



Localized strong energy conversion regions and their use for identifying reconnection sites

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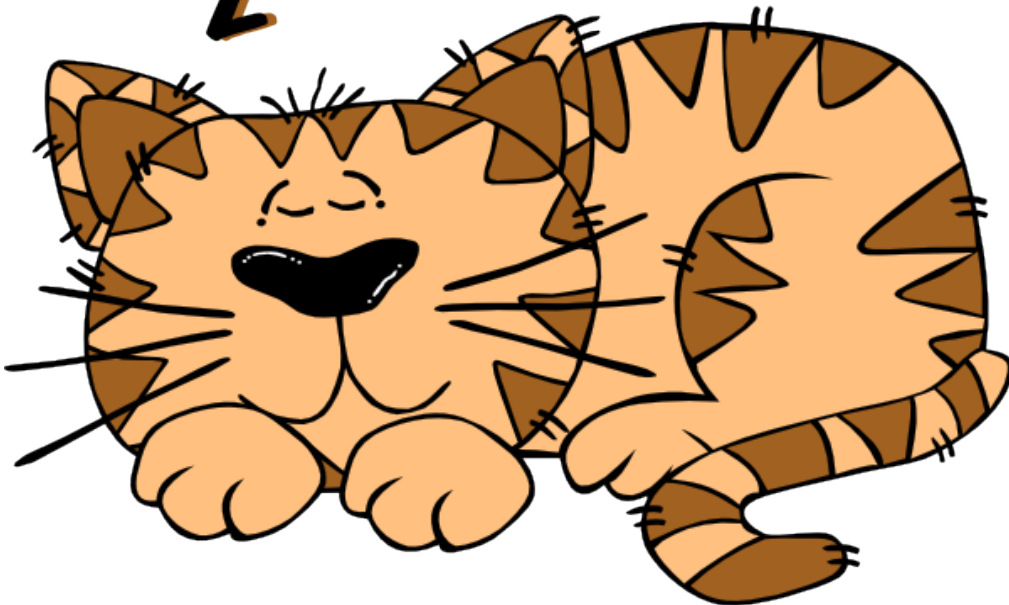
L. Andersson, A. Vaivads, T. Pitkänen

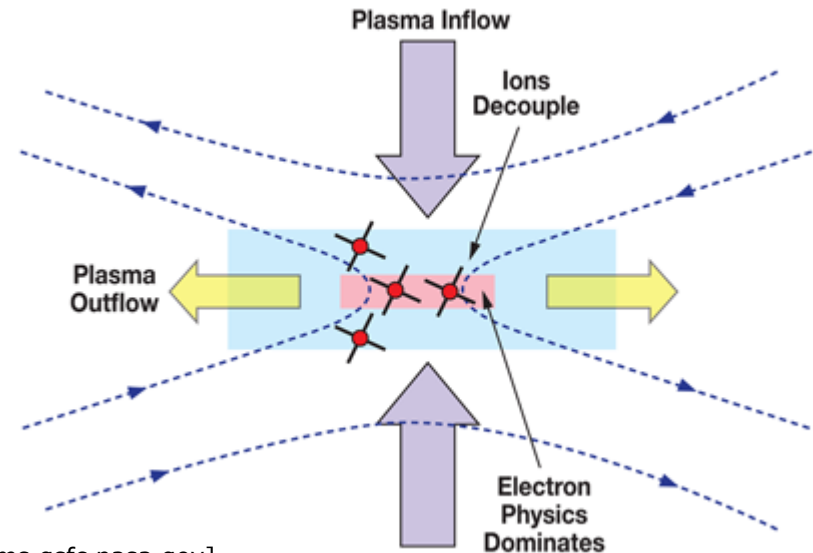
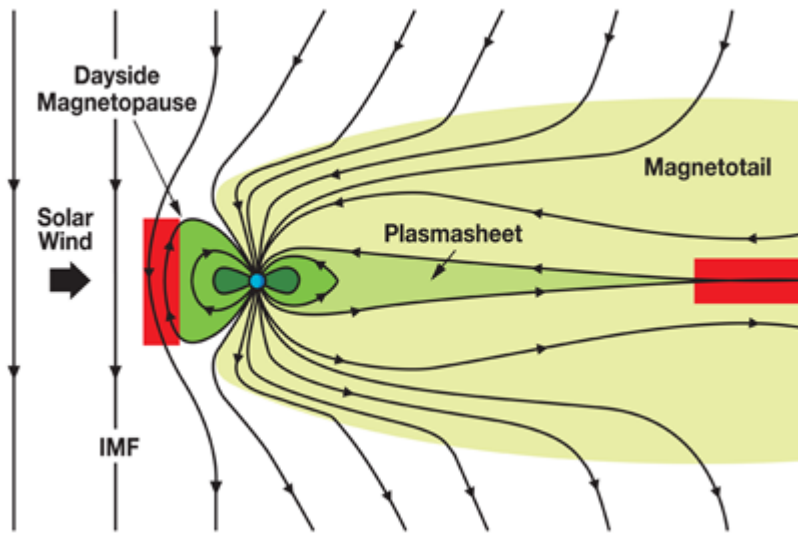
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Key points

- Use of power density $\mathbf{E} \cdot \mathbf{J}$ as a method (primary indicator) for identifying potential reconnection events from large amounts of measured spacecraft data.
- Approximate criterion for tail reconnection: $\mathbf{E} \cdot \mathbf{J}$ positive and large, $\mathbf{E} \cdot \mathbf{J} > \sim 20 \text{ pW/m}^3$ (consistent with theoretical arguments).
- $\mathbf{E} \cdot \mathbf{J}$ easily computed from multi-spacecraft data, e.g. MMS and Cluster.





[<http://mms.gsfc.nasa.gov>]

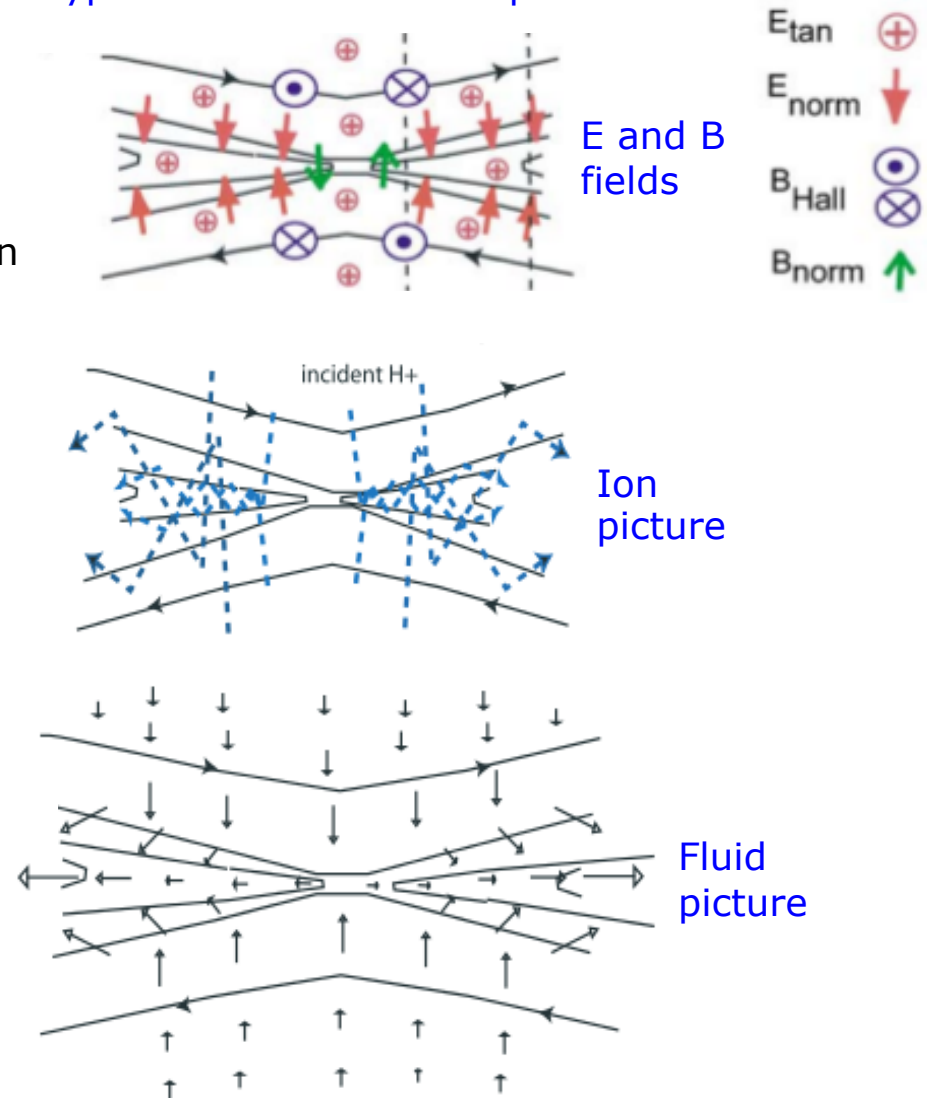
- Reconnection is a fundamental process.
- Change in magnetic topology. Magnetic \rightarrow kinetic energy.
- Goal for MMS, Magnetospheric Multiscale Mission: Understanding reconnection.
- How to identify potential events from large sets of data?
 - Look for jet reversals, Hall E and B fields, ...
 - Look for signatures of plasma acceleration!



