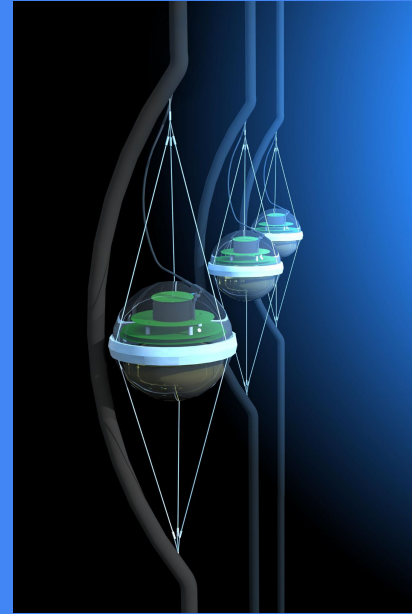


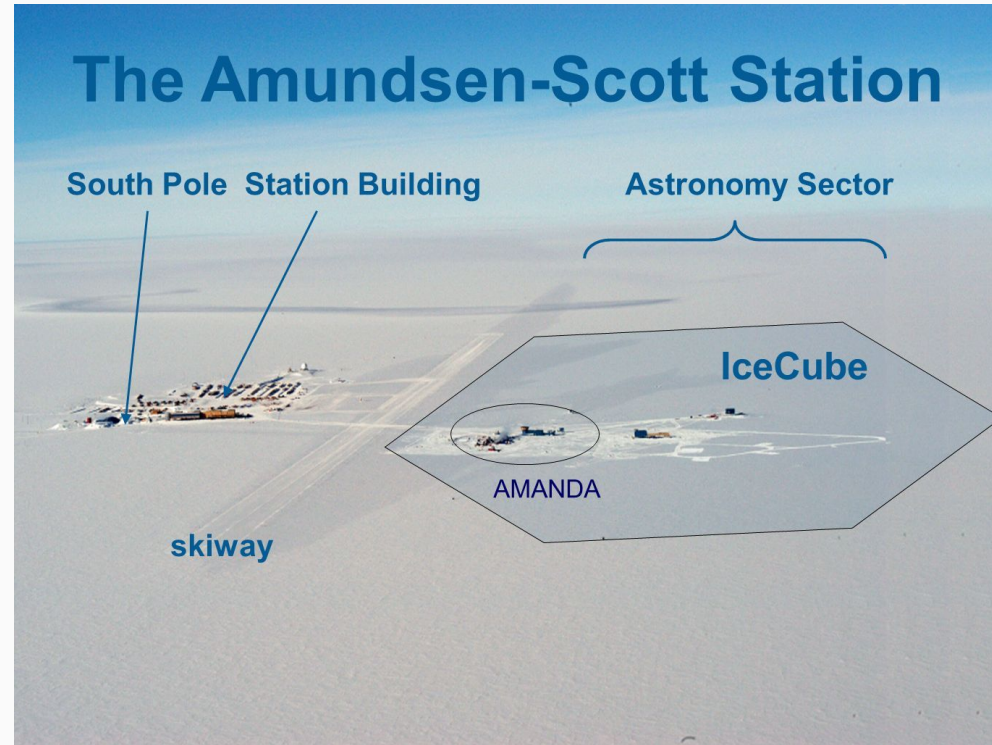
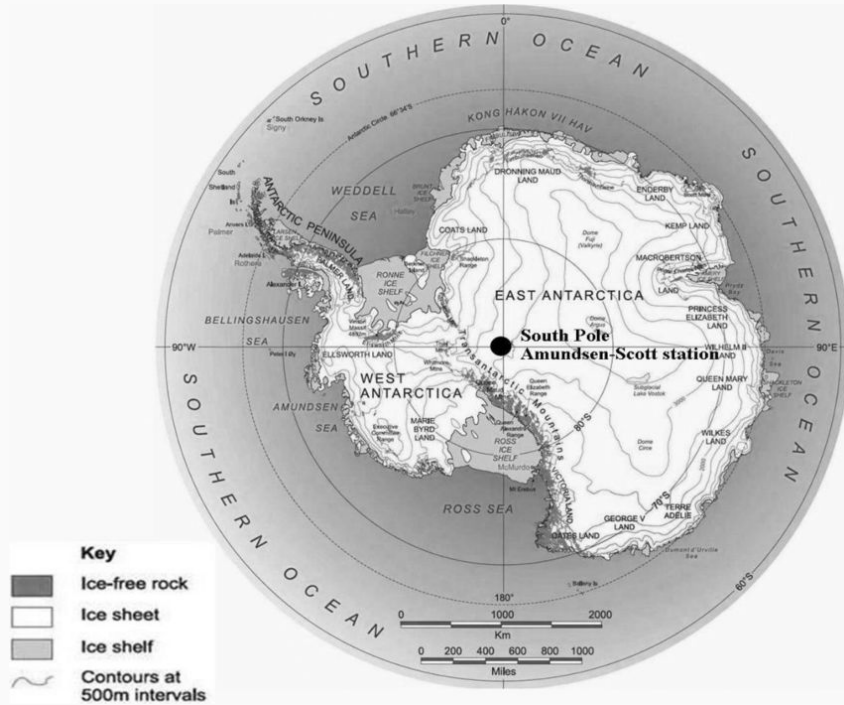
# Deep learning in the IceCube detector status - prospects - challenges



