QCD meets Gravity 2018

Gravitational Radiation from Color-kinematics Duality

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based on 1806.07388

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Simple Problem

- Relativistic scattering of classical point sources in post-Minkowskian regime
 - blackholes, neutron stars, etc.
 - hyperbolic orbit + perturbation in G



- Action:
 - Consistent with symmetries, diffeomorphism, reparametrization invariance, etc.
- Classical equations of motion

Deviation from straight-line order by order

... not so simple calculation

- Even 3pt vertex ~ 100 terms
- Infinite tower
 of higher point vertices!
- Modern solution to this classical problem?

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Wolfram <i>Mathematica</i>	STUDENT EDITION	Demonstrations Ma	thWorld Wolfram Community H
Functions $_{_{\mathbb{V}}}$ Sections $_{_{\mathbb{V}}}$	 ↔ Update 		Debug Run Packag
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r ^5 ∗eeString	[eE3]*eeString[eE1, eE2]*keString]	[p1, p1] +	
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(r ^5 *eeString	g[eE1]*eeString[eE2]*eeString[eE3]]*keString[p2, p2])/2 +	
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3∗r^5∗eeStrir	ng[eE2, eE3, eE1]*keString[p2, p3]] -	
(r ^5 ∗eeString	<pre>g[eE1] * eeString[eE2] * eeString[eE3]</pre>]*keString[p3, p1])/4 +	
r ^5 ∗eeString	[eE3]*eeString[eE1, eE2]*keString]	[p3, p1] +	
(3*r^5*eeStri	ing[eE2]*eeString[eE3, eE1]*keStr	ing[p3, p1])/4 -	
3∗r^5∗eeStrir	ng[eE3, eE1, eE2]*keString[p3, p1]] -	
(r ^5 *eeString	<pre>g[eE1] *eeString[eE2] *eeString[eE3]</pre>]∗keString[p3, p2])/4 +	
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3∗r^5∗eeStrir	ng[eE3, eE2, eE1]*keString[p3, p2] –	

Simplicity in Gravity?

• On-shell 3pt amplitudes: simple!



Gravity = Yang-Mills²



[Bern, Carrasco, Johansson '08] [Bern, Carrasco, Johansson; +Dennen, Huang, Kiermaier '10]

QCD meets Classical Gravity?

Double Copy in Exact Classical Solutions

[Monteiro, O'Connell, White; Luna, Nicholson, Berman, Chacon, Bahjat-Abbas; Ridgway, Wise; Carrillo-Gonzalez, Penco, Trodden,...]

> See Bern's review; talks by Steinhoff, Gonzalez, Nicholson, Casali

Double Copy in Perturbative Classical Radiation

[Goldberger, Ridgway, Prabhu, Thompson, Li; Luna, Nicholson, Ochirov, O'Connell, Westerberg, White; Chester; Carrillo-Gonzalez, Penco, Trodden; CHS; Adamo, Casali, Mason, Nekovar]

- Accessible to study higher order/non-linear effects
- Accessible to study broader range of theories/sources

Outline

- Setup and Review
- Next-to-leading Order
- Double Copy
- Conclusion & Outlook



Setup & Review

Theories



Gravity = kinematics^2

Yang-Mills = color x kinematics

See Bern's and Roiban's talk for the full zoology

Theories

Theory	Radiation	Point sources DoF	
Dilaton Gravity	graviton+dilaton	trajectory	
Yang-Mills	gluon	trajectory + charge	
Bi-adjoint	bi-adjoint scalar	trajectory + charge/dual charge	

[Goldberger, Ridgway '16] [Goldberger, Prabhu, Thompson '17]

Theories

$$\begin{aligned}
S_{pp} &= -m^{2} \int d\tau' e^{\kappa \phi} \\
S_{GR} &= \int d^{d}x \sqrt{g} \left[-\frac{2}{\kappa^{2}} R + 2(d-2) g^{\mu\nu} D_{\mu} \phi D_{\nu} \phi \right] \\
\hline
S_{pp} &= -m^{2} \int d\tau' + \int d\tau' \psi^{\dagger} i p \cdot D\psi \\
S_{YM} &= -\frac{1}{4} \int d^{d}x F^{a\mu\nu} F^{a}_{\mu\nu} \\
\hline
S_{pp} &= -m^{2} \int d\tau' + \int d\tau' \left(\psi^{\dagger} i p \cdot \partial \psi + y \phi^{a\bar{a}} c^{a} \tilde{c}^{\bar{a}} \right) \\
S_{BS} &= \int d^{d}x \left(\frac{1}{2} (\partial_{\mu} \phi^{a\bar{a}})^{2} - \frac{y}{3} f^{abc} \tilde{f}^{\bar{a}\bar{b}\bar{c}} \phi_{a\bar{a}} \phi_{b\bar{b}} \phi_{c\bar{c}} \right)
\end{aligned}$$

