

Gravity OS-Diagrams

Enrico Herrmann

Based on: arXiv:1604.03479

in collaboration with:

Jaroslav Trnka

see also: Heslop, Lipstein: arXiv:1604.03046

June 23, 2016

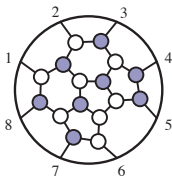


Caltech

Motivation

planar $\mathcal{N} = 4$ sYM \Rightarrow Hydrogen atom of the 21st century!

- Dual Conformal Invariance [Drummond, Henn, Smirnov, Sokatchev, Korchemsky,...]
- Relation to Wilson loops and Correlation Functions [Mason, Skinner, Caron-Huot, Alday, Eden, Korchemsky, Maldacena, Sokatchev, ...]
- Yangian Invariance and Integrability [Drummond, Henn, Plefka, Beisert, Staudacher, Alday, Viera, Basso,...]



Mathematical structures beyond planar $\mathcal{N} = 4$ sYM?

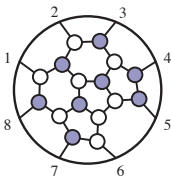
non-planar $\mathcal{N} = 4$ sYM

$\mathcal{N} = 8$ SUGRA

Motivation

planar $\mathcal{N} = 4$ sYM \Rightarrow Hydrogen atom of the 21st century!

- Dual Conformal Invariance [Drummond, Henn, Smirnov, Sokatchev, Korchemsky,...]
- Relation to Wilson loops and Correlation Functions [Mason, Skinner, Caron-Huot, Alday, Eden, Korchemsky, Maldacena, Sokatchev, ...]
- Yangian Invariance and Integrability [Drummond, Henn, Plefka, Beisert, Staudacher, Alday, Viera, Basso,...]



Mathematical structures beyond planar $\mathcal{N} = 4$ sYM?

non-planar $\mathcal{N} = 4$ sYM

$\mathcal{N} = 8$ SUGRA ✓

Outline

- 1 On-Shell Diagrams in $\mathcal{N} = 4$ sYM
 - Why On-Shell Diagrams?
 - Three-Point Amplitudes
 - Grassmannian Formulation for On-Shell Diagrams
- 2 On-Shell Diagrams in Gravity
 - A First Look at Gravity On-Shell Diagrams
 - Grassmannian Formula for Gravity
- 3 Comments-Conclusion

On-Shell methods, Cuts of loop amplitudes and Generalized Unitarity

[Britto, Cachazo, Feng, Witten; Bern, Dixon, Kosower; ...]

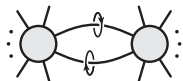
- core idea: on-shell amplitudes break up into products of simpler amplitudes on all factorization channels
- amplitudes are **fixed** from their singularities

locality: $\frac{1}{P^2}$ propagators

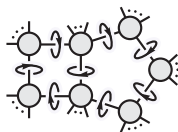
unitarity: factorization on poles

$$\partial \left| \begin{array}{c} \bullet \bullet \bullet \\ \textcircled{A}_n^\ell \\ \bullet \bullet \bullet \end{array} \right| = \sum_{L,R} \begin{array}{c} \bullet \bullet \bullet \\ \textcircled{L} \text{---} \textcircled{R} \\ \bullet \bullet \bullet \end{array} + \sum_a \begin{array}{c} \bullet \bullet \bullet \\ \textcircled{A}_{n+2}^\ell \\ \textcircled{a+1} \text{---} \textcircled{a} \\ \bullet \bullet \bullet \end{array}$$

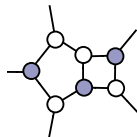
iterative cuts:



\Rightarrow



Maximal Cut



OS-Diagram

Three-Point Amplitudes

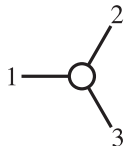
massless theories in $D = 4$: 3pt-amplitude completely fixed by Poincaré invariance

$$p^\mu \rightarrow p^{\alpha\dot{\alpha}} \quad \Rightarrow \quad p^2 = \det p^{\alpha\dot{\alpha}} = 0 \quad \Rightarrow \quad p^{\alpha\dot{\alpha}} = \lambda^\alpha \tilde{\lambda}^{\dot{\alpha}}$$

On-shell conditions have two solutions:

$\overline{\text{MHV}}, k = 1$

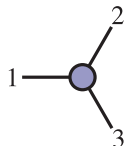
$$\lambda_1 \sim \lambda_2 \sim \lambda_3$$



$$\mathcal{A}_3^{\overline{\text{MHV}}} = \frac{\delta^4(P) \delta^4([\!12\!] \tilde{\eta}_3 + [\!23\!] \tilde{\eta}_1 + [\!31\!] \tilde{\eta}_2)}{[\!12\!][\!23\!][\!31\!]}$$

$\text{MHV}, k = 2$

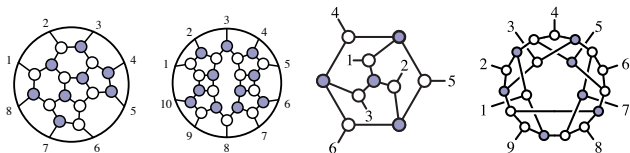
$$\tilde{\lambda}_1 \sim \tilde{\lambda}_2 \sim \tilde{\lambda}_3$$



$$\mathcal{A}_3^{\text{MHV}} = \frac{\delta^4(P) \delta^8(\lambda_1 \tilde{\eta}_1 + \lambda_2 \tilde{\eta}_2 + \lambda_3 \tilde{\eta}_3)}{\langle 12 \rangle \langle 23 \rangle \langle 31 \rangle}$$

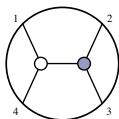
On-Shell Diagrams as Gluing of 3pt-amplitudes

use 3pt-amplitudes as building blocks for more complicated On-Shell diagrams



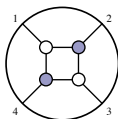
results are functions of external kinematics,

$$P > 4L$$



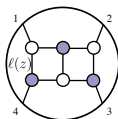
Extra δ -functions

$$P = 4L$$



Leading Singularity

$$P < 4L$$



Unfixed Parameters

Grassmannian Formulation for On-Shell Diagrams

[Arkani-Hamed, Bourjaily, Cachazo, Goncharov, Postnikov, Trnka]

deep connection: **On-Shell Diagrams** \Leftrightarrow **Grassmannian** $G(k, n)$

$G(k, n)$: space of k -planes in n -dimensions

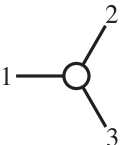
$\cong (k \times n)$ -matrices modulo $GL(k)$

planar OS-diags \Rightarrow positive Grassmannian \rightarrow connection to
combinatorics & algebraic geometry [Lusztig, Postnikov, Speyer, Williams, Knutson, Lam,...]

Motivation from physics: linearize momentum conservation!

Grassmannian Formulation for On-Shell Diagrams

Remember 3pt-amplitude ($k = 1$):

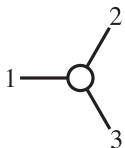


$$\mathcal{A}_3^{\overline{\text{MHV}}} = \frac{\delta^4(\sum_{i=1}^3 \lambda_i \tilde{\lambda}_i) \delta^4([12]\tilde{\eta}_3 + [23]\tilde{\eta}_1 + [31]\tilde{\eta}_2)}{[12][23][31]}$$

- mom. conservation is **quadratic** constraint on λ and $\tilde{\lambda}$
- **not** manifest that $\lambda_1 \sim \lambda_2 \sim \lambda_3$