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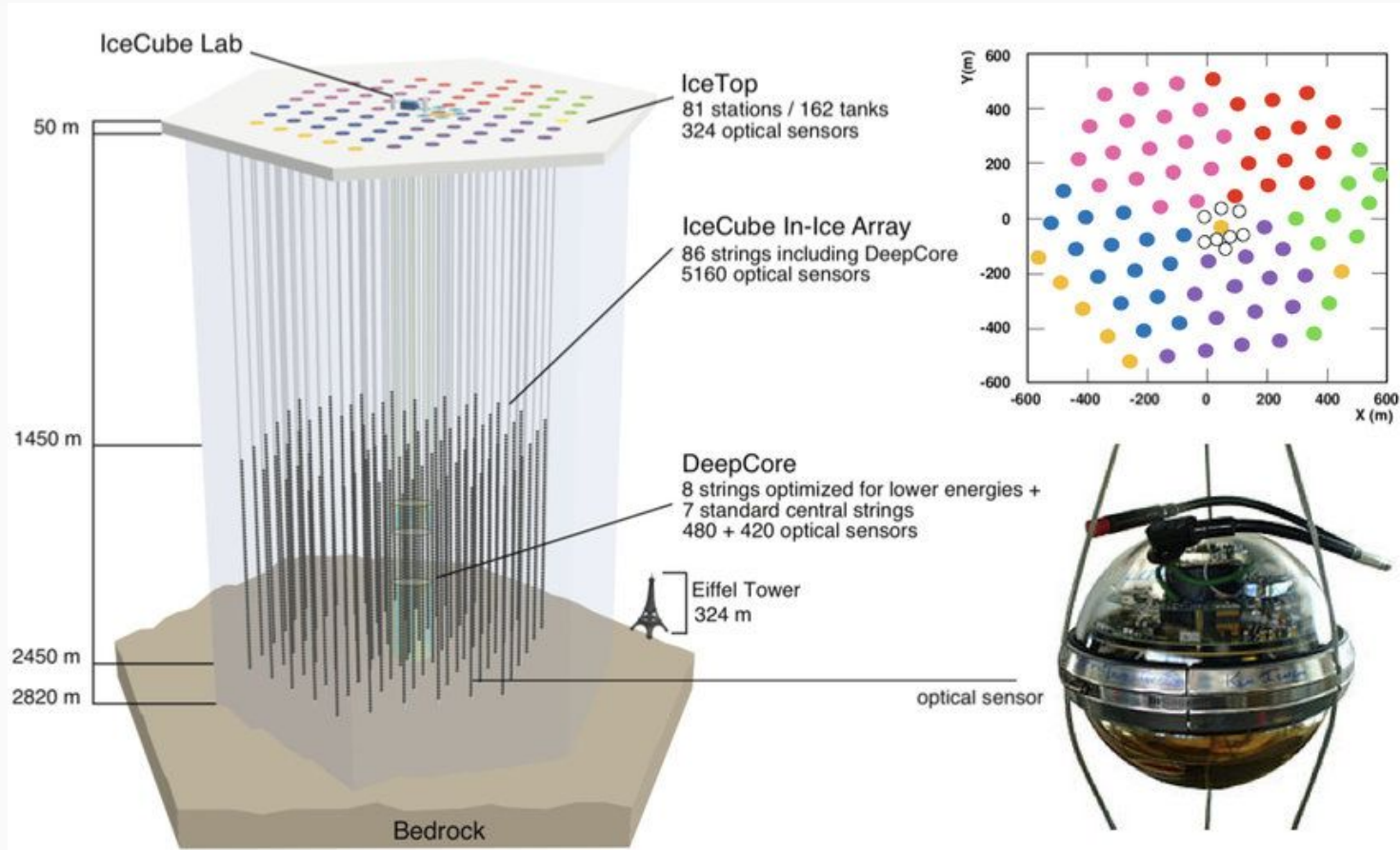
Thorsten Glüsenkamp, 16.6. 2023