Scales

Internal size	r ≳ GM	
Orbital size	b~l/q	
Mass	Μ	b
Coupling	G	r

Point source approximation: r ~GM << b
 Classical angular momentum: M b >> 1 (M >> q)

Observables

Observables: radiation at spatial infinity (r → ∞)
 on-shell (k^2=0) and transverse components

$$A^{\pm a}(t,\vec{n}) = \frac{1}{4\pi r} \int \frac{d\omega}{2\pi} e^{-i\omega t_r} \lim_{k^2 \to 0} \left[-k^2 \left(e^{\pm}_{\mu} \cdot A^{\mu a}(k) \right) \right]$$
can be viewed as the amplitude

• GR=YM^2 only refer to this observable part!

Example: YM

• EoM for radiation:

 $\Box A^{\mu a}(x) = -\frac{\delta S_{int}}{\delta A^a_{\mu}} + g \, J^{\mu a}(x)$ Berends-Giele Recursion in QCD

$$\begin{aligned} A^{\mu}_{a}(k) &= -\frac{g}{k^{2}} \int d\tau' \, e^{ik \cdot (b+p\tau')} \, c^{a} p^{\mu} \\ &= -\frac{g}{k^{2}} \, 2\pi \delta(k \cdot p) \, e^{ik \cdot b} \, c^{a} p^{\mu}, \end{aligned}$$



Example: YM

• EoM for radiation:

$$\Box A^{\mu a}(x) = -\frac{\delta S_{int}}{\delta A^a_{\mu}} + g \, J^{\mu a}(x)$$

$$\frac{d^2 x^{\mu}}{d\tau'^2} = g p^{\nu} \left(c \cdot F^{\mu}_{\nu} \right)$$
$$\frac{dc^a}{d\tau'} = ig p^{\mu} \left[c, A_{\mu} \right]$$

$$\begin{aligned} A^{\mu}_{a}(k) &= -\frac{g}{k^{2}} \int d\tau' \, e^{ik \cdot (b+p\tau')} \, c^{a} p^{\mu} & \qquad \delta x^{\mu} \propto \frac{e^{-ik \cdot (b+p\tau')}}{(k \cdot p)^{2}} \, p^{\nu} \, \left(c \cdot F^{\mu}_{\nu} \right) \\ &= -\frac{g}{k^{2}} \, 2\pi \delta(k \cdot p) \, e^{ik \cdot b} \, c^{a} p^{\mu}, & \qquad \delta c^{a} \propto \frac{e^{-ik \cdot (b+p\tau')}}{(ik \cdot p)} \, p^{\nu} \left[c, A_{\nu} \right] \end{aligned}$$



Leading Order

- Pioneered by Goldberger and Ridgway
- YM:
- Gravity+dilaton:



Leading Order

Goldberger and Ridgway: beautiful relation from YM to GR

$$c^a \to p^{\mu}$$
$$f^{abc} \to V^{\mu\nu\rho}_{\rm YM}$$

- Reproduced by S-matrix double copy in classical limit [Luna, Nicholson, Monteiro, O'Connell, White '17]
- Generalization to bound states: quadrupole & potential [Goldberger and Ridgway '17; Steinhoff, Plefka, Wormsbecker'18]
- Generalization to Einstein gravity, spinning sources, scalar theories [Goldberger, Prabhu, Thompson, Li; Chester '17; Carrillo-Gonzalez, Penco, Trodden '18]

See Steinhoff's and Gonzalez's talk

QCD meets Classical Gravity?

- Higher orders are crucial for connection to LIGO
- Challenges at NLO
 - Systematic construction? Gravity for *free*?
 - What are "propagators" ?
 - Color-kinematics duality?



