

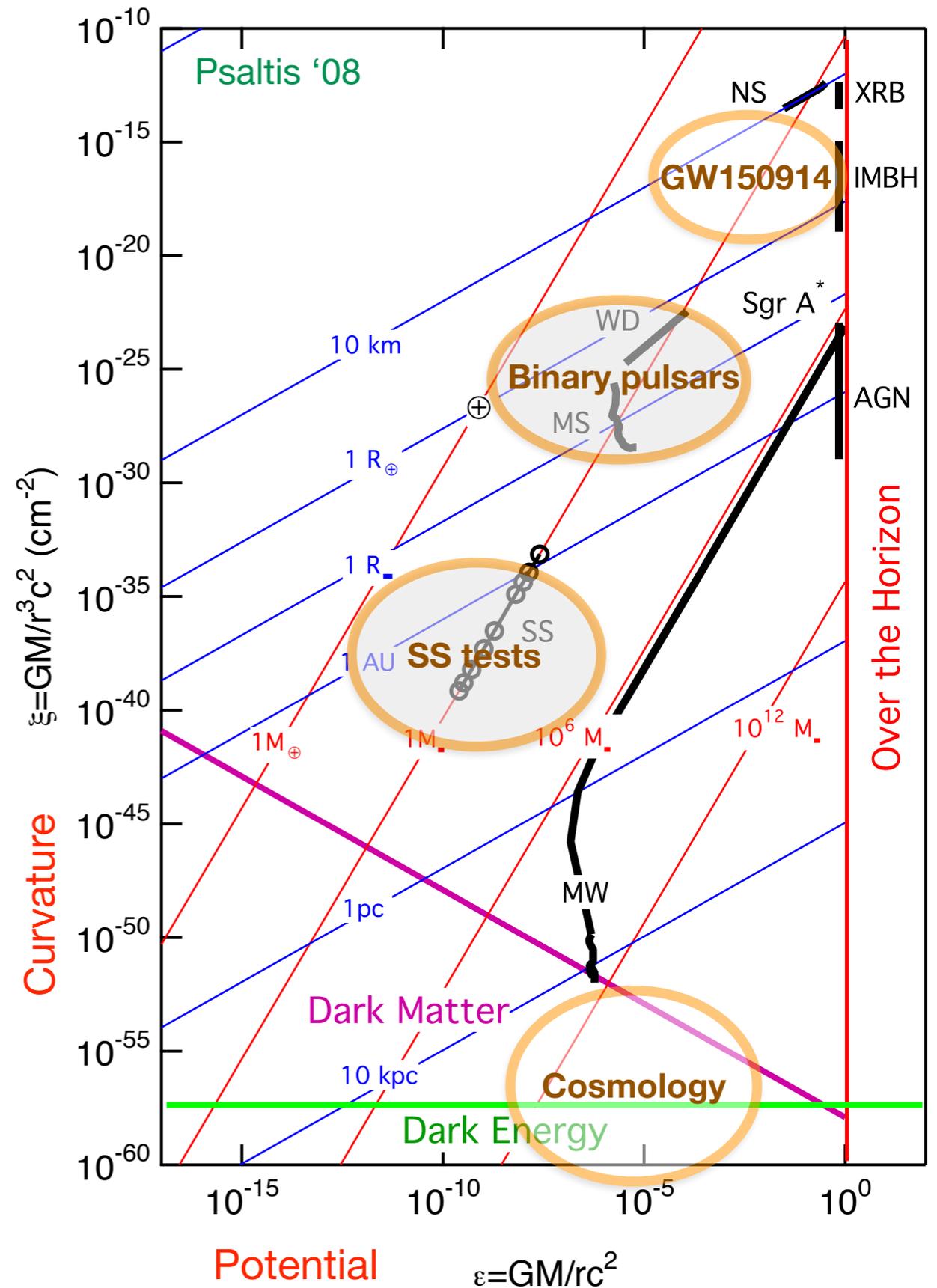
Effective Theory of Dark Energy

Filippo Vernizzi - IPhT, CEA Saclay

Advances in Theoretical Cosmology in Light of Data
Nordita, Stockholm - July 12, 2017

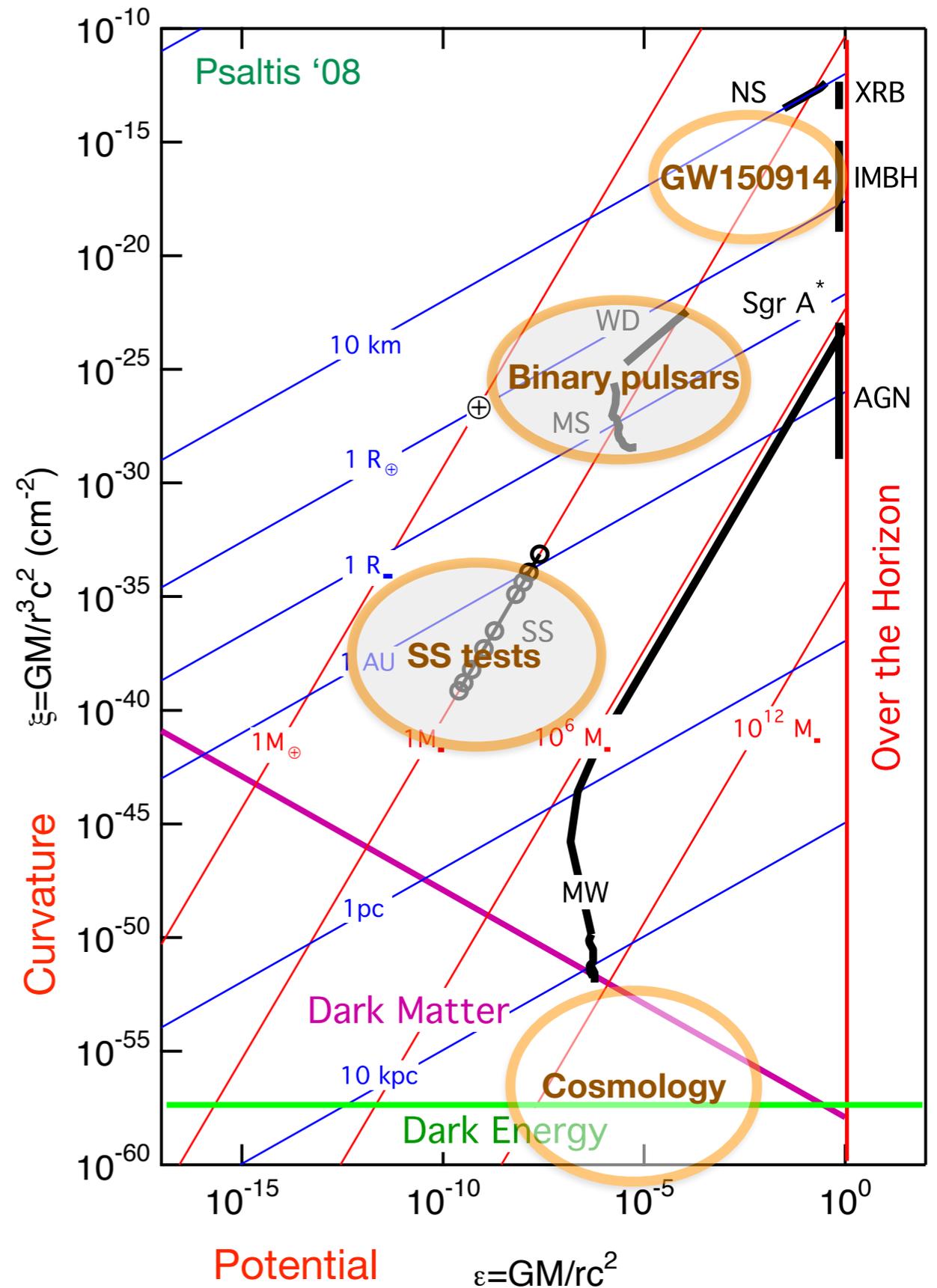
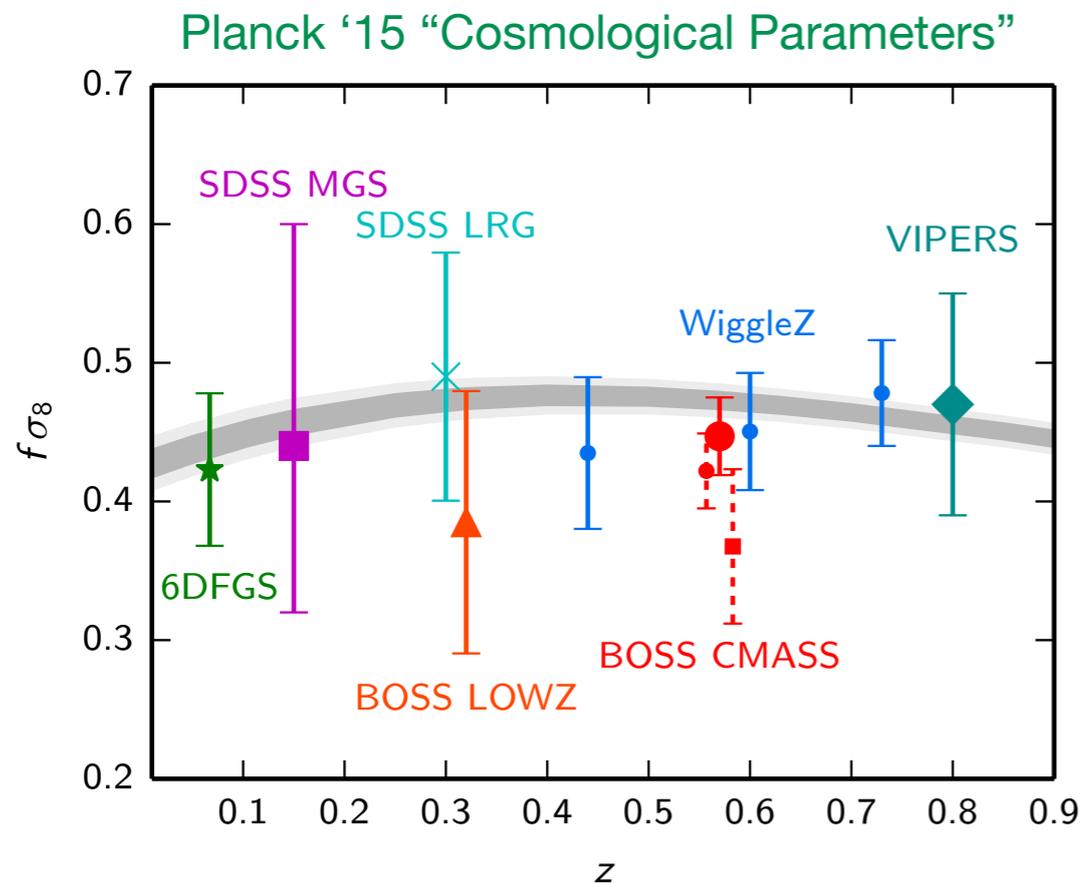
Motivations