Reconnection load

- Physical process: Plasma energization at reconnection.
- Right: Acceleration of ions in ion/fluid picture.
- Strong localized energy conversion regions can be probed with $\mathbf{E} \cdot \mathbf{J}$.
- Reconnection region should manifest itself as a load $\mathbf{E} \cdot \mathbf{J} > 0$ (accelerated jets).

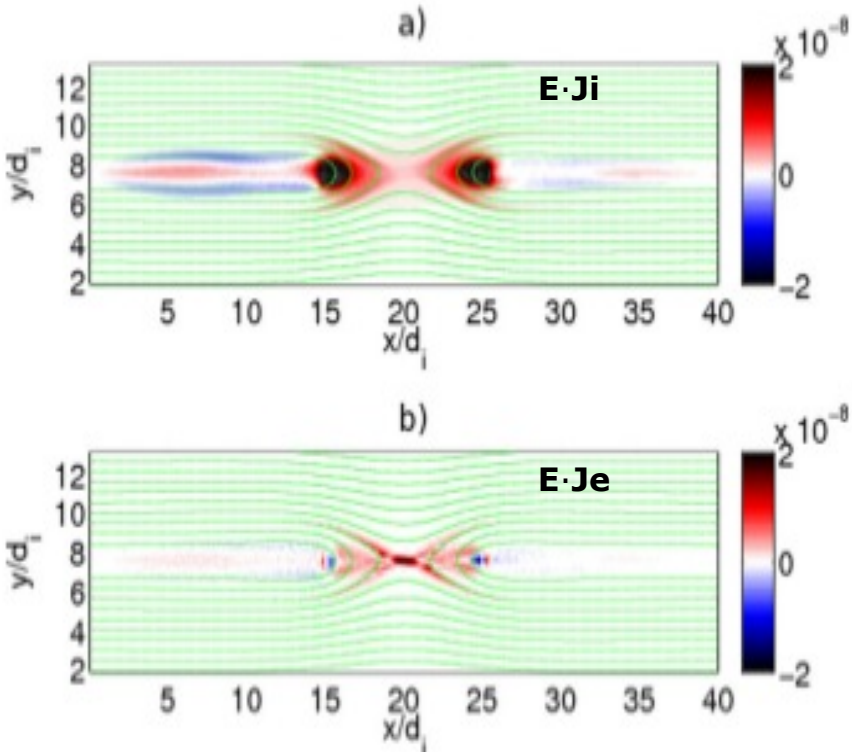
Typical tail reconnection picture:



[Wygant et al., JGR 2005]

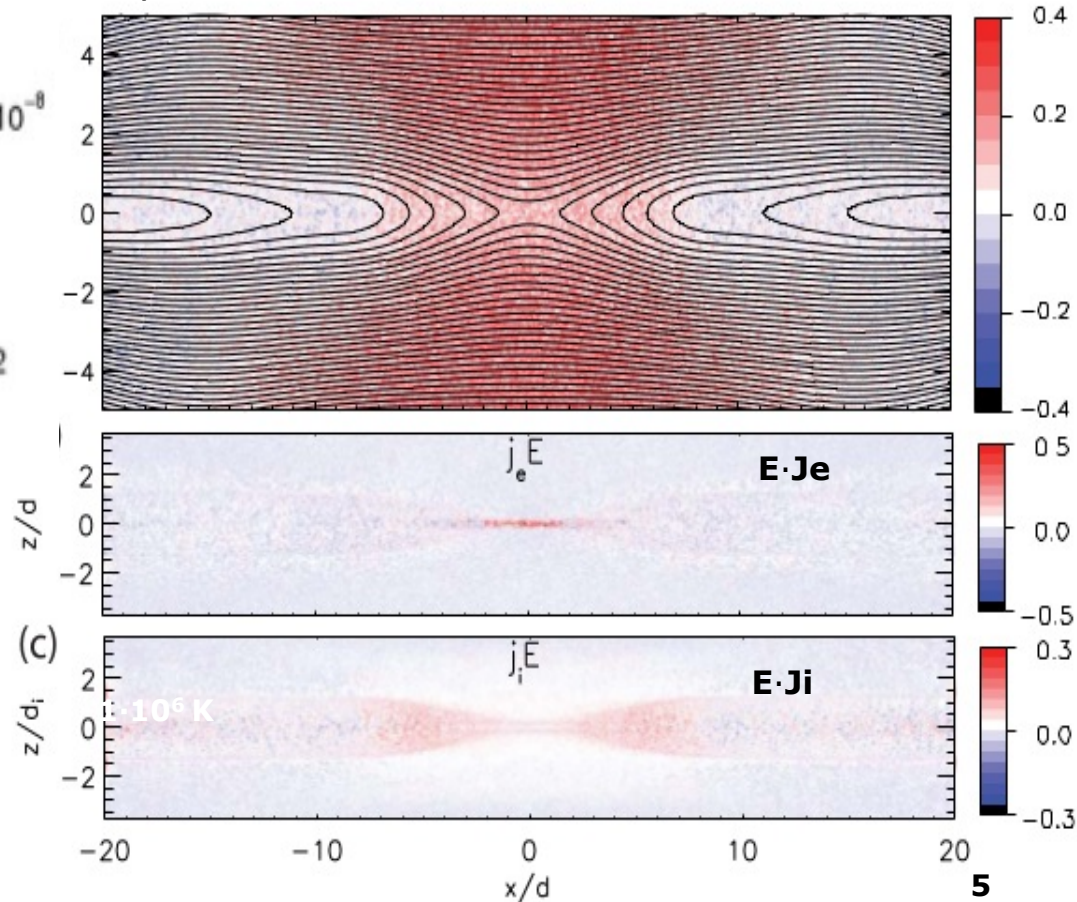
Reconnection load in simulations

- Ion and electron $\mathbf{E} \cdot \mathbf{J}$ in numerical simulations from cross-tail \mathbf{E} and \mathbf{J} .
- Width $\mathbf{E} \cdot \mathbf{J}_i \sim$ a few ion scales.
- (Electron region smaller.)

