Recent Progresses on laser driven magnetic reconnection experiments





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Outline

- 1. Magnetic reconnection Phenomena
- 2. Magnetic reconnection experiments on SGII
- 3. Opportunities for Laser driven reconnection
- 4. Summary

Magnetic reconnection Phenomena In solar flares and space environment





Observations from NASA's Solar Dynamic Observatory provide compelling evidence for the central role of magnetic reconnection in solar flares.



Y. Su et al (2013)

Magnetic reconnection Phenomena - In Lab plasmas

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Magnetically driven systems





MRX at Princeton for reconnection studies

ne= 1-10 x10¹³ cm⁻³, Te~5-15 eV, B~100-500 G

Magnetic reconnection Phenomena In Lab plasmas

Flow driven systems

Omega laser facility, Univ. of Rochester

> 60 arms, 30 kJ, 1/3 μm, 1-10 ns, ~mm scale targets (E/V ~ 10¹⁴ erg/cm³)

Laser on from 0 – 1 ns

0.67 ns

1.42 ns

C. K. Li et al. (2007)

0.04 ns



2. Magnetic reconnection experiments on SGI



• What is SGII



SG II Lasers Beams: 8 Pulse: 1 ns Energy: 260 J/beam for 3ω Focus spot: 50-100 μm Intensity: 1-5×10¹⁵ W/cm²

ne= 1-10 x10¹⁹ cm⁻³, Te~1keV, B~10⁶ G

• What is SGII





Diagnostics: X-rays: pinhole, framing,... Optical devices: shadowgraphy, Interferometry, Faraday, ... Particles: EMS, RCF...

- Like other Advanced HEDP Facilities: NIF, OMEGA, Z,... ...
- SG II is used to study fusion, laboratory astrophysics,
- In China, 20% shots of Shenguang II for each year are given for the study of Laboratory astrophysics

High energy density laboratory astrophysics (HEDLA)similar with IPELS

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Web site http://hedla2014.sciencesconf.org







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HEDLA 2014

10th International Conference on High Energy Density Laboratory Astrophysics 12-16 May 2014 Bordeaux (France)

MEETING TOPICS

Plasma physics Stellar explosions Magnetized HED laboratory astrophysics Astrophysical disks, jets and outflows Stellar, solar and nuclear astrophysics **Computations in HED physics Radiative hydrodynamics** Warm dense matter

Roberto Mancini

Tomasz Plewa

Marc Pound

Thomas Ray

Steven Rose

James Stone

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Scientific Organisation Committee

John Bally Yutong Li James Bailey Tony Bell Mikhail Medvedev Paul Bellan Claire Michaut Nancy Brickhouse Peter Norreys Serge Bouquet (co-chair) Norimasa Ozaki Andrea Ciardi Werner Dappen Melissa Douglas Paul Drake Adam Frank Patrick Hartigan Dongsu Rvu John Hillier Youichi Sakawa Luis Silva Hantao Ji Michel Koenig Hideaki Takabe Yasuhiro Kuramitsu Feilu Wang Carolyn Kuranz Michael Wiescher George Kyrala

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Magnetic field in ns laser plasmas



Laser driven magnetic reconnection opens a new territory of study for magnetic reconnection in a parameter regime not covered before



Parameters	Laser-produced plasmas	Coil plasmas
Length (cm)	10^{-1}	10^{-1} - 1
Time (s)	10 ⁻⁹ ,	10^{-9} - 10 ⁻⁸
Pressure (Pa)	10^{7} .	
Density (cm ⁻³)	$\sim 10^{19} - 10^{20}$	$< 10^{18}$
Velocity (km s ⁻¹)	~ 100 €	10,0-1,000,
Magnetic field (G)	$\sim 10^{6}$	10 ^{,6-7} ,

 $\beta = 4.03 \times 10^{-11} nTB^{-2}$ 0.001-100

B is frozen in Plasmas



B is separated with Plasmas



The coil will be ionized by the radiation from the laser produced plasmas, while the density is lower.



Experiments on Laser-Flow driven MR



$\rho \frac{du}{dt} = -\nabla (\frac{B^2}{2} + p) + B \cdot \nabla B$ Magnetic reconnection Magnetic repulsion Hydro Anti-Parallel B fields Parallel B fields $\nabla(\frac{B^2}{2} + p) + B \cdot \nabla B$ $\rho \frac{du}{dt} =$ du

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Experiments on Laser-Flow driven MR

1. Reconnection effects in hydro-interaction

The shadowgraph together with X-ray images show the topology of magnetic reconnection with laser driven flow plasmas, confirmed by MHD simulation.



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Anti-Parallel B fields



Optical shadows



Experiments on Laser-Flow driven MR

2. Reconnection effects for guide fields

The toroidal magnetic field on the target will produce a magnetic field component at the reconnection plane which is similar provide an applied guide field with the level of the reconnection field.

Guide field do affect the reconnection both for the current sheet and outflows.

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Without guide field

Optical shadows

X-ray image

MHD +test particles simulation

With guide field

Optical shadows

X-ray image

MHD +test particles simulation

The observed energy spectra of energetic electrons shows the similar distribution with the one observed in space plasmas.

The spectra are obviously of the so-called Kappa distribution that is quasi-Maxwellian at low and thermal energies, while its nonthermal tail decreases as a power law at high energies, as generally observed in space plasmas.

Experiments on Laser-magnetically driven MR

Side view

Front view

Shadowgraphy shows the shape of plasmas exhibiting with a X-type reconnection. A magnetostatics computer code Radia is applied to construct magnetic topology of the present coil target.

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Magnetic fields from "long pulse" laser to "short pulse" laser

ns lasers

C.K. Li et al 2007

U. Wagner et al 2004

10⁶ Gauss to 10⁹ Gauss

Opportunities for Laser driven reconnection

After

Before

Three-dimensional fast magnetic reconnection driven by *two ultraintense femtosecond laser pulses* is investigated by relativistic particle-in-cell simulation 1902

Omega EP

Key challenges:

- co-time of multi beams
- effect from beam itself
- diagnostic ...

Summary

- Magnetic reconnection is a basic dissipation process in laboratory, space and astrophysical plasmas.
- Some basic reconnection physics is being studied in *flow driven systems, such as guide field, electron acc.*
- New opportunities for HED experiments to study reconnection.
- Relativistic intensity short pulse lasers are hoping to perform the reconnection experiments with GG-B field.
- With special target, such as capacity target, realizing
 magnetically driven systems and *flow driven systems*, enable
 to extend a large rang of parameter space for reconnection.