- ◆ Gravity only been tested over special ranges of scales and masses
- ◆ Cosmology is a window for testing gravity on very large distances



Motivations

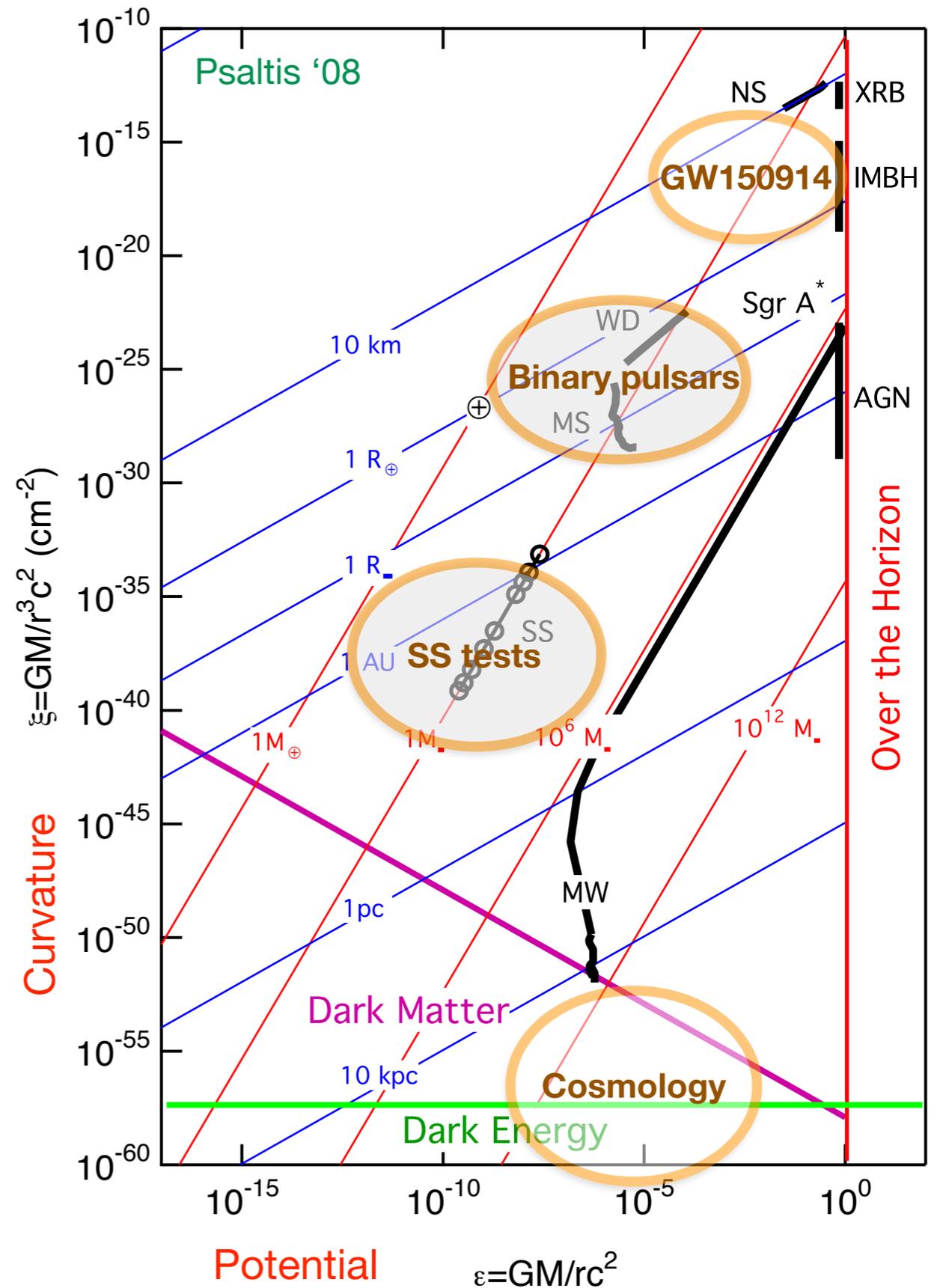
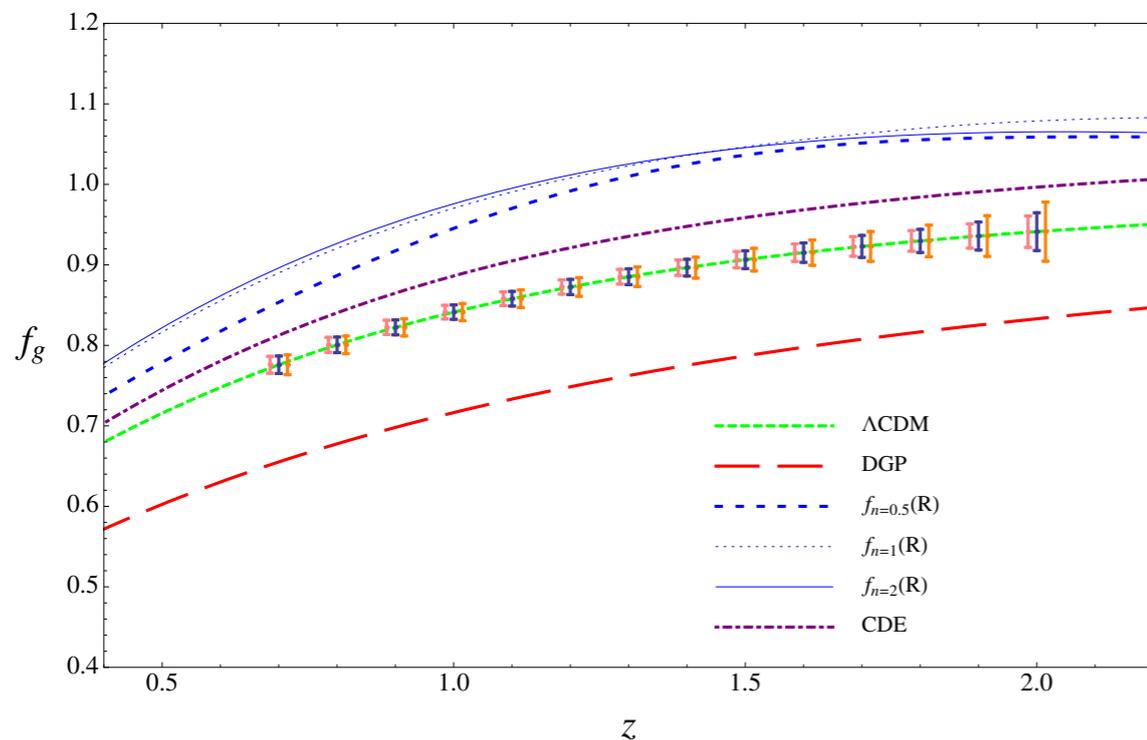
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- ◆ Standard model (GR): LCDM



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Euclid Theory Working Group '16



Tessa Baker

Einstein-Dilaton-Gauss-Bonnet

Cascading gravity

Lorentz violation
Hořava-Lifschitz

Conformal gravity

Strings & Branes

DGP

$$f\left(\frac{R}{\square}\right) \quad R_{\mu\nu} \square^{-1} R^{\mu\nu}$$

$$f(G)$$

Randall-Sundrum I & II

Some degravitation scenarios

Higher-order

Higher dimensions

Non-local

$$f(R)$$

General $R_{\mu\nu}R^{\mu\nu}$, $\square R$, etc.

Kaluza-Klein

Modified Gravity

Generalisations of S_{EH}

Vector

Einstein-Aether
Lorentz violation

TeV S — Add new field content

Massive gravity

Bigravity

Gauss-Bonnet

Lovelock gravity

Scalar-tensor & Brans-Dicke

Ghost condensates

Galileons

the Fab Four

Scalar

Chern-Simons

Cuscuton

Chaplygin gases

Tensor

EBI

Bimetric MOND

KGB

Coupled Quintessence

Horndeski theories

$f(T)$
Einstein-Cartan-Sciama-Kibble

Torsion theories

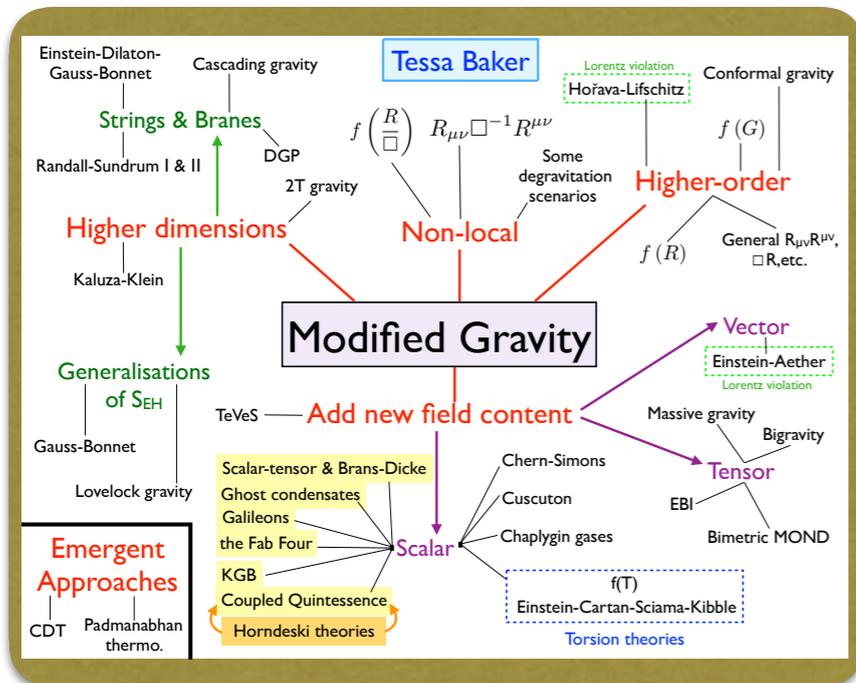
Emergent Approaches

CDT

Padmanabhan thermo.

