ \begin{array}{l} \text{MSUGRA/CMSSM} \\ \text{MSUGRA/CMSSM} \\ \text{MSUGRA/CMSSM} \\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow q \tilde{\chi}_{10}^0 \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow q \bar{q} \tilde{\chi}_{1}^0 \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow q \bar{q} \tilde{\chi}_{1}^1 \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow q q \tilde{\chi}_{1}^\pm \rightarrow q W^\pm \tilde{\chi}_{1}^0 \\ \text{GMSB} (\tilde{\ell} \text{ NLSP}) \\ \text{GMSB} (\tilde{\ell} \text{ NLSP}) \\ \text{GGM} (\text{bino NLSP}) \\ \text{GGM} (\text{higo NLSP}) \\ \text{GGM} (\text{higosino NLSP}) \\ \text{GGM} (\text{higosino NLSP}) \\ \text{GGM} (\text{higosino NLSP}) \\ \text{GGM} (\text{higosino NLSP}) \\ \text{Gravitino LSP} \\ \end{array} $	$\begin{array}{c} 0 \\ 1 \ e, \mu \\ 0 \\ 0 \\ 1 \ e, \mu \\ 2 \ e, \mu \\ 1 - 2 \ \tau \\ 2 \ \gamma \\ 1 \ e, \mu + \gamma \\ \gamma \\ 2 \ e, \mu (Z) \\ 0 \end{array}$	2-6 jets 3-6 jets 7-10 jets 2-6 jets 2-6 jets 3-6 jets 0-3 jets 2-4 jets 0-2 jets 1 b 0-3 jets mono-jet	Yes Yes Yes Yes Yes Yes Yes Yes Yes Yes	20.3 20.3 20.3 20.3 20.3 20.3 20.3 4.7 20.7 20.7 20.3 4.8 4.8 5.8 10.5	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	ATLAS-CONF-2013-047 ATLAS-CONF-2013-062 1308.1841 ATLAS-CONF-2013-047 ATLAS-CONF-2013-047 ATLAS-CONF-2013-062 ATLAS-CONF-2013-026 ATLAS-CONF-2013-026 ATLAS-CONF-2014-001 ATLAS-CONF-2012-144 1211.1167 ATLAS-CONF-2012-152 ATLAS-CONF-2012-152
	3 rd gen. <u>§</u> med.	$ \begin{split} \tilde{g} &\rightarrow b \bar{b} \tilde{\chi}_{1}^{0} \\ \tilde{g} &\rightarrow t \bar{t} \tilde{\chi}_{1}^{0} \\ \tilde{g} &\rightarrow t \bar{t} \tilde{\chi}_{1} \\ \tilde{g} &\rightarrow b \bar{t} \tilde{\chi}_{1}^{+} \end{split} $	0 0 0-1 <i>e</i> ,μ 0-1 <i>e</i> ,μ	3 <i>b</i> 7-10 jets 3 <i>b</i> 3 <i>b</i>	Yes Yes Yes Yes	20.1 20.3 20.1 20.1	ğ 1.2 TeV m($\tilde{\chi}_1^0$)<600 GeV // ğ 1.1 TeV m($\tilde{\chi}_1^0$)<350 GeV // ğ 1.34 TeV m($\tilde{\chi}_1^0$)<400 GeV // ğ 1.3 TeV m($\tilde{\chi}_1^0$)<600 GeV //	ATLAS-CONF-2013-061 1308.1841 ATLAS-CONF-2013-061 ATLAS-CONF-2013-061
atural SUSY	3 rd gen. squarks direct production	$ \frac{\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow b \tilde{x}_1^0}{\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow t \tilde{x}_1^{\pm}} $ $ \frac{\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow t \tilde{x}_1^{\pm}}{\tilde{t}_1 \tilde{t}_1 (\text{light}), \tilde{t}_1 \rightarrow b \tilde{x}_1^{\pm}} $ $ \tilde{t}_1 \tilde{t}_1 (\text{light}), \tilde{t}_1 \rightarrow b \tilde{t}_1^{\pm} $ $ \tilde{t}_1 \tilde{t}_1 (\text{medium}), \tilde{t}_1 \rightarrow t \tilde{x}_1^0 $ $ \tilde{t}_1 \tilde{t}_1 (\text{meavy}), \tilde{t}_1 \rightarrow t \tilde{x}_1^0 $ $ \tilde{t}_1 \tilde{t}_1 (\text{heavy}), \tilde{t}_1 \rightarrow t \tilde{x}_1^0 $ $ \tilde{t}_1 \tilde{t}_1 (\text{heavy}), \tilde{t}_1 \rightarrow t \tilde{x}_1^0 $ $ \tilde{t}_1 \tilde{t}_1 (\text{neavy}) = \delta \tilde{t}_1^0 $ $ \tilde{t}_1 \tilde{t}_1 (\text{neavy}) = \delta \tilde{t}_1^0 $ $ \tilde{t}_1 \tilde{t}_1 (\text{neavy}) = \delta \tilde{t}_1^0 $ $ \tilde{t}_1 \tilde{t}_1 (\text{neav}) = \delta \tilde{t}_1^0 $	$\begin{matrix} 0 \\ 2 \ e, \mu \ (SS) \\ 1-2 \ e, \mu \\ 2 \ e, \mu \\ 2 \ e, \mu \\ 0 \\ 1 \ e, \mu \\ 0 \\ 0 \\ 3 \ e, \mu \ (Z) \end{matrix}$	2 b 0-3 b 1-2 b 0-2 jets 2 jets 2 b 1 b 2 b nono-jet/c-ta 1 b 1 b	Yes Yes Yes Yes Yes Yes Yes Yes Yes Yes	20.1 20.7 4.7 20.3 20.3 20.1 20.7 20.5 20.3 20.3 20.3 20.3	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	1308.2631 ATLAS-CONF-2013-007 1208.4305, 1209.2102 1403.4853 1403.4853 1308.2631 ATLAS-CONF-2013-037 ATLAS-CONF-2013-024 ATLAS-CONF-2013-068 1403.5222 1403.5222
