

Hydrodynamic Modes of Incoherent Black Holes

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Based on work in collaboration with A. Donos, J. Gauntlett
[arXiv: 1708.05412, 1710.xxxxx]

October 4, 2017



Transport in inhomogeneous media

- Conductivity matrix characterizes linear response of system to external sources.
- Apply *constant, time-dependent* electric field E_j and thermal gradient $\zeta_j \leftrightarrow (\nabla_j T) / T$ and read off $U(1)$ and heat currents:

$$\begin{pmatrix} J^i \\ Q^i \end{pmatrix} = \begin{pmatrix} \sigma^{ij} & T\alpha^{ij} \\ T\bar{\alpha}^{ij} & T\bar{\kappa}^{ij} \end{pmatrix} \begin{pmatrix} E_j \\ \zeta_j \end{pmatrix}$$

- Translation invariance \Rightarrow divergent DC conductivities $\sim (\delta(\omega) + \frac{i}{\omega})$ as $\omega \rightarrow 0$.
- For physical, finite conductivities, we need a mechanism for momentum to dissipate \rightarrow Explicit breaking of translational invariance by spatially dependent sources.
- Spontaneous breaking of translational invariance is also physically relevant (e.g. charge density waves).

Breaking translational invariance in holography

Explicitly, by sources at infinity (“holographic lattice”).¹

- Analytic results for perturbative lattices, (hard) numerical constructions of strong lattices.
- Easier if we retain homogeneity of metric but break translations in the scalar sector (linear axions,² Q-lattices³).
- Hydrodynamics dominated by energy and charge diffusion.

¹[Horowitz, Santos, Tong '12, ...]

²[Andrade, Withers '14]

³[Donos, Gauntlett '14]

Breaking translational invariance in holography

Spontaneously, if there is an instability towards a broken phase with spatial modulation.⁴

- The boundary is homogeneous but the bulk (and the horizon) is *not*.
- Numerical constructions in Einstein-Maxwell-Dilaton theories, possibly with Chern-Simons(-like) terms (holographic charge density waves,⁵ helical black holes⁶).
- Energy and charge diffusive modes + sound modes corresponding to the Goldstone modes +

⁴[Nakamura, Ooguri, Park '10, . . .]

⁵[Donos, Gauntlett '13]

⁶[Donos, Gauntlett '12]

Approach

- Study diffusive modes in a framework accounting for both cases of breaking translations
- Construct a “thermodynamic” zero mode perturbation
- Construct the diffusive modes in long wavelength perturbative expansion
- Obtain dispersion relations by focusing on the near horizon region → no need for explicit bulk solutions
- For holographic lattices, connect with DC conductivities → obtain generalised form of Einstein relations

Outline

Introduction & Motivation

Bulk Diffusive Mode

- Setup and horizon constraints

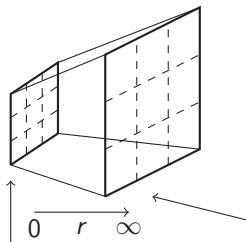
- Constructing the diffusive mode

- Einstein relations

Comments & Outlook

Holographic setup

$$\text{Bulk action: } S = \int d^D x \sqrt{-g} \left(R - V(\phi) - \frac{Z(\phi)}{4} F^2 - \frac{1}{2} (\partial\phi)^2 \right)$$



Regular horizon

$$g_{ij} = g_{ij}^{(0)}(x) + \dots$$

$$U = 4\pi T r + \dots$$

$$a_t = r a_t^{(0)}(x) + \dots$$

$$\phi = \phi^{(0)}(x) + \dots$$

Background static black holes:

$$ds^2 = -U(r)G dt^2 + \frac{F}{U(r)} dr^2 + g_{ij} dx^i dx^j$$

$$A = a_t dt$$

AdS boundary asymptotics

$$g_{ij} = r^2 \bar{g}_{ij}(x) + \dots$$

$$U = r^2 + \dots$$

$$a_t = \mu(x) + \dots$$

$$\phi = r^{\Delta-d-1} \phi^{(\infty)}(x) + \dots$$

Constraints on the horizon

We consider a generic *sourceless* perturbation $\delta\psi(t, r, x^i)$ with time dependence $\sim e^{-i\omega\nu_{EF}}$.

