

Max-Planck-Institut für Physik
(Werner-Heisenberg-Institut)

FROGS
FRont Of pro-Galician Scientists

Disordered Holographic Superconductors

In collaboration with

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A. Scardicchio (ICTP, Italy)

[1308.1920, 1407.7526, ...]

Daniel Areán
Reykjavík, August 2014

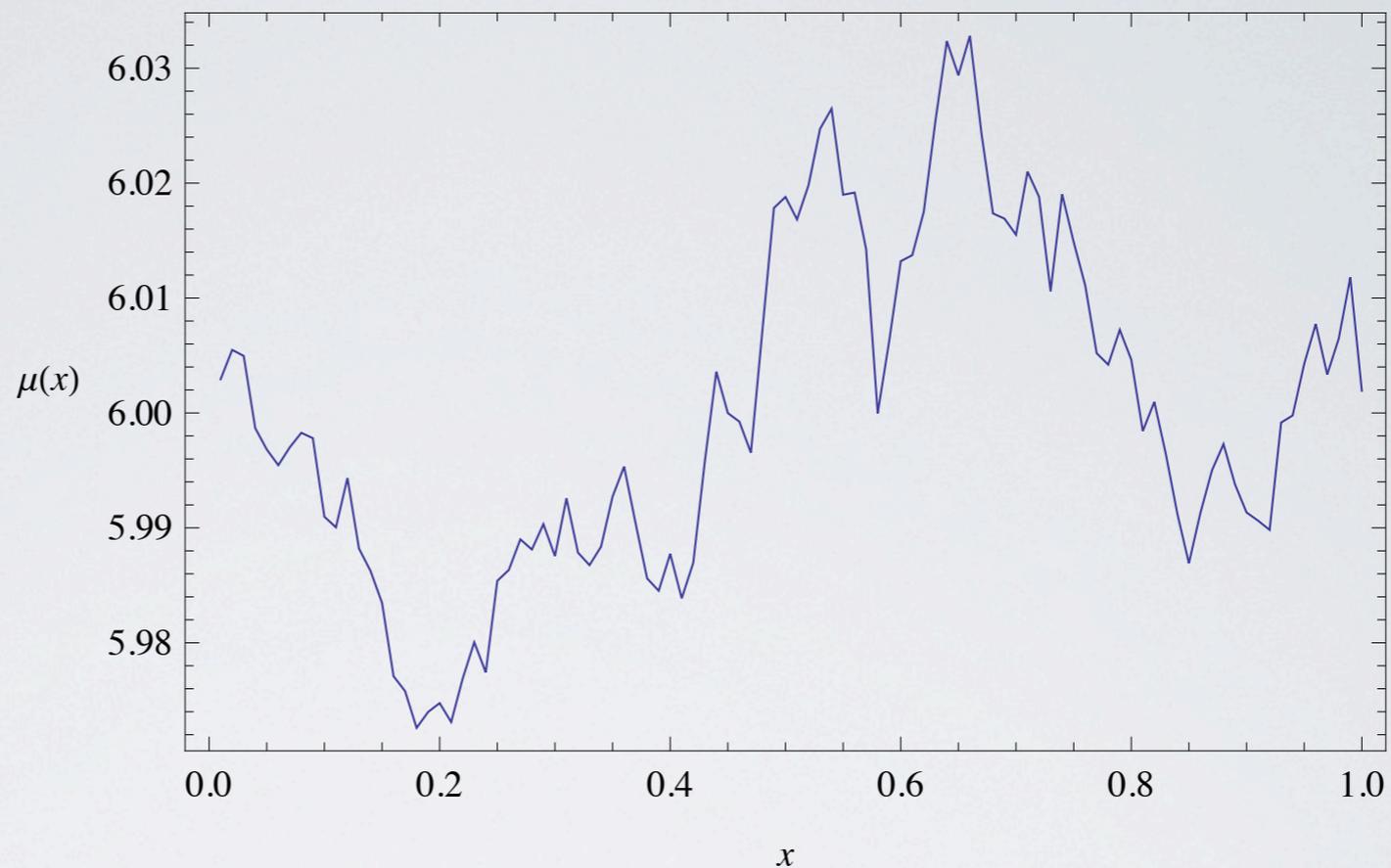


Noise

[charged impurities]