Next-to-leading Order

NLO

~ one-loop process



- For classical EoM, one more source inserted
 - Convenient to imagine this as a 3-body problem
- No clean diagrammatic expression

$$\delta(e^{ikx}) = ikx^{(2)} + \frac{1}{2!} \left(ikx^{(1)}\right)^2 \qquad \boxed{l_2 \underbrace{}} x \underbrace{l_3 \underbrace{}}_{l_3 \underbrace{}}$$

Colors

• 21 color structures

$$\begin{bmatrix} C_1 = (c_1 \cdot c_2)(c_1 \cdot c_3)c_1^a \\ C_2 = (c_1 \cdot c_2)(c_2 \cdot c_3)c_2^a \\ C_3 = (c_1 \cdot c_3)(c_2 \cdot c_3)c_3^a \\ C_4 = (c_1 \cdot [c_2, c_3]) c_1^a \\ C_5 = (c_1 \cdot [c_2, c_3]) c_2^a \\ C_6 = (c_1 \cdot [c_2, c_3]) c_3^a \end{bmatrix} \begin{bmatrix} C_7 = (c_1 \cdot c_2)(c_2 \cdot c_3)c_1^a \\ C_8 = (c_1 \cdot c_2)(c_1 \cdot c_3)c_2^a \\ C_9 = (c_1 \cdot c_2)(c_1 \cdot c_3)c_2^a \\ C_{10} = (c_1 \cdot c_3)(c_2 \cdot c_3)c_2^a \\ C_{11} = (c_1 \cdot c_2)(c_2 \cdot c_3)c_3^a \\ C_{12} = (c_1 \cdot c_2)(c_1 \cdot c_3)c_3^a \end{bmatrix} \begin{bmatrix} C_{13} = (c_1 \cdot c_3)[c_1, c_2]^a \\ C_{14} = (c_2 \cdot c_3)[c_1, c_2]^a \\ C_{15} = (c_1 \cdot c_2)[c_2, c_3]^a \\ C_{16} = (c_1 \cdot c_3)[c_2, c_3]^a \\ C_{17} = (c_1 \cdot c_2)[c_3, c_1]^a \\ C_{18} = (c_2 \cdot c_3)[c_3, c_1]^a \end{bmatrix} \begin{bmatrix} C_{19} = [c_1, [c_2, c_3]]^a \\ C_{20} = [c_2, [c_3, c_1]]^a \\ C_{21} = [c_3, [c_1, c_2]]^a \end{bmatrix}$$

Jacobi identity

 $[c_1, [c_2, c_3]]^a + [c_2, [c_3, c_1]]^a + [c_3, [c_1, c_2]]^a = 0$

Results

 All the trajectories, colors, and radiation are calculated at NLO

 Complexity: 	BS	YM	GR
	20kb	80kb	

Checked Ward identity, conservation laws



Results

 All the trajectories, colors, and radiation are calculated at NLO

 Complexity: 	BS	YM	GR
	20kb	80kb	70Mb

Checked Ward identity, conservation laws



Bi-adjoint amplitudes

Partial amplitudes in bi-adjoint scalar theory

$$\mathcal{J}_{BS}^{a\tilde{a}}(2,C_{7}) = \frac{l_{123} \cdot l_{23}}{l_{3}^{2} l_{23}^{2} (l_{3} \cdot p_{2}) (p_{1} \cdot l_{23})^{2}} \left(\tilde{C}_{4} + \frac{l_{3} \cdot l_{23}}{l_{3} \cdot p_{2}} \tilde{C}_{7}\right) + \frac{1}{l_{3}^{2} l_{23}^{2} (l_{3} \cdot p_{2}) (p_{1} \cdot l_{23})} \left(\tilde{C}_{19} + \frac{l_{3} \cdot l_{23}}{l_{3} \cdot p_{2}} \tilde{C}_{14}\right),$$

