

# Why building a muon collider

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DI PADOVA



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If you want to know more, look here:

**Muon Collider Physics Summary** [[2203.07256](#)]

**The Muon Smasher's Guide** [[2103.14043](#)]

**The Physics Case of a 3 TeV Muon Collider Stage** [[2203.07261](#)]

# (Emerging) Conclusions

## Muon Colliders BSM Physics Pillars:

- **High Energy** available for direct particles production
- **High Rate** available for **Precision** measurements
- **Energy and Precision**

All this, at a single collider with feasible timescale

## WIMP Dark Matter

- **Higgsino/Wino “very directly” accessible** (more is coming on WIMP at 10 TeV)

## Explaining the origin of the Weak scale

- $\Delta=10 \rightarrow \Delta=80$  , simply from direct searches
- Probe  $\Delta=1/\xi=1000$ , in CH, in two ways

## How much is the Higgs radius?

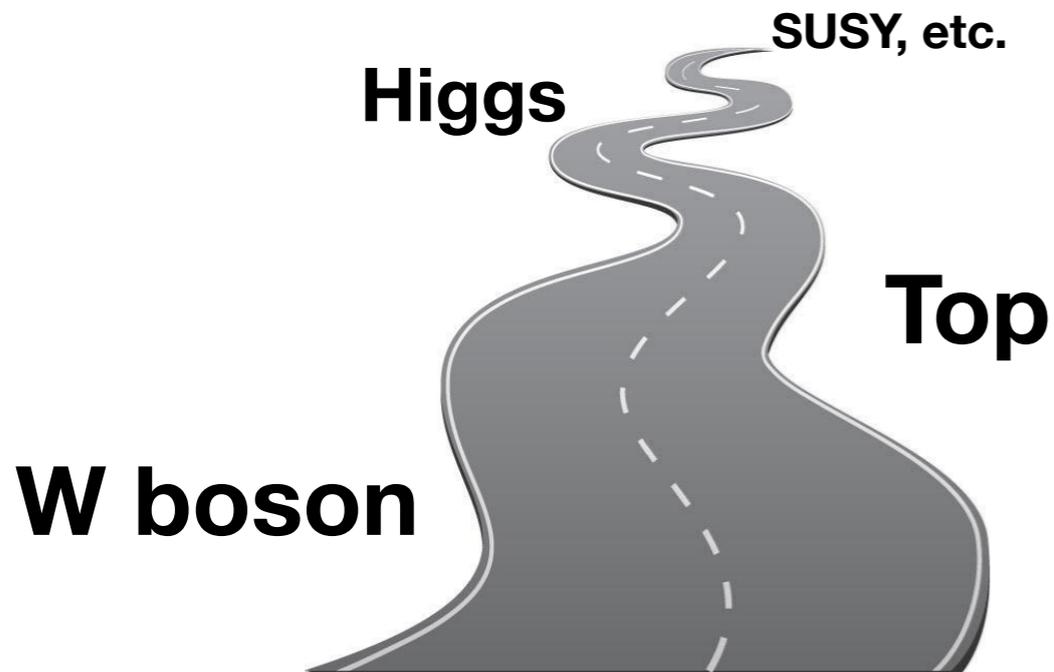
- If as “large” as  $1/(50\text{TeV})$ , we can tell

## Added value from colliding muons [\[2203.07261\]](#)

- Current anomalies are in muons. If coincidence, a fortunate one: illustrates obvious potential of colliding  $\mu$  for the first time!

# The High Energy Physics Landscape

## HEP Yesterday



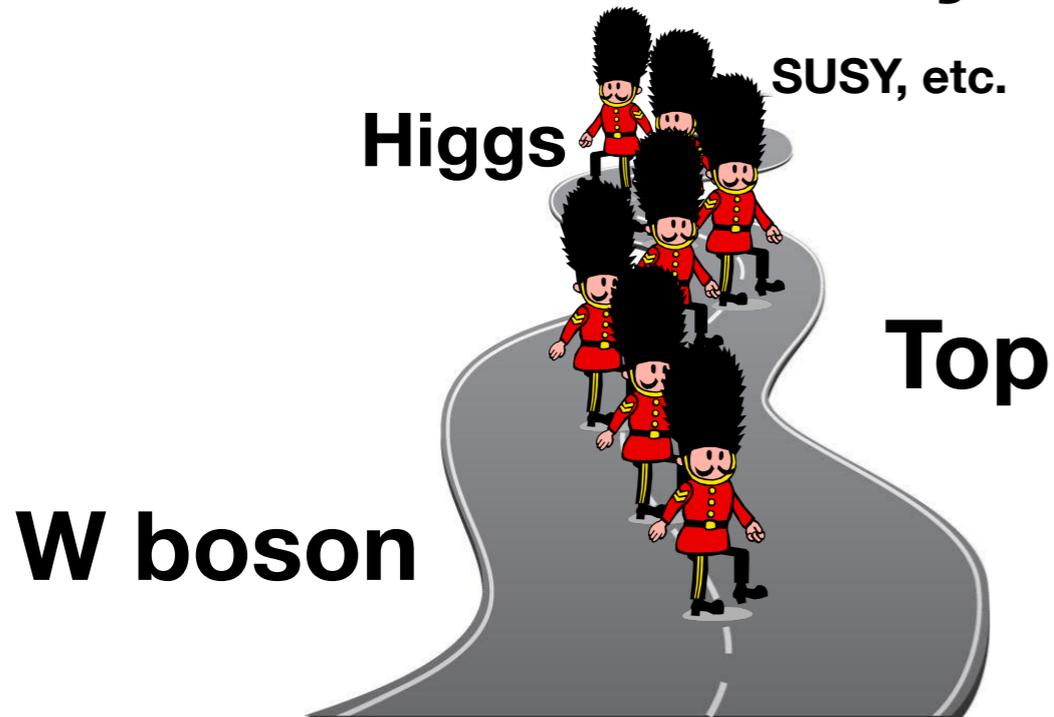
LHC

## HEP Today



# The High Energy Physics Landscape

**HEP Yesterday**



**LHC**

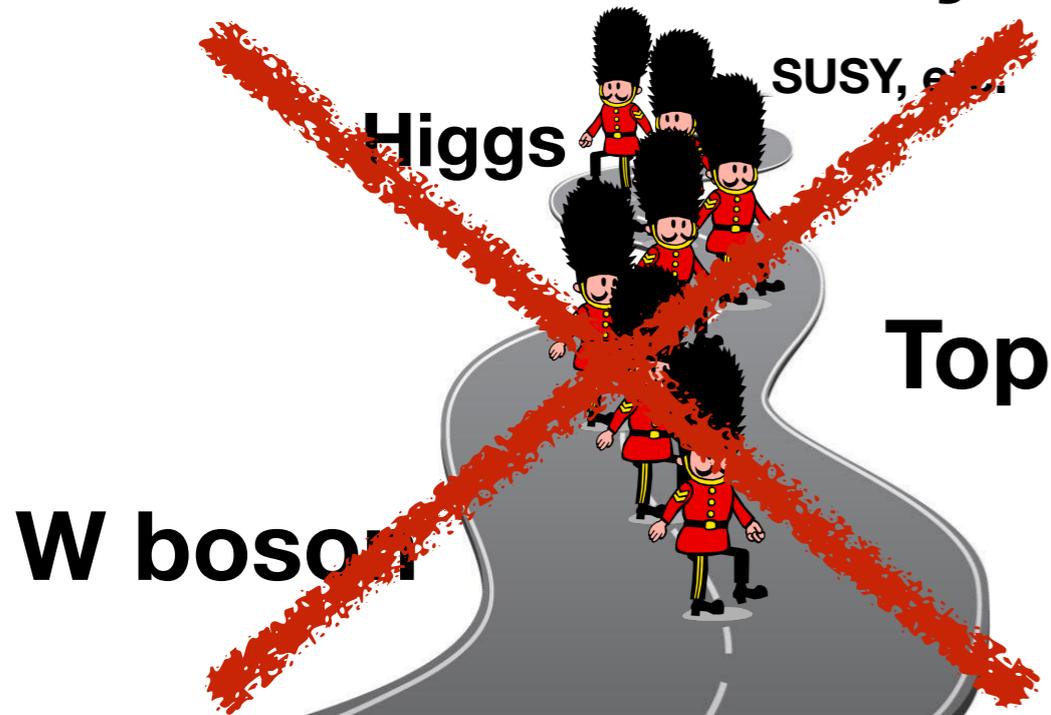
**HEP Today**



Yesterday, HE-Physicists were used to **follow a road.**

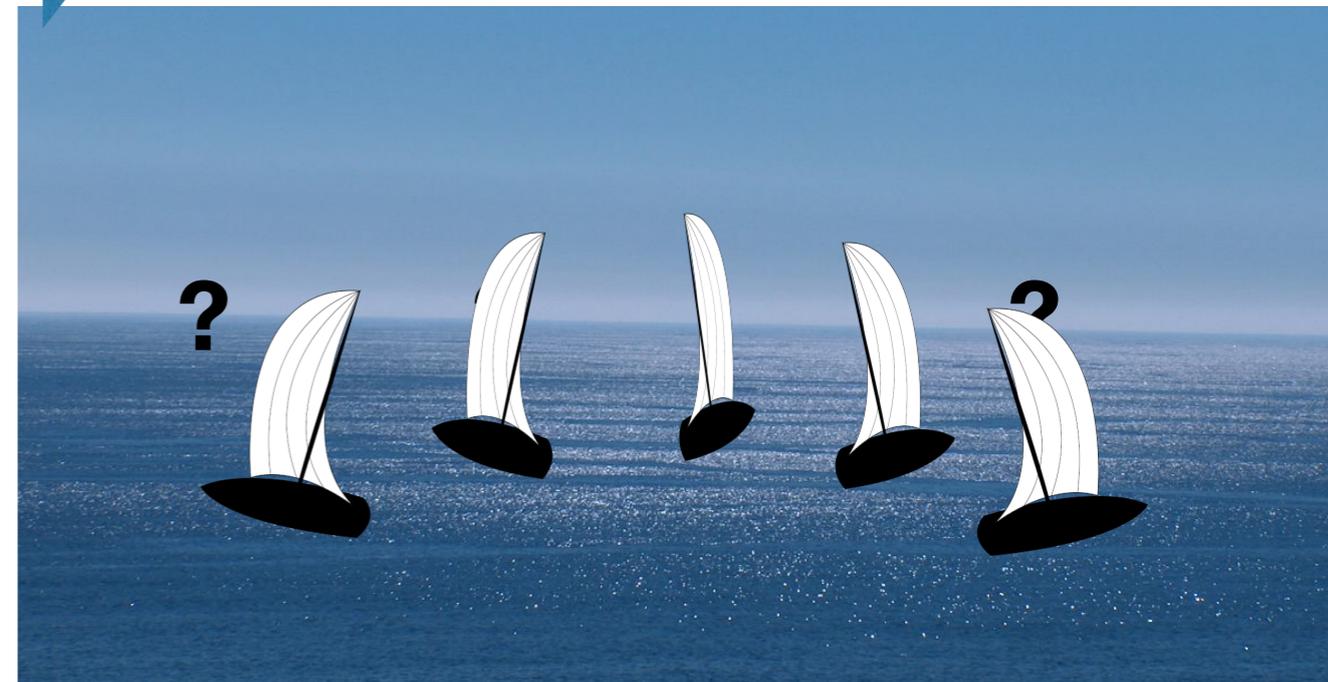
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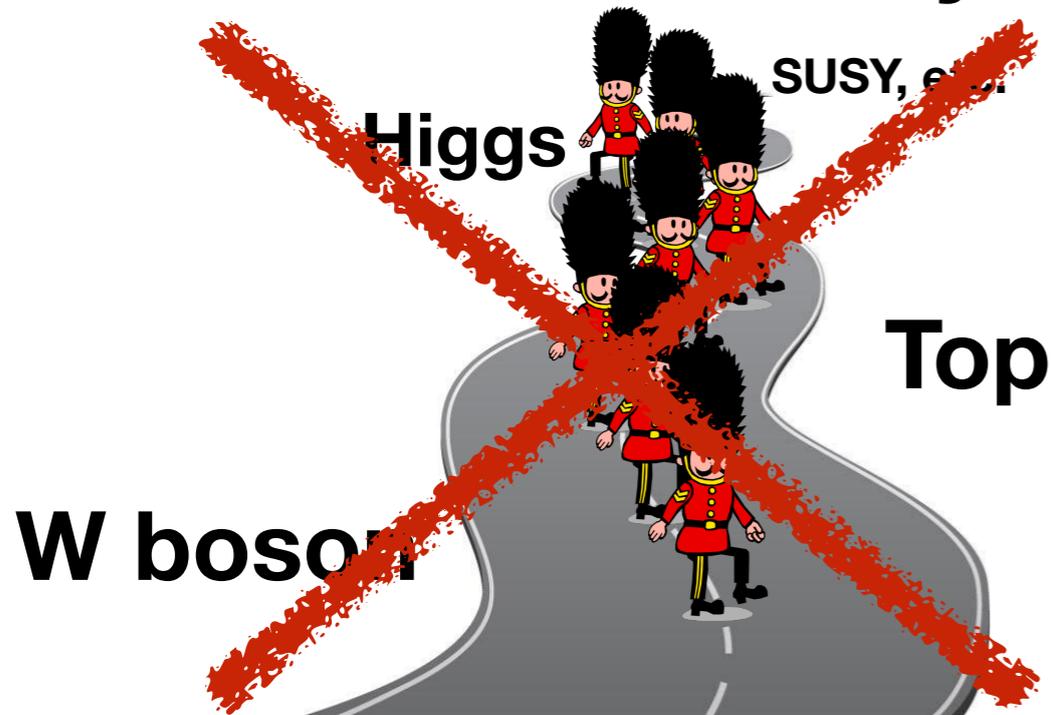
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Today, the mission is to **explore uncharted territory\***

\*Which is **good!**

It means that the next discovery will be more revolutionary than the Higgs one

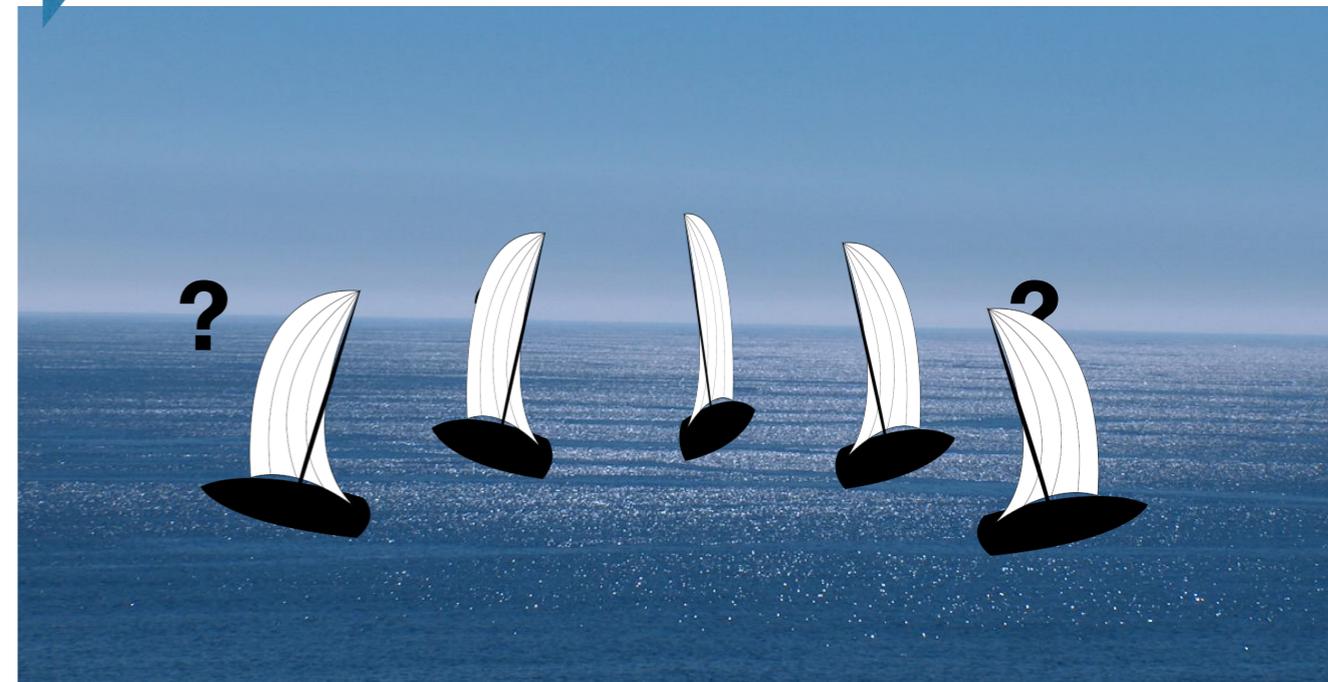
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This is why we started speaking about **Frontiers**

