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$n \rightarrow n'$ Transformation and Neutron Transition Magnetic Moment

Based on the paper to be submitted to arXiv by Zurab Berezhiani, Riccardo Biondi (L'Aquila U & LNGS) YK (U. Tennessee), Louis Varriano (Chicago U.)

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General theoretical framework:

neutron can spontaneously transform to "mirror" (dark, sterile) neutron with degenerate mass (Berezhiani and Bento: 2006):

$$m_n = m_{n'}$$

Mirror neutron can be part of Mirror Matter that can be a fraction or a whole of Dark Matter (Berezhiani, Comelli, Villante: 2001, ...)

$$\mathcal{L}_{tot} = \mathcal{L}_{SM} + \mathcal{L}_{SM'} + \mathcal{L}_{mix}$$

with gravity common for both sectors

 $\mathcal{L}_{mix} \rightarrow \text{SM}$ neutral singlets can transform between $SM \leftrightarrow SM'$, therefore $n \leftrightarrow n'$ violating Baryon number by 1 unit is possible.

Yesterday Z. Berezhiani discussed a variation of this model where $\Delta m = |m_n - m_{n'}| \sim 10^{-7} \text{eV}$. This conjecture will be easily tested (see talk of Leah Broussard). Here we will be back to the initial Mirror Matter assumption: $m_n = m_{n'}$

Will discuss:

- what is neutron Transition Magnetic Moment (TMM)
- how it changes $n \rightarrow n'$ oscillations in magnetic field
- how TMM can influence *n* lifetime measurement
- how TMM effect can be measured



 μ makes *n* spin precession in magnetic field

Example of Transition Magnetic Moment



Interaction with magnetic photon $U = \mu_{\Sigma^0 \Lambda} \boldsymbol{\sigma} \cdot \boldsymbol{B}$

leads to decay
$$\Sigma^0 \rightarrow \Lambda^0 + \gamma$$

Possible mechanism of $n \rightarrow n'$

 ϵ -mixing term in Hamiltonian $\begin{pmatrix} m & \epsilon \\ \epsilon & m \end{pmatrix}$



If this process exists, then photon γ and mirror photon γ' can interact with n and n'

Neutron Transition Magnetic Moment TMM



Neutron and mirror neutron TMM $\kappa_n = \kappa_{n'} = \kappa$

is
$$\kappa = \kappa(\epsilon)$$
?

Origin of *n*TMM can be more complicated



ϵ and κ can be two independent parameters

Magnitudes of ϵ and κ a priori not known, but can be found or limited experimentally

Notice, TMM for $n \rightarrow \overline{n}$ transformation is zero

Non-zero TMM for $n \rightarrow \overline{n}$ would violate rotational invariance (<u>Berezhiani & Vainshtein 2018</u>)

General form of $n \leftrightarrow n'$ Hamiltonian in the presence of **B** and **B**'

$$\mathcal{H} = \mathcal{H}_{0} + \mathcal{H}_{int}$$

$$= \begin{pmatrix} E & \epsilon \\ \epsilon & E' \end{pmatrix} + \begin{pmatrix} \mu \sigma \cdot B & \kappa \sigma \cdot [B \pm B'] \\ \kappa \sigma \cdot [B \pm B'] & \mu' \sigma \cdot B' \end{pmatrix}$$
interaction of (n, n') system with B and B'

$$= 0$$
 if $B = B' = 0$

 ϵ — mixing parameter of $n \leftrightarrow n'$: $\epsilon = 1/\tau$, where τ — oscillation time;

 $\mu = \mu'$ usual neutron magnetic moment $\mu = -1.91 \ \mu_N$

 κ – transitional magnetic moment, sign is unknown, assume that $|\kappa| \ll |\mu|$

 \pm here is due to unknown possibly different relative parity of **B** and **B**'

