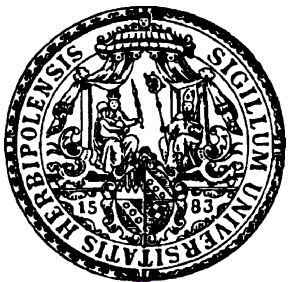
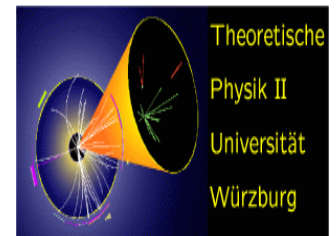


Testing supersymmetric neutrino mass models at LHC (and beyond)

Werner Porod



Universität Würzburg



- Introduction
- Flavour violating decays of SUSY particles
 - Dirac neutrinos
 - Majorana neutrinos
 - Neutrino masses via R-parity violation
- "Reconstruction of supersymmetric models" or
"Can we access heavy right-handed neutrinos"
- Conclusions

- local SUSY implies gravity
- gauge coupling unification
- provides dark matter candidates
- solves hierarchy problem
- might help to explain observed baryon asymmetry

Neutrinos: tiny masses

$$\Delta m_{atm}^2 \simeq 3 \cdot 10^{-3} \text{ eV}^2$$

$$\Delta m_{sol}^2 \simeq 7 \cdot 10^{-5} \text{ eV}^2$$

$$^3\text{H decay: } m_\nu \lesssim 2 \text{ eV}$$

Neutrinos: large mixings

$$|\tan \theta_{atm}|^2 \simeq 1$$

$$|\tan \theta_{sol}|^2 \simeq 0.4$$

$$|U_{e3}|^2 \lesssim 0.05$$

strong bounds for charged leptons

$$BR(\mu \rightarrow e\gamma) \lesssim 1.2 \cdot 10^{-11}$$

$$BR(\mu^- \rightarrow e^- e^+ e^-) \lesssim 10^{-12}$$

$$BR(\tau \rightarrow e\gamma) \lesssim 1.1 \cdot 10^{-7}$$

$$BR(\tau \rightarrow \mu\gamma) \lesssim 6.8 \cdot 10^{-8}$$

$$BR(\tau \rightarrow lll') \lesssim O(10^{-8}) \quad (l, l' = e, \mu)$$

$$|d_e| \lesssim 10^{-27} \text{ e cm}, |d_\mu| \lesssim 1.5 \cdot 10^{-18} \text{ e cm}, |d_\tau| \lesssim 1.5 \cdot 10^{-16} \text{ e cm}$$

SUSY contributions to anomalous magnetic moments

$$|\Delta a_e| \leq 10^{-12}, \quad 0 \leq \Delta a_\mu \leq 43 \cdot 10^{-10}, \quad |\Delta a_\tau| \leq 0.058$$

analog to leptons or quarks

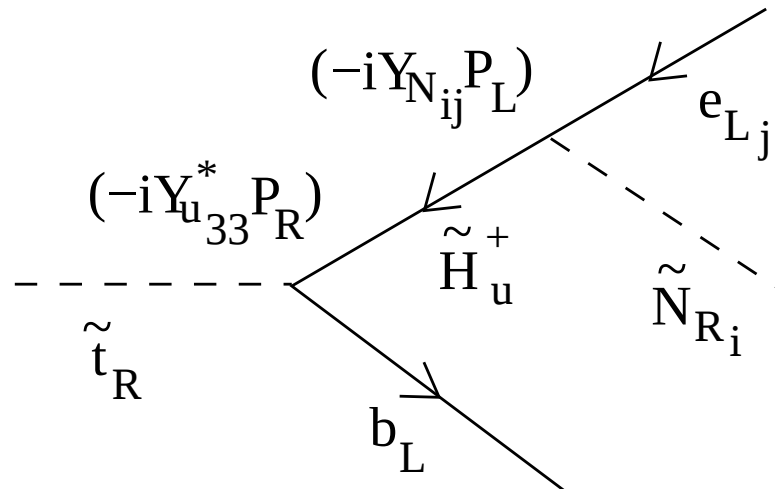
$$Y_\nu H \bar{\nu}_L \nu_R \rightarrow Y_\nu v \bar{\nu}_L \nu_R = m_\nu \bar{\nu}_L \nu_R$$

requires $Y_\nu \ll Y_e$

⇒ no impact for future collider experiments

Exception: $\tilde{\nu}_R$ is LSP and thus a candidate for dark matter

⇒ long lived NLSP



NLSP	decay modes	comments
$\tilde{\chi}_1^0$	$\nu\tilde{\nu}_R$	invisible
$\tilde{\chi}_1^+$	$l^+\tilde{\nu}_R$	mainly inside detector but finite decay length
$\tilde{f}$	$f\nu\tilde{\nu}_R, f'l\tilde{\nu}_R$	inside/outside detector
$\tilde{g}$	$q\bar{q}\nu\tilde{\nu}_R, q'\bar{q}l\tilde{\nu}_R$	mainly outside detector

in mSUGRA: $m_{\tilde{\nu}_R} = m_0$

in GMSB: $m_{\tilde{\nu}_R} \simeq 0$

Assume the existence of very heavy ν_R :

$$\Rightarrow \begin{pmatrix} \nu_L & \nu_R \end{pmatrix} \begin{pmatrix} 0 & Y_\nu v \\ Y_\nu v & M_R \end{pmatrix} \begin{pmatrix} \nu_L \\ \nu_R \end{pmatrix}$$

$Y_\nu v$ in the range 1 MeV - 100 GeV.

$$\begin{aligned} m_1 &\simeq -\frac{h_\nu^2 v^2}{M_R} \\ m_2 &\simeq M_R \end{aligned}$$

expect: $10^7 \text{ GeV} \lesssim M_R \lesssim 10^{14} \text{ GeV}$.

3 generations:

$$m_\nu \simeq -(Y_\nu^T v) M_R^{-1} (Y_\nu v)$$

much more parameters than observables

* P. Minkowski, Phys. Lett. B **67** (1977) 421; T. Yanagida, KEK-report 79-18 (1979);
M. Gell-Mann, P. Ramond, R. Slansky, in *Supergravity*, North Holland (1979), p. 315

Y_ν contributes to RGEs :

$$\begin{aligned} 16\pi^2 \dot{A}_e &= \dots + 2Y_\nu A_\nu \\ 16\pi^2 \dot{M}_L^2 &= \dots + M_L^2 Y_\nu^\dagger Y_\nu + Y_\nu^\dagger Y_\nu M_L^2 \end{aligned}$$

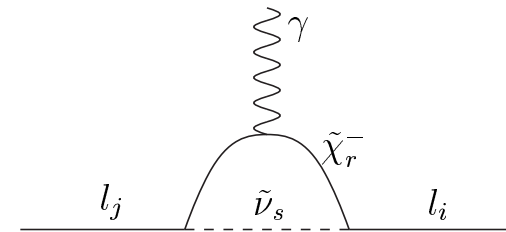
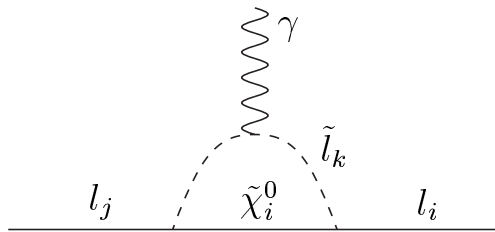
Off-diagonal elements

- rare lepton decays
- flavour violating SUSY decays

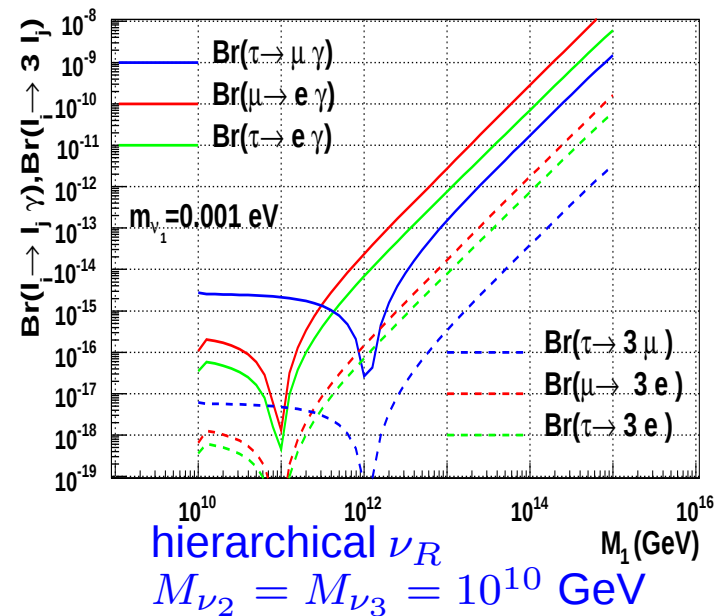
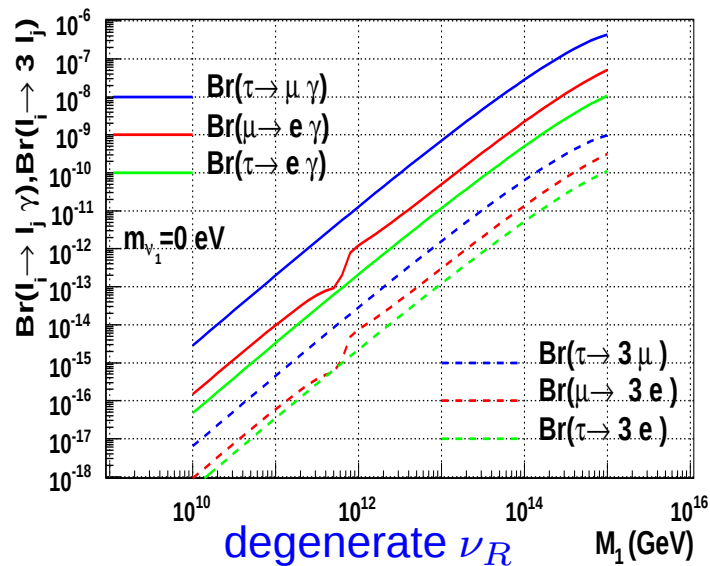
Change diagonal elements $\Rightarrow$ ratio of masses change

