



Concluding Talk

The Spacetime Odyssey Continues

celebrating the accomplishments and identifying the challenges



2 to 5 June 2015

Stockholm, Sweden

Michael S. Turner

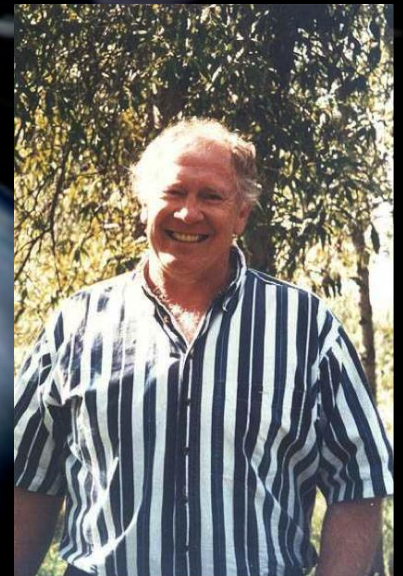
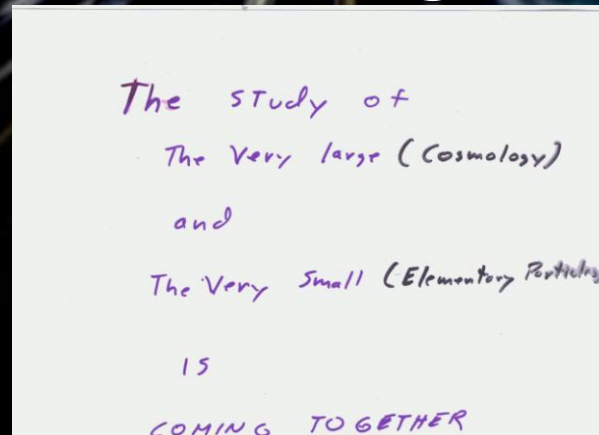
Kavli Institute for Cosmological Physics

The University of Chicago

Concluding talks

- Definition
 - |Summary> = coherent sum of what was said + what I heard + what should have been said + my opinions
 - What you hear: $|\text{Summary}|^2 \approx |\text{what should have been said}|^2 + |\text{my opinions}|^2$
- Dangerous job
 - Thorlacius/Silk/Aprile/Turner \rightarrow Silk/Aprile/Turner \rightarrow Aprile/Wilczek/Silk/Turner \rightarrow Aprile/Wilczek/Turner \rightarrow Turner (survival of the fittest?)
 - ... right up there with being #2 at Al-Qaida (7 killed since 2006!)

Big ideas (quarks and the cosmos)
and powerful instruments
(accelerators & telescopes) have
pushed both fields forward
shoulder to shoulder
Golden Age??



Circa 1984 (when Katie was finishing her thesis)



- Observational cosmology summary: $H_0 = 50 \pm 1$ or 100 ± 1 ; $\Omega_0 = \Omega_M \sim 0.1$; CMB: dipole + $\delta T/T < 10^{-4}$
- HST 6 years from launch
- Birth of modern inflationary paradigm
- HDM falling and CDM rising with axion as first CDM candidate
- SUSY DM coming soon



Celebrating the
accomplishments
we've come a long way

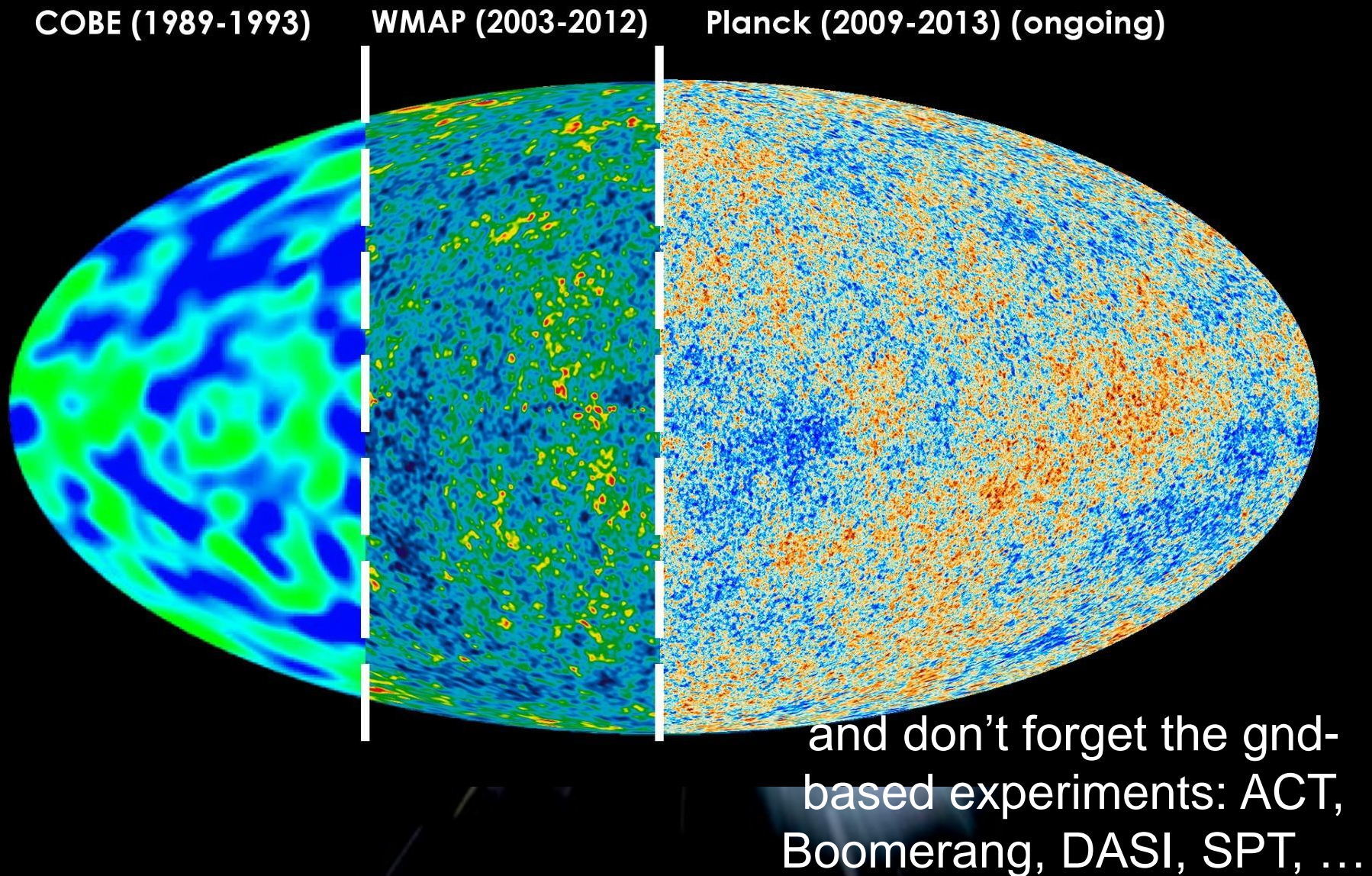
Consensus Cosmology

precision, accuracy, full accounting and consistency

- From quark soup to nuclei and atoms to galaxies and large-scale structure
- Flat, accelerating Universe
- Atoms, exotic dark matter & dark energy
- Consistent with inflation
- Precision cosmological parameters
 - $\Omega_0 = 1.005 \pm 0.005$ (uncurved = flat)
 - $\Omega_M = 0.315 \pm 0.01$
 - $\Omega_B = 0.048 \pm 0.001$
 - $\Omega_{DE} = 0.685 \pm 0.01$
 - $H_0 = 67 \pm 0.5$ km/s/Mpc
 - $t_0 = 13.80 \pm 0.02$ Gyr
 - $n_s = 0.965 \pm 0.005$
 - $N_v = 3.0 \pm 0.33$