Ž	EW direct	$ \begin{array}{c} \tilde{\ell}_{L_{\mathbf{R}}} \tilde{\ell}_{L,\mathbf{R}}, \tilde{\ell} \rightarrow \ell \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\ell} \nu (\ell \tilde{\nu}) \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\tau} \nu (\tau \tilde{\nu}) \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{2}^{0} \rightarrow \tilde{\ell}_{L} \nu \tilde{\ell}_{L} \ell (\tilde{\nu}\nu), \ell \tilde{\nu} \tilde{\ell}_{L} \ell (\tilde{\nu}\nu) \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} Z \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} h \tilde{\chi}_{1}^{0} \end{array} $	2 e,μ 2 e,μ 2 τ 3 e,μ 2-3 e,μ 1 e,μ	0 0 - 0 2 b	Yes Yes Yes Yes Yes Yes	20.3 20.3 20.7 20.3 20.3 20.3	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	1403.5294 1403.5294 ATLAS-CONF-2013-028 1402.7029 1403.5294, 1402.7029 ATLAS-CONF-2013-093
+ RPV	Long-lived particles	Direct $\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}$ prod., long-lived $\tilde{\chi}_{1}^{\pm}$ Stable, stopped \tilde{g} R-hadron GMSB, stable $\tilde{\tau}, \tilde{\chi}_{1}^{0} \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu})_{+}\tau(e,$ GMSB, $\tilde{\chi}_{1}^{0} \rightarrow \gamma \tilde{G}$, long-lived $\tilde{\chi}_{1}^{0}$ $\tilde{q}\tilde{q}, \tilde{\chi}_{1}^{0} \rightarrow qq\mu$ (RPV)	Disapp. trk 0 μ) 1-2 μ 2 γ 1 μ, displ. vtx	1 jet 1-5 jets - -	Yes Yes - Yes -	20.3 22.9 15.9 4.7 20.3	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	ATLAS-CONF-2013-069 ATLAS-CONF-2013-057 ATLAS-CONF-2013-058 1304.6310 ATLAS-CONF-2013-092
ed LLP	ΝPV	$ \begin{array}{l} LFV pp \rightarrow \widetilde{v}_{\tau} + X, \widetilde{v}_{\tau} \rightarrow e + \mu \\ LFV pp \rightarrow \widetilde{v}_{\tau} + X, \widetilde{v}_{\tau} \rightarrow e(\mu) + \tau \\ Bilinear \ RPV \ CMSSM \\ \widetilde{\chi}_1^+ \widetilde{\chi}_1^-, \widetilde{\chi}_1^+ \rightarrow W \widetilde{\chi}_1^0, \widetilde{\chi}_1^0 \rightarrow ee \widetilde{\nu}_{\mu}, e\mu \widetilde{\nu}_e \\ \widetilde{\chi}_1^+ \widetilde{\chi}_1^-, \widetilde{\chi}_1^+ \rightarrow W \widetilde{\chi}_1^0, \widetilde{\chi}_1^0 \rightarrow \tau \tau \widetilde{\nu}_e, e\tau \widetilde{\nu}_{\tau} \\ \widetilde{g}^+ \overline{q} \overline{q} \\ \widetilde{g} \rightarrow \widetilde{q} \overline{q} \\ \widetilde{g} \rightarrow \widetilde{\eta}_1 t, \widetilde{\tau}_1 \rightarrow bs \end{array} $	$\begin{array}{c} 2 \ e, \mu \\ 1 \ e, \mu + \tau \\ 1 \ e, \mu \\ 4 \ e, \mu \\ 3 \ e, \mu + \tau \\ 0 \\ 2 \ e, \mu \ (SS) \end{array}$	- 7 jets - - 6-7 jets 0-3 b	- Yes Yes Yes - Yes	4.6 4.7 20.7 20.7 20.3 20.7	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	1212.1272 1212.1272 ATLAS-CONF-2012-140 ATLAS-CONF-2013-036 ATLAS-CONF-2013-036 ATLAS-CONF-2013-091 ATLAS-CONF-2013-007
MSSM	Other	Scalar gluon pair, sgluon $\rightarrow q\bar{q}$ Scalar gluon pair, sgluon $\rightarrow t\bar{t}$ WIMP interaction (D5, Dirac χ)	$ \begin{array}{c} 0\\ 2 e, \mu (SS)\\ 0\\ \end{array} $	4 jets 2 b mono-jet	- Yes Yes	4.6 14.3 10.5	sgluon 100-287 GeV incl. limit from 1110.2693 sgluon 350-800 GeV m(χ)<80 GeV, limit of<687 GeV for D8	1210.4826 ATLAS-CONF-2013-051 ATLAS-CONF-2012-147
		$v_s = r + v_s$	artial data	$v_s = c$ full d	data		10 ⁻¹ Mass scale [TeV]	

