Is "single component" enough for the thermal dust emission?

Hao Liu, Sebastian von Hausegger, Pavel Nashelsky For Nordita 2017 in Stockholm



CMB



A good foreground estimation is required to get a good CMB detection

Some estimations of the thermal dust emission

- Planck 2013
 - Planck 2013 results. XI. All-sky model of thermal dust emission, 2014 A&A 571, A11
- Planck 2015
 - Planck 2015 results. X. Diffuse component separation: Foreground maps, 2016, A&A, 594, A10
- GNILC
 - Planck intermediate results. XLVIII. Disentangling Galactic dust emission and cosmic infrared background anisotropies, 2016, A&A 596, A109
- Different estimations, but same assumption: **single component** thermal dust
 - Example of multi-component assumption: "Foreground Bias From Parametric Models of Far-IR Dust Emission", 2016, A. Kogut, D.J. Fixsen.

- Line of sight → We know for sure that the real thermal dust emission is **not** single component.
- Therefore, a meaningful discussion should be:
 - Is "single component" good enough as an approximation?
 - Especially for extrapolation down to ~100 GHz
- Simplification of the discussion
 - Multi dust component?
 - Multi other component?
 - Variation of the parameter as function of frequency?
 - Many other possibilities, but we discuss only

What happens if we don't answer this question

- Real case is multi, but assume single?
 - For a few high freq. bands: still good fit
 - But extrapolation will not be reliable
- Real case is single, but assume multi?
 - Can fit even the noise, systematics...
 - Over fitting
- Therefore, we should have a clear answer to "Is single component good enough"?
 - If good enough: Fine! we can continue to use the current estimations
 - If not good enough: Multi-frequency experiments will be preferred.

The idea

- Use a local region from two adjacent bands to derive the ratio • of dust amplitude between them – R: ratio by data. R0: ratio by existing dust model.
- Use only the regions with very high cross-correlation coefficients between two adjacent bands to cast limits on the error R-R0



857 GHz



(40.0, 0.0) Galact



Let R be the ratio given by linear regression, and R_0 be the expected ratio, then we have:

$$R_{0} = \frac{S_{xy}}{S_{x}}$$
(4)

$$R = \frac{S_{x'y'}}{S_{x'}}$$

$$C(x, y) = -\frac{S_{xy}}{\sigma_{x}\sigma_{y}}$$



$$S_{x'} - S_x(1 + 2k_x\beta_x + \beta_x^2)$$

$$S_{y'} = S_y(1 + 2k_y\beta_y + \beta_y^2)$$

$$S_{x'y'} = S_{xy}(1 + k_x\beta_x + k_y\beta_y + \beta_x\beta_y)$$

We get:

$$1 \approx \frac{1 + k_x \beta_x + k_y \beta_y + \beta_x \beta_y k_\Delta}{\sqrt{(1 + 2k_x \beta_x + \beta_x^2)(1 + 2k_y \beta_y + \beta_y^2)}}$$
$$R = \frac{S_{x'y'}}{S_{x'}} \approx \frac{\sigma_y'}{\sigma_x'} = \frac{(1 + 2k_y \beta_y + \beta_y^2)}{(1 + 2k_x \beta_x + \beta_x^2)} R_0$$

By Taylor expansion, we see that 1-R/R0 and 1-C(x',y') are same level small numbers \rightarrow This is the limitation we need!





Using r=10-deg disc as patches, CC is at least 0.95

Note that the CMB map (SMICA) is excluded in advance



FIG. 4: Similar to Fig. 3 but use only the $|b| > 30^{\circ}$ region. There is no essential difference to Fig.3.



FIG. 5: Similar to Fig. 3, but with 5° smoothing to the input maps and use only the $|b| > 30^{\circ}$ region. There is no essential difference to Fig.3.





FIG. 6: Similar to Fig. 3, but with 5° smoothing to the input maps and use the polarization data. Note that there is no polarization data for 545 and 857 GHz.

Is the deviation due to...

- Color correction
- CIB
- Free-free
- Zodiac light
- Systematic + Residual CMB
- Due to the limitation given by "High band-band correlation":
 - Either unaffected by these issues
 - Or these issues must provide strong correlation to thermal dust

Current conclusion

- "single component" is likely not a good assumption for the thermal dust emission
- Especially for the middle frequencies (~100 GHz)
- Measurement of more frequency bands will be great for a reliable estimation of the termal dust emission.

Thanks!