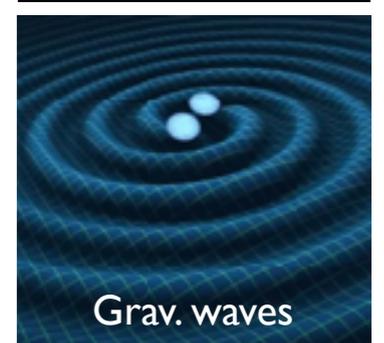
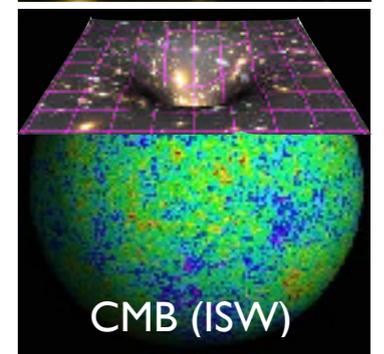
Motivations

Models



Many theoretical models of modified gravity, each with their own motivation and phenomenology

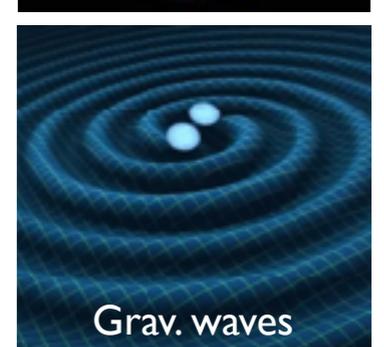
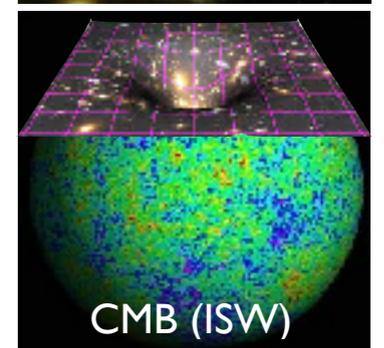
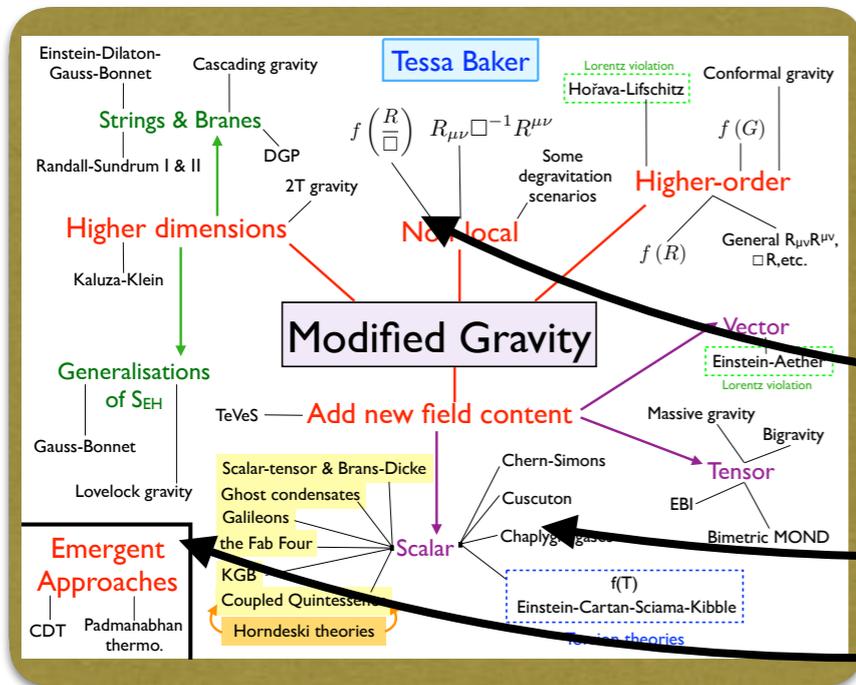
Observations



Motivations

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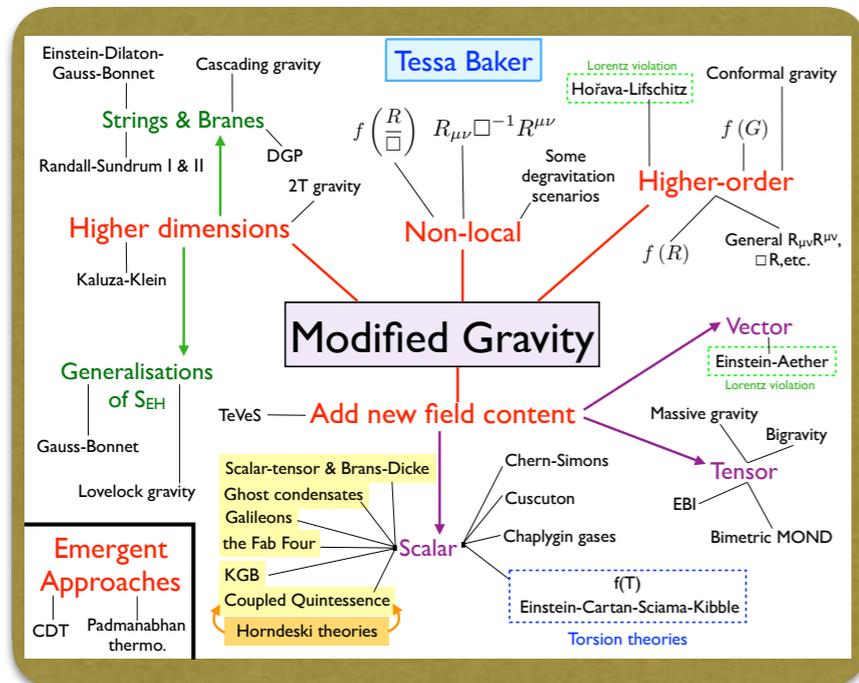
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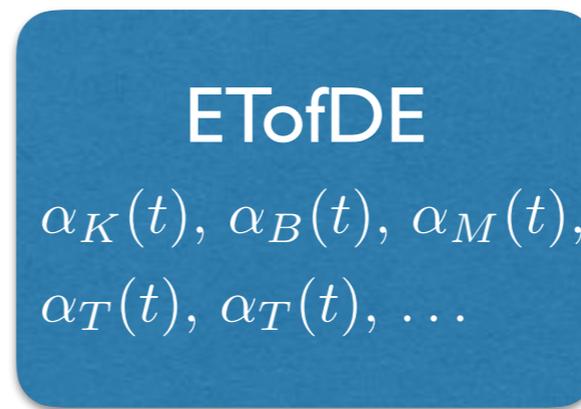
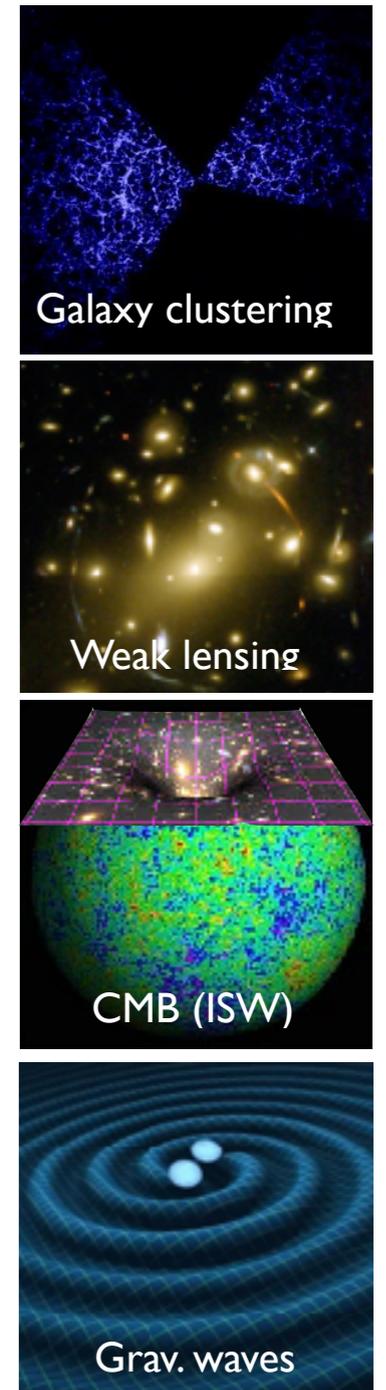
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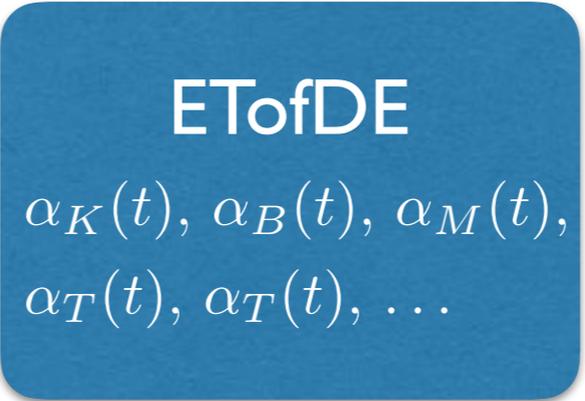
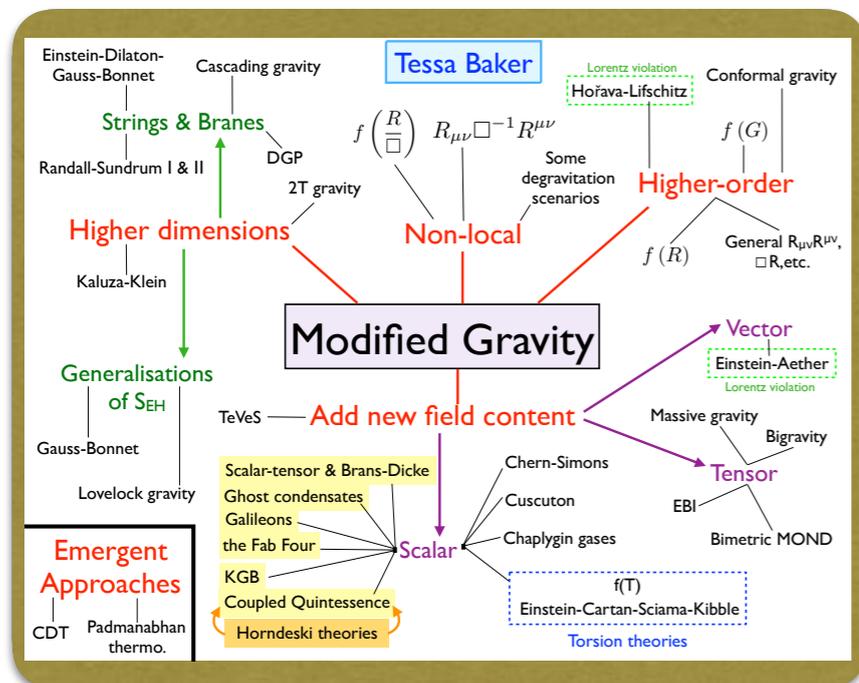
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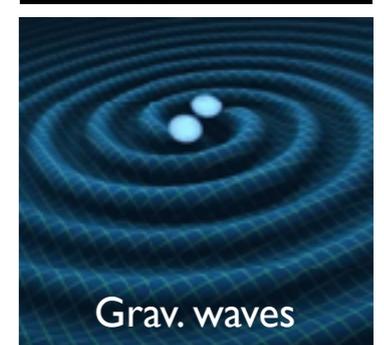
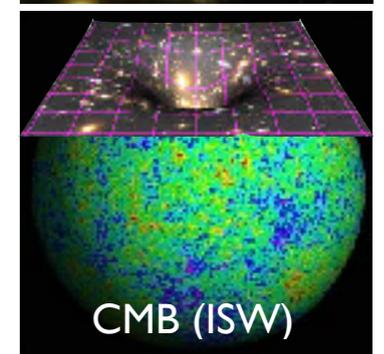
Effective approach to bridge models with observations in a minimal and systematic way