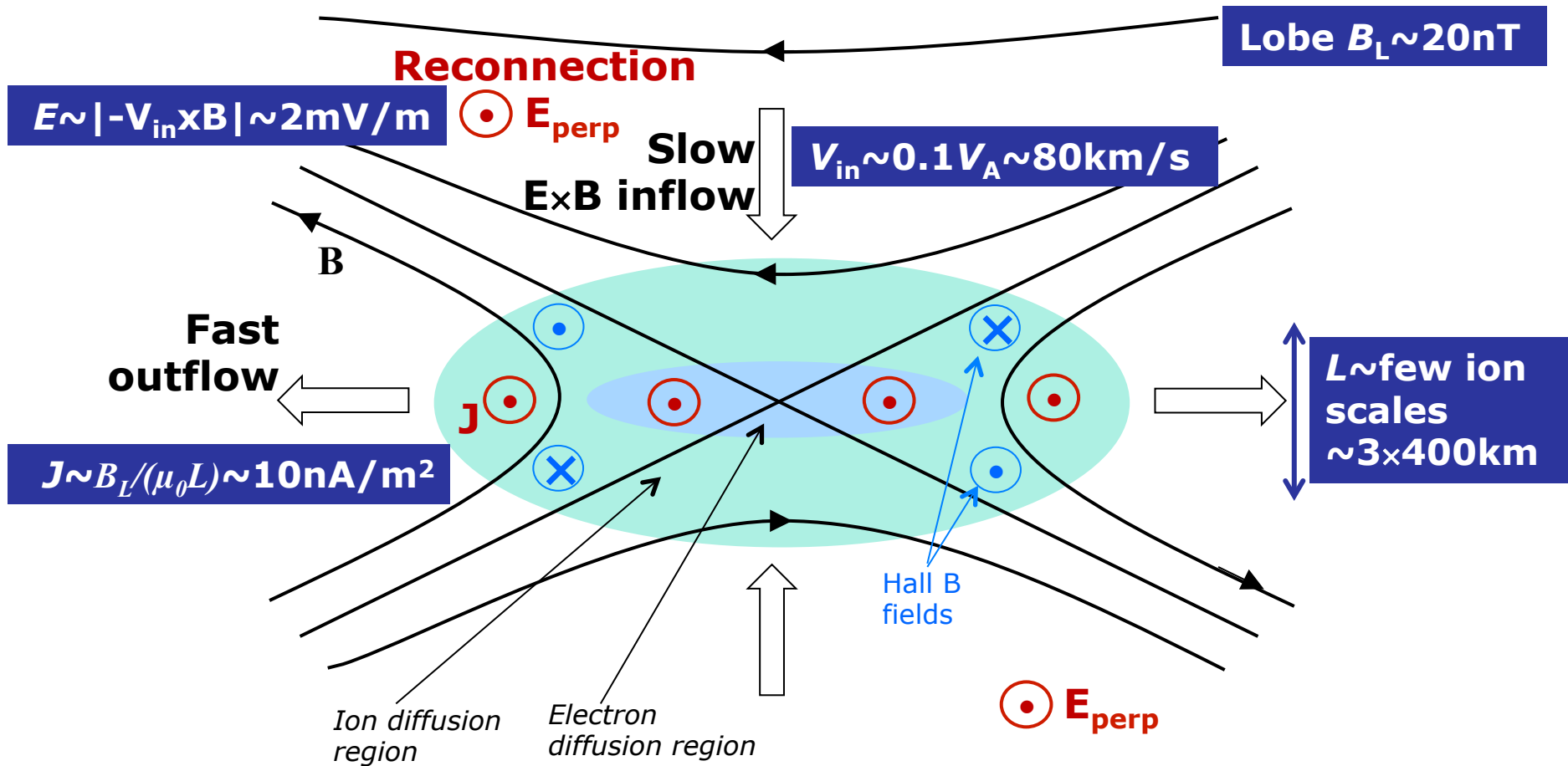


3D particle simulation PIC. [Lapenta et al., Phys. Plasmas, 2014]

2D particle simulation PIC. [Sitnov & Svisdak, JGR 2011]



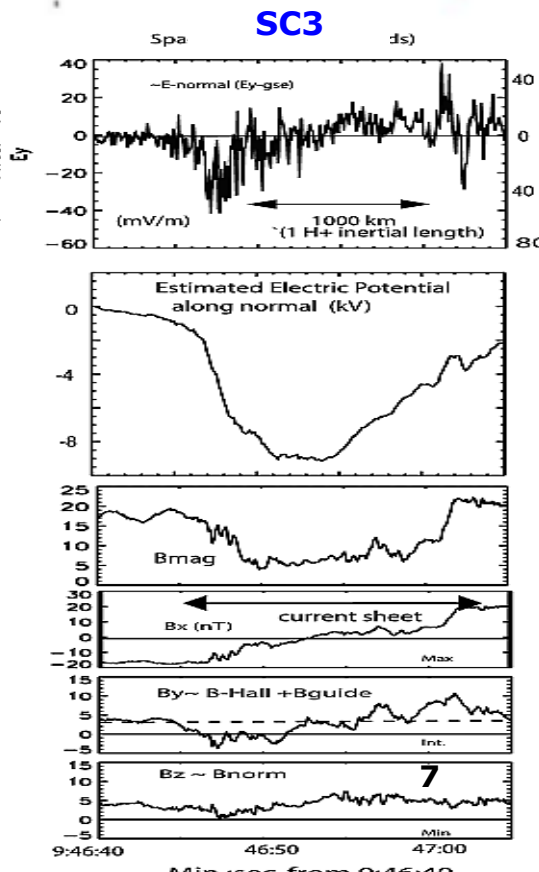
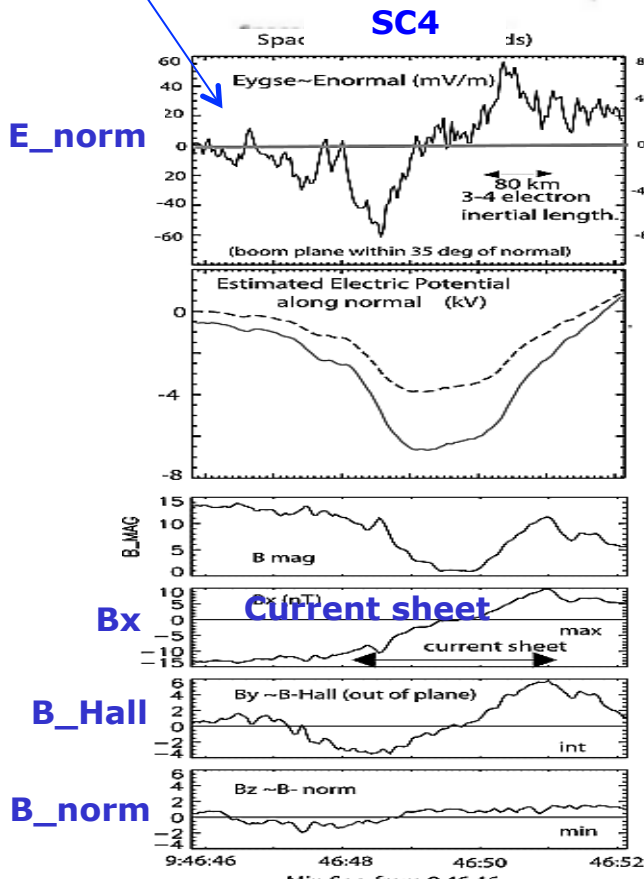
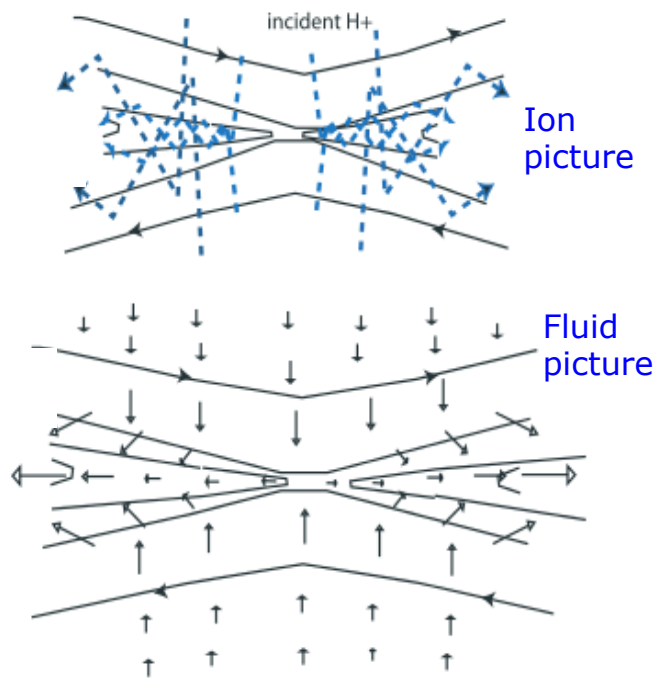
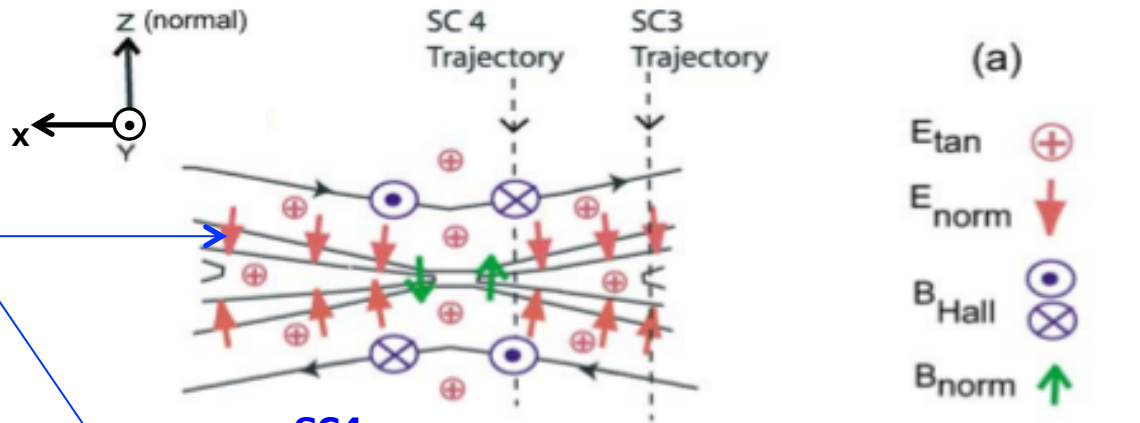
Tail reconnection load (typical energy change)



- Reconnection load (ion scales): $E \cdot J \sim (10 \text{ nA/m}^2) \cdot (2 \text{ mV/m}) \sim 20 \text{ pW/m}^3$
- This is larger than typical values in plasma sheet [Hamrin et al., JGR 2011]

