

# Fermions, Gravity & Chiral Symmetry

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& Astrid Eichhorn New J.Phys. 13 (2011) 125012,  
& Stefan Lippoldt Phys.Rev. D87 (2013) 104026,  
Phys.Rev. D89 (2014) 6, 064040, Phys.Lett. B (2015) [arXiv:1502.00918]

Probing the Fundamental Nature of Spacetime with the Renormalization Group, Nordita, March 23-27 2015

☉ Gravity

☉ Fermions:

☉ Chiral Symmetry:

☉ Gravity

... see workshop title

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... most of the luminous matter in the universe

☉ Chiral Symmetry:

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... see workshop title

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... most of the luminous matter in the universe

## ☉ Chiral Symmetry:

☞ symmetry of all known fermions

☞ protects mass  $\implies$  light fermions

☞ non-chiral fermions expected to be heavy

☞ chiral symmetry breaking ( $\chi$ SB): by strong & EW force

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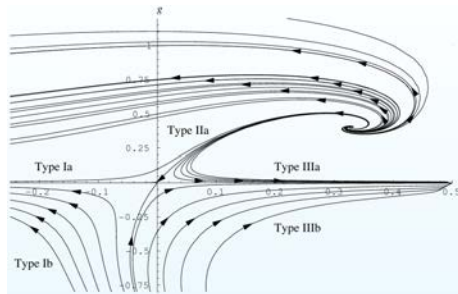
☞ chiral symmetry breaking ( $\chi$ SB): by strong & EW force

... what about gravity?

# Asymptotically safe metric quantum gravity

☉ e.g., flow in Einstein-Hilbert theory space:

$$\Gamma_k = \frac{1}{16\pi G_k} \int d^D x \sqrt{g} (-R + 2\bar{\Lambda}_k)$$



(REUTER'96)

(DOU,PERCACCI'97)

(SOUMA'99)

(LAUSCHER,REUTER'01'02)

(REUTER,SAUERESSIG'01)

(NIEDERMAIER'02)

(LITIM'03; PAWLOWSKI'03)

(CODELLO,PERCACCI'06)

(CODELLO,PERCACCI,RAHMEDE'07'08)

(MACHADO,SAUERESSIG'07)

(BENEDETTI,MACHADO,SAUERESSIG'09)

(EICHHORN,GIES,SCHERER'09)

(BENEDETTI,GROH,MACHADO,SAUERESSIG'10)

(DONKIN,PAWLOWSKI'12)

(.....)

⇒ so far remarkably **robust**

# Criteria for Asymptotic Safety

(WEINBERG'76'79)

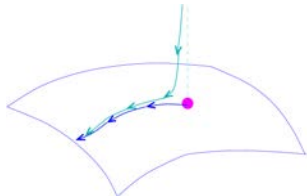
- Existence of a (non-Gaußian) fixed point  $g_*$

$$\beta(g_*) = 0$$

- # relevant directions  $\dim S < \infty$

$$\{\Theta^l\} = \text{spectr} \left( - \left. \frac{\partial \beta_i}{\partial g_j} \right|_{g_*} \right)$$

$$\dim S = \#(\Theta^l > 0)$$



⇒ Predictivity

# Asymptotically safe quantum gravity in nature

- ⊙ testable physical predictions

... nice to have ...

(TALKS: A. BONANNO, A CODELLO)

- ⊙ consistency with existing observations

... “matter matters” ...



# Consistency with observations

☉ Gravity & gauge sectors:

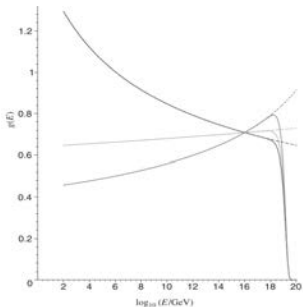
☞ nonabelian asymptotic freedom preserved (or enhanced)

(WILCZEK,ROBINSON'06)

(PIETRYKOWSKI'07;TOMS'07'10;EBERT,PLEFKA,RODIGAST'08)

(TANG,WU'08;DAUM,HARST,REUTER'10)

(FOLKERTS,LITIM,PAWLOWSKI'12)