Grassmannian Formulation for On-Shell Diagrams

Remember 3pt-amplitude ($k = 1$):



$$\mathcal{A}_3^{\overline{\text{MHV}}} = \frac{\delta^4(\sum_{i=1}^3 \lambda_i \tilde{\lambda}_i) \delta^4([12]\tilde{\eta}_3 + [23]\tilde{\eta}_1 + [31]\tilde{\eta}_2)}{[12][23][31]}$$

- mom. conservation is **quadratic** constraint on λ and $\tilde{\lambda}$
- **not** manifest that $\lambda_1 \sim \lambda_2 \sim \lambda_3$

cure:

$$\lambda_i^\alpha = \alpha_i \rho^\alpha, \quad \rho: \text{reference spinor}$$

$$\delta\left(\sum \lambda_i \tilde{\lambda}_i\right) \xrightarrow{\lambda_i = \alpha_i \rho} \delta\left(\rho \sum \alpha_i \tilde{\lambda}_i\right) \Rightarrow \alpha_1 \tilde{\lambda}_1 + \alpha_2 \tilde{\lambda}_2 + \alpha_3 \tilde{\lambda}_3 = 0$$

$$\alpha_1 = [23], \quad \alpha_2 = [31], \quad \alpha_3 = [12]$$

suggests:

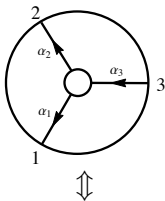
$$\mathcal{A}_3^{\overline{\text{MHV}}} \stackrel{?}{=} \int \frac{d\alpha_1}{\alpha_1} \frac{d\alpha_2}{\alpha_2} \frac{d\alpha_3}{\alpha_3} \int d^2\rho \left(\prod_{i=1}^3 \delta^2(\alpha_i \rho - \lambda_i) \right) \delta^2\left(\sum \alpha_i \tilde{\lambda}_i\right) \delta^4\left(\sum \alpha_i \tilde{\eta}_i\right)$$

Grassmannian Formulation for On-Shell Diagrams

$$\mathcal{A}_3^{\overline{\text{MHV}}} \stackrel{?}{=} \int \frac{d\alpha_1}{\alpha_1} \frac{d\alpha_2}{\alpha_2} \frac{d\alpha_3}{\alpha_3} \int d^2\rho \left(\prod_{i=1}^3 \delta^2(\alpha_i \rho - \lambda_i) \right) \delta^2(\sum \alpha_i \tilde{\lambda}_i) \delta^4(\sum \alpha_i \tilde{\eta}_i)$$

GL(1)-redundancy: $\alpha_i \rightarrow t\alpha_i$, $\rho \rightarrow \frac{1}{t}\rho \Rightarrow$ gauge-fix!

Encode **linear relations** in terms of $(k \times n)$ -matrix $C \bmod GL(k)$.



$$C = \begin{pmatrix} \alpha_1 \alpha_3 & \alpha_2 \alpha_3 & 1 \end{pmatrix}$$

$$\delta(P) \equiv \delta(\lambda \cdot \tilde{\lambda}) \rightarrow \delta(C \cdot \tilde{\lambda}) \delta(C^\perp \cdot \lambda)$$

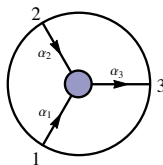
C^\perp : orthogonal matrix to C

arrows \leftrightarrow gauge-fixing

$$\mathcal{A}_3^{\overline{\text{MHV}}} = \frac{1}{\text{vol}(GL(1))} \int \frac{d\alpha_1}{\alpha_1} \frac{d\alpha_2}{\alpha_2} \frac{d\alpha_3}{\alpha_3} \delta(C \cdot \tilde{\lambda}) \delta(C^\perp \cdot \lambda) \delta(C \cdot \tilde{\eta})$$

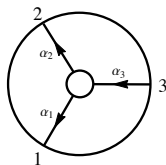
Grassmannian Formulation for On-Shell Diagrams

MHV($k = 2$)



$$C = \begin{pmatrix} 1 & 0 & \alpha_1 \alpha_3 \\ 0 & 1 & \alpha_2 \alpha_3 \end{pmatrix}$$

$\overline{\text{MHV}}(k = 1)$



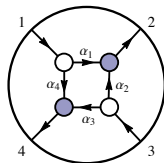
$$C = \begin{pmatrix} \alpha_1 \alpha_3 & \alpha_2 \alpha_3 & 1 \end{pmatrix} .$$

Note: for MHV-amplitude we **partially gauge-fixed** $GL(2)$!