Example: October 1, 2001

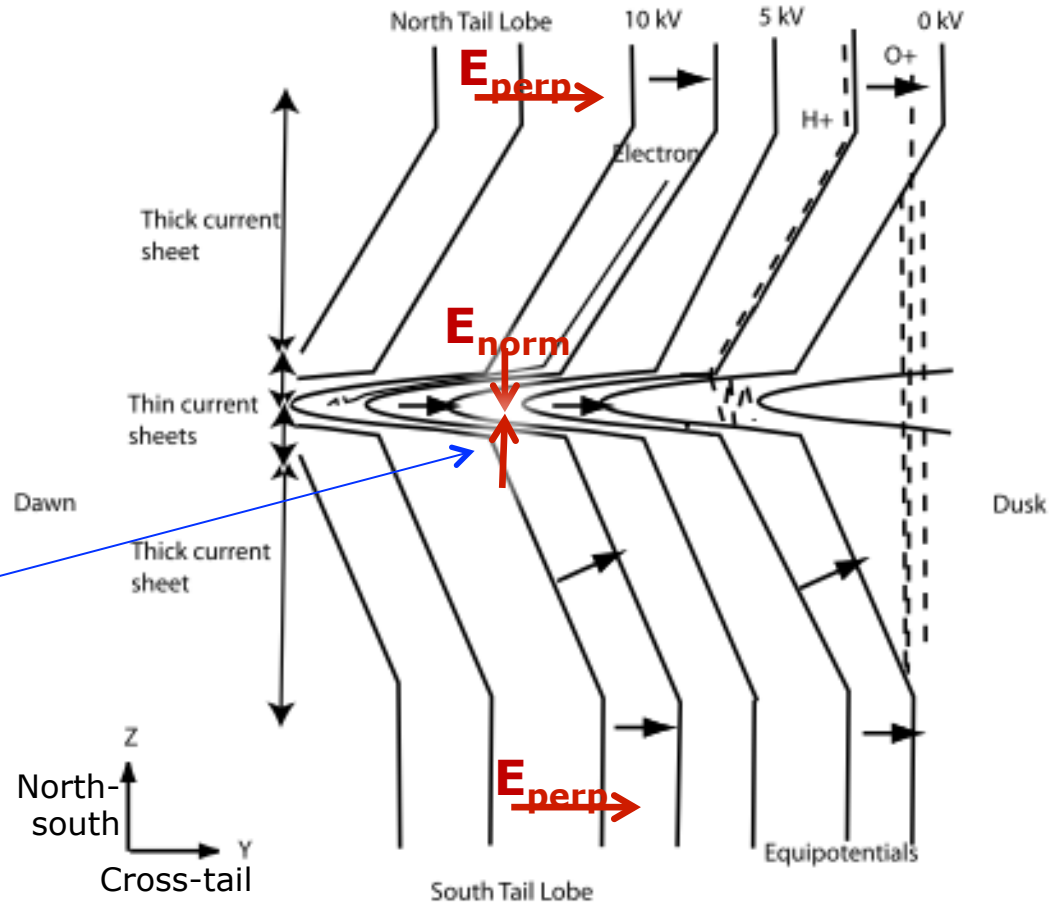
- Cluster observations, ~ 22.4 MLT. Several Cluster crossings due to tail flappings. [Wygant+, JGR 2005]
- Intense \mathbf{E}_{norm} (± 60 mV/m) observed in the ion decoupling region. 4-6 kV potential drop
- Ballistic acceleration of ions.
- Evidence for accelerated ion beams by \mathbf{E}_{norm} .



Example: October 1, 2001

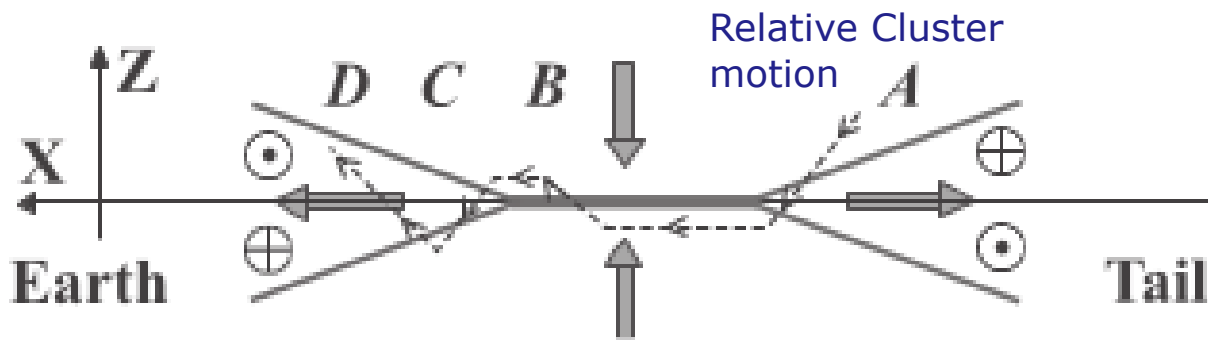
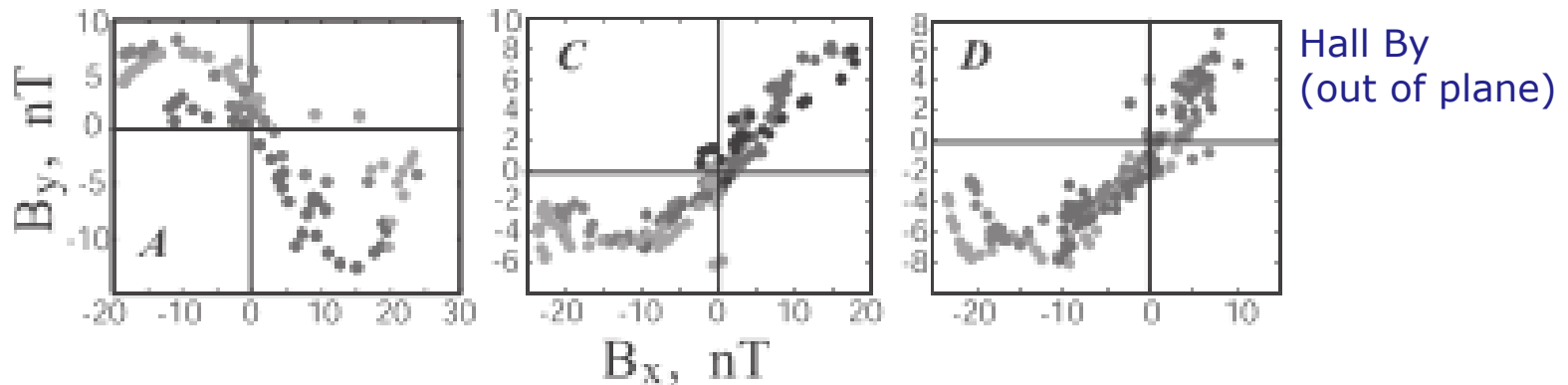
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- Ions accelerated by \mathbf{E}_{norm}

