

Perturbative construction of gravity action via BRST ghosts

Silvia Nagy

University of Nottingham

based on **Phys.Rev.Lett. 121 (2018) no.21, 211601**
with A. Anastasiou, L. Borsten, M. Duff, and M. Zoccali

December 14, 2018

Many promising results

Examples of solutions in classical gravitational theories from double copy piling up [Bahjat-Abbas, Bern, Bjerrum-Bohr, Carrillo Gonzalez, Cheung, Damgaard, Festuccia, Goldberger, Kosower, Li, Luna, Maybee, Monteiro, Nicholson, Ochirov, O'Connell, Peneduto, Penco, Prabhu, Ridgway, Rothstein, Shen, Solon, Thompson, Trodden, Vanhove, Westerberg, White]

How general is the double copy?

- Can we construct the full space of the gravity theory in a *gauge-independent* way ?
- Could we *derive* the gravity action from YM ?

How general is the double copy?

- Can we construct the full space of the gravity theory in a *gauge-independent* (and respectively *coordinate independent*) way ?
- Could we *derive* the gravity action from YM ?
- First and second order in perturbation theory for pure gravitational theory.

Minimal YM^2

- Schematically:

$$A_\mu * \tilde{A}_\nu \equiv h_{\mu\nu} + B_{\mu\nu} + \eta_{\mu\nu}\phi$$

- Want to construct the **full theory** with **arbitrary boundary conditions** of graviton, dilaton and two-form from the double copy.
- If this is achievable, it should be possible to extract (pure) gravitational solutions by requiring $B_{\mu\nu} = \phi = 0$
- Start by constructing Lorenz-covariant dictionaries for all the fields compatible with **symmetries** and **equations of motion**.

The set-up

- We build gravitational fields as convolutions of gauge fields with the bi-adjoint scalar [Anastasiou, Borsten, Duff, Hughes, SN 2014]:

$$Z_{\mu\nu} = A_{\mu}^i \star [\Phi_{ii'}]^{-1} \star \tilde{A}_{\nu}^{i'} \equiv A_{\mu} \star \tilde{A}_{\nu}$$

- The convolution is defined as

$$[f \star g](x) = \int d^4y f(y)g(x - y).$$

- Importantly, is **doesn't** obey the Leibnitz rule when the functions f, g have appropriate fall-off:

$$\partial_{\mu}(f \star g) = (\partial_{\mu}f) \star g = f \star (\partial_{\mu}g)$$

Symmetry requirement

- Reproduces the correct local transformations at linear level:

$$\delta h_{\mu\nu} = \partial_\mu \xi_\nu + \partial_\nu \xi_\mu$$

$$\delta B_{\mu\nu} = \partial_\mu \Lambda_\nu - \partial_\nu \Lambda_\mu$$

$$\delta \phi = 0$$

from

$$\delta A_\mu^a = \partial_\mu \alpha^a + f_{bc}^a A_\mu^b \theta^c$$

$$\delta \phi_{a\tilde{a}} = -f_{ac}^b \phi_{b\tilde{a}} \theta^c - f_{\tilde{a}\tilde{c}}^{\tilde{b}} \phi_{a\tilde{b}} \theta^{\tilde{c}}$$

Dynamics requirement

- Reproduce gravitational eom

$$R_{\mu\nu} = j_{\mu\nu}^{(h)}$$

$$\partial^\rho H_{\rho\mu\nu} = j_{\mu\nu}^{(B)}$$

$$\square\phi = j^{(\phi)}$$

from

- Yang-Mills

$$\partial^\mu F_{\mu\nu}^i = j_\nu^i, \quad \partial^\mu (*F_{\mu\nu}^i) = 0, \quad \partial^\mu j_\mu^i = 0$$

Sources have a dual role : they ensure proper fall-off for the fields to be convoluted and they will illuminate a fundamental issue of the double copy.

