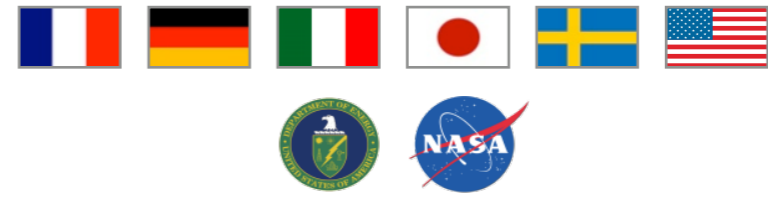
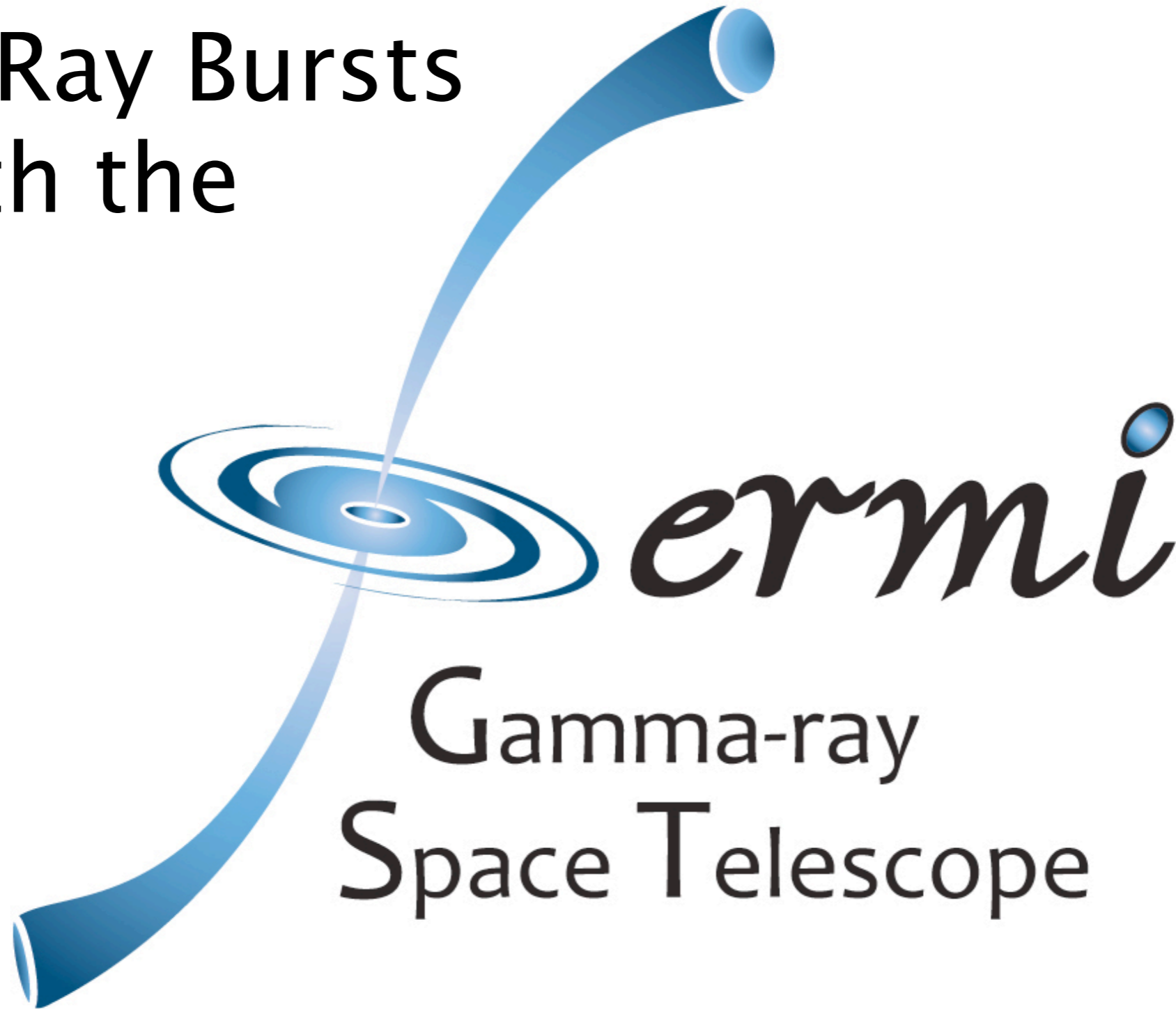