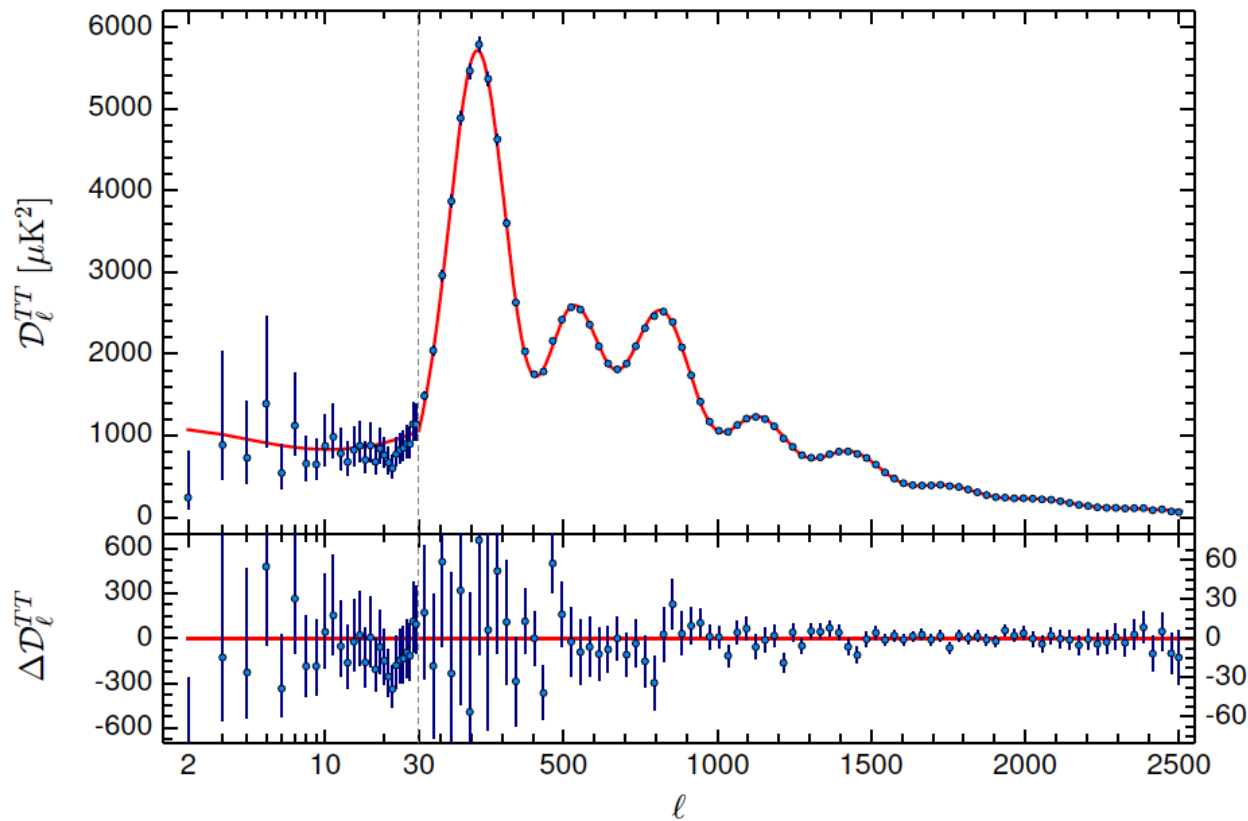
Consistent with all
data, laboratory
and cosmological!

The Universe circa 380,000 yrs

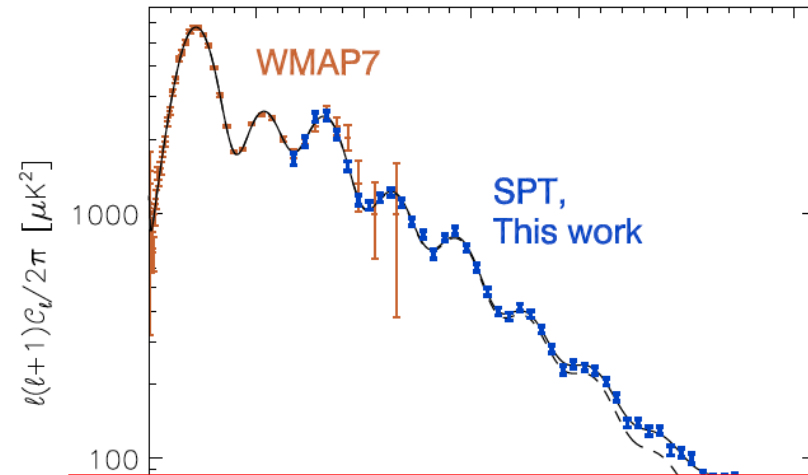
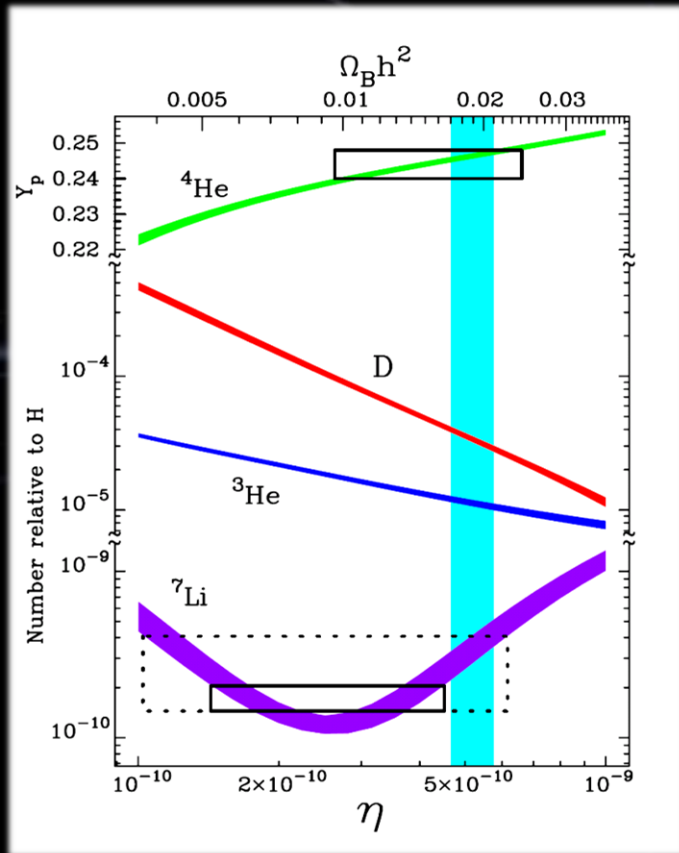


Planck Power Spectrum

a thing of beauty!