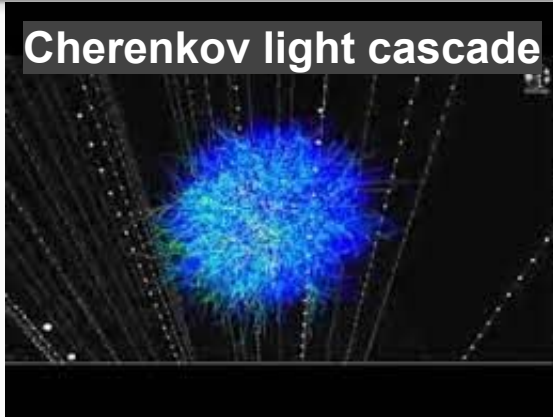
# IceCube - the detector



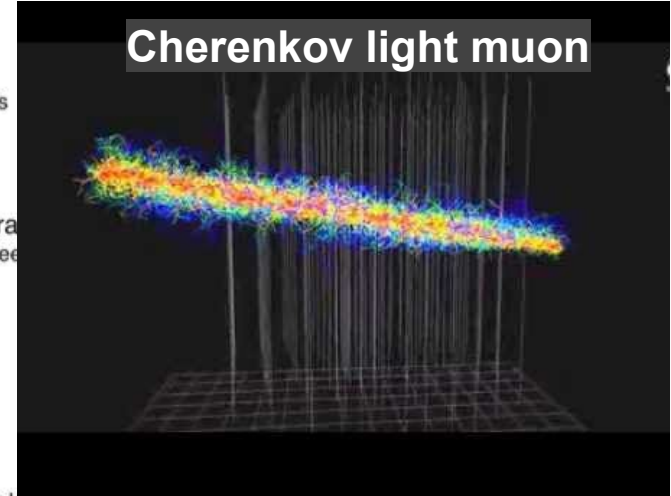
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## Cherenkov light cascade



## Cherenkov light muon



$v_e$

e-cascade

$\mu$

1450 m

2450 m

2820 m



IceTop

81 stations / 162 tanks  
324 optical sensors

IceCube In-Ice Array

86 strings including DeepCore  
5160 optical sensors

DeepCore

8 strings optimized for lower energies +  
7 standard central strings  
480 + 420 optical sensors