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 σ theoretical signal cross section uncertainty.

David Berge (GRAPPA) / TeVPA 2013

Natural SUSY searches Search for 3rd generation squarks, leptons, b-jets, jets, missing Et

Probe SUSY as solution to the hierarchy problem



2: Dark Matter Pair Searches at Colliders

"mono-X" = $\gamma/jet/W/Z$

Assume contact interaction with mass suppression scale M*.

Dark matter mass m_x, plus higher-dimensional operators (e.g. D1-D11), fix cross sections.

Name	Initial state	Type	Operator
D1	qq	scalar	$rac{m_q}{M_\star^3}ar\chi\chiar q q$
D5	qq	vector	$rac{1}{M_\star^2}ar{\chi}\gamma^\mu\chiar{q}\gamma_\mu q$
D8	qq	axial-vector	$rac{1}{M_\star^2}ar\chi\gamma^\mu\gamma^5\chiar q\gamma_\mu\gamma^5 q$
D9	qq	tensor	$rac{1}{M_\star^2} ar\chi \sigma^{\mu u} \chi ar q \sigma_{\mu u} q$
D11	gg	scalar	$rac{1}{4M_\star^3}ar\chi\chilpha_s(G^a_{\mu u})^2$











- Data

Тор

Z(vv)+jet $W/Z(e/\mu/\tau)$ +jet

Diboson ///// uncertainty

> D5(u=d) x100 D5(u=-d) x1

D5(u=d) x20

 $D5(u=-d) \times 0.2$

100

110 m_{iet} [GeV]

<u>Mono-Z→II</u>



- 8 TeV, 20 fb⁻¹, PRD (submitted), arXiv:1404.0051
- Two leptons, Z invariant mass cut, balanced by missing ET
- Veto on >2 leptons and jets
- Main background SM ZZ/WZ (from NLO MC)



Galaxy cluster Abell 2744

 M_* and m_{χ} fix dark-matter nucleon scattering and annihilation cross sections (for given interaction type)!





SM χ SM

arXiv:1210.4491

- M_{*} fixes coupling SM-χ
- Coupling fixes annihilation cross section χ to SM
- Relic density (WMAP) corresponds to cross section
- Probe thermal relic dark matter particle at LHC!
- Red above green line: conflict!
 - Exclusion for this operator, or
 - Additional annihilation channels (leptons!), or
 - Wrong assumptions

Mono-W/Z→qq vs Monojets



Constructive interference with potentially large effect!

Mono-Z→II – EFT validity?



Effective Field Theory / assumptions of contact interactions not always valid (cf. C.McCabe's talk) \rightarrow start to explore simplified models!

Limits on χ - η coupling constant f shown. Above black line, lower limit on f below f from relic abundance calculations.





Direct search for invisible Higgs decays PRL (submitted), arXiv:1402.3244 Constrain invisible Higgs width to <75% of the full SM width



 \overline{q}

Z

'Light' mediator, simplest connection of SM and dark sector.

Events / 30 GeV 10²

E

10

2⊧

.5

0.5₽

Data / Expected

 χ

A. Djouadi *et al.*, Phys. Lett. B 709, 65 (2012), arXiv:1112.3299



This, together with the relic density, basically rules out the simplest Higgs portals for DM up to half the Higgs mass, except right at the Higgs pole!

Conclusions

• The one model independent result we have about dark matter:

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We have not discovered a dark matter candidate particle yet at the LHC!