$$\mathcal{J}_{BS}^{a\tilde{a}}(2,C_{13}) = -\frac{1}{l_{2}^{2} l_{3}^{2} (l_{2} \cdot p_{1}) (l_{3} \cdot p_{1})} \left(\tilde{C}_{17} \begin{pmatrix} l_{123} \cdot l_{3} \\ l_{3} \cdot p_{1} \end{pmatrix} \tilde{C}_{1}\right) - \frac{1}{l_{2}^{2} l_{3}^{2} (l_{3} \cdot p_{1}) (p_{1} \cdot l_{23})} \left(\tilde{C}_{4} + \frac{l_{2} \cdot l_{3}}{l_{3} \cdot p_{1}} \tilde{C}_{1}\right) \\ + \frac{1}{l_{3}^{2} l_{13}^{2} (l_{3} \cdot p_{1}) (p_{2} \cdot l_{13})} \left(\tilde{C}_{5} - \frac{l_{3} \cdot l_{13}}{l_{3} \cdot p_{1}} \tilde{C}_{9}\right) - \frac{2}{l_{2}^{2} l_{3}^{2} l_{13}^{2} (l_{3} \cdot p_{1})} \left(\tilde{C}_{20} + \frac{l_{3} \cdot l_{13}}{l_{3} \cdot p_{1}} \tilde{C}_{13}\right).$$

$$(53)$$

Double poles, simple poles, composite poles,...
 kinematic numerators...



- Gauge invariance:
 - Need to impose color-kinematics duality

$$C_i \pm C_j = 0 \quad \leftrightarrow \quad N_i \pm N_j = 0$$
$$C_i \pm C_j \pm C_k = 0 \quad \leftrightarrow \quad N_i \pm N_j \pm N_k = 0$$

- Gauge invariance: color-kinematics duality
- Locality:
 - Fuzzy for classical radiation, even nontrivial structure for bi-adjoint theory!
 - Still, the structure should be universal

$$\Box A^{\mu a}(x) = -\frac{\delta S_{int}}{\delta A^a_{\mu}} + g J^{\mu a}(x)$$

$$\frac{d^2 x^{\mu}}{d\tau'^2} = g p^{\nu} \left(c \cdot F^{\mu}_{\nu} \right)$$
$$\frac{dc^a}{d\tau'} = ig p^{\mu} \left[c, A_{\mu} \right]$$

- Gauge invariance: color-kinematics duality
- Locality: build from bi-adjoint scalar theory
 - Fuzzy for classical radiation, even nontrivial structure for bi-adjoint theory!
 - Still, the structure should be universal

$$\Box A^{\mu a}(x) = -\frac{\delta S_{int}}{\delta A^a_{\mu}} + g J^{\mu a}(x)$$

$$\frac{d^2 x^{\mu}}{d\tau'^2} = g p^{\nu} \left(c \cdot F^{\mu}_{\nu} \right)$$
$$\frac{dc^a}{d\tau'} = ig p^{\mu} \left[c, A_{\mu} \right]$$

1. Build the propagator matrix from bi-adjoint theory



1. Build the propagator matrix from bi-adjoint theory

$$A_{BS}(n) = (-1)^n \sum_{i,j} C_i P_{ij} \tilde{C}_j$$
$$A_{YM}(n) = \sum_{i,j} C_i P_{ij} N_j$$

2. Impose color-kinematics duality for gauge invariance

$$C_i \pm C_j = 0 \quad \leftrightarrow \quad N_i \pm N_j = 0$$
$$C_i \pm C_j \pm C_k = 0 \quad \leftrightarrow \quad N_i \pm N_j \pm N_k = 0$$

3. Gravity for free from YM^2

$$A_{BS}(n) = (-1)^n \sum_{i,j} C_i P_{ij} \tilde{C}_j$$

$$A_{YM}(n) = \sum_{i,j} C_i P_{ij} N_j$$

$$A_{GR}(n) \clubsuit (-1)^n \sum_{i,j} N_i P_{ij} \tilde{N}_j,$$

$$C_i \pm C_j = 0 \quad \leftrightarrow \quad N_i \pm N_j = 0$$

$$C_i \pm C_j \pm C_k = 0 \quad \leftrightarrow \quad N_i \pm N_j \pm N_k = 0$$

Perfectly works at next-to-leading order

$$A_{BS}(n) = (-1)^n \sum_{i,j} C_i P_{ij} \tilde{C}_j$$

$$A_{YM}(n) = \sum_{i,j} C_i P_{ij} N_j$$

$$C_i \pm C_j = 0 \quad \leftrightarrow \quad N_i \pm N_j = 0$$

$$C_i \pm C_j \pm C_k = 0 \quad \leftrightarrow \quad N_i \pm N_j \pm N_k = 0$$

$$A_{GR}(n) = (-1)^n \sum_{i,j} N_i P_{ij} \tilde{N}_j,$$

• Perfectly works at NLO: hidden simplicity

$$\begin{split} A_{\rm BS}(n) &= (-1)^n \sum_{i,j} C_i \, P_{ij} \, \tilde{C}_j \\ A_{\rm YM}(n) &= \sum_{i,j} C_i \, P_{ij} \, N_j \\ A_{\rm GR}(n) &= (-1)^n \sum_{i,j} N_i \, P_{ij} \, \tilde{N}_j, \end{split}$$