Electric equipotentials in YZ plane



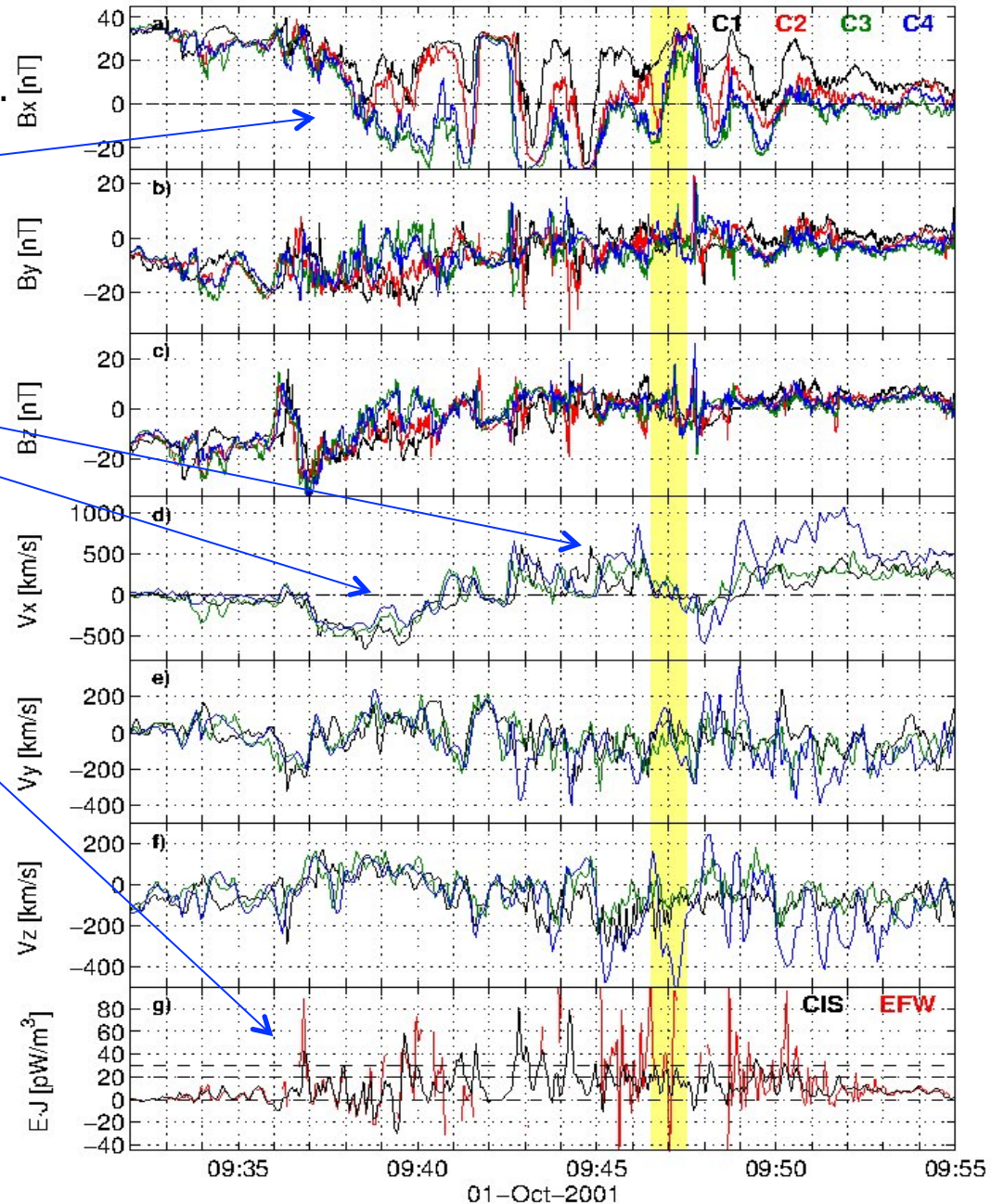
Example: October 1, 2001

- Specific event 09:47-09:51 UT also studied by Runov et al. [GRL, 2003].
- Observed:
 - Plasma jet reversals,
 - Bifurcated current sheet
 - Hall B_y at passages A, C, D
- Verified: current density can be computed with multi-spacecraft Curlometer method $\mathbf{J} \sim \nabla \times \mathbf{B} / \mu_0$.



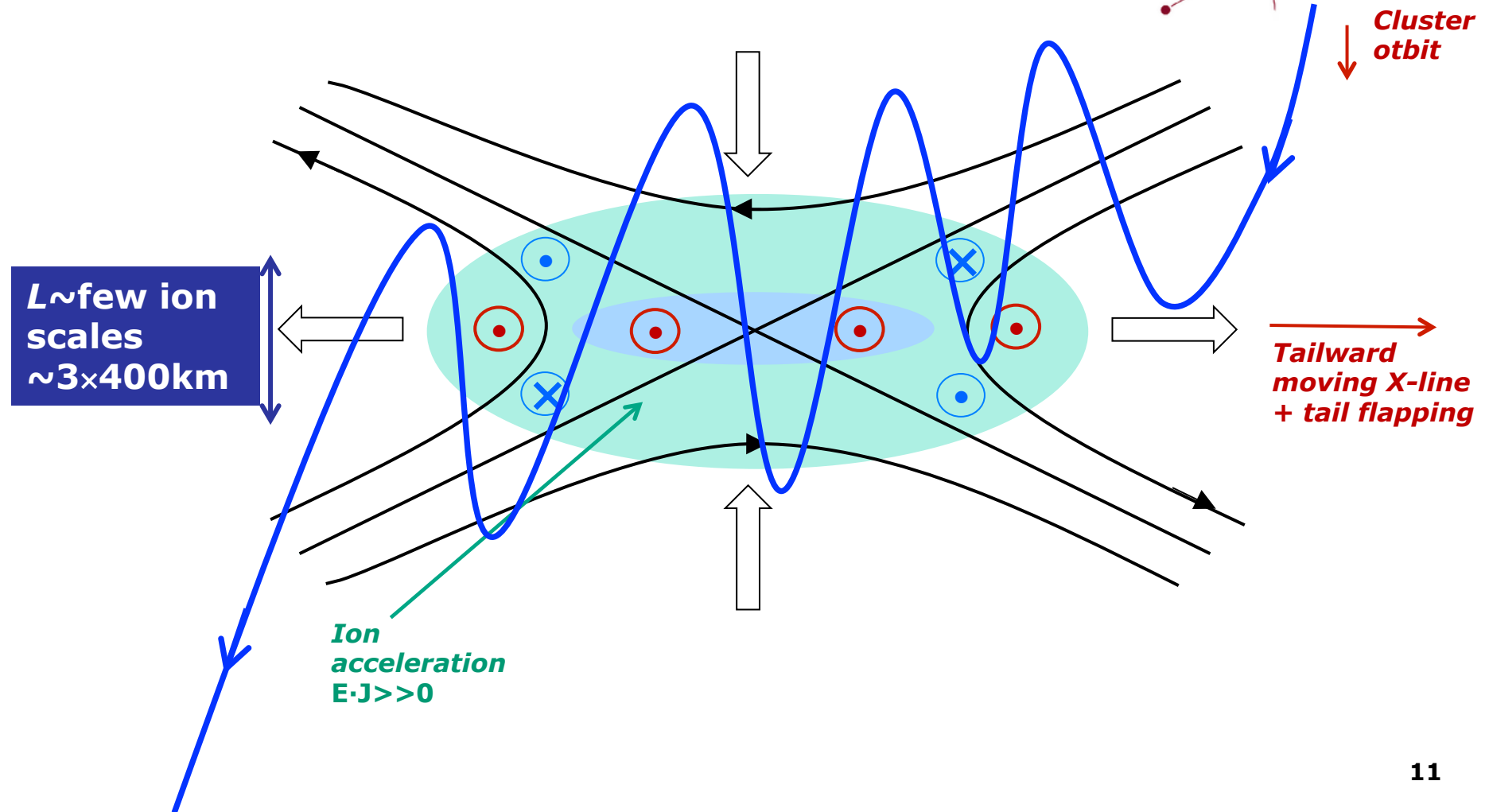
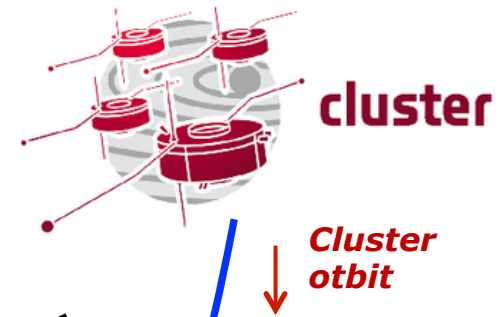
Details: October 1, 2001

- Main Wygant and Runov event in yellow.
- Tail flappings
- Jet reversals
- Power density $\mathbf{E} \cdot \mathbf{J}$ using
 - \mathbf{J} from Curlometer and \mathbf{E} from
 - CIS $\mathbf{E} \sim -\mathbf{V} \times \mathbf{B}$ (frozen-in) [black]
 - EFW (direct \mathbf{E} measurements) [red]
- **Power density is positive and large!**
- $\mathbf{E} \cdot \mathbf{J} > 20$ pW/m³ in general.



Typical scenario

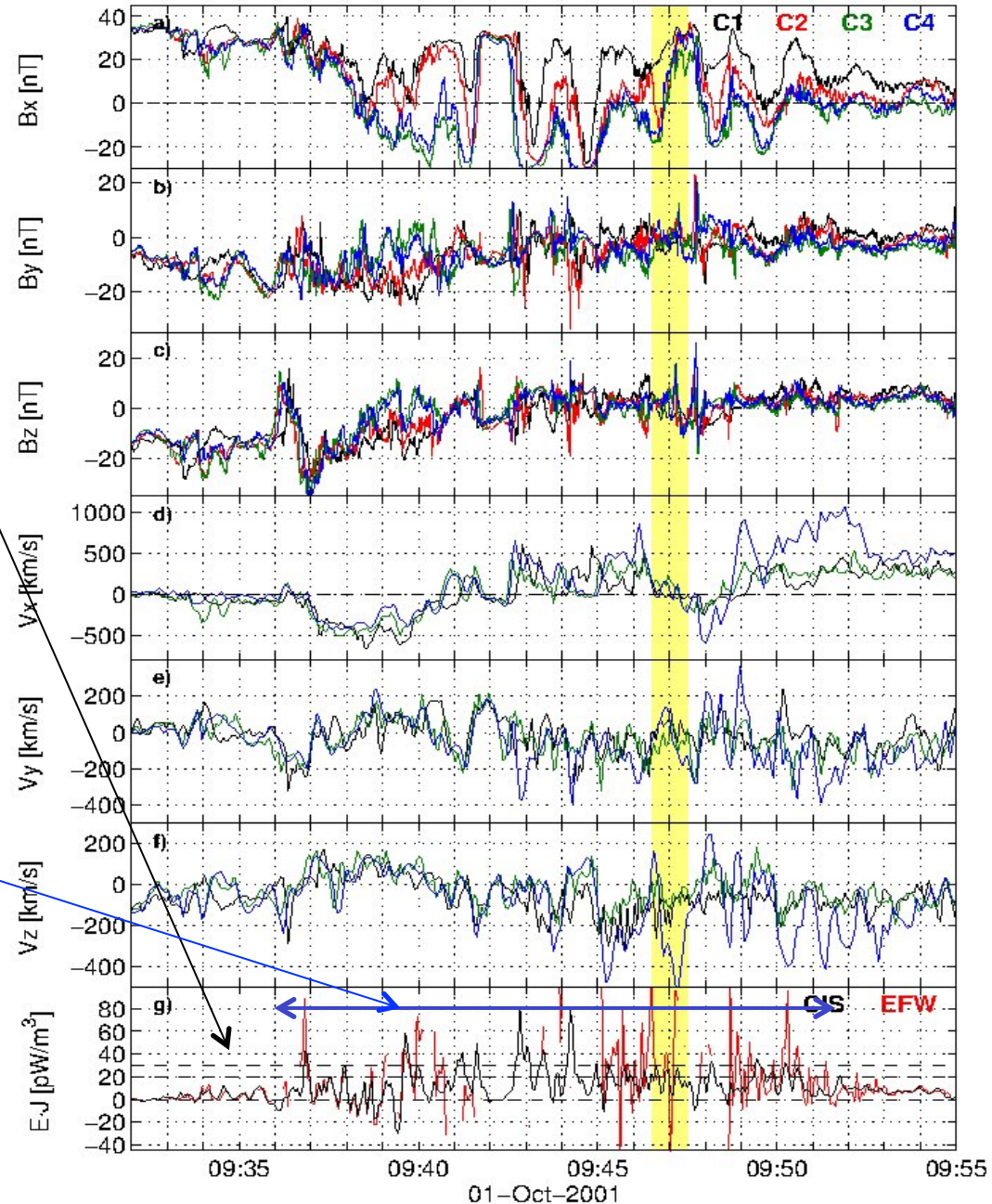
- Assume $L \sim 1200\text{km}$
- $V_{\text{sat}} \sim 1\text{km/s}$
- Time for Cluster to pass through diffusion region:
 $T \sim L/V_{\text{sat}} \sim 20\text{min}$