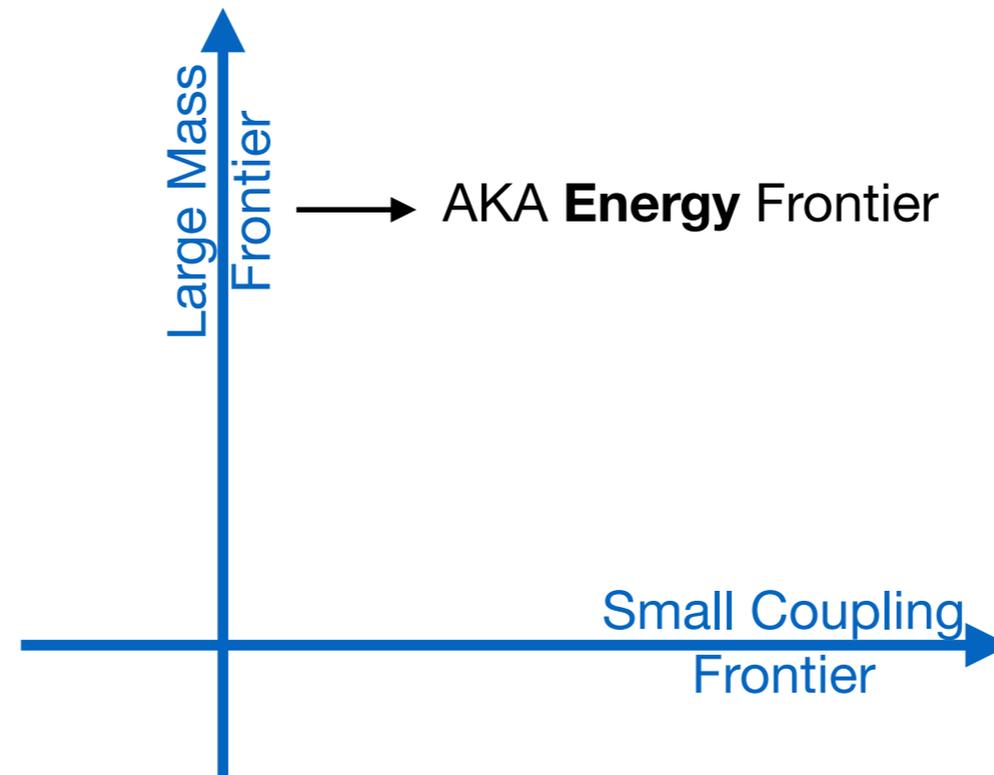
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# The High Energy Physics Landscape

Our **Frontiers** are the **directions** in which (i.e., reasons why) New Physics might hide

W boson



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**No single experiment** can explore all directions at once.

**None** (in heaven or earth) can **guarantee discoveries** of new fundamental laws of nature.

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Accurate measurements of great variety of observables.

Under precisely known experimental conditions.

+

Accurate predictions within the Standard Model of Particle Physics.

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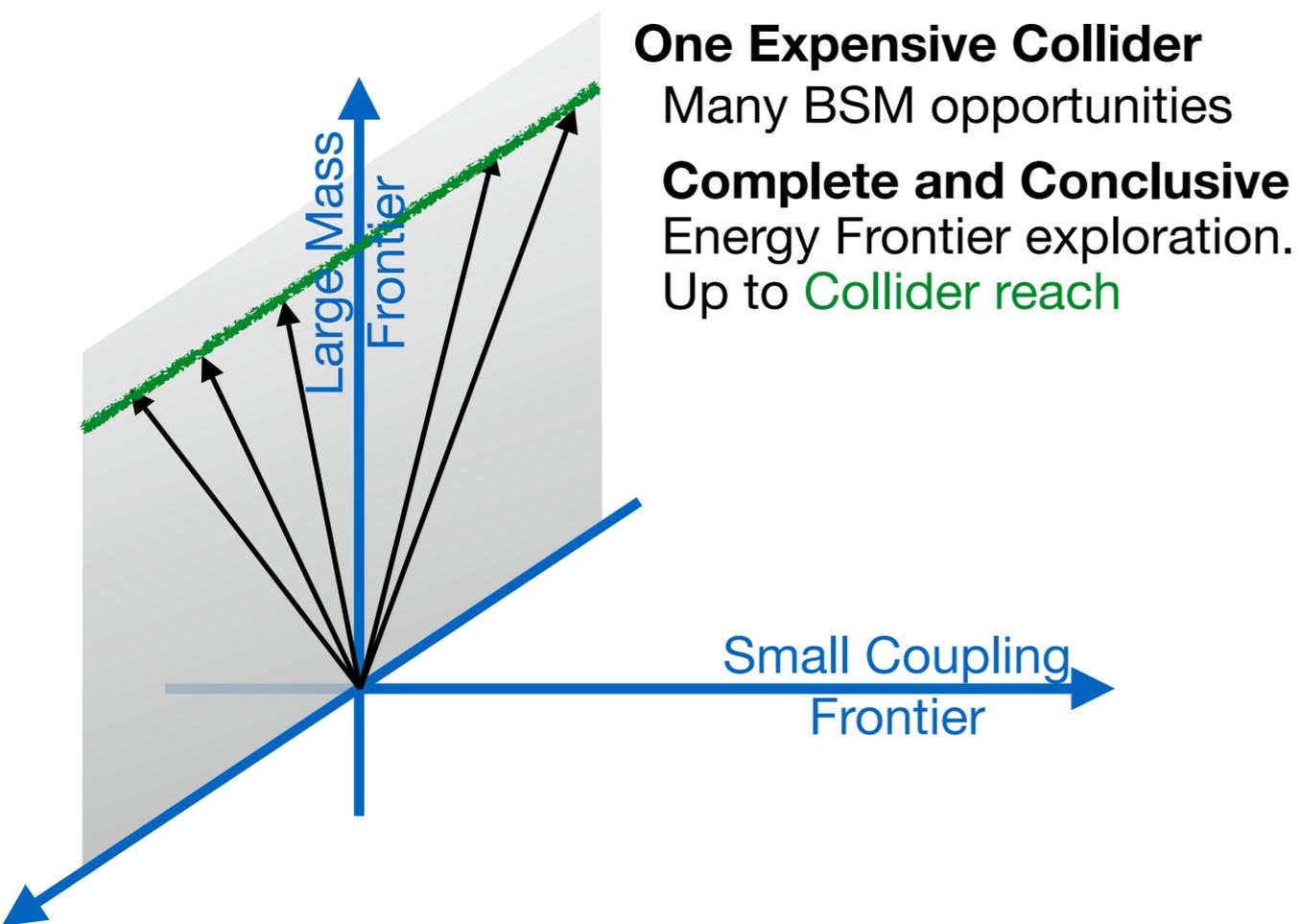
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— 2) Sharp answers to well-posed **Beyond the SM questions**

Only one drawback: they are **Expensive**.

# Why Colliders?

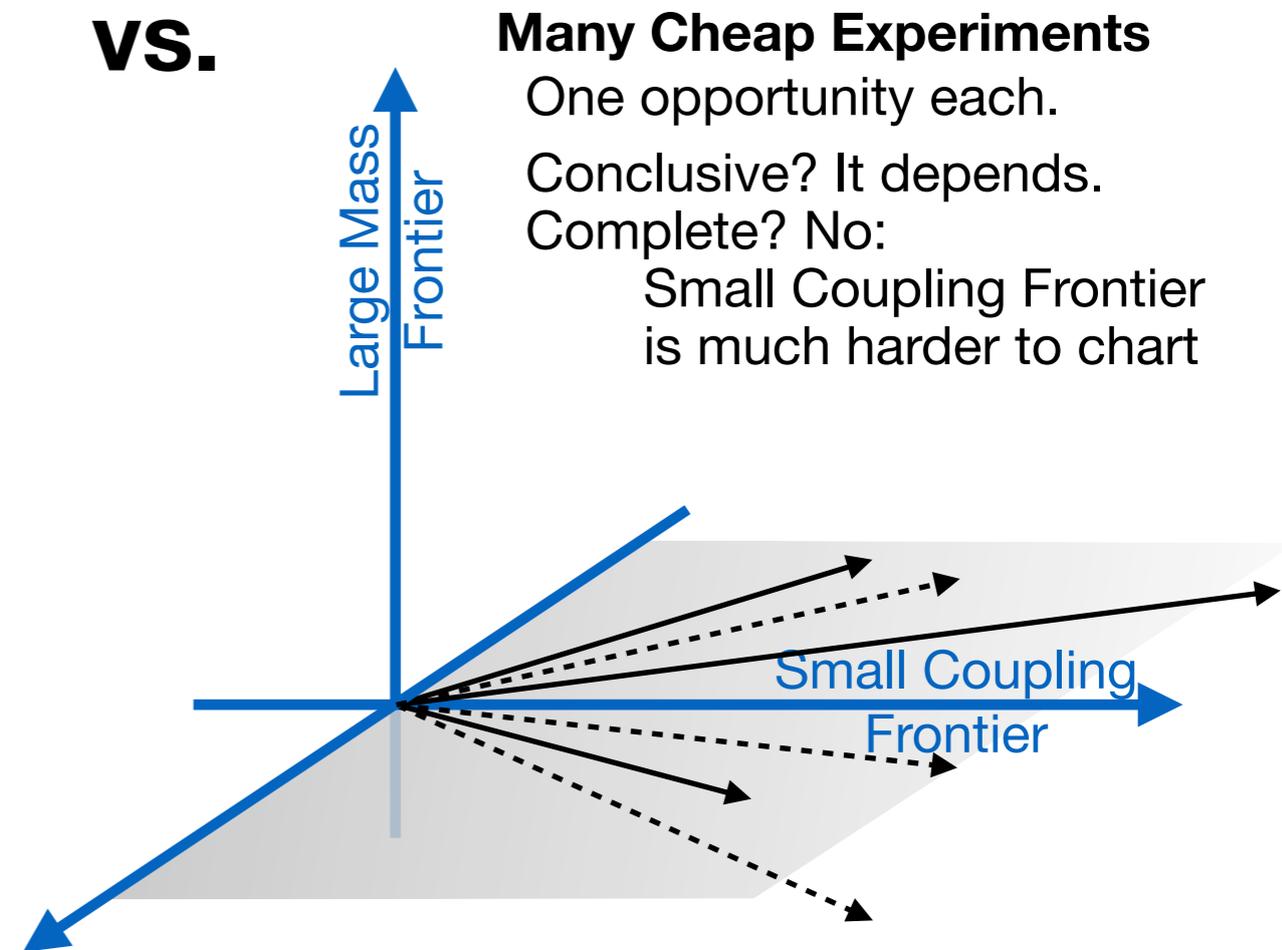
**Expensive?** Yes, no doubt, but ...



**One Expensive Collider**  
Many BSM opportunities  
**Complete and Conclusive**  
Energy Frontier exploration.  
Up to **Collider reach**

One of the many additional axes that characterise New Physics complexity

**VS.**

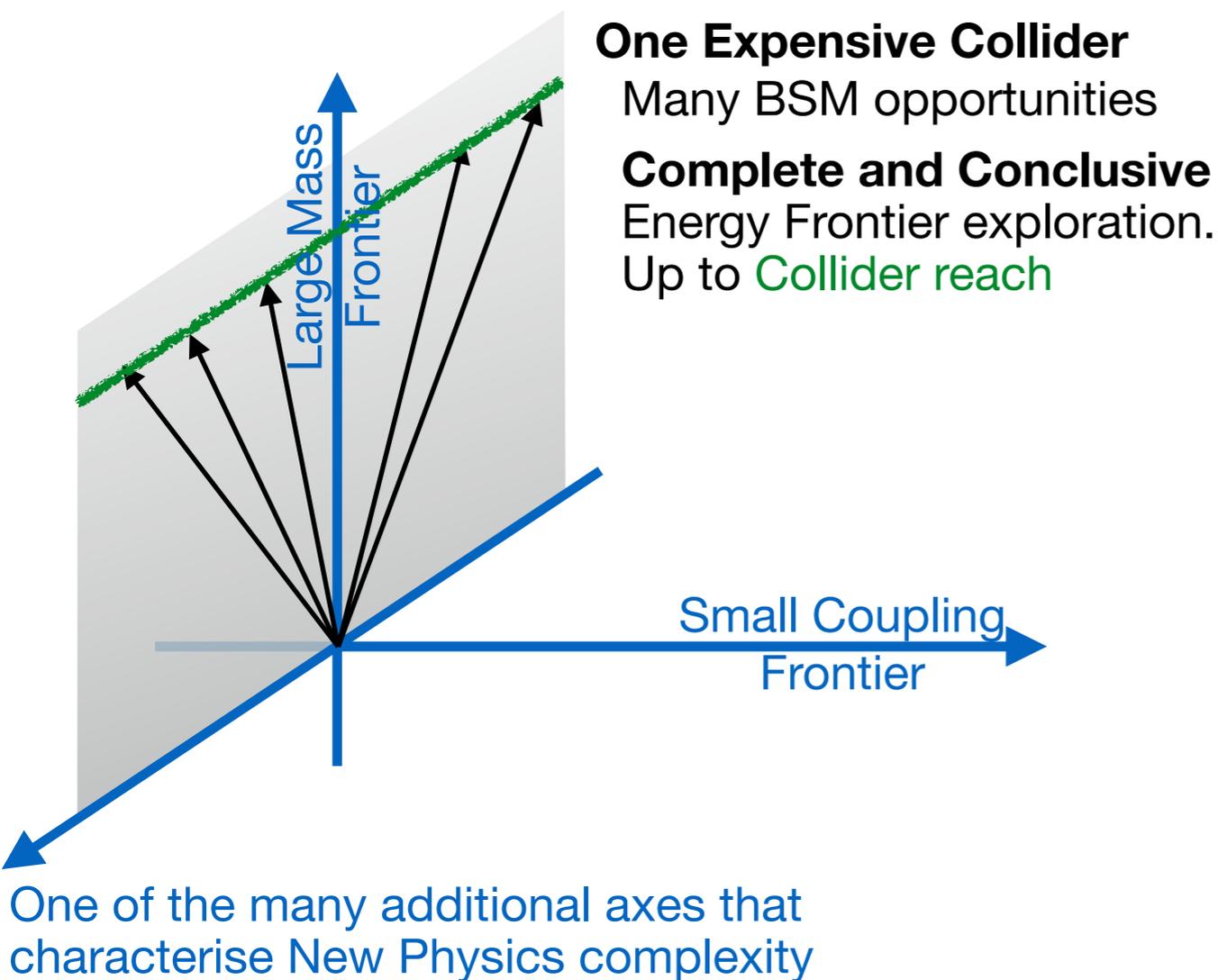


**Many Cheap Experiments**  
One opportunity each.  
Conclusive? It depends.  
Complete? No:  
Small Coupling Frontier  
is much harder to chart

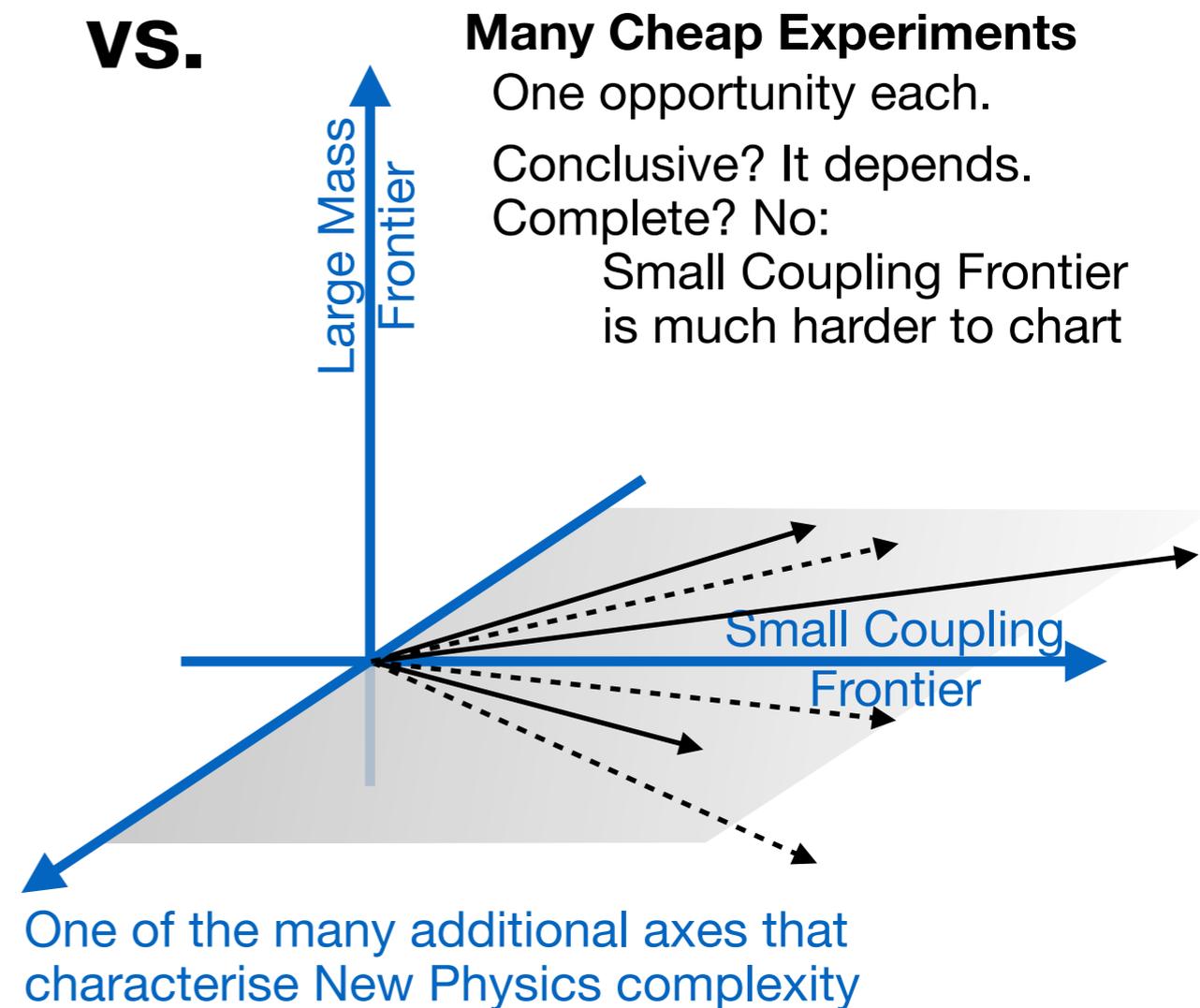
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# Why Colliders?