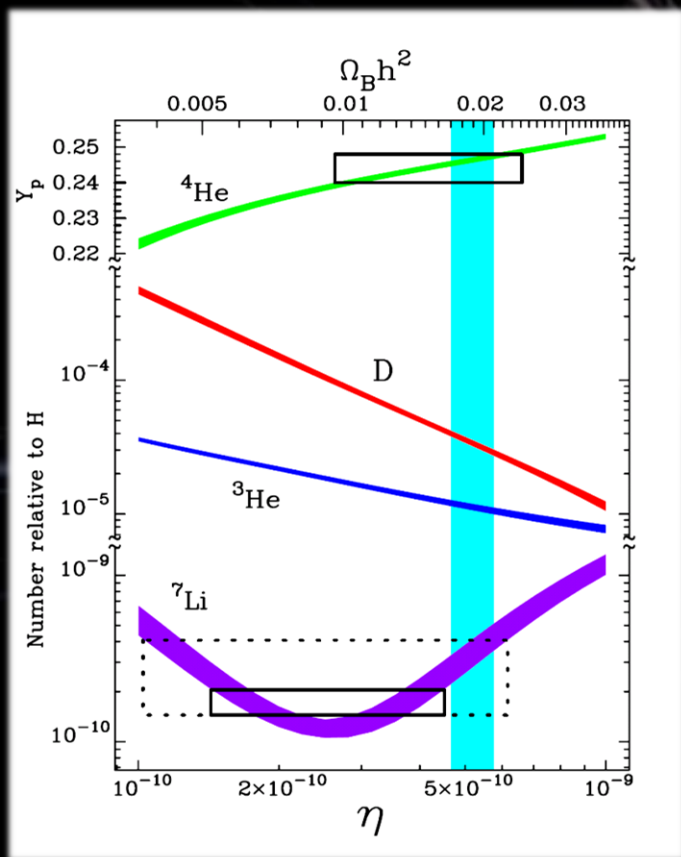


Airtight Evidence for Nonbaryonic Dark Matter

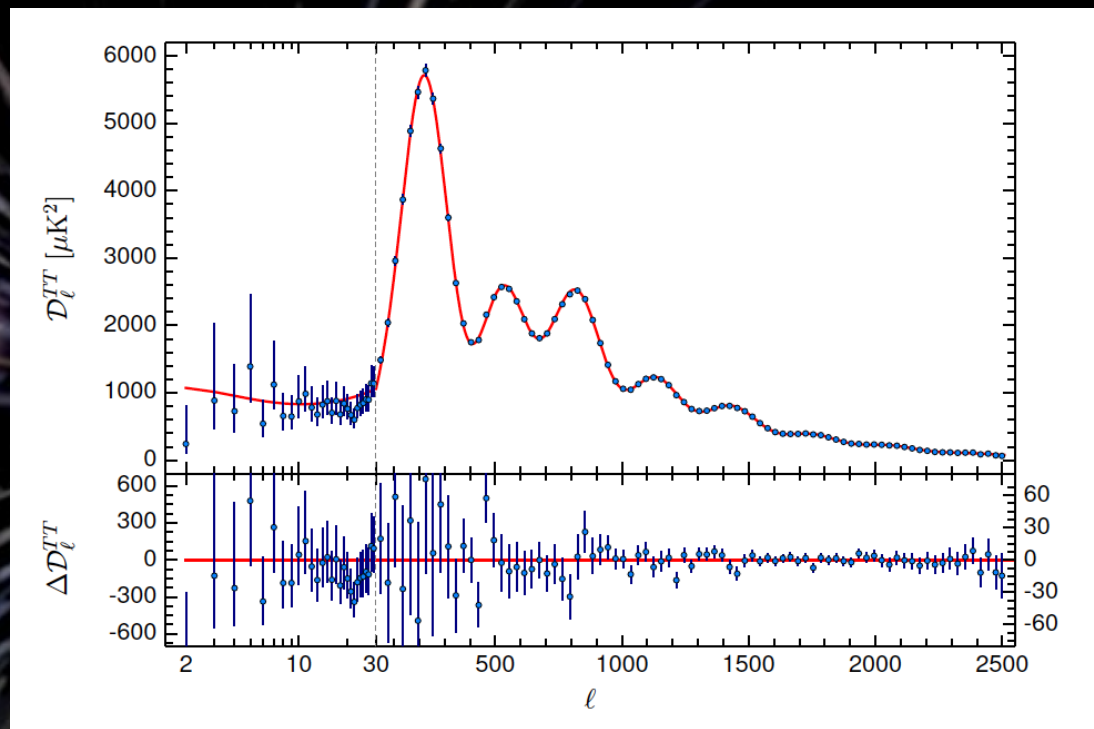


CMB & BBN
 $\Omega_b h^2 = 0.0222 \pm 0.0002$
vs.
CMB/SDSS
 $\Omega_M h^2 = 0.142 \pm 0.0013$
40+ σ discrepancy

Cosmic consistency



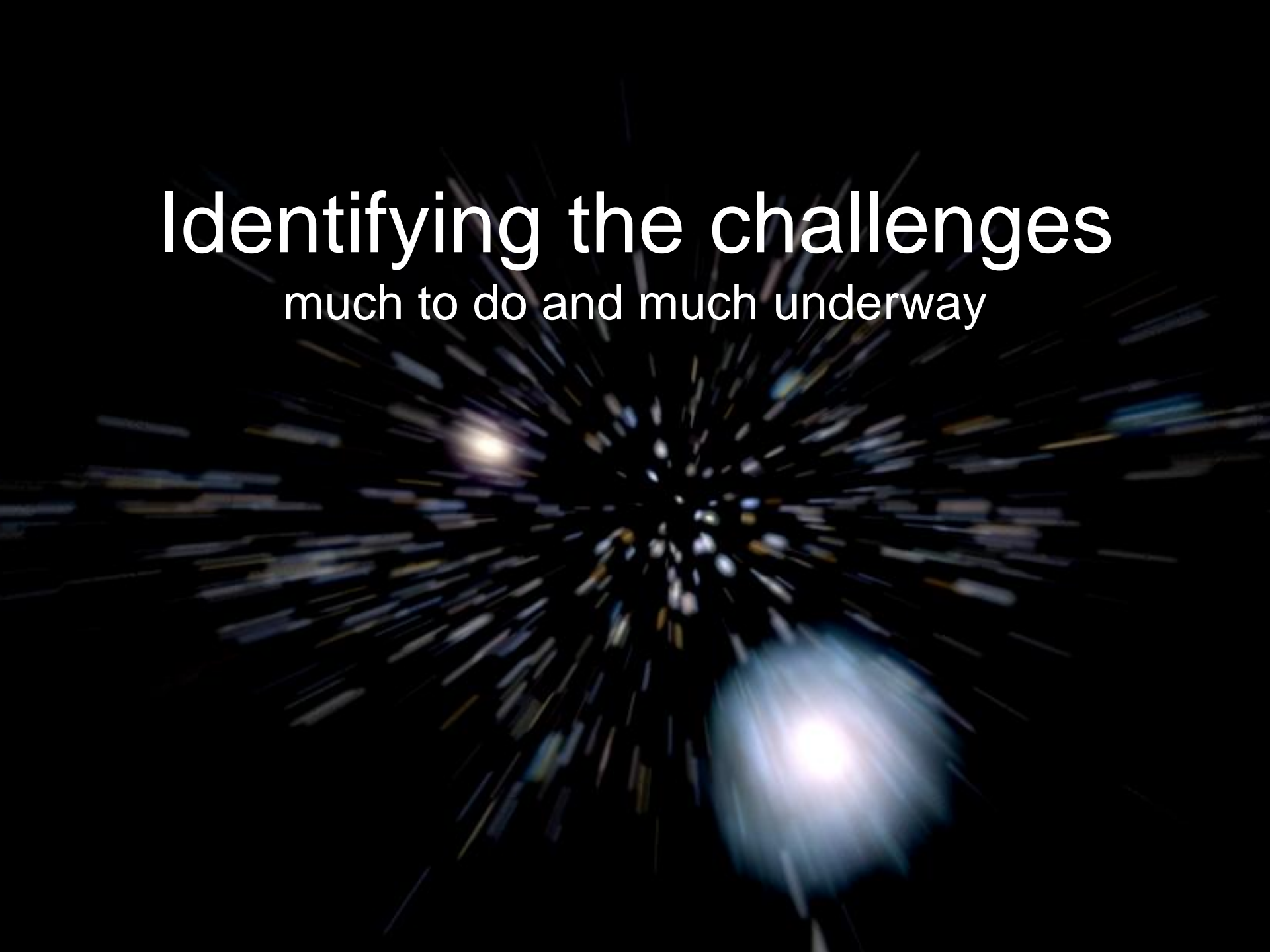
D/H + Nuclear physics at $t \sim 1$ sec $\rightarrow \Omega_b h^2 = 0.022 \pm 0.002$



CMB + Gravity driven acoustic oscillations at $t = 380,000$ yrs $\rightarrow \Omega_b h^2 = 0.0222 \pm 0.0002$

Identifying the challenges

much to do and much underway



Consensus Cosmology



Rests upon three mysterious pillars
All implicate new physics!



Full Court Press!!