Infalling conditions at the horizon imply that

$$\delta g_{ri}^{(0)} = \delta g_{ti}^{(0)} \equiv -v_i(x), \quad \delta g_{tr}^{(0)} = \frac{1}{2} \left(\delta g_{rr}^{(0)} + \delta g_{rr}^{(0)} \right) \equiv -\rho(x)/(4\pi T)$$

$$\delta a_r^{(0)} = \delta a_t^{(0)} \equiv w(x)$$

Local “currents” on the horizon

$$Q_{(0)}^i = 4\pi T \sqrt{g^{(0)}} v^i, \quad J_{(0)}^i = \sqrt{g^{(0)}} Z^{(0)} \left(\partial^i w + a_t^{(0)} v^i + i\omega \delta a^{(0)i} \right)$$

Constraints on the horizon

The Hamiltonian constraints on the (stretched) horizon have the form

$$\begin{aligned} \partial_i Q_{(0)}^i &= \omega(\dots), & \partial_i J_{(0)}^i &= \omega(\dots), \\ -2 \nabla^i \nabla_{(i} v_{j)} + \nabla_j p - Z^{(0)} a_t^{(0)} \nabla_i w + \nabla_i \phi^{(0)} v^j \nabla_j \phi^{(0)} &= \omega(\dots) \end{aligned}$$

with the brackets involving various other fields.

The above set of equations is *not* closed.

Diffusive mode

Next, we note that there is a zero mode solution corresponding to perturbing the temperature T by a constant amount δT and the zero lattice mode of the chemical potential $\bar{\mu} \equiv \oint \mu(x)$ by a constant $\delta \bar{\mu}$

$$\delta\psi^{RT} = \frac{\partial\psi}{\partial T} \delta T + \frac{\partial\psi}{\partial\bar{\mu}} \delta\bar{\mu},$$

Can be brought to a form that

- does not introduce boundary sources,
- is regular at horizon,
- satisfies horizon constraints.

Diffusive mode

To construct a bulk diffusive mode, we need to consider a time-dependent sourceless perturbation. We take⁷

$$\delta\psi = e^{-i\omega\nu_{EF}} e^{i\varepsilon k_i x^i} [\delta\psi^{RT} + \varepsilon\delta\psi_{[1]} + \dots]$$

The functions $\delta\psi_{[\alpha]}(x)$ are periodic on the lattice \Rightarrow Bloch decomposition. In particular

$$\omega = \varepsilon\omega_{[1]} + \varepsilon^2\omega_{[2]} + \dots, \quad p = e^{i\varepsilon k_i x^i} (4\pi\delta T + \varepsilon p_{[1]} + \dots),$$

$$v^i = e^{i\varepsilon k_i x^i} (\varepsilon v_{[1]}^i + \varepsilon^2 v_{[2]}^i + \dots), \quad w = e^{i\varepsilon k_i x^i} (-\delta\bar{\mu} + \varepsilon w_{[1]} + \dots),$$

$$\delta\mathbf{g}_{ij}^{(0)} = e^{i\varepsilon k_i x^i} \left(\frac{\partial\mathbf{g}_{ij}^{(0)}}{\partial T} \delta T + \frac{\partial\mathbf{g}_{ij}^{(0)}}{\partial\bar{\mu}} \delta\bar{\mu} + \dots \right),$$

$$\delta\phi^{(0)} = e^{i\varepsilon k_i x^i} \left(\frac{\partial\phi^{(0)}}{\partial T} \delta T + \frac{\partial\phi^{(0)}}{\partial\bar{\mu}} \delta\bar{\mu} + \dots \right)$$

⁷[Donos, Gauntlett, VZ '17]

Diffusive mode

We need to determine the dispersion relation $\omega(\varepsilon k)$ in order to show that this is indeed a diffusive mode.

At order $\mathcal{O}(\varepsilon)$ the horizon constraints imply that $\omega_{[1]} = 0$ and

$$\begin{aligned}\nabla_i v_{[1]}^i &= 0, & \nabla_j \left(Z^{(0)} \left(-i k^j \delta \bar{\mu} + \nabla^j w_{[1]} + v_{[1]}^j a_t^{(0)} \right) \right) &= 0, \\ -2 \nabla^j \nabla_{(j} v_{[1]i)} &+ \nabla_i p_{[1]} - Z^{(0)} a_t^{(0)} \nabla_i w_{[1]} + v_{[1]}^j \nabla_j \phi^{(0)} \nabla_i \phi^{(0)} \\ &+ i k_i 4\pi \delta T + Z^{(0)} a_t^{(0)} i k_i \delta \bar{\mu} &= 0\end{aligned}$$

This system has a unique solution for $v_{[1]}^i$, $p_{[1]}$ and $w_{[1]}$, so we can write

$$\begin{aligned}4\pi T i \oint \sqrt{g^{(0)}} v_{[1]}^i &= T \alpha^{ij} k_j \delta \bar{\mu} + \bar{\kappa}^{ij} k_j \delta T, \\ i \oint \sqrt{g^{(0)}} Z^{(0)} \left(-i k^j \delta \bar{\mu} + \nabla^j w_{[1]} + v_{[1]}^j a_t^{(0)} \right) &= \sigma^{ij} k_j \delta \bar{\mu} + \alpha^{ij} k_j \delta T\end{aligned}$$