**Expensive?** Yes, no doubt, but ...



**VS.**



Still, no doubt that next big project, to have a chance, must be ambitious enough to make **great jump ahead** in exploration of **multiple directions** [even better if constructed with **revolutionary technology**]

# Why Muons?

**Leptons** are the ideal probes of short-distance physics:

All the energy is stored in the colliding partons

No energy “waste” due to parton distribution functions

High-energy physics probed with much smaller collider energy

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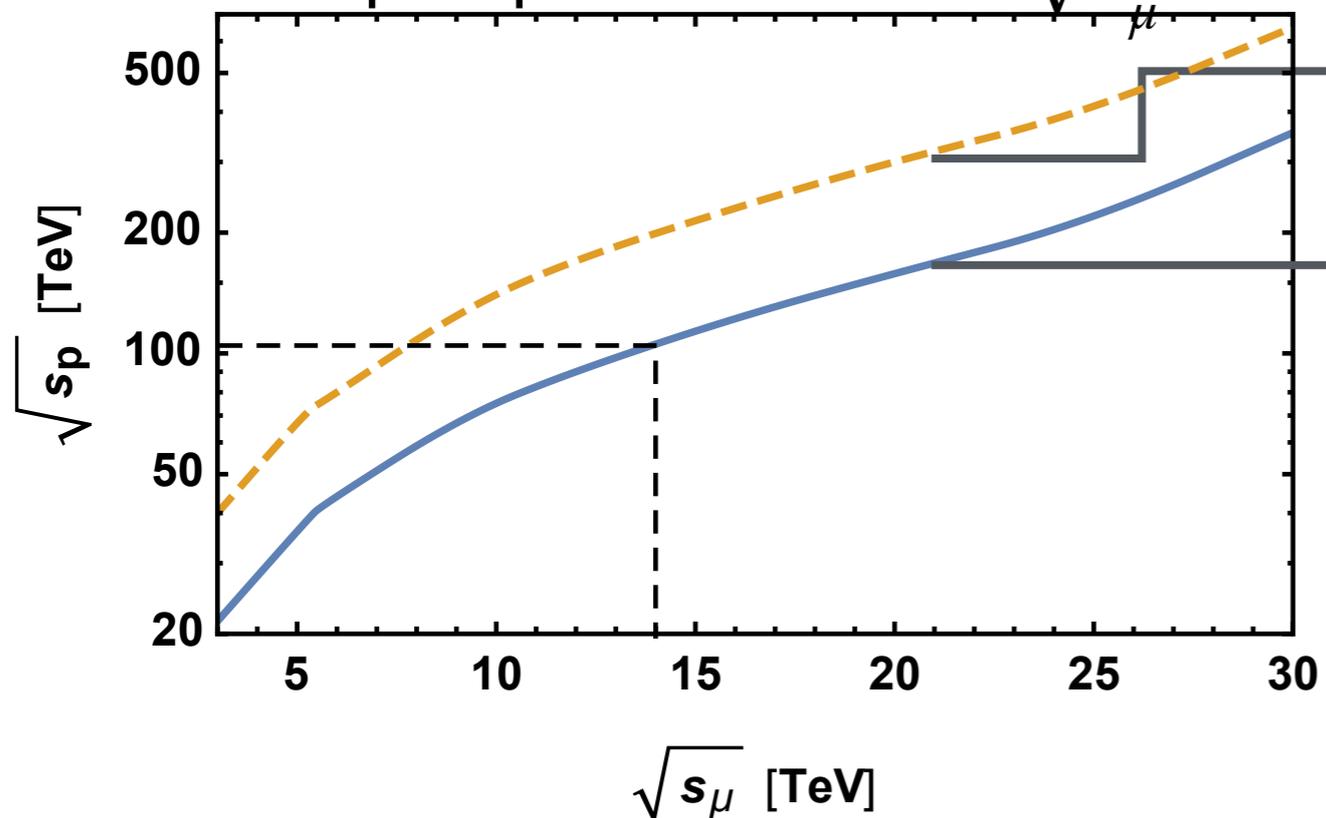
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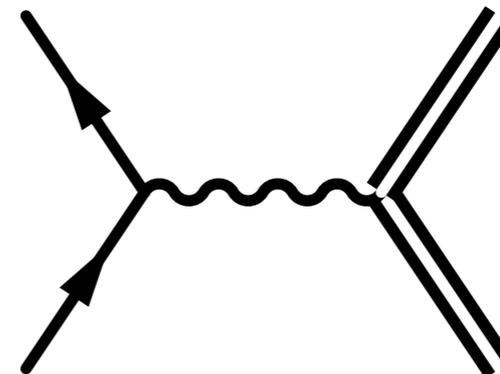
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pp  $\sqrt{s}$  at which  $\sigma_{pp} = \sigma_{\mu\mu}$   
for pair prod. with  $M \sim \sqrt{s}$



Estimate for EWK-only  
charged particles

Estimate for EWK+QCD-  
charged particles



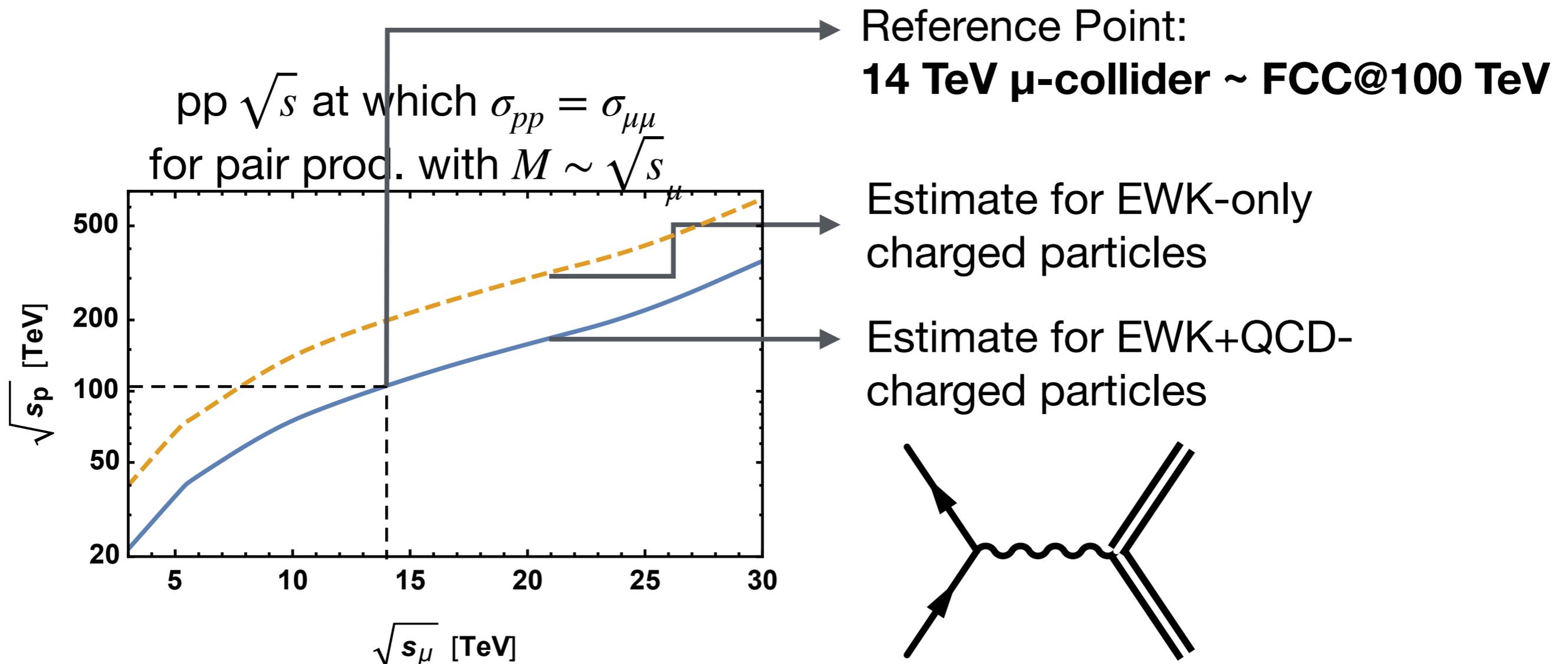
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[cannot accelerate them in rings above few 100 GeV]

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**Muon Colliders**

Input to the European Particle Physics Strategy Update

**The Muon Collider Working Group**

Jean Pierre Delahaye<sup>1</sup>, Marcella Diemoz<sup>2</sup>, Ken Long<sup>3</sup>, Bruno Mansoulié<sup>4</sup>, Nadia Pastrone<sup>5</sup> (chair),  
Lenny Rivkin<sup>6</sup>, Daniel Schulte<sup>1</sup>, Alexander Skrinsky<sup>7</sup>, Andrea Wulzer<sup>1,8</sup>



**Deliberation Document**

**on the 2020 update of the European Strategy for Particle Physics**

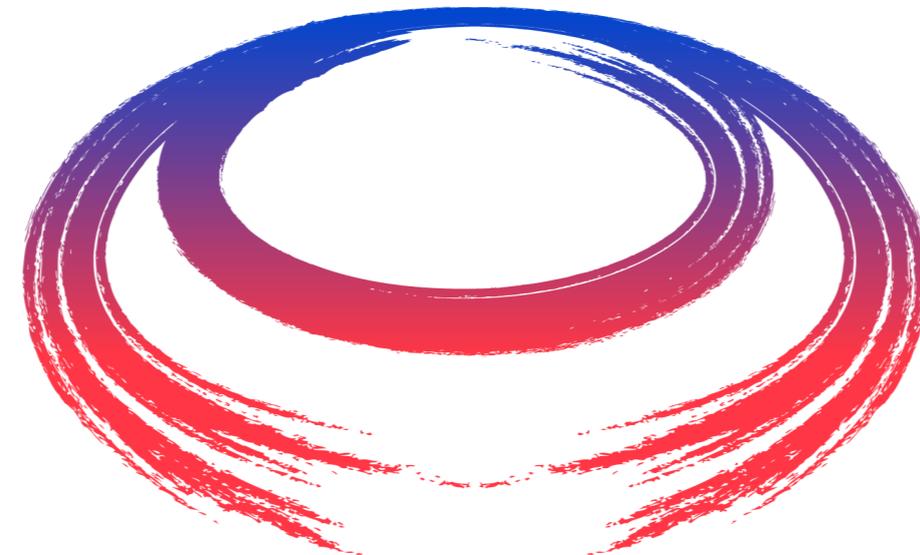
- an international design study for a muon collider, as it represents a unique opportunity to achieve a multi-TeV energy domain beyond the reach of  $e^+e^-$  colliders, and potentially within a more compact circular tunnel than for a hadron collider. The biggest challenge remains to produce an intense beam of cooled muons, but novel ideas are being explored;

# Why Muons?

**Leptons** ... physics:

All the e  
No ene  
High-er

**Electro**



**M** International  
UON Collider  
Collaboration

[muoncollider.web.cern.ch](http://muoncollider.web.cern.ch)

GeV]

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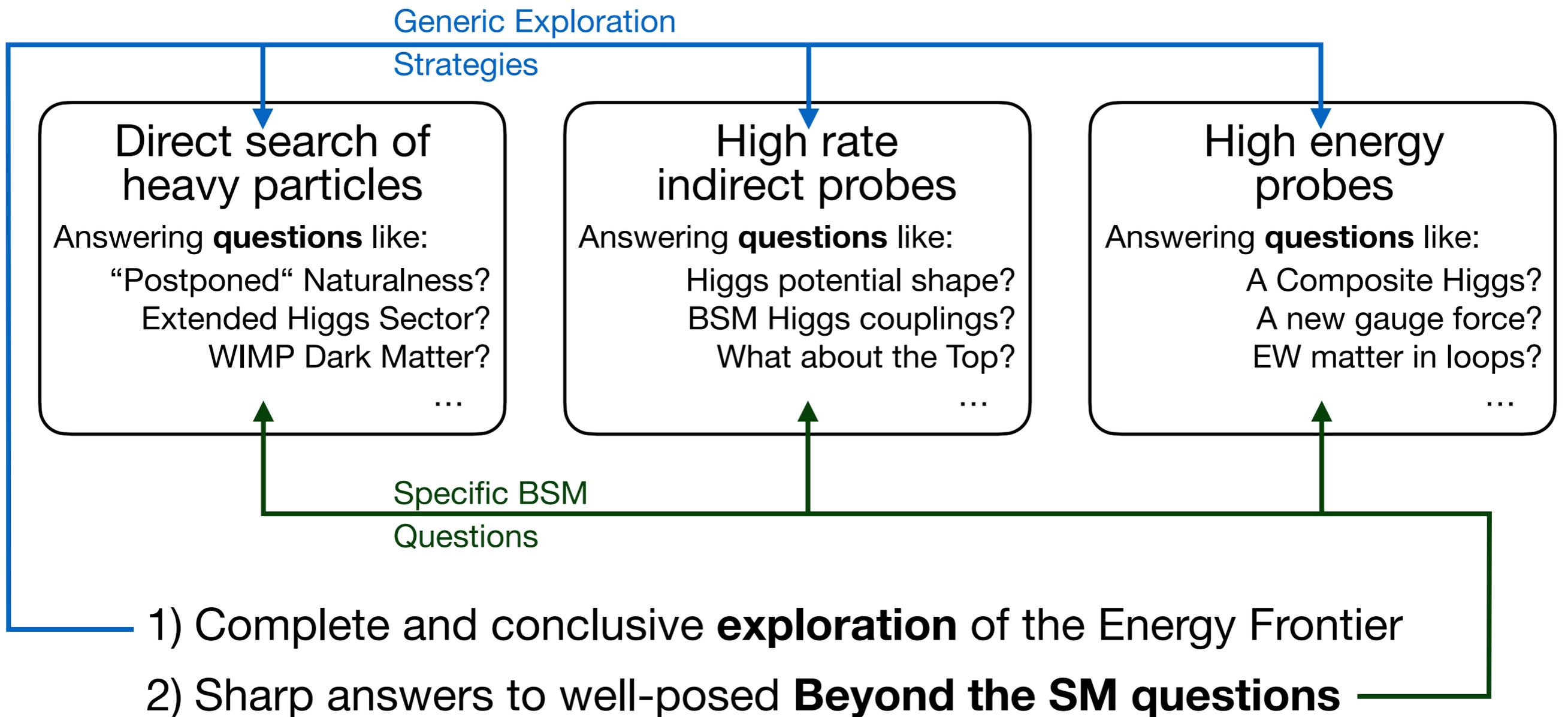
IMCC Design Targets:

- **~3 TeV MuC with  $\sim 1/\text{ab}$** , in 5 yrs for 1 IP
- **10 TeV MuC with  $10/\text{ab}$** , in 5 yrs for 1 IP
- **E<sub>max</sub> MuC = ?**, to be assessed