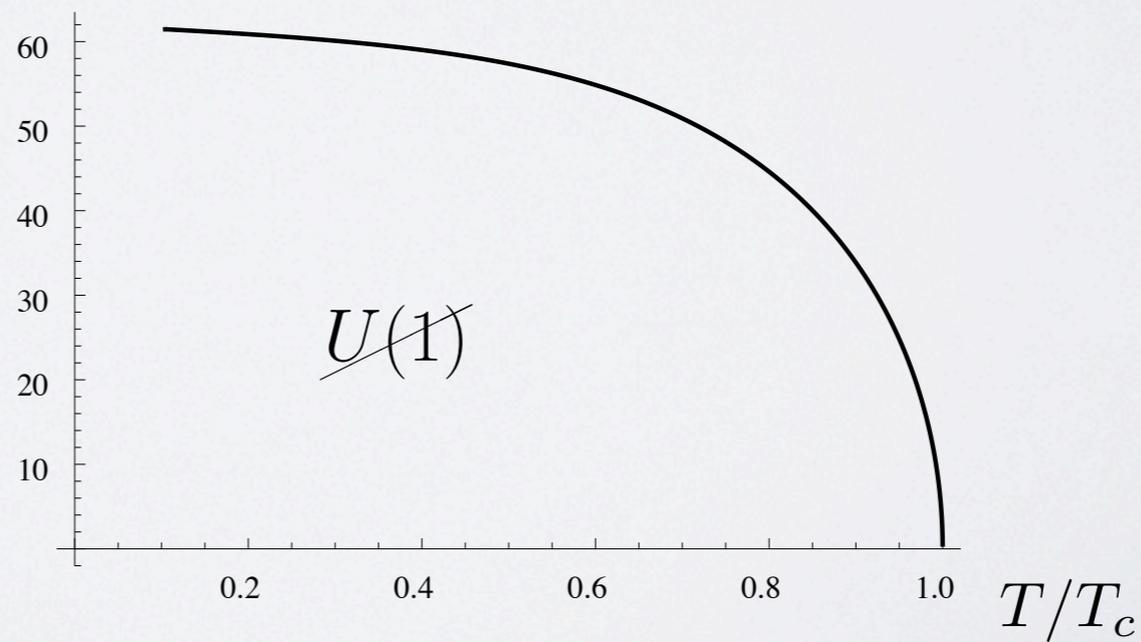
in holography

$$\mu(x)$$

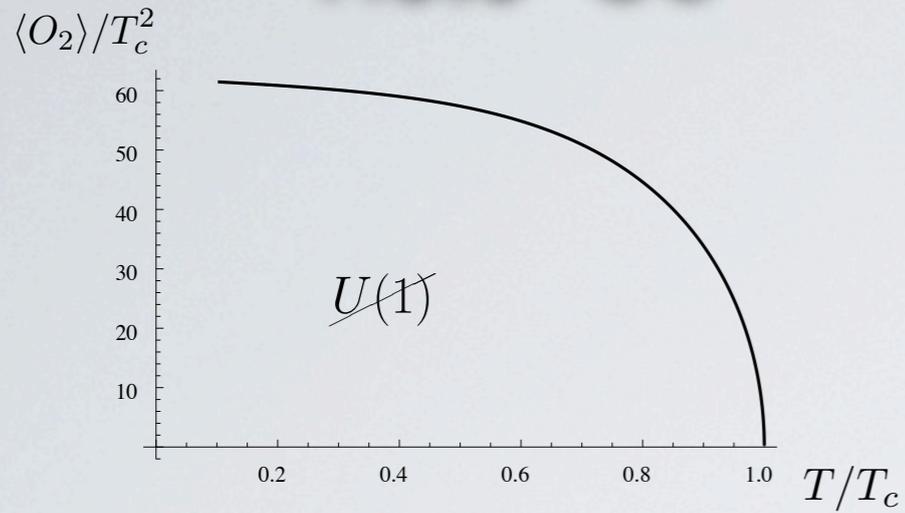


Holo-SC

$$\langle O_2 \rangle / T_c^2$$



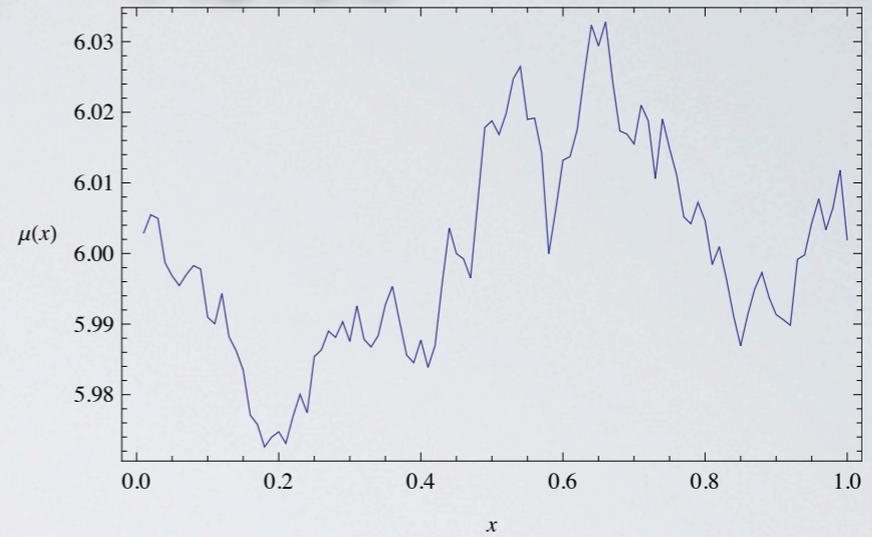
Holo-SC



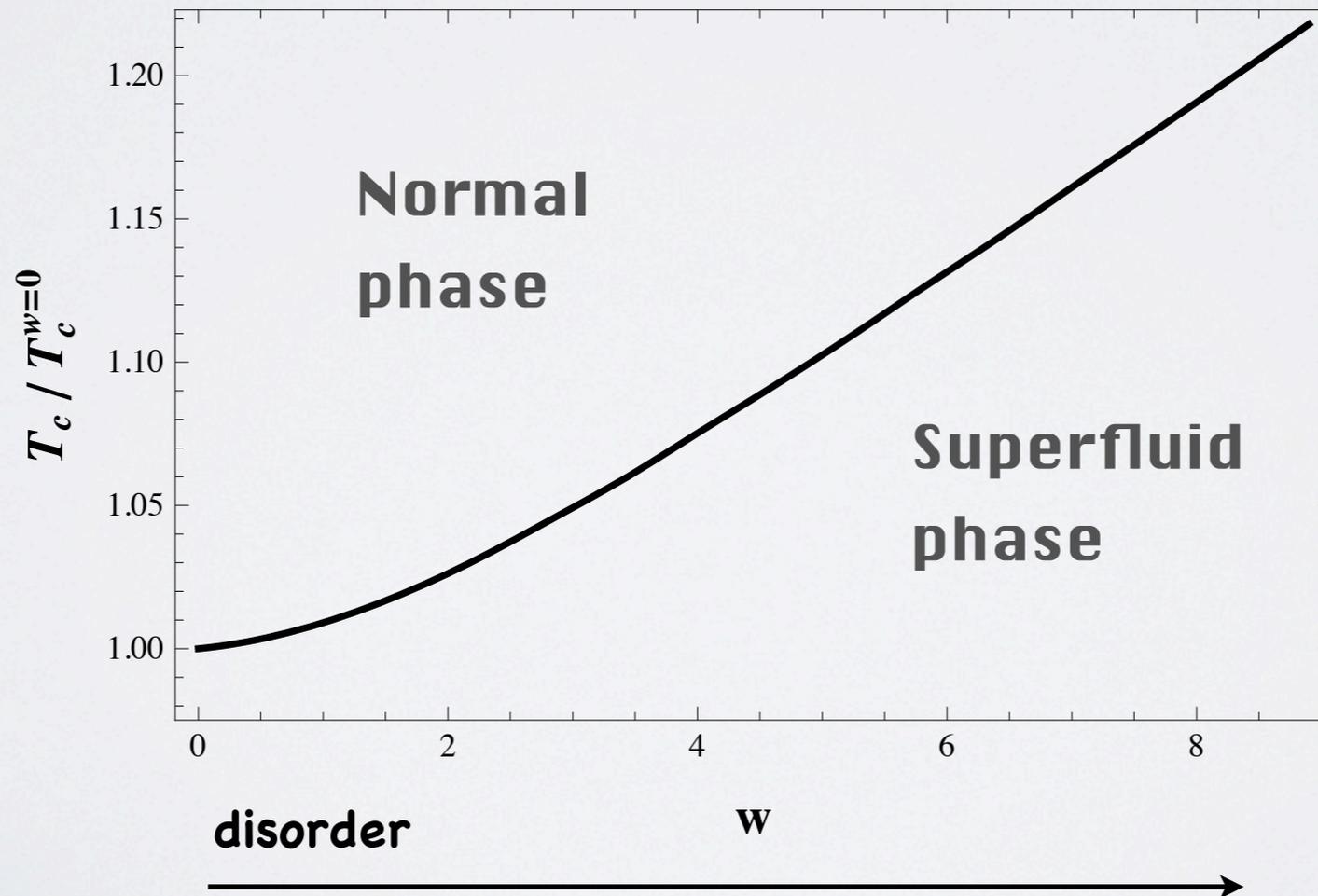
+

Noise

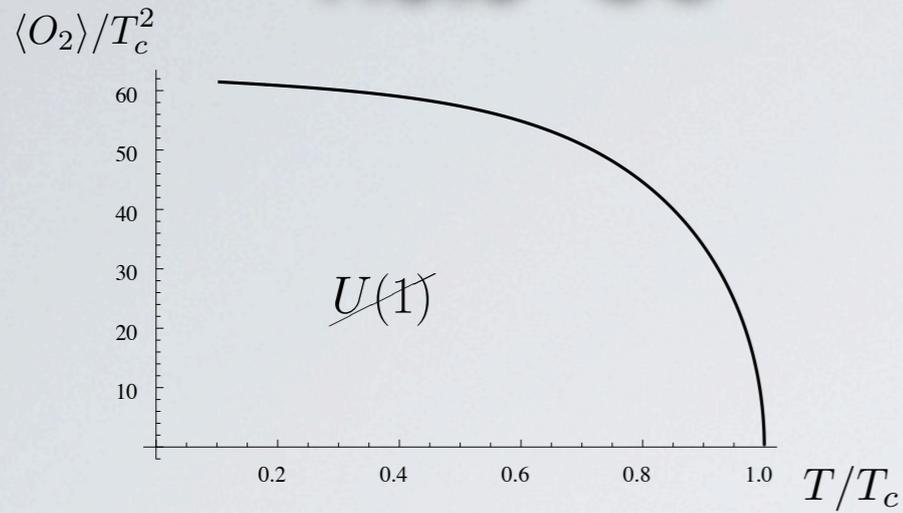
$\mu(x)$



= ★ Enhancement of SC

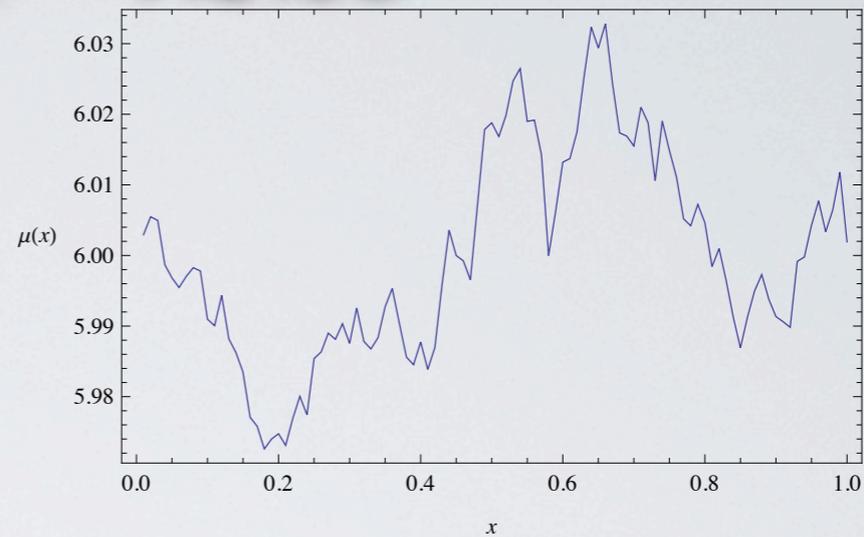


Holo-SC



+

Noise $\mu(x)$

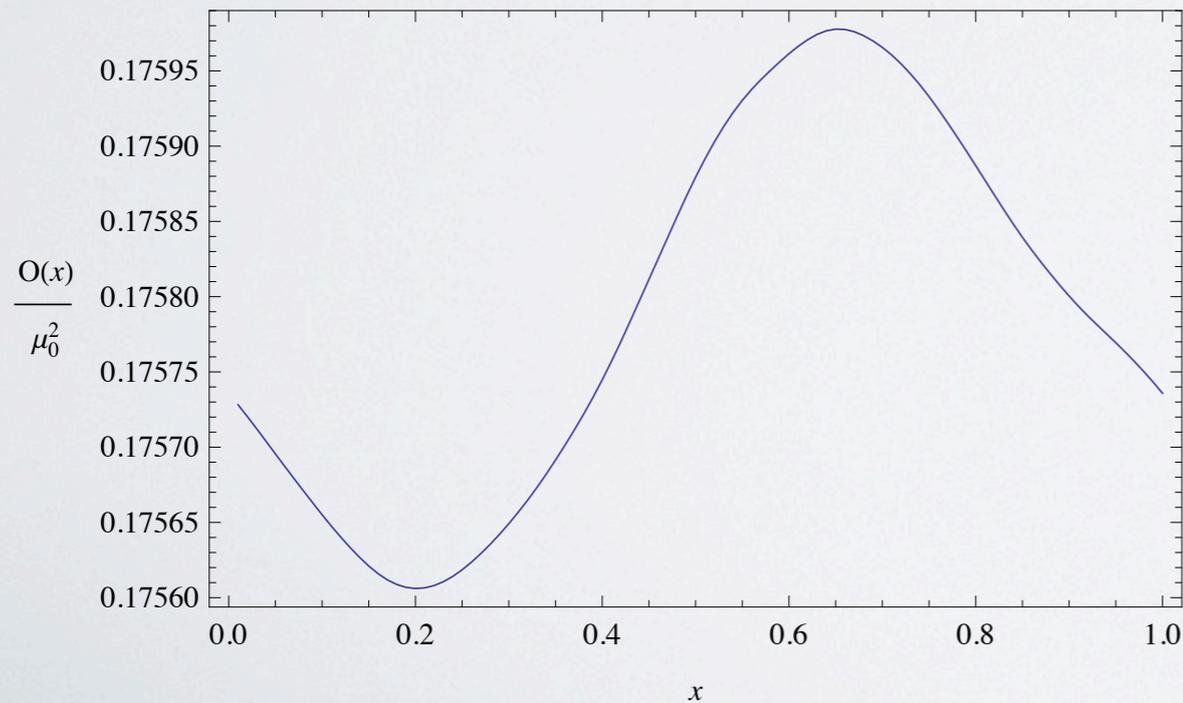


$$S_k = \frac{1}{k^{2\alpha}}$$

= ★ Spectrum 'renormalization'

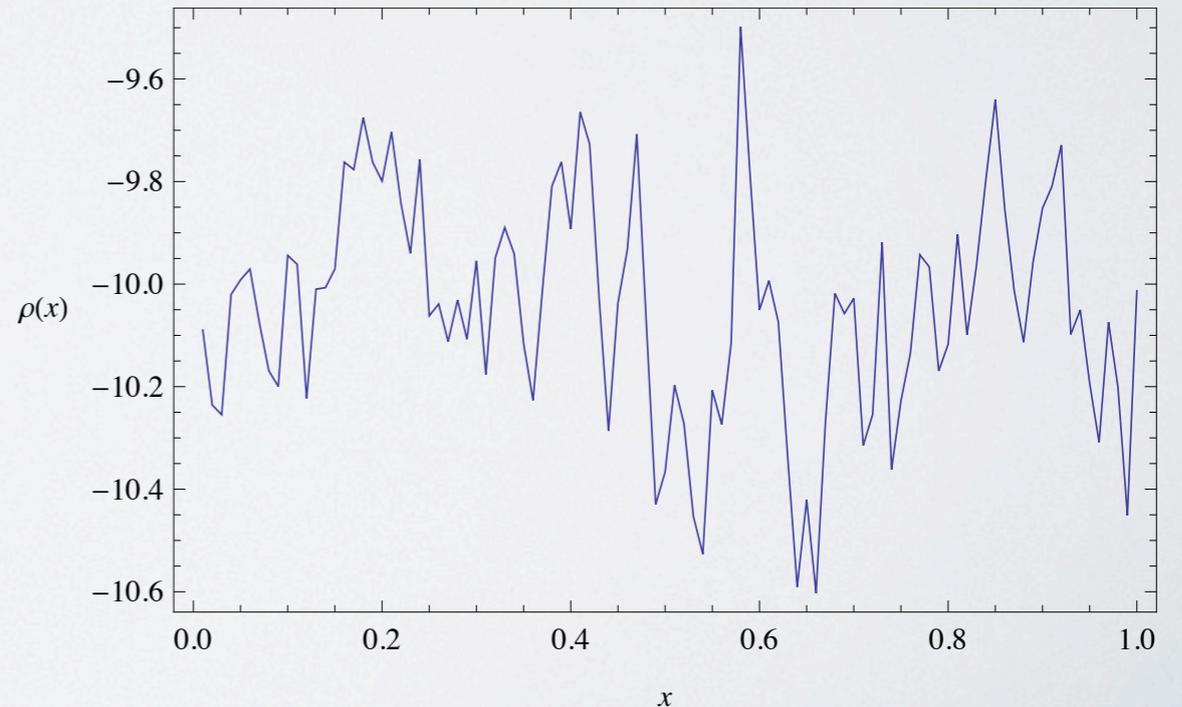
Condensate

$$S_k = \frac{1}{k^{2\alpha+4}}$$



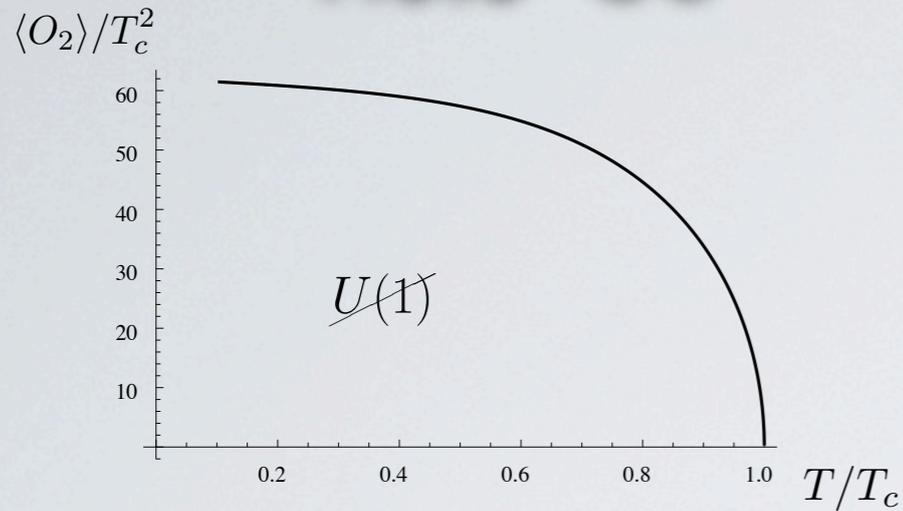
Charge density

$$S_k = \frac{1}{k^{2\alpha-2}}$$



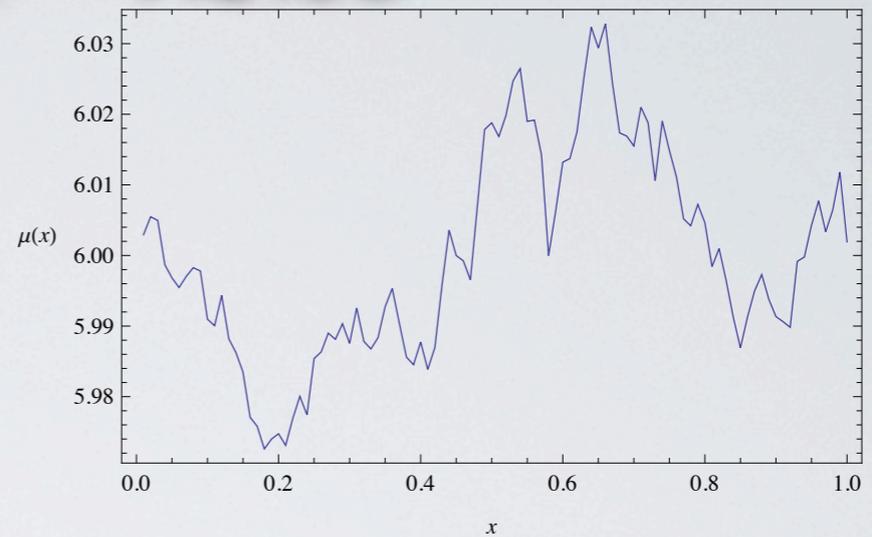
[also in brane intersections]

Holo-SC



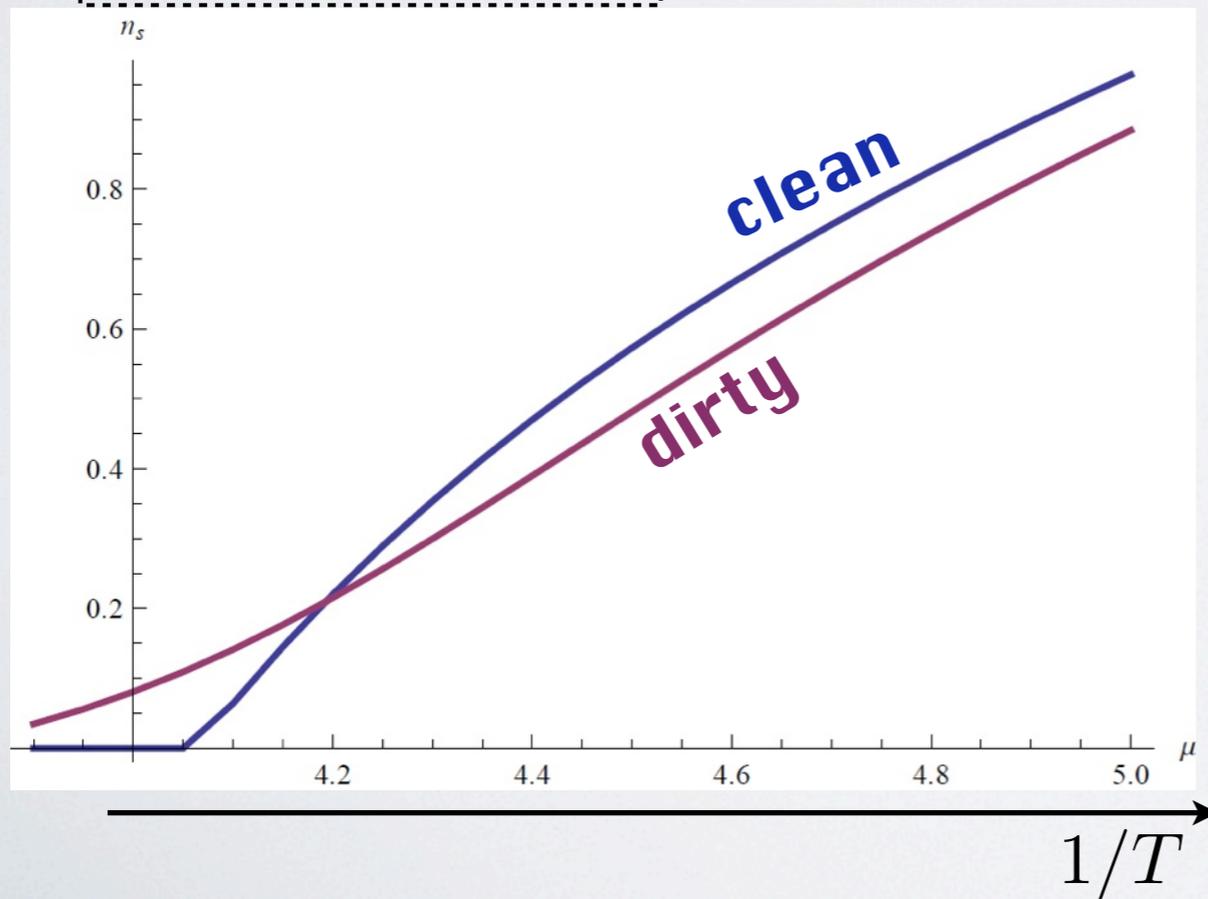
+

Noise $\mu(x)$



= ★ **Conductivities of disordered systems** [for branes too]

Superfluid density



‘noise lowers the conductivity’

OUTLINE

- > **Motivation: Strong coupling, Disorder, Superconductors**
- > **Review: Holographic Superconductors**
- > **Dirty Holographic (p-wave) Superconductors**
- > **Results: Phase diagram, spectrum, (some) noisy σ**
- > **Future: Dirty Thin Films (islands of SC?), noisy σ , ...**

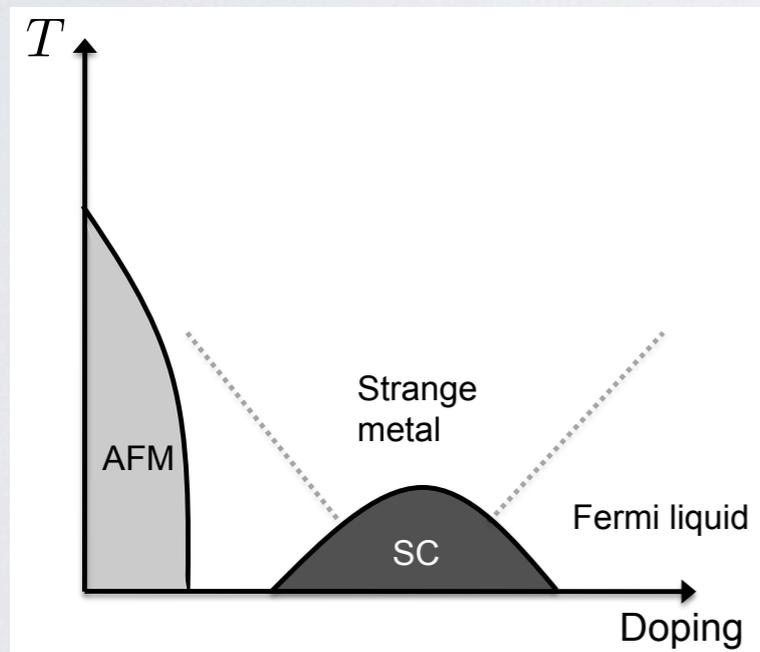
> Challenges in Condensed Matter:

- > **Strong Coupling:** High T_c Superconductors (strange metals), heavy fermions, ...
- > **Disorder + Interactions:** Anderson localization in many body int. systems

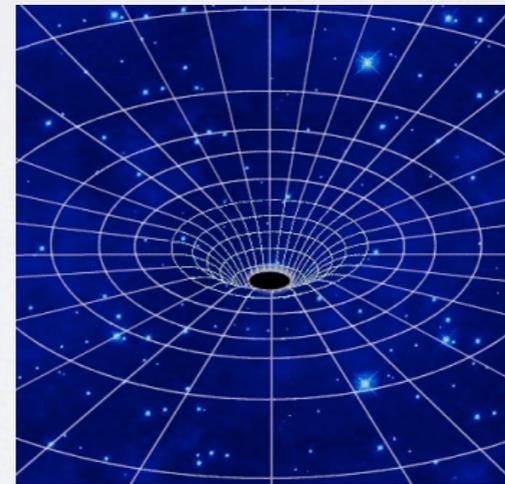
★ **AdS/CFT** weak / strong coupling duality

> High T_c Superconductors

[‘gravities’ + matter in \sim AdS]