Eiffel Tower  
324 m

optical sensor

Bedrock

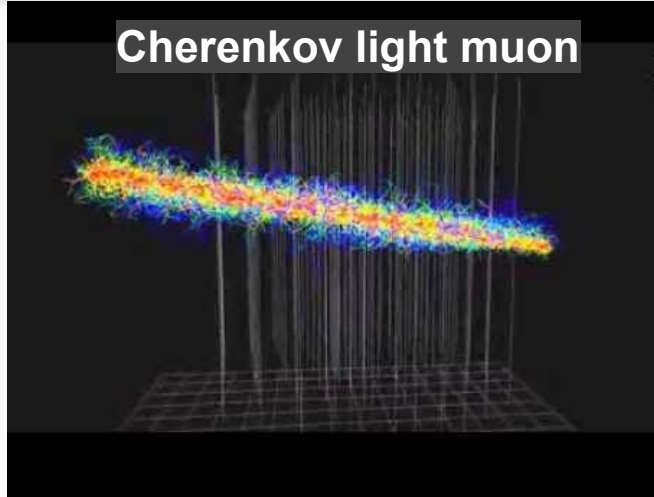
velocity of charged particle larger than speed of light in ice  
( $\sim 0.75 \cdot c$ )  $\rightarrow$  emission of Cherenkov light



Digital  
Optical  
Module

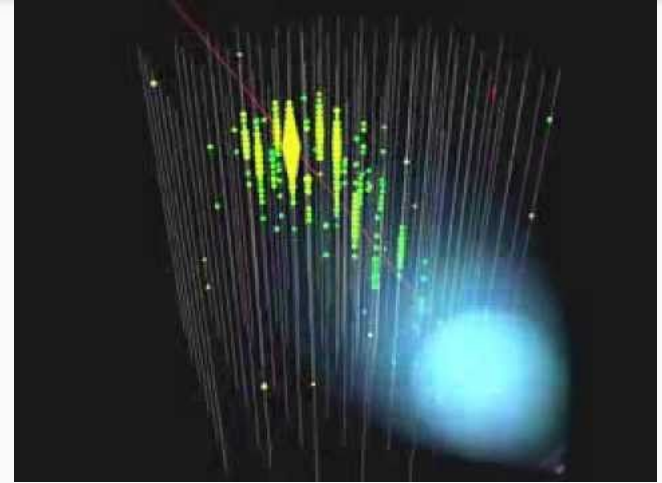


## Cherenkov light muon

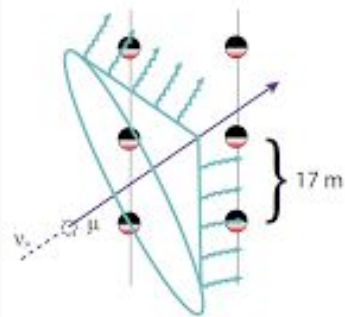


**Blob size:**  
**Number of photons**

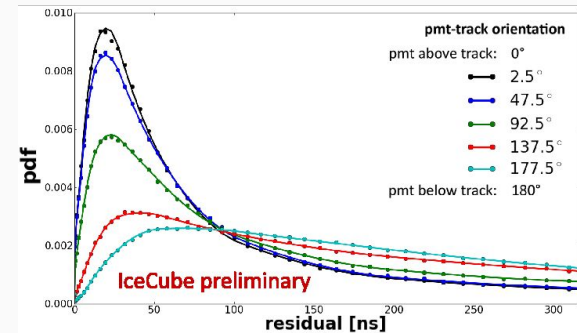
**Blob color:**  
**Photon arrival time**