Most focus on 10 TeV and  $10/\text{ab}$  in what follows

# Why Muons?

## Muon Collider Physics Potential Pillars

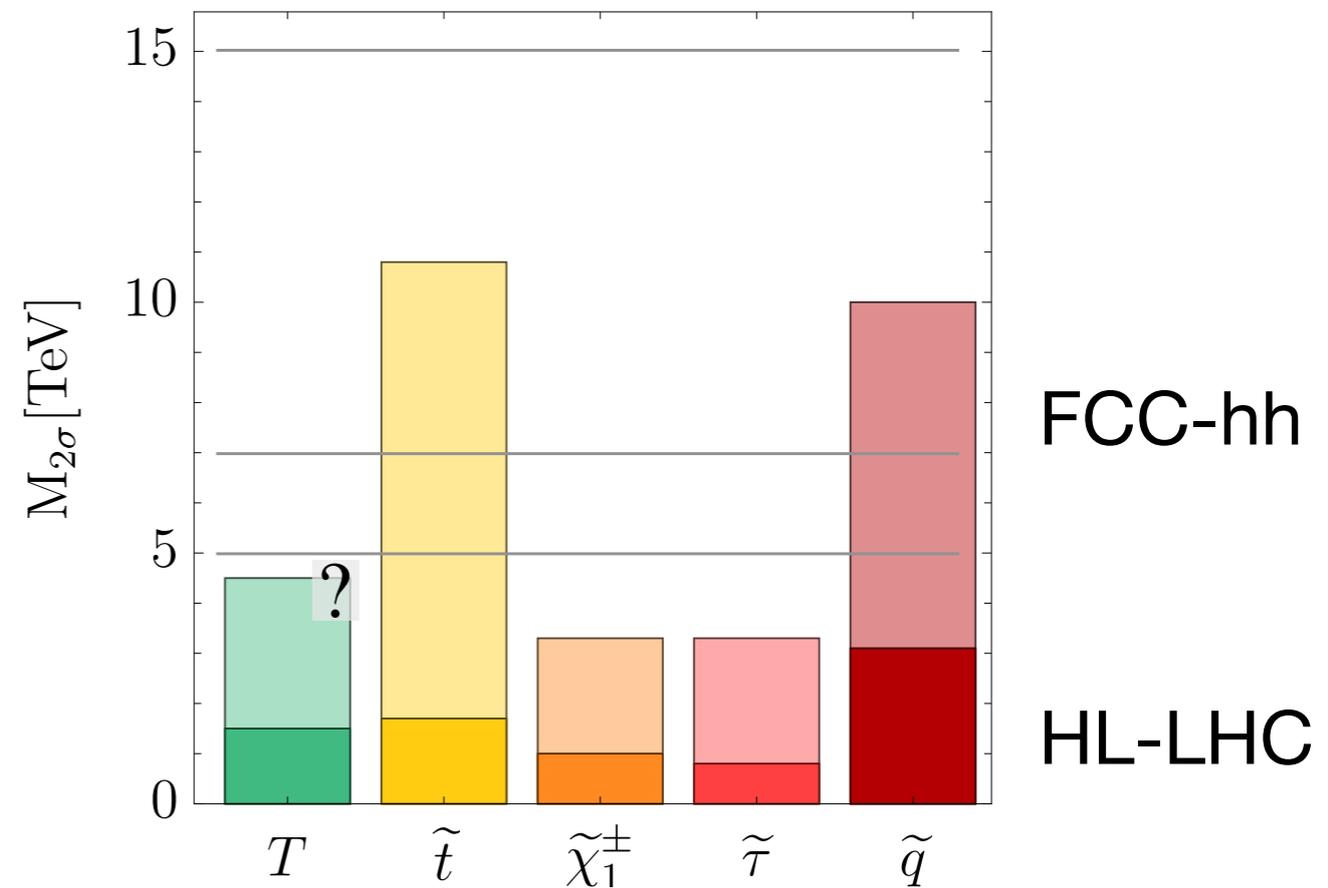
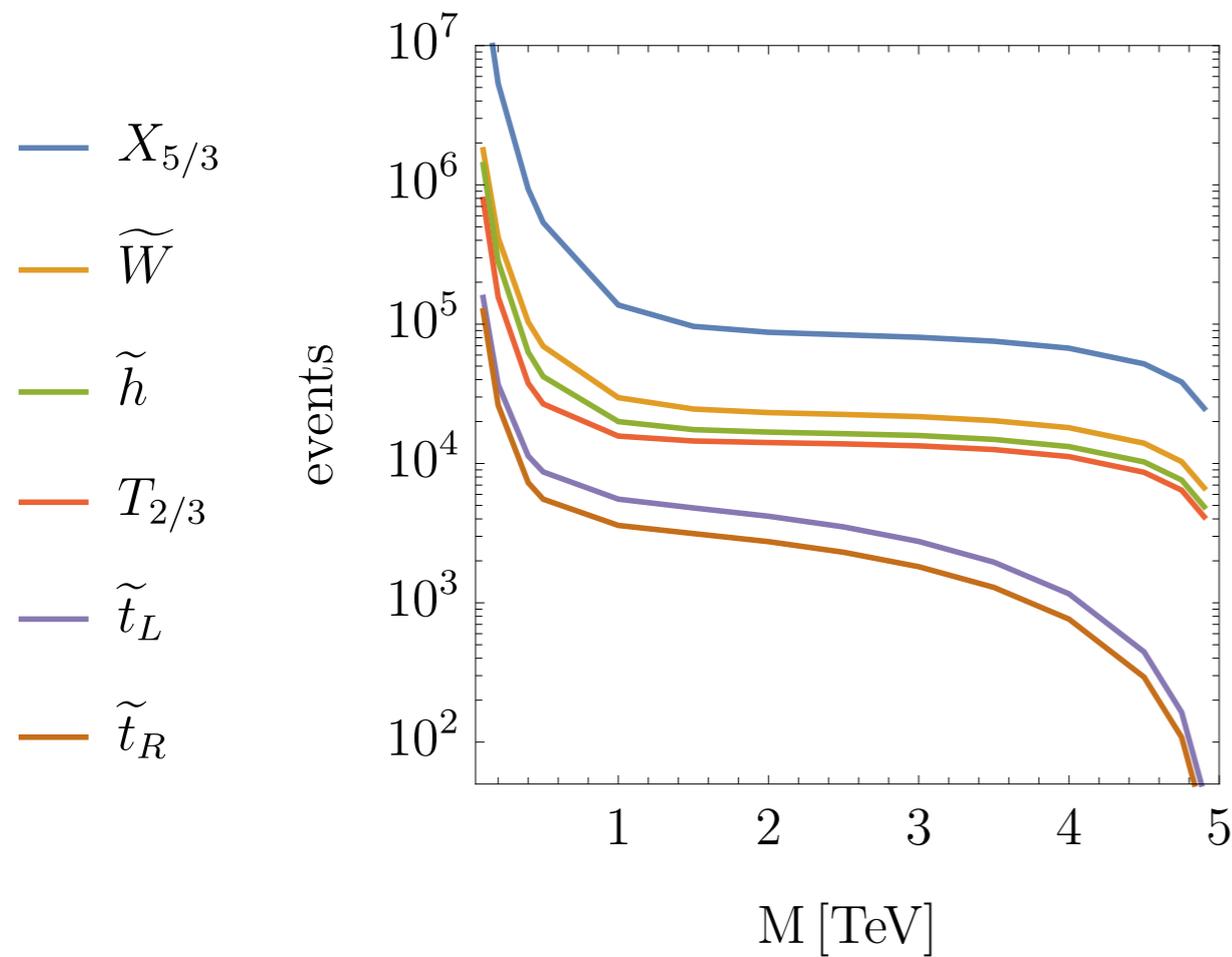


# The case for direct searches

## EW particles discovered up to kinematical threshold

say,  $0.9 E_{\text{cm}}/2 = 4.5 \text{ TeV}$

$10 \text{ TeV } \mu^+ \mu^- , L_{\text{int}} = 10 \text{ ab}^{-1}$



# The case for direct searches

EW particles discovered up to kinematical threshold

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Naturalness implications\*: (“minimal tuning” scenario:  $\Delta=10$  today)

- $\Delta=80$ , from Stops (SUSY) and from Top Partners (CH)
- $\Delta=2500$ , from generic Higgsinos (tree-level tuning)

\*though less popular today than in the past, the question on the origin of the EW scale will keep being asked until when solution found!

# The case for direct searches

EW particles discovered up to kinematical threshold ...  
say,  $0.9 E_{\text{cm}}/2 = 4.5 \text{ TeV}$

... but only if decay final states not “difficult” to see.

Relevant “difficult” cases:

## **Compressed spectra:**

Not studied yet. Perspectives to cover LHC “holes” already at 3 TeV

## **Minimal WIMP Dark Matter:**

Possibly Higgsino, but more general.

“Very direct” signatures from disapp. tracks [realistic BIB included! [2009.11287](#)]

Probed in mono-X [[2009.11287](#), [2107.09688](#), [2203.07351](#)]

“Indirect” probes **above mass-reach** [[1810.10993](#)]

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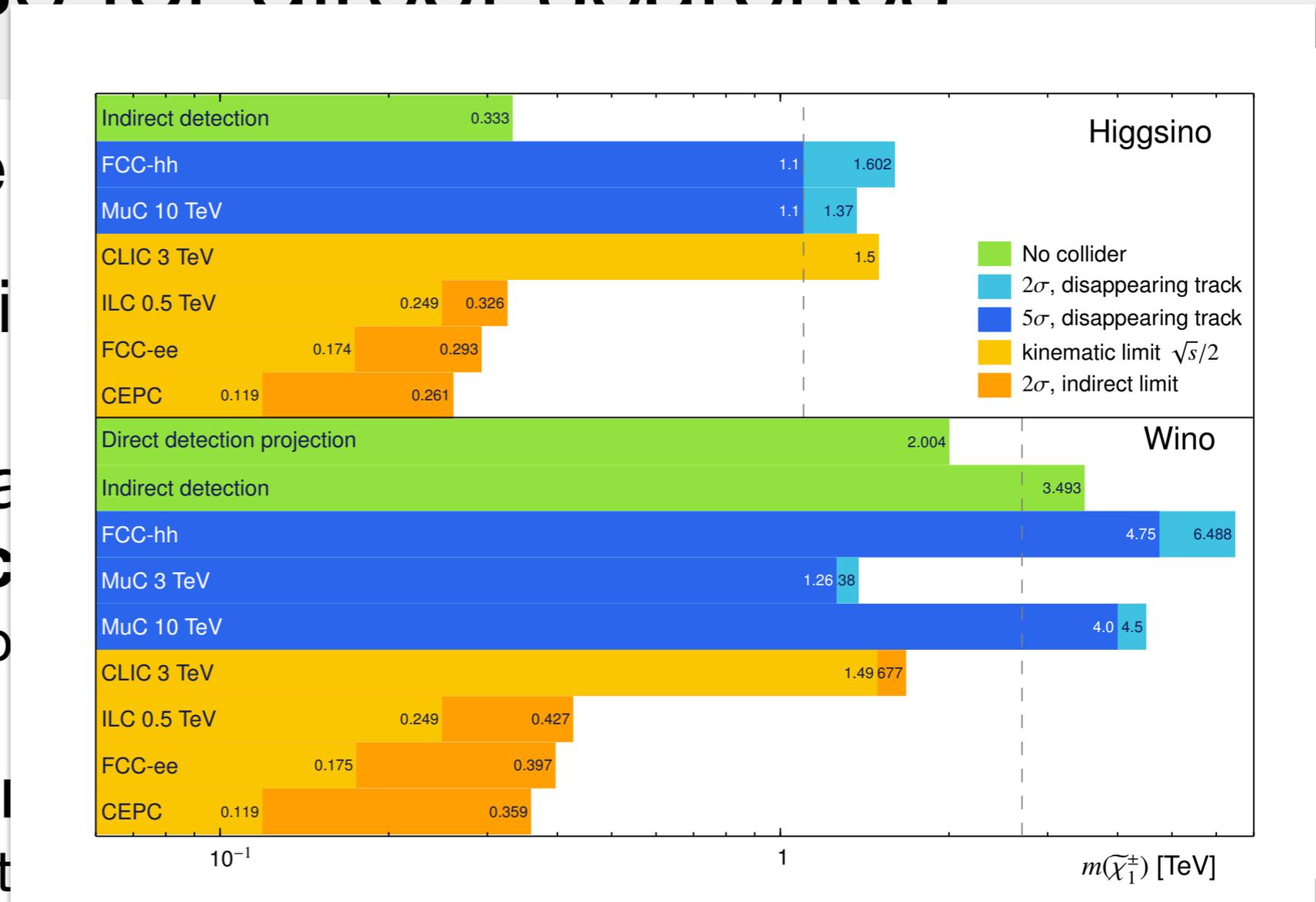
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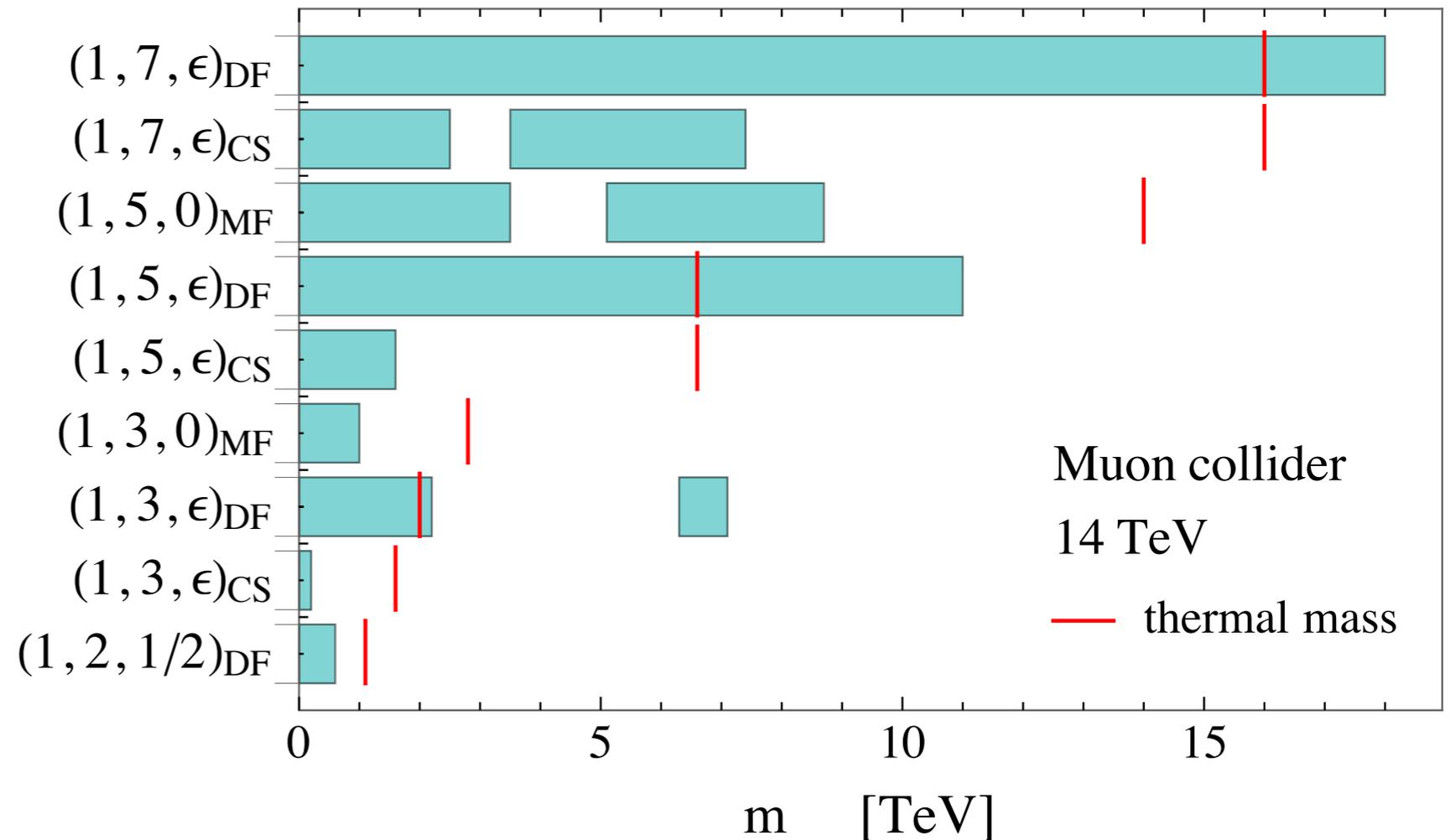
## Minimal WIMP Dark Matter

Possibly Higgsino, bino

“Very direct” signatures

Probed in mono-X [2010.04404]