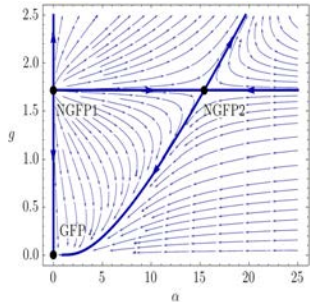


☞ abelian sector:

(HARST,REUTER'11)

similarly UV controlled (NGFP1)

or even higher predictivity (NGFP2)



# Consistency with observations

☉ Gravity & matter sectors:

Properties/Existence of gravity fixed point depend on matter content

(PERCACCI,PERINI'03)

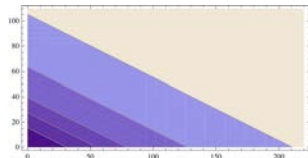
(CODELLO,PERCACCI,RAHMEDE'09;VACCA,ZANUSSO'10; DONÀ,PERCACCI'13)

one-loop:

(DONÀ,EICHHORN,PERCACCI'13)

$$g_* = \frac{12\pi}{46 + 4N_V - 2N_D - N_S}$$

$$\lambda_* = \frac{3}{4} \frac{2 + 2N_V - 4N_D + N_S}{31 + 4N_V - 2N_D - N_S}$$



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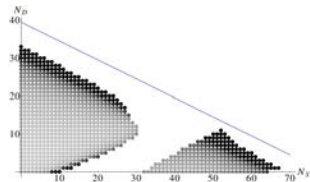
(CODELLO,PERCACCI,RAHMEDE'09;VACCA,ZANUSSO'10; DONÀ,PERCACCI'13)

(DONÀ,EICHHORN,PERCACCI'13)

refined analysis:

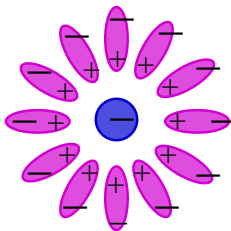
using bi-metric input

(CODELLO,D'ODORICO,PAGANI'13)



## Similarity to many-flavor QCD

⊙ charge screening:

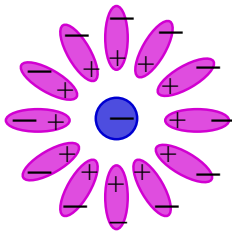


⊙  $\beta$  function

$$\beta = -2 \left( \frac{11}{3} N_c - \frac{2}{3} N_D \right) \frac{g^4}{16\pi^2} - 2 \left( \frac{34N_c^3 + 3N_D - 13N_c^2 N_D}{3N_c} \right) \frac{g^6}{(16\pi^2)^2} + \dots$$

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$$\text{for } N_D > \frac{34N_c^3}{13N_c^2 - 3} \stackrel{\text{SU}(3)}{\simeq} 8.05$$

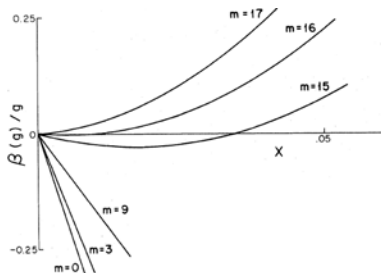
# Many-flavor QCD

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⊙ e.g., SU(3): IR fixed point  $\alpha_*$

(CASWELL'74; BANKS&ZAKS'82)



[CASWELL@PHYS.REV.LETT.33:244,1974]