\sim

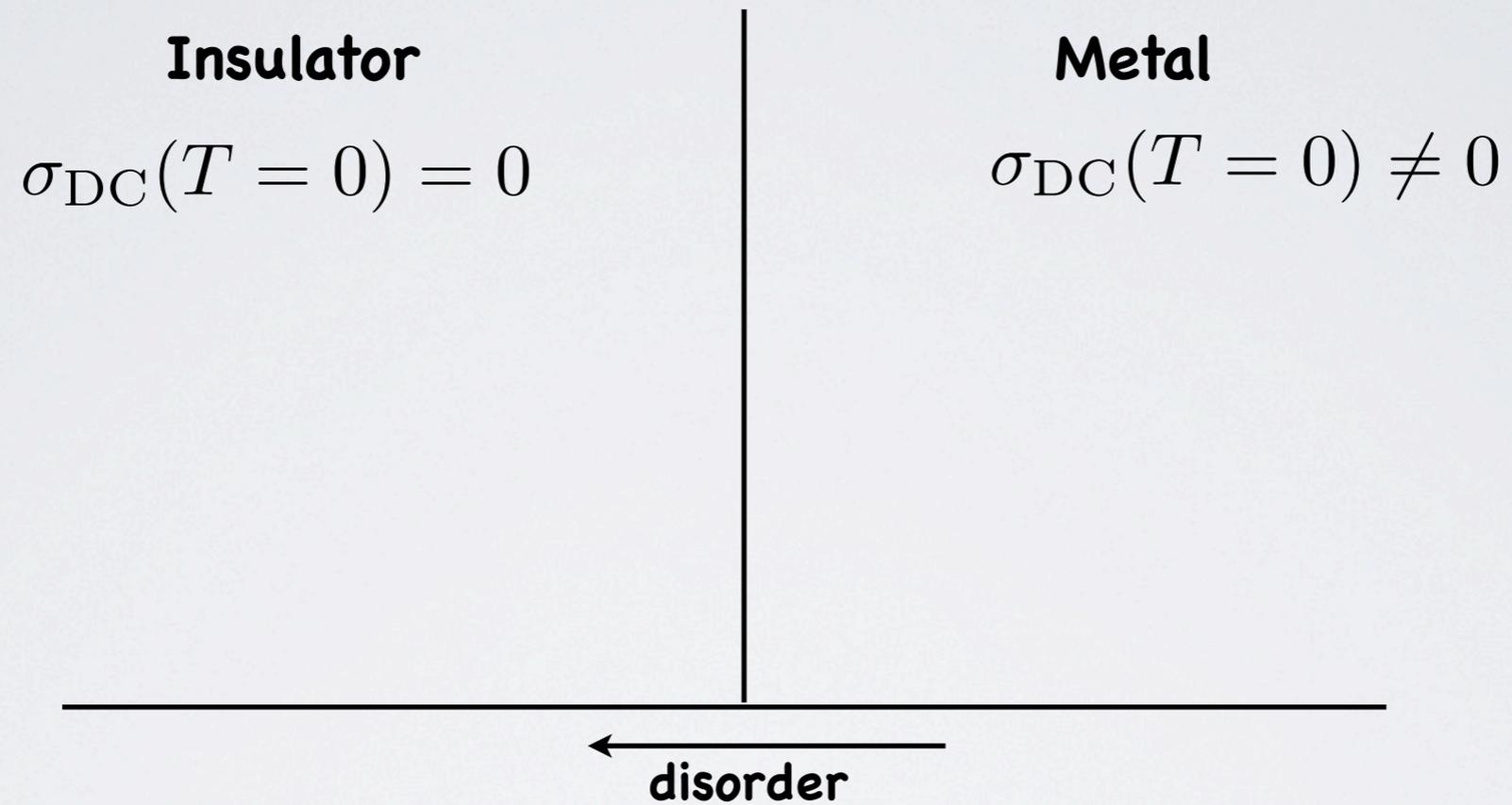


> **Black holes with hair, domain wall geometries, electron stars ...**

- Superconducting phase \rightarrow ~~BCS~~
- Strange metal \rightarrow Non-Fermi liquid

> Disorder and interactions

> Anderson Localization '58: disorder suppresses conductivity

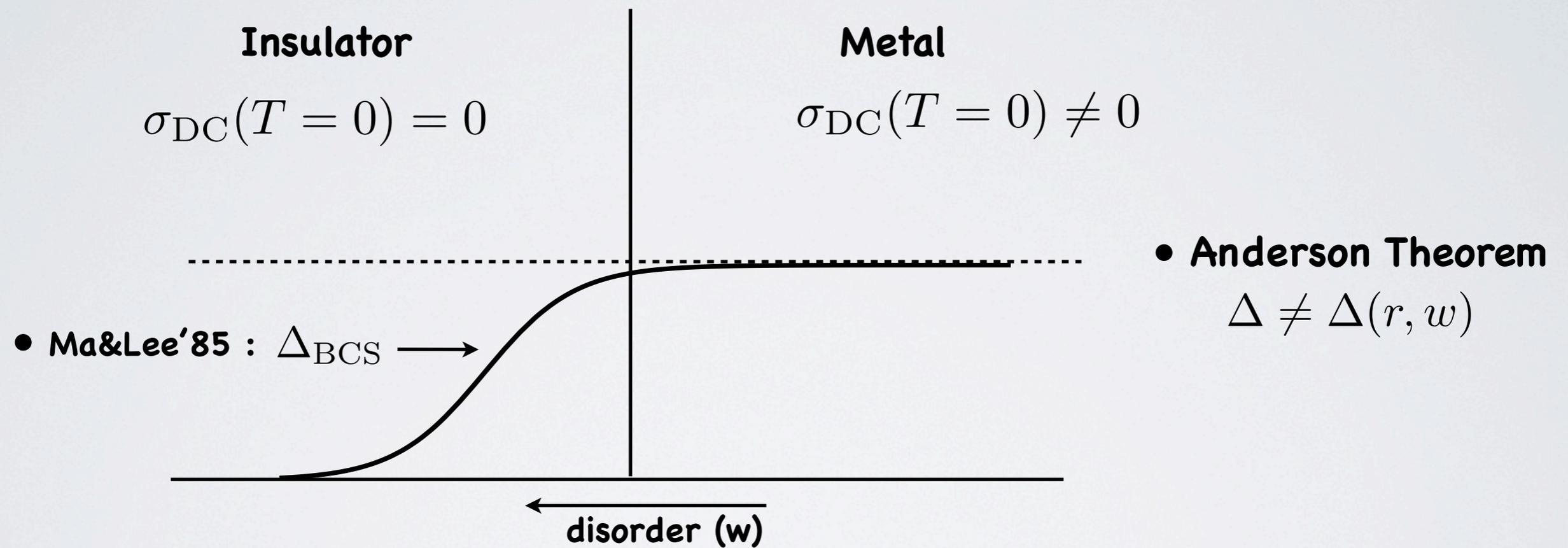


Disorder and superconductors?

> Disorder and interactions

> Anderson Localization '58: disorder suppresses conductivity

> Disorder and superconductors?

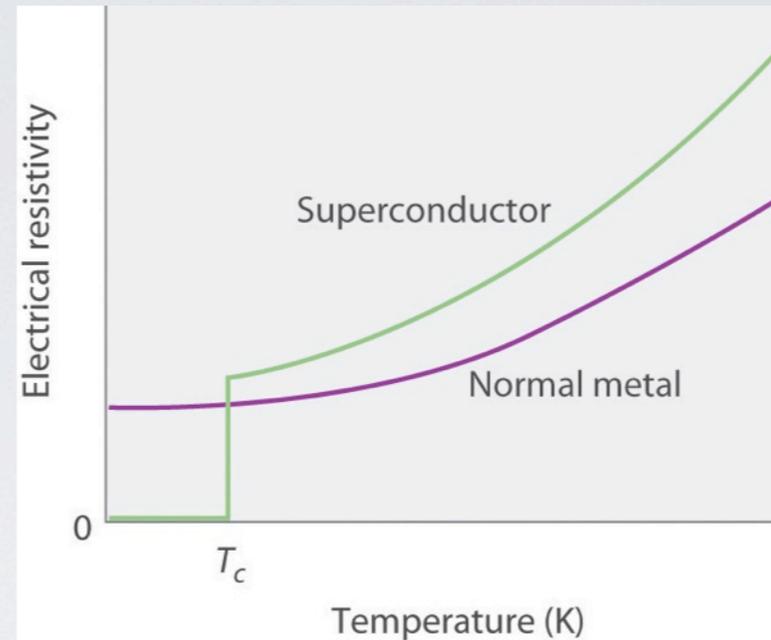


> Disorder + many body interacting system → difficult! (see cond-mat/0506617)

→ **AdS/CFT ?** → **Dirty Holographic Superconductors !**

★ Holographic Superconductors (Hartnoll, Herzog, Horowitz, '08)

- Black Hole gets hair \sim

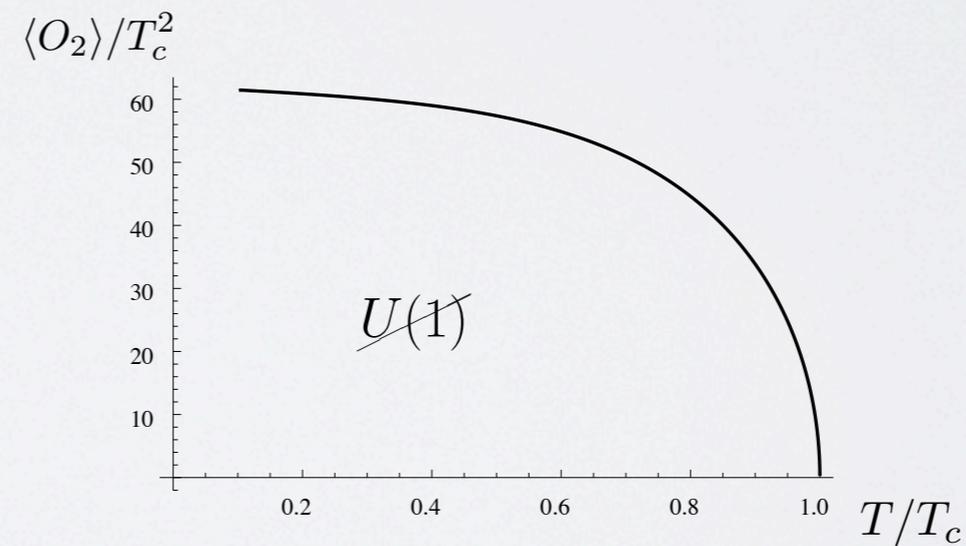


$$\langle U(1) \rangle$$

◆ Charged BH unstable against scalar condensation (Gubser'08)

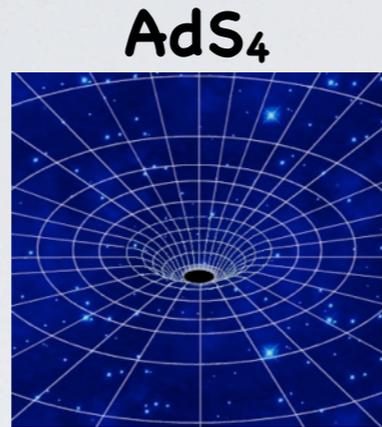


\sim



> Holographic p-wave Superconductor (Gubser'08)

$SU(2) F_{ab} F^{ab}$ in

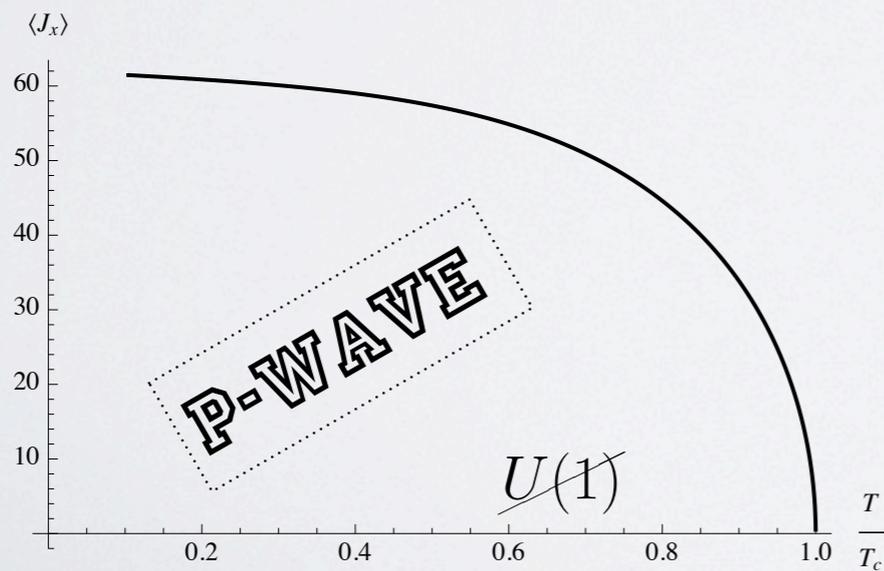


$\sim 2 + 1$ CFT ($T \neq 0$)

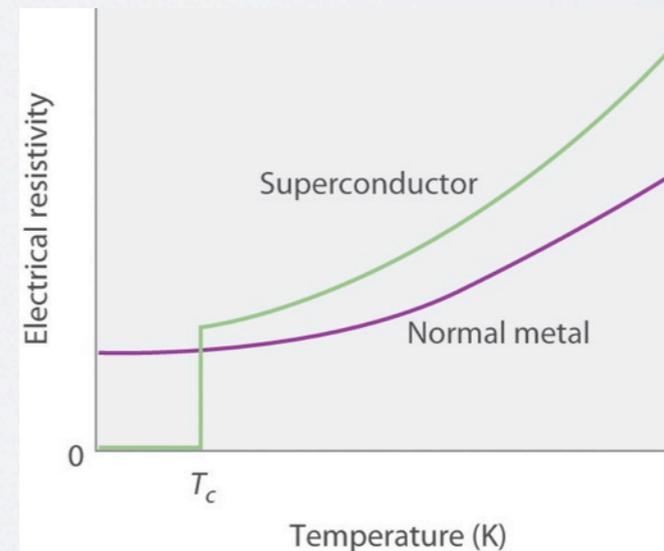
with

chemical potential // $U(1) \subset SU(2)$ $A_t^3(z) \sim \mu$ [$SU(2) \rightarrow U(1)$]

p-wave condensate $A_x^1(z) \sim \langle \mathcal{J}_x^1 \rangle$ [$\langle U(1) \rangle$] [~~rotational invariance~~]



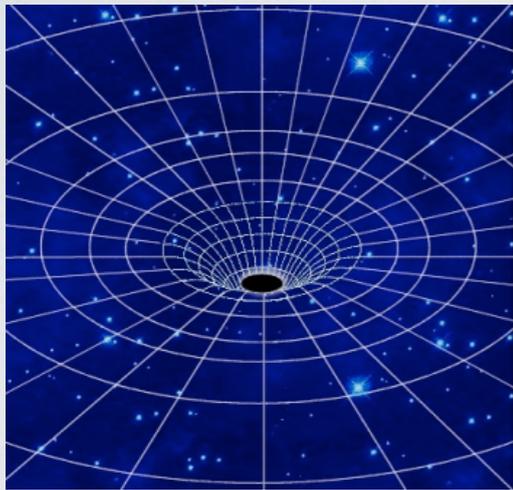
\sim



Adding Impurities!

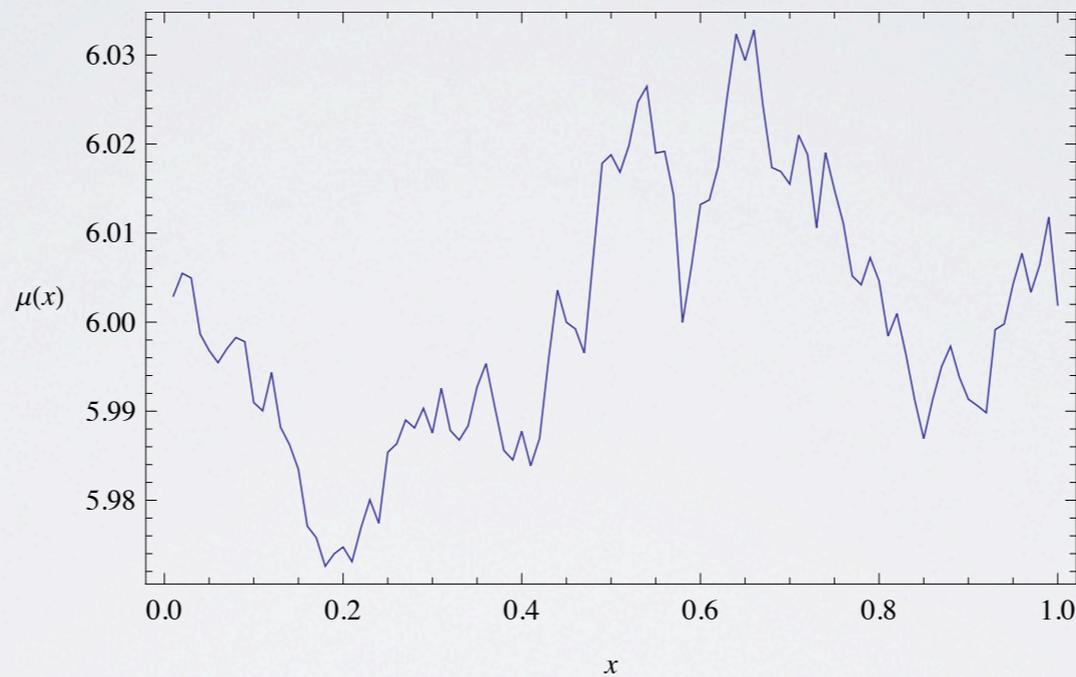
> 2+1 Holographic Superconductors + **Noisy chemical potential** $\mu = \mu(x)$

p-wave



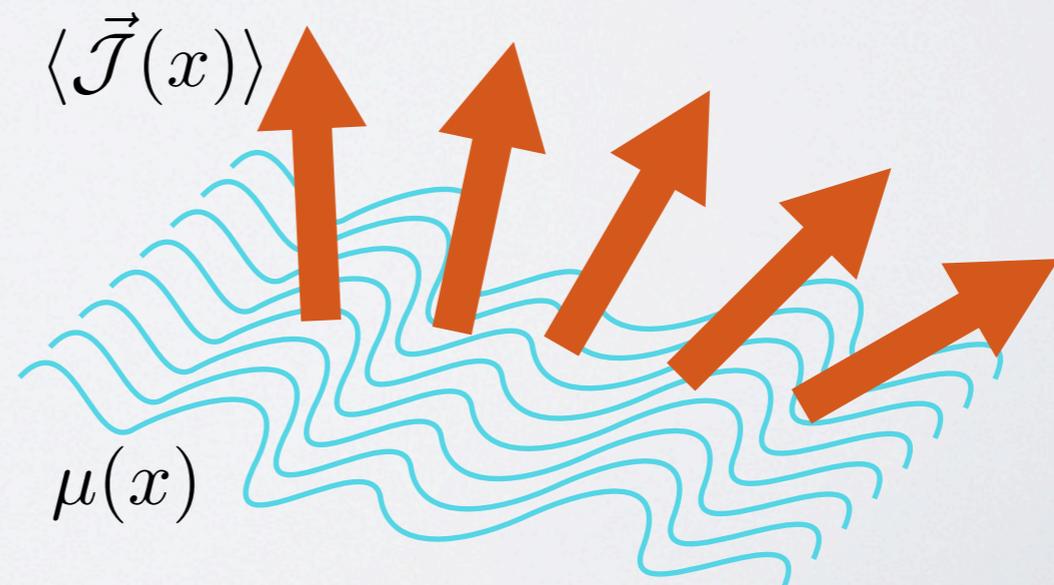
+

Disorder $\mu(x)$



? Condensate
? Phase diagram
? Spectrum

→ p-wave condensate
picks a direction?