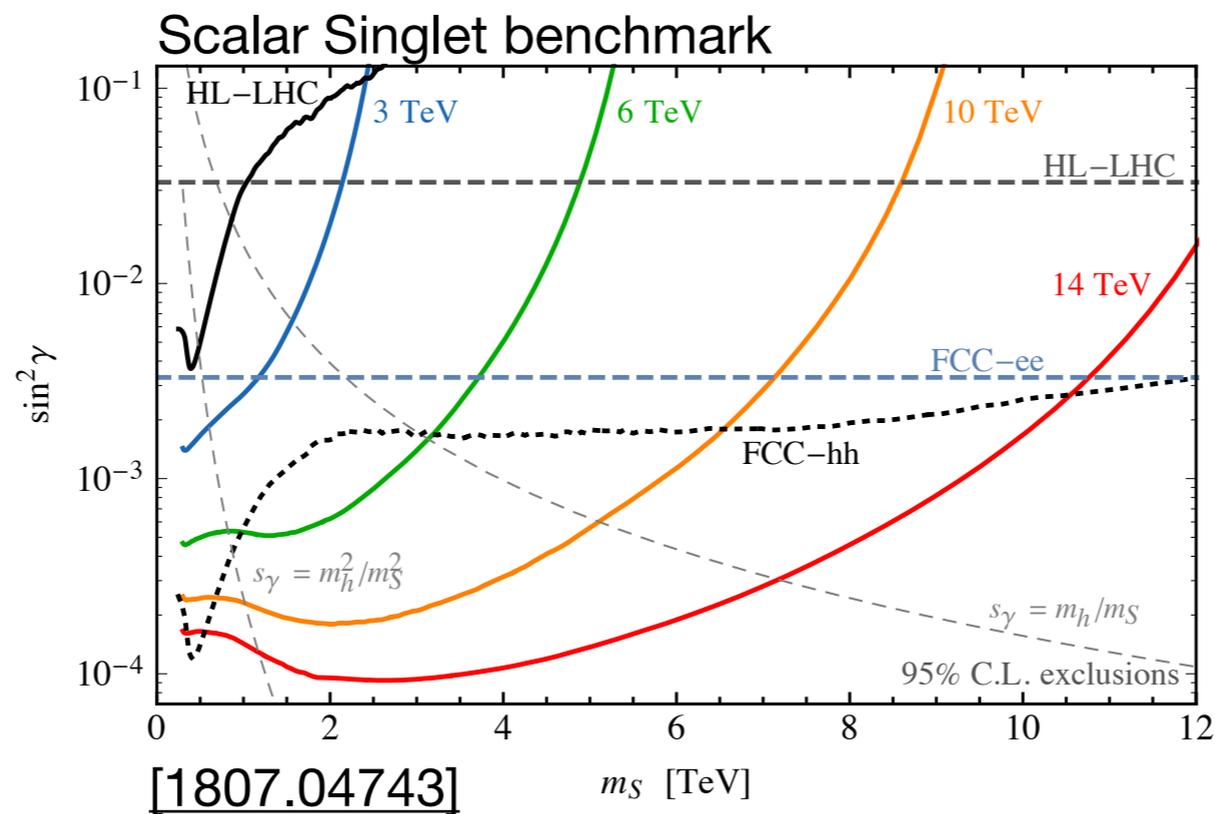
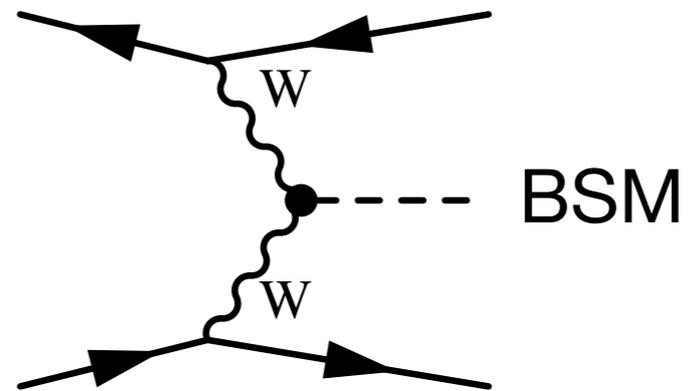
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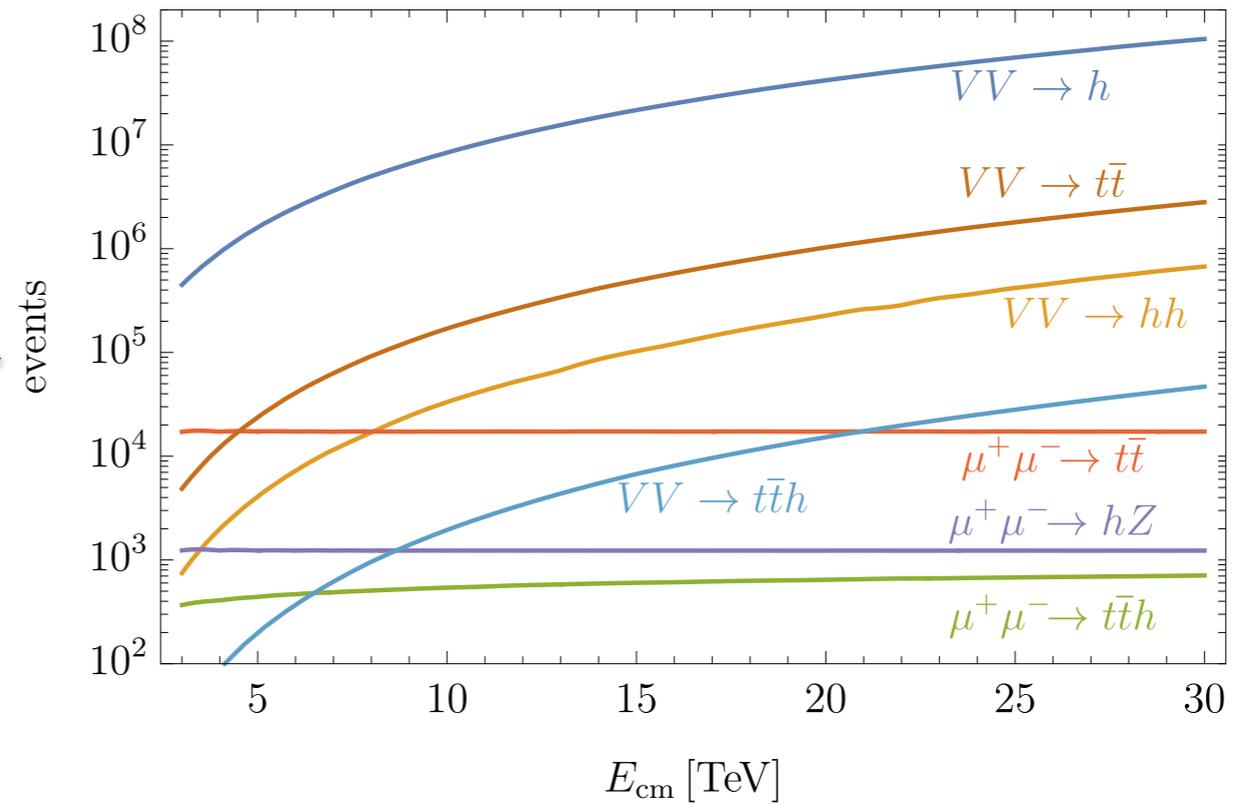
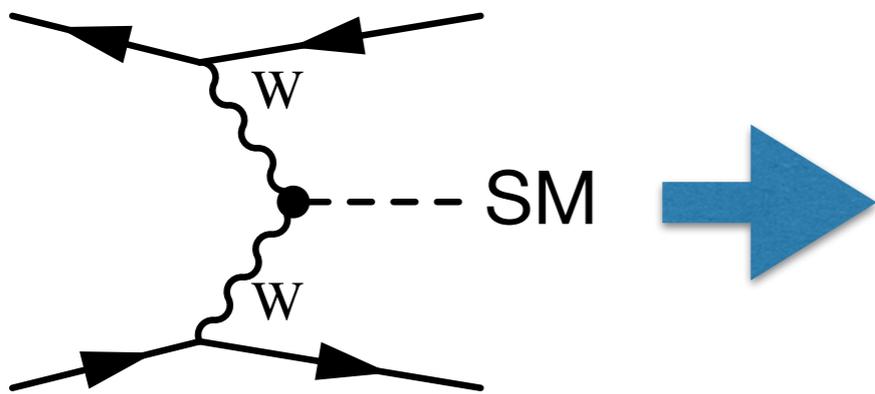
# The case for direct searches

EW-Neutral, **Higgs portal-coupled** particles:

Great Reach, from VBF:



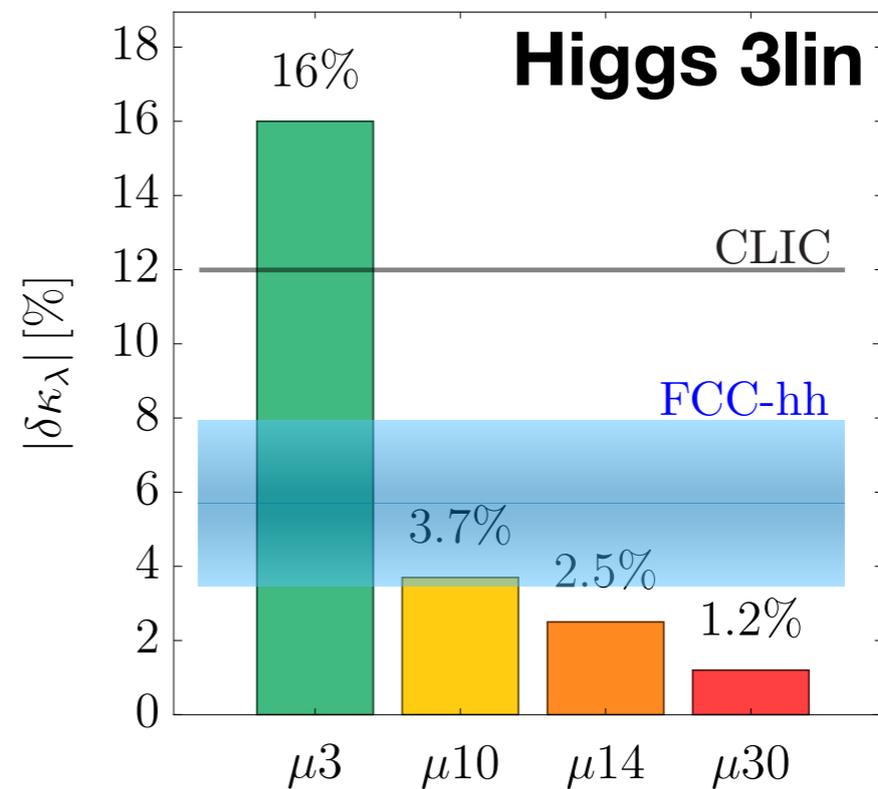
# A vector bosons collider



## Single Higgs couplings

	HL-LHC	HL-LHC +10 TeV	HL-LHC +10 TeV + ee
$\kappa_W$	1.7	0.1	0.1
$\kappa_Z$	1.5	0.4	0.1
$\kappa_g$	2.3	0.7	0.6
$\kappa_\gamma$	1.9	0.8	0.8
$\kappa_c$	-	2.3	1.1
$\kappa_b$	3.6	0.4	0.4
$\kappa_\mu$	4.6	3.4	3.2
$\kappa_\tau$	1.9	0.6	0.4
$\kappa_{Z\gamma}^*$	10	10	10
$\kappa_t^*$	3.3	3.1	3.1

\* No input used for  $\mu$  collider



# High energy probes

[Buttazzo, Franceschini, AW, 2020]

As simple as this:

$$\frac{\Delta\sigma(E)}{\sigma_{\text{SM}}(E)} \propto \frac{E^2}{\Lambda_{\text{BSM}}^2} \quad [\text{say, } \Lambda_{\text{BSM}} = 100 \text{ TeV}]$$

**=**

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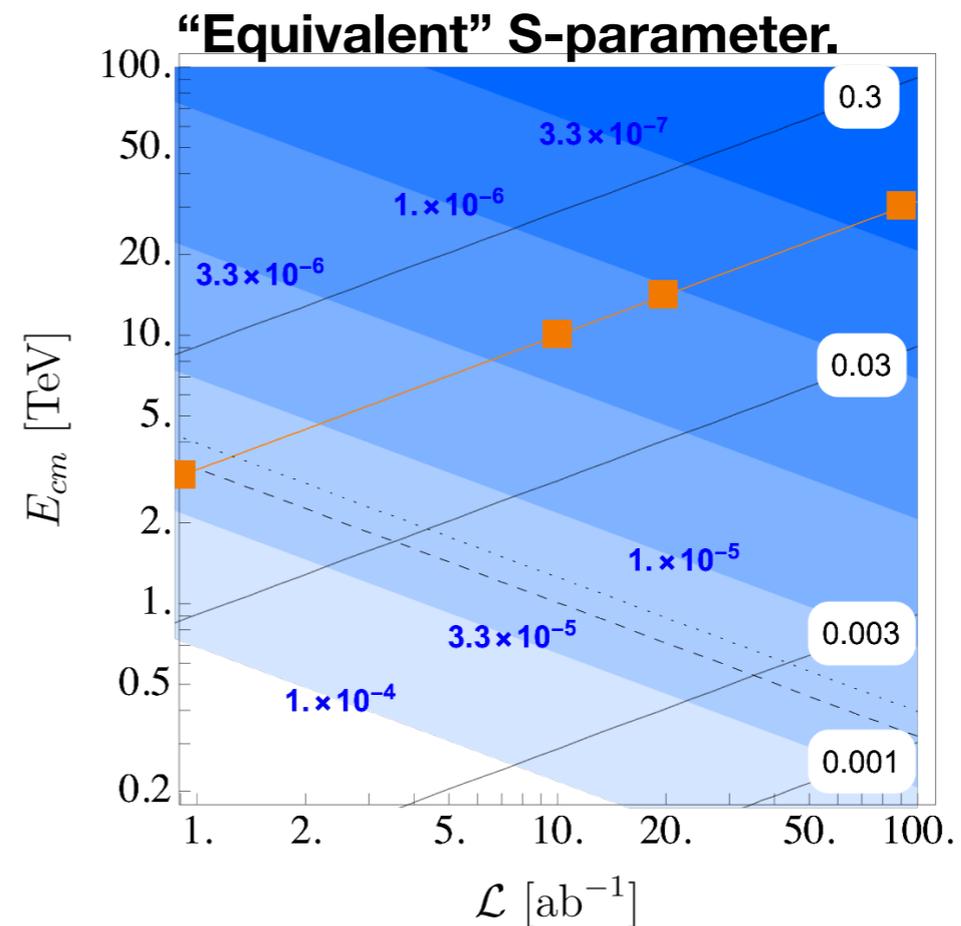
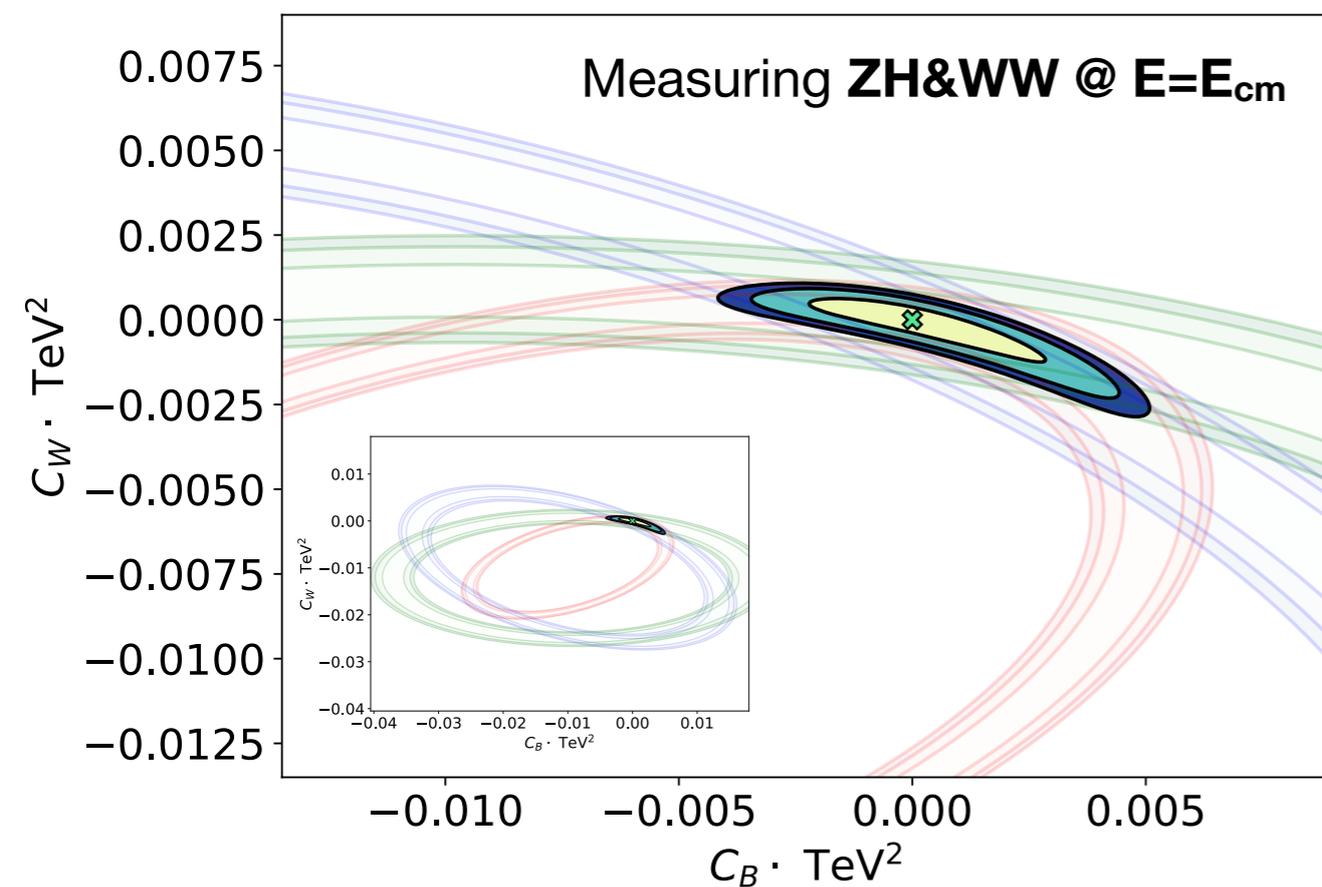
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[Buttazzo, Franceschini, AW, 2020]