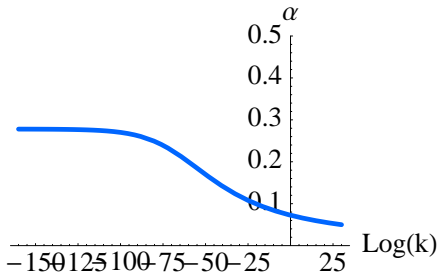
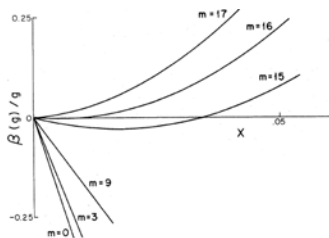
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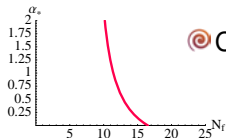
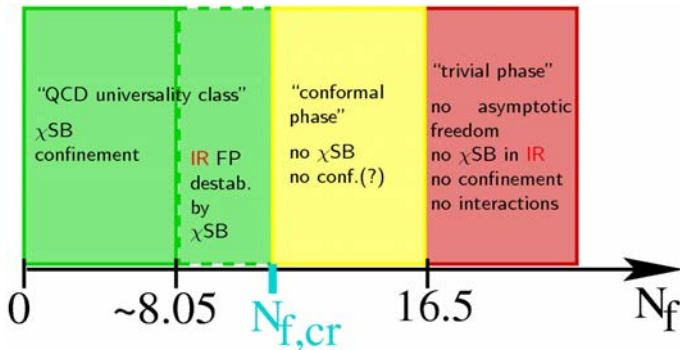
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⊙ e.g.  $N_D = 14$

⇒ IR fixed point



# Many-flavor QCD



⊙ Caswell-Banks-Zaks fixed point destabilized by  $\chi_{SB}$

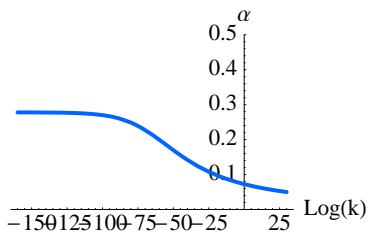
⇒ conformal phase  $\neq$  real life QCD



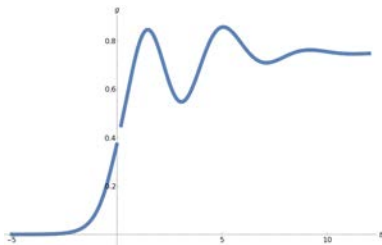
# Many flavor QCD vs. Gravity

☉ UV to IR flow:

many-flavor QCD



Gravity



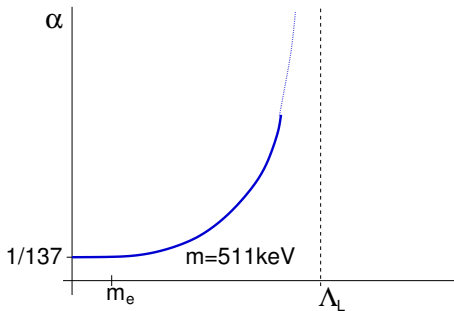
⇒ check status of chiral symmetry in UV regime!

# Comparison to QED

⊙ coupling  $\alpha$  grows towards UV

(GELL-MANN,LOW'54)

... Landau-Pole (LANDAU'55)

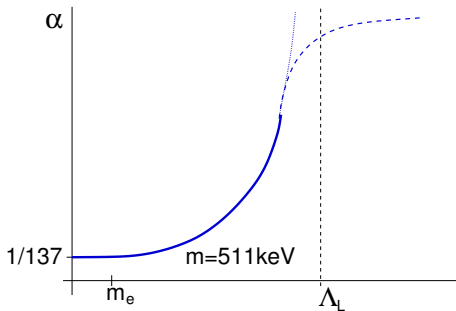


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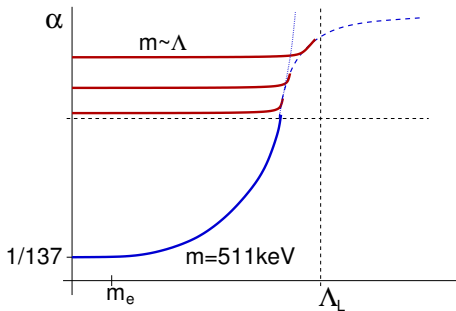
⊙ ... even if UV was controlled ...