$$\tilde{\chi}_i^0 = (\tilde{\gamma}, \tilde{z}^0, \tilde{h}_d^0, \tilde{h}_d^0)$$

$$\tilde{\chi}_i^- = (\tilde{w}^-, \tilde{h}^-)$$



⇒ anomalous magnetic moments, electric dipole moments, rare lepton decays



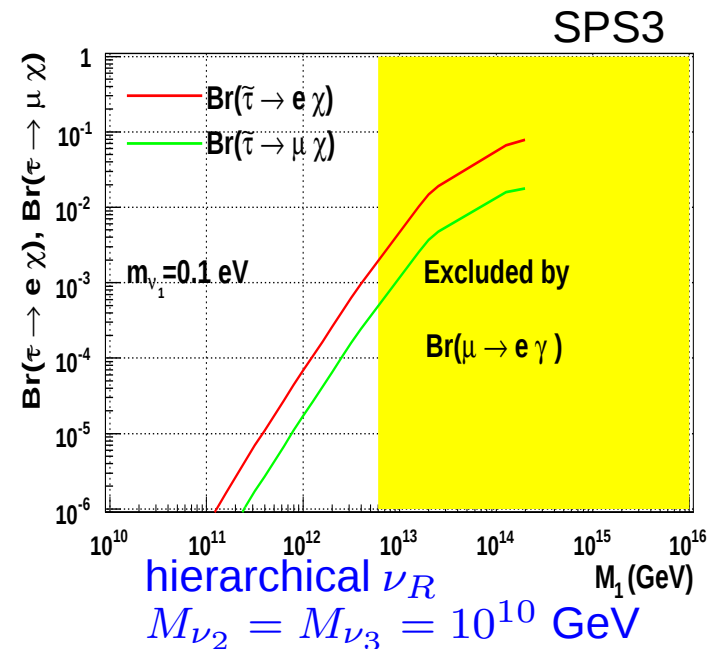
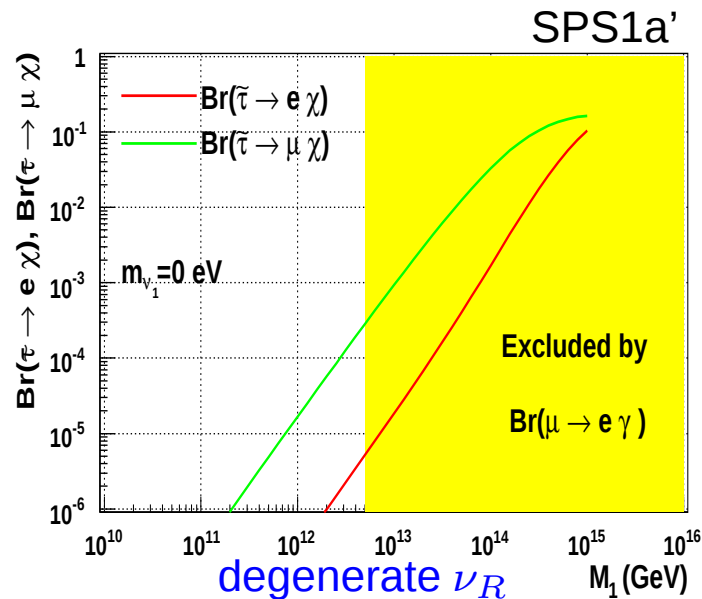
M. Hirsch et al., arXiv:0804.4072

$$(\Delta M_{\tilde{L}}^2)_{ij} = -\frac{1}{8\pi^2}(3m_0^2 + A_0^2)(Y_\nu^\dagger LY_\nu)_{ij} \ , \ (\Delta A_l)_{ij} = -\frac{3}{8\pi^2}A_0 Y_{l_i}(Y_\nu^\dagger LY_\nu)_{ij} \ ,$$

$$L_{kl} = \log\left(\frac{M_{GUT}}{M_k}\right)\delta_{kl}$$

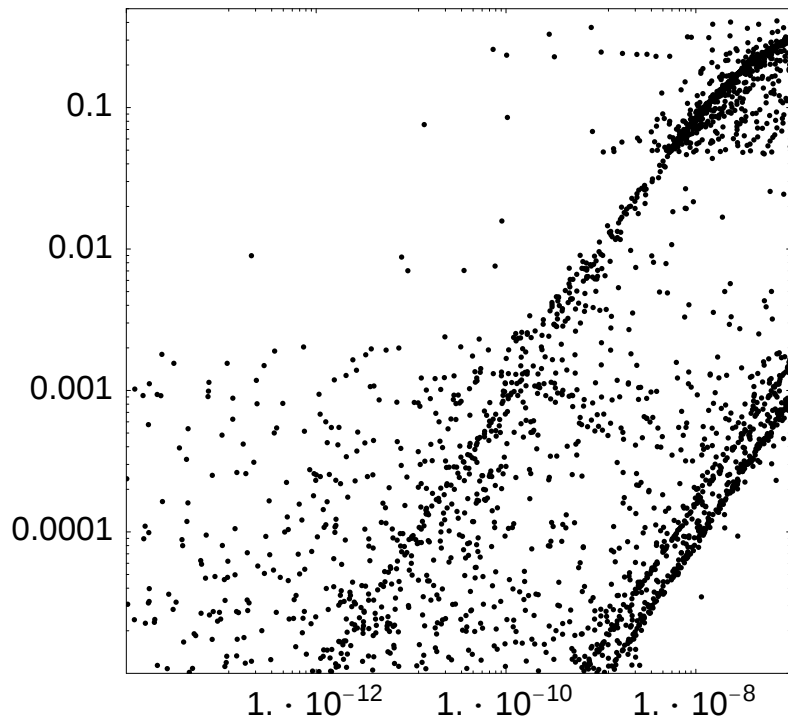
$$\frac{BR(\tilde{\tau}_2 \rightarrow e + \chi_1^0)}{BR(\tilde{\tau}_2 \rightarrow \mu + \chi_1^0)} \simeq \left(\frac{\theta_{13}}{\theta_{23}}\right)^2 \simeq \left(\frac{(\Delta M_{\tilde{L}}^2)_{13}}{(\Delta M_{\tilde{L}}^2)_{23}}\right)^2 \ ,$$

(neglecting L - R mixing)