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$$\frac{\Delta\sigma(E)}{\sigma_{\text{SM}}(E)} \propto \frac{E^2}{\Lambda_{\text{BSM}}^2} \quad [\text{say, } \Lambda_{\text{BSM}} = 100 \text{ TeV}] = \begin{cases} 10^{-6} \text{ at EW [FCC-ee] energies} \\ 10^{-2} \text{ at muon collider energies} \end{cases}$$



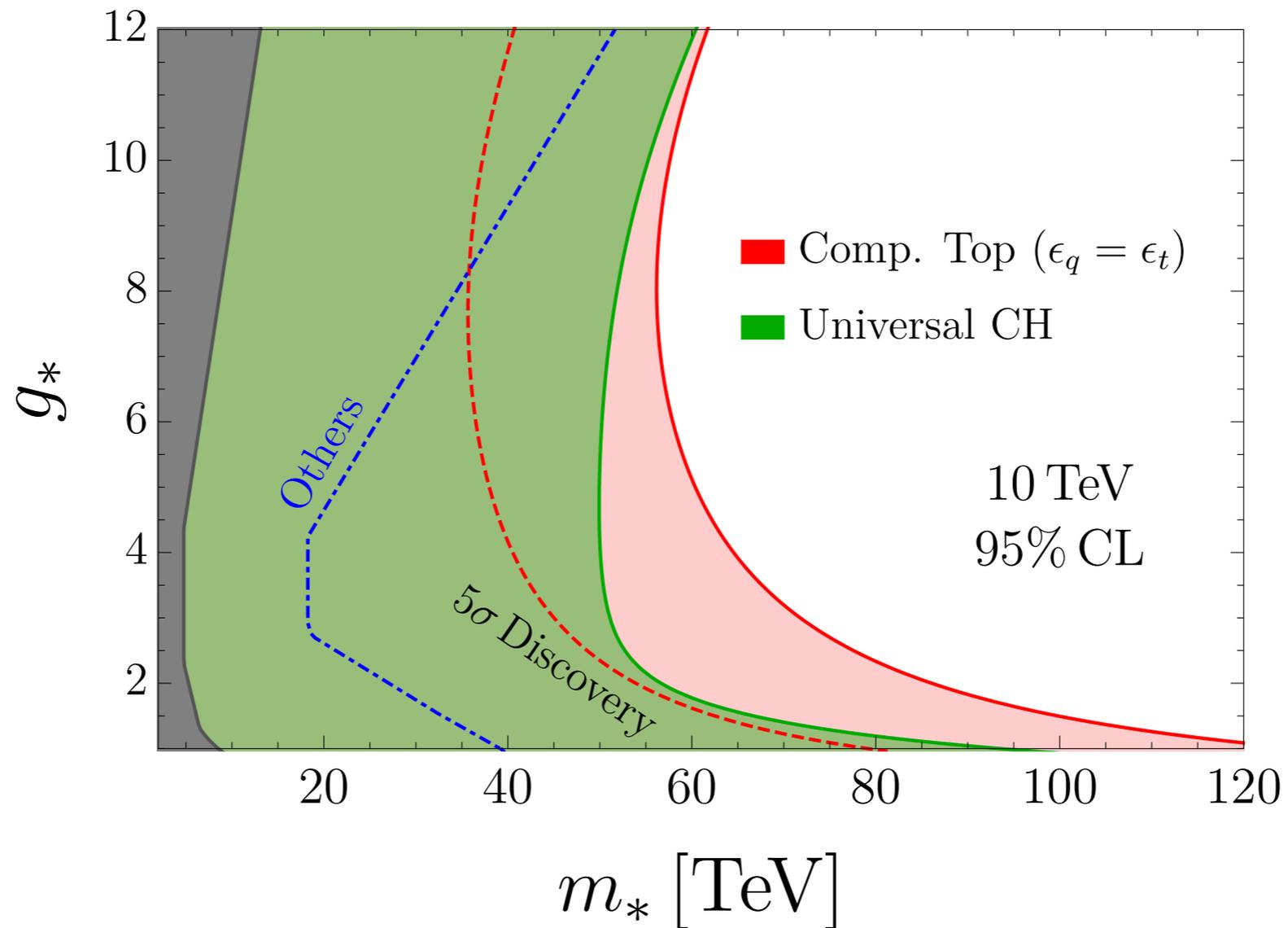
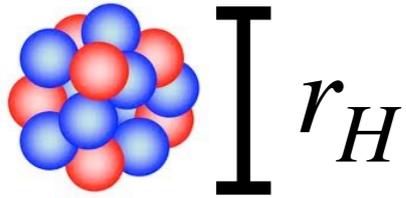
Today:  
**10<sup>-3</sup>**

**MUC14**

**10<sup>-6</sup>**

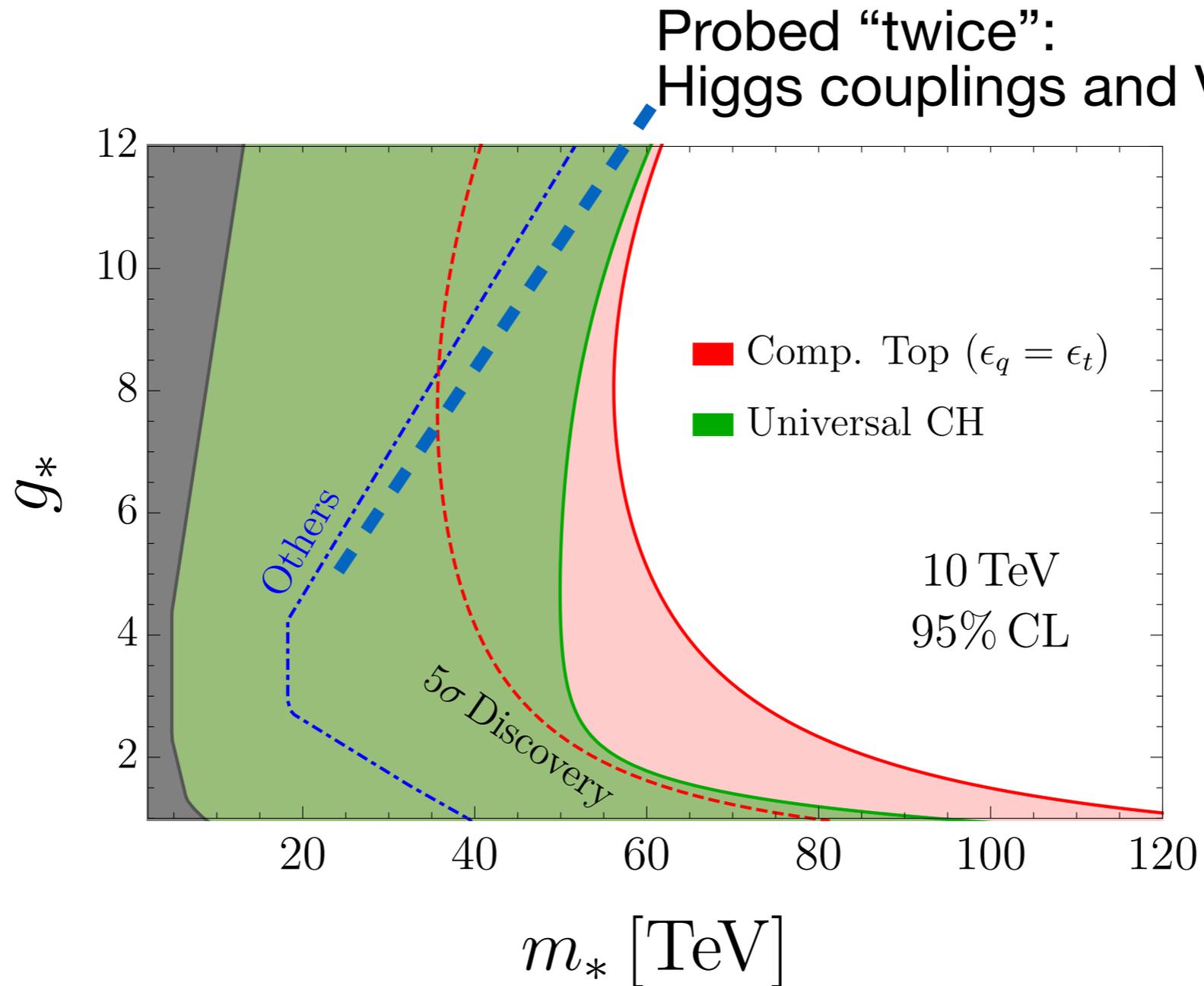
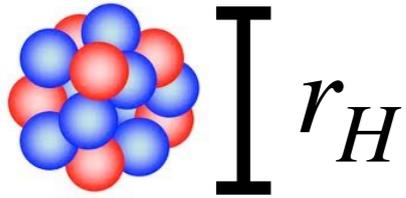
# Composite Higgs, indirect

[Chen, Glioti, Rattazzi, Ricci, AW]