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- Intense search program optimised for dark matter (mono-X) or SUSY (R parity!)
- LHC Run-2 will hopefully tell us more at 13 or 14 TeV
- ATLAS public results:
 - Journal publications: <u>https://twiki.cern.ch/twiki/bin/view/AtlasPublic/Publications</u>
 - Conference notes with preliminary results, internally reviewed: <u>https://twiki.cern.ch/twiki/bin/view/AtlasPublic/CONFnotes</u>

The Large Hadron Collider



	2012 performance	Design performance
Colliding bunches	1331	2808
Energy	4 TeV x 4 TeV	7 TeV x 7 TeV
Bunch spacing	50 ns	25 ns
Luminosity	7.7 x 10 ³³ cm ⁻² s ⁻¹	10 ³⁴ cm ⁻² s ⁻¹
Pile-up interactions	~35	~25



Difficulty: event pile-up

ATLAS-CONF-2013-082



 $Z \rightarrow \mu\mu$ event in ATLAS with 25 reconstructed vertices



With pile-up suppression algorithms

Going after the Dark at the LHC

Galaxy cluster Abell 2744

Data combinations and interpretations require assumptions about the red bubble!

Particle Dark Matter Searches based on: SM SM X χ χ χ SM SM SM SN χ Indirect Direct Colliders

χ









$\chi \longrightarrow SM$ $\chi \longrightarrow SM$

arXiv:1210.4491

- M_* fixes coupling SM- χ
- Coupling fixes annihilation cross section χ to SM
- Relic density (WMAP) corresponds to cross section
- Probe thermal relic dark matter particle at LHC!
- Red above green line: conflict!
 - Exclusion for this operator, or
 - Negative interference needed, or
 - Additional annihilation (leptons!)

Reminder:

	Name	Initial state	Type	Operator
Ĩ	D1	qq	scalar	$rac{m_q}{M_\star^3}ar\chi\chiar q q$
	D5	qq	vector	$rac{1}{M_\star^2}ar\chi\gamma^\mu\chiar q\gamma_\mu q$
	D8	qq	axial-vector	$rac{1}{M_\star^2}ar\chi\gamma^\mu\gamma^5\chiar q\gamma_\mu\gamma^5 q$
	D9	qq	tensor	$rac{1}{M_{\star}^2}ar{\chi}\sigma^{\mu u}\chiar{q}\sigma_{\mu u}q$
	D11	gg	scalar	$rac{1}{4M_\star^3}ar\chi\chilpha_s(G^a_{\mu u})^2$

Where do we start?

Huge parameter space, but guiding principles



SUSY searches strategy driven by cross section and luminosity





Nada! Set limits...



<u>CMS EXO-12-048</u>

David Berge / GRAPPA & Nikhef

CMS mono-jet, 20 fb⁻¹ at 8 TeV

Nada! Set limits...

arxiv:1209.4625



ATLAS mono- γ , 4.6 fb⁻¹ at 7 TeV

How do we search for SUSY ?



SUSY search analyses look for tails in distributions of observables sensitive to new heavy particles.

Can only exploit tails if detector response and SM backgrounds in signal region are understood.

How do we search for SUSY ? A brief primer ...

 $CR \rightarrow SR$ transfer factors taken from MC simulation (many systematic effects cancel).



LHC Mission: Complete the Standard Model and Discover the Dark Sector (and more...)







How to search for SUSY at the LHC

SUSY (more than) duplicates spectrum of particle states wrt. Standard Model

Sparticles decay in jets, leptons, taus, photons, invisible (MET), ...

1. <u>*R*-parity conserving (RPC) signatures:</u>

- Sparticles produced in pairs, each decays to (WIMP) LSP, mostly lightest neutralino or gravitino
- One invisible LSP per decay chain \rightarrow MET
- 2. <u>R-parity violating (RPV) signatures:</u>
 - Resonances or multijets / multileptons: single sparticle production or LSP decay
 - Displaced vertices from late LSP decay
- 3. <u>Long-lived particles from:</u>
 - Weak couplings (eg, RPV, gravitino)
 - High virtuality from heavy mediator sparticles (eq, heavy squarks in split SUSY)
 - Mass degeneracy (eg, m (chargino) ~ m (LSP) in AMSB)





Can still search for pairs of "ISR-tagged" dark-matter particles, not particular SUSY signatures



Amount of missing E_T depends on mass difference!







Data combinations require assumptions about the red bubble!



Name	Initial state	Type	Operator
D1	qq	scalar	$rac{m_q}{M_\star^3}ar\chi\chiar q q$
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D9	qq	tensor	$rac{1}{M_\star^2}ar\chi\sigma^{\mu u}\chiar q\sigma_{\mu u}q$
D11	gg	scalar	$rac{1}{4M_\star^3}ar\chi\chilpha_s(G^a_{\mu u})^2$



45

Data combinations require assumptions about the red bubble!