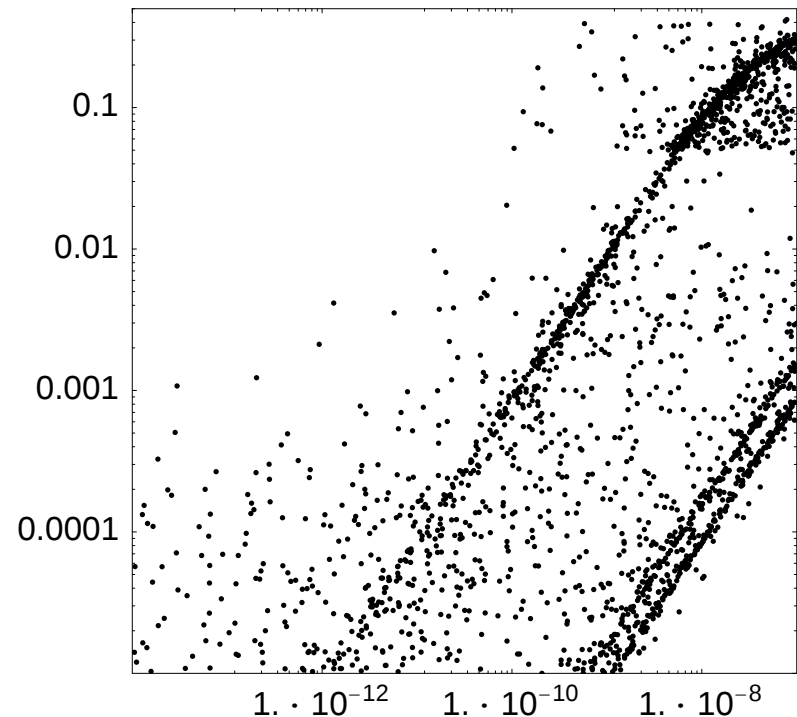
M. Hirsch et al., arXiv:0804.4072

$\text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 e^\pm \tau^\mp)$



$\text{BR}(\tau \rightarrow e\gamma)$

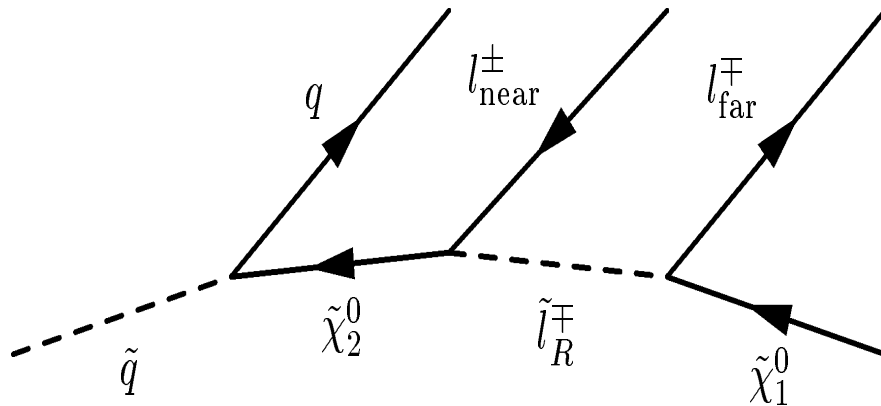
$\text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 \mu^\pm \tau^\mp)$



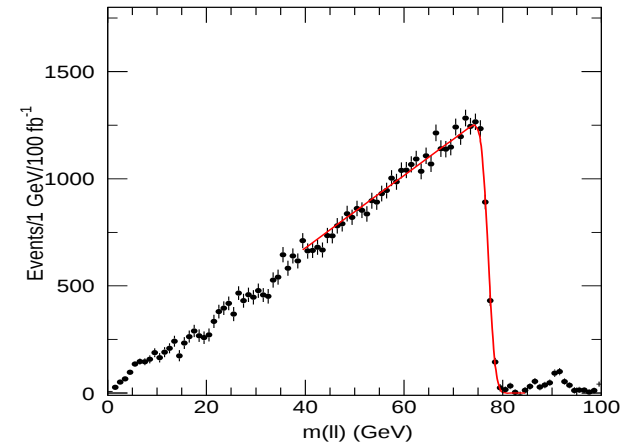
$\text{BR}(\tau \rightarrow \mu\gamma)$

Variations around SPS1a

$(M_0 = 100 \text{ GeV}, M_{1/2} = 250 \text{ GeV}, A_0 = -100 \text{ GeV}, \tan \beta = 10)$



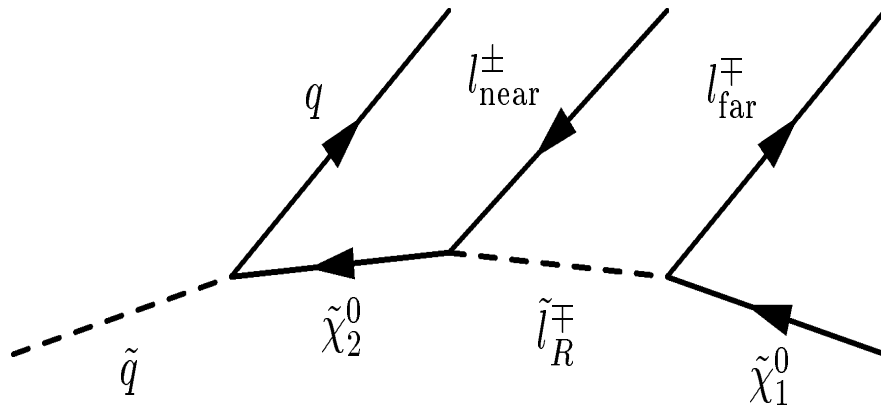
G. Polesello



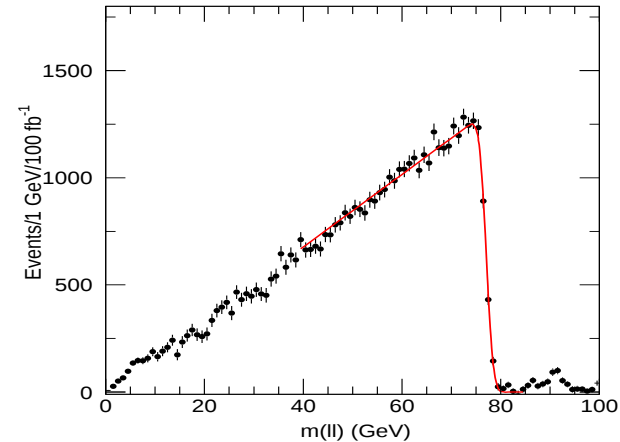
5 kinematical observables depending on 4 SUSY masses

e.g.: $m(ll) = 77.02 \pm 0.05 \pm 0.08$

$\Rightarrow$ mass determination within 2-5%



G. Polesello



5 kinematical observables depending on 4 SUSY masses

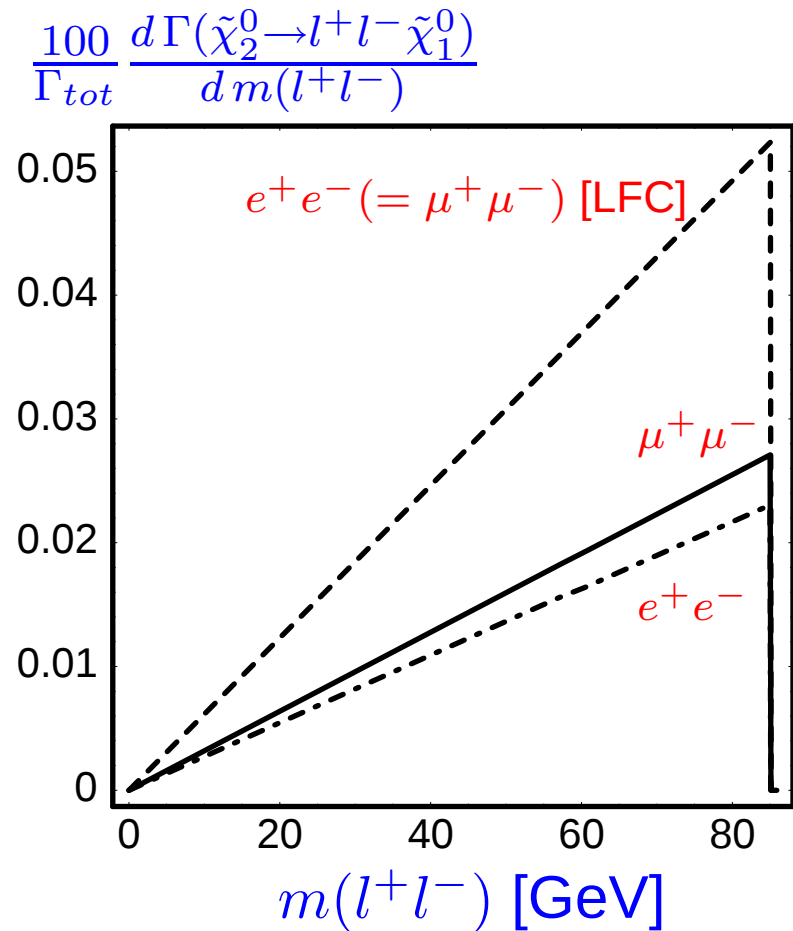
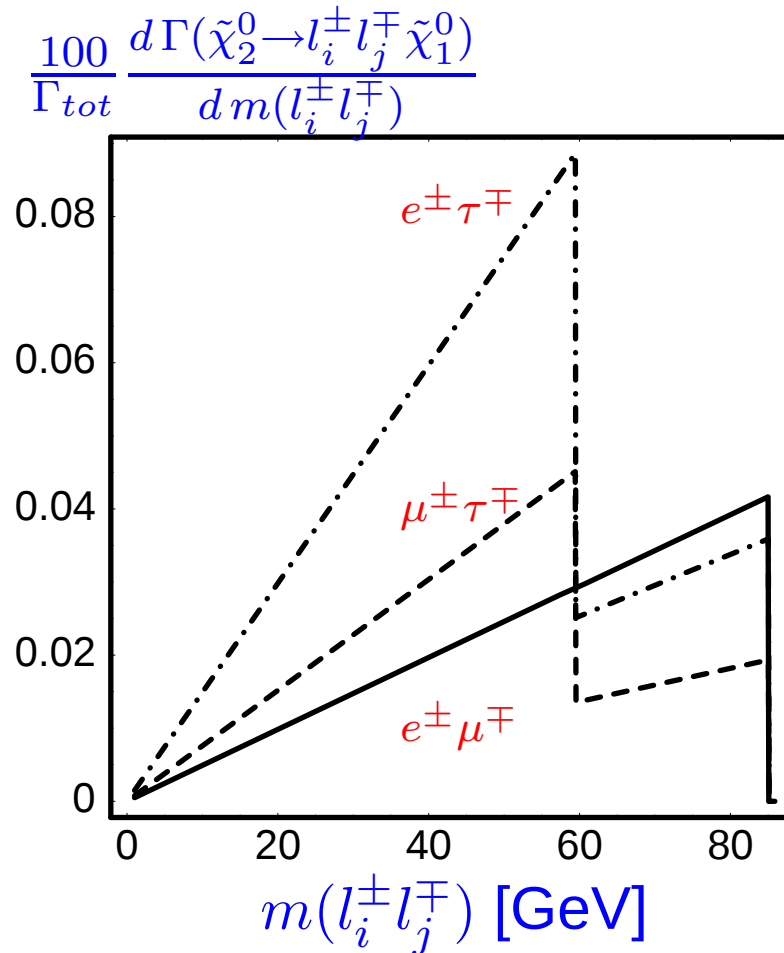
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For background suppression