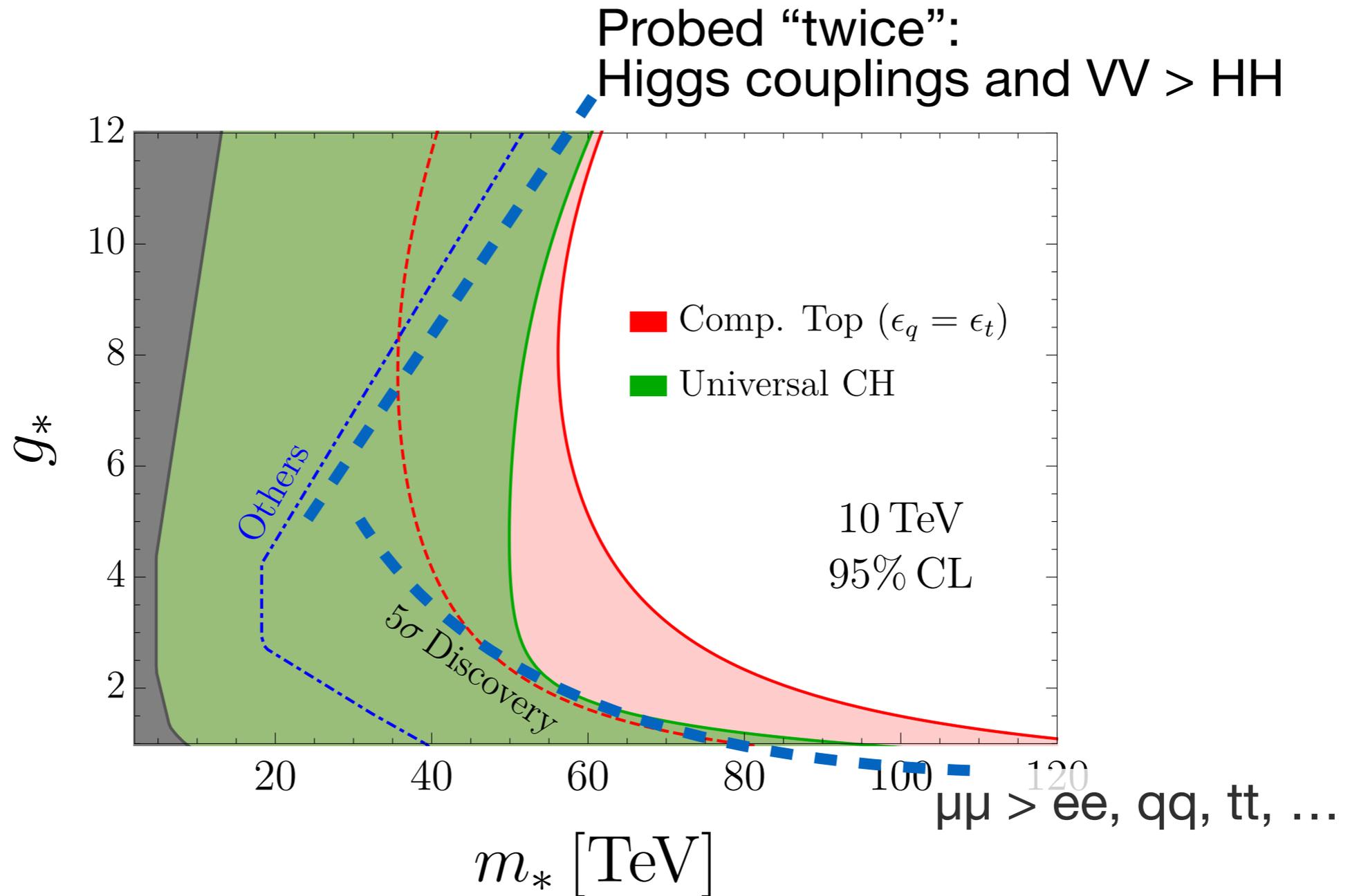
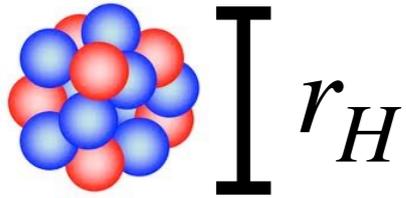
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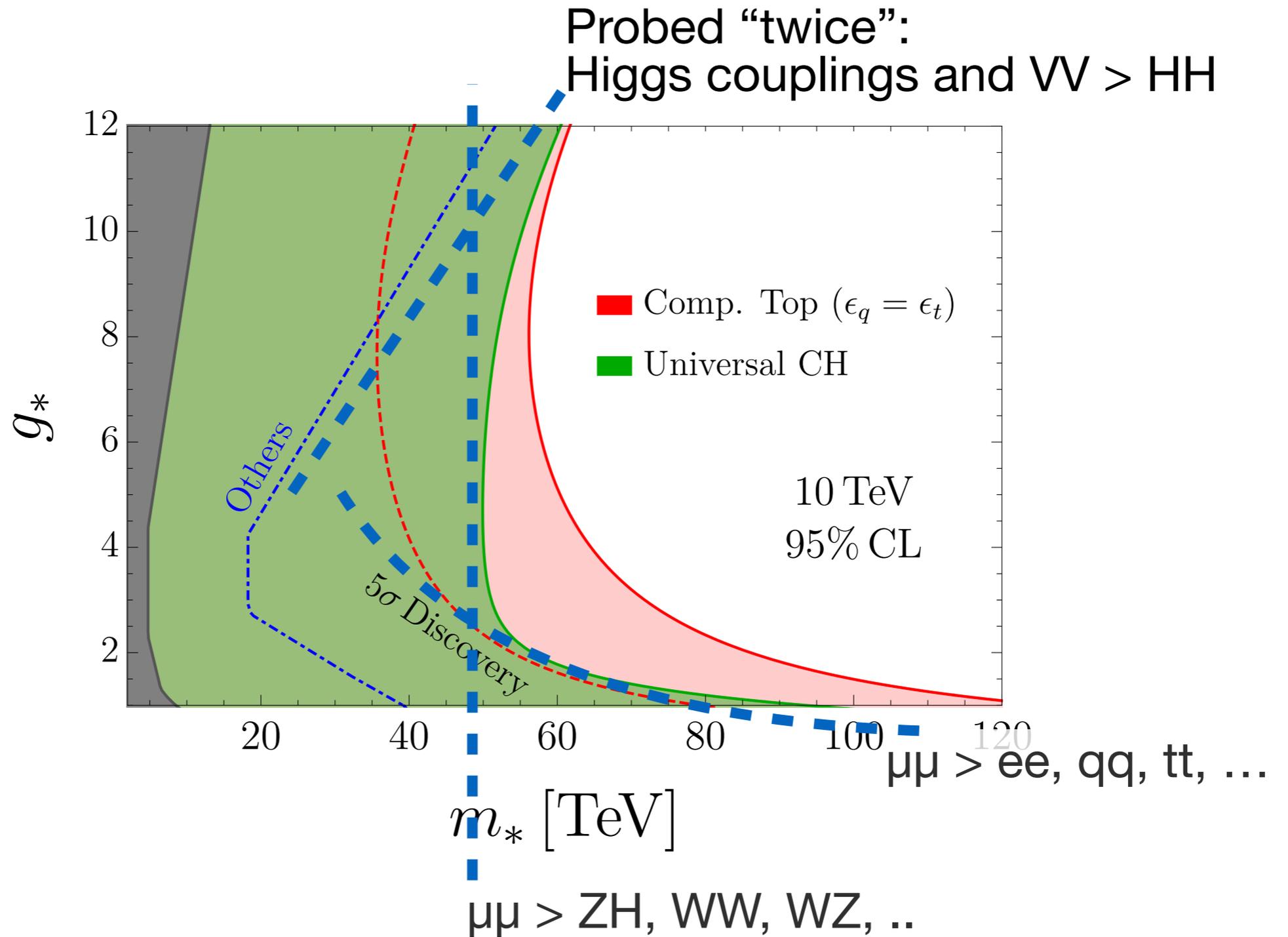
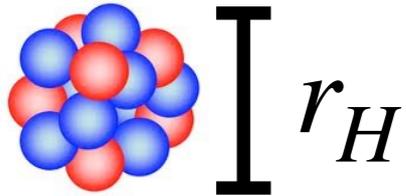
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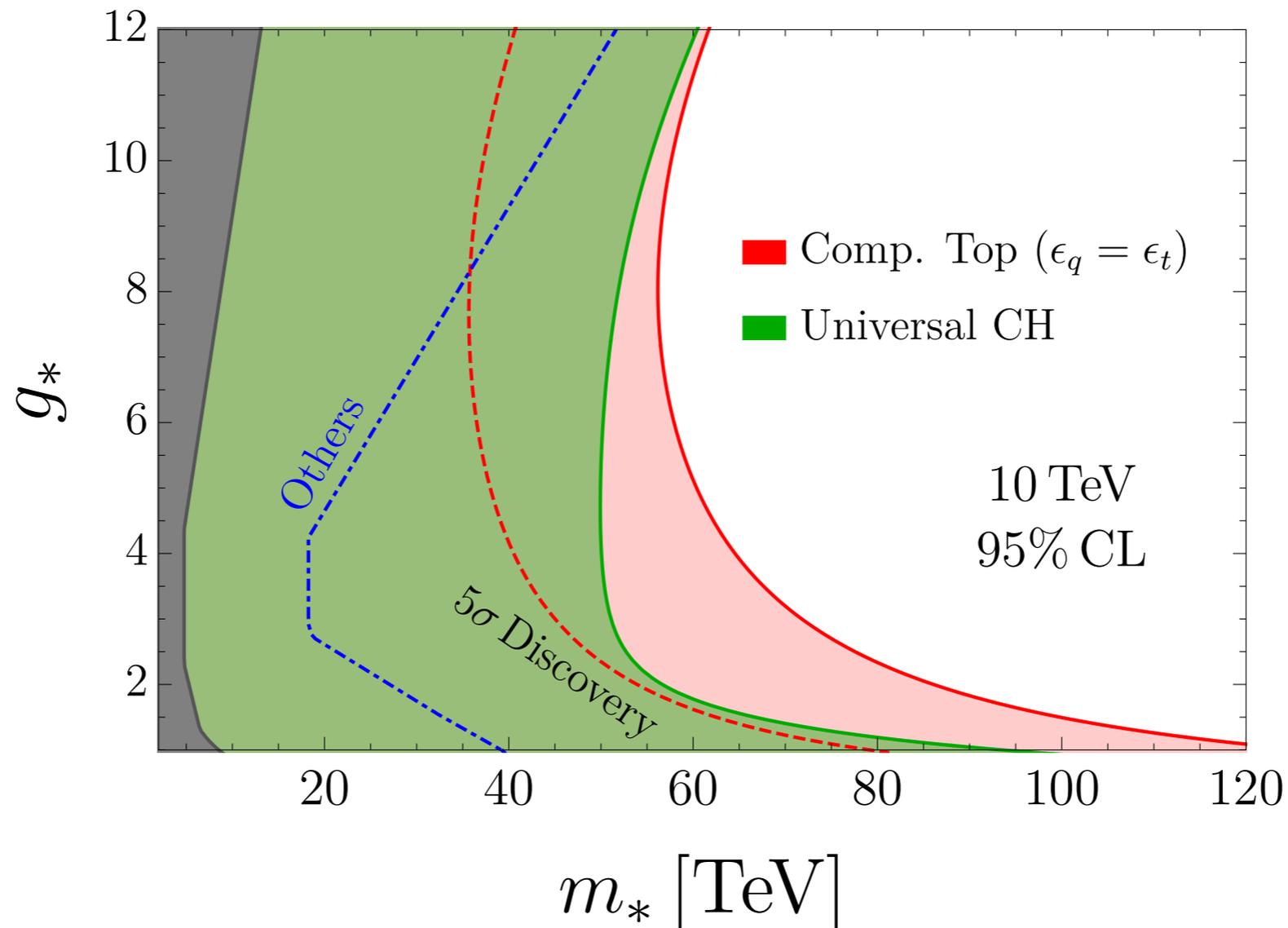
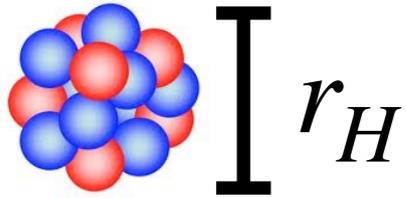
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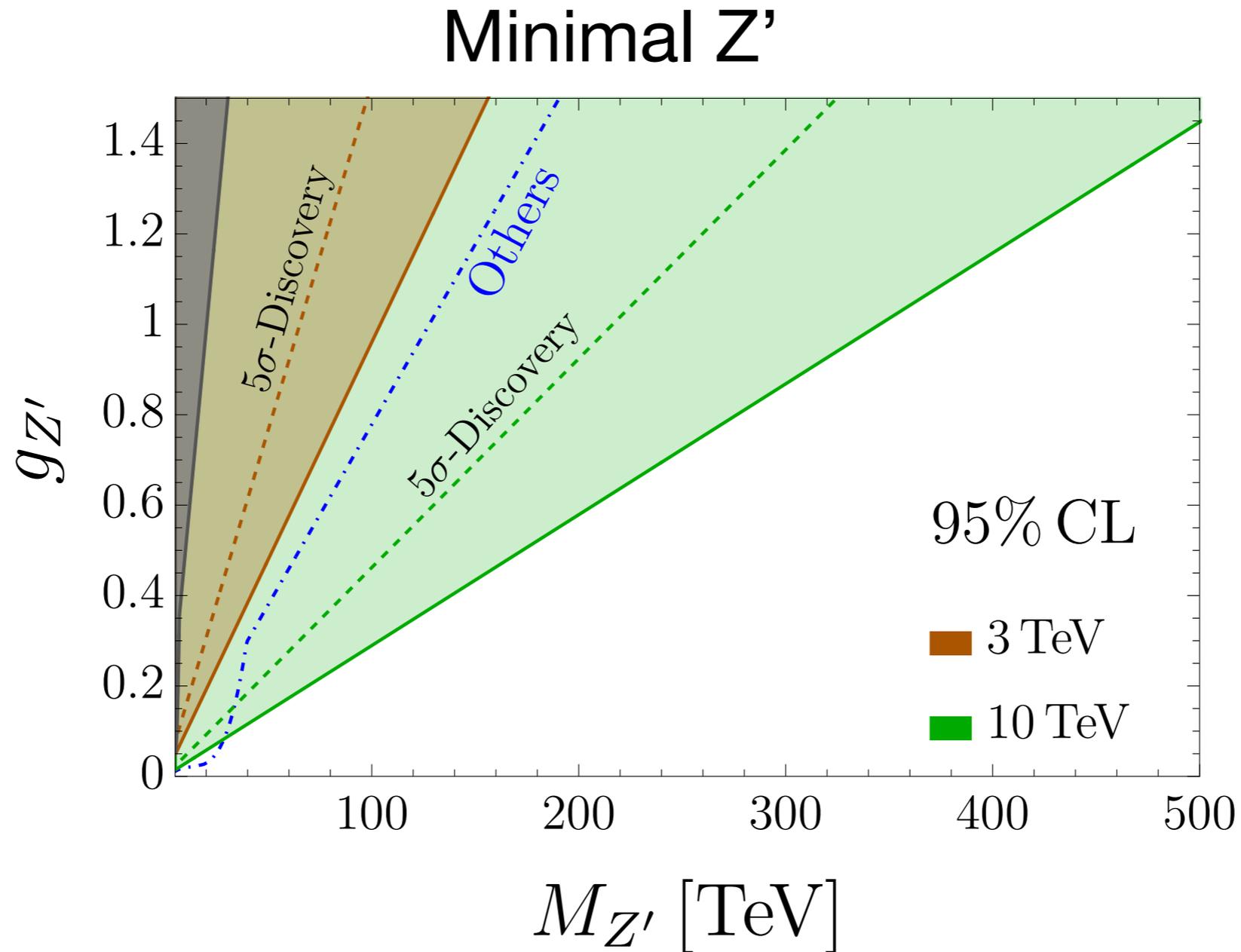


**Superior** reach on Higgs radius from measurements at the high collider energy, where Higgs size effects are enhanced.

Proton size discovered this way: raising energy until close enough to  $r_P^{-1}$   
We might be lucky again!

# Same mechanism, simpler model

[Chen, Glioti, Rattazzi, Ricci, AW]



Generically, we can test EW interactions at  $> 100$  TeV scale.

# Is this potential **real**?

Remember what we expect from colliders:

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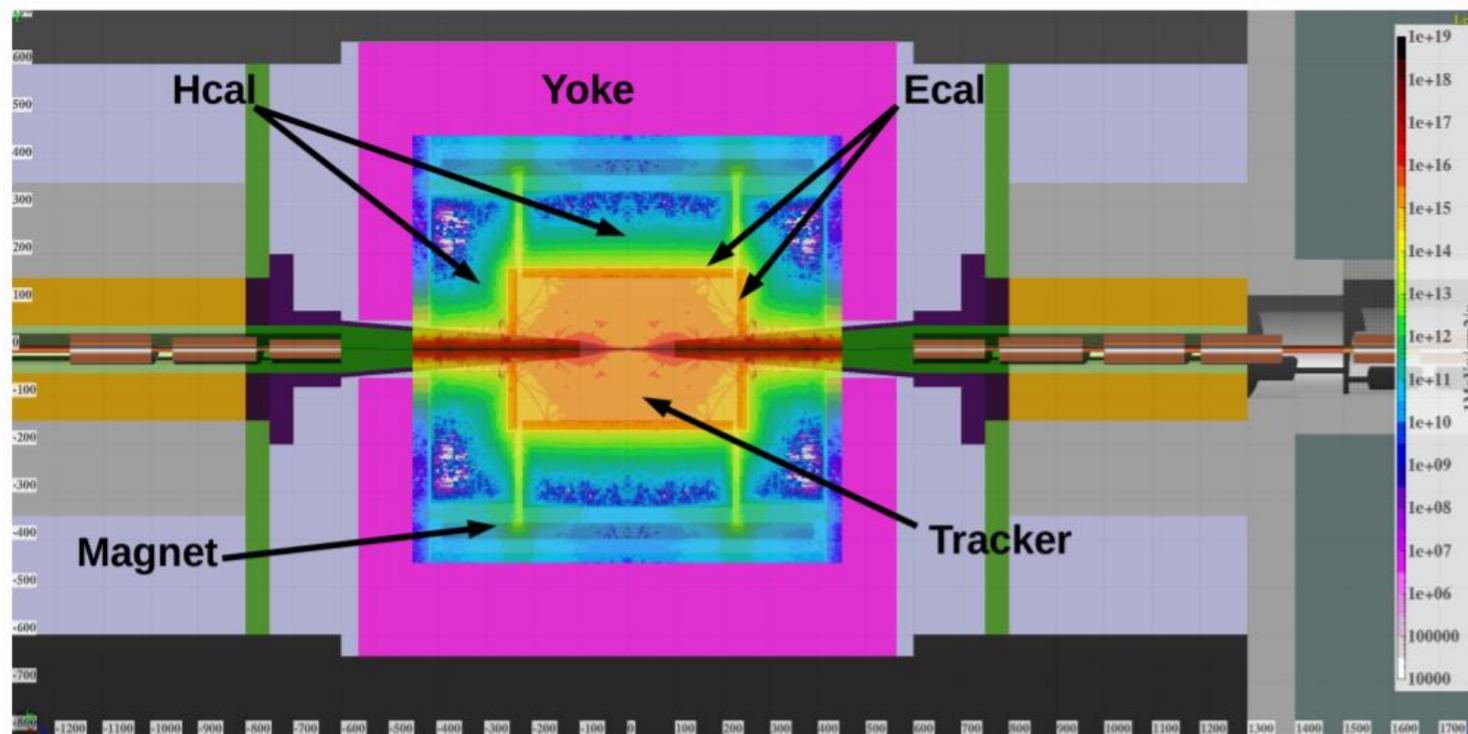
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Background from decaying muons (BIB)



FLUKA @ 1.5 TeV

**New Challenge** for  
Detector@Analysis design

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→ **EW Infrared** logarithms are **order one** at MUC energies

Accurate **resummation** is needed.

[[Manohar and Waalewijn, 2018, ...](#)]

As well as accurate **EW showering**.

[[Chen, Han, Tweedie, 2016; Han, Ma and Xie, 2021, ...](#)]

**NOT** an easy extension of QED/QCD radiation treatment

Because of the peculiarities of broken gauge theories

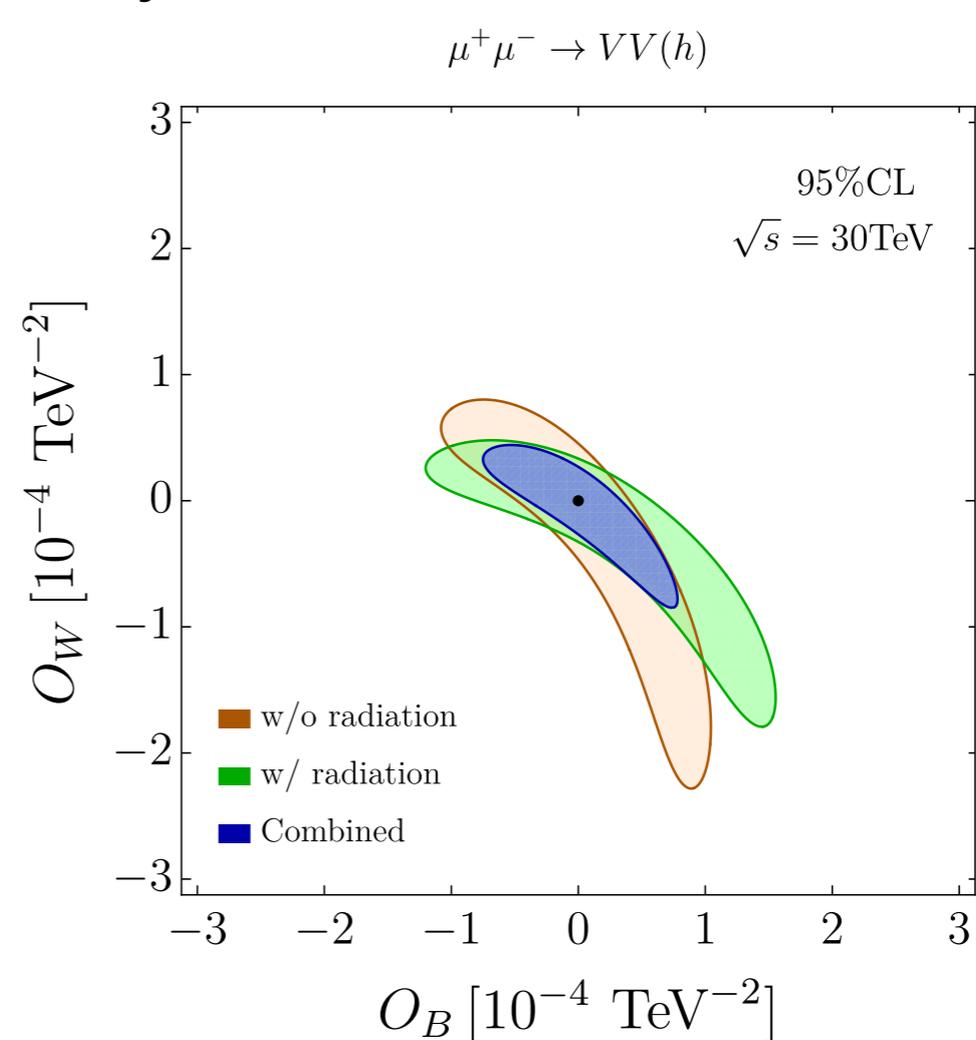
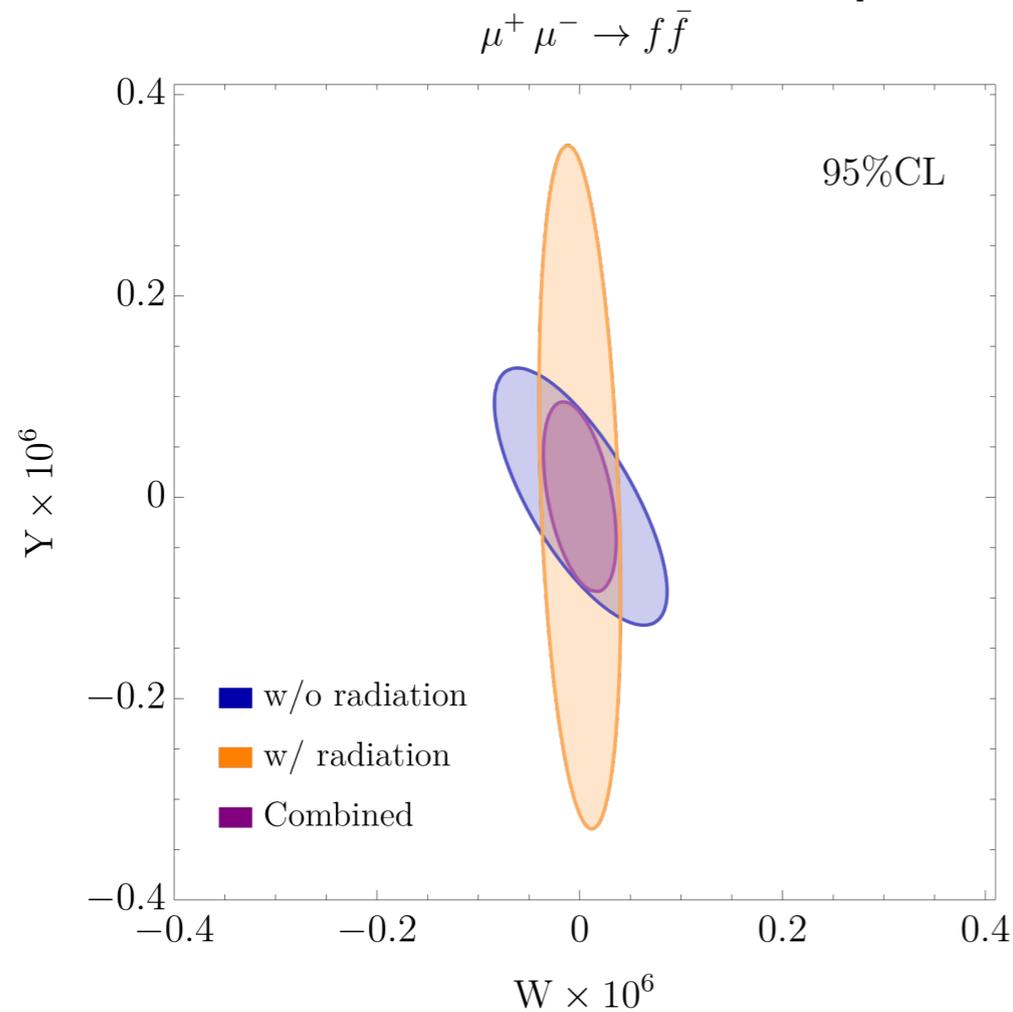
Because of the accuracy we need