[./Tech Specs/pwave_1]

Probe Limit

• **Action** $S = \int d^4x \sqrt{-g} \left(\boxed{-\frac{1}{4} F_{\mu\nu}^c F_c^{\mu\nu}} + \frac{R}{\mathcal{K}} + \frac{6}{\mathcal{K} L^2} \right)$

• **Field content** $A_t^3(x, z) \sim \mu(x)$ $(A_x^1(x, z), A_y^1(x, z)) \sim (\langle \mathcal{J}_x^1(x) \rangle, \langle \mathcal{J}_y^1(x) \rangle)$

$A_t^2(x, z)$

2nd 'charge density'

• **UV boundary conditions (z=0)**

$A_t^3(x, z) = \mu(x) + \dots$

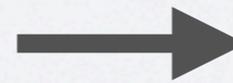
$A_i^1(x, z) = \cancel{w_i^{(0)}}(x) + \langle \mathcal{J}_i^{(1)}(x) \rangle z + \dots$

$A_t^2(x, z) = \cancel{\mu_2}(x) - \rho_2(x) z + \dots$

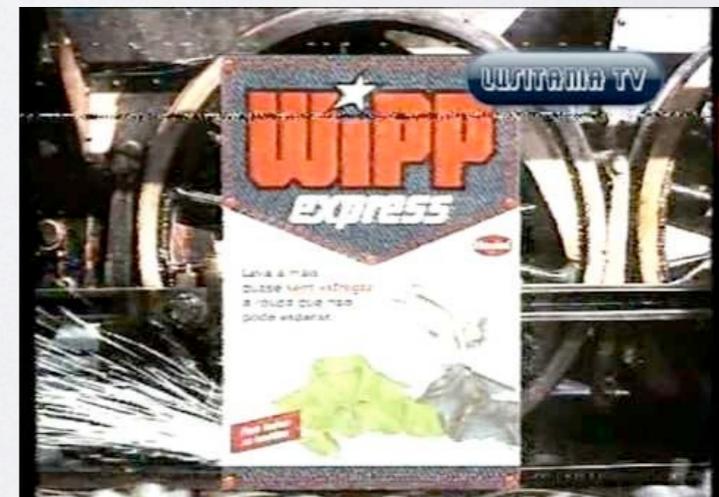


4 Coupled PDEs

s-wave: 2 PDEs



Numerics



Charged impurities >>> Noisy chemical potential

- NOISE THROUGH RANDOM PHASES

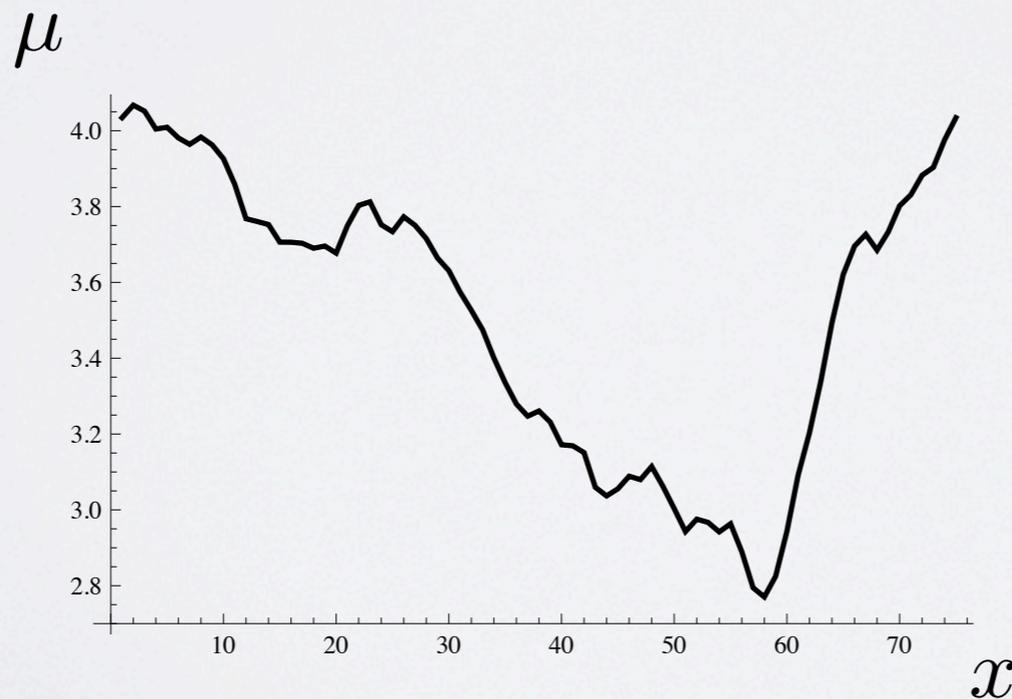
$$\mu(x) = \mu_0 + \epsilon \sum_{k=k_0}^{k_*} \sqrt{S_k} \cos(kx + \delta_k) = \mu_0 + \epsilon \sum_{k=k_0}^{k_*} \frac{1}{k^\alpha} \cos(kx + \delta_k)$$

Power spectrum
Random phases

Scales: $k_0, k_*, (\epsilon)$

Strength of noise $w = 25\epsilon/\mu_0$
 [see also Scardicchio cond-mat/0505050]

- $\mu_0 = 3.50, \alpha = 1.50, w = 3.50$ [$\mu_0 < \mu_c = 3.66$]



$L_x = 2\pi \rightarrow K_0 = 1$
 $N_x = 75$

Charged impurities >>> Noisy chemical potential

- NOISE THROUGH RANDOM PHASES

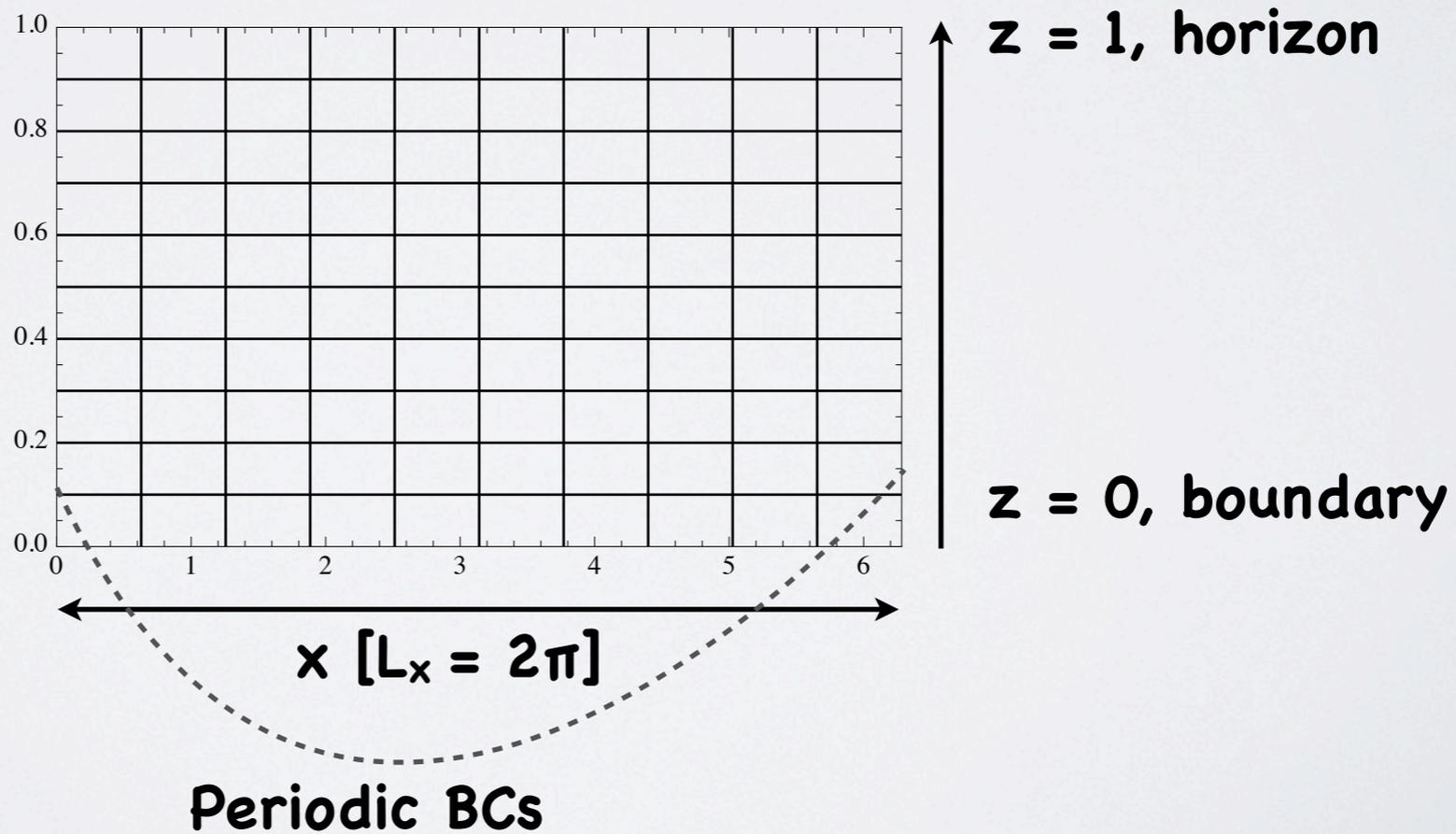
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Power spectrum
Random phases

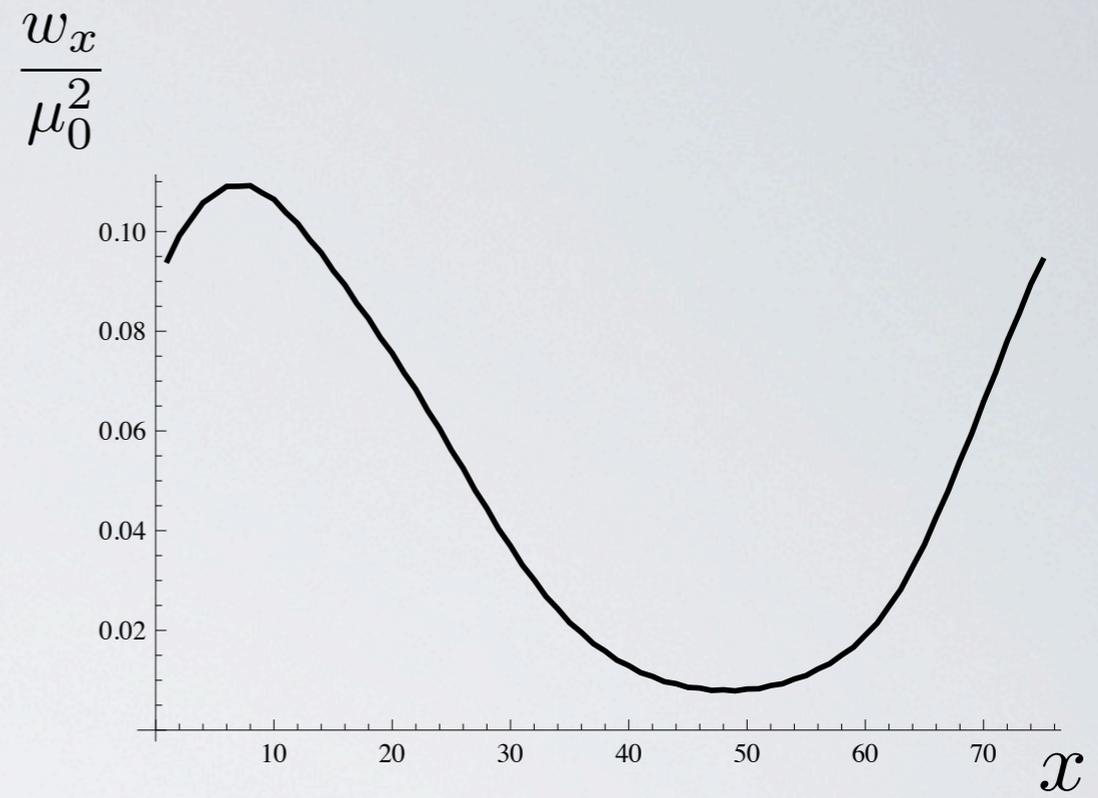
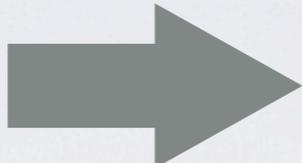
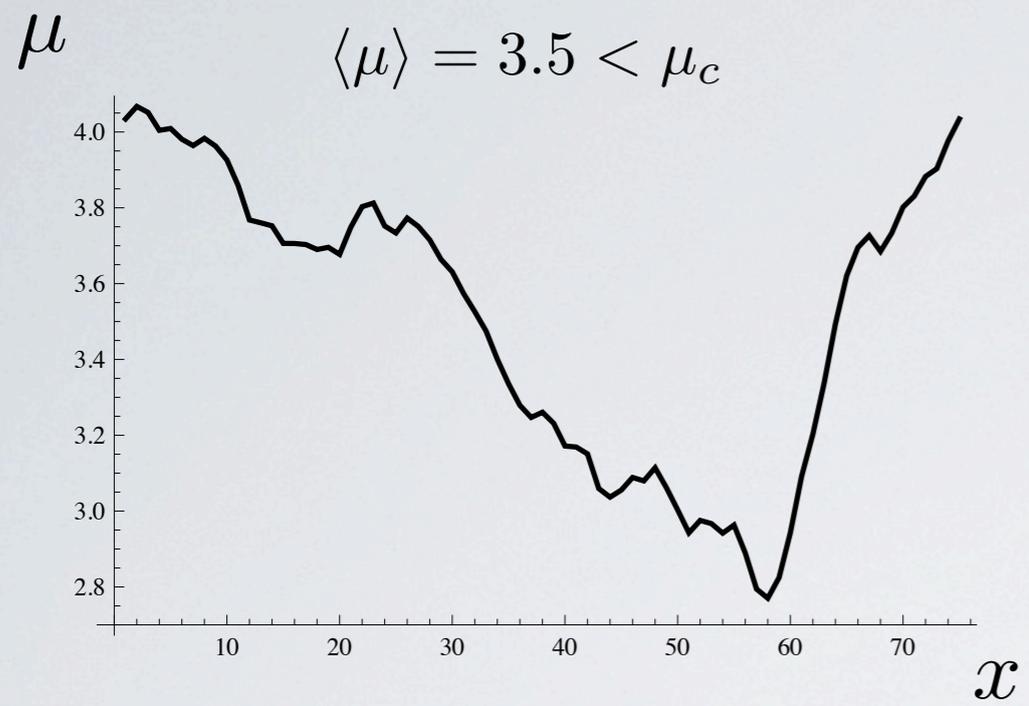
Scales: $k_0, k_*, (\epsilon)$

Strength of noise $w = 25\epsilon/\mu_0$
 [see also Scardicchio cond-mat/0505050]