- Produce at LHC
- Detect particles in our halo
- Detect annihilation products

OF MOOSE DIAGRAM DARK MATTER CANDIDATES

MT90

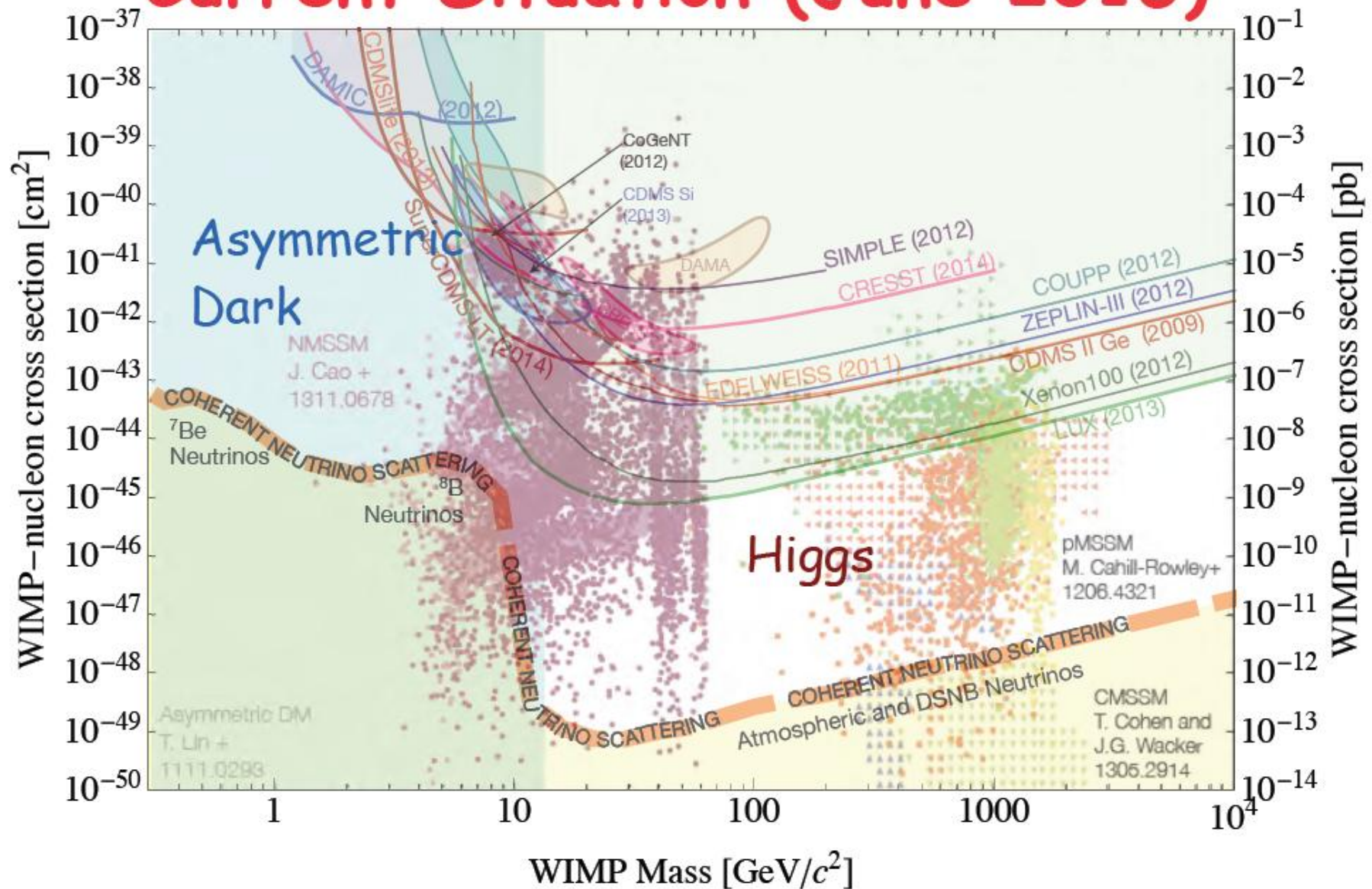


Neutrinos
contribute a
few 0.1%

79th Nobel Symposium 1990 (Gräfvallen, SE)

Direct Detection

Current Situation (June 2015)



Lots of intriguing hints

- Puzzles to resolve: Pamela/AMS, “the Hooperon”, DAMA/Libra, how much do neutrinos contribute? (cosmology seems to have the best shot)
- Importance of complementarity
 - Backgrounds: indirect (astrophysical), direct (radioactivity) & accelerator (SM/new physics)
 - Triple signature to convince the skeptic
- Time for new ideas? Cast the net more broadly? (Ugh, I hate the dark sector)

DARK ENERGY

MAY BE THE MOST

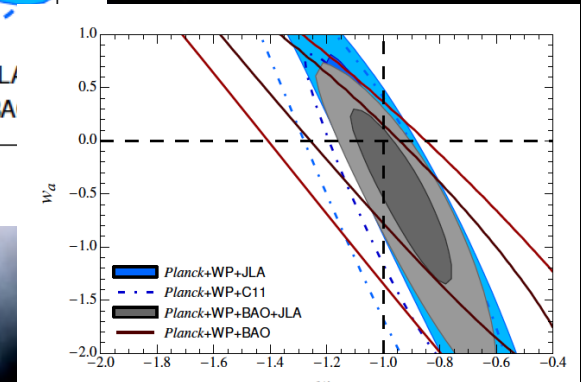
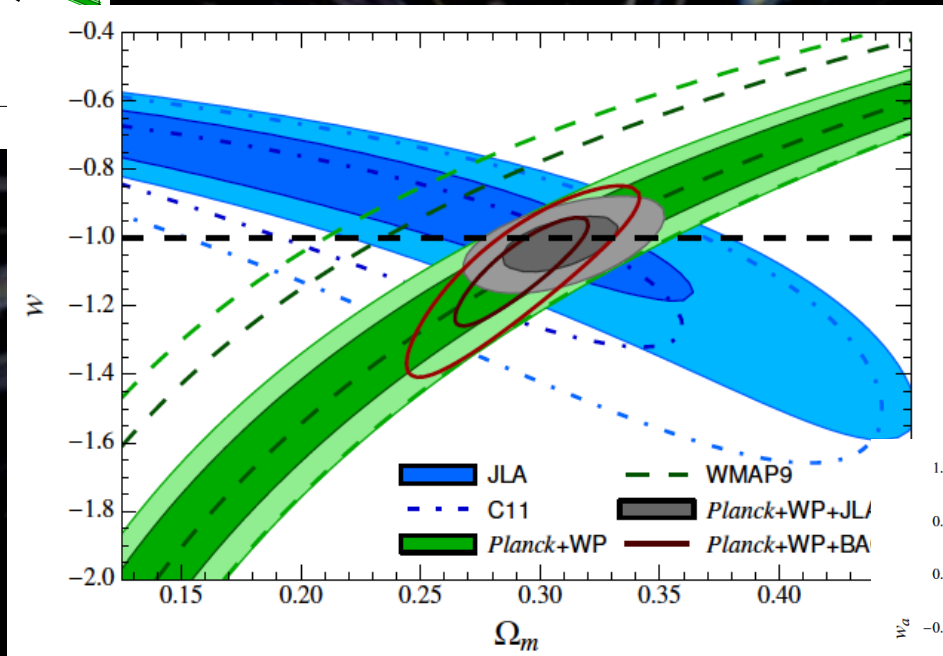
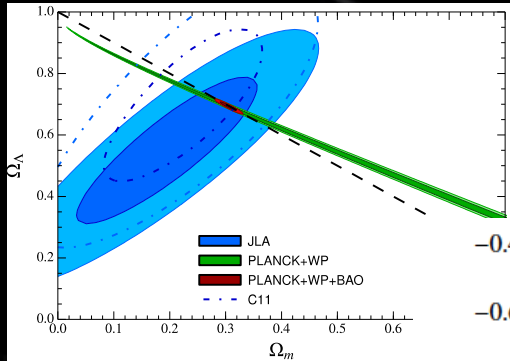
PROFOUND PROBLEM

IN ALL OF SCIENCE TODAY

So why does the solution look so simple? Λ

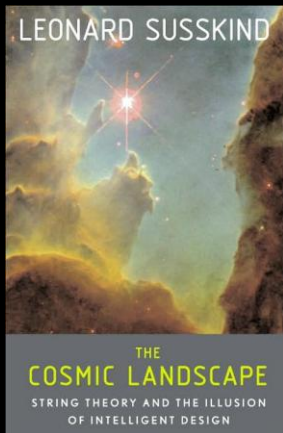
Wow!

$$w = -1.018 \pm 0.06$$



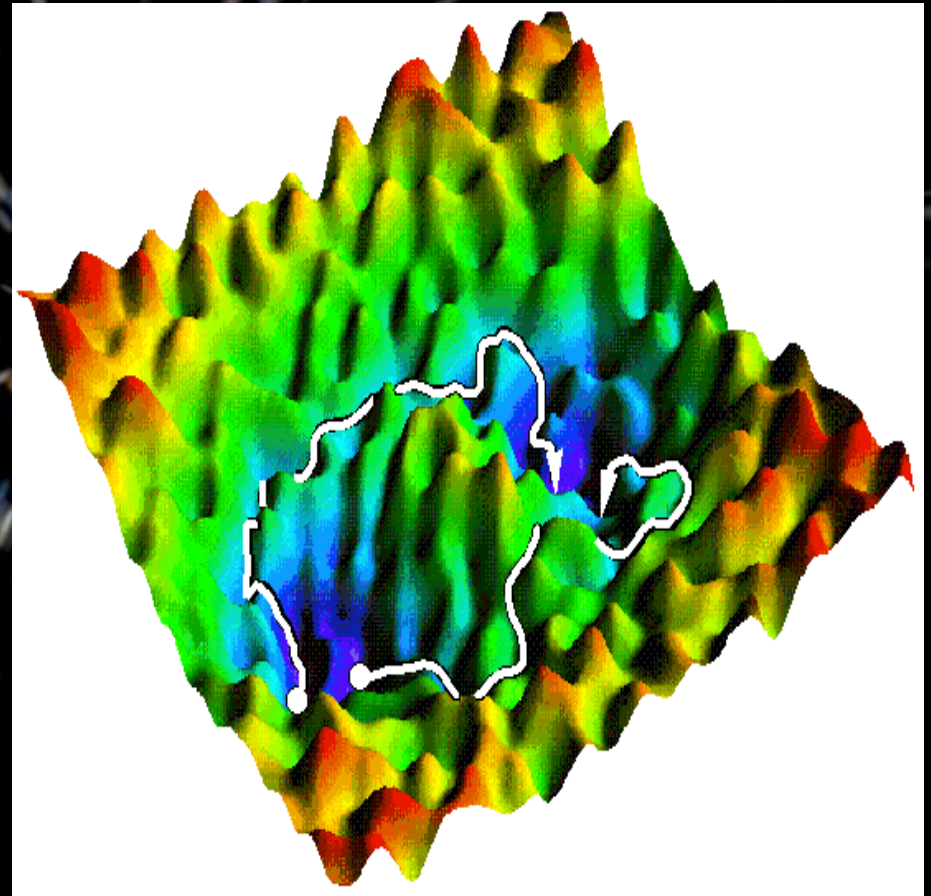


H.L. Mencken: “For every complex problem, there is a solution that is simple, elegant and wrong!”