# Comparison to QED

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⊙  $\chi$ SB leads to heavy fermions  $m \sim \Lambda$  if  $\alpha > \alpha_{cr}$

(MIRANSKY'85)

⇒ Strong QED and “real” QED not on a line of constant physics

(GOCKELER ET AL.'98; HG,JAECKEL'04)

# Chiral Symmetry Breaking ( $\chi$ SB) in fermionic systems

☉ e.g., Gross-Neveu model in 3 dimensions:

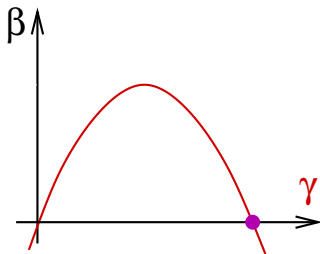
$$\Gamma_k = \int d^3x \bar{\psi} i \not{\partial} \psi + \frac{1}{2} \bar{\gamma} (\bar{\psi} \psi)^2 + \dots \quad , \quad [\bar{\gamma}] = -1$$

$\mathcal{M}$  dim'less coupling  $\gamma = k \bar{\gamma}$

$$\partial_t \gamma = \gamma - c \gamma^2$$

$\mathcal{M}$  UV fixed point  $\gamma_* = 1/c$

$\mathcal{M}$  critical exponent  $\Theta = 1$



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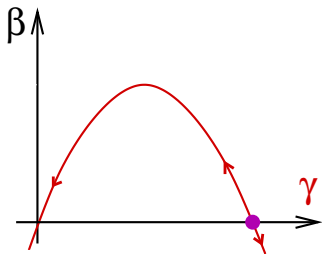
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$\implies$  asymptotically safe

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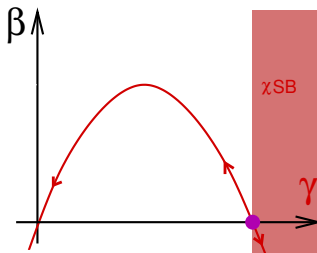
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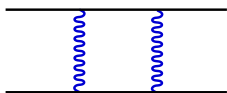
$\implies$  fixed point = **critical point**  $\gamma_{cr}$  of  $\chi$ SB quantum phase transition

# Interaction-induced $\chi$ SB

⊙ e.g., gauge interactions:

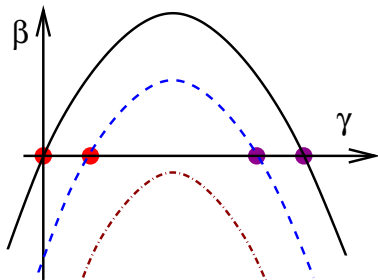
$$S_{\text{int}} \sim \int d^d x \dots \bar{g} \bar{\psi} \mathbf{A} \psi$$

🌀 gauge-boson exchange:



🌀 coupling flow:

$$\partial_t \gamma = \gamma - c \gamma^2 - c_G g^4$$



⇒ **fixed-point collision:** dynamical  $\chi$ SB for  $g^2 > g_{\text{cr}}^2$



# Gravity-induced $\chi$ SB?

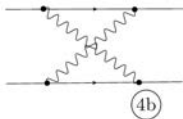
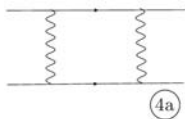
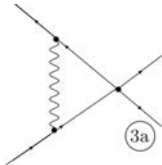
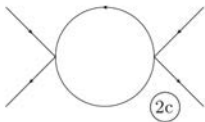
(EICHHORN, HG'11)

⊙ chiral fermions and quantum gravity,  $SU(N_D)_R \times SU(N_D)_L$  in  $D = 4$ :

$$\Gamma_{\text{int},k} = \int d^4x \sqrt{g} \frac{1}{2} \int d^4x \sqrt{g} [\bar{\lambda}_-(V - A) + \bar{\lambda}_+(V + A)]$$

$$V = (\bar{\psi}^i \gamma_\mu \psi^i) (\bar{\psi}^j \gamma^\mu \psi^j), \quad A = -(\bar{\psi}^i \gamma_\mu \gamma^5 \psi^i) (\bar{\psi}^j \gamma^\mu \gamma^5 \psi^j)$$

⊙ Diagrammar:



cancellations!

# Gravity-induced $\chi$ SB?

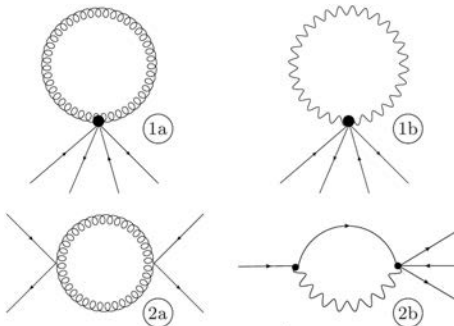
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⊙ Diagrammar:



specific  
to  
gravity!