Field dictionary [Cardoso, Inverso, SN, Nampuri]

$$h_{\mu\nu} = A_\mu \star \tilde{A}_\nu + A_\nu \star \tilde{A}_\mu + q\eta_{\mu\nu} \left[A_\rho \star \tilde{A}^\rho - \frac{1}{\square} (\partial \cdot A) \star (\partial \cdot \tilde{A}) \right]$$

$$B_{\mu\nu} = A_\mu \star \tilde{A}_\nu - A_\nu \star \tilde{A}_\mu$$

$$\phi = A_\rho \star \tilde{A}^\rho - \frac{1}{\square} (\partial \cdot A) \star (\partial \cdot \tilde{A})$$

Source dictionary

$$j_{\mu\nu}^{(h)} = -j_{(\mu} \star \tilde{j}_{\nu)} + (q+1) \frac{\partial_\mu \partial_\nu}{\square} j_\rho \star \tilde{j}^\rho$$

$$j_{\mu\nu}^{(B)} = 2j_{[\mu} \star \tilde{j}_{\nu]}$$

$$j^{(\phi)} = j_\rho \star \tilde{j}^\rho$$

Source dictionary and constraint

$$j_{\mu\nu}^{(h)} = -j_{(\mu} \star \tilde{j}_{\nu)} + (q+1) \frac{\partial_\mu \partial_\nu}{\square} j_\rho \star \tilde{j}^\rho$$

$$j_{\mu\nu}^{(B)} = 2j_{[\mu} \star \tilde{j}_{\nu]}$$

$$j^{(\phi)} = j_\rho \star \tilde{j}^\rho$$

Note that our theory is constrained:

$$j^\phi \propto -T^\rho_\rho$$

Obstruction to obtaining pure gravity - even for the most general dictionary !!

Source dictionary and constraint

$$j_{\mu\nu}^{(h)} = -j_{(\mu} \star \tilde{j}_{\nu)} + (q+1) \frac{\partial_\mu \partial_\nu}{\square} j_\rho \star \tilde{j}^\rho$$

$$j_{\mu\nu}^{(B)} = 2j_{[\mu} \star \tilde{j}_{\nu]}$$

$$j^{(\phi)} = j_\rho \star \tilde{j}^\rho$$

Note that our theory is constrained:

$$j^\phi \propto -T^\rho{}_\rho$$

Obstruction to obtaining pure gravity - even for the most general dictionary !!

- issue with off-shell d.o.f. counting: $A_\mu \star \tilde{A}_\nu$ has 9, whereas $h_{\mu\nu}$ has 6, $B_{\mu\nu}$ has 3 and ϕ has 1.

Is there a more natural home for the double copy?

Is there a more natural home for the double copy?

- Siegel('95) conjectured in that a string-inspired double-copy would be manifested at the level of BRST systems (physical fields+ghosts)

	physical	ghost
physical	physical	ghost
ghost	ghost	physical + gh-gh

- Originally considered because it resolves the apparent conflict between on-shell and off-shell d.o.f (in a way that extends naturally to supergravity theories).

BRST

- introduced to avoid issues caused by gauge symmetry in path integrals
- perform a gauge-fixing

$$\mathcal{L}^{BRST} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + b\mathcal{G}(A_{\mu}^i, b, \xi_{(A)})$$

- replace the gauge symmetry by a fermionic global symmetry by introducing a ghost-antighost pair for each original gauge parameter. The symmetries have the same form as the original ones, but with gauge parameters replaced by ghosts:

$$Q_{BRST} A_{\mu}^i = \partial_{\mu} c^i$$

$$Q_{BRST} c = 0$$

$$Q_{BRST} \bar{c} = b$$

$$Q_{BRST} b = 0$$

BRST

- introduced to avoid issues caused by gauge symmetry in path integrals
- perform a gauge-fixing

$$\mathcal{L}^{BRST} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - b\mathcal{G}(A_{\mu}^i, b, \xi_{(A)})$$

- replace the gauge symmetry by a fermionic global symmetry by introducing a ghost-antighost pair for each original gauge parameter and write a new Lagrangian which manifests this symmetry by the prescription:

$$\mathcal{L}^{BRST} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - b\mathcal{G}(A_{\mu}^i, b, \xi_{(A)}) - \bar{c} \int d^4y \frac{\delta\mathcal{G}}{\delta A_{\mu}} \partial^{\mu} c$$

- At classical level this is equivalent with the original one!