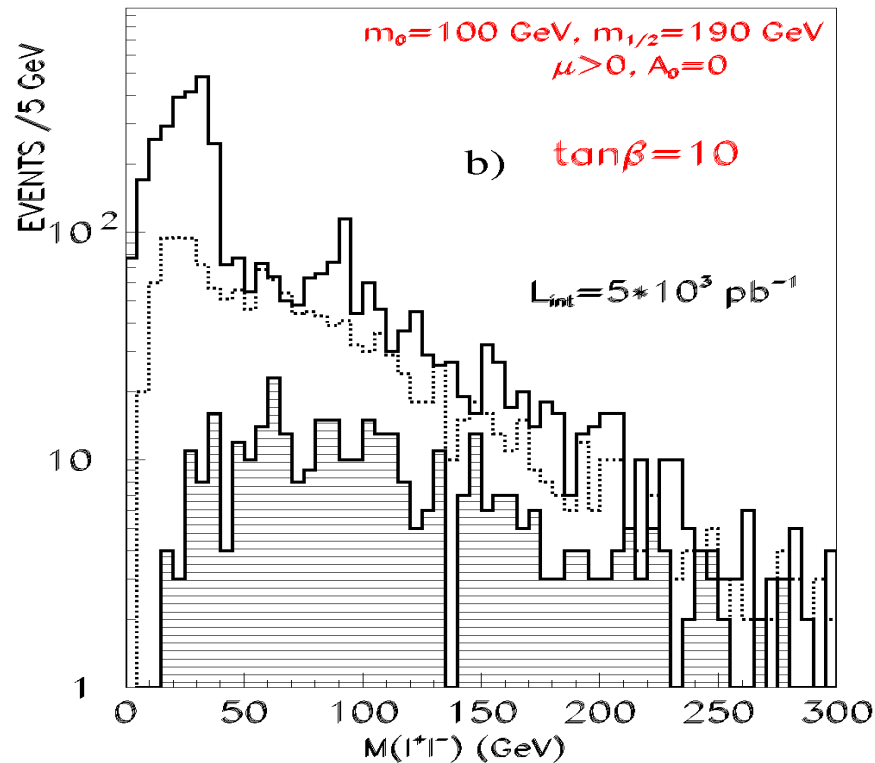
$$N(e^+e^-) + N(\mu^+\mu^-) - N(e^+\mu^-) - N(\mu^+e^-)$$

assumptions: — no flavour violation

$$\text{— BR}(\tilde{\chi}_2^0 \rightarrow \tilde{e}_R e) = \text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\mu}_R \mu)$$



A. Bartl et al., Eur. Phys. J. C **46** (2006) 783



D. Denegri, W. Majerotto und L. Rurua
PRD **60** (1999) 035008

increase of experimental uncertainty
on the edges by an order of magnitude

see also I. Hinchliffe and F. E. Paige, PRD **63** (2001) 115006;
J. Hisano, R. Kitano and M. M. Nojiri, PRD **65** (2002) 116002

R-parity: $(-1)^{3(B-L)+2S}$

bilinear R-parity violation:

$$W = W_{MSSM} + \epsilon_i \hat{L}_i \hat{H}_u$$

⇒ mixings between SM and SUSY particles

Gravitino as dark matter*

generic prediction of GMSB : light Gravitino LSP

relic density requires: $100 \text{ eV} \lesssim m_{3/2} \lesssim 1000 \text{ eV}$

NLSP: $\tilde{\chi}_1^0$ oder $\tilde{l}_R$ ($l = e, \mu, \tau$)

*S. Borgani, A. Masiero, M. Yamaguchi, PLB**386** (1996) 189

F. Takayama and M. Yamaguchi, PLB **485** (2000) 388

M. Hirsch, W. P., D. Restrepo, JHEP **0503**, 062 (2005)

Neutrino physics*:

- neutrino masses via ν - $\tilde{\chi}_i^0$ mixing
- neutrino mixing angles in terms of R-parity violating couplings

* M. Hirsch, et al. Phys. Rev. D **62**, 113008 (2000)

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- neutrino mixing angles in terms of R-parity violating couplings

Neutralino decays:

- dominant R-parity violating decays: $\tilde{\chi}_1^0 \rightarrow W^\pm l_i^\mp$, $\tilde{\chi}_1^0 \rightarrow Z \nu_i$, $\tilde{\chi}_1^0 \rightarrow \nu \tau^+ l_i^-$
- R-parity conserving decay: $\Gamma(\tilde{\chi}_1^0 \rightarrow \tilde{G} \gamma) \propto \left(\frac{m_{\tilde{\chi}_1^0}}{100 \text{ GeV}} \right)^5 \left(\frac{100 \text{ eV}}{m_{3/2}} \right)^2$
- $\Gamma(\tilde{\chi}_1^0) \simeq (10^{-4} - 10^{-2}) \text{ eV} \Rightarrow$ decay length of $O(10 \text{ } \mu\text{m}) - O(1 \text{ mm})$

* M. Hirsch, et al. Phys. Rev. D **62**, 113008 (2000)

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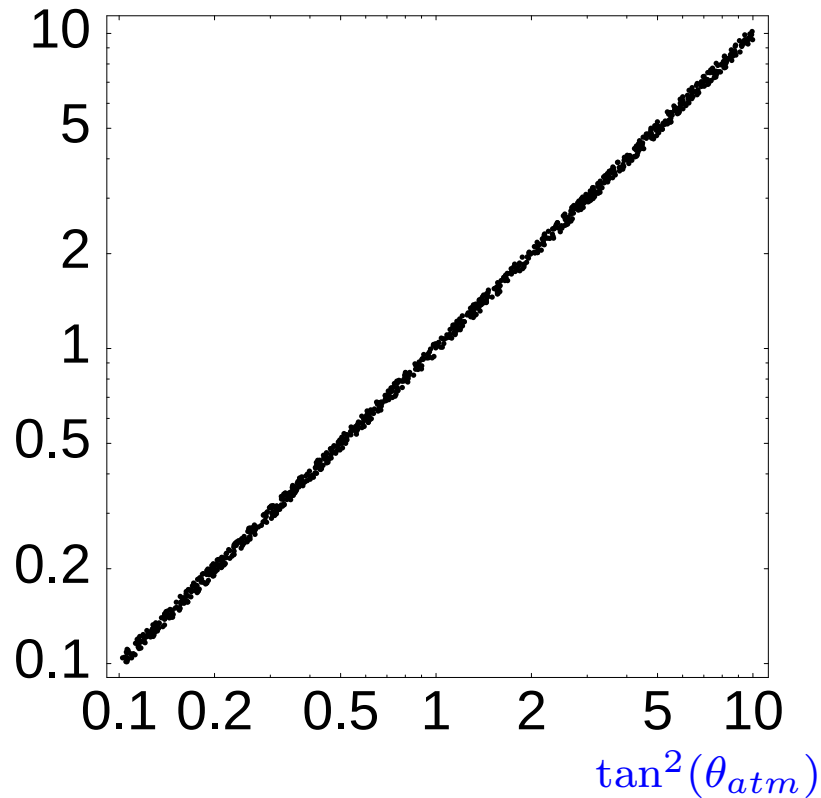
Gravitino:

- $\tilde{G}$ eventually decays into $\nu \gamma$
but: $\tau(\tilde{G}) \sim O(10^{30})$ · life time of the universe