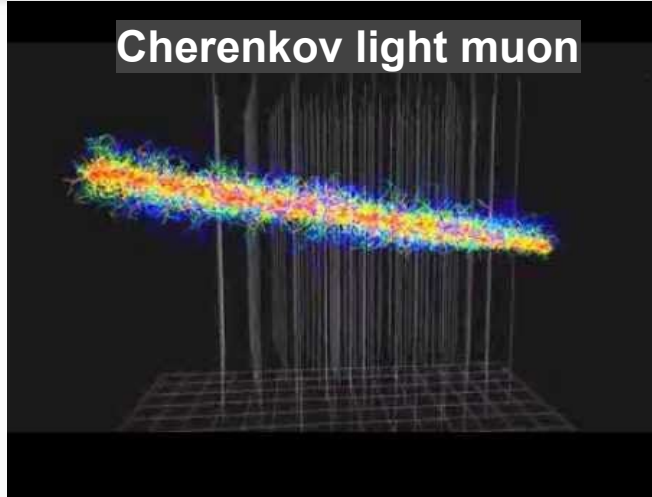
~ km-long muon tracks from  $\nu_\mu$



**Track orientation determines photon arrival PDF**

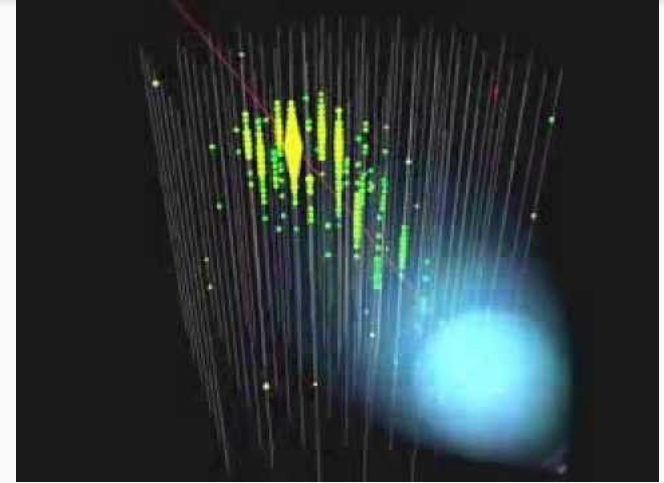


## Cherenkov light muon

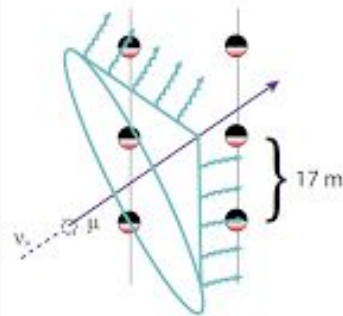


**Blob size:**  
**Number of photons**

**Blob color:**  
**Photon arrival time**



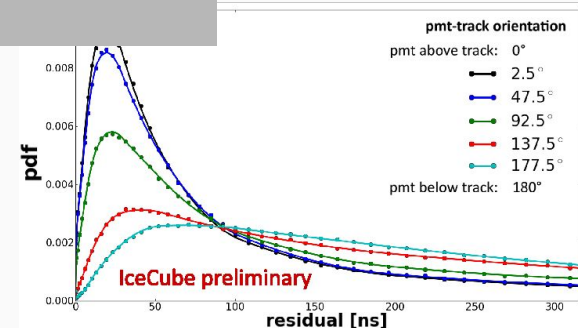
~ km-long muon tracks from  $\nu_\mu$



Use data + algorithm (typically likelihood Ansatz) to determine parameters

$$L_{\text{ext}} = \prod_j \frac{e^{-\lambda_j} \lambda_j^{q_j}}{q_j!} \prod_i \frac{e^{-N_{\text{hit}}}}{N_{\text{hit}}!} [p_j(t_i)]^{q_i},$$

**Energy, Direction, ...**



# The first point source - TXS 0506+056

[ Previous | Next | ADS ]

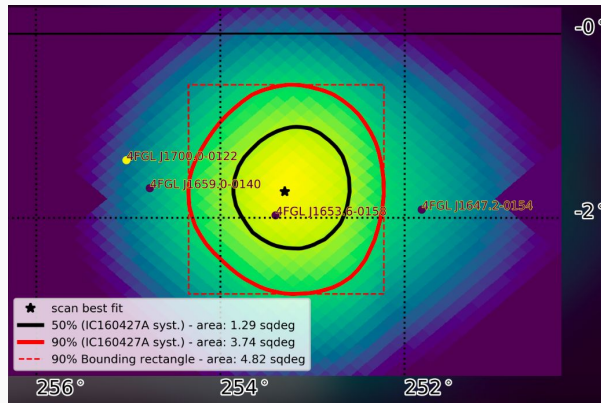
## Fermi-LAT detection of increased gamma-ray activity of TXS 0506+056, located inside the IceCube-170922A error region.