Sinead McGlynn
Particle & Astroparticle Physics
KTH



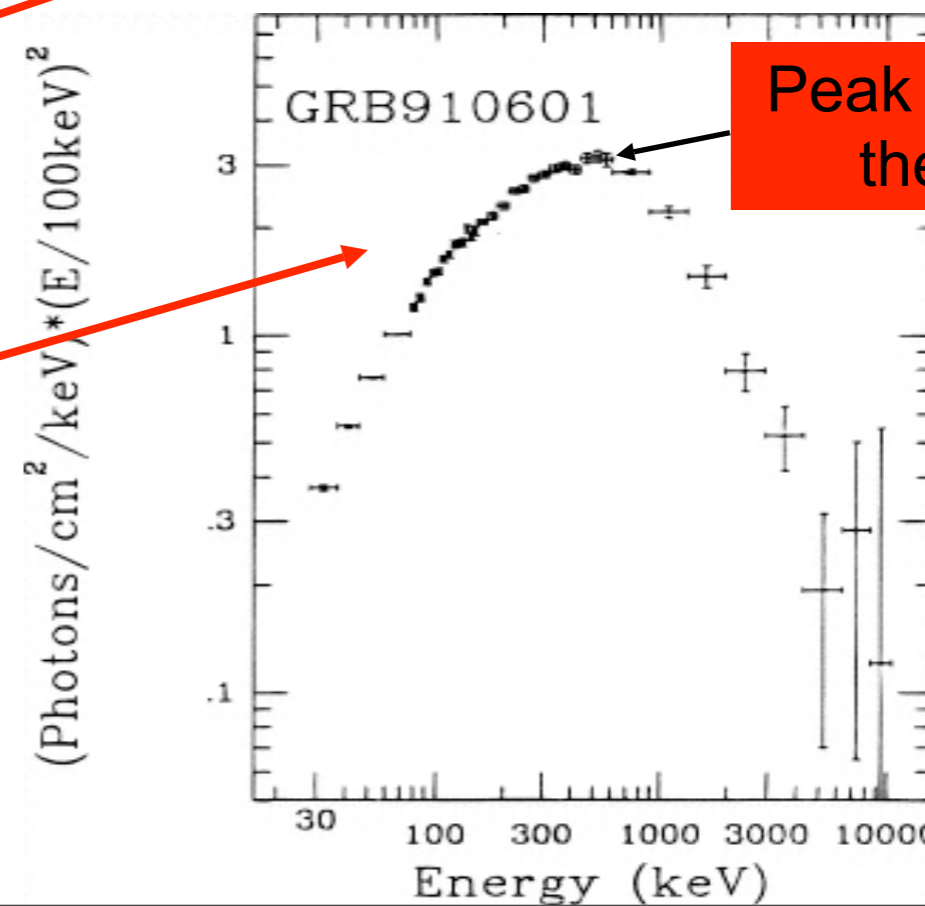
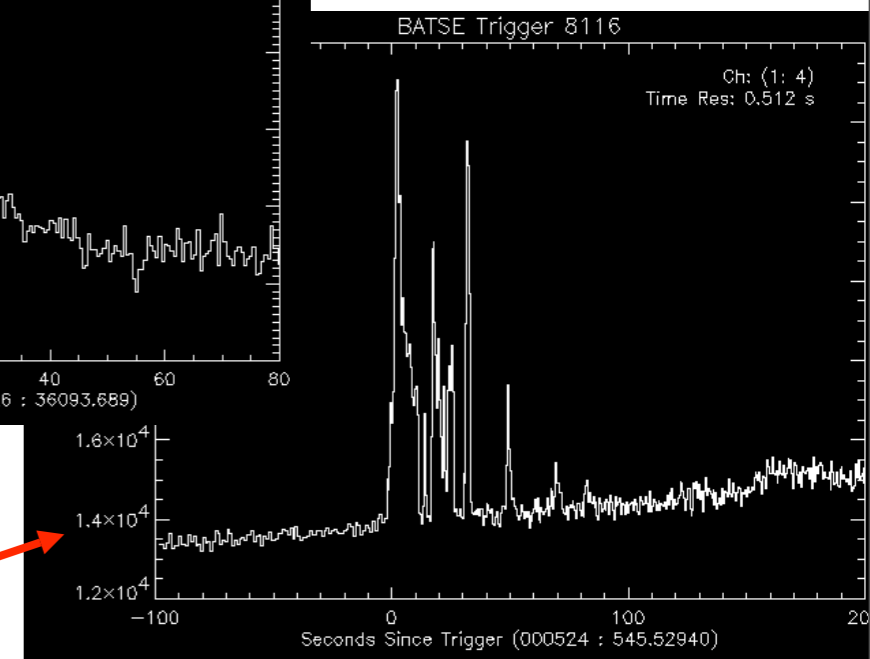
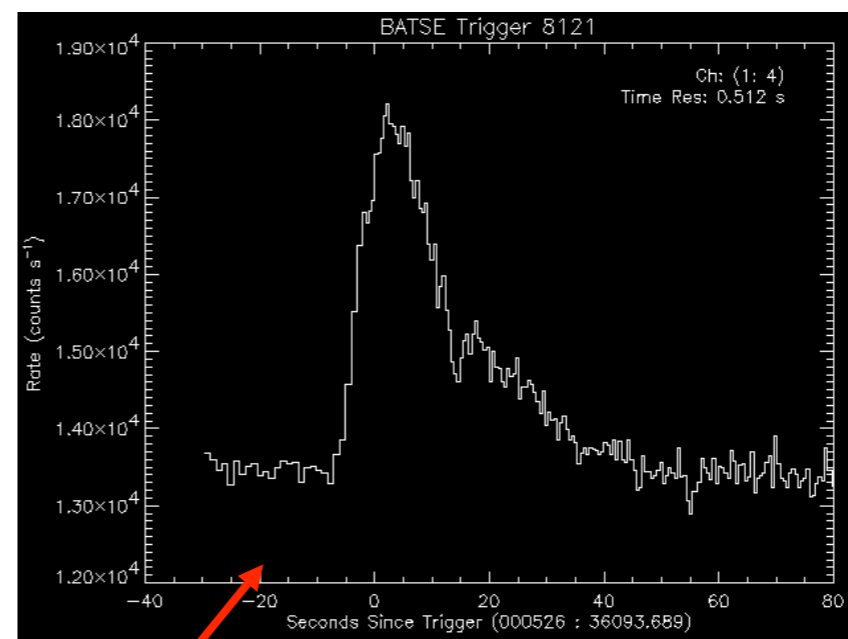
Gamma Ray Bursts with the



Gamma-ray
Space Telescope

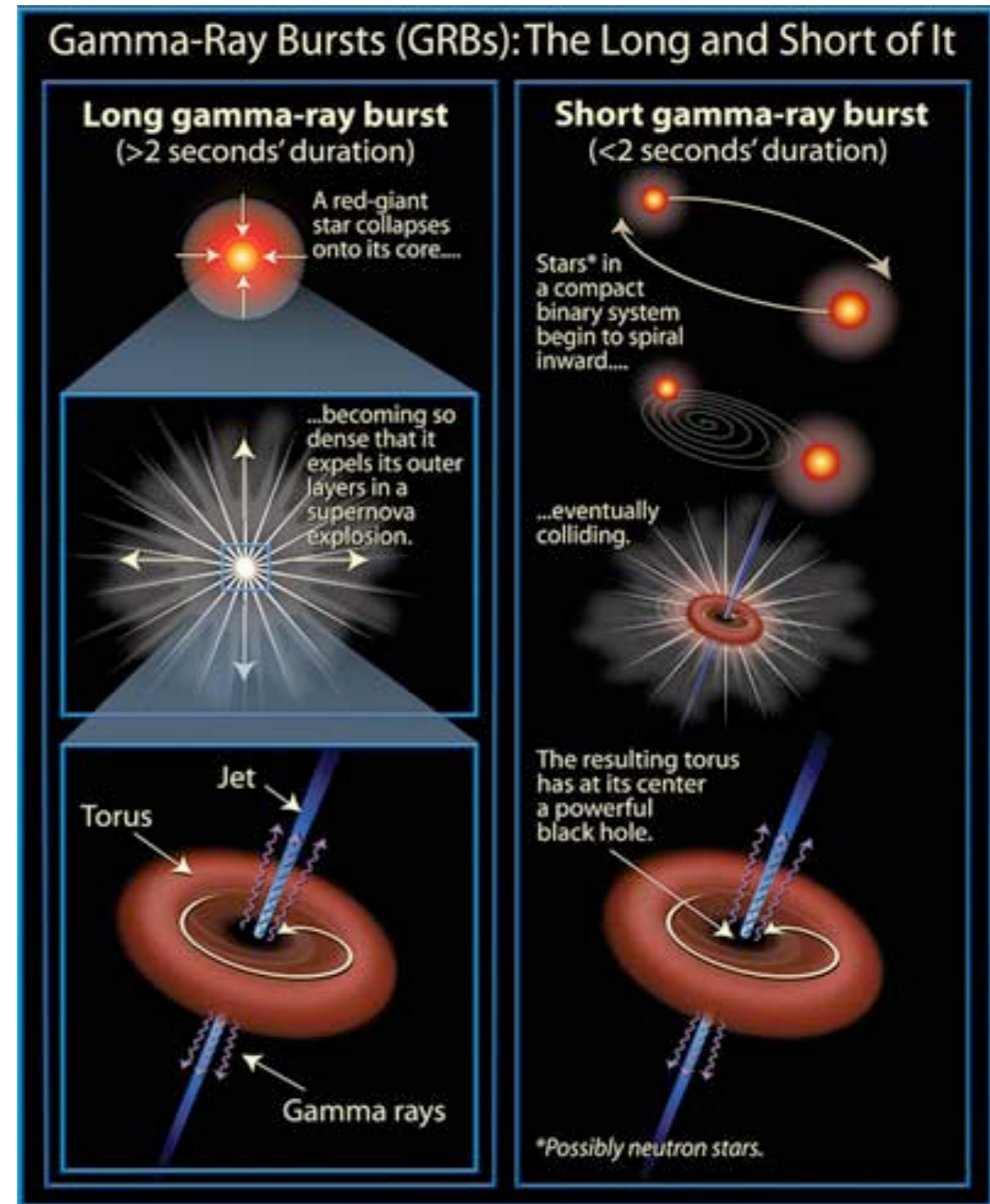
What is a gamma ray burst?