Recurrent Cluster sampling diffusion region

- Tail flappings → Recurrent Cluster observations of the diffusion region.

- Region with $\mathbf{E} \cdot \mathbf{J} > 0$ is of the order of 20 min!



Can we trust the Cluster measurements?

- Electric field:**

- **EFW:**

- Direct measurements of \mathbf{E} .
 - Full vectors not always available.
 - Cannot measure \mathbf{E}_{para} . ($\mathbf{E} \cdot \mathbf{B} = 0$)

- **CIS (HIA, CODIF):**

- $\mathbf{E} = -\mathbf{V} \times \mathbf{B}$ frozen-in condition.
 - Generally available.
 - Cannot measure \mathbf{E}_{para} .

- Current density:**

- **Curlometer:**

- Generally operational. Simple to use.
 - Smoothing: Resolves Cluster scales.

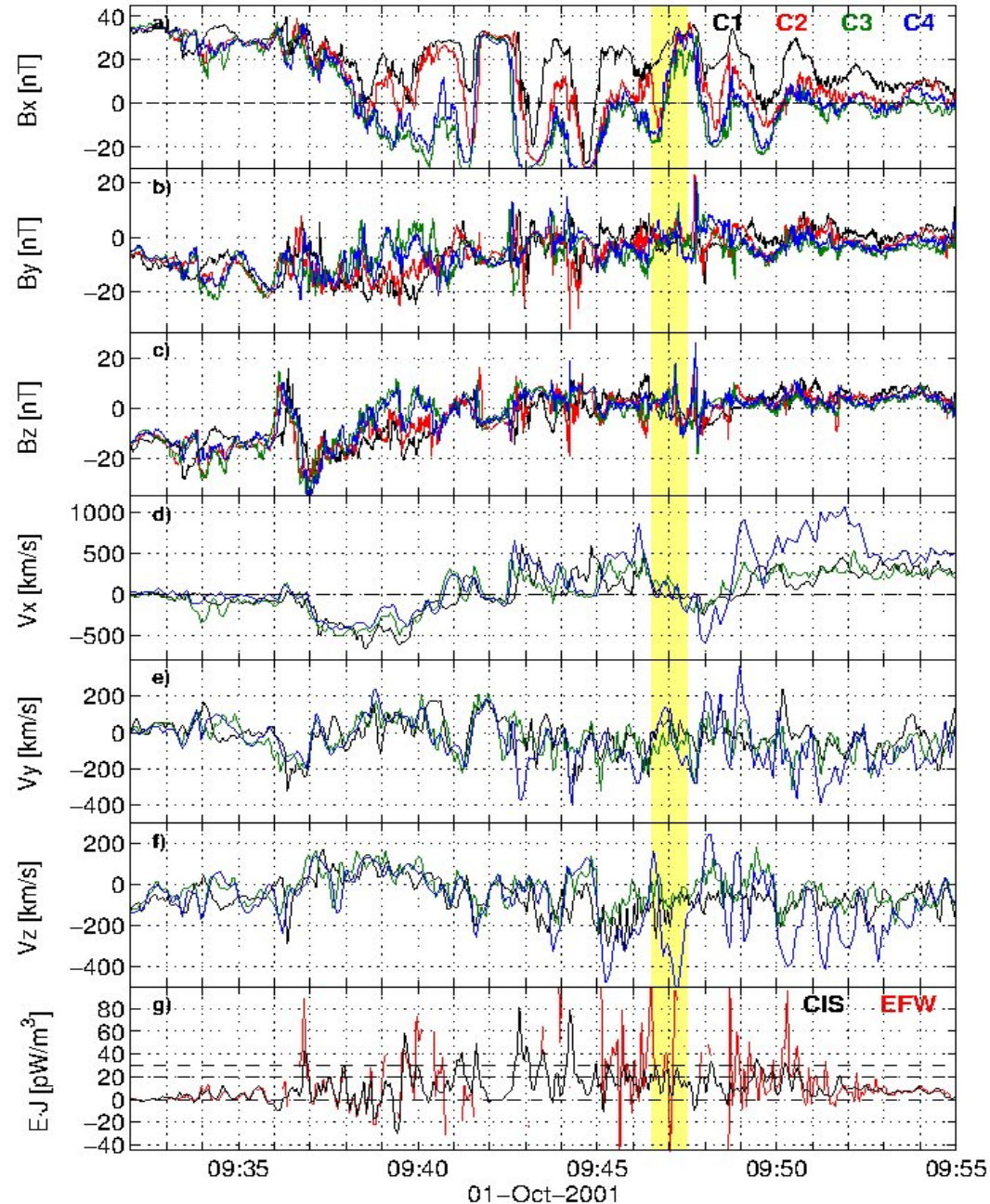
- **Single-spacecraft method:**

- Only operational at current sheets.
 - Higher resolution.