Motivations

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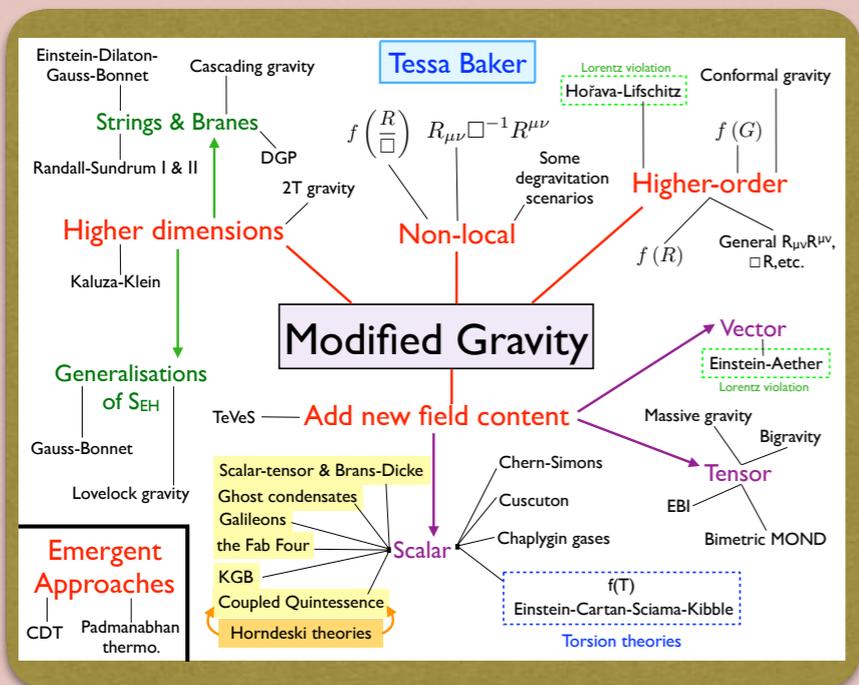
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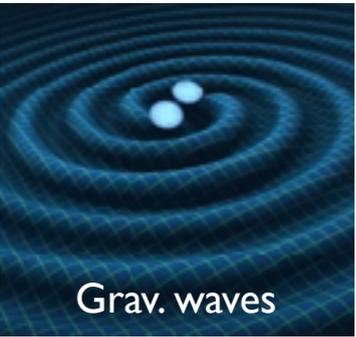
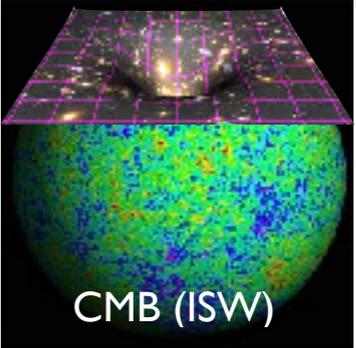
Theoretically motivated: locality, causality, diff invariance, unitarity, etc...

Models



ETofDE
 $\alpha_K(t), \alpha_B(t), \alpha_M(t),$
 $\alpha_T(t), \alpha_T(t), \dots$

Observations



Scalar-tensor theories

◆ Simplest models of modified gravity are based on a single scalar field

◆ Old school theories (Quintessence, Brans-Dicke, K-essence, ...) $\mathcal{L}(\phi, \partial_\mu \phi)$

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second-order
equations of motion

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- ◆ Horndeski theories:

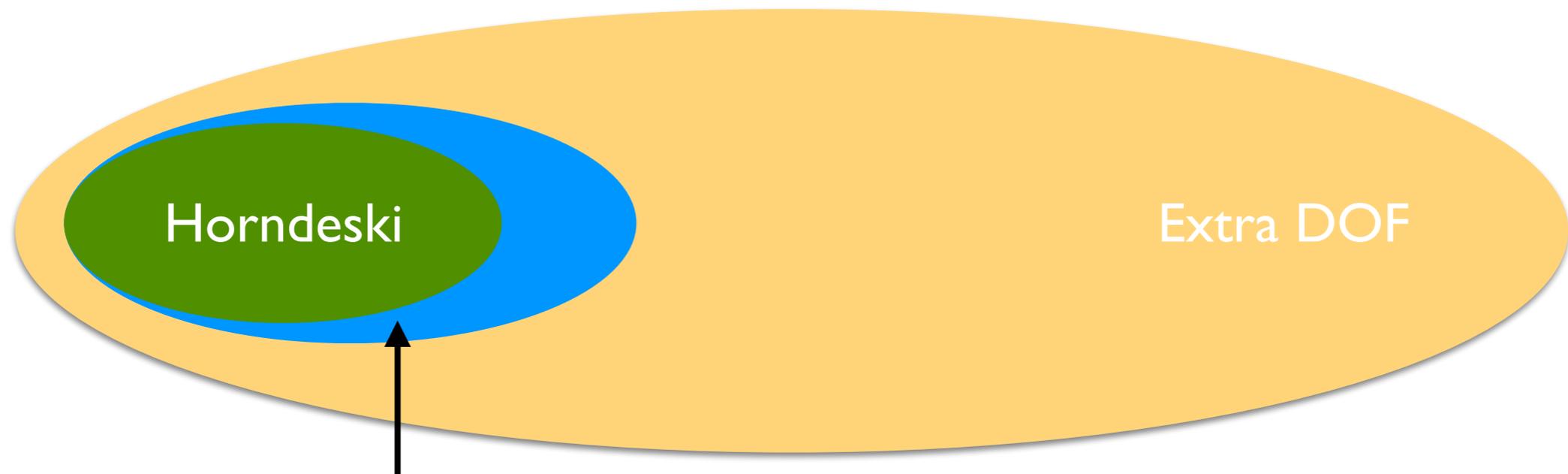
Horndeski '73, see also Deffayet et al. '11

$$X \equiv \phi_{;\mu} \phi^{;\mu} \equiv \nabla_\mu \phi \nabla^\mu \phi$$

$$\begin{aligned}
 L_H = & G_2(\phi, X) + G_3(\phi, X) \square \phi + \\
 & + G_4(\phi, X) {}^{(4)}R - 2G_{4,X}(\phi, X) [(\square \phi)^2 - \phi_{;\mu\nu} \phi^{;\mu\nu}] \\
 & + G_5(\phi, X) {}^{(4)}G^{\mu\nu} \phi_{;\mu\nu} + \frac{1}{3} G_{5,X}(\phi, X) [(\square \phi)^3 - 3\square \phi \phi_{;\mu\nu} \phi^{;\mu\nu} + 2\phi_{;\mu\nu} \phi^{;\nu\lambda} \phi_{;\lambda}^{;\mu}]
 \end{aligned}$$

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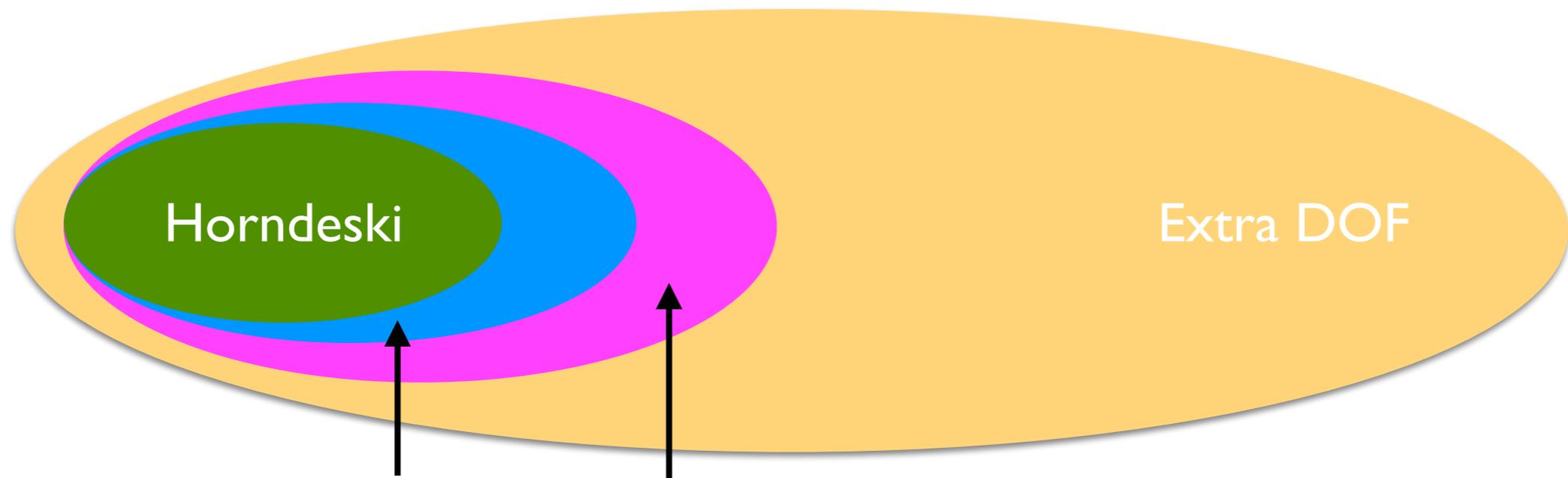


beyond Horndeski

Zumalacarregui, Garcia-Bellido '13
with Gleyzes, Langlois, Piazza '14;