BRST for linearised YM

- Pick a Lorenz covariant ξ -gauge for YM field:

$$\mathcal{G}(A_\mu, b, \xi_{(A)}) = \partial^\mu A_\mu + \frac{\xi_{(A)}}{2} b$$

- and corresponding one for graviton

$$\mathcal{G}(h_{\mu\nu}, b_\mu, \xi_{(h)}) = h_{\mu\nu} - \frac{1}{2} \partial_\mu h + \frac{\xi_{(h)}}{2} b_\mu$$

with BRST transformations

$$Q_{BRST} h_{\mu\nu} = \partial_\mu c_\nu + \partial_\nu c_\mu$$

$$Q_{BRST} c_\mu = 0$$

$$Q_{BRST} \bar{c}_\mu = b_\mu$$

$$Q_{BRST} b_\mu = 0$$

- Similarly for the two-form $B_{\mu\nu}$, though more complicated: we have ghosts d_μ, \bar{d}_μ ghosts-for-ghosts d, \bar{d} and a physical spin-0 field η in the gauge-fixing term [Kimura '80].

BRST dictionary

- *symmetry requirement*: reproduce BRST symmetries of gravitational fields, ghosts, ghosts-for-ghosts from those of YM.
- *dynamics requirement*: reproduce gravitational e.o.m. in the BRST form

$$\begin{aligned} \square h_{\mu\nu} - \frac{\xi^{(h)+2}}{\xi^{(h)}} \left(2\partial^\rho \partial_{(\mu} h_{\nu)\rho} - \partial_\mu \partial_\nu \right) h &= j_{\mu\nu}^{(h)} \\ \square B_{\mu\nu} + \frac{\xi^{(B)+2}}{\xi^{(B)}} 2\partial^\rho \partial_{[\mu} B_{\nu]\rho} &= j_{\mu\nu}^{(B)} \\ \square \varphi &= j^{(\varphi)} \end{aligned}$$

from YM equations in BRST form

$$\square A_\mu - \frac{\xi+1}{\xi} \partial_\mu \partial A = j_\mu, \quad \square c^\alpha = j^\alpha$$

We find

$$\xi_{(h)} = \xi_{(B)} = \xi$$

and

$$\begin{aligned} h_{\mu\nu} = & A_{(\mu} * A_{\nu)} + \frac{1}{1-\xi} \frac{\partial_\mu \partial_\nu}{\square} A * A + \frac{1+\xi}{2(1-\xi)} \frac{\partial_\mu \partial_\nu}{\square} c * c \\ & - \frac{1}{2} \frac{1}{\square} \left(\partial A * \partial_{(\mu} A_{\nu)} + \partial_{(\mu} A_{\nu)} * \partial A \right) \\ & + \frac{\xi}{(2-D)(\xi-1)} \eta_{\mu\nu} \left(A * A + \frac{1}{\xi} c * c + \left(\frac{1}{\xi^2} - 1 \right) \frac{1}{\square} \partial A * \partial A \right) \\ B_{\mu\nu} = & A_{[\mu} * A_{\nu]} - \frac{1}{2} \frac{1}{\square} \left(\partial A * \partial_{[\mu} A_{\nu]} - \partial_{[\mu} A_{\nu]} * \partial A \right) \\ \phi = & A * A + \frac{1}{\xi} c * c + \left(\frac{1}{\xi^2} - 1 \right) \frac{1}{\square} \partial A * \partial A \end{aligned}$$