• Leading order:

$$c^{a} \to p^{\mu}$$
 $f^{abc} \to V^{\mu\nu\rho}_{\rm YM}$

11

 α

Non-trivially generalized at NLO

$C_1 = (c_1 \cdot c_2)(c_1 \cdot c_3)c_1^a$	$N_1 = (p_1 \cdot p_2)(p_1 \cdot p_3)(p_1 \cdot e)$
$C_4 = (c_1 \cdot [c_2, c_3]) c_1^a$	$N_4 = \left((p_1 \cdot [p_3, p_2] \cdot l_{23}) + \frac{1}{2} (p_2 \cdot p_3) (p_1 \cdot q_{23}) \right) (p_1 \cdot e) + \frac{1}{2} (p_1 \cdot l_{23}) (p_1 \cdot [p_2, p_3] \cdot e)$
$C_7 = (c_1 \cdot c_2)(c_2 \cdot c_3)c_1^a$	$N_7 = (p_1 \cdot p_2)(p_2 \cdot p_3)(p_1 \cdot e)$
$\frac{O(1 + O_2)(O_2 + O_3)O_1}{O(1 + O_2)(O_2 + O_3)O_1}$	$N_{13} = (p_1 \cdot p_3) \left((l_{123} \cdot [p_1, p_2] \cdot e) - (p_1 \cdot p_2) (l_2 \cdot e) \right)$
$C_{13} = (c_1 \cdot c_3) [c_1, c_2]$	$N_{19} = \left((l_1 \cdot [p_2, p_3] \cdot l_{23}) - \frac{1}{2} (p_2 \cdot p_3) (l_{123} \cdot q_{23}) \right) (p_1 \cdot e) + (p_1 \cdot l_{23}) (l_{23} \cdot [p_2, p_3] \cdot e)$
$C_{19} = [c_1, [c_2, c_3]]^a$	$+ (p_1 \cdot [p_2, p_3] \cdot l_{23}) (l_{23} \cdot e) - (p_2 \cdot p_3) (p_1 \cdot [l_2, l_3] \cdot e) - \frac{1}{2} (l_{23} \cdot l_{123}) (p_1 \cdot [p_2, p_3] \cdot e)$

• Leading order:

$$c^{a} \to p^{\mu}$$
 $f^{abc} \to V^{\mu\nu\rho}_{\rm YM}$

11

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Non-trivially generalized at NLO

$N_1 = (p_1 \cdot p_2)(p_1 \cdot p_3)(p_1 \cdot e)$
$N_4 = \left((p_1 \cdot [p_3, p_2] \cdot l_{23}) + \frac{1}{2} (p_2 \cdot p_3) (p_1 \cdot q_{23}) \right) (p_1 \cdot e) + \frac{1}{2} (p_1 \cdot l_{23}) (p_1 \cdot [p_2, p_3] \cdot e)$
$N_7 = (p_1 \cdot p_2)(p_2 \cdot p_3)(p_1 \cdot e)$
$N_{13} = (p_1 \cdot p_3) \left((l_{123} \cdot [p_1, p_2] \cdot e) - (p_1 \cdot p_2) (l_2 \cdot e) \right)$
$N_{19} = \left((l_1 \cdot [p_2, p_3] \cdot l_{23}) - \frac{1}{2} (p_2 \cdot p_3) (l_{123} \cdot q_{23}) \right) (p_1 \cdot e) + (p_1 \cdot l_{23}) (l_{23} \cdot [p_2, p_3] \cdot e)$
$+ (p_1 \cdot [p_2, p_3] \cdot l_{23}) (l_{23} \cdot e) - (p_2 \cdot p_3) (p_1 \cdot [l_2, l_3] \cdot e) - \frac{1}{2} (l_{23} \cdot l_{123}) (p_1 \cdot [p_2, p_3] \cdot e)$

• Leading order:

$$c^{a} \to p^{\mu}$$
 $f^{abc} \to V^{\mu\nu\rho}_{YM}$

0. 11.