ATel #10791; *Yasuyuki T. Tanaka (Hiroshima University), Sara Buson (NASA/GSFC), Daniel Kocevski (NASA/MSFC) on behalf of the Fermi-LAT collaboration*  
on 28 Sep 2017; 10:10 UT  
Credential Certification: David J. Thompson (David.J.Thompson@nasa.gov)

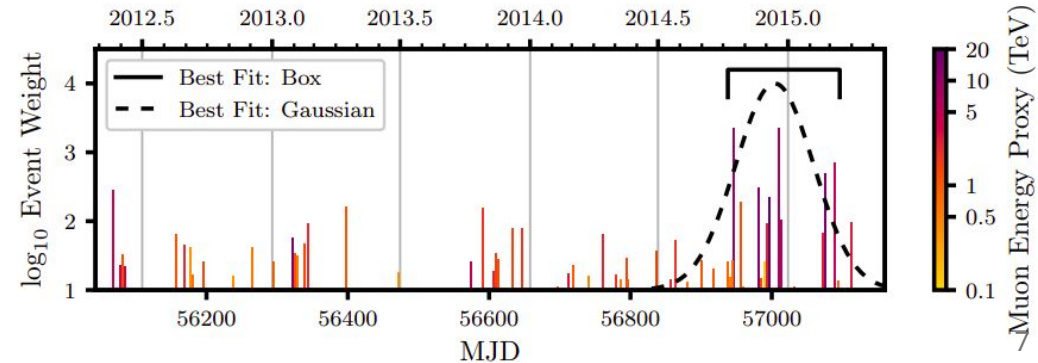
Subjects: Gamma Ray, Neutrinos, AGN



## Neutrino coincident with gamma rays (2017)



## Historical data at same loc. shows excess in 2014/2015



# The first point source - TXS 0506+056

[ Previous | Next | ADS ]

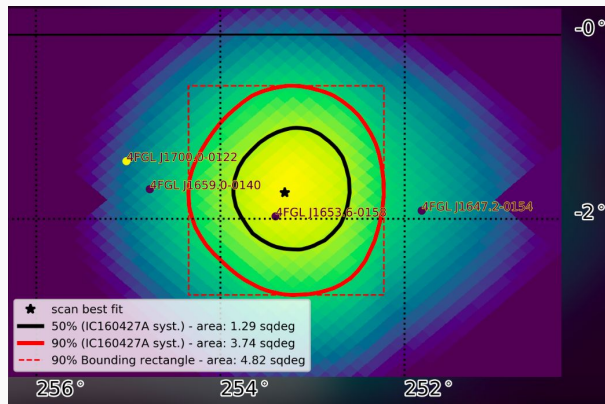
**Fermi-LAT detection of increased gamma-ray activity of TXS 0506+056, located inside the IceCube-170922A error region.**