Diffusive mode

Finally, integrating the constraints at order $\mathcal{O}(\varepsilon^2)$ we get

$$\begin{aligned}i\omega_{[2]} (T^{-1}c_\mu \delta T + \xi \delta \bar{\mu}) - T^{-1}\bar{\kappa}^{ij} k_i k_j \delta T - \alpha^{ij} k_i k_j \delta \bar{\mu} &= 0, \\i\omega_{[2]} (\xi \delta T + \chi \delta \bar{\mu}) - \alpha^{ij} k_i k_j \delta T - \sigma^{ij} k_i k_j \delta \bar{\mu} &= 0\end{aligned}$$

In the above we have written the thermodynamic susceptibilities

$$\begin{aligned}\delta s &\equiv T^{-1}c_\mu \delta T + \xi \delta \bar{\mu}, \\ \delta \rho &\equiv \xi \delta T + \chi \delta \bar{\mu}\end{aligned}$$

as horizon integrals given a family of background solutions, since we have

$$s = 4\pi \oint_H \sqrt{g^{(0)}}, \quad \rho = \oint_H \sqrt{g^{(0)}} Z^{(0)} a_t^{(0)}$$

Diffusive mode

So we can solve for $\omega_{[2]}$ and obtain the two eigenfrequencies $i\omega_{[2]}^{\pm}$ associated with the diffusion modes. Defining

$$\bar{\kappa}(k) = \bar{\kappa}^{ij} k_i k_j, \quad \alpha(k) = \alpha^{ij} k_i k_j, \quad \sigma(k) = \sigma^{ij} k_i k_j \quad (1)$$

we have⁸

$$\begin{aligned} i\omega_{[2]}^+ i\omega_{[2]}^- &= \frac{\kappa(k) \sigma(k)}{c_\rho \chi}, \\ i\omega_{[2]}^+ + i\omega_{[2]}^- &= \frac{\kappa(k)}{c_\rho} + \frac{\sigma(k)}{\chi} + \frac{T [\chi \alpha(k) - \xi \sigma(k)]^2}{c_\rho \chi^2 \sigma(k)} \end{aligned} \quad (2)$$

where

$$c_\rho \equiv c_\mu - \frac{T \xi^2}{\chi}, \quad \kappa(k) \equiv \bar{\kappa}(k) - \frac{\alpha^2(k) T}{\sigma(k)} \quad (3)$$

The quantities $\bar{\kappa}$, α , σ and the susceptibilities are fixed by horizon data, given a family of solutions.

⁸[Hartnoll '15]

Einstein relations

For a holographic lattice, one can show that $\bar{\kappa}$, α , σ are the DC conductivities of the boundary theory and using the following facts:⁹

- Apply time-independent thermal and electric boundary sources $\delta g_{tj} \leftrightarrow \bar{\zeta}_j$, $\delta a_i \leftrightarrow \bar{E}_i$ and obtain linearised Stokes equations horizon:

$$\begin{aligned}\partial_i Q_{(0)}^i &= 0, & \partial_i J_{(0)}^i &= 0, \\ -2 \nabla^i \nabla_{(i} v_{j)} &- Z^{(0)} a_t^{(0)} \nabla_j w + \nabla_j \phi^{(0)} \nabla_i \phi^{(0)} v^i + \nabla_j p \\ &= 4\pi T \bar{\zeta}_j + Z^{(0)} a_t^{(0)} \bar{E}_j\end{aligned}$$

- In general the *local* horizon current differs from the *local* boundary current, but their *fluxes* are equal

$$\oint_H Q_{(0)}^i = \bar{Q}^i, \quad \oint_H J_{(0)}^i = \bar{J}^i$$

Thus we obtain a set of generalised Einstein relations.

⁹[Donos, Gauntlett '15], [Banks, Donos, Gauntlett '15]

Comments & outlook

- Thermodynamic instability \Rightarrow dynamical instability
- Diffusion poles in retarded Green's functions for spatially inhomogeneous QFTs¹⁰
- Analogous construction of diffusive modes within hydrodynamics on curved manifolds¹⁰
- Sound and Goldstone modes in models with spontaneous symmetry breaking?
- Bounds on diffusion and relation to chaos?¹¹

¹⁰[Donos, Gauntlett, VZ '17]

¹¹[Hartnoll '15, ...], [Blake '16, ...]

Thank you for your attention!