

The role of stability conditions in single field Quintessence

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based on: Peirone, MM, Raveri, Silvestri
arXiv:1702.06526

Outline

- 1 Tensions in data and deviations from Λ CDM
- 2 Theoretical conditions
- 3 Effect on data analysis
- 4 Conclusions

Overview

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2 Theoretical conditions

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Tensions between data

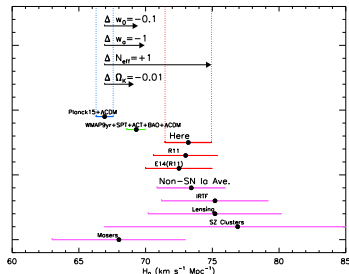
Recent improvement in observation allows to constrain with high precision our cosmological model.

Moreover, tensions between low and high redshift measurements were found

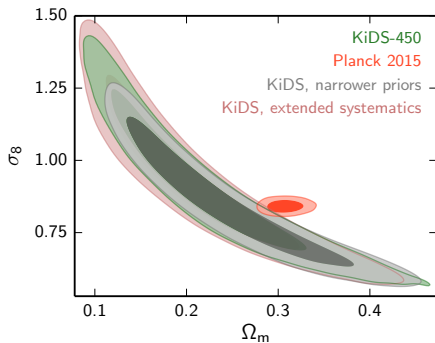
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Riess et al. 2016



Joudaki et al. 2016

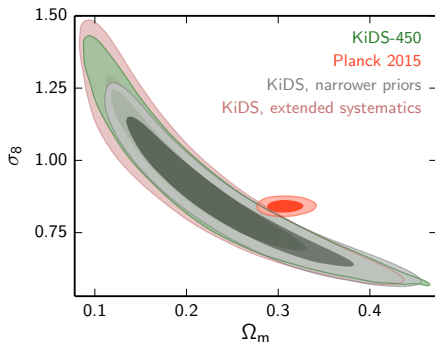
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$$T(S_8) = \frac{|S_8(CMB) - S_8(WL)|}{\sqrt{\sigma_{S_8}^2(CMB) + \sigma_{S_8}^2(WL)}} = 2.1\sigma$$

$$S_8 = \sigma_8 \sqrt{\frac{\Omega_m}{0.3}}$$



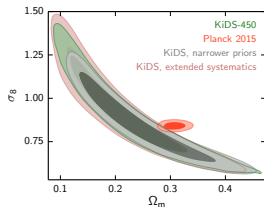
Joudaki et al. 2016

Systematics or extensions to Λ CDM

One can blame systematics or think this is an hint for new/non-standard physics

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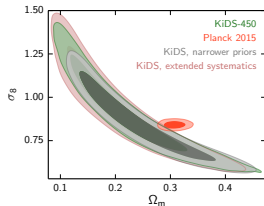
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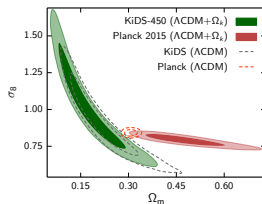
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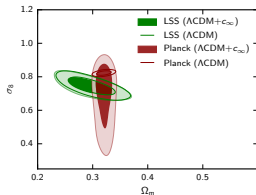
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Camera, MM, Bertacca 2016

Dynamical Dark Energy to the rescue

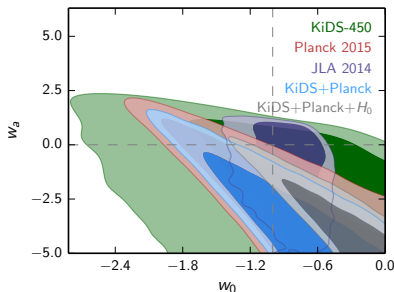
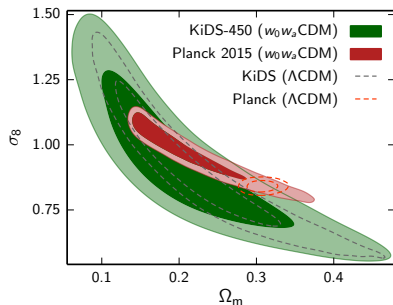
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Dynamical Dark Energy to the rescue

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$$T(S_8) = 0.91\sigma \text{ with } \Delta DIC = -6.4$$

Joudaki et al. 2016

Underlying assumption

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Probes deviations from Λ CDM without assumptions on any underlying theoretical model.