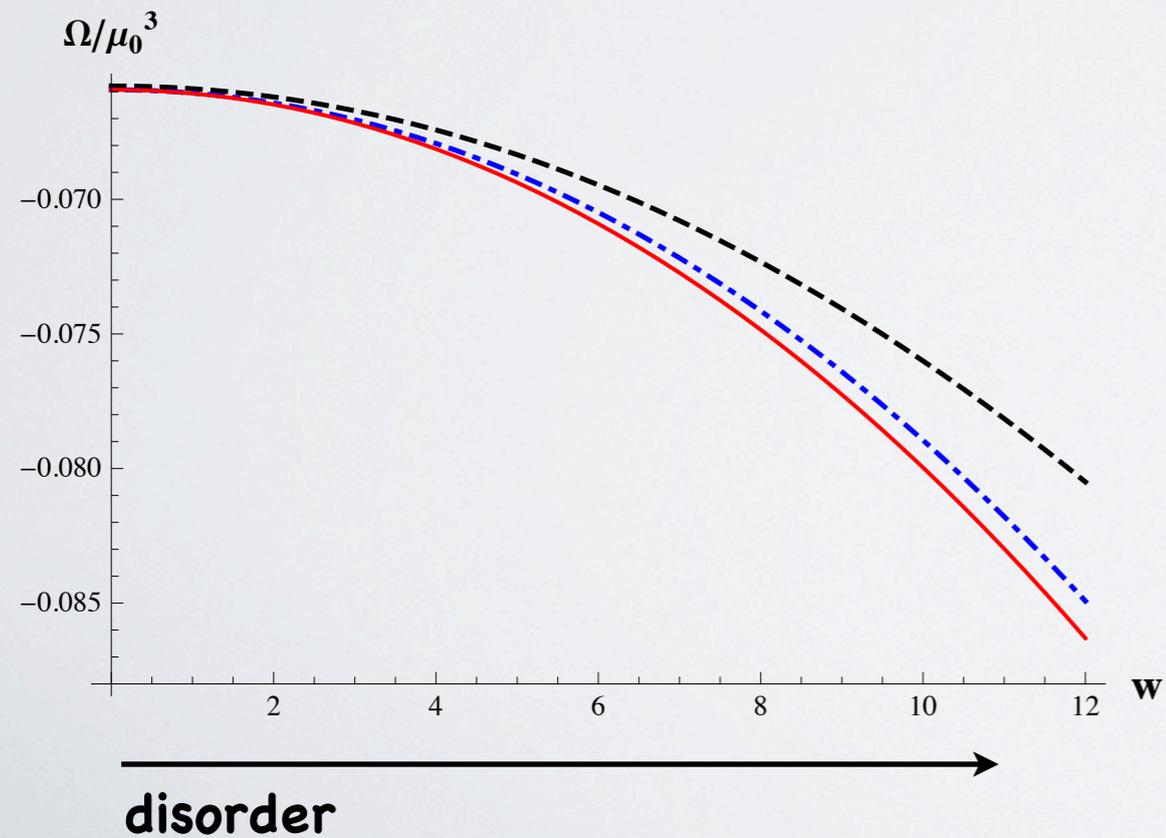
- SYSTEM ON A GRID



Condensate likes noise



> Free energy of competing solutions



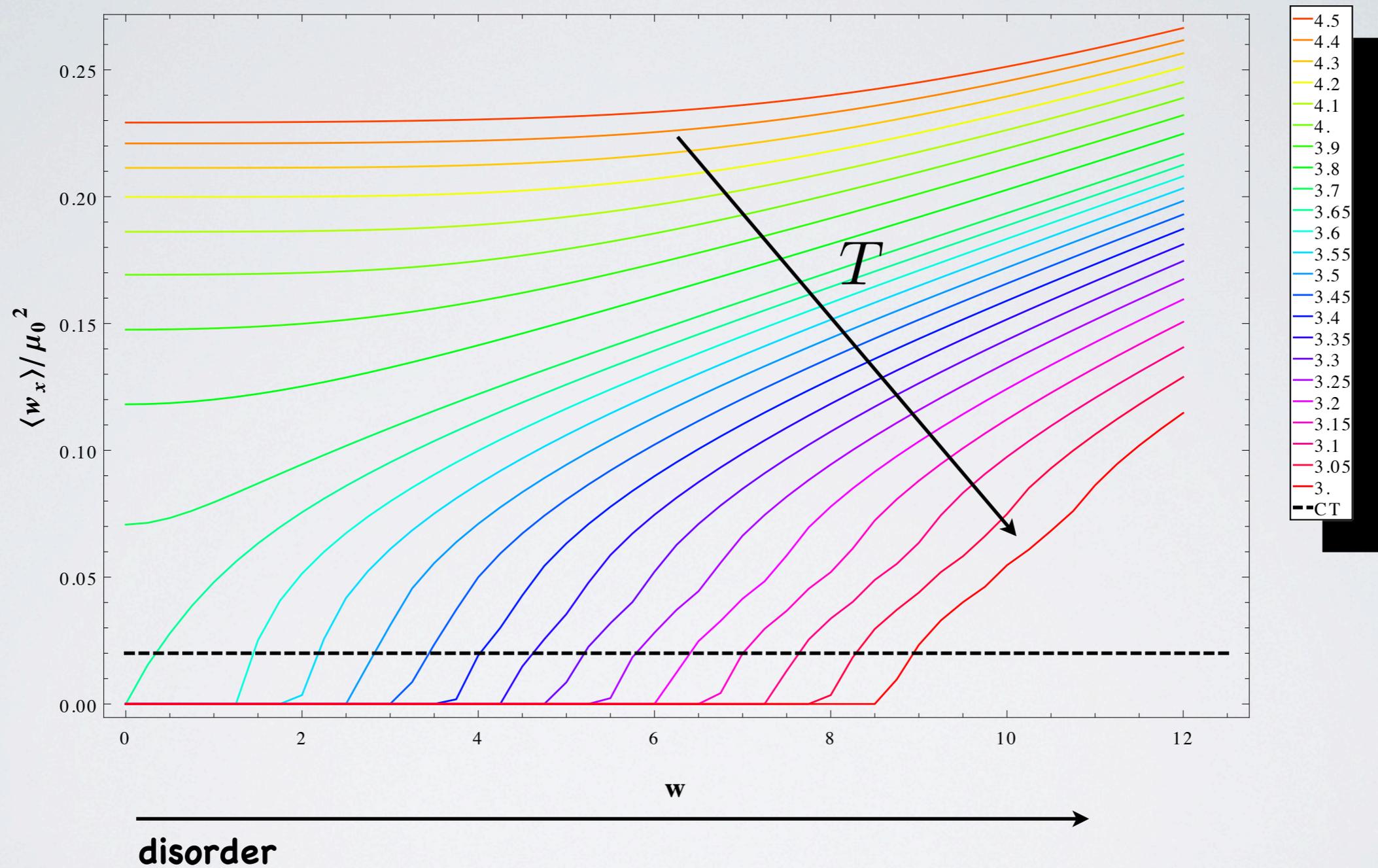
↳ *p-wave picks x:*

- Normal phase
- ... Condensate \perp Noise
- Condensate // Noise**

always wins!

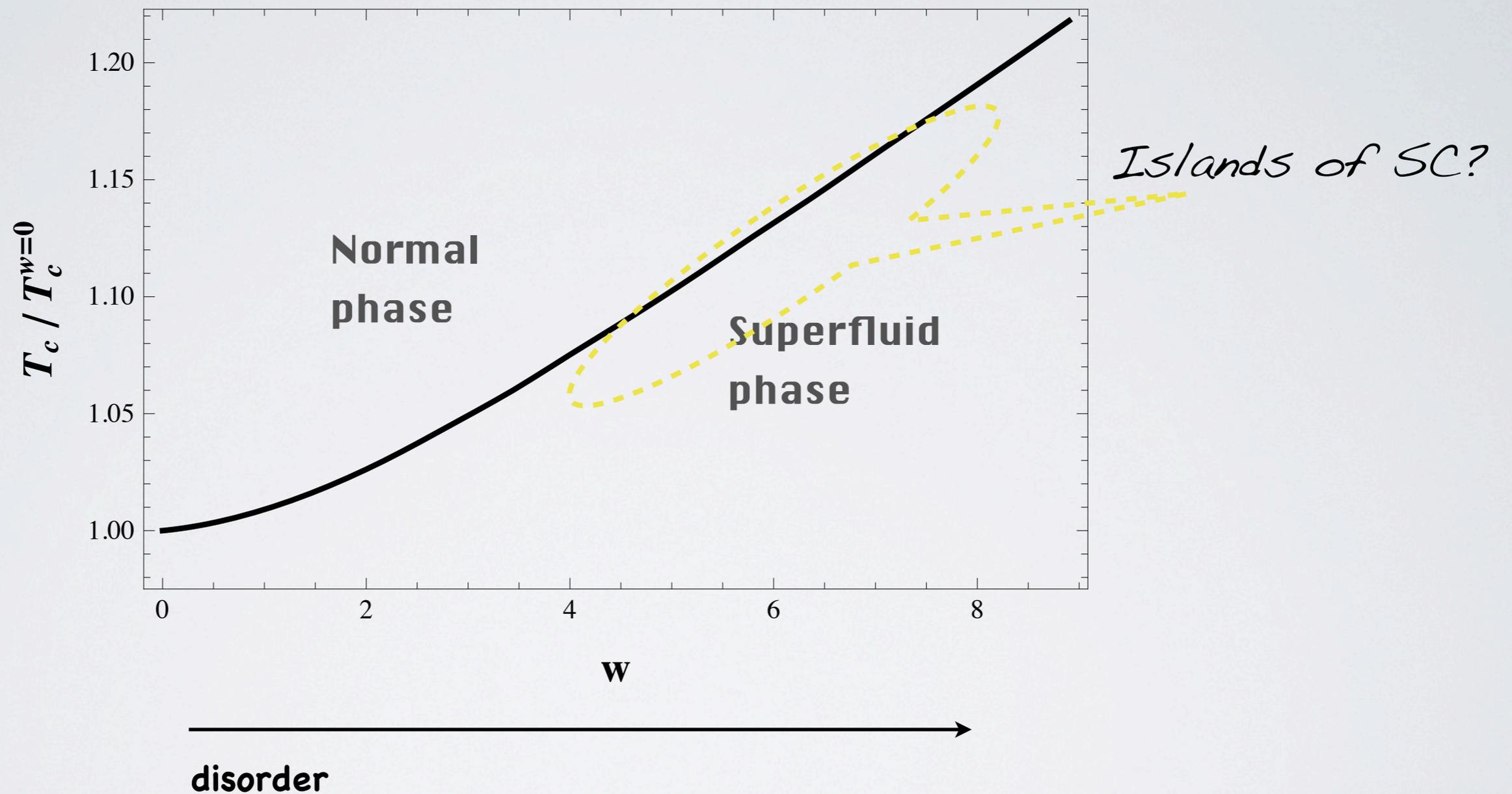
★ Enhancement of SC

Spatial average of the condensate



★ Enhancement of SC

Phase Diagram



Seen before in CM (hard-core bosons)

- 'Disorder-induced superfluidity', Dang et al, Phys. Rev. B 79, 214529

★ Spectrum 'renormalization'

>>> Noisy chemical potential

$$\mu(x) = \mu_0 + \epsilon \sum_{k=k_0}^{k_*} \sqrt{S_k} \cos(kx + \delta_k) = \mu_0 + \epsilon \sum_{k=k_0}^{k_*} \frac{1}{k^\alpha} \cos(kx + \delta_k)$$

Power spectrum
Random phases
Strength of noise

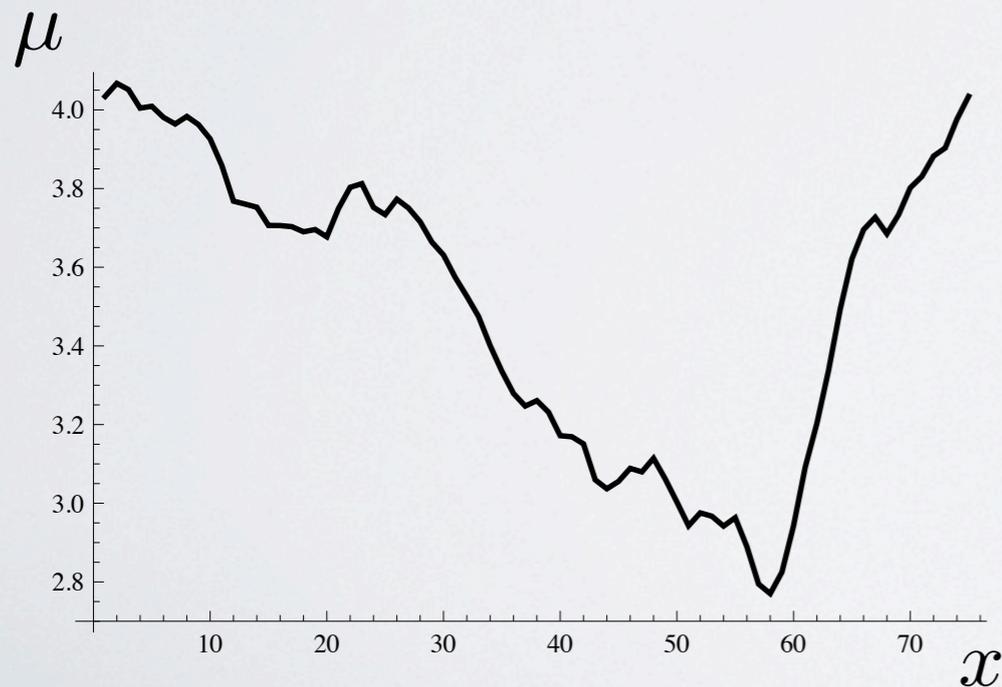
>input spectrum

$$S_k = \frac{1}{k^{2\alpha}}$$

>output spectra

Condensate
Charge density

$$S_k = \frac{1}{k^\Gamma} \quad ?$$



★ Spectrum 'renormalization'

>input spectrum

$$S_k = \frac{1}{k^{2\alpha}}$$

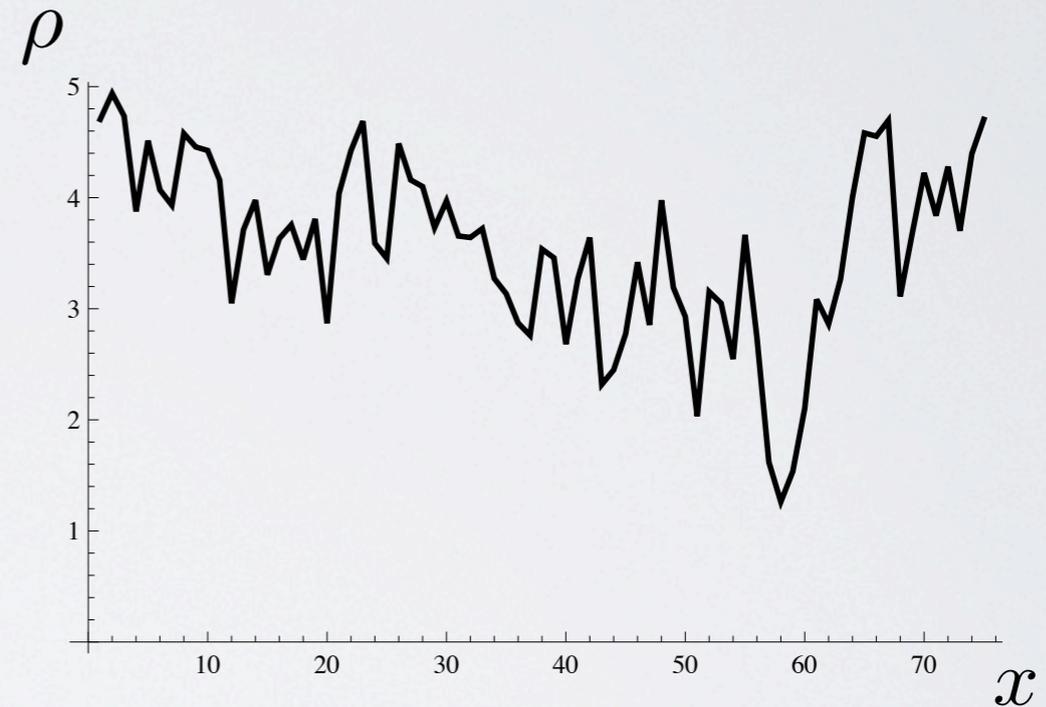
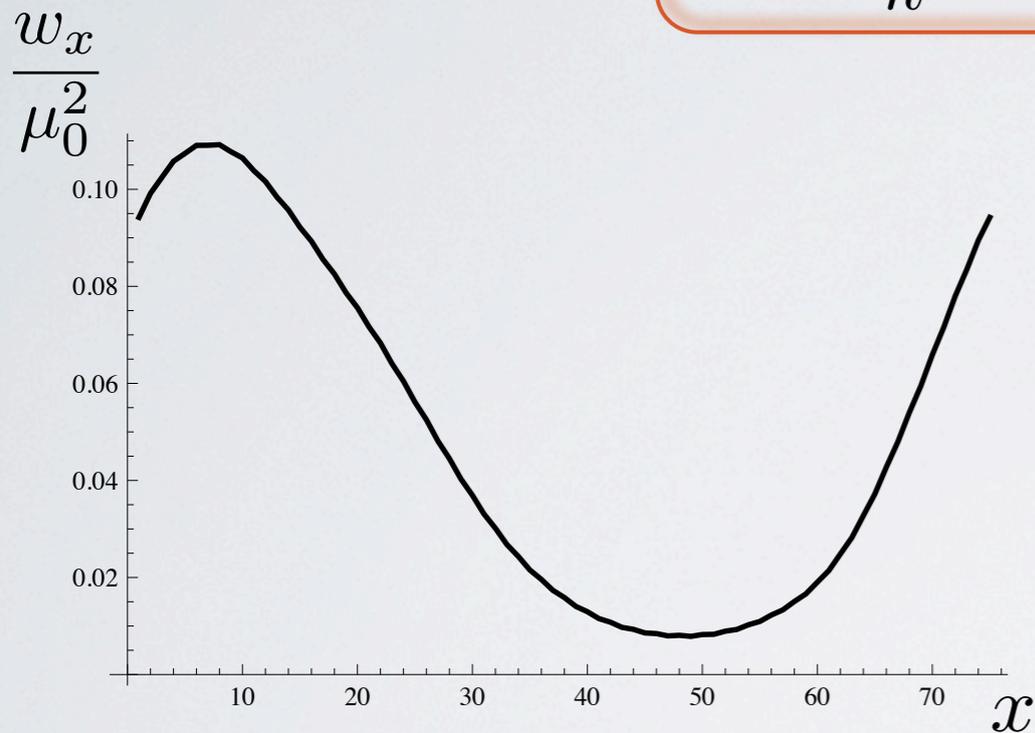
> OUTPUT