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⊙ RG flow, e.g.  $\lambda_+$ :

$$\partial_t \lambda_+ = 2\lambda_+ + c\lambda_+^2 + b_{G\lambda_+} \frac{g\lambda_+}{(1-2\lambda)^2} + c_G \frac{g^2}{(1-2\lambda)^3}$$

⌘ attractive gravitational binding  $\sim c_G > 0$

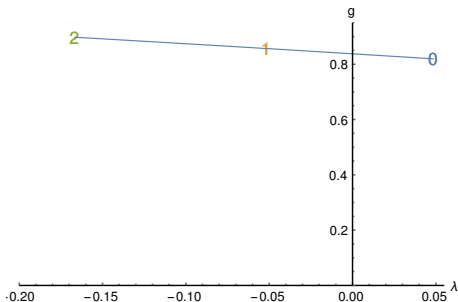
⌘ effective repulsive gravity-fermion interplay  $\sim b_{G\lambda_+} > 0$

⌘ gravitational decoupling for  $\lambda < 0$

# Gravity-induced $\chi$ SB?

⊙  $N_D$  dependence of gravitational fixed point

(DONÀ, EICHHORN, PERCACCI '13)



*m* e.g., one-loop

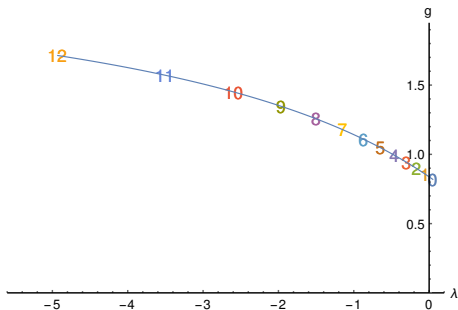
⇒ attraction, repulsion, decoupling @ work

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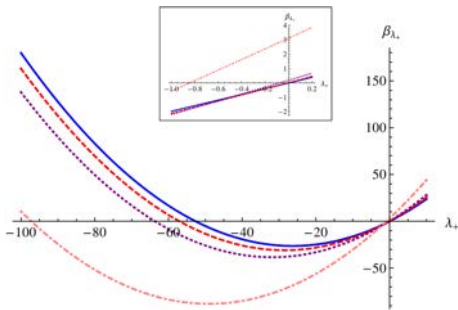


⇒ attraction, repulsion, decoupling @ work

# Gravity-induced $\chi$ SB?

⊙ no fixed-point collision from gravity fluctuations

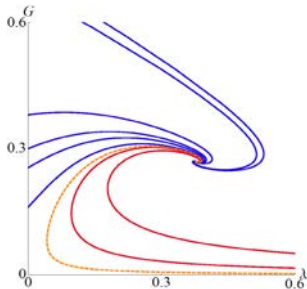
(EICHORN, HG'13)



- ⌘  $D = 4$ : repulsive gravity-fermion interplay dominates
- ⌘ existence of light fermions compatible with asymptotic safety
- ⌘ non-Gaussianity:  $\implies$  shifted Gaussian fixed point
- ⌘ so far no constraint on matter

# Asymptotic Safety as a quantum phase transition

⊙ dS/AdS quantum phase transition



⊙ spacetime structure in the fixed-point regime:

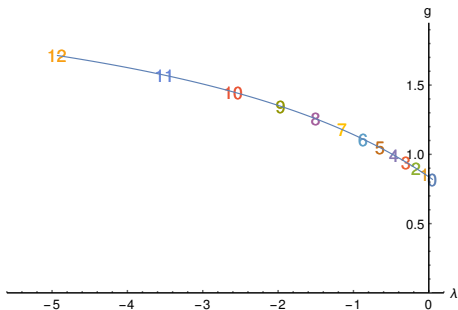
(LAUSCHER,REUTER'05)

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = -\bar{\Lambda}g_{\mu\nu} \quad \rightarrow \quad R = \frac{2d}{d-2}\bar{\Lambda}$$

..... Einstein-Hilbert ...)