The Extravagant Solution

- M-theory $\rightarrow 10^{500}$ vacua
- $\rho_{\text{vac}} \sim O(m_{\text{Pl}}^4) \pm O(m_{\text{Pl}}^4)$
- Universe has a multiverse structure & we were lucky (narcissistic principle)
- Testable?



New grand principle?

NO DARK ENERGY

NEW ASPECT OF GRAVITY

→ "EMPTY" UNIVERSE
UNDERGOES ACCELERATED
EXPANSION!

AVERAGE MATTER DENSITY TODAY $\approx 10^{-29} \text{ g/cm}^3$
 $\approx 10^{-100} \times \text{DENSITY AFTER INFLATION}$

Two big questions

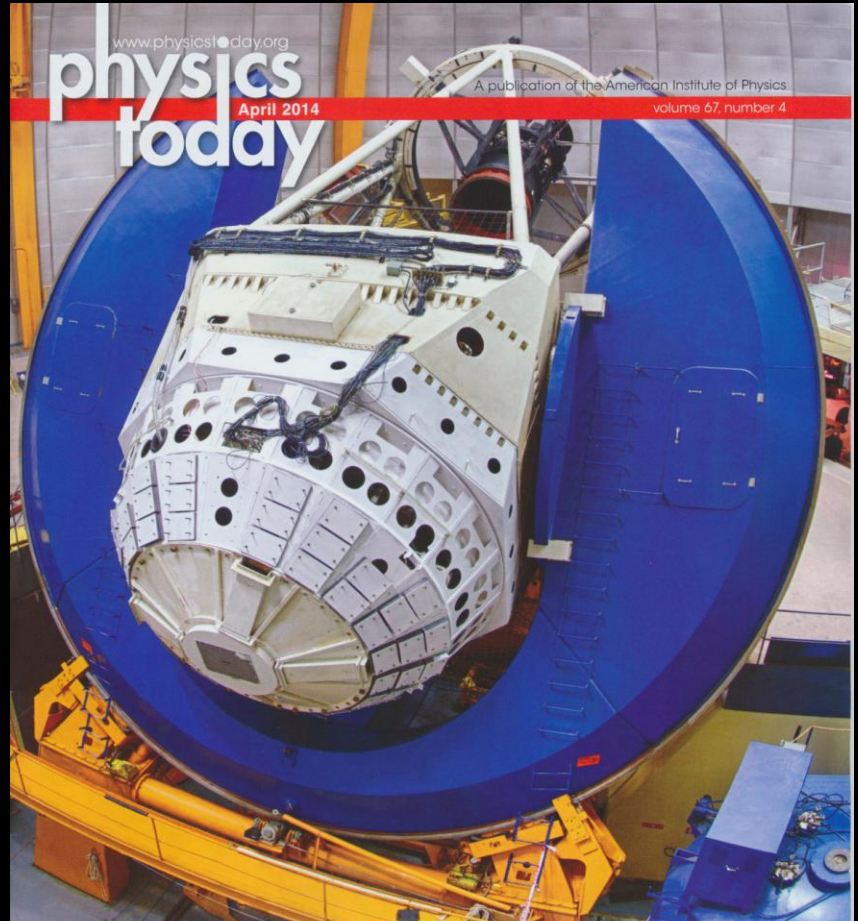
1. Does dark energy vary with time?

w and friends

1. Does general relativity self consistently describe accelerated expansion?

Growth of structure ...

Dark Energy Survey



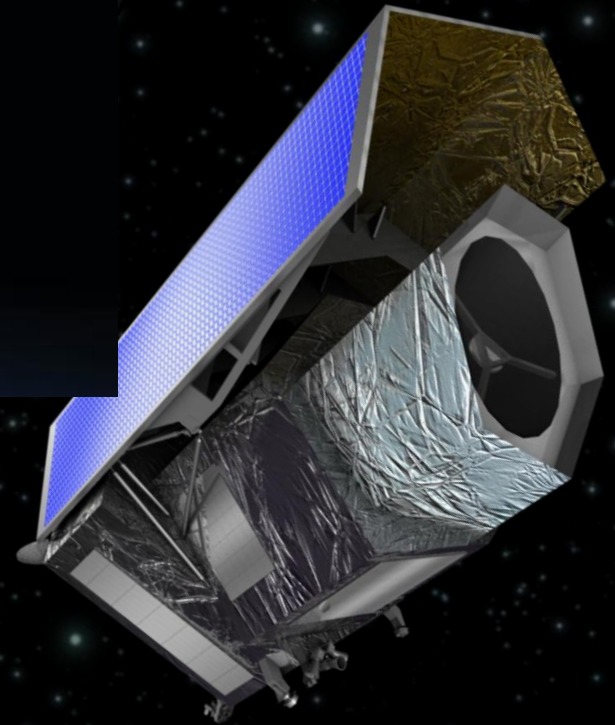
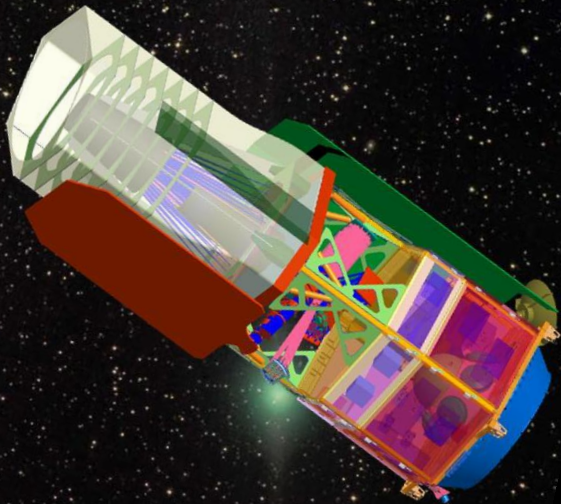
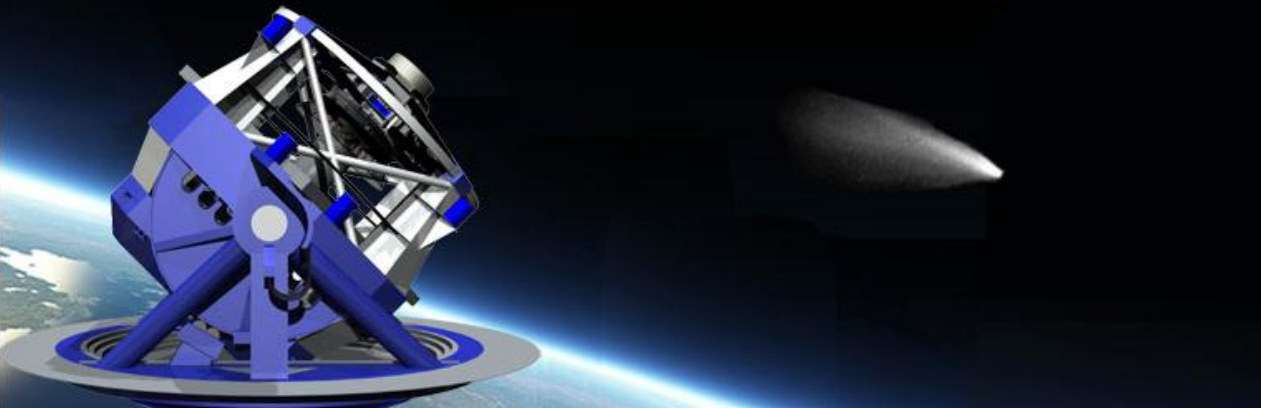
Hunting for dark energy

also:

- The electron's electric dipole moment ◀
- Nukes in the 21st century ◀
- A stratospheric skydive ◀