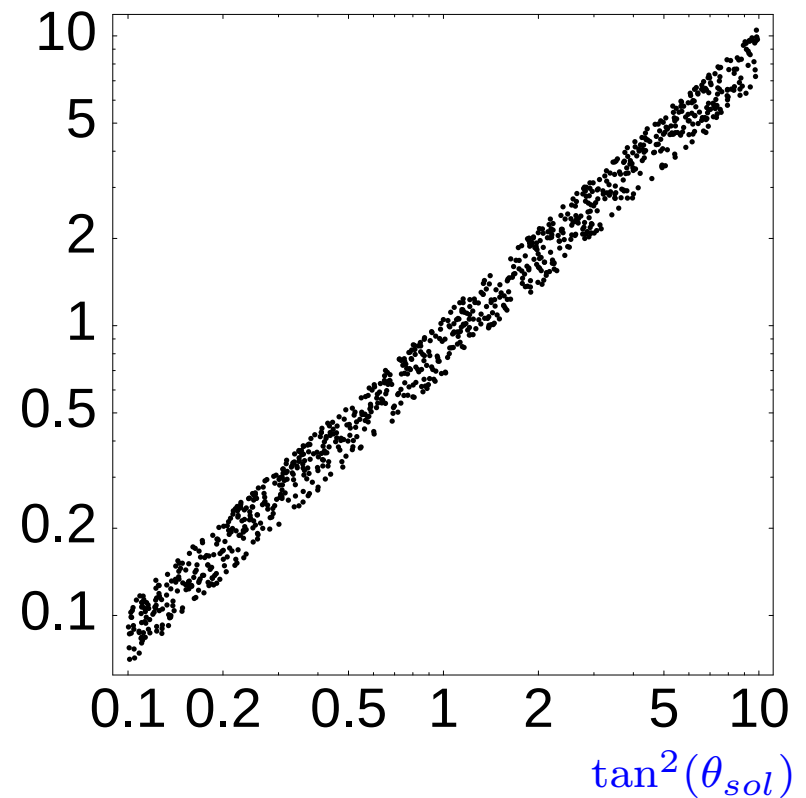
* M. Hirsch, et al. Phys. Rev. D **62**, 113008 (2000)

Correlations

$$\text{BR}(\tilde{\chi}_1^0 \rightarrow W\mu) / \text{BR}(\tilde{\chi}_1^0 \rightarrow W\tau)$$

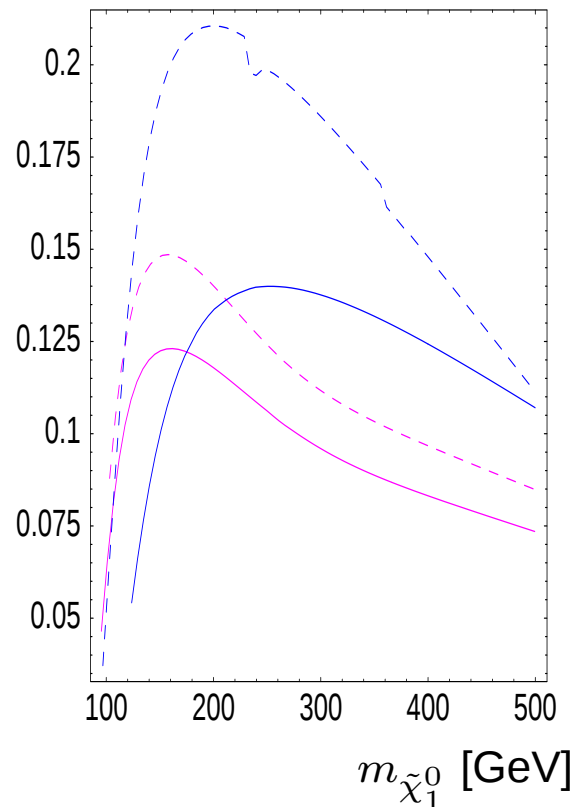


$$\text{BR}(\tilde{\chi}_1^0 \rightarrow \nu e \tau) / \text{BR}(\tilde{\chi}_1^0 \rightarrow \nu \mu \tau)$$

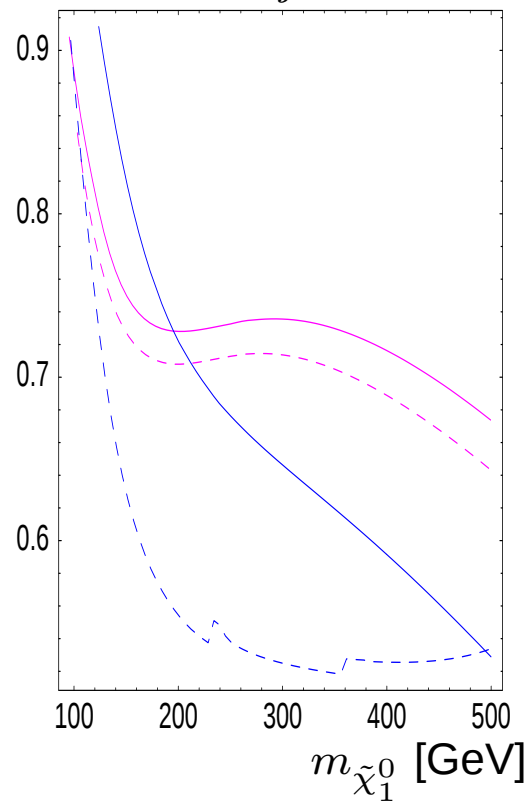


W. P. et al., Phys. Rev. D **63**, 115004 (2001)

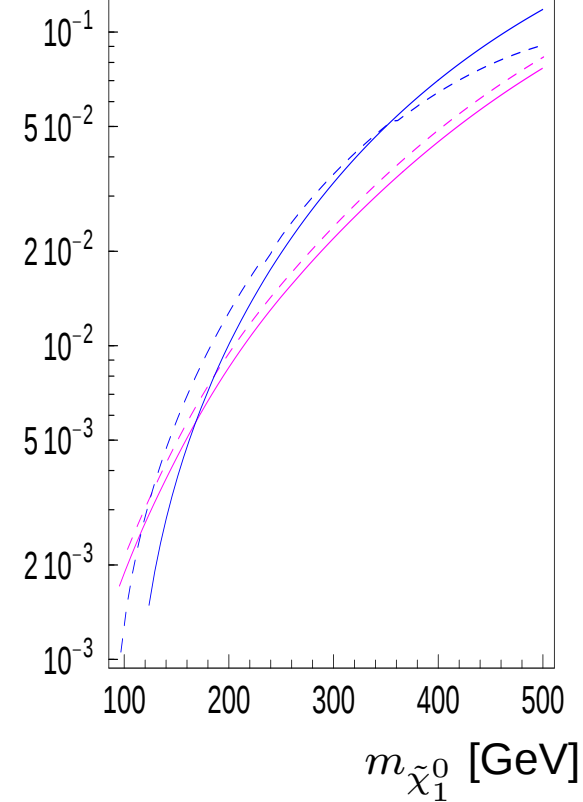
$$\text{BR}(\tilde{\chi}_1^0 \rightarrow \sum_i W l_i)$$



$$\text{BR}(\tilde{\chi}_1^0 \rightarrow \sum_{ij} \nu_i \tau l_j)$$



$$\text{BR}(\tilde{\chi}_1^0 \rightarrow \tilde{G} \gamma)$$



— $\tan \beta = 10, \mu > 0$, - - $\tan \beta = 10, \mu < 0$, — $\tan \beta = 35, \mu > 0$, - - $\tan \beta = 35, \mu < 0$