## Spacetime in fixed point regime

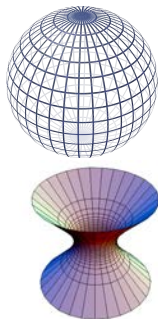
⊙ spacetime curvature in fixed-point regime: matter dependent:



⇒ Chiral symmetry needs to be studied on curved background.



# Gravitational Catalysis



⊙ positive curvature screens IR fluctuations

⊙ negative curvature enhances IR fluctuations

⊙ effective dimensional reduction of Dirac spectrum on  $H^d$

(GORBAR'08)

$$d + 1 \rightarrow 1 + 1$$

⇒ negative curvature supports  $\chi$ SB

(BUCHBINDER,KIRILLOVA'89)

(INAGAKI,MUTO,ODINTSOV'93'97; WACHS,WIPF'93; ELIZALDE ET AL.'96; MIELE,VITALE'96'98; GORBAR,GUSYNIN'07)

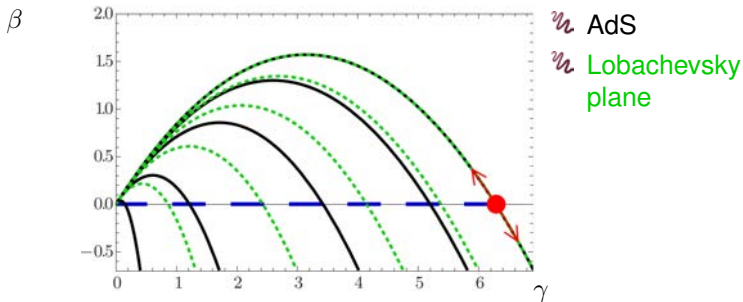
opposite effect for bosons(BENEDETTI'14)

# Gravitational Catalysis á la RG

⊙ e.g., Gross-Neveu model in  $D = 3$ :

(HG,LIPPOLDT'13)

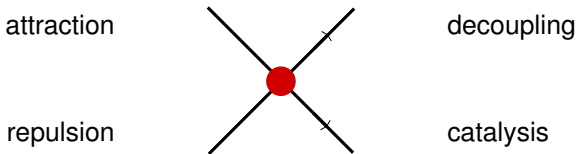
$$\partial_t \gamma = \beta_\gamma = \gamma - N_D \frac{\gamma^2}{\pi} \left( 1 + \frac{|R|}{24k^2} \right), \quad R < 0$$



$\mathcal{M}$   $\chi$ SB even for the **weakest attraction** (for fixed  $R < 0$ )

$\mathcal{M}$  fermion mass, e.g.,  $m_f \sim k_{UV}$  (weak coupling, large fixed  $R < 0$ )

## Gravity-induced $\chi$ SB, Part II?



© Simple model for gravity-induced  $\chi$ SB?

$\mathcal{M}$   $D = 4$ : flavor singlet channels:  $(V \pm A)$

catalysis in vector-channels not yet studied

$\mathcal{M}$   $D = 3$ : gravity preserves  $U(2N_D)$  symmetry

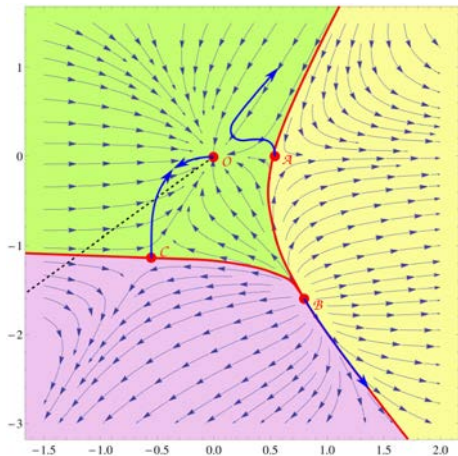
singlet channels:  $(\bar{\psi}\gamma_{45}\psi)^2$ ,  $(\bar{\psi}\gamma_{\mu}\psi)^2$

no gravity-induced attraction in singlet channel

## Gravity-induced $\chi$ SB, Part II?

©  $D = 3$ : fixed-point collision in mixed channel expected

$\implies$  symmetry breaking pattern:  $U(2N_D) \rightarrow U(N_D) \times U(N_D)$



e.g., QED<sub>3</sub>

(BRAUN, HG, JANSEN, ROSCHER '14)