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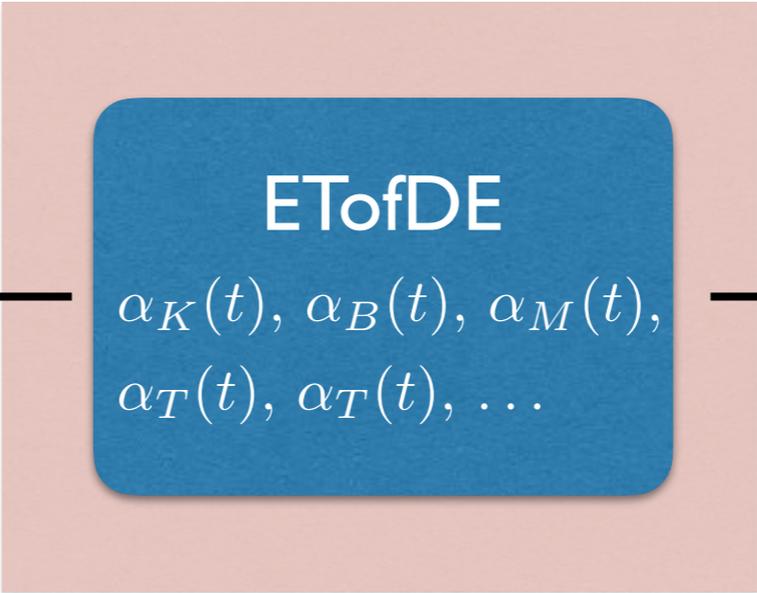
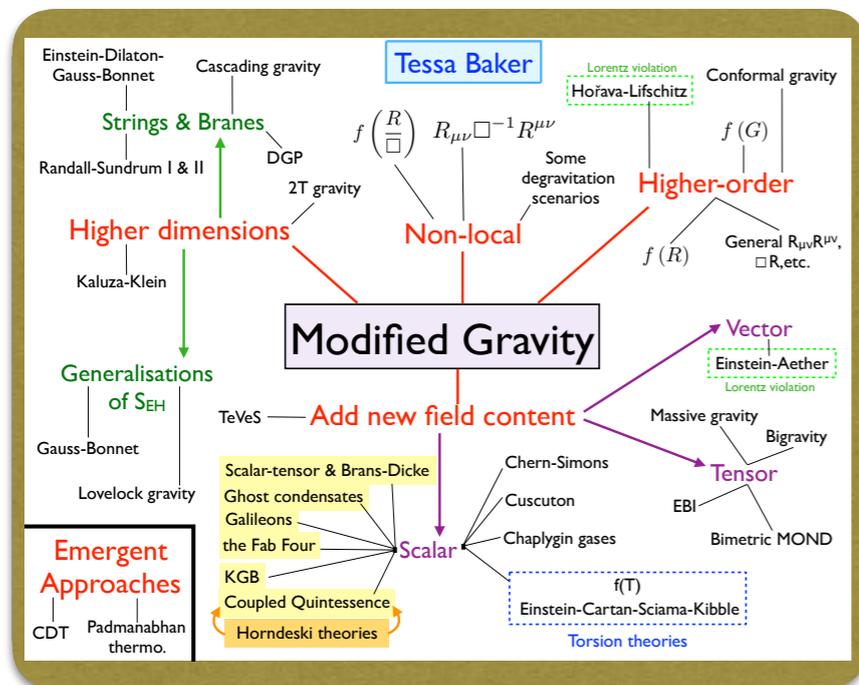
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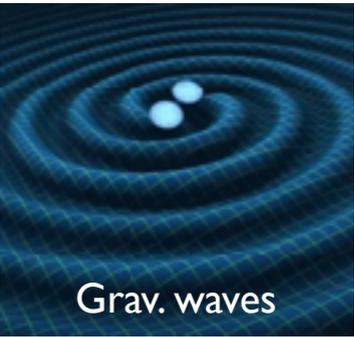
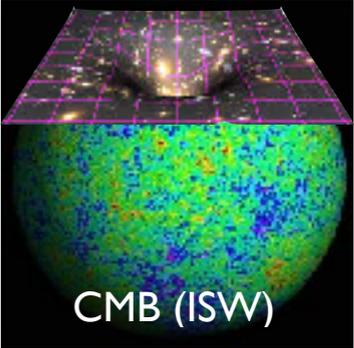
**Degenerate Higher-Order
Scalar-Tensor theories**

Langlois, Noui '15, '16;
Crisostomi, Hull et al. '16;
Crisostomi, Koyama, Tasinato '16;
Achour et al. '16

Models



Observations

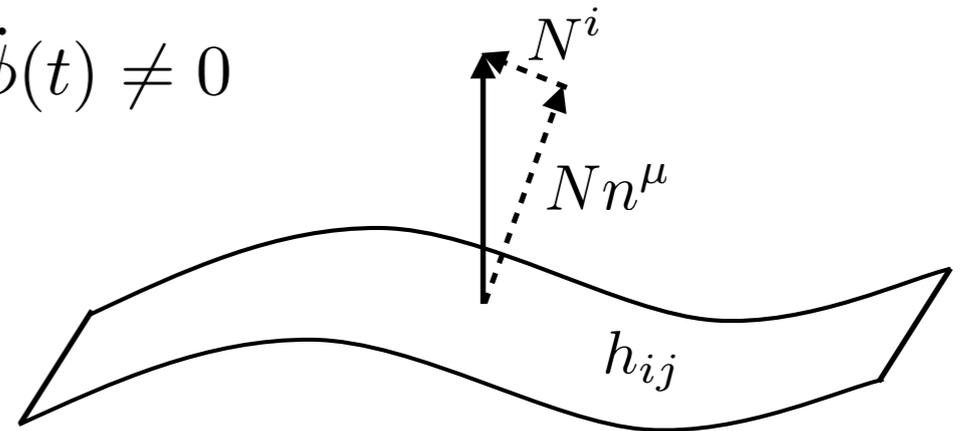


Constructing the action

- ◆ Use metric quantities in uniform scalar field slicing $\dot{\phi}(t) \neq 0$

▶ ADM decomposition

$$ds^2 = -N^2 dt^2 + h_{ij}(N^i dt + dx^i)(N^j dt + dx^j)$$

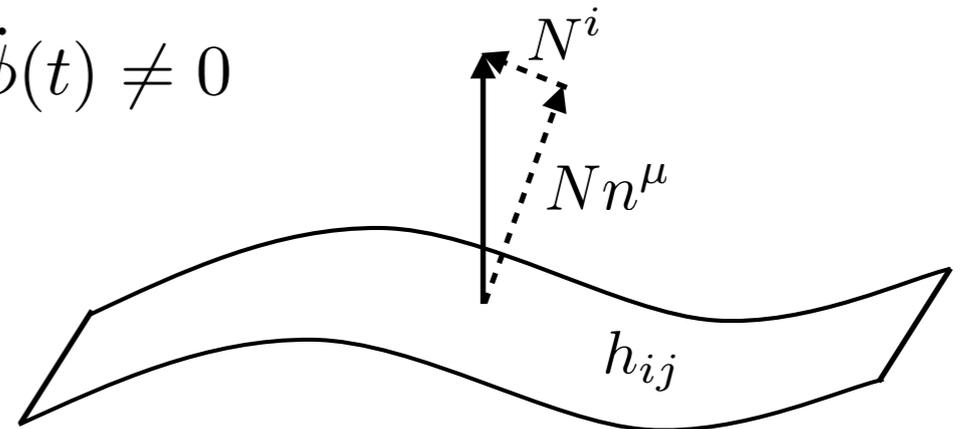


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- ◆ Lagrangian contains all possible scalars under spatial diffs, **ordered by number of perturbations and derivatives** Cheung et al. '07

$$S = \int d^4x \sqrt{-g} L[t; N, K_j^i, {}^{(3)}R_j^i, \dots]$$

▶ Lapse

$$N \sim \dot{\phi}$$

$$(\partial\phi)^2 = -\dot{\phi}_0^2(t)/N^2$$

▶ Extrinsic curvature

$$K_{ij} \sim \partial_t g_{ij}$$

$$K_{ij} = \frac{1}{2N} (\dot{h}_{ij} - \nabla_i N_j - \nabla_j N_i)$$

▶ Intrinsic curvature

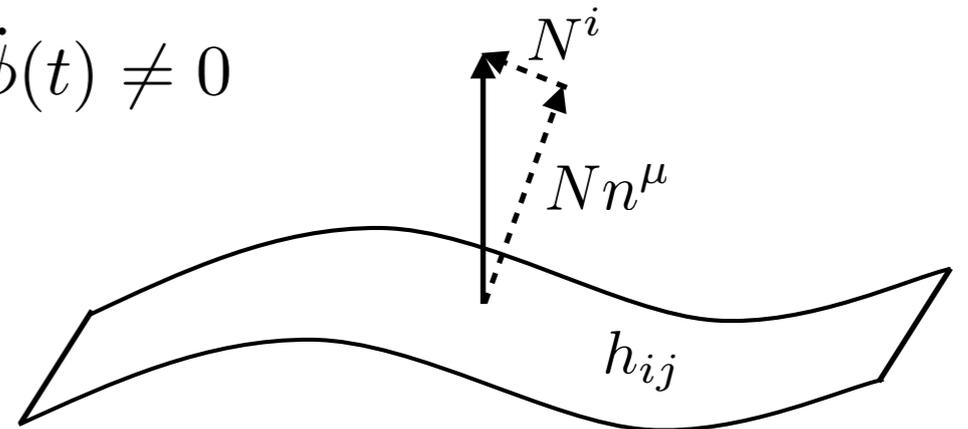
$${}^{(3)}R_{ij} \sim \partial^2 g_{ij}$$

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$$S = \int d^4x \sqrt{-g} L[t; N, K_j^i, {}^{(3)}R_j^i, \dots]$$

- ◆ Expand the action

$$\delta N \equiv N - 1, \quad \delta K_{ij} \equiv K_{ij} - H h_{ij}, \quad {}^{(3)}R_{ij}$$

$$L(N, K_j^i, R_j^i, \dots) = \bar{L} + L_N \delta N + \frac{\partial L}{\partial K_j^i} \delta K_j^i + \frac{\partial L}{\partial R_j^i} \delta R_j^i + L^{(2)} + \dots$$

Building blocks of linear perts

$$S^{(2)} = \int d^4x \sqrt{h} \frac{M^2}{2} \left[\delta K_i^j \delta K_j^i - \delta K^2 + {}^{(3)}R + \delta N {}^{(3)}R + \sum_i \alpha_i(t) \mathcal{O}_i^{(2)}(\delta N, \delta K, \dots) \right]$$

with Gleyzes, Langlois, Piazza '13 (see also Bloomfield '13)

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- ◆ Deviation from GR (LCDM) parameterized by time-dependent functions independent from background evolution