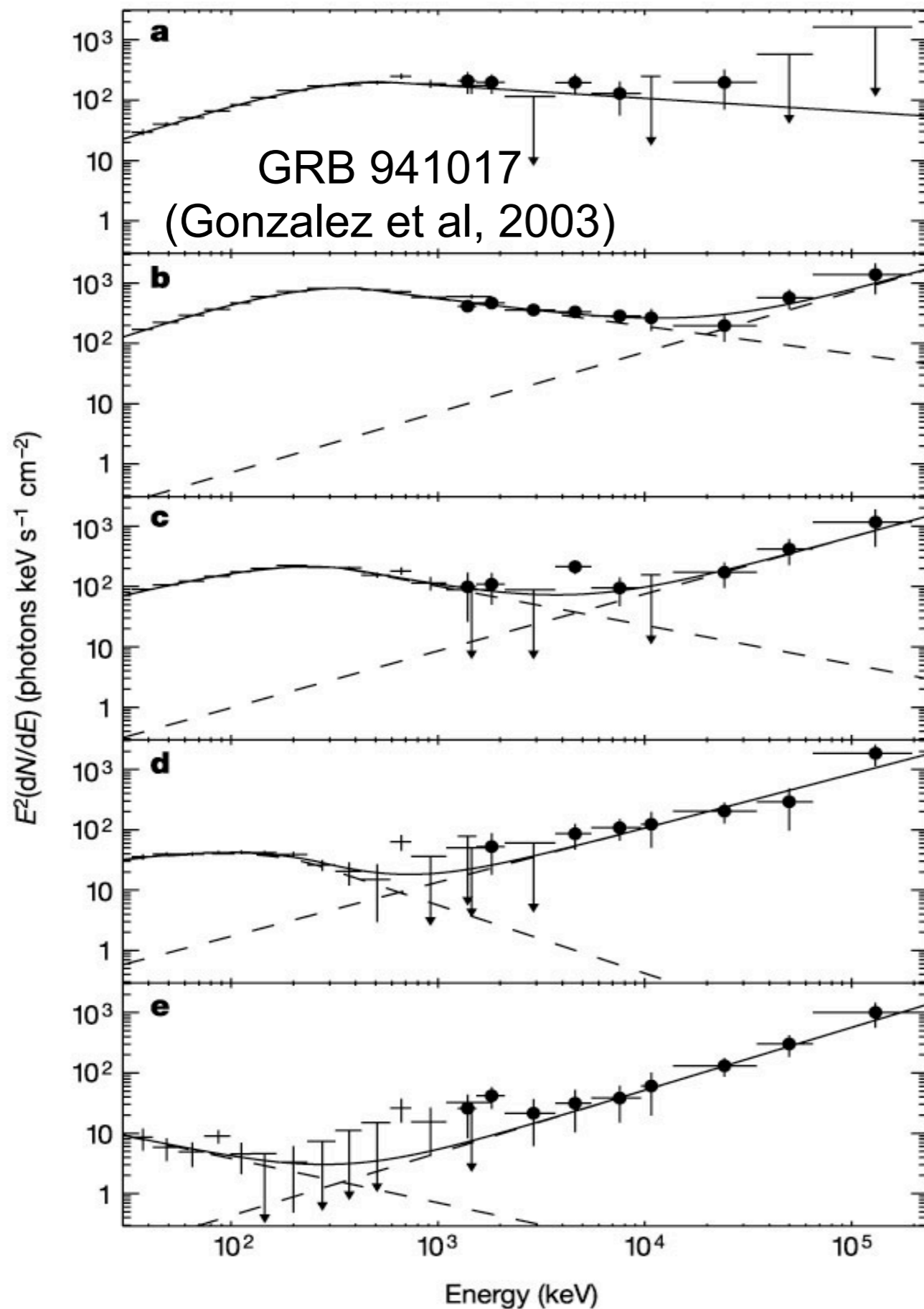
- emission (over seconds) of a large number of high energy (gamma ray) photons
- Named after date: i.e. GRB 081015a – 1st burst on 15 Oct 2008
- cosmological origin, from $z \sim 0.1$ - $z \sim 6.7$
- temporal structure varies a lot – short, long, spiky, smooth, multiple pulses, quiescent (quiet) phases
- similar spectral shapes, smooth featureless continuum



What causes a gamma ray burst?

- Long and short bursts have slightly different origins
- caused by the collapse of a massive star (long) or the merger of 2 stellar remnants (short) to a black hole and an accretion disk
- long bursts occur in star-forming regions in a galaxy
- gamma rays are emitted in beamed jets





- Previous high energy observations (EGRET on CGRO) showed extended high energy emission
- What processes are involved?
- Possibly inverse Compton scattering: synchrotron photons scattered by relativistic electrons
- or hadronic - pion decays following the photo-pion interaction of ultrarelativistic hadrons