... as difficult as  $D = 4$

## Gravity-induced $\chi$ SB, Part II

⊙ a generic one-channel model ( $D = 3$ ):

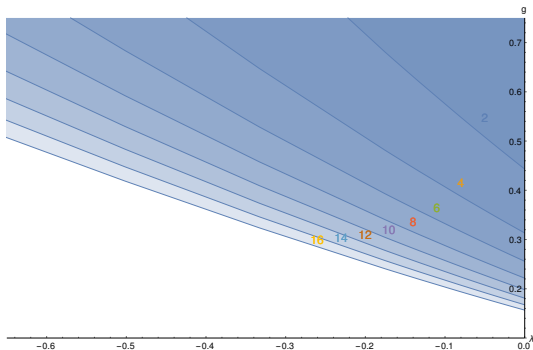
$$\partial_t \gamma = \beta_\gamma = \gamma - N_D \frac{\gamma^2}{\pi} \left( 1 + \frac{|R|}{24k^2} \right) - c_G \frac{g^2}{(1 - 2\lambda)^3}, \quad R < 0$$

## Gravity-induced $\chi$ SB, Part II

⊙ a generic one-channel model ( $D = 3$ ) near the fixed point:

$$\partial_t \gamma = \beta_\gamma = \gamma - N_D \frac{\gamma^2}{\pi} (1 - c_{\text{Cat}} \lambda) - c_G \frac{g^2}{(1 - 2\lambda)^3}, \quad \lambda < 0$$

⊙ fixed-point collision  $\implies$   $\chi$ SB for  $N_D$  flavors, e.g.  $c_{\text{Cat}} = \frac{1}{4}$ ,  $c_G = 2$

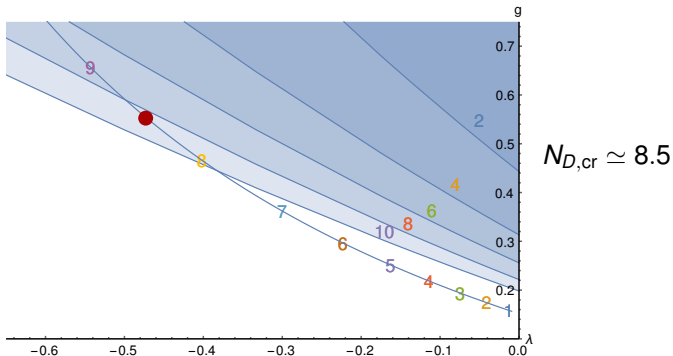


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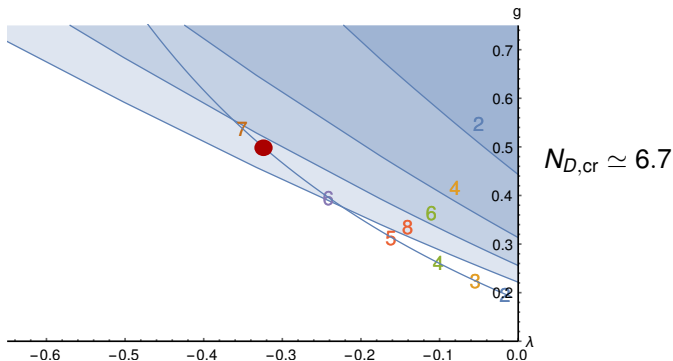


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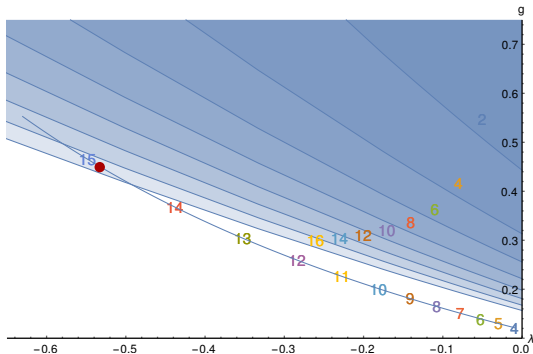