Notation of Bellini, Sawicki '14 for the alphas

5 functions of time instead of 5 functions of $\phi, (\partial\phi)^2$; minimal number of parameters

α_i	α_K	α_B	α_M	α_T	α_H
$\mathcal{O}_i^{(2)}$	δN^2	$\delta N \delta K$	$\frac{dM^2}{d \ln a}$	${}^{(3)}R$	$\delta N {}^{(3)}R$
quintessence, k-essence	✓				
Cubic Galileon	✓	✓			
Brans-Dicke, f(R)	✓	✓	✓		
Horndeski	✓	✓	✓	✓	
Beyond Horndeski	✓	✓	✓	✓	✓

Building blocks of linear perts

$$S^{(2)} = \int d^4x \sqrt{h} \frac{M^2}{2} \left[\delta K_i^j \delta K_j^i - \delta K^2 + {}^{(3)}R + \delta N {}^{(3)}R + \sum_i \alpha_i(t) \mathcal{O}_i^{(2)}(\delta N, \delta K, \dots) \right]$$

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- ◆ Deviation from GR (LCDM) parameterized by time-dependent functions independent from background evolution
- ◆ Locality, diff invariance preserved. Stability:

	Scalar	Tensor
No ghosts	$\alpha_K + 6\alpha_B^2$	$M^2 > 0$
No gradient instability	$c_s^2(\alpha_i) \geq 0$	$\alpha_T \geq -1$

Higher-Order theories

$$S^{(2)} = \int d^4x \sqrt{h} \frac{M^2}{2} \left[\delta K_i^j \delta K_j^i - \delta K^2 + {}^{(3)}R + \delta N {}^{(3)}R + \sum_i \alpha_i(t) \mathcal{O}_i^{(2)}(\delta N, \delta K, \dots) \right]$$

- ◆ All operators up to two derivatives

with Langlois, Mancarella, Noui '17

α_i	α_K	α_B	α_M	α_T	α_H	α_L	β_1	β_2	β_3
$\mathcal{O}_i^{(2)}$	δN^2	$\delta N \delta K$	$\frac{dM^2}{d \ln a}$	${}^{(3)}R$	$\delta N {}^{(3)}R$	δK^2	$\delta \dot{N}^2$	$\delta \dot{N} \delta K$	$(\partial_i \delta N)^2$

- ◆ Generic scalar dispersion relation: $\mathcal{E}_1 \omega^4 + \mathcal{E}_2 \omega^2 k^2 + \mathcal{E}_3 \omega^2 + \mathcal{E}_4 k^4 + \mathcal{E}_5 k^2 = 0$

- ◆ Two types of degeneracy conditions lead to $\omega^2 - c_s^2 k^2 = 0$

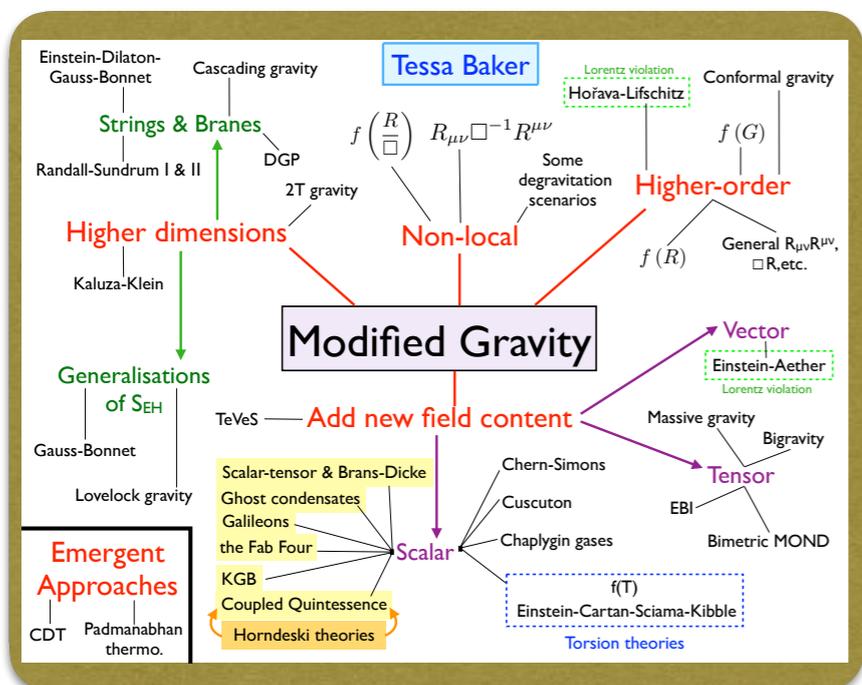
$$\mathcal{C}_I : \quad \alpha_L = 0, \quad \beta_2 = f_2(\beta_1), \quad \beta_3 = f_3(\beta_1)$$

related to Horndeski by metric redefinitions (that change the matter couplings)

$$\mathcal{C}_{II} : \quad \beta_1 = f_1(\alpha_T, \alpha_H, \alpha_L), \quad \beta_2 = f_2(\alpha_T, \alpha_H, \alpha_L), \quad \beta_3 = f_3(\alpha_T, \alpha_H, \alpha_L)$$

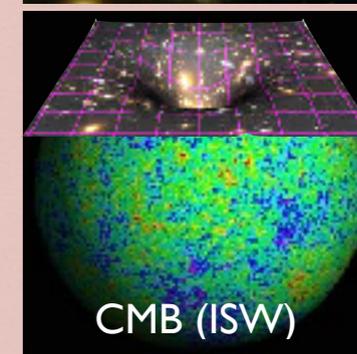
$$c_s^2 \propto -c_T^2 \quad \text{ruled out!}$$

Models



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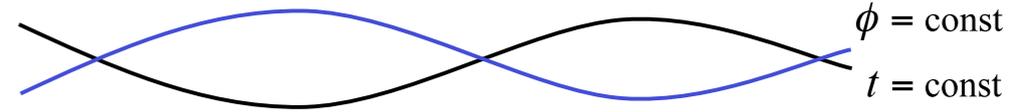
Observations



Phenomenology

- Undo unitary gauge:

$$t \rightarrow t + \pi(t, \vec{x})$$



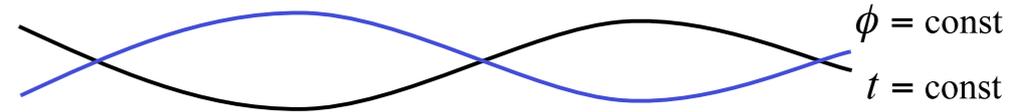
- Newtonian gauge (scalar flucts):

$$dt^2 = -(1 + 2\Phi)dt^2 + a^2(t)(1 - 2\Psi)d\vec{x}^2$$

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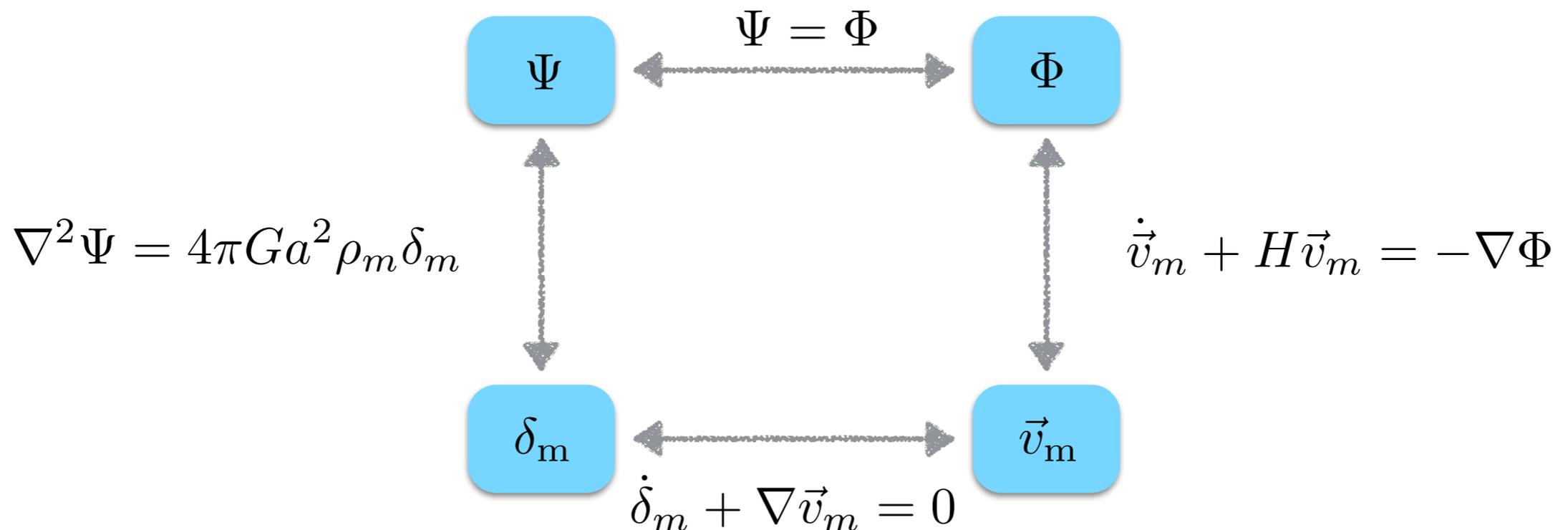
$$dt^2 = -(1 + 2\Phi)dt^2 + a^2(t)(1 - 2\Psi)d\vec{x}^2$$

- Quasi-static approximations — valid on scales $k \gg aHc_s^{-1}$.

Sawicki, Bellini '15

E.g., for surveys such as Euclid $c_s \gtrsim 0.1$.