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The CPL $w(z)$ crosses the phantom divide ($w = -1$), where equations for DE perturbations become singular.

Parametrized Post-Friedmann (PPF) prescription, which assumes smooth Dark Energy field(s), stabilizes the model.

Fang, Hu, Lewis 2008

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We use the EFT approach to DE/MG (see Filippo's talk).

These requirements give conditions on the EFT functions which translates to limitations on the parameter space.

Stability conditions are implemented in EFTCAMB

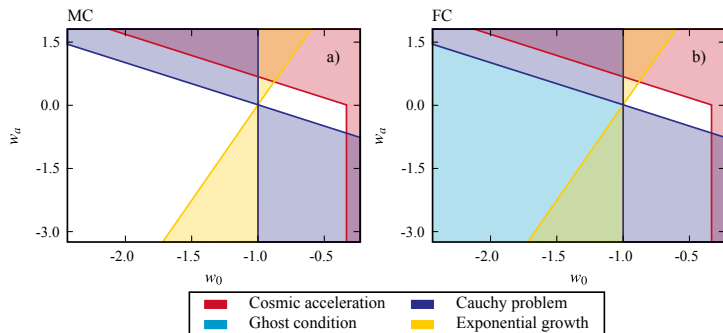
Hu, Raveri, Frusciante, Silvestri 2014 and 2015

Single field Quintessence

We specify to a class of models: minimally coupled single field Quintessence.
We assume the CPL parametrizes the EoS of this class.
Our requirements limit the $(w_0 - w_a)$ parameter space.

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Peirone et al. 2017

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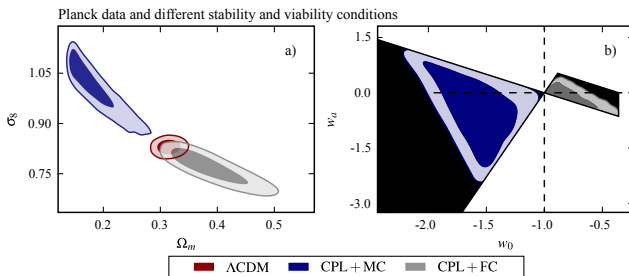
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Theoretical priors

These conditions are not imposed a posteriori on the analysis, but rather included in the prior probability $P(\theta)$.

$$P(\theta|D) \propto L(D|\theta)P(\theta)$$

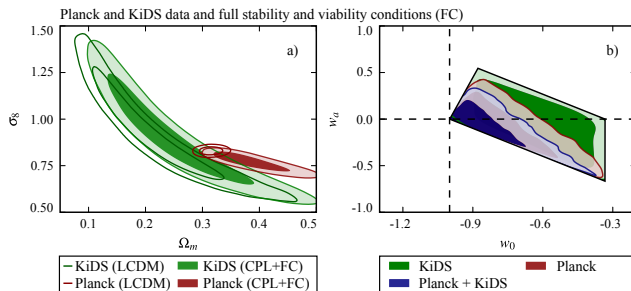


Tension and single field quintessence

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$$T(S_8) = 1.3\sigma \text{ with } \Delta DIC = 4.6$$

Assuming single field Quintessence, Λ CDM is again the better model

Peirone et al. 2017

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- Tensions between cosmological datasets can be seen as hints of failure of Λ CDM and used to test for extended models
- Phenomenological parameterizations allow to probe departure from standard paradigm, without assuming any specific model.
- if one wants to stick to physically viable theories, theoretical priors should be included in the analysis
- Simple example: imposing $w(a)$ comes from single field quintessence, hints for departures from Λ CDM are removed.