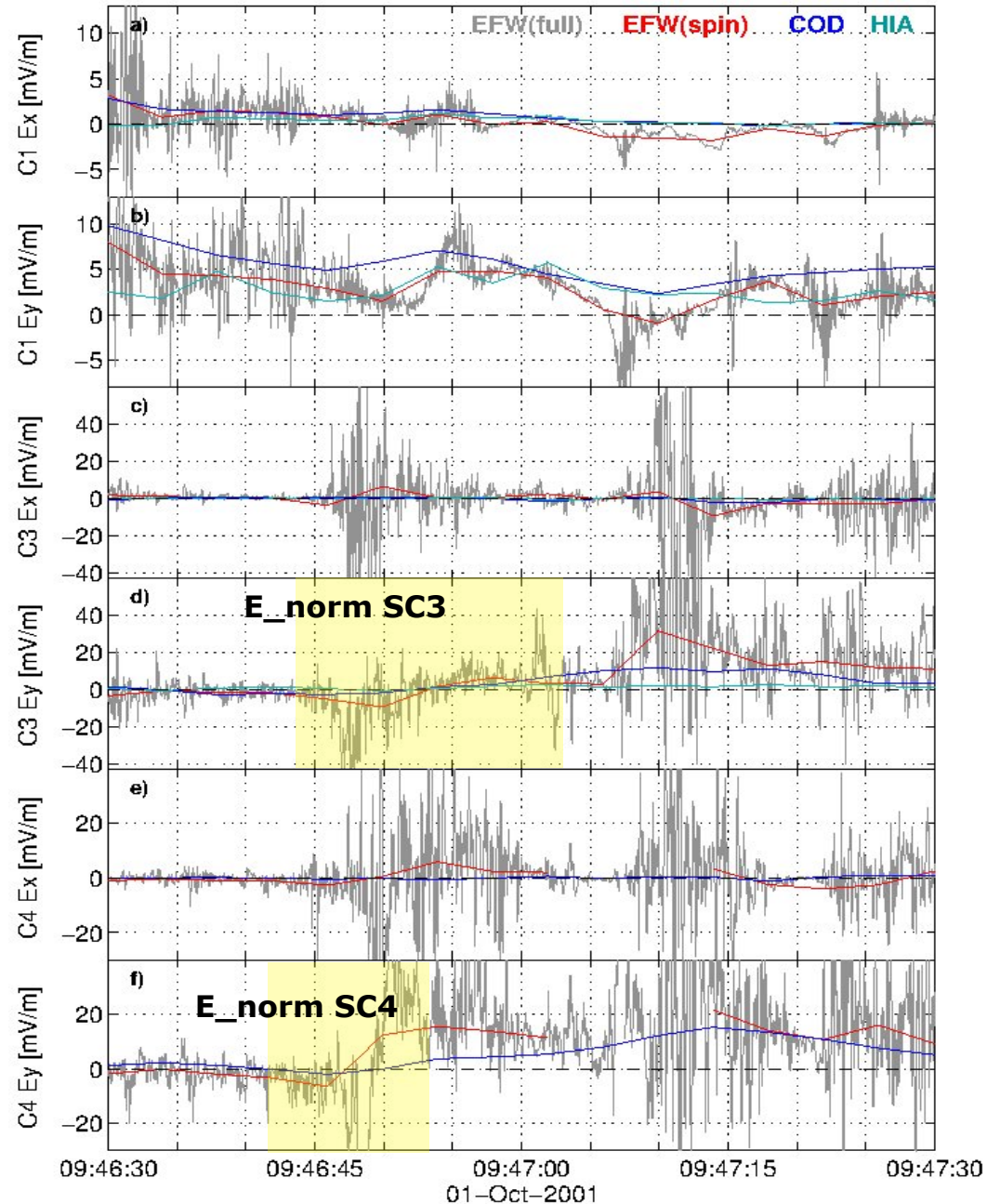
- Improvements with MMS?**

- 3D \mathbf{E} field measurements.
 - Smaller tetrahedron: more accurate Curlometer for small scale regions.



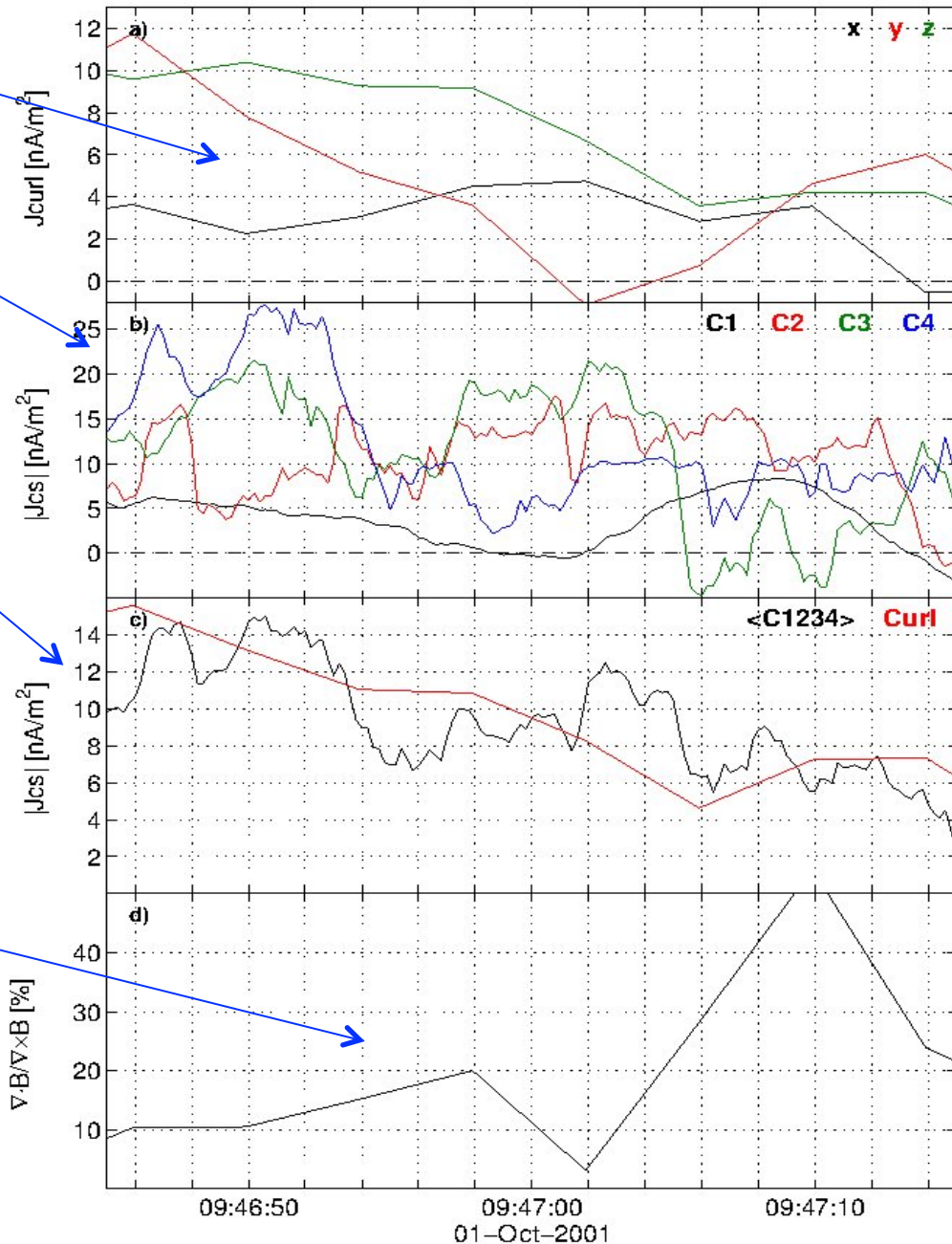
Cluster details: Electric fields

- October 1, 2001: Main Wygant and Runov event in yellow.
- Full resolution \mathbf{E} (25Hz) from EFW [gray]
- CIS \mathbf{E} [blue/cyan] underestimates EFW spin-averaged estimate [red] but follows general trend.
- **Qualitative agreement.** Same sign.



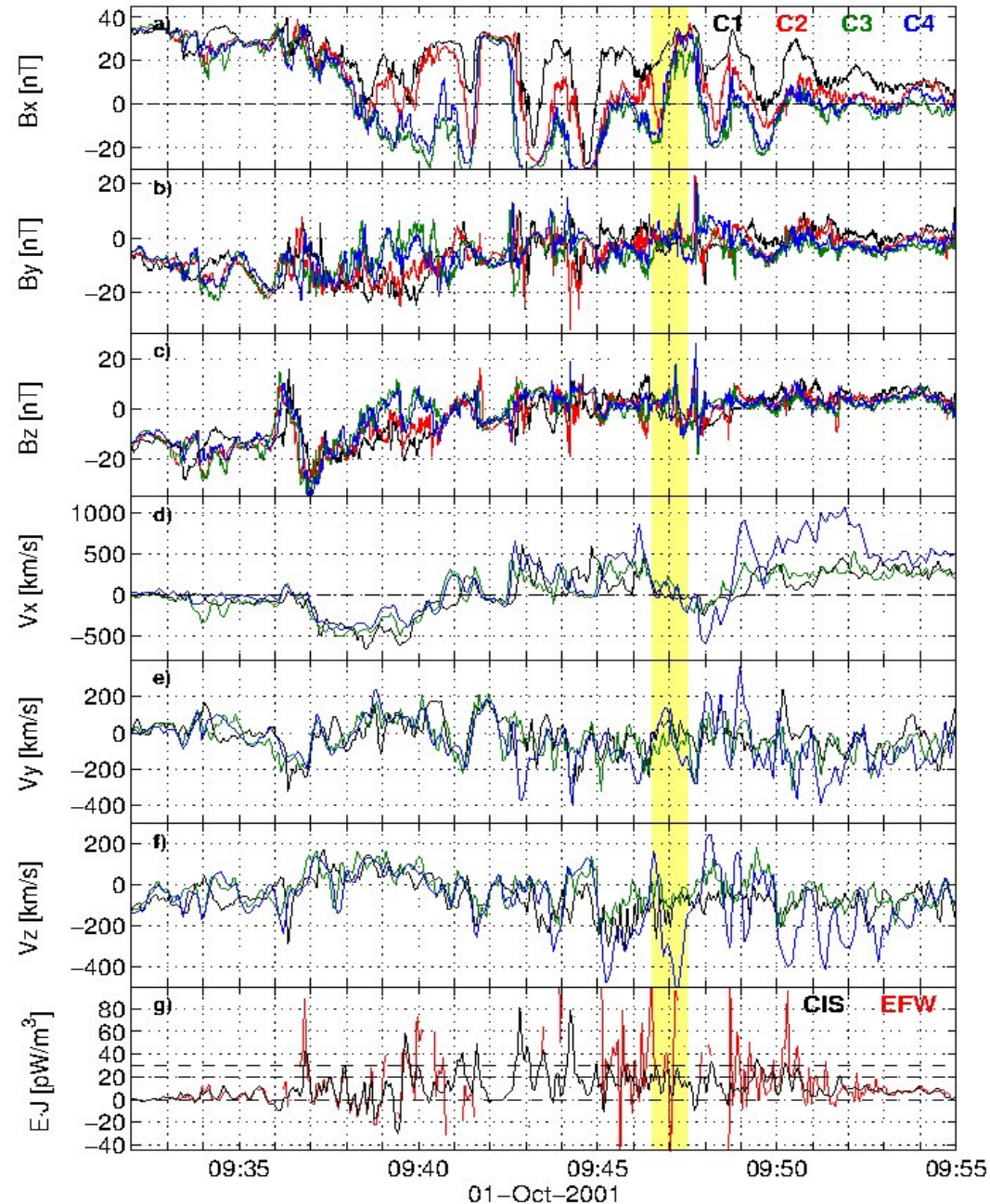
Cluster details: Current densities

- Multi-spacecraft Curlometer $\mathbf{J} \sim \nabla \times \mathbf{B} / \mu_0$.
- Single-spacecraft currents (5vec/s).
- Cross-tail Curlometer current [**red**] and single-spacecraft current averaged over C1234 [**black**].
- Curlometer and single-spacecraft estimate are consistent!
But Curlometer is more smooth.
- Curlometer quality estimate, $\nabla \cdot \mathbf{B} / \nabla \times \mathbf{B}$, is generally small.
- **Curlometer method can be used.**
(cf. Runov 2003).
- **Qualitative agreement.** Same sign.