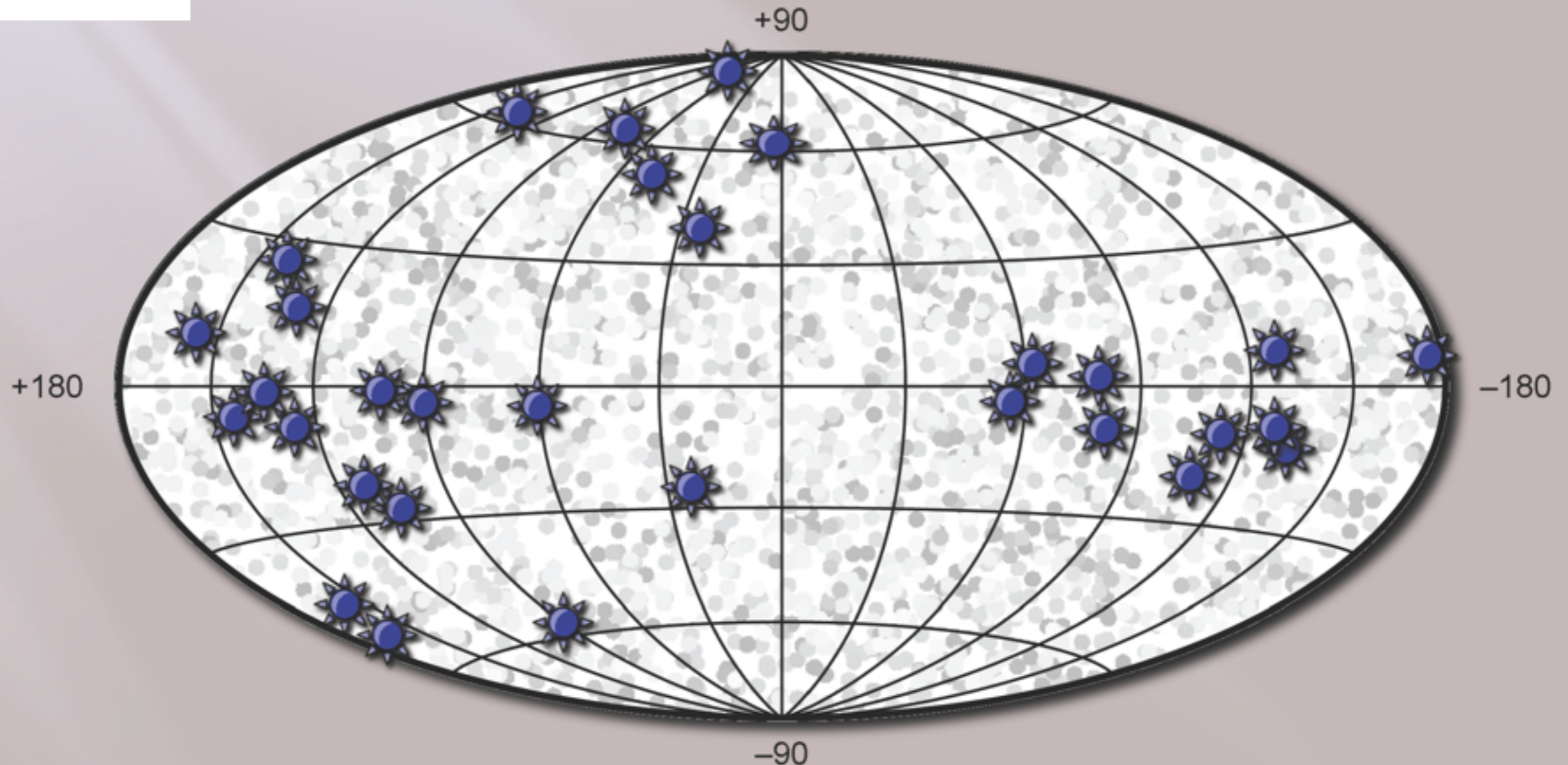
Instruments on *Fermi*

(See also: earlier talk by C. Meurer)

- Formerly known as *GLAST* (Gamma-ray Large Area Space Telescope), renamed *Fermi* after launch
- Large Area Telescope (LAT): energy range 20 MeV - 300 GeV, 4x4 tracker array with calorimeters
- Gamma Ray Burst Monitor (GBM): energy range 8 keV - 30 MeV, 12 NaI/ 2 BGO scintillator detectors
- Both instruments can be triggered by a GRB, but currently only GBM trigger is operational