- Non-trivially generalized at NLO
- Still, it can be chosen s.t.

$$c_i \cdot c_j \to p_i \cdot p_j$$

Another example of kinematic algebra?

[Monteiro, O'Connell '11, '13; Cheung, CHS, '16; Cheung, Remmen, CHS, Wen '17; Fu, Vanhove, Wang '18]

BCJ and KLT relations

• BCJ relations: non-trivial consequence of the duality

$$A_{\rm YM}(n) = \sum_{i,j} C_i P_{ij} N_j$$

- **e.g. LO:** $(p_1 \cdot l_2)A(1, C_1) + (p_2 \cdot l_1)A(1, C_2) + \frac{1}{2}(l_1^2 l_2^2)A(1, C_3) = 0.$
- KLT relations: GR= YM^2 at physical level

$$A_{\rm GR}(n) = (-1)^n \sum_{i,j} N_k P_{ki} P_{ij}^{-1} P_{jl} N_l$$

= $(-1)^n \sum_{i,j} A_{\rm YM}(n, C_i) S_{ij} A_{\rm YM}(n, C_j), \checkmark$

• e.g. LO: $S_{11} = l_2^2 (p_1 \cdot l_2)^2 / (l_2 \cdot l_{12}), S_{22} = l_1^2 (p_2 \cdot l_1)^2 / (l_1 \cdot l_{12}),$



Conclusion & Outlook

Conclusion

• In the classical radiation @ NLO, we find



Gravity for free; Systematic approach works

Color-kinematics duality w/ BCJ + KLT relations

Midden simplicity

Double copy for Einstein gravity?

Bound orbit?

• Higher order?

- Double copy for Einstein gravity?
 - Done at LO by adding "ghost" in YM [Luna, Nicholson, Monteiro, O'Connell, White '17]
- Bound orbit?

• Higher order?

- Double copy for Einstein gravity?
 - Done at LO by adding "ghost" in YM [Luna, Nicholson, Monteiro, O'Connell, White '17]
- Bound orbit?
 - Direct confirmation at LO [Goldberger and Ridgway '17]
 - Universality: match to EFT action via S-matrix see Cheung's talk [Neill, Rothstein '13; Vaidya '14; Cheung, Rothstein, Solon '18; Bern, Cheung, Roiban, CHS, Solon, Zeng]
- Higher order?

- Double copy for Einstein gravity?
 - Done at LO by adding "ghost" in YM [Luna, Nicholson, Monteiro, O'Connell, White '17]
- Bound orbit?
 - Direct confirmation at LO [Goldberger and Ridgway '17]
 - Universality: match to EFT action via S-matrix see Cheung's talk [Neill, Rothstein '13; Vaidya '14; Cheung, Rothstein, Solon '18; Bern, Cheung, Roiban, CHS, Solon, Zeng]
- Higher order?
 - Understand the connection to S-matrix see O'Connell's talk; Proof? [Luna, Nicholson, Monteiro, O'Connell, White '17; Kosower, Maybee, O'Connell '18]







Backups

Perturbative Solutions

General structure:

$$A^{\mu a}(k) = \sum_{n} \left(-\frac{g^{2n+1}}{k^2} \sum_{\text{perm}} \int_1 \int_2 \cdots \int_{n+1} (2\pi)^d \delta^d (k - \sum_{i=1}^{n+1} l_i) \mathcal{J}^{\mu a}(n) \right)$$

$$phase \text{ space} \qquad Mom. \text{ conservation} \qquad \text{``Integrand''}$$

$$\int_i = \int \frac{d^d l_i}{(2\pi)^d} \frac{2\pi \delta(p_i \cdot l_i) e^{il_i b_i}}{(2\pi)^d} \qquad \text{``on-shell''} \qquad \text{``phase shift.}$$

• Zero in YM but not zero in GR?



• Zero in YM but not zero in GR?



• Well-known examples in multi-loop gravity [Bern, Carrasco, Dixon, Johansson, Roiban, '12]

$YM: C \times N = 0 \times N \rightarrow GR: N^*N \neq 0$

3-body treatment naturally avoid this degeneracy