LST

Large Synoptic Survey Telescope



WFIRST/Euclid

Serious testing of Inflation has *begun*

Key Predictions



Flat Universe



Almost scale-invariant, Gaussian perturbations:

$$|(n-1)| \sim 0.1 \text{ and } |dn/d\ln k| \sim 0.001$$

- Gravity waves: spectrum, but not amplitude

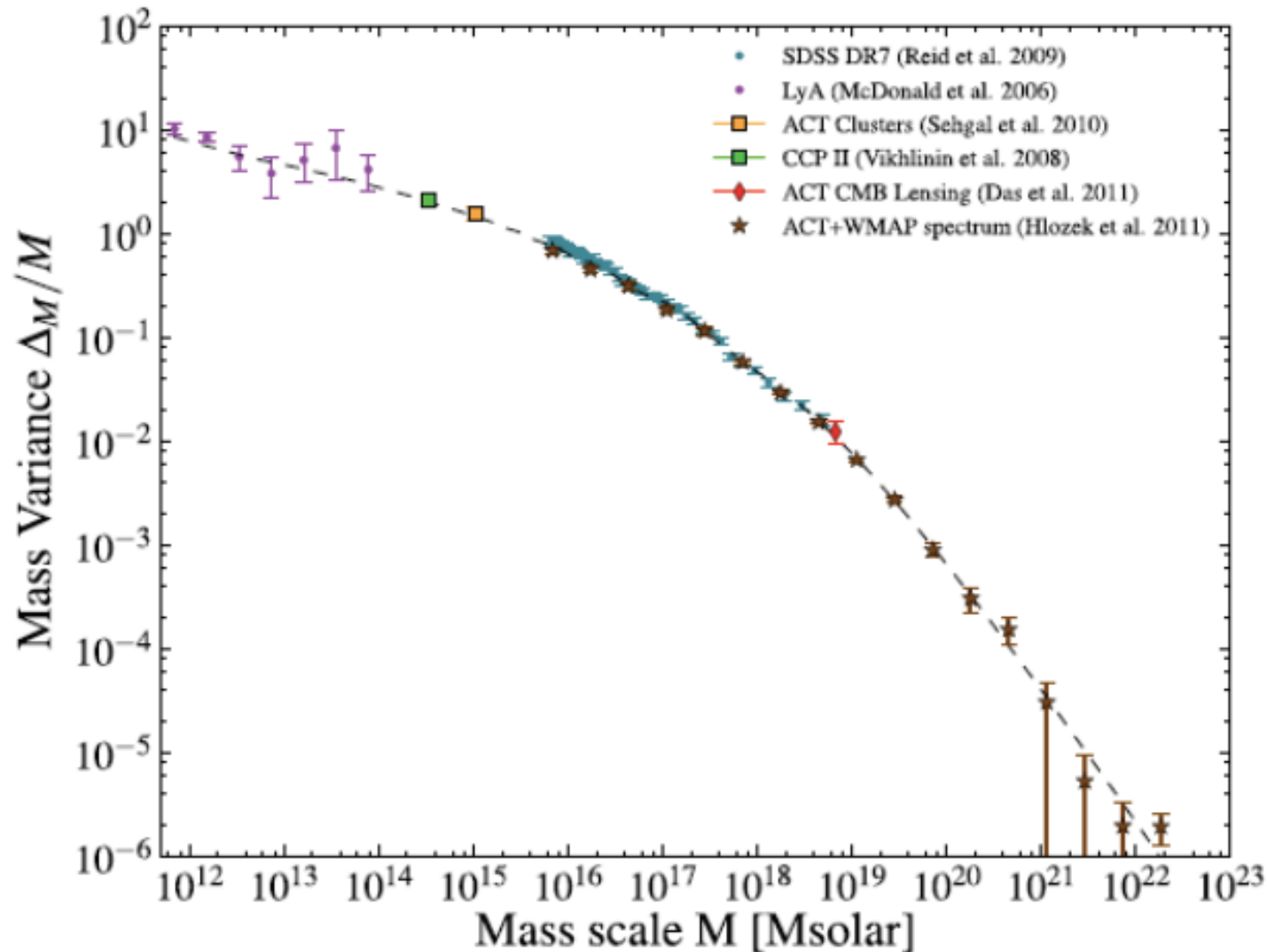



Cold Dark Matter Scenario

Key Results

- $\Omega_0 = 1.00 \pm 0.006$
- $n = 0.965 \pm 0.005$; $dn/d\ln k = -0.002 \pm 0.013$; no evidence for nonGaussianity
- $r < 0.10$ (95% cl)

Success of CDM (from Spergel)





WILD GOOSE CHASE



B-modes detected!

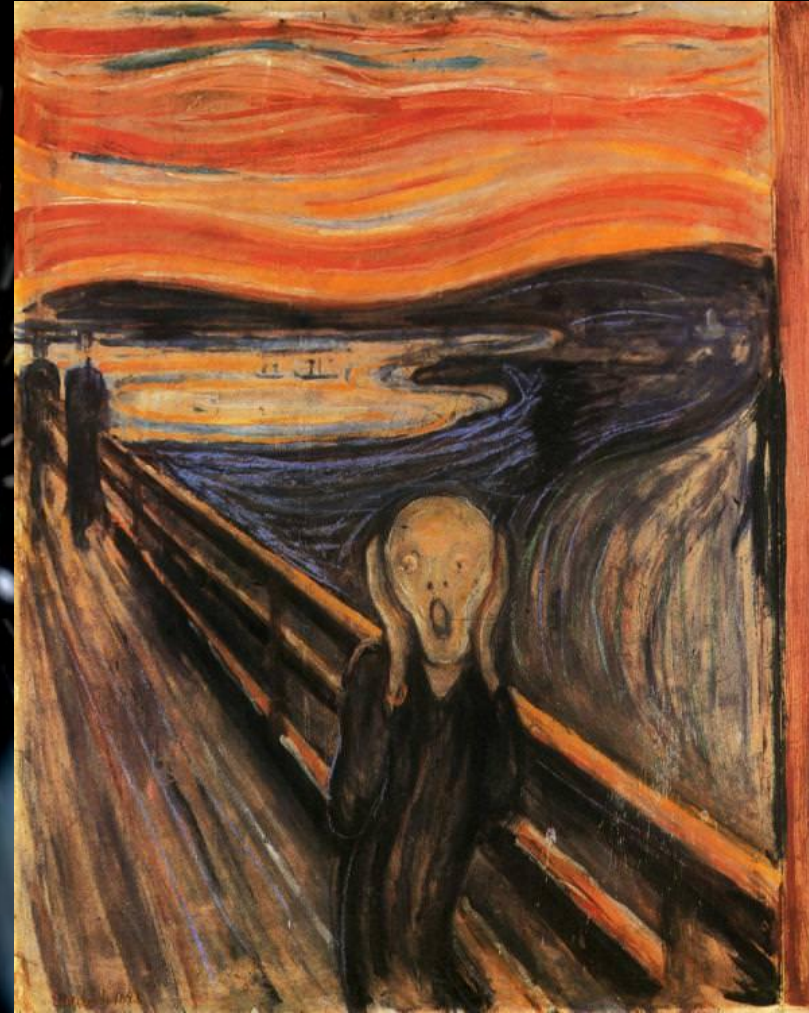
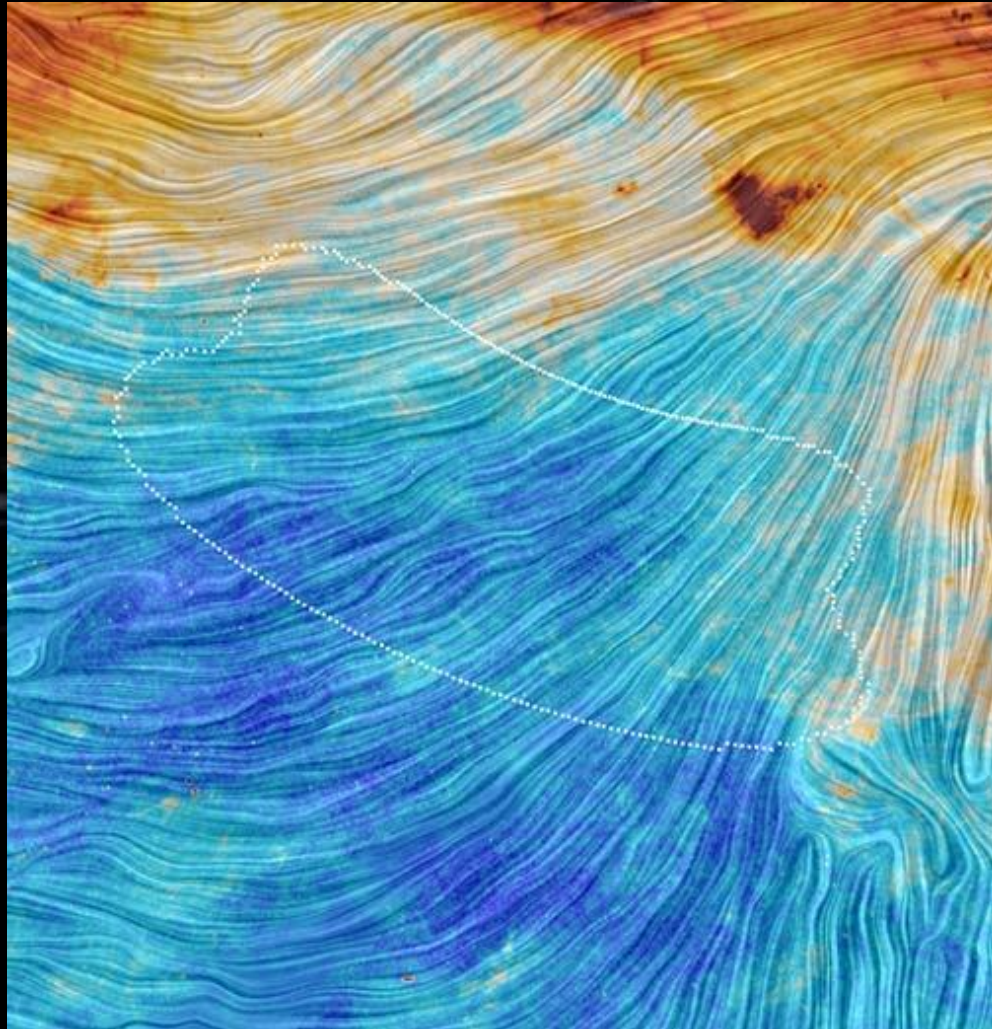
17 March 2014/BICEP2