Convenient to keep $GL(1)_v$ redundancy in each vertex v

Grassmannian Formulation for On-Shell Diagrams

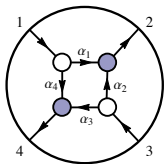
- glue 3pt-amplitudes together \rightarrow build bigger on-shell diagrams
- encode momentum conservation by $(k \times n)$ -matrix C
- n : number of external legs,
 $k = 2n_B + n_W - n_I$: MHV-degree



$$C = \begin{pmatrix} 1 & \alpha_1 & 0 & \alpha_4 \\ 0 & \alpha_2 & 1 & \alpha_3 \end{pmatrix}$$

Grassmannian Formulation for On-Shell Diagrams

- glue 3pt-amplitudes together \rightarrow build bigger on-shell diagrams
- encode momentum conservation by $(k \times n)$ -matrix C
- n : number of external legs,
 $k = 2n_B + n_W - n_I$: MHV-degree



$$C = \begin{pmatrix} 1 & \alpha_1 & 0 & \alpha_4 \\ 0 & \alpha_2 & 1 & \alpha_3 \end{pmatrix}$$

math

α_i definite sign $\Rightarrow C$ has pos. minors \Rightarrow **pos. Grassmannian**

physics

$$\Omega = \frac{d\alpha_1}{\alpha_1} \frac{d\alpha_2}{\alpha_2} \frac{d\alpha_3}{\alpha_3} \frac{d\alpha_4}{\alpha_4} \delta(C \cdot Z)$$

$$\delta(C \cdot Z) \equiv \delta(C \cdot \tilde{\lambda}) \delta(C^\perp \cdot \lambda) \delta(C \cdot \tilde{\eta})$$

equal to product of 3pt amplitudes!

Note: $d \log$ -measure! \rightarrow special for Yang-Mills
 relation to amplitudes: hidden symmetries, geometric formulation,...

[see Trnka talk at Amplitudes]

A first look at gravity on-shell diagrams

- 3pt-amplitudes: **squaring relation**

$$A_3 = \frac{\langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \langle 31 \rangle} \rightarrow M_3 = \left(\frac{\langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \langle 31 \rangle} \right)^2$$

- general on-shell diagram (product of 3pt amplitudes)

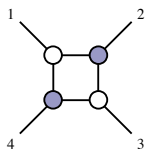
$$(\text{YM})^2 = (\text{GR}) \times (\phi^3)$$

“don't square the propagators”

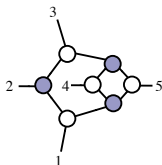
- (ϕ^3) factor changes expressions drastically

A first look at gravity on-shell diagrams

study some data (MHV leading singularities)



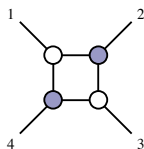
$$\frac{[13][24]}{\langle 12 \rangle \langle 13 \rangle \langle 14 \rangle \langle 23 \rangle \langle 24 \rangle \langle 34 \rangle}$$



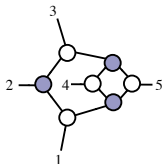
$$\frac{[12][23][45]^2}{\langle 12 \rangle \langle 14 \rangle \langle 15 \rangle \langle 23 \rangle \langle 34 \rangle \langle 35 \rangle}$$