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⊙ gravity fixed point for  $N_S = 4$ ,  $N_V = 4$ ,



$$N_{D,\text{cr}} \simeq 15$$

# Gravity-induced $\chi$ SB

⊙ details matter

🌀 Fierz completeness

🌀 finite  $\mathcal{O}(1)$  corrections vs. large- $N_D$  expansion

🌀 residual interactions of scalar, gauge, etc. sectors

(EICHHORN'12)

shifted Gaußian fixed point

🌀 anomalous dimensions

🌀 bimetric treatment

... weakens link between decoupling and catalysis

🌀 depends on quantum degrees of freedom

... e.g., catalysis irrelevant for unimodular gravity

(EICHHORN'13'15)

# Gravity, Fermions & Chiral Symmetry

⊙ mechanisms of gravity-induced  $\chi$ SB

... (not only) put a constraint on  $N_D$

⊙ gravity-induced  $\chi$ SB + asymptotically safe gravity

... yields a dynamical theory of  $N_D$

⌘ Given a UV theory with  $N_{D,UV} > N_{D,cr}$

⌘  $\chi$ SB  $\implies N_{D,light} < N_{D,cr}$

... depending on breaking patten

CAVE: Goldstone bosons!

⌘ Observation:  $N_{D,SM} \leq N_{D,light}$

# Is there a theory of the discrete numbers of nature?

$N_D$



gravity + fermions +  $\chi$ SB

$N_V$



YM theories + Higgs mechanism

# Is there a theory of the discrete numbers of nature?

$N_D$



gravity + fermions +  $\chi$ SB

$N_V$



YM theories + Higgs mechanism

$N_S$



# Is there a theory of the discrete numbers of nature?

$N_D$



gravity + fermions +  $\chi$ SB

$N_V$



YM theories + Higgs mechanism

$N_S$



$D$



Why  $D = 4$ ?



$$N_O = 3$$



$$N_O = 3$$



$$N_O = 3$$



Thanks for all your efforts!!!\*

\* uncontroversial statement



# Hen or Egg?

⊙ quantum fields  $\hat{=}$  representations of Lorentz group

🌀 fundamental (nontrivial) building block: spin  $\frac{1}{2}$

🌀 ... perhaps all fields are made of fermions

🌀 ... perhaps also the graviton

⇒ fermions first!

(HEBECKER, WETTERICH'03; DIAKONOV'11)

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$$\{\gamma_\mu, \gamma_\nu\} = 2g_{\mu\nu}$$

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... or Dirac matrices first?

# Fluctuating fermions and metric gravity?

⊙ fermions:  $\mathcal{D}\bar{\psi}\mathcal{D}\psi$ ; gravity:  $\mathcal{D}g, \mathcal{D}e, \mathcal{D}\Gamma, \dots?$

(CAPPOVILLA, JACOBSON, DELL'91; DAUM, REUTER'10; HARST, REUTER'12; DONÀ, PERCACCI'12)

⊙ guiding principle: simplicity

(LIPPOLDT, HG'14'15)

⌘ symmetries preserving the Clifford algebra:  $\{\gamma_\mu, \gamma_\nu\} = 2g_{\mu\nu}$

⌘ Diff's:  $g'_{\mu\nu} = \frac{\partial x^\rho}{\partial x'^\mu} \frac{\partial x^\lambda}{\partial x'^\nu} g_{\rho\lambda}$

⌘ spin-base transformation:  $\mathcal{S} \in \text{SL}(4, \mathbb{C})$

(SCHRÖDINGER'32; BARGMANN'32)

$$\psi \rightarrow \mathcal{S}\psi, \quad \bar{\psi} \rightarrow \bar{\psi}\mathcal{S}^{-1}$$

⌘ most general Dirac matrix transformation

(PAULI'36)

$$\gamma_\mu \rightarrow \gamma'_\mu = \frac{\partial x^\rho}{\partial x'^\mu} \mathcal{S}\gamma_\rho\mathcal{S}^{-1}$$

⊙ spin-base group factors out of functional integral

(HG, LIPPOLDT'14)

⇒ metric-based quantum gravity with fermions (“as usual”)