Condensate

$$S_k = \frac{1}{k^{2\alpha+4}}$$

Charge density

$$S_k = \frac{1}{k^{2\alpha-2}}$$



Hints of universality

- **S-wave [1308.1920]**
- **[Hartnoll&Santos 1402.0872]**
- **Fundamental matter (D3-D5) [w/ M. Araújo, J. Lizana, I.S. Landea]**
- **FT: noisy U(1) @ finite T [D. Musso, I.S. Landea]**

A taste of 'disordered conductivities'

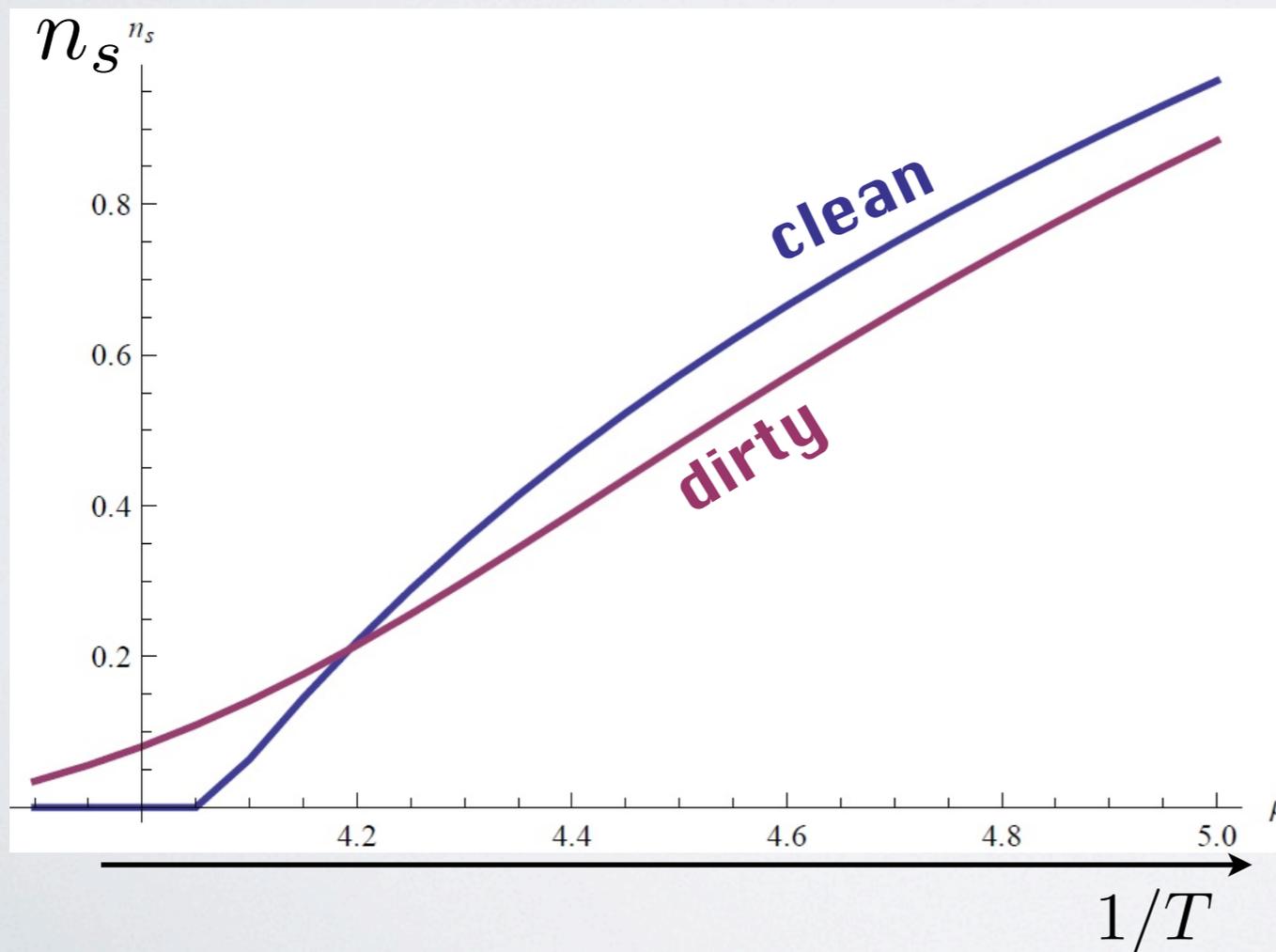
[WORK IN PROGRESS!]

[see also Ryu et al 1103.6068]

- **STUDY FLUCTUATIONS** ($\mathbf{a}_x \sim \mathbf{j}_x$) [in the SC phase they'll see the noise, even in the probe limit]

- **AVERAGED CONDUCTIVITY** $\sigma_x(\omega) = \frac{\langle j_x(x, \omega) \rangle}{E_x(\omega)}$

- **SC PHASE:** $\sigma_{DC} \rightarrow \infty$. **SUPERFLUID DENSITY** n_s : $\sigma_x \approx n_s \left(\pi \delta(\omega) + \frac{i}{\omega} \right)$



> **s-wave holo SC**

> **fixed noise strength**

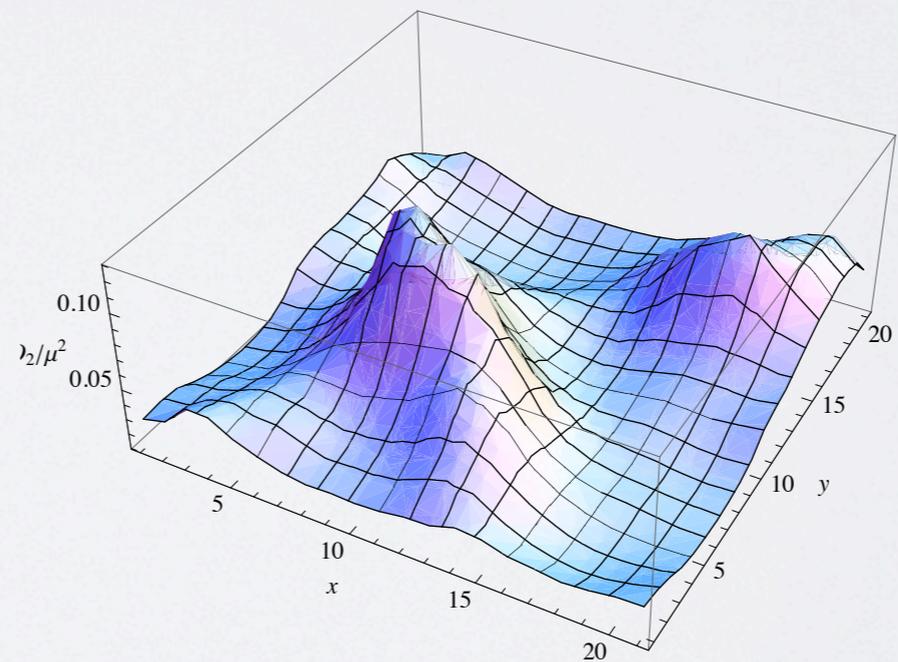
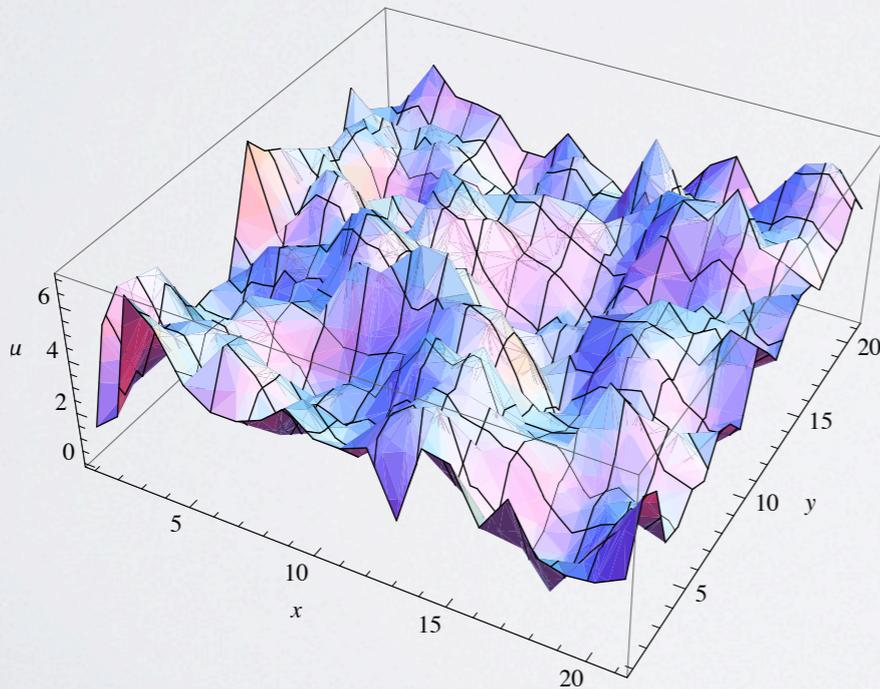
> FUTURE & ONGOING

>Disordered holo SCs: both s- and p-wave ✓

>Enhancement of SC & 'spectrum renormalization' ✓
(thermo limit 0K)

>Conductivity of disordered strongly coupled systems [....% % % %]

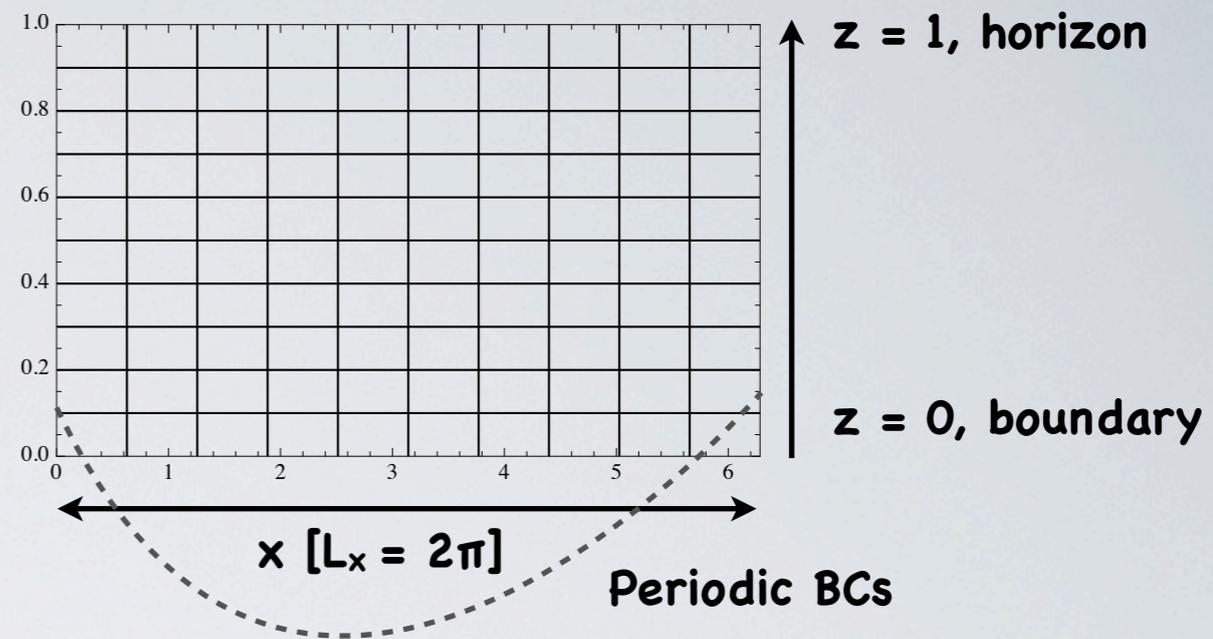
> Dirty Thin Films (islands of superfluidity?) [....% % % %]





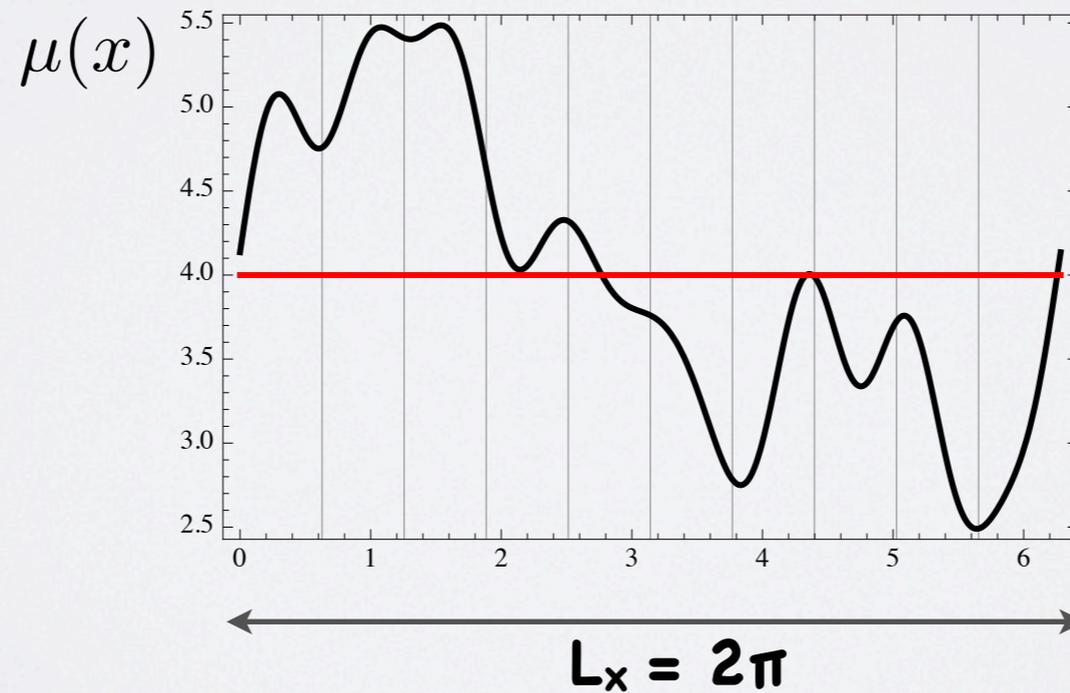
> Noisy chemical potential

• GRID



$$\mu(x) = \mu_0 + \epsilon \sum_{k=k_0}^{k_*} \sqrt{S_k} \cos(kx + \delta_k) = \mu_0 + \epsilon \sum_{k=k_0}^{k_*} \frac{1}{k^\alpha} \cos(kx + \delta_k)$$