Most exciting interpretation

- Smokin' gun for inflation
 - $r = \text{tensors/scalars} = 0.2^{+0.07}_{-0.05}$
- The when and clues about how **10^{-38} sec**
 - 10^{-38} sec
 - Energy scale of 2×10^{16} GeV = 10^{12} x LHC
 - “Planck scale physics”:
 - $\Delta\phi \sim m_{\text{pl}}$
 - $V \sim (2 \times 10^{16} \text{ GeV})^4$, $V' \sim 1.2V/m_{\text{pl}}$, $V'' \sim 0.6V/m_{\text{pl}}^2$
- Remember, “extraordinary results require extraordinary evidence” ... Carl Sagan

The bane of astronomy: dust



When all the “dust” settled: BICEP2/Keck/Planck joint analysis

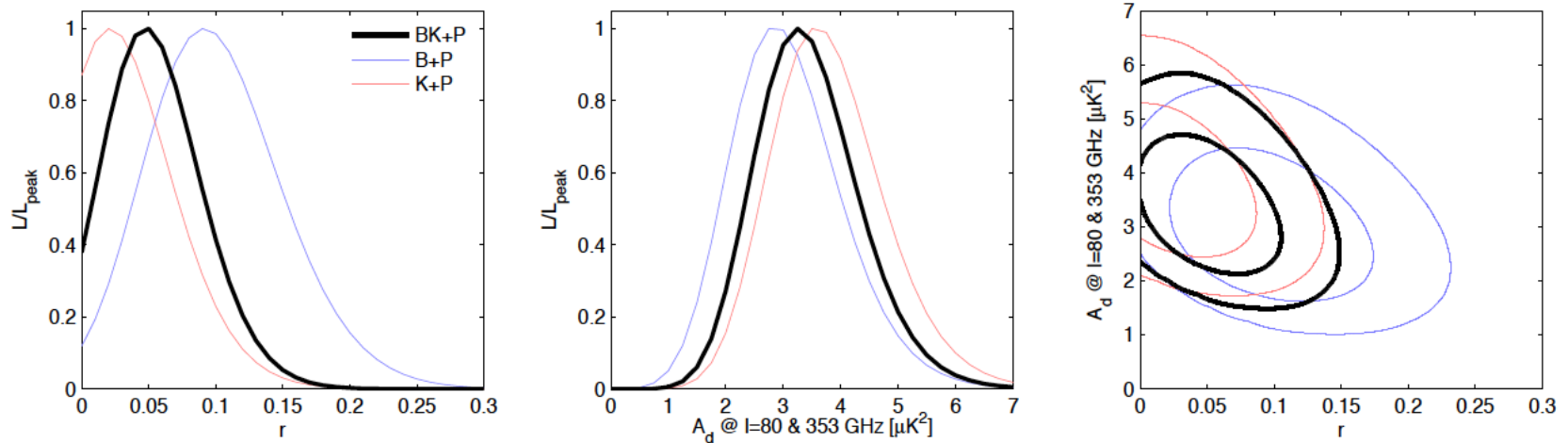


FIG. 6. Likelihood results from a basic lensed- Λ CDM+ r +dust model, fitting BB auto- and cross-spectra taken between maps at 150 GHz, 217, and 353 GHz. The 217 and 353 GHz maps come from *Planck*. The primary results (heavy black) use the 150 GHz combined maps from BICEP2/*Keck*. Alternate curves (light blue and red) show how the results vary when the BICEP2 and *Keck Array* only maps are used. In all cases a Gaussian prior is placed on the dust frequency spectrum parameter $\beta_d = 1.59 \pm 0.11$. In the right panel the two dimensional contours enclose 68% and 95% of the total likelihood.

B-modes detected! Dust detected at $> 3\sigma$
Evidence for something else (GWs)

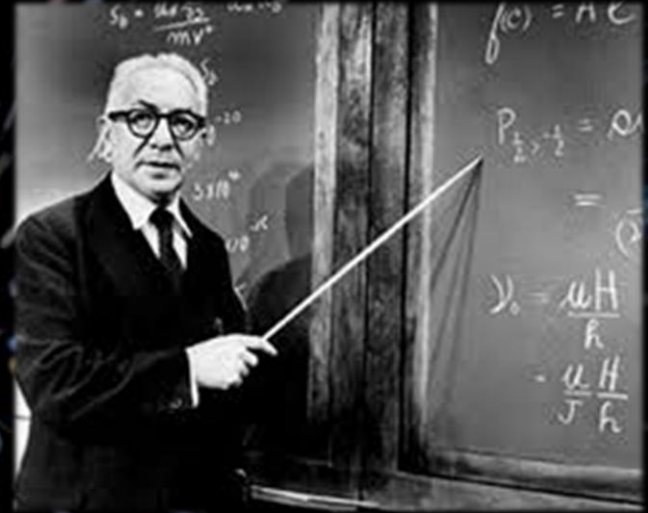
Grand challenges and big surprises

plenty to motivate the most talented



The Complicated Universe

- Atoms only:
Democritus to 1964
- + Photons: 1964
- + Neutrinos (e, μ):
1967
- + Exotic dark matter:
1981
- + CDM: 1983/4
- + Massive neutrinos:
1998
- + Dark energy: 1998
- + τ neutrino: 2000
- + ?? : 20??



I.I. Rabi: Who ordered that?
Origin of dimensionless
ratios?
More relics?

WHY SHOULD BARYONS AND EXOTIC RELIC PARTICLES HAVE COMPARABLE DENSITIES?