Because from radiation structure we can learn about New Physics!

[[Chen, Glioti, Rattazzi, Ricci, AW](#)]

# Is this potential real?

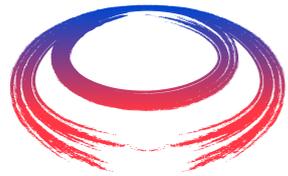
We can, e.g., access charged current int. from  $W$  in.state radiation.  
Exclusive/semi-inclusive complementarity in EFT interactions sensitivity.



Because from radiation structure we can learn about New Physics!

[Chen, Glioti, Ricci, Rattazzi, AW, to appear]

# Is the collider feasible?



International  
UON Collider  
Collaboration

Objective:

In time for the next European Strategy for Particle Physics Update, the study aims to **establish whether the investment into a full CDR and a demonstrator is scientifically justified.**

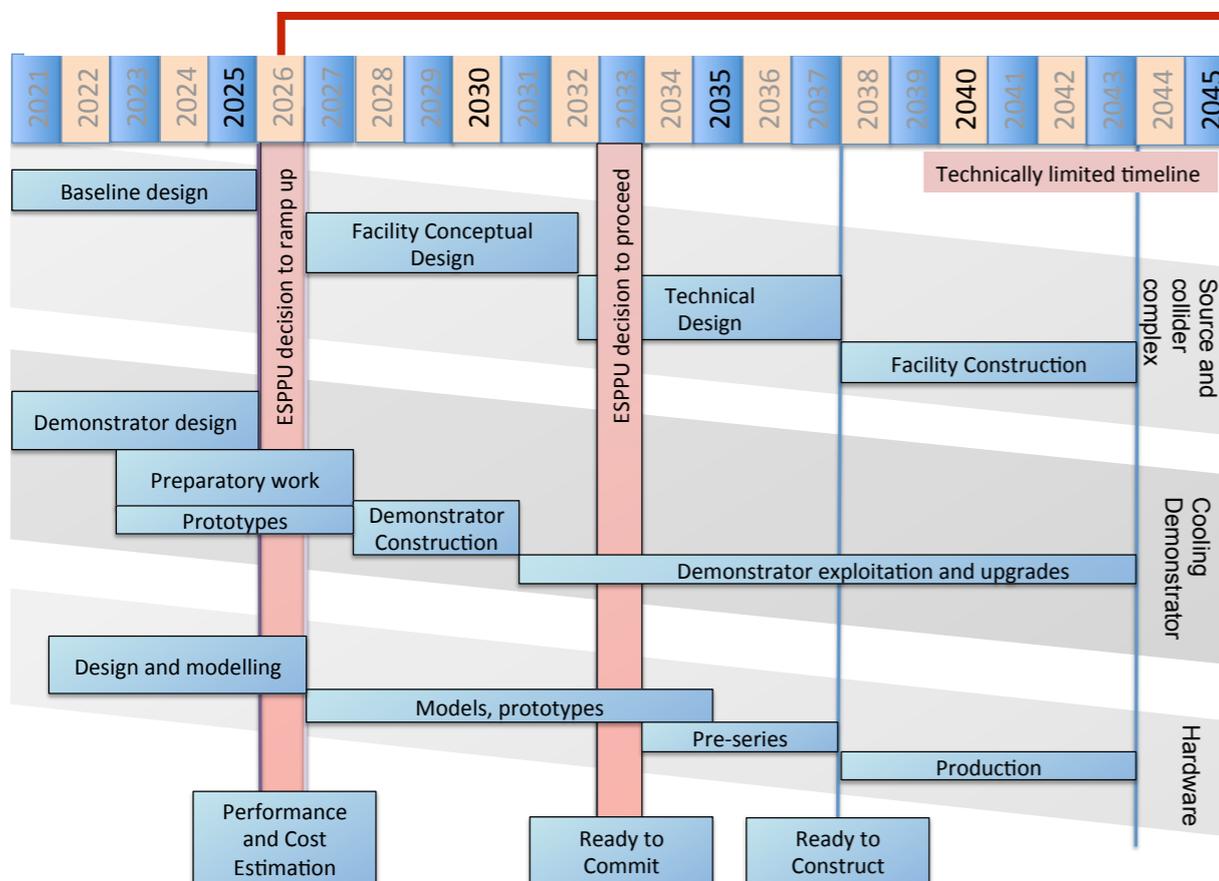
[Daniel Schulte, IMCC head. [link](#)]

Scope:

- Focus on two energy ranges:
  - **3 TeV**, if possible with technology ready for construction in 10-20 years
  - **10+ TeV**, with more advanced technology, **the reason to chose muon colliders**

[Daniel Schulte, IMCC head. [link](#)]

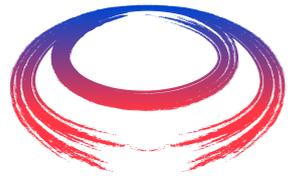
Tentative Target for Aggressive Timeline



We might know the answer in few years

A summary, and a work plan, available [here](#)

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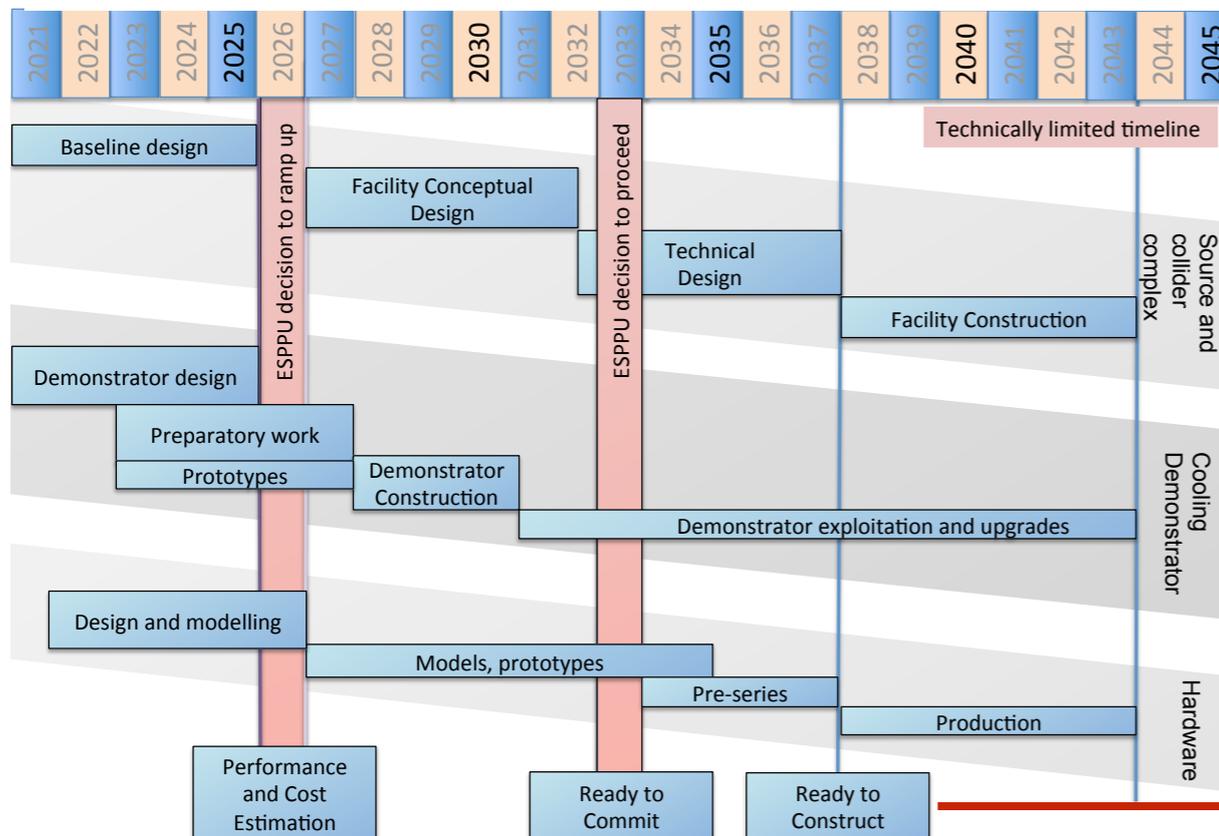
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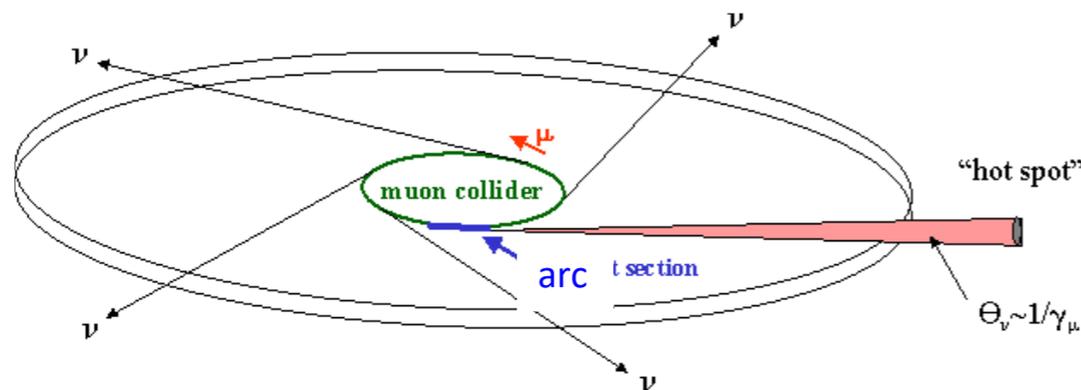
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A first MUC, maybe at 3 TeV, could start being built as early as 2038!

# Is the collider feasible?

One famous possible issue is radiation from h.e. neutrinos

## Neutrino Flux Mitigation

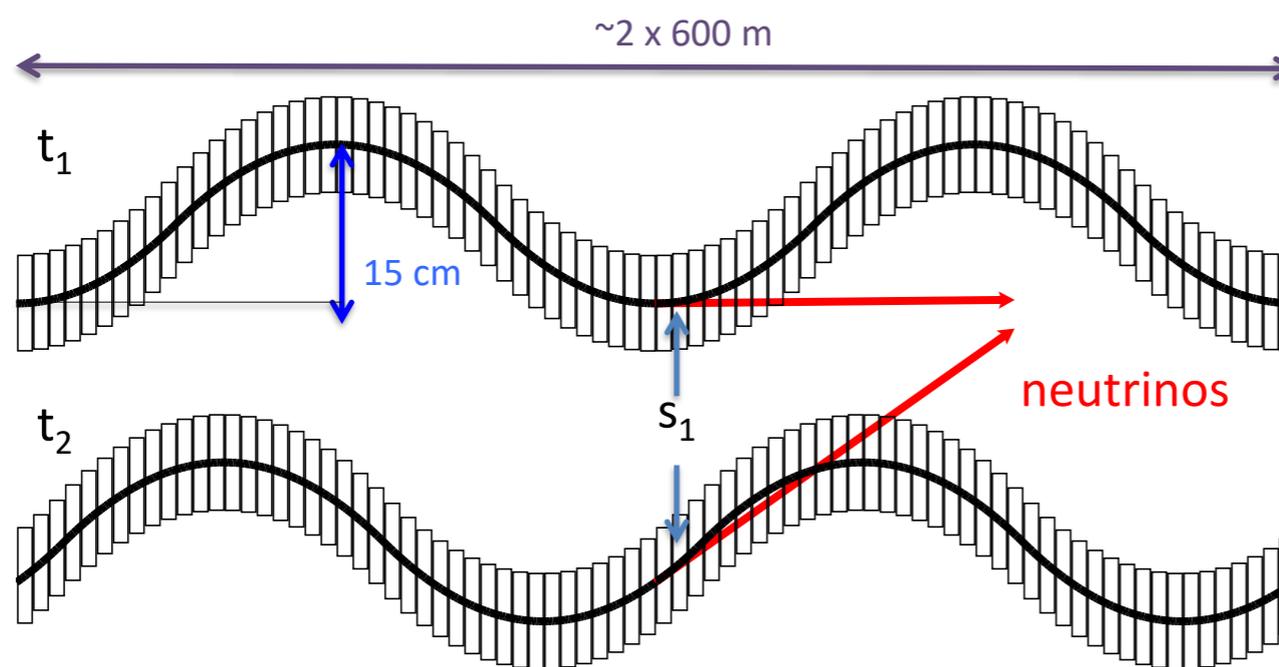


Concentrate neutrino cone from arcs can approach legal limits for 14 TeV

**Goal is to reduce to level similar to LHC**

**3 TeV, 200 m deep tunnel is about OK**

**Need mitigation of arcs at 10+ TeV:** idea of Mokhov, Ginneken to move beam in aperture  
Our approach: move collider ring components, e.g. vertical bending with 1% of main field



Opening angle  $\pm 1$  mradian

**14 TeV, in 200 m deep tunnel comparable to LHC case**

**Need to study mover system, magnet, connections and impact on beam**

**Working on different approaches for experimental insertion**

# (Emerging) Conclusions

For 10 TeV MuC! Much better if higher Energy!

## Muon Colliders BSM Physics Pillars:

- **High Energy** available for direct particles production
- **High Rate** available for **Precision** measurements
- **Energy and Precision** → probing EW in 100 TeV ballpark

All this, at a single collider with feasible timescale

## WIMP Dark Matter

- **Higgsino/Wino “very directly” accessible** (more is coming on WIMP at 10 TeV)

## Explaining the origin of the Weak scale

- $\Delta=10 \rightarrow \Delta=80$  , simply from direct searches
- Probe  $\Delta=1/\xi=1000$ , in CH, in two ways

## How much is the Higgs radius?

- If as “large” as  $1/(50\text{TeV})$ , we can tell

## Added value from colliding muons [2203.07261]

- Current anomalies are in muons. If coincidence, a fortunate one: illustrates obvious potential of colliding  $\mu$  for the first time!

Thank You