Power spectrum
Random phases
Strength of noise



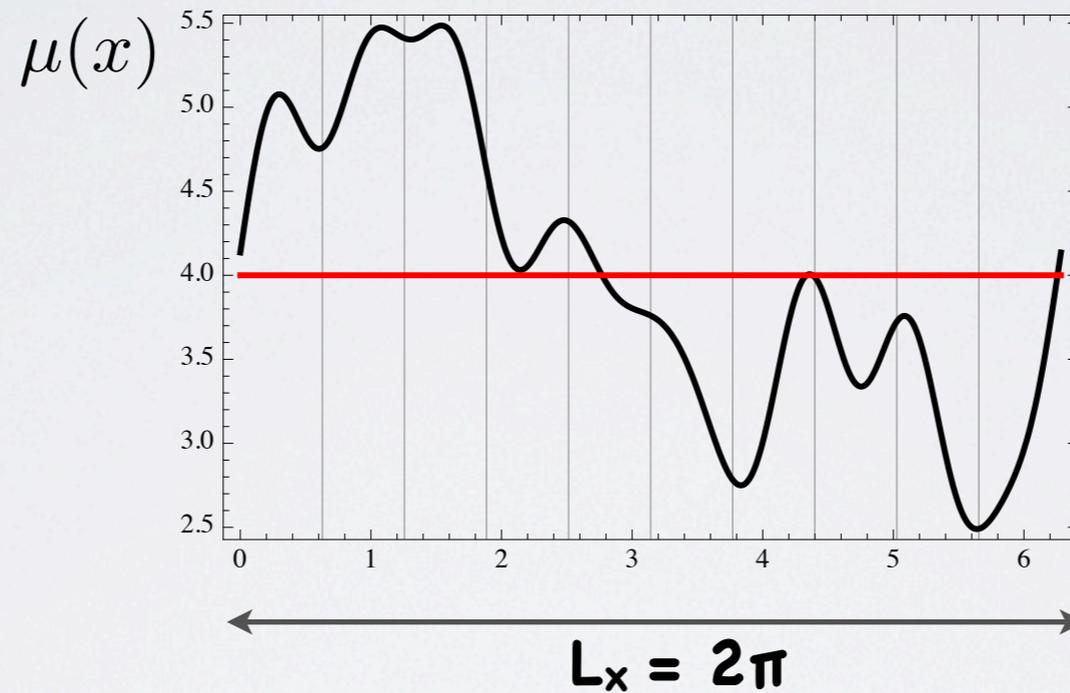
> Noisy chemical potential

- NOISE THROUGH RANDOM PHASES

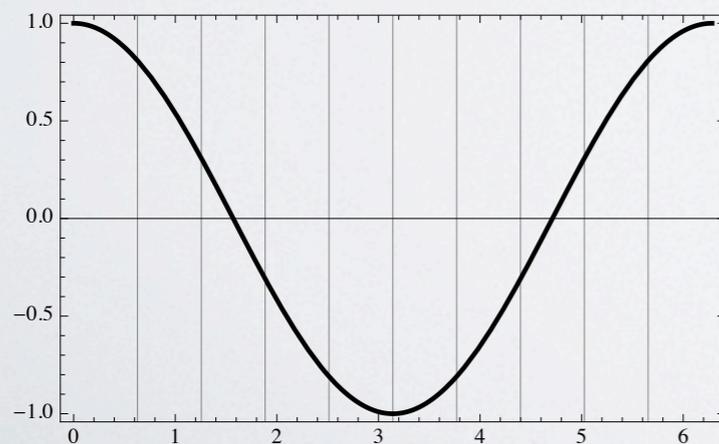
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Power spectrum
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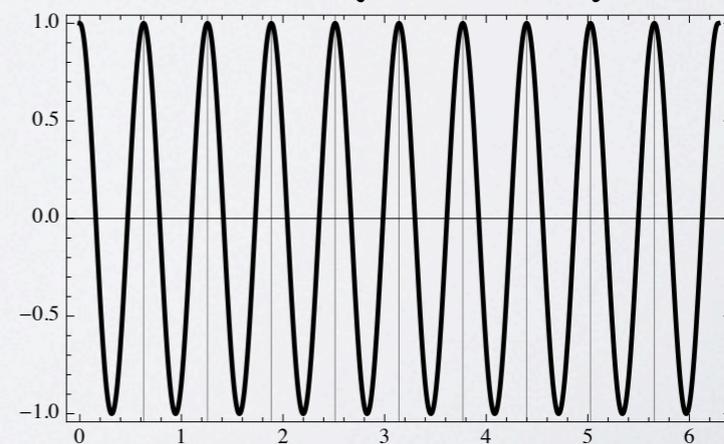
- SYSTEM ON A GRID



K_0 (IR scale)



K^* (UV scale)



> Thermodynamic limit

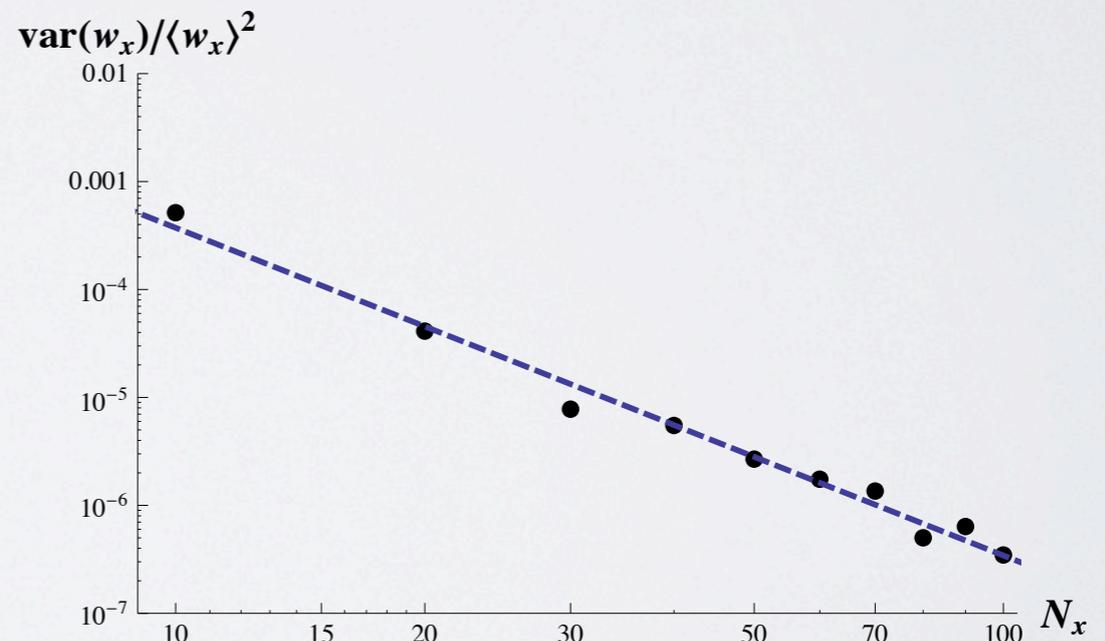
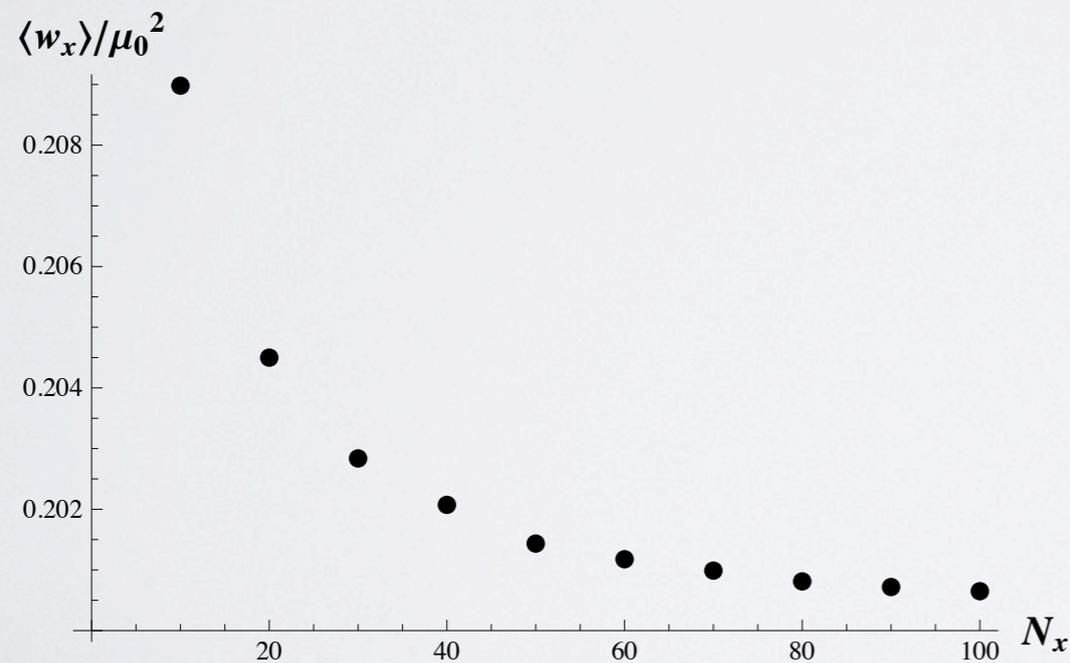
- Thermo limit: Noise correlation length \ll System length

> Flat spectrum noise: correlation length $\propto 1 / (\text{grid size})$

- Condensate and Charge density are self-averaging in the thermo limit:

> X_n is self-averaging when
$$\frac{\langle X_n^2 \rangle - \langle X_n \rangle^2}{\langle X_n \rangle^2} \rightarrow 0$$

Condensate



$$\log(\text{var}(w_x) / \langle w_x \rangle^2) = -0.90 - 3.03 \log(N_x)$$

> Thermodynamic limit

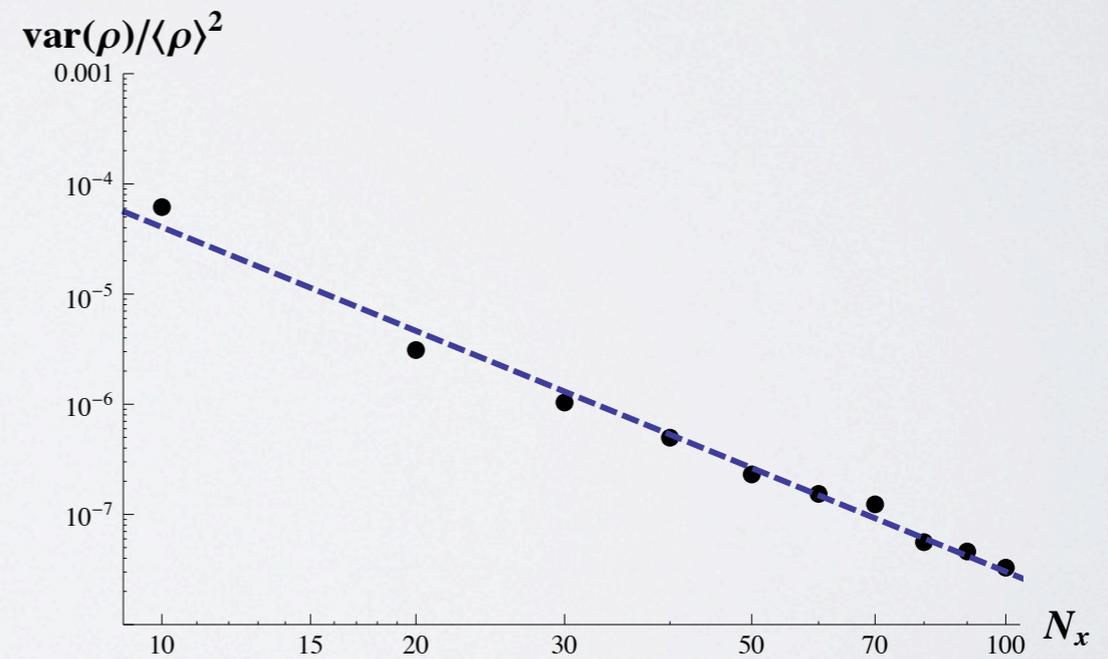
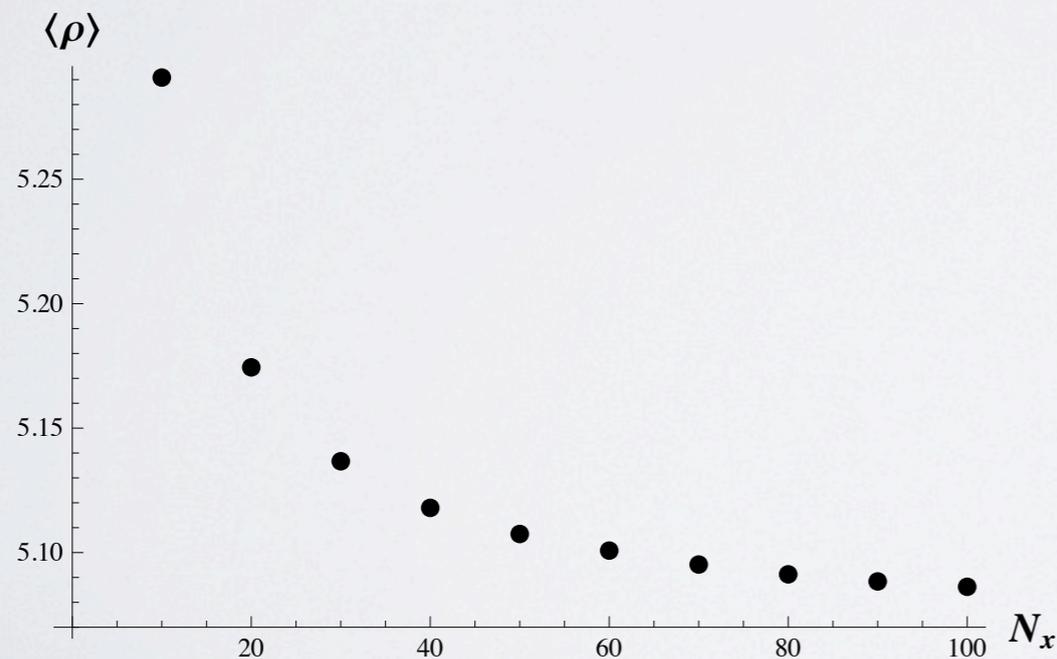
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Charge density



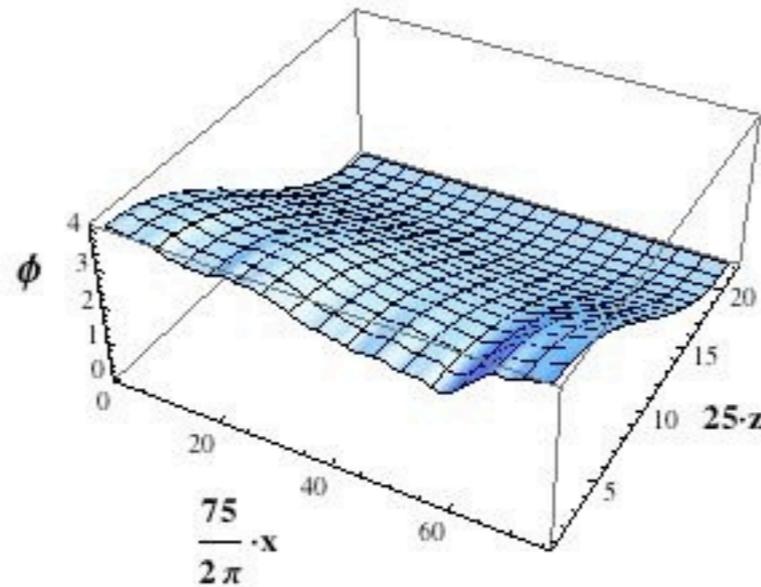
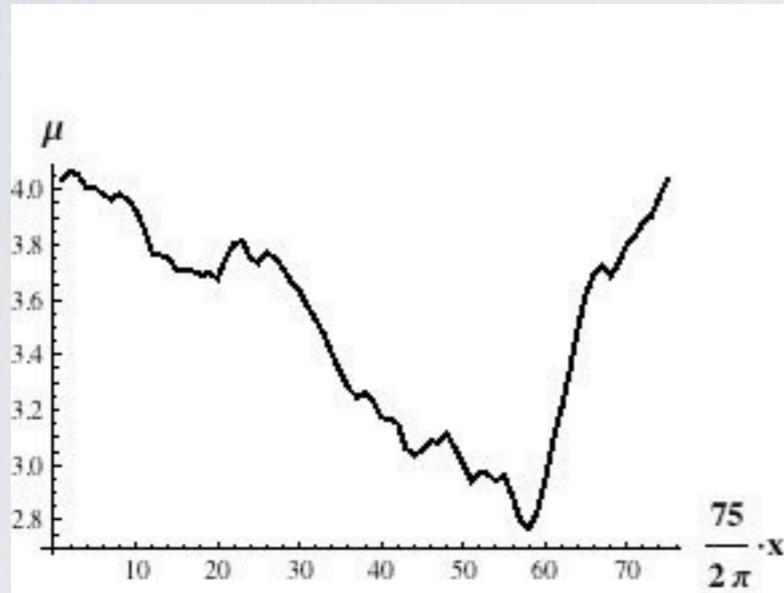
$$\log(\text{var}(\rho)/\langle \rho \rangle^2) = -2.92 - 3.13 \log(N_x)$$

> Simulation #1

$$\mu(x) = \mu_0 + \epsilon \sum_{k=k_0}^{k_*} \sqrt{S_k} \cos(kx + \delta_k) = \mu_0 + \epsilon \sum_{k=k_0}^{k_*} \frac{1}{k^\alpha} \cos(kx + \delta_k)$$