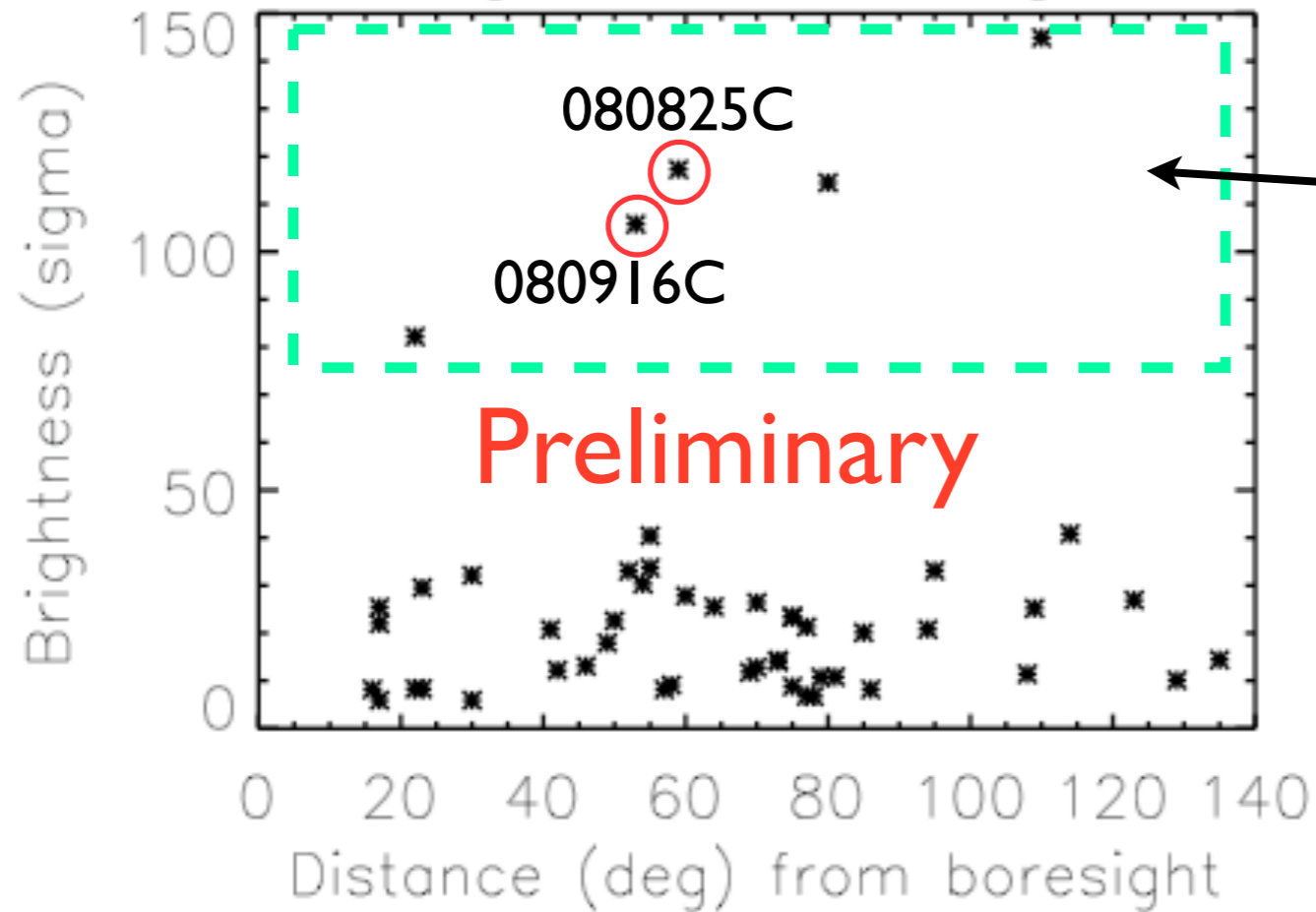


GBM Spots Gamma Ray Bursts Daily

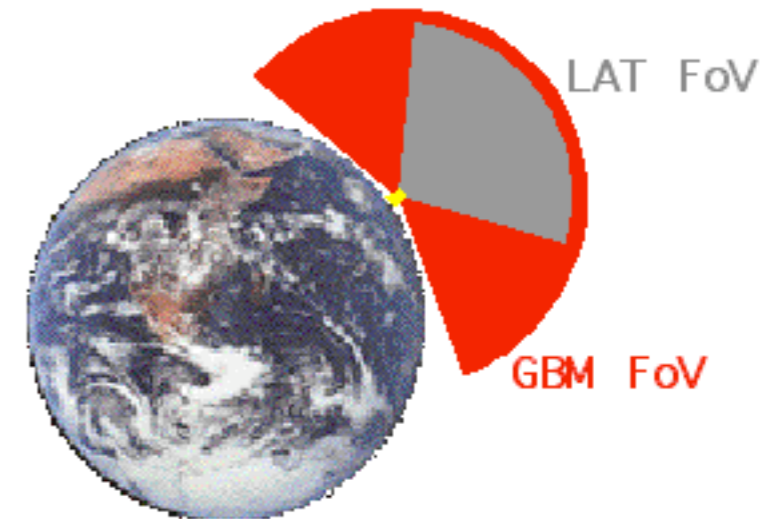


- 31 GRBs seen in first month of operations
- Activation phase complete; all working well
- Sensitivity as predicted
- GRB locations within a few degrees of Swift calculations

GBM GRB – brightness vs angle to bc



Bright bursts, expect LAT to see these



- Several GBM bursts too far off axis for LAT to see
- LAT has observed 2 bursts so far, 080825c and 080916c, out of ~70 GBM triggers in total

GRBs observed by LAT

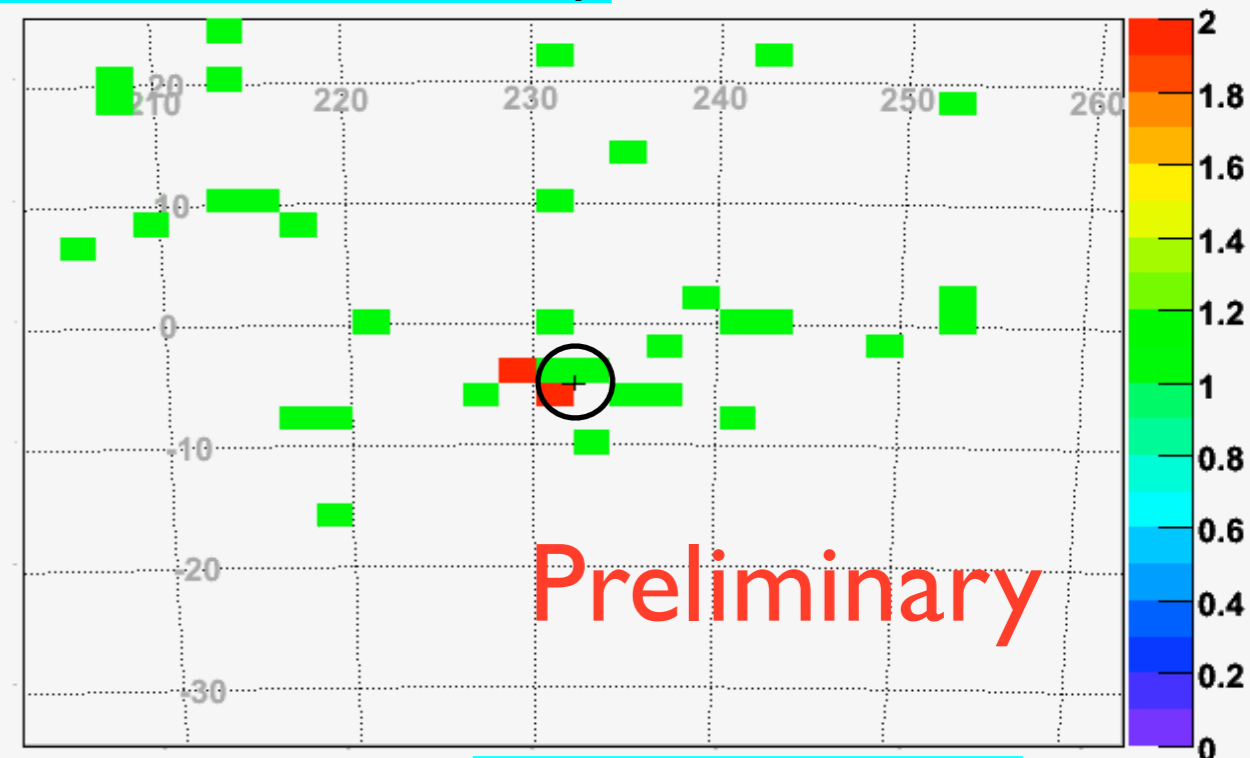
080825c:

RA = 232.2, Dec = -4.6 deg
~ 20 photons seen in LAT

080916c:

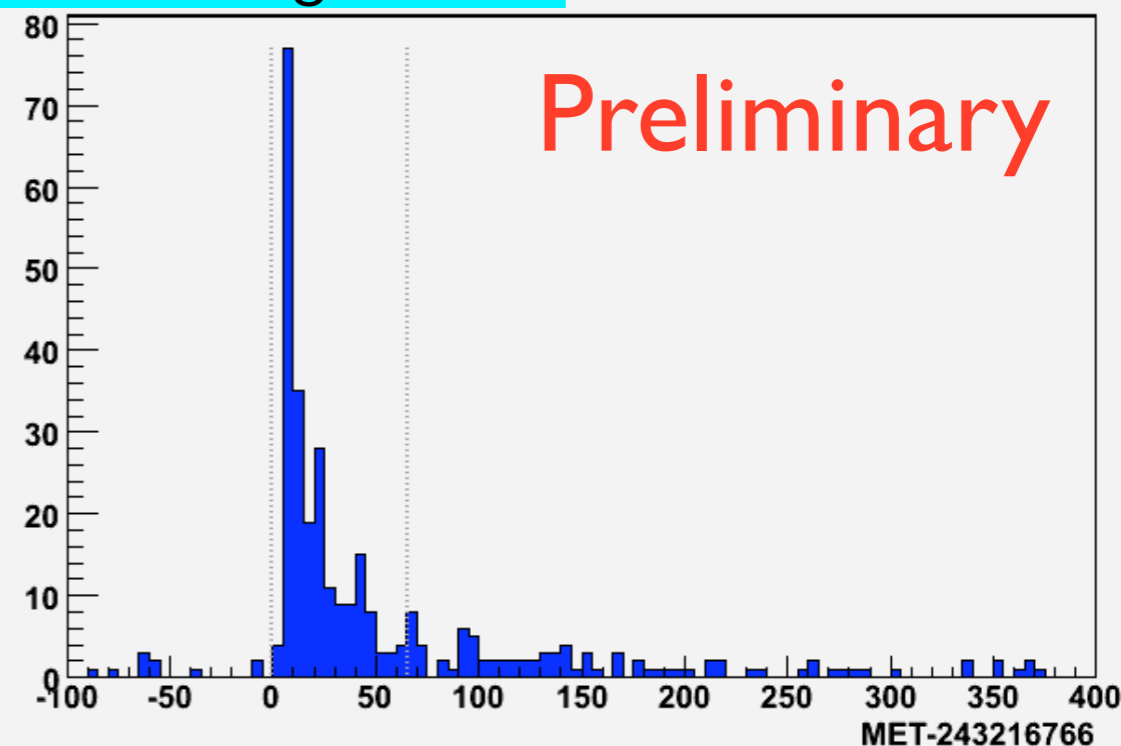
RA = 119.9, Dec = -56.6 deg
Enough counts to perform
time-resolved spectral analysis