We knew already that in quintessence $w(a) \geq -1$... Is this useful?

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- for more complex theories we don't have a simple intuition... theoretical conditions can avoid exploring unphysical parts of parameter space
- conditions can be generalized to broader class of theories and to other theory health requirements
(see *De Felice, Frusciante, Papadomanolakis 2016*)

SAFETY SLIDES

Quintessence in the EFT formalism

Starting from the EFT action

$$S \propto \int d^4x \sqrt{-g} \left\{ \frac{m_0^2}{2} [1 + \Omega(\tau)] R + \Lambda(\tau) - a^2 c(\tau) \delta g^{00} + \dots \right\}$$

Assuming a specific $w(z)$ completely specifies the $\Lambda(\tau)$ and $c(\tau)$ functions.

$$\frac{c(\tau) a^2}{m_0^2} = \frac{1}{2} \frac{a^2 \rho_{DE}}{m_0^2} (1 + w_{DE})$$

$$\frac{\Lambda(\tau) a^2}{m_0^2} = \frac{a^2 \rho_{DE}}{m_0^2} w_{DE}$$

The assumption of a minimally coupling quintessence sets $\Omega(\tau) = 0$ (together with the perturbations operators).

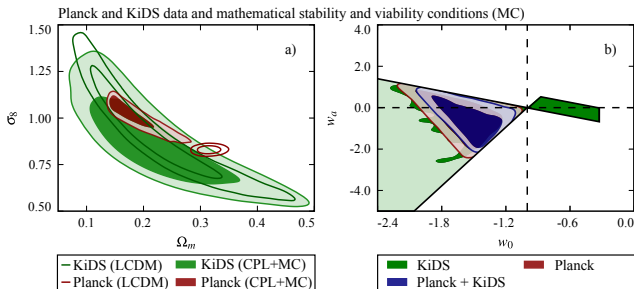
Mathematical and Physical conditions translate on constraints on the 2 free functions and, consequently, on w_0 and w_a .

Hu, Raveri, Frusciante, Silvestri 2014 and 2015

Removing the physical conditions

Without the PC, we are implicitly assuming that ghost instabilities do not develop on the time scale of interest.

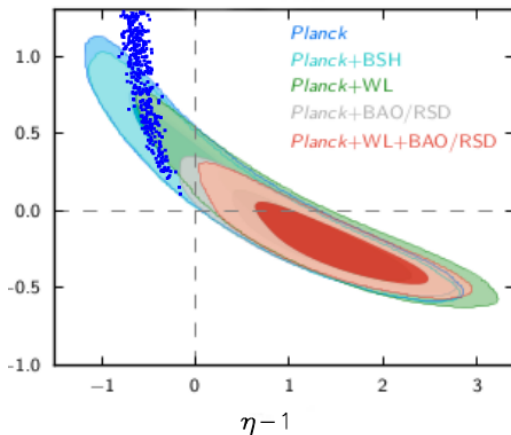
This scenario is approximately corresponding to the PPF case



Peirone et al. 2017

Conditions on Modified Gravity

Stability requirements applies to all kinds of phenomenological parameterizations



Perenon et al. 2015

Deviance Information Criterion

The DIC is used as a model comparison tool.

$$DIC \equiv \chi_{\text{BF}}^2 + 2p_D$$

where p_D is a term accounting for the complexity of the model

$$p_D = \bar{\chi}^2 - \chi_{\text{BF}}^2$$

The results show $\Delta DIC = DIC^{\text{ext}} - DIC^{\Lambda\text{CDM}}$, thus $\Delta DIC < 0$ highlights a preference of the data for the extended model.

A better estimate would be given by the Bayesian Evidence, but it's generally complicated to compute, specially for a non trivial prior volume.