Assumptions:

if $n \rightarrow n'$ oscillations period is τ Serebrov at al. (2009): $\tau > 448$ s at B = 0 assuming B'=0It follows then that $\epsilon = \hbar/\tau$ is small $\epsilon \leq 10^{-18} eV$ In the B = 0.5G Earth magnetic field $\mu B \cong 3 \times 10^{-12} eV$ Let's assume that $\epsilon \ll \kappa B$ and we can neglect ϵ . Also $\kappa B \ll \mu B$, so κ can be e.g. $\kappa \sim (10^{-3} - 10^{-5})\mu$ For simplification we also assume that B'=0 and consider neutrons with only one polarization

$$\mathcal{H} = \begin{pmatrix} E + \mu B & \epsilon + \kappa B \\ \epsilon + \kappa B & E' \end{pmatrix} = \begin{pmatrix} x & d \\ d & y \end{pmatrix}$$

This is real Hermitian matrix with eigenvalues determining the energy levels of two oscillating components:

$$\mathcal{E}_{1,2} = \frac{x+y}{2} \pm \sqrt{\left(\frac{x-y}{2}\right)^2 + d^2} \qquad d = \epsilon + \kappa B \cong \kappa B$$

$$\Omega = \frac{\mathcal{E}_1 - \mathcal{E}_2}{2} = \sqrt{\left(\frac{x-y}{2}\right)^2 + d^2} \qquad \text{oscillation frequency}$$

$$\frac{x-y}{2} = \frac{E + \mu B - E'}{2} \cong \frac{\mu B}{2} \qquad \text{in constant mag. field}$$

B

$$\Omega = \sqrt{(\mu B/2)^2 + (\kappa B)^2} \cong \mu B/2 \qquad d \cong \kappa B$$

$$P_{n \to n'}(t) = \frac{d^2}{\Omega^2} sin^2(\Omega t)$$

Solution in a constant magnetic field for initial neutron condition $\begin{pmatrix} 1\\ 0 \end{pmatrix}$

• For B = 0: $P_{n \to n'}(t) = sin^2(\epsilon t)$ oscillation with $\tau > 448$ s

• In UCN storage experiments $\langle \Delta t \rangle \sim 0.1$ s between collisions with walls, and $\langle sin^2(\Omega t) \rangle = 1/2$ at B > few mG

$$P_{n \to n'}(t) = \frac{2\kappa^2}{\mu^2} = 2 \cdot (10^{-3} - 10^{-5})^2$$

Interesting Result: In MM model with *n*TMM

for B > few mG

$$P_{n \to n'}(t) = \frac{2\kappa^2}{\mu^2}$$

- Probability of $n \rightarrow n'$ is constant independent on B;
- $P_{n \rightarrow n'} \lesssim 10^{-6}$;
- Oscillation frequency Ω increases with B;
- Probability is the same for both *n* polarizations;
- If B = 0, but B' > few mG same effect;
- If $B' \neq 0$ resonance at B = B' is possible enhanced by $\kappa B' > \epsilon$

Can we relate TMM effect to *n* lifetime experiments?



Halbach Magnetic Array in UCN τ experiment



Neutron bouncing in non-uniform magnetic field without touching walls





Decoherence measures nn' system

What minimum *B* – gradient is needed to "measure" *nn*['] system?



nn' in unusual QM system where magnetic field gradient is exerted only on one component of the system.

At some value of gradient G = dB/dz decoherence will "measure" the nn' system. We can assume that this will occur when the separation of energy levels due to gradient will exceed the energy width of nn' wave packet. What minimum *B* – gradient is needed to "measure" *nn*['] system?

$$\begin{pmatrix} 1 \\ 0 \end{pmatrix} \leftarrow \mathbf{G} = d\mathbf{B}/dz \begin{pmatrix} n(t_{12}) \\ n'(t_{12}) \end{pmatrix}$$

$$\frac{\mathbf{D}}{1} \qquad \mathbf{v}_n \rightarrow \qquad 2 \qquad \Delta E_{wp} \sim 1/\Delta t_{12}$$

If for time Δt_{12} : $\mu \Delta B > \Delta E_{wp} \quad \leftarrow \text{Decoh}$

$$\mu G \upsilon \Delta t_{12} > \frac{\hbar}{\Delta t_{12}}$$
Minimal gradient:
$$G > \frac{\hbar}{\mu \upsilon (\Delta t_{12})^2} = \frac{\hbar \upsilon}{\mu (\Delta z_{12})^2}$$





Simplified *B*-description of UCN τ (*LANL*)



With $\kappa \cong 4 \times 10^{-5} \mu$ this can explain the *n* lifetime difference between UCN τ and NIST experiments (even, if $B' \neq 0$).