080825c counts map



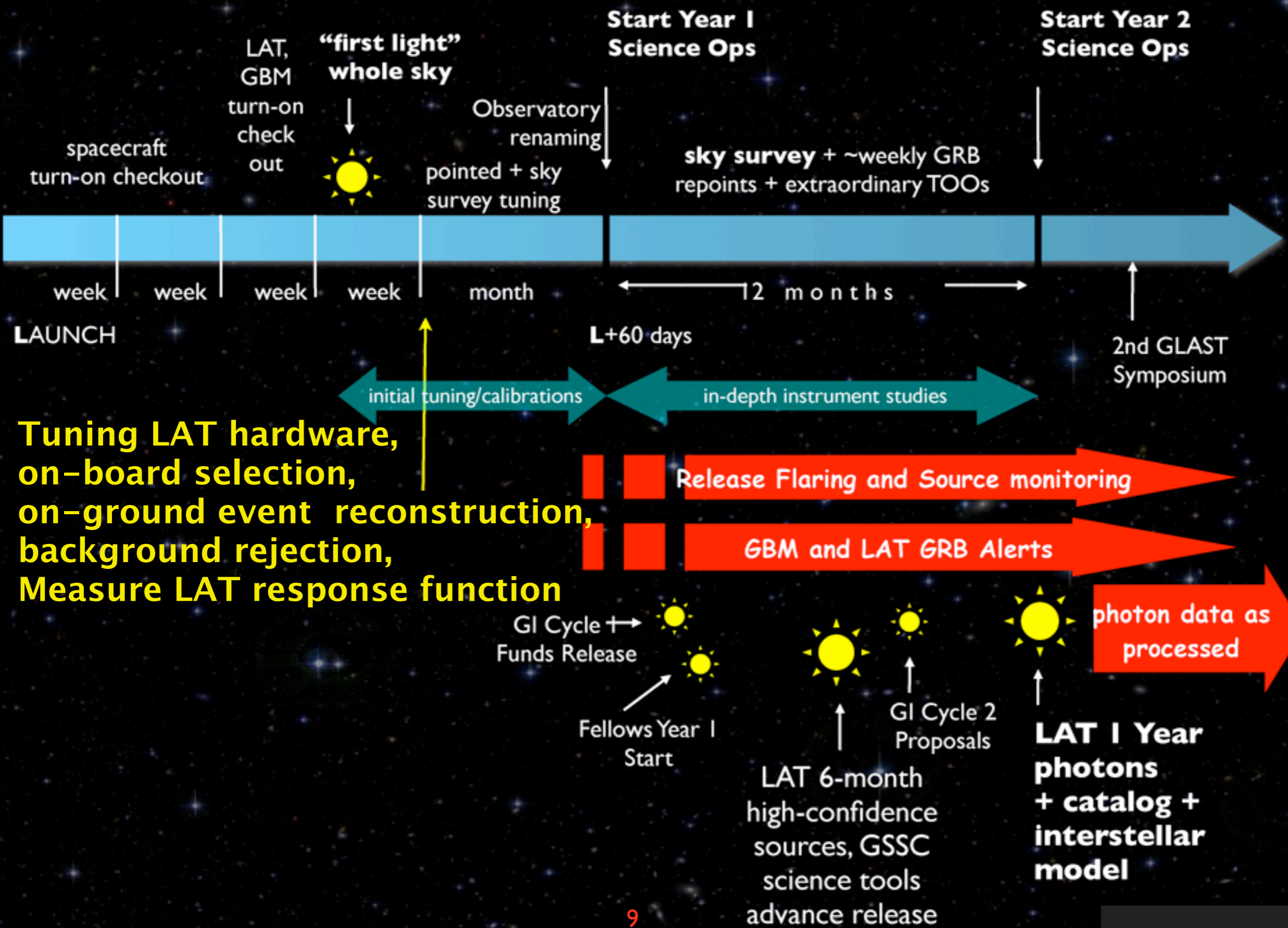
GBM location in circle

080916c lightcurve



Timeline

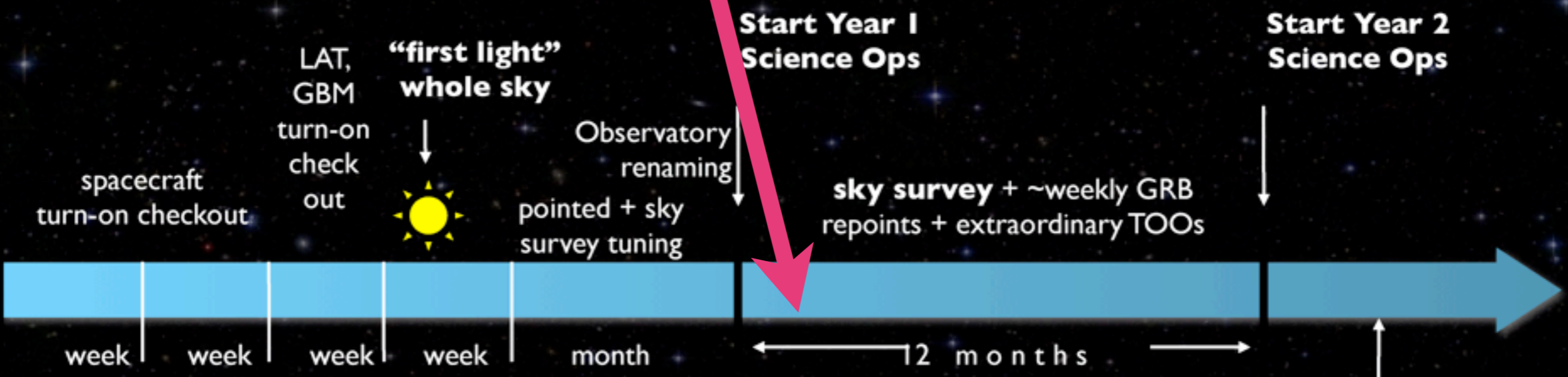
We are ~here



Tuning LAT hardware, on-board selection, on-ground event reconstruction, background rejection, Measure LAT response function

Timeline

We are ~here



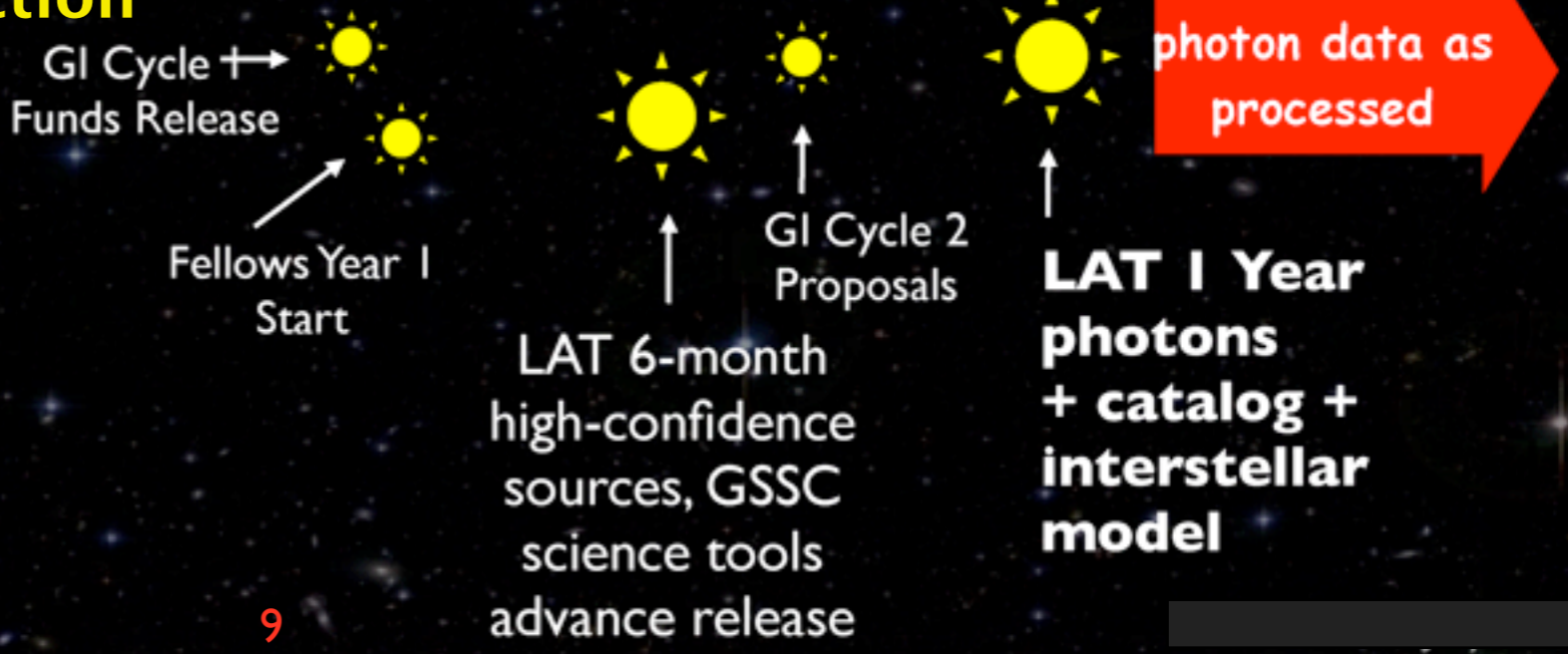
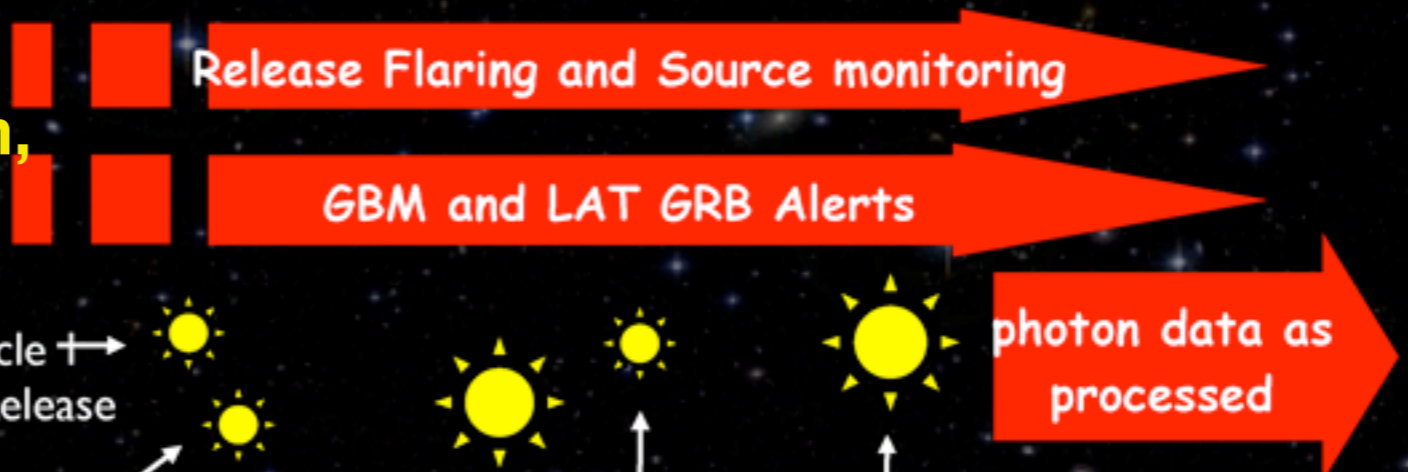
LAUNCH