$m_{3/2} = 100 \text{ eV}, n_5 = 1$

M. Hirsch, W. P. und D. Restrepo, JHEP 0503, 062 (2005)

Comments



$$\frac{m_{\tilde{\tau}_1}}{m_{\tilde{\chi}_1^0}} \propto \frac{1}{\sqrt{n_5}}$$

$\Rightarrow$ for $n_5 \geq 3$ hardly points with $\tilde{\chi}_1^0$ NLSP



$\tilde{l}_R$ NLSPs: $\text{BR}(l\nu) > \text{BR}(l\tilde{G})$

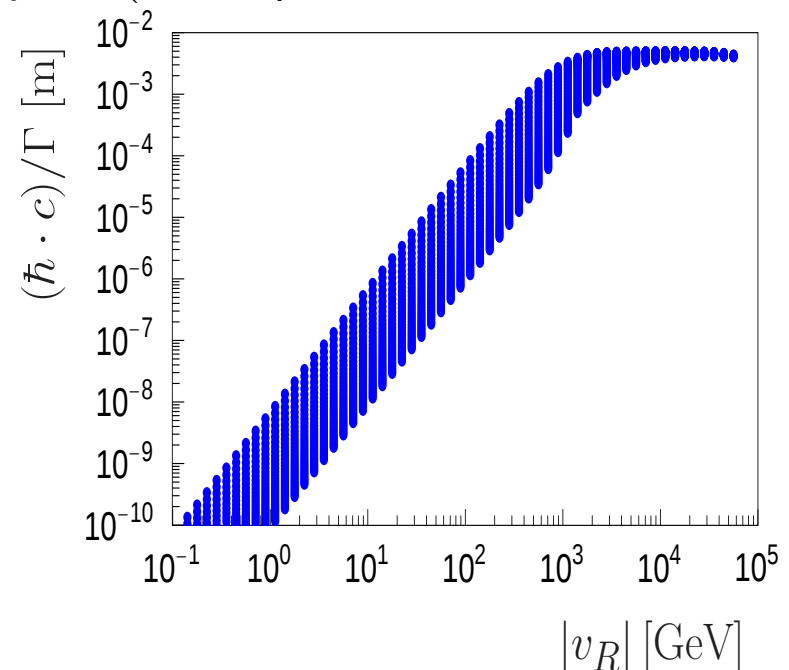
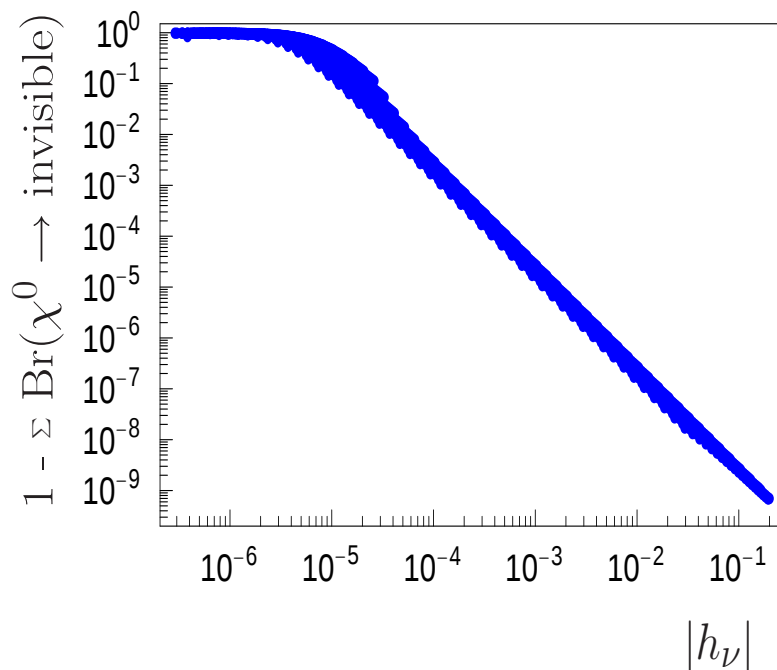


$n_5 = 2$: $\text{BR}(\tilde{G}\gamma)$ reduced by a factor 2-3

Spontaneous R-Parity violation:

$$\begin{aligned}\mathcal{W} = & h_U^{ij} \hat{Q}_i \hat{U}_j \hat{H}_u + h_D^{ij} \hat{Q}_i \hat{D}_j \hat{H}_d + h_E^{ij} \hat{L}_i \hat{E}_j \hat{H}_d \\ & + h_\nu^i \hat{L}_i \hat{\nu}^c \hat{H}_u - h_0 \hat{H}_d \hat{H}_u \hat{\Phi} + h \hat{\Phi} \hat{\nu}^c \hat{S} + \frac{\lambda}{3!} \hat{\Phi}^3\end{aligned}$$

SPS1a', $\epsilon_i = h_{\nu,i} v_R = O(1 \text{ GeV})$



M. Hirsch, W.P., PRD74 (2006) 055003

Two-scale picture of nature:



Two-scale picture of nature:



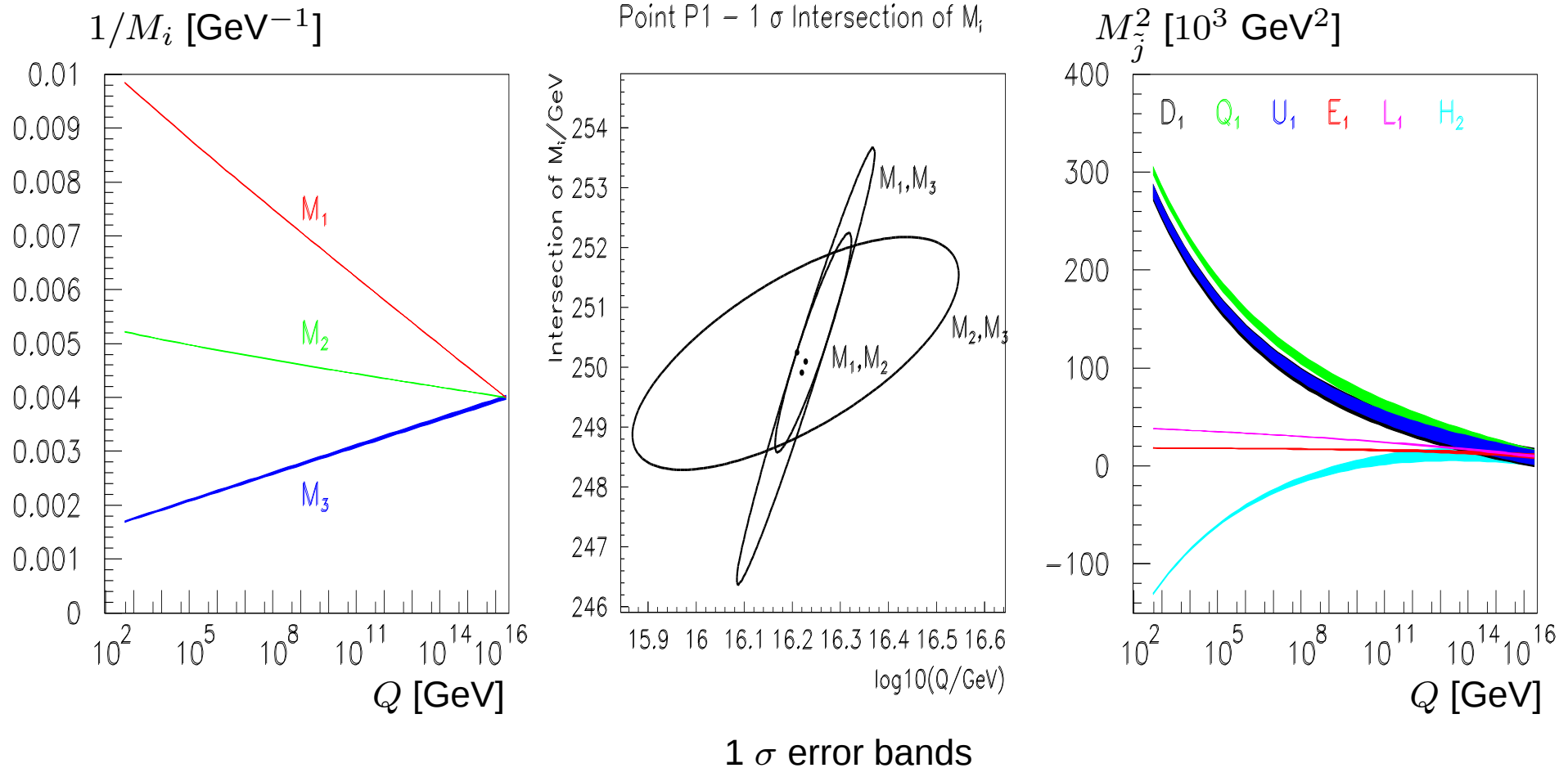
program to explore the high scale[†]:

- LHC, ILC: masses, mixings, couplings ...
 - ⊕ low energy observables: $b \rightarrow s\gamma$, $(g - 2)_\mu$, ...
- extract SUSY Lagrange parameters
- extrapolation to GUT/Planck scale*

[†] „Supersymmetry Parameter Analysis” project

* G. Blair, W. P., P. Zerwas, Phys. Rev. D **63**, 017703 (2001); Eur. Phys. J. C **27**, 263 (2003)