Gravity dictionary

Kill $B_{\mu\nu}$ by setting $A_\mu = \tilde{A}_\mu$ and kill dilaton by setting

$$\frac{1}{\xi} c * c = -A * A \left(1 - \frac{1}{\xi^2}\right) \frac{1}{\square} \partial A * \partial A$$

so graviton dictionary reduces to

$$h_{\mu\nu} = A_{(\mu} * A_{\nu)} + \frac{1}{1-\xi} \frac{\partial_\mu \partial_\nu}{\square} A * A + \frac{1+\xi}{2(1-\xi)} \frac{\partial_\mu \partial_\nu}{\square} c * c \\ - \frac{1}{2} \frac{1}{\square} \left(\partial A * \partial_{(\mu} A_{\nu)} + \partial_{(\mu} A_{\nu)} * \partial A \right)$$

E.o.m. dictionary

Source dictionary for graviton-dilaton

$$\begin{aligned}
 j_{\mu\nu}^{(h)} &= a_6 \frac{1}{\square} j_{(\mu} * \tilde{j}_{\nu)} + a_7 \frac{\partial_\mu \partial_\nu}{\square^2} j^\rho * \tilde{j}_\rho + a_8 \frac{\partial_\mu \partial_\nu}{\square^2} j^a * \tilde{j}_a \\
 &\quad + \eta_{\mu\nu} \frac{1}{\square} \left[b_4 j^\rho * \tilde{j}_\rho + b_5 j^a * \tilde{j}_a \right] \\
 j^\phi &= \frac{1}{\square} \left[j^\rho * \tilde{j}_\rho + \frac{1}{\xi} j^a * \tilde{j}_a \right]
 \end{aligned}$$

with

$$a_6 = \frac{1}{1-\xi}, \quad a_7 = \frac{(1+\xi)^2}{\xi(1-\xi)}, \quad a_8 = \frac{(1+\xi)^2}{\xi^2(1-\xi)}, \quad b_4 = \frac{\xi}{2(1-\xi)}, \quad b_5 = \frac{1}{2(1-\xi)}$$

E.o.m. dictionary

Kill dilaton source this reduces to

$$j_{\mu\nu}^{(h)} = \frac{1}{1-\xi} \frac{1}{\square} j_{(\mu} \star \tilde{j}_{\nu)}$$

of

$$e.o.m.(h_{\mu\nu}) = \frac{1}{1-\xi} \frac{1}{\square} [e.o.m.(A_\mu) \star e.o.m.(A_\nu)]$$

Can construct gravity Lagrangian directly from YM fields !

Higher order

- First obstacle: how are the gravity and YM gauge fixing terms \mathcal{G} related?
- At linear level, we made a guess that generalised Lorenz corresponds to generalised De Donder, which turned out to work, because of the simplicity of the action.
- At second order, could be more complicated, even restricting to Lorenz covariant gauges
- Writing most general \mathcal{G} can be very ugly!

Higher order

- First obstacle: how are the gravity and YM gauge fixing terms \mathcal{G} related?
- At linear level, we made a guess that generalised Lorenz corresponds to generalised De Donder, which turned out to work, because of the simplicity of the action.
- at second order, could be more complicated, even restricting to Lorenz covariant gauges
- use BCJ double copy to work out relation between $\mathcal{G}(h)^{(2)}$ and $\mathcal{G}(A)^{(2)}$

Second order

Sticking with generalised Lorenz for YM, we get, at second order

$$\square A_{\mu}^{(2)a} - \frac{\xi+1}{\xi} \partial_{\mu} \partial A^{(2)a} = -f^{abc} \left[\partial A^{(1)b} A_{\mu}^{(1)c} + 2A_{\rho}^{(1)b} \partial^{\rho} A_{\mu}^{(1)c} + \partial_{\mu} (\bar{c}^{(1)b} c^{(1)c}) \right]$$