A first look at gravity on-shell diagrams

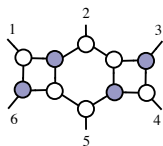
study some data (MHV leading singularities)



$$\frac{[13][24]}{\langle 12 \rangle \langle 13 \rangle \langle 14 \rangle \langle 23 \rangle \langle 24 \rangle \langle 34 \rangle}$$



$$\frac{[12][23][45]^2}{\langle 12 \rangle \langle 14 \rangle \langle 15 \rangle \langle 23 \rangle \langle 34 \rangle \langle 35 \rangle}$$

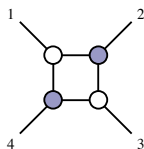


$$\frac{\langle 5 | Q_{16} | 2 \rangle \langle 2 | Q_{34} | 5 \rangle [16]^2 [34]^2}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 45 \rangle \langle 56 \rangle \langle 61 \rangle \langle 25 \rangle^2}$$

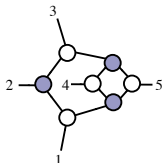
- nontrivial numerators!
- higher power poles possible! (unlike YM)

A first look at gravity on-shell diagrams

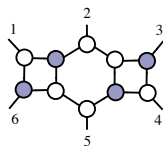
study some data (MHV leading singularities)



$$\frac{[13][24]}{\langle 12 \rangle \langle 13 \rangle \langle 14 \rangle \langle 23 \rangle \langle 24 \rangle \langle 34 \rangle}$$



$$\frac{[12][23][45]^2}{\langle 12 \rangle \langle 14 \rangle \langle 15 \rangle \langle 23 \rangle \langle 34 \rangle \langle 35 \rangle}$$

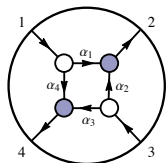


$$\frac{\langle 5 | Q_{16} | 2 \rangle \langle 2 | Q_{34} | 5 \rangle [16]^2 [34]^2}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 45 \rangle \langle 56 \rangle \langle 61 \rangle \langle 25 \rangle^2}$$

Detailed analysis

- nontrivial numerators! → **collinearity condition** in the vertices
- higher power poles possible! (unlike YM) → **infinite momenta**

Grassmannian Formula for Gravity

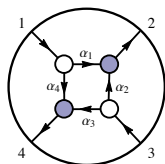


$$C = \begin{pmatrix} 1 & \alpha_1 & 0 & \alpha_4 \\ 0 & \alpha_2 & 1 & \alpha_3 \end{pmatrix}$$

Inspired by the explicit data, can “discover” the gravity formula

- Yang–Mills: $\Omega = \frac{d\alpha_1}{\alpha_1} \frac{d\alpha_2}{\alpha_2} \frac{d\alpha_3}{\alpha_3} \frac{d\alpha_4}{\alpha_4} \delta(C \cdot Z)$
- Gravity: $\Omega = \frac{d\alpha_1}{\alpha_1^3} \frac{d\alpha_2}{\alpha_2^3} \frac{d\alpha_3}{\alpha_3^3} \frac{d\alpha_4}{\alpha_4^3} (\prod_v \Delta_v) \delta(C \cdot Z)$

Grassmannian Formula for Gravity



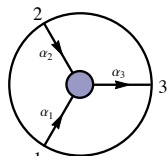
$$C = \begin{pmatrix} 1 & \alpha_1 & 0 & \alpha_4 \\ 0 & \alpha_2 & 1 & \alpha_3 \end{pmatrix}$$

Inspired by the explicit data, can “discover” the gravity formula

- Yang–Mills: $\Omega = \frac{d\alpha_1}{\alpha_1} \frac{d\alpha_2}{\alpha_2} \frac{d\alpha_3}{\alpha_3} \frac{d\alpha_4}{\alpha_4} \delta(C \cdot Z)$
- Gravity: $\Omega = \frac{d\alpha_1}{\alpha_1^3} \frac{d\alpha_2}{\alpha_2^3} \frac{d\alpha_3}{\alpha_3^3} \frac{d\alpha_4}{\alpha_4^3} (\prod_v \Delta_v) \delta(C \cdot Z)$
- special numerator Δ_v for each vertex
- α_i^3 poles