Extrapolation of gaugino- and scalar mass parameters

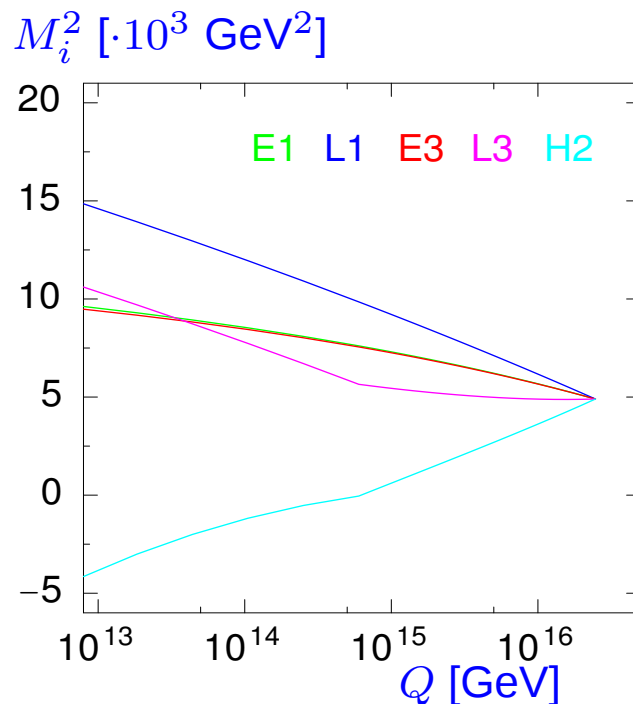


SPA: $M_{1/2} = 250$ GeV, $M_0 = 70$ GeV, $A_0 = -300$ GeV, $\tan \beta = 10$, $\text{sign}(\mu) = +$

$m_\nu \neq 0$ neutrino masses via seesaw mechanism

$\Rightarrow$ Seesaw scale $M[\nu_R] \sim 10^{10}/10^{14}$ GeV in SO(10)

$\Rightarrow$ affects evolution of 3rd generation parameters $\Rightarrow$ **kink**



A. Freitas, W. P. and P. M. Zerwas, Phys. Rev. D **72** (2005) 115002

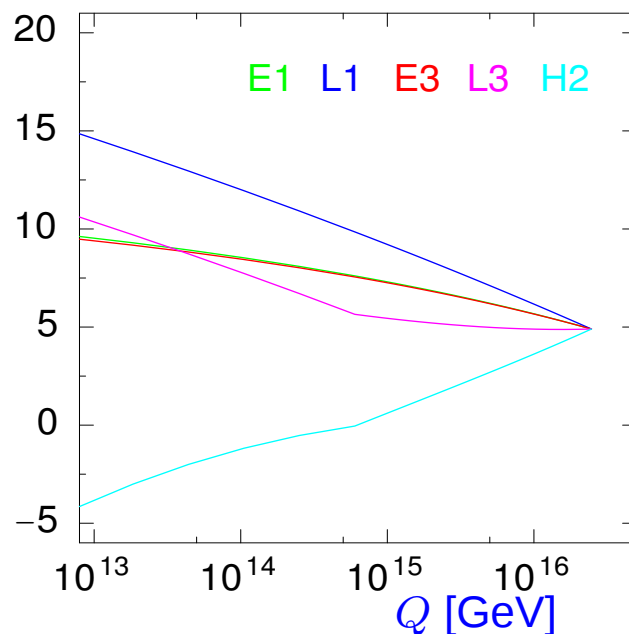
A. Freitas, F. Deppisch, W. P. and P. M. Zerwas, arXiv:0712.0361

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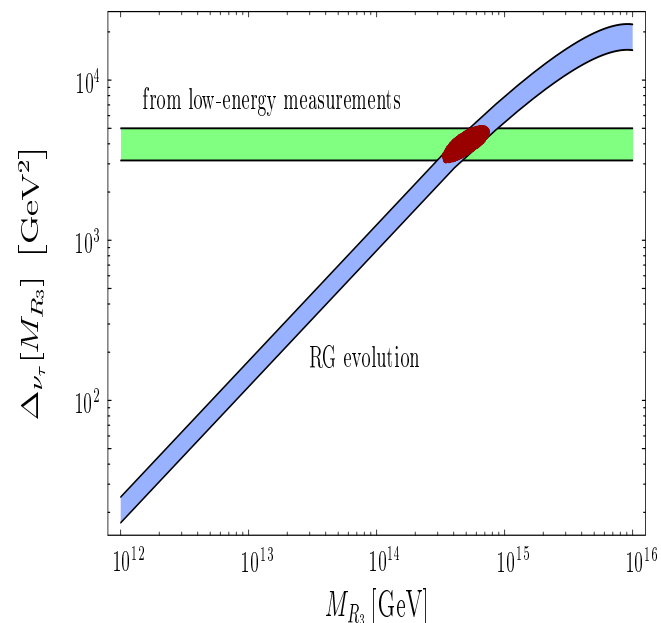
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$M_i^2 [\cdot 10^3 \text{ GeV}^2]$



$\oplus M_\nu \Rightarrow M[\nu_{R3}] \sim 10^{14}$ GeV [30%]



A. Freitas, W. P. and P. M. Zerwas, Phys. Rev. D **72** (2005) 115002

A. Freitas, F. Deppisch, W. P. and P. M. Zerwas, arXiv:0712.0361

- Neutrino physics requires extension of the Standard Model
- Supersymmetry is a very promising and well motivated extension
- $\Rightarrow$ in general lepton flavour violating decays of SUSY particles
- can be measured at the LHC
but things might turn out to be more complicated

talk by I. Borjanovic at 'Flavour in the era of LHC', Nov.'05, CERN

L=100 fb⁻¹

Fit results

Edge	Nominal Value	Fit Value	Syst. Error Energy Scale	Statistical Error
$m(ll)^{\text{edge}}$	77.077	77.024	0.08	0.05
$m(qll)^{\text{edge}}$	431.1	431.3	4.3	2.4
$m(ql)^{\text{edge}}_{\text{min}}$	302.1	300.8	3.0	1.5
$m(ql)^{\text{edge}}_{\text{max}}$	380.3	379.4	3.8	1.8
$m(qll)^{\text{thres}}$	203.0	204.6	2.0	2.8

Mass reconstruction

5 endpoints measurements, 4 unknown masses

$$\chi^2 = \sum \chi_j^2 = \sum \left[\frac{E_j^{\text{theory}}(\vec{m}) - E_j^{\text{exp}}}{\sigma_j^{\text{exp}}} \right]^2$$

$$E_j^i = E_j^{\text{nom}} + a_j^i \sigma_j^{\text{fit}} + b_j^i E_{\text{scale}}$$

$$m(\chi_1^0) = 96 \text{ GeV}$$

$$m(l_R) = 143 \text{ GeV}$$

$$m(\chi_2^0) = 177 \text{ GeV}$$

$$m(q_L) = 540 \text{ GeV}$$

$$\Delta m(\chi_1^0) = 4.8 \text{ GeV}, \quad \Delta m(\chi_2^0) = 4.7 \text{ GeV},$$

$$\Delta m(l_R) = 4.8 \text{ GeV}, \quad \Delta m(q_L) = 8.7 \text{ GeV}$$

Gjelsten, Lytken, Miller, Osland, Polesello, ATL-PHYS-2004-007

1-loop solutions to RGEs

$$m_{\tilde{e}_R}^2 = M_0^2 + D_{\mathcal{U}} + \alpha_R M_{1/2}^2 - \frac{6}{5} S' - 2s_W^2 D_{EW}$$

$$m_{\tilde{e}_L}^2 = M_0^2 - 3D_{\mathcal{U}} + \alpha_L M_{1/2}^2 + \frac{3}{5} S' - c_{2W} D_{EW}$$

$$m_{\tilde{\nu}_{eL}}^2 = M_0^2 - 3D_{\mathcal{U}} + \alpha_L M_{1/2}^2 + \frac{3}{5} S' + D_{EW}$$

$$\alpha_L = \frac{3}{10} f_1 + \frac{3}{2} f_2, \quad \alpha_R = \frac{6}{5} f_1$$

$$f_i = \frac{1}{b_i} \left(1 - \left[1 + \frac{\alpha_{\mathcal{U}}}{4\pi} b_i \log \frac{\Lambda_{\mathcal{U}}^2}{\tilde{M}^2} \right]^{-2} \right) \quad \text{with } (b_1, b_2) = \left(\frac{33}{5}, 1 \right)$$