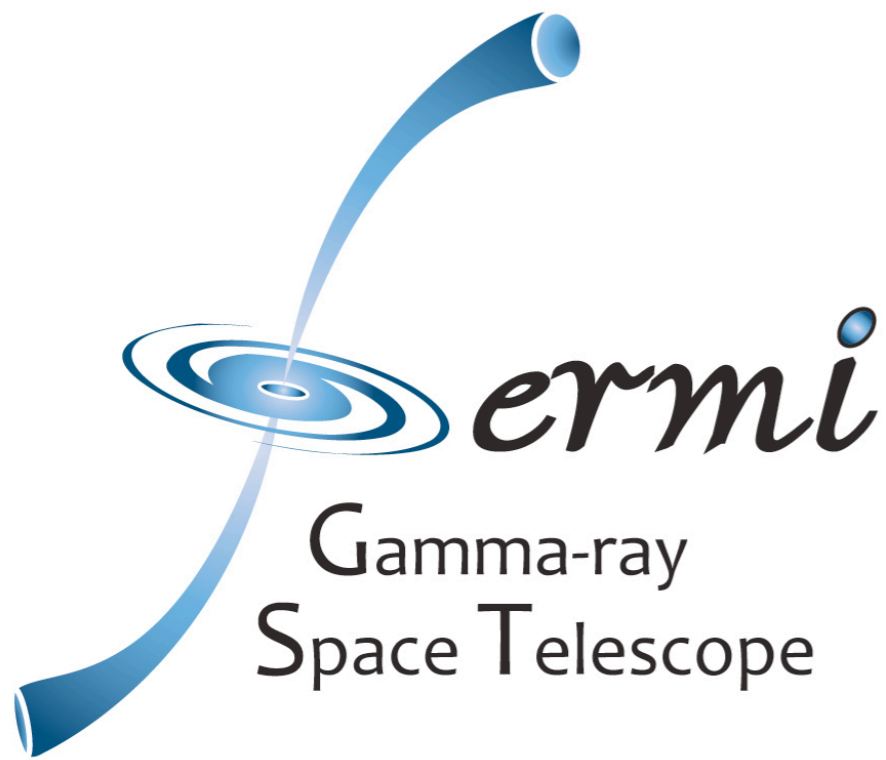
L+60 days

2nd GLAST Symposium



Tuning LAT hardware, on-board selection, on-ground event reconstruction, background rejection, Measure LAT response function

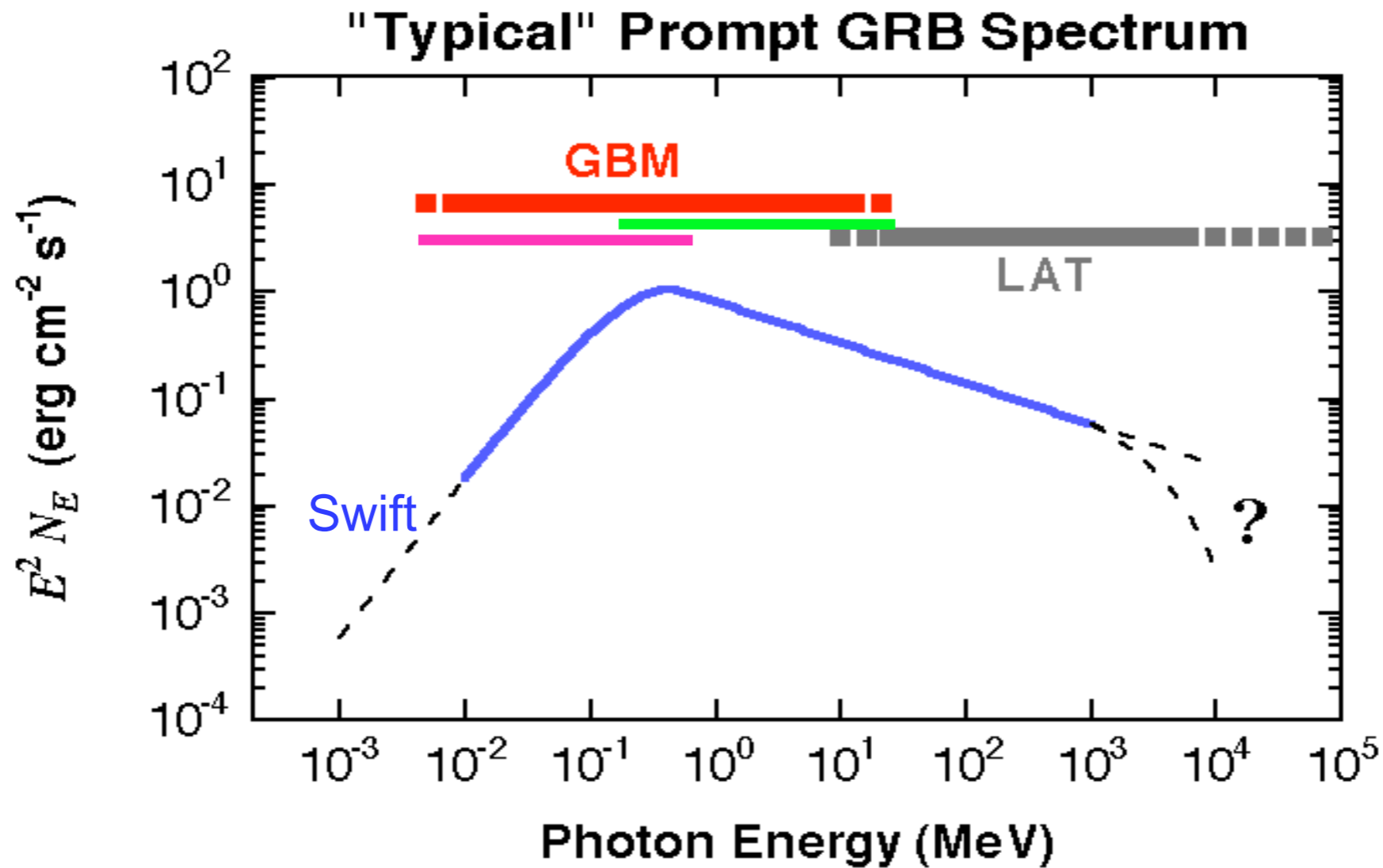




Conclusions:

- Instruments working extremely well so far
- 2 GRBs observed with LAT to date, more expected!
- More systematic testing needed, event selections, optimisation of background selection etc
- Many exciting results to come...





- *Fermi* is able to detect emission from the prompt and afterglow phase of GRBs
- Can follow up high energy emission of GRBs from other instruments, joint spectrum over 7 decades of energy
- High resolution at very high energies - helps to constrain spectral models & emission mechanisms