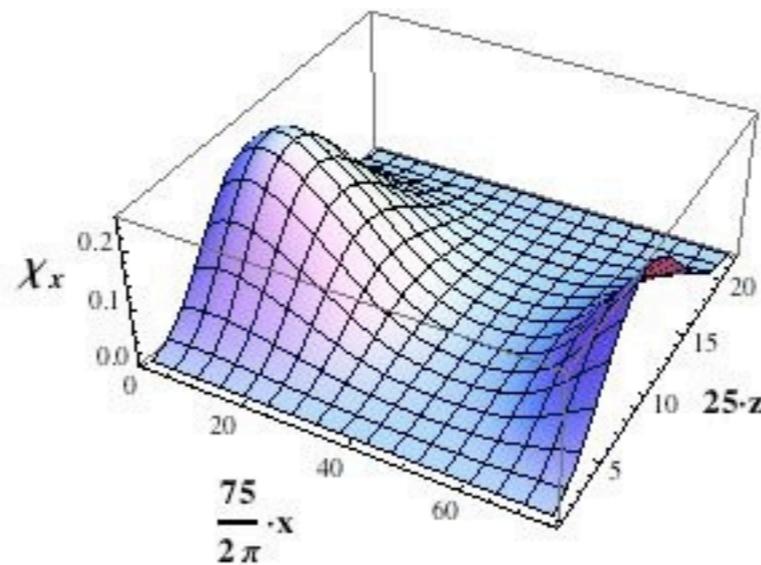
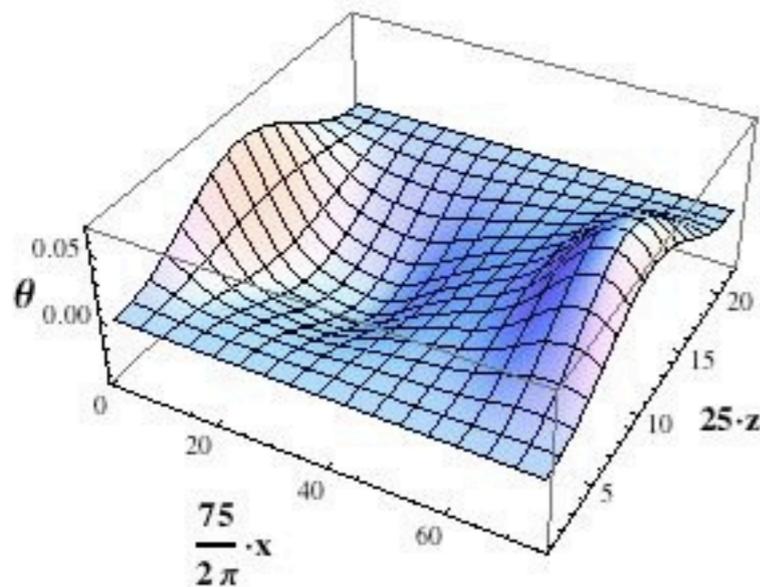
$$w = 25\epsilon/\mu_0$$

- $\mu_0 = 3.50$, $\alpha = 1.50$, $w = 3.50$ [$\mu_0 < \mu_c = 3.66$]



$$L_x = 2\pi \rightarrow K_0 = 1$$

$$N_z \times N_x = 25 \times 75$$

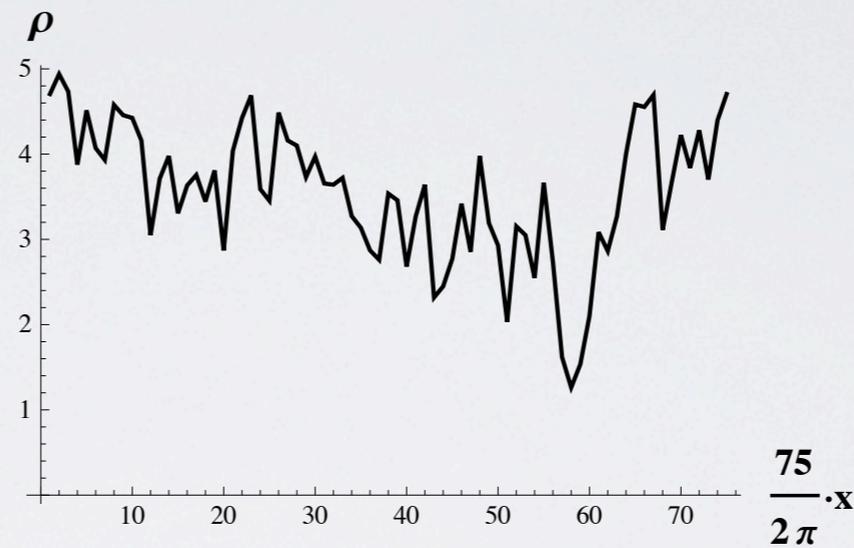
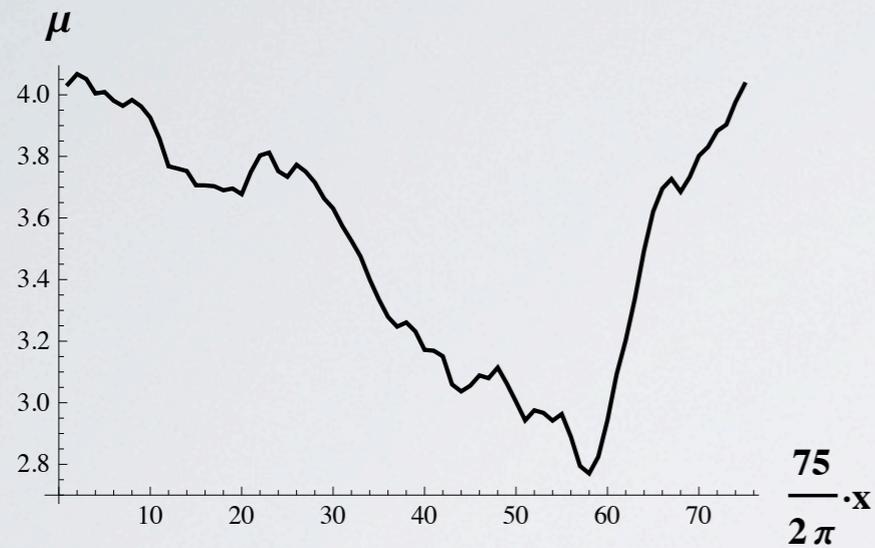


> Simulation #1

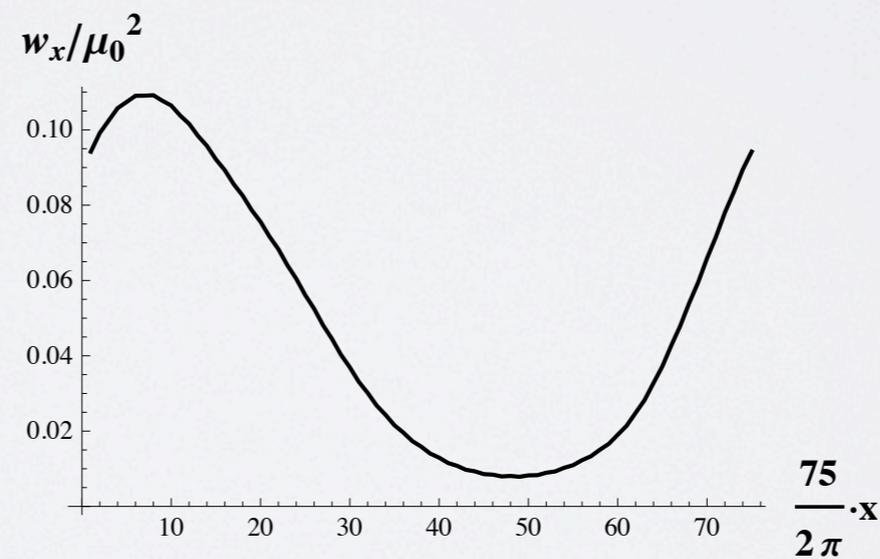
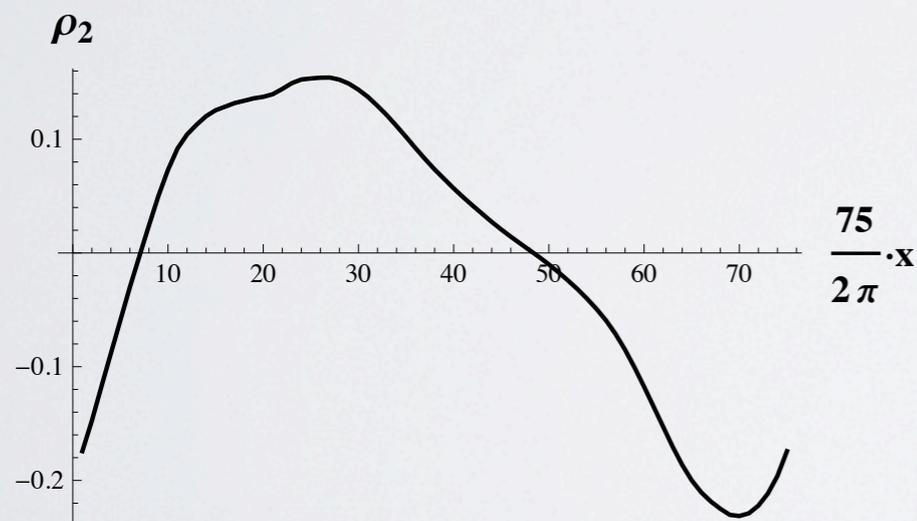
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$L_x = 2\pi \rightarrow K_0 = 1$
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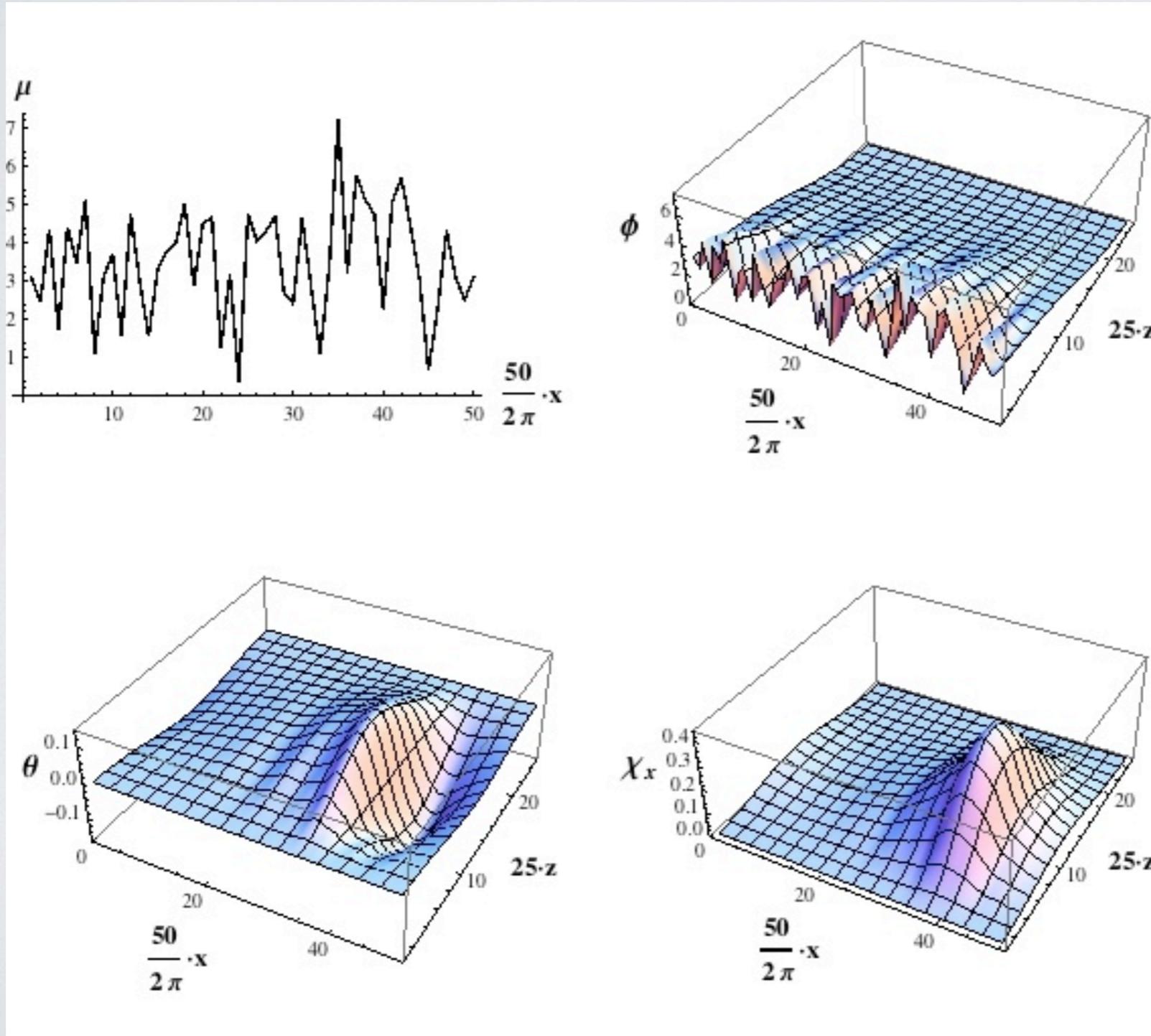
> Simulation #2

Flat Noise

$$\mu(x) = \mu_0 + \epsilon \sum_{k=k_0}^{k_*} \sqrt{S_k} \cos(kx + \delta_k) = \mu_0 + \epsilon \sum_{k=k_0}^{k_*} \frac{1}{k^\alpha} \cos(kx + \delta_k)$$

$$w = 25\epsilon/\mu_0$$

- $\mu_0 = 3.50$, $\alpha = 0$, $w = 3.50$ [$\mu_0 < \mu_c = 3.66$]



$$L_x = 2\pi \rightarrow K_0 = 1$$

$$N_z \times N_x = 25 \times 75$$

> Simulation #2

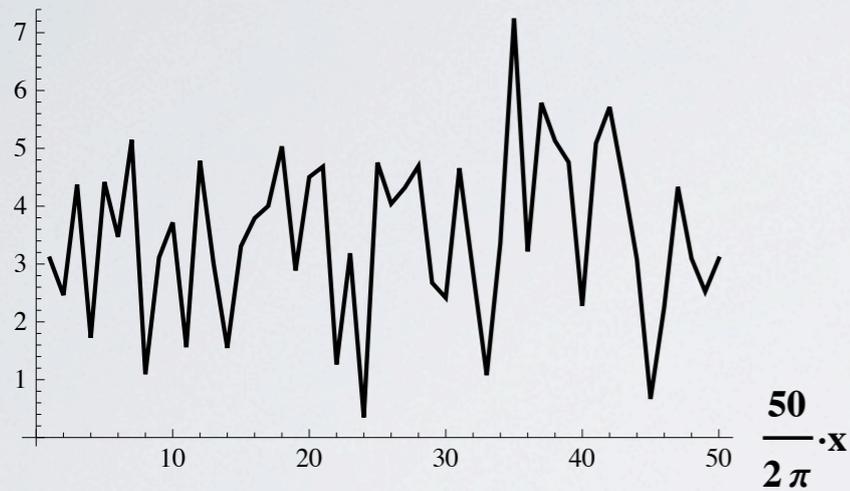
Flat Noise

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μ



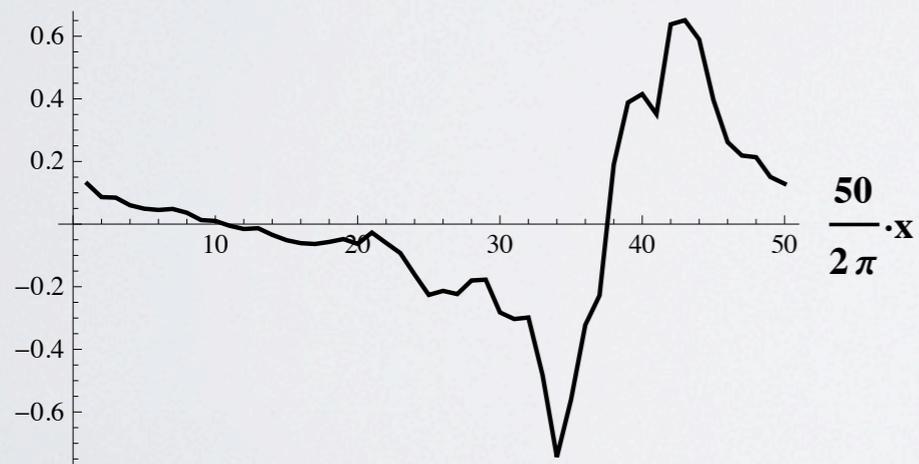
ρ



$$L_x = 2\pi \rightarrow K_0 = 1$$

$$N_z \times N_x = 25 \times 75$$

ρ_2



w_x/μ_0^2