$$m_{\tilde{\tau}_R}^2 - m_{\tilde{e}_R}^2 = m_{\tau}^2 - 2\Delta_{\tau}$$

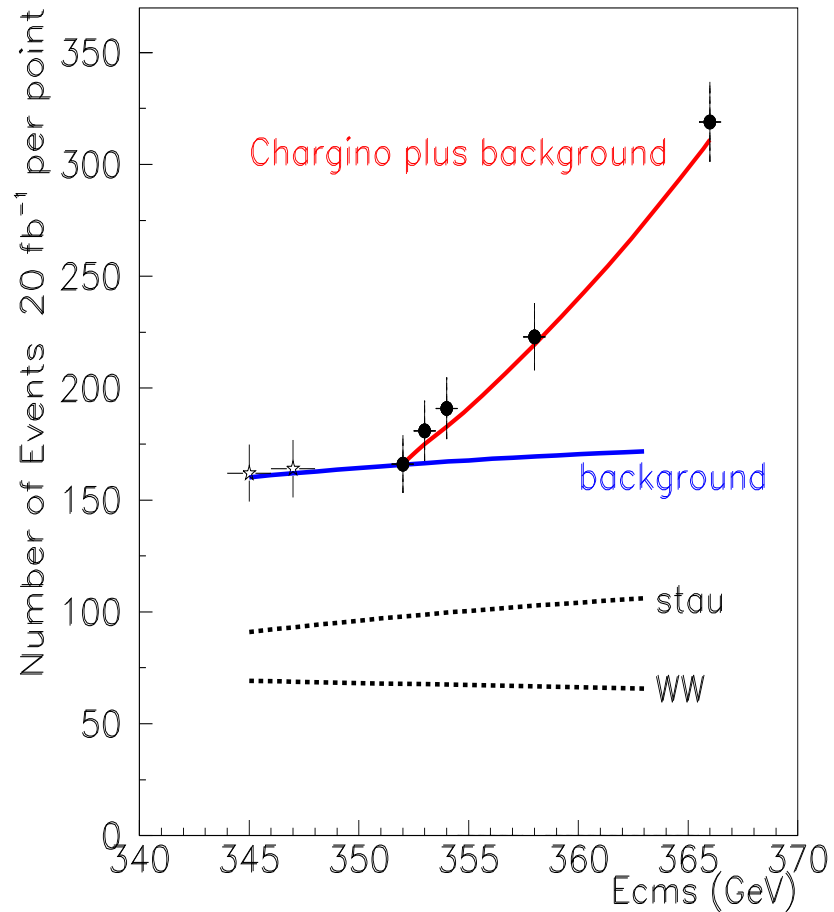
$$m_{\tilde{\tau}_L}^2 - m_{\tilde{e}_L}^2 = m_{\tau}^2 - \Delta_{\tau} - \Delta_{\nu_{\tau}}$$

$$m_{\tilde{\nu}_{\tau L}}^2 - m_{\tilde{\nu}_{eL}}^2 = -\Delta_{\tau} - \Delta_{\nu_{\tau}}$$

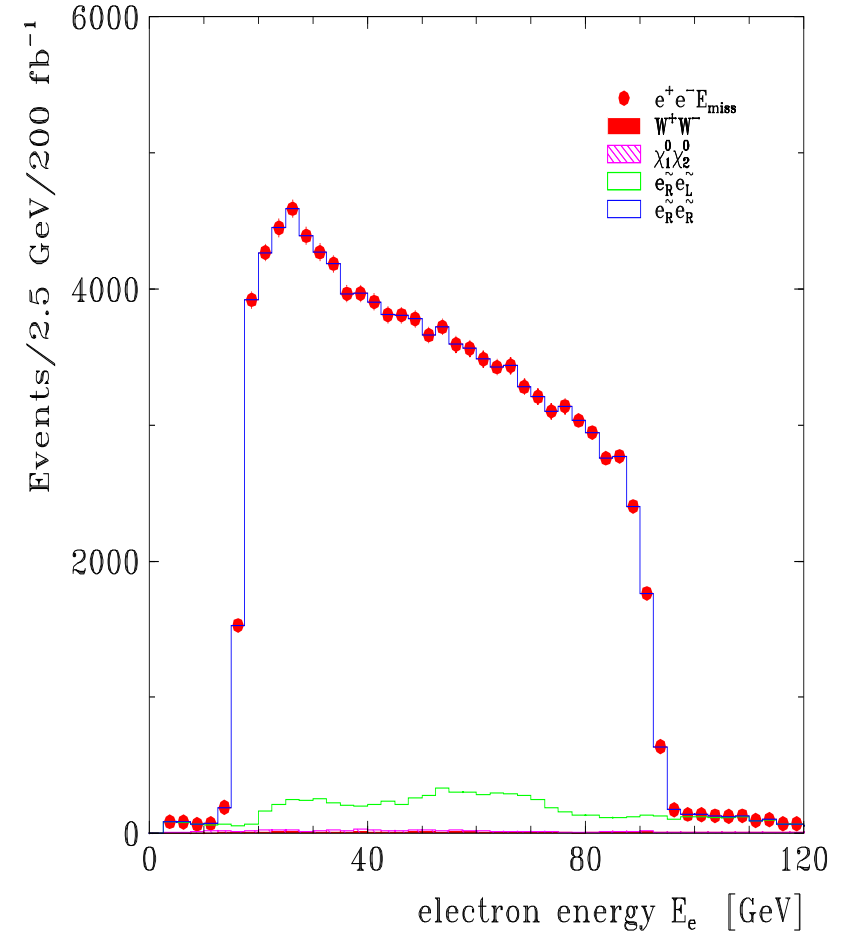
$$\Delta_{\tau} \approx \frac{m_{\tau}^2(\Lambda_{\mathcal{U}})}{8\pi^2 v_d^2} (3M_0^2 + A_0^2) \log \frac{\Lambda_{\mathcal{U}}^2}{\tilde{M}^2}$$

$$\Delta_{\nu_{\tau}} \approx \frac{m_t^2(\Lambda_{\mathcal{U}})}{8\pi^2 v_u^2} (3M_0^2 + A_0^2) \log \frac{\Lambda_{\mathcal{U}}^2}{M_{\nu_{R3}}^2}$$

$$e_R^+ e_L^- \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^-$$



$$e_L^+ e_R^- \rightarrow \tilde{e}_R^+ \tilde{e}_R^- \rightarrow e^+ \tilde{\chi}_1^0 e^- \tilde{\chi}_1^0$$



G. Blair, U. Martyn

Expected experimental accuracies

	Mass, ideal	“LHC”	“LC”	“LHC+LC”
h^0	116.0	0.25	0.05	0.05
H^0	425.0		1.5	1.5
$\tilde{\chi}_1^0$	97.7	4.8	0.05	0.05
$\tilde{\chi}_2^0$	183.9	4.7	1.2	0.08
$\tilde{\chi}_4^0$	413.9	5.1	3-5	2.5
$\tilde{\chi}_1^\pm$	183.7		0.55	0.55
$\tilde{e}_R$	125.3	4.8	0.05	0.05
$\tilde{e}_L$	189.9	5.0	0.18	0.18
$\tilde{\tau}_1$	107.9	5-8	0.24	0.24
$\tilde{q}_R$	547.2	7-12	-	5-11
$\tilde{q}_L$	564.7	8.7	-	4.9
$\tilde{t}_1$	366.5		1.9	1.9
$\tilde{b}_1$	506.3	7.5	-	5.7
$\tilde{g}$	607.1	8.0	-	6.5

$$m_0 = 70 \text{ GeV}$$

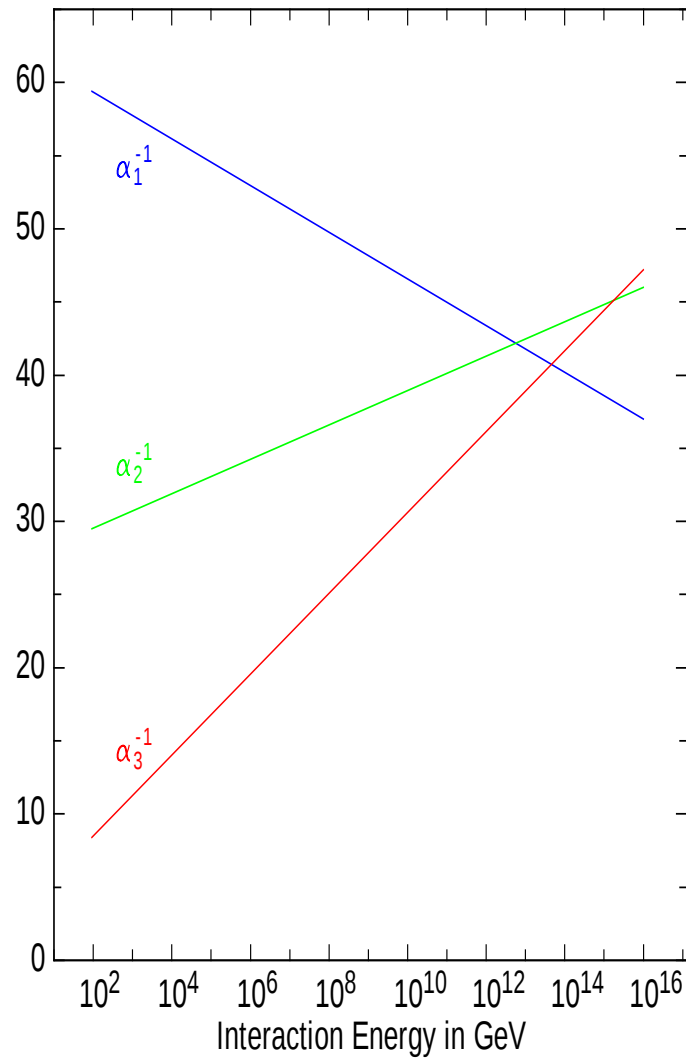
$$m_{1/2} = 250 \text{ GeV}$$

$$A_0 = -300 \text{ GeV}$$

$$\tan \beta = 10$$

$$\text{sign}(\mu) = +$$

SM



MSSM