Can motivate this formula by looking at **3pt amplitudes**

Grassmannian Formula for Gravity

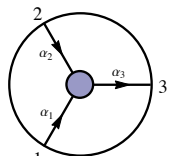


$$C = \begin{pmatrix} 1 & 0 & \alpha_1\alpha_3 \\ 0 & 1 & \alpha_2\alpha_3 \end{pmatrix}, \quad C^\perp = \begin{pmatrix} -\alpha_1\alpha_3 & -\alpha_2\alpha_3 & 1 \end{pmatrix}$$

- Need to modify measure by some dimensionful, permutation invariant object Δ

$$\delta(C^\perp \cdot \lambda) \Rightarrow -\overbrace{\alpha_1\lambda_1}^{\lambda_A} - \overbrace{\alpha_2\lambda_2}^{\lambda_B} + \overbrace{\frac{1}{\alpha_3}\lambda_3}^{\lambda_E} = 0 \Rightarrow \Delta = \langle AB \rangle = \langle BE \rangle = \langle EA \rangle$$

Grassmannian Formula for Gravity



$$C = \begin{pmatrix} 1 & 0 & \alpha_1\alpha_3 \\ 0 & 1 & \alpha_2\alpha_3 \end{pmatrix}, \quad C^\perp = \begin{pmatrix} -\alpha_1\alpha_3 & -\alpha_2\alpha_3 & 1 \end{pmatrix}$$

- Need to modify measure by some dimensionful, permutation invariant object Δ

$$\delta(C^\perp \cdot \lambda) \Rightarrow -\overbrace{\alpha_1\lambda_1}^{\lambda_A} - \overbrace{\alpha_2\lambda_2}^{\lambda_B} + \overbrace{\frac{1}{\alpha_3}\lambda_3}^{\lambda_E} = 0 \Rightarrow \Delta = \langle AB \rangle = \langle BE \rangle = \langle EA \rangle$$

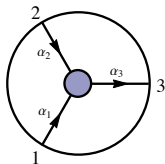
Ansatz:

$$\Omega_3^{\text{MHV}} = \kappa \frac{\Delta^\rho d\alpha_1 d\alpha_2 d\alpha_3}{\alpha_1^{\sigma_1} \alpha_2^{\sigma_2} \alpha_3^{\sigma_3}} \delta^{2 \times 2}(C \cdot \tilde{\lambda}) \delta^{1 \times 2}(C^\perp \cdot \lambda) \delta^{2 \times \mathcal{N}}(C \cdot \tilde{\eta})$$

- Impose: a) maximal SUSY $\mathcal{N} = 4s$ b) permutation invariance
 c) independence of α_3 ($\Omega \sim \frac{d\alpha_3}{\alpha_3}$)

$$\rho = s - 1, \quad \sigma_1 = \sigma_2 = \sigma_3 = 2s - 1$$

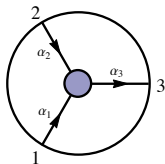
Grassmannian Formula for Gravity



$$\Omega_3^{\text{MHV},s} = \kappa \frac{\Delta^{s-1} d\alpha_1 d\alpha_2 d\alpha_3}{\alpha_1^{2s-1} \alpha_2^{2s-1} \alpha_3^{2s-1}} \overbrace{\delta^{2 \times 2}(C \cdot \tilde{\lambda}) \delta^{1 \times 2}(C^\perp \cdot \lambda) \delta^{2 \times 4s}(C \cdot \tilde{\eta})}^{=\delta(C \cdot Z)}$$

What does this mean for $s > 2$?