Second order

Pick $\xi = -1$ for simplicity

$$\square A_{\mu}^{(2)a} = -f^{abc} \left[\partial A^{(1)b} A_{\mu}^{(1)c} + 2A_{\rho}^{(1)b} \partial^{\rho} A_{\mu}^{(1)c} + \partial_{\mu} (\bar{c}^{(1)b} c^{(1)c}) \right]$$

Fourier transform and write in BCJ form

$$-p^2 A^{(2)\mu a}(-p) = if^a{}_{bc} \int \bar{d}k \bar{d}q \bar{d}\delta \left[n^{\mu\nu\sigma} A_{\nu}^{(1)b}(k) A_{\sigma}^{(1)c}(q) \right. \\ \left. + n^{\mu} c^{(1)\alpha b}(k) c^{(1)\alpha c}(q) \right]$$

with

$$n^{\mu\nu\sigma} = \frac{1}{2} [(q - k)^{\mu} \eta^{\nu\sigma} + (k - p)^{\sigma} \eta^{\mu\nu} + (p - q)^{\nu} \eta^{\sigma\mu}], \\ n^{\mu} = \frac{1}{2} (k^{\mu} + q^{\mu}).$$

second order

we get

$$\begin{aligned}
 -p^2 H^{(2)\mu\tilde{\mu}} = i \int dkdq\delta \left[n^{\tilde{\mu}}{}_{\tilde{\nu}\tilde{\sigma}} n^{\mu\nu\sigma} H_{\nu}^{(1)\tilde{\nu}} H_{\sigma}^{(1)\tilde{\sigma}} + n^{\tilde{\mu}}{}_{\tilde{\nu}\tilde{\sigma}} n^{\mu} D^{(1)\alpha\tilde{\nu}} D^{(1)}{}_{\alpha}{}^{\tilde{\sigma}} \right. \\
 \left. + n^{\tilde{\mu}} n^{\mu\nu\sigma} D_{\nu}^{(1)\tilde{\alpha}} D_{\sigma\tilde{\alpha}}^{(1)} + n^{\tilde{\mu}} n^{\mu} C^{(1)\alpha\tilde{\alpha}} C^{(1)}{}_{\alpha\tilde{\alpha}} \right]
 \end{aligned}$$

with

$$H_{\mu\tilde{\mu}}^{(i)} = A_{\mu}^{(i)} \star \tilde{A}_{\tilde{\mu}}^{(i)} \quad (1a)$$

$$D_{\mu\tilde{\alpha}}^{(i)} = A_{\mu}^{(i)} \star \tilde{c}_{\tilde{\alpha}}^{(i)} \quad (1b)$$

$$D_{\alpha\tilde{\mu}}^{(i)} = c_{\alpha}^{(i)} \star \tilde{A}_{\tilde{\mu}}^{(i)} \quad (1c)$$

$$C_{\alpha\tilde{\alpha}}^{(i)} = c_{\alpha}^{(i)} \star \tilde{c}_{\tilde{\alpha}}^{(i)} \quad (1d)$$

Second order

- Could stop at this point and study solutions
- We want general second order equation: invert linear dictionary to obtain expressions for $A \star A$ etc and use these to extract the gravitational gauge fixing term for our chosen YM \mathcal{G} :

$$\mathcal{G}(h)^{(2)} = [\partial_\mu (\sqrt{-g} g^{\mu\nu})]^{(2)} + \frac{1}{\square} f(h)$$

- Sanity check: the ghost terms in the e.o.m. as they come from the amplitude d.c. have to match the terms as they come from acting on \mathcal{G} with Q_{BRST}

In progress

- Is it possible to give a direct fields dictionary at second order, now that we understand how to match e.o.m. ? (bi-adjoint scalar will play a role here)
- quantum applications (anomalies etc)
- more speculative: understanding non-perturbative double copy systematically will require Batalin Vilkovisky formalism.

Thank You !