 $P_{n'}(t) \cong rac{2\eta^2}{\mu^2} = 3.2 imes 10^{-9}$ (realistic UCNau simulations can be useful)

By braking of coherence, the wave packet will collapse (n, n') system to pure interaction states \rightarrow either to remain as n or become n'

$$\frac{dB}{dx} > \frac{\hbar}{\mu \nu (\Delta t)^2} = \frac{\hbar \nu}{\mu (\Delta x)^2}$$

For UCN gravity trap: $\Delta t = 0.1 s$ and v = 3 m/sdecoherence collapse requires dB/dx > 3.7 mG/m

In the magnetically un-shielded UCN gravity trap expts. such gradients can lead to in-volume loss of neutrons due to $n \rightarrow n'$.

Magnetic gradients inside the buildings:

In my Stockholm hotel room:

$$\frac{dB}{dx} = 33 \ \frac{mG}{m}; \quad \frac{dB}{dy} \uparrow = 170 \ \frac{mG}{m}; \quad \frac{dB}{dz} = 120 \ \frac{mG}{m}$$

In this lecture hall:

$$\frac{dB}{dx} = 20 \ \frac{mG}{m}; \quad \frac{dB}{dy} \uparrow = 140 \ \frac{mG}{m}; \quad \frac{dB}{dz} = 50 \ \frac{mG}{m}$$

How TMM effects can be measured?

First method: with high-intensity cold beam by the regeneration method using high magnetic field gradients:



E.g. high-field superconducting magnets with neutron absorber in the center, or special gradient magnets (discussed by Leah Broussard today). Due to TMM the probability of $n \rightarrow n'$ transformation doesn't depend on mag. field *B*.

Second TMM detection method:

All previous was in vacuum. Now, let's consider another possibility: neutron propagation in gas, e.g. in air

$$\mathcal{H} = \begin{pmatrix} V_F - \mu B & \epsilon + \kappa B \\ \epsilon + \kappa B & 0 \end{pmatrix} \rightarrow \begin{pmatrix} 0 & \epsilon + \kappa B \\ \epsilon + \kappa B & 0 \end{pmatrix}$$

$$V_F(air) \sim 0.12$$
 neV; at $B^* = V_F/\mu pprox 20G
ightarrow \Omega = d$

With constant and uniform magnetic field for one of polarizations of neutron beam one can compensate effect of V_F

Elastic scattering at angle $\theta \neq 0$ and absorption will work decoherently. However, in air at STP these result in ~5% intensity loss per m

$$P_{n \to n'}(t) = \frac{d^2}{\Omega^2} \sin^2(\Omega t) = \sin^2[(\epsilon + \kappa B)t]$$

It is like "quasi-free" unsuppressed oscillations

$$P_{n \to n'}(t) = \sin^2[(\epsilon + \kappa B)t] \approx \sin^2[\kappa Bt]$$

Probability grows quadratically with B (with gas compensation) and with time without decoherence collapse; it will have a maximum at $\kappa Bt/\hbar = \pi/2$



With $\kappa = 10^{-5}\mu$ at one polarization one can see large oscillation effect with UCN at short distances a la Serebrov's $\bar{\nu}$

Should be also easy measurable effect with cold beam e.g. at 1000 m/s. Beam-gas elastic scattering and absorption will result is some losses. Hopefully effect still can be easy visible. Gas/pressure can be optimized. In MM model $n \rightarrow n'$ with *n*TMM + compensated gas

$$P_{n \to n'}(t) = \sin^2[(\epsilon + \kappa B)t]$$

- $P_{n \rightarrow n'} \lesssim 10^{-6}$ in vacuum without magnetic field;
- Probability increases with B as $\sim B^2$;
- Probability is for only one gas-compensated polarization;
- More interesting effects if $B' \neq 0$ will be found

ESS ANNI beam at initial low intensity opens up interesting possibilities for $n \rightarrow n'$ search with simple and non-expensive experiments (more details in the talk of Josh Barrow today)