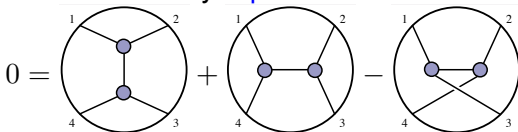
Grassmannian Formula for Gravity



$$\Omega_3^{\text{MHV},s} = \kappa \frac{\Delta^{s-1} d\alpha_1 d\alpha_2 d\alpha_3}{\alpha_1^{2s-1} \alpha_2^{2s-1} \alpha_3^{2s-1}} \overbrace{\delta^{2 \times 2}(C \cdot \tilde{\lambda}) \delta^{1 \times 2}(C^\perp \cdot \lambda) \delta^{2 \times 4s}(C \cdot \tilde{\eta})} = \delta(C \cdot Z)$$

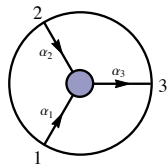
What does this mean for $s > 2$?

OS-diags in $\mathcal{N} = 4$ sYM satisfy **equivalence moves!**



$$\Gamma_s(\langle 12 \rangle \langle 34 \rangle)^{s-1} + \Gamma_t(\langle 14 \rangle \langle 23 \rangle)^{s-1} = \Gamma_u(\langle 13 \rangle \langle 24 \rangle)^{s-1}$$

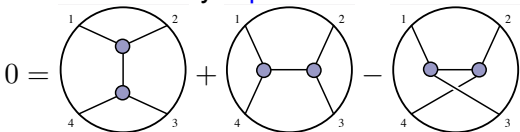
Grassmannian Formula for Gravity



$$\Omega_3^{\text{MHV},s} = \kappa \frac{\Delta^{s-1} d\alpha_1 d\alpha_2 d\alpha_3}{\alpha_1^{2s-1} \alpha_2^{2s-1} \alpha_3^{2s-1}} \overbrace{\delta^{2 \times 2}(C \cdot \tilde{\lambda}) \delta^{1 \times 2}(C^\perp \cdot \lambda) \delta^{2 \times 4s}(C \cdot \tilde{\eta})}^{=\delta(C \cdot Z)}$$

What does this mean for $s > 2$?

OS-diags in $\mathcal{N} = 4$ sYM satisfy **equivalence moves!**

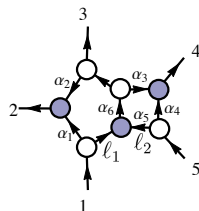


$$\Gamma_s(\langle 12 \rangle \langle 34 \rangle)^{s-1} + \Gamma_t(\langle 14 \rangle \langle 23 \rangle)^{s-1} = \Gamma_u(\langle 13 \rangle \langle 24 \rangle)^{s-1}$$

Two solutions:

- 1 $s = 1$: $\Gamma_s + \Gamma_t = \Gamma_u \Rightarrow$ **Jacobi identity**
 $\Gamma_s = f^{12a} f^{34a}$, $\Gamma_t = f^{14a} f^{23a}$, $\Gamma_u = f^{13a} f^{24a}$
- 2 $s = 2$: $\Gamma_s = \Gamma_t = \Gamma_u \Rightarrow$ **Shouten identity**
 universality of gravitational coupling

Grassmannian Formula for Gravity



$$\Omega_{\text{Gr}} = \frac{d\alpha_1}{\alpha_1^3} \cdots \frac{d\alpha_m}{\alpha_m^3} \left(\prod_v \Delta_v \right) \delta(C \cdot Z)$$

Features of the formula:

- higher poles present
- reduces to single pole if edge is erasable
- diagram vanishes if momenta in a given vertex become collinear

Can read off the λ 's and $\tilde{\lambda}$'s required for Δ similar to C -matrix!

Comments-Conclusion

- Extended Grassmannian formulation for On-Shell Diagrams to gravity
- Formulas for $\mathcal{N} < 8$ SUGRA only slightly modified
- presence of **poles at infinity**
- collinear vanishing of on-shell diagrams

Interesting other directions (not in this talk) [see Trnka talk at Amplitudes]

- relation to properties of **amplitudes**
- recursion relations
- UV-structure of gravity