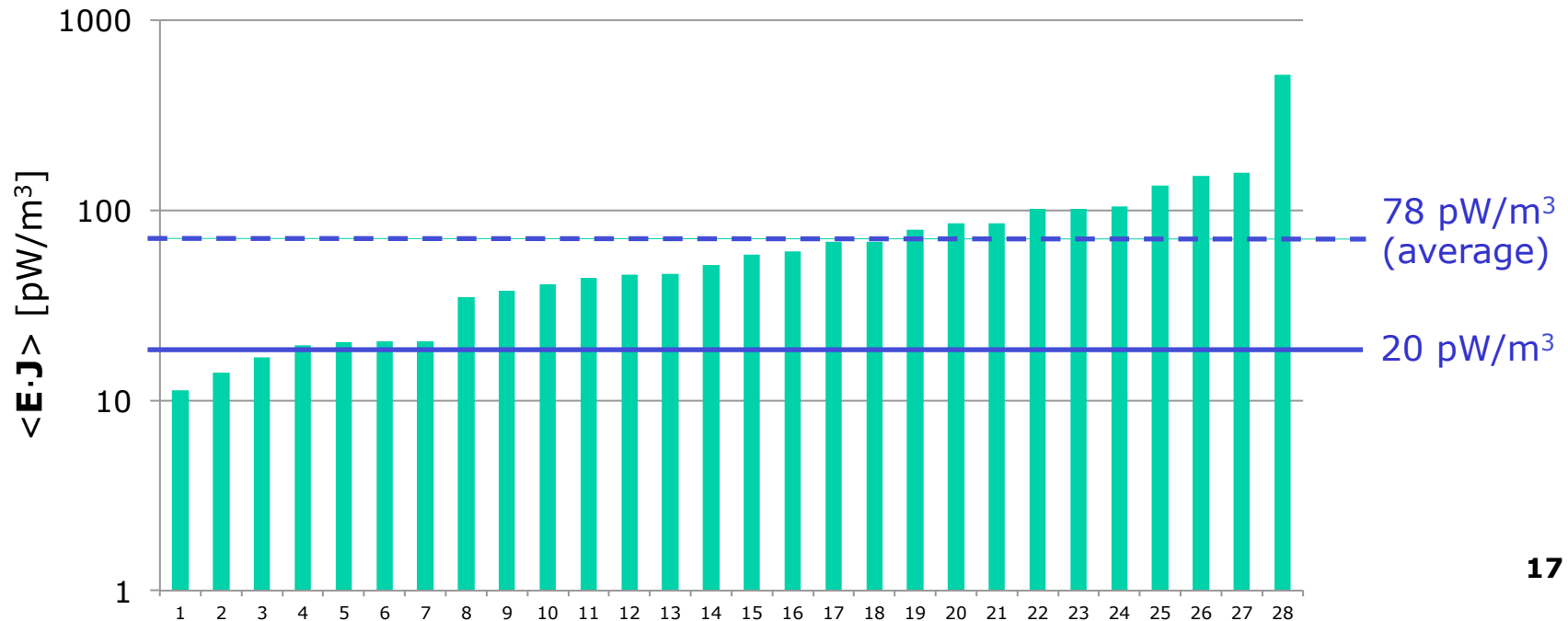


Recap: The reconnection load

- Using the Culometer **J** and CIS $\mathbf{E} = -\mathbf{V} \times \mathbf{B}$ we find that $\mathbf{E}(\text{CIS}) \cdot \mathbf{J}(\text{Curl}) \gg 0$ ($> \sim 20 \text{ pW/m}^3$).
- $|\mathbf{E}(\text{CIS})| < |\mathbf{E}(\text{EFW})|$, same sign
- $|\mathbf{J}(\text{Curl})| < |\mathbf{J}(1\text{-sc})|$, same sign
- “True” $\mathbf{E} \cdot \mathbf{J}$ is possibly even larger than $\sim 20 \text{ pW/m}^3$.



- Cluster observations 2001.
- 28 documented Cluster reconnection events in the literature: Borg et al. [Ann. Geophys. 2012], Eastwood et al. [JGR 2010], Østgaard et al. [JGR 2009].
- $-14 < X_{\text{GSM}} \leq -19R_E$, $|Z_{\text{GSM}}| \leq 3.5R_E$, $|Y_{\text{GSM}}| \leq 10R_E$ (except 2 events slightly outside).
- $\mathbf{E} \cdot \mathbf{J} > 20 \text{ pW/m}^3$ using CIS \mathbf{E} (3 events lightly lower).
- $\mathbf{E} \cdot \mathbf{J}$ generally $\gg 20 \text{ pW/m}^3$ when using EFW \mathbf{E} . (But full EFW \mathbf{E} not always available.)



Additional Cluster events 2001

- 14 additional magnetotail reconnection events (or multiple events) found when searching all 2001 Cluster data.
- Cluster data 2002-2004 can also be scanned, but in >2005 Cluster is in multiscale mode. (But more care needed due to successive instrumental degradations.)

Event	Date	T_1	T_2	X/Y/Z [R_E]	Max $E \cdot J$ [pW/m ³] CIS/EFW	Samples $E \cdot J > 20$ CIS/EFW	Reconnection?	Comment
1	2001-08-03	10:40	11:15	-17/-8.5/2.2	48/48	38/21	Yes	
2	2001-08-05	13:00	14:10	-17/-7.8/5.7	35/27	6/4	Can be	Several sub-events
3	2001-08-07	23:40	00:05	-17/-8.4/2.1	28/38	6/7	Yes	
4	2001-08-12	17:00	18:50	-18/-6.7/3.1	54/47	47/27	Yes	Several sub-events
5	2001-08-17	13:00	15:00	-18/-4.7/2.6	114/71	55/77	Can be	Several sub-events
6	2001-08-31	17:10	17:20	-19/-0.8/2.5	51/79	9/10	Yes	
7	2001-09-02	23:45	00:35	-19/-0.7/3.5	30/26	2/3	Yes	Several sub-events
8	2001-09-07	21:30	22:00	-19/1.2/1.1	21/27	2/4	Yes	
9	2001-09-12	15:10	15:16	-19/2.8/0.5	36/27	3/3	Yes	
10	2001-09-15	03:49	04:35	-19/3.5/-2.0	56/47	44/10	Yes	Several sub-events
11	2001-09-24	00:10	00:20	-14/3.4/ 7.2	30/NaN	5/-	Can be	
12	2001-09-26	22:20	23:10	-18/7.1/-0.4	25/59	2/7	No	
13	2001-10-01	11:40	11:50	-17/8.0/-0.4	26/49	17/24	Can be	
14	2001-10-08	14:40	14:50	-16/9.7/-0.3	28/24	2/2	Can be	

Summary and conclusions

- Regions of strong and localized energy conversion may be used as an indicator for magnetic reconnection:
 $\mathbf{E} \cdot \mathbf{J}$ as a method for identifying potential reconnection events from large amounts of measured spacecraft data.
- Signatures should be verified manually by investigating e.g. jet reversals and Hall fields.
- Criterion for tail reconnection: **$\mathbf{E} \cdot \mathbf{J}$** is large ($> \sim 20 \text{ pW/m}^3$ consistent with theoretical arguments).
- **$\mathbf{E} \cdot \mathbf{J}$** easily computed from multi-spacecraft data, e.g. MMS and Cluster.
 - **\mathbf{J}** : Curlometer method
 - **\mathbf{E}** : ion moments $-\mathbf{V} \times \mathbf{B}$ (e.g. CIS) or direct measurements (e.g. EFW)
- From a statistical investigation using already documented events (Cluster 2001) we find:
 - **$\mathbf{E} \cdot \mathbf{J}$** $> \sim 20 \text{ pW/m}^3$ using CIS ion moments
 - **$\mathbf{E} \cdot \mathbf{J}$** generally $\gg 20 \text{ pW/m}^3$ using EFW measurements.
- Several additional events found when scanning all Cluster plasma sheet data from 2001.



Thank you!

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