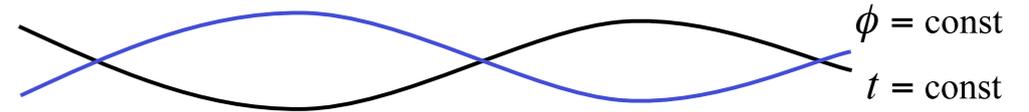
$$\nabla^2(\Psi + \Phi) = 8\pi G a^2 \rho_m \delta_m$$



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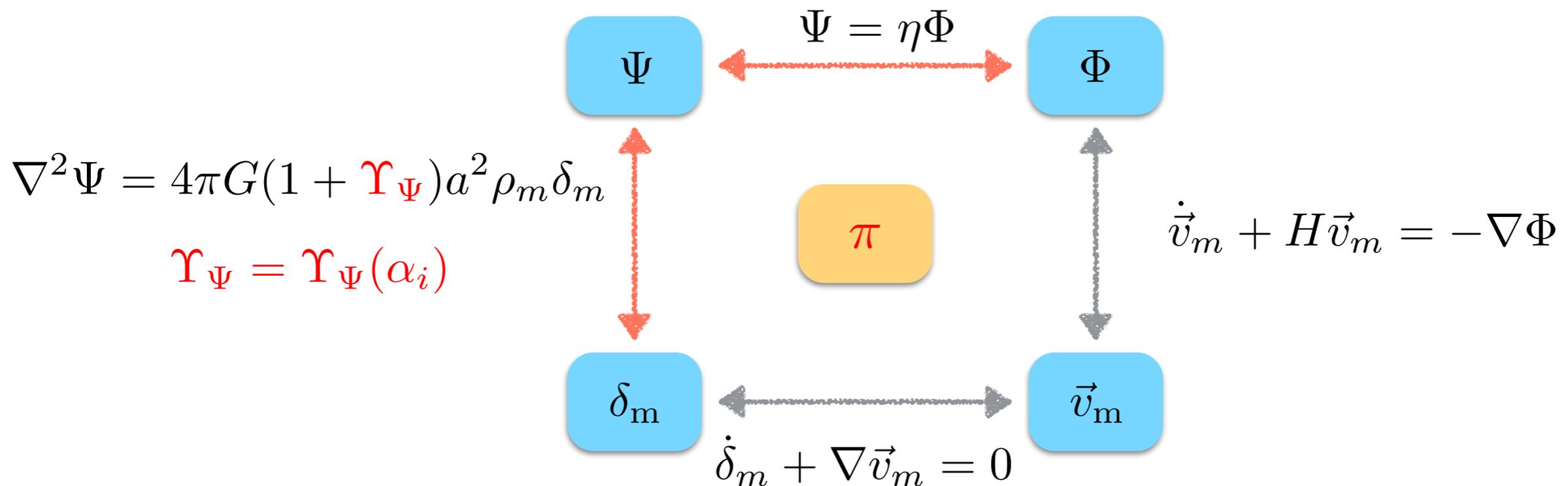
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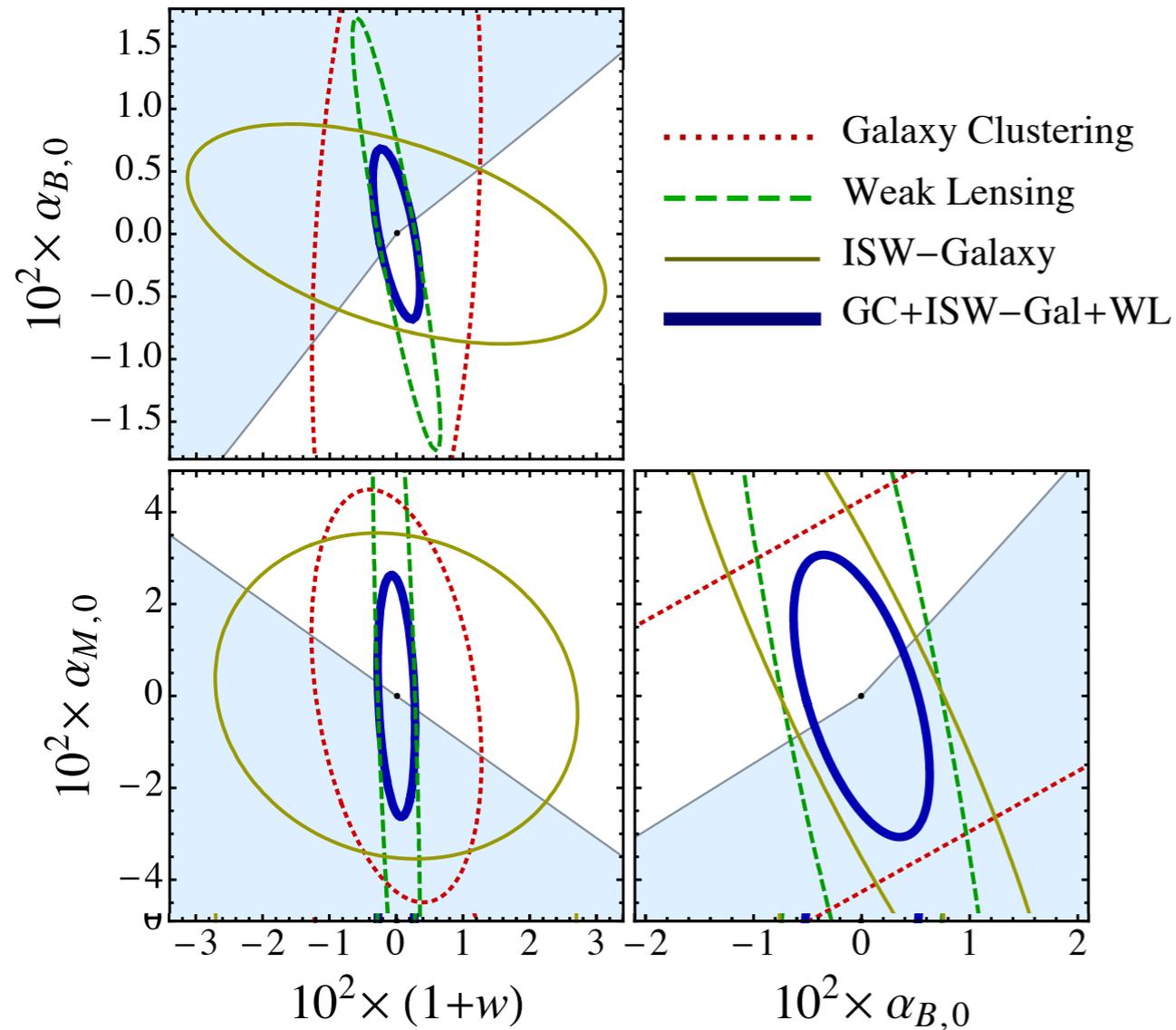
$$\nabla^2(\Psi + \Phi) = 8\pi G(1 + \Upsilon_{\text{lens}})a^2\rho_m\delta_m$$

$$\Upsilon_{\text{lens}} = \Upsilon_{\text{lens}}(\alpha_i)$$



Fisher matrix analysis

with Gleyzes, Langlois, Mancarella '15



Euclid specifications (LCDM fiducial)
Quasi-static approximation

◆ Background parametrization:

$$H^2 = H_0^2 \left[\Omega_{m0} a^{-3} + (1 - \Omega_{m0}) a^{-3(1+w)} \right]$$

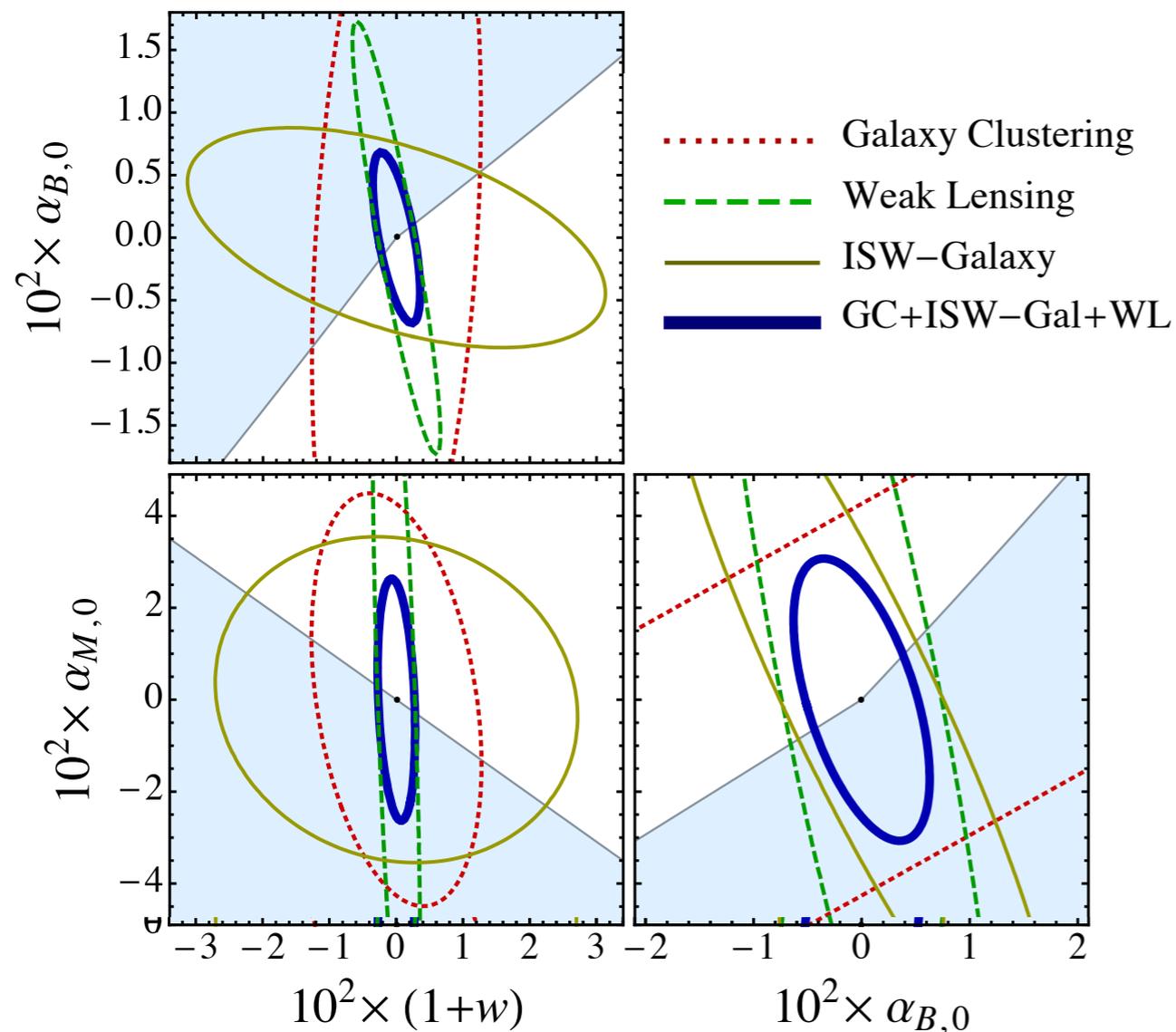
◆ Free functions parametrization:

$$\alpha_I(t) = \alpha_{I,0} \frac{1 - \Omega_m(t)}{1 - \Omega_{m,0}}$$

Piazza et al. '13, Bellini, Sawicki '14
see also Alonso et al. '16

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◆ Time dependence of free functions α 's still critical issue

see e.g. Linder '16, Gleyzes '17, Kennedy et al. '17

Boltzmann codes

◆ We want to go beyond the quasi-static approximation:

• Full Boltzmann solver: $\frac{df_I}{d\eta} = C_I[f_I]$, $I = \gamma, \nu, b, \text{CDM}$

$$\frac{\delta S^{(2)}}{\delta \pi} = 0 \quad \& \quad G_{ij}^{\text{modified}} = 8\pi G \sum_I T_{ij}^{(I)}$$

- EFTCAMB (Hu, Raveri, Frusciante, Silvestri et al.)
- hi_class (Zumalacarregui, Bellini, Sawicki, Lesgourgues et al.)
- COOP (Zhiqi Huang) (with D'Amico, Huang and Mancarella)

Boltzmann codes

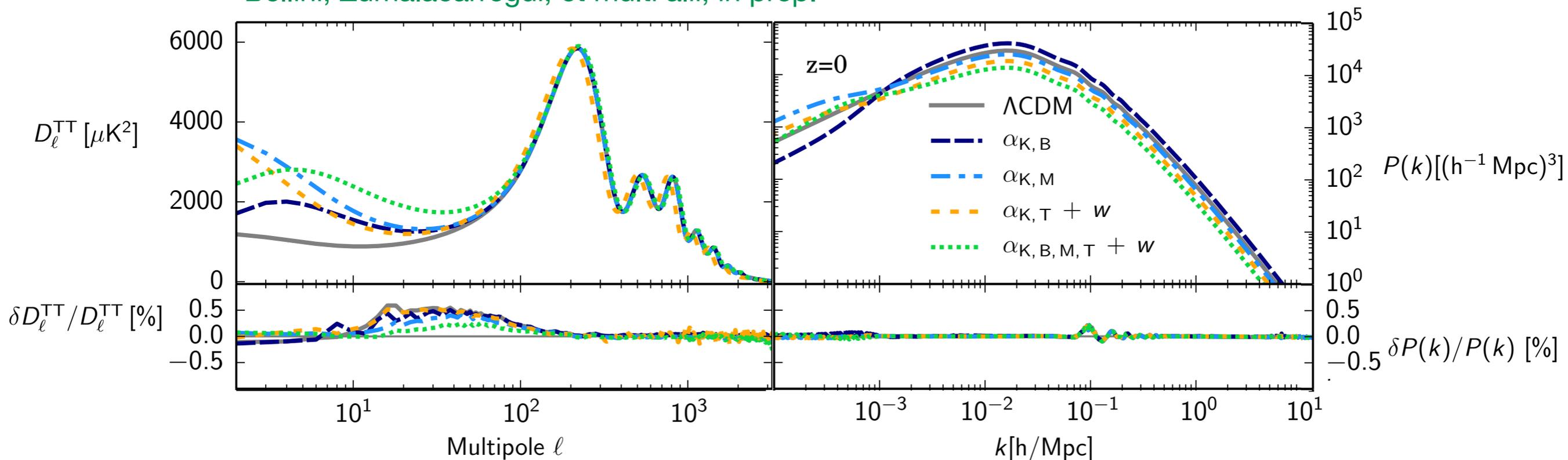
◆ We want to go beyond the quasi-static approximation:

• Full Boltzmann solver: $\frac{df_I}{d\eta} = C_I[f_I]$, $I = \gamma, \nu, b, \text{CDM}$

$$\frac{\delta S^{(2)}}{\delta \pi} = 0 \quad \& \quad G_{ij}^{\text{modified}} = 8\pi G \sum_I T_{ij}^{(I)}$$

- EFTCAMB (Hu, Raveri, Frusciante, Silvestri et al.)
- hi_class (Zumalacarregui, Bellini, Sawicki, Lesgourgues et al.)
- COOP (Zhiqi Huang) (with D'Amico, Huang and Mancarella)

Bellini, Zumalacarregui, et multi alii, in prep.



Nonlinear ET of DE

$$S = \int d^4x \sqrt{h} \frac{M^2}{2} \left[\delta K_i^j \delta K_j^i - \delta K^2 + {}^{(3)}R + \delta N {}^{(3)}R + \sum_i \alpha_i(t) \mathcal{O}_i^{(2)} + \sum_i \alpha_i(t) \mathcal{O}_i^{(3)} \right]$$

- ◆ In the Newtonian limit, a finite number of operators dominate

Example: Horndeski has **only 3 cubic operators and nothing more**

Bellini, Jimenez, Verde '15

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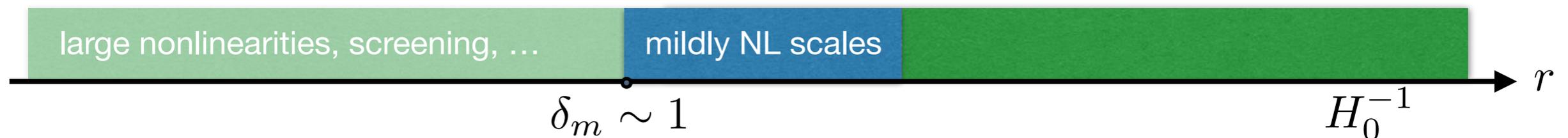
- ◆ Standard Perturbation Theory

$$\dot{\delta}_m + \nabla \cdot [(1 + \delta_m) \vec{v}_m] = 0$$

$$\dot{\vec{v}}_m + H \vec{v}_m + \vec{v}_m \cdot \nabla \vec{v}_m = -\nabla \Phi$$

- ◆ Modifications of gravity encoded in Poisson-like equation

$$k^2 \Phi = -\frac{3}{2} a^2 H^2 \Omega_m \mu_{\Phi,1} \delta_m - \frac{9}{4} a^2 H^2 \Omega_m^2 \mu_{\Phi,2}(\vec{k}_1, \vec{k}_2) \delta_m(\vec{k}_1) \star \delta_m(\vec{k}_2) + \dots$$



EFT of DE and LSS combined

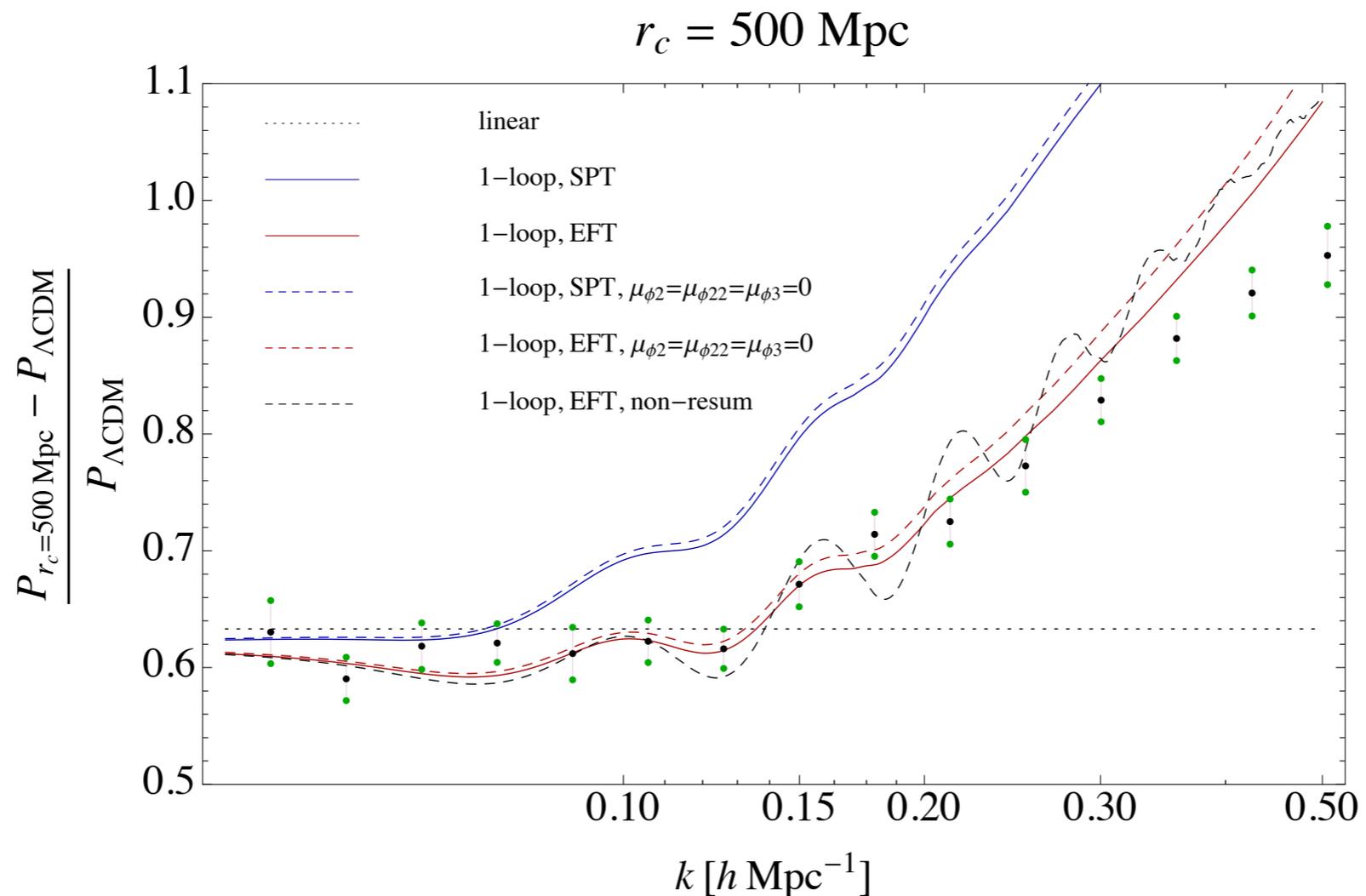


- ◆ Naturally incorporated into the Effective Field Theory of Large Scale Structures

$$\dot{v}_m^i + H v_m^i + v_m^j \nabla_j v_m^i + \nabla^i \Phi = \frac{1}{\rho_m} \partial_j \tau^{ji} \sim \frac{c_{m,s}^2}{k_{\text{NL}}^2} \nabla^i \delta_m$$

Baumann et al. '10,
Senatore et al. '12, etc...

- ◆ Example: **1-loop Power Spectrum with IR resummation**. Comparison with Fabian Schmidt '09 simulations of **nDGP**, 3 realizations, from 400 to 128 Mpc/h box size:



Conclusions

- * Unifying description for scalar-tensor theories, including higher-order ones (and more)
- * Analysis of (degenerate higher-order) theories highly simplified
- * Linear regime worked out! Issue of time dependence of α 's
- * Straightforward connection to mildly nonlinear and fully nonlinear regime