MICHAEL S. TURNER

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and

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and

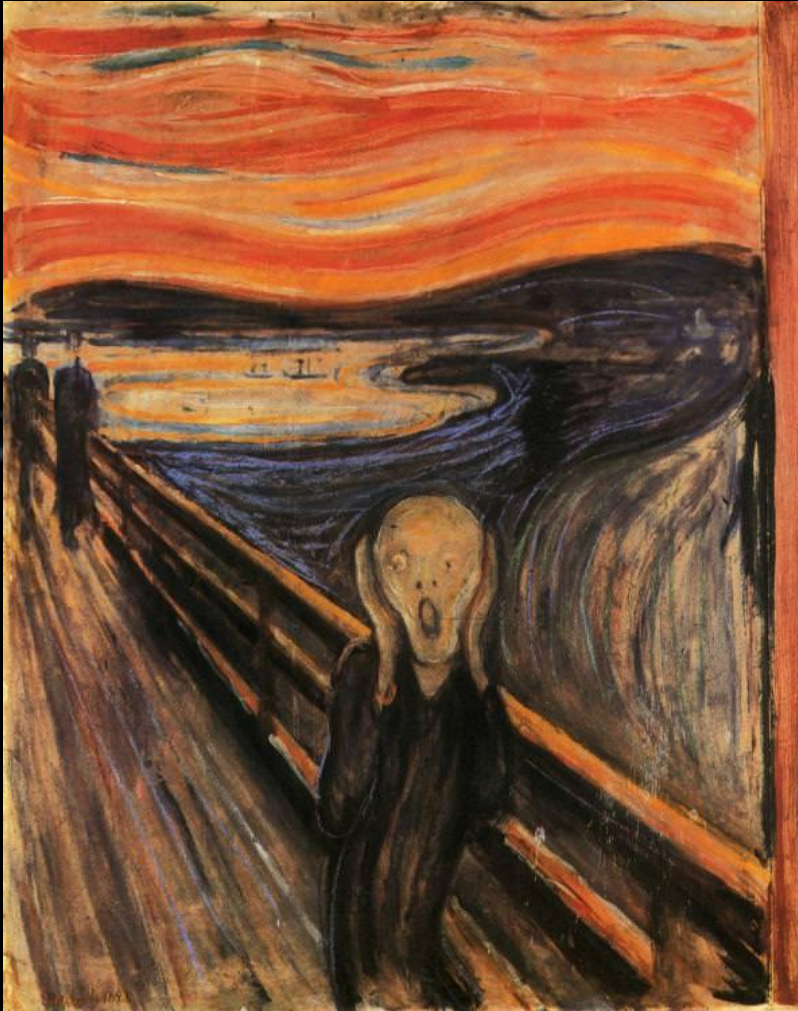
BERNARD J. CARR

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Mile End Road, London E1 4NS, England*

Received 23 December 1986

Observations suggest that the mass density of the Universe is dominated, not by ordinary matter, but by exotic particles which are a relic of the Big Bang. In this case, a new dimensionless cosmological ratio arises, the ratio of the mass density in ordinary matter to that in exotic matter, whose value is about 0.1. *A priori*, it might seem remarkable that this ratio should be so close to unity. However, we point out that, for many exotic dark matter candidates, the ratio is related to the fundamental scales of particle physics. A value of order unity arises naturally provided rather simple relationships exist between these scales. If the exotic particles are of a kind whose relic abundance is determined by annihilations (e.g., the photino or a heavy neutrino), then the required relationship is already satisfied for independent, cosmological reasons.

What to do about the multiverse



- Most important discovery since Copernicus?
- Is it science? (not testable)
- Many true believers (left coast) and not enough doubters

A woman with long, thick brown dreadlocks is shown from the chest up. She is wearing a bright orange long-sleeved shirt and a colorful, patterned headband. Her eyes are closed, and she has a peaceful, upward-looking expression. The background is blurred, showing other people in a crowd.

**I BELIEVE IN
SCIENCE**

**SO I DON'T BELIEVE IN MAGIC OR THE
MULTIVERSE**

More grand challenges

- Origin of ordinary matter (baryo/leptogenesis)
- “Completion of inflation” (Landau-Ginzburg \rightarrow BCS) and/or development of a worthy competitor
- Experimental
 - $21(1+z)$ cm science: next CMB?
 - More from the CMB than multipoles?
 - SUSY or something better at LHC
 - Cosmic neutrinos (never forget neutrinos!)

Game changing idea, but not time for a coronation

- No fundamental theory
- Does not address initial singularity or initial conditions
- Does not address cosmological constant problem
- Like “duct tape”, very useful but ...
 - Only postpones appearance of inhomogeneity
 - not all initial conditions inflate
- Unsettling, uncertain predictions: eternal inflation; multiverse;

SPACE-TIME GEOMETRY

our initial geometry



smooth, small
ripples



15 Gyr



generic initial
geometry



10^{-43} sec



NOT LOGICAL INCONSISTENCY!



DILEMMA OF INITIAL DATA
(WHY SO SPECIAL?)

Crack in the cosmic egg?

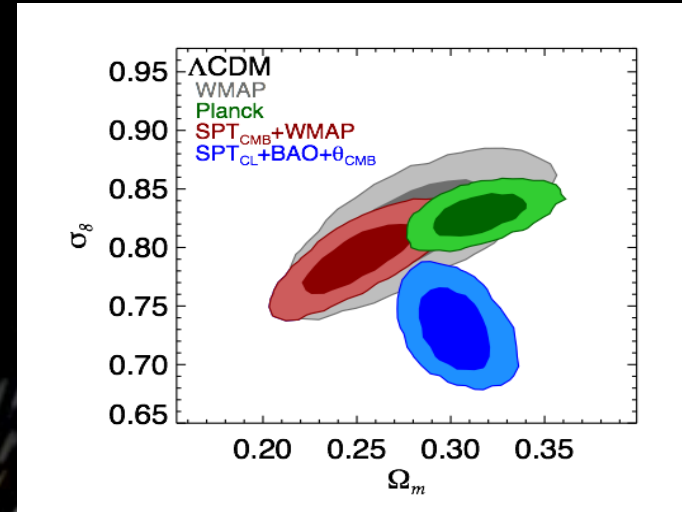
... which we have been waiting for?!#



- CDM anomalies
 - on small scales
 - SIDM? WDM? MDM?WMD?
- CMB anomalies
 - $H_0/\sigma_8/N_v$
 - Low ℓ
- Time varying constants, dark flows, ...

Theorem: real cracks get bigger – not smaller – with time! (2nd law)

“Planck/WMAP” H_0 trade space



- Planck: “low” H_0 , high σ_8 & $N_v = 3$
(\rightarrow Planck predicts 3X clusters SPT observes)
- WMAP: “consensus H_0 ”, “right number of clusters” & $N_v = 3.8$

Cluster abundance exponentially sensitive to σ_8

Cosmological death spiral?

- Continued success of Λ CDM, no discovery of DM or understanding of DE
 - akin to success of SM or Ptolemy
 - continued overemphasis on the multiverse
- A personal observation
 - 1980s: lots of powerful, new ideas & little data
 - 2000s: lots of data & fewer powerful, new ideas (related to lots of data?)

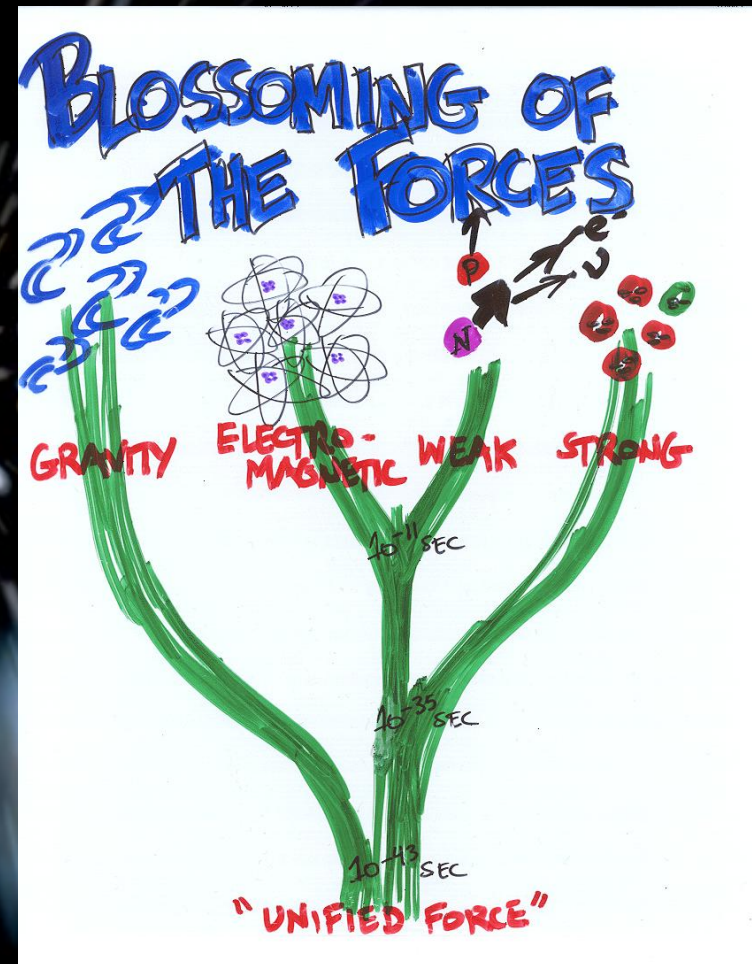


I am bullish on the future!

The mysteries of cosmology
and of particle physics are
inextricably tied together

The combined tools and
intellectual capital will push
both fields forward

The best is yet to come!



Thanks to our hostess with the mostest
for the meeting with the mostest! And bon
voyage on your journey with Nordita