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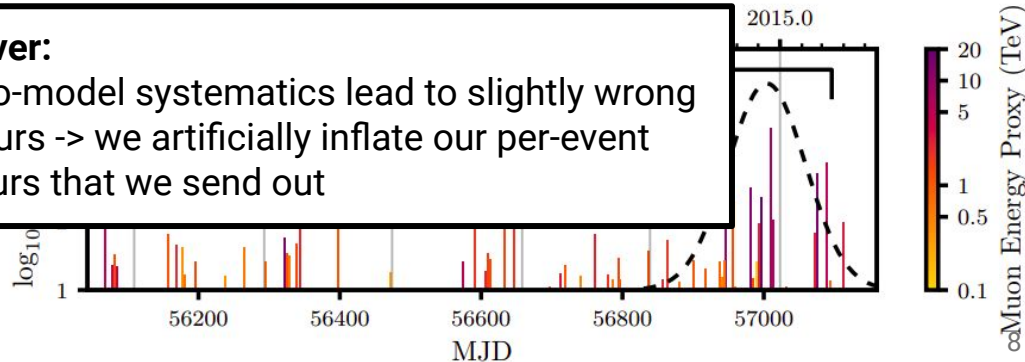
Neutrino coincident with gamma rays (2017)



Historical data at same loc. shows excess in 2014/2015

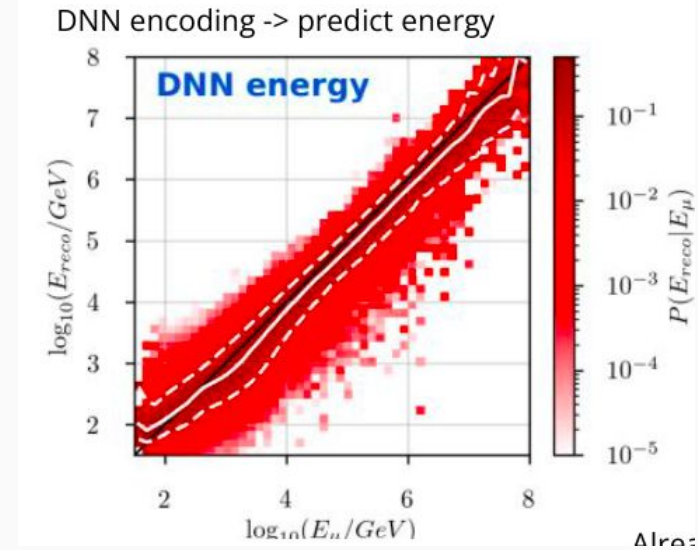
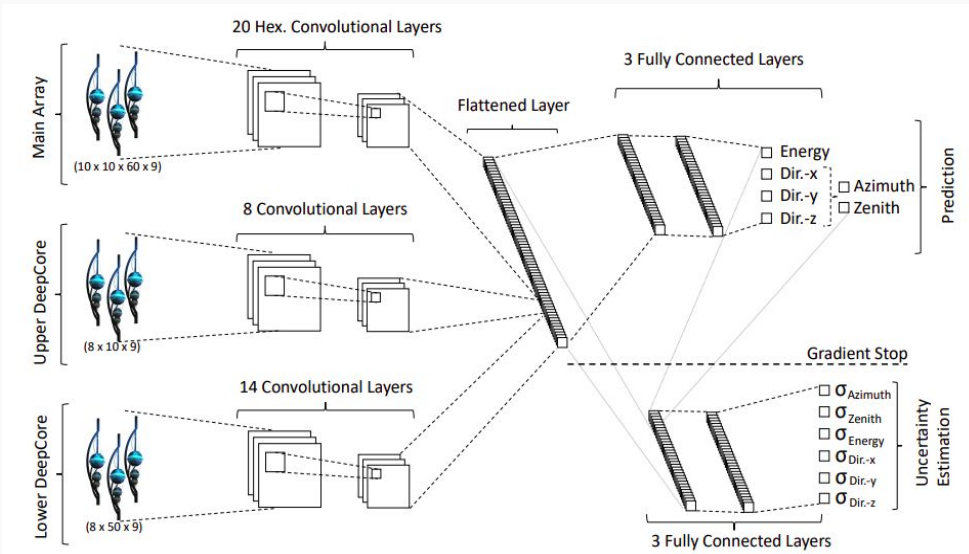
**However:**

hard-to-model systematics lead to slightly wrong Contours -> we artificially inflate our per-event contours that we send out





## Neural-network based muon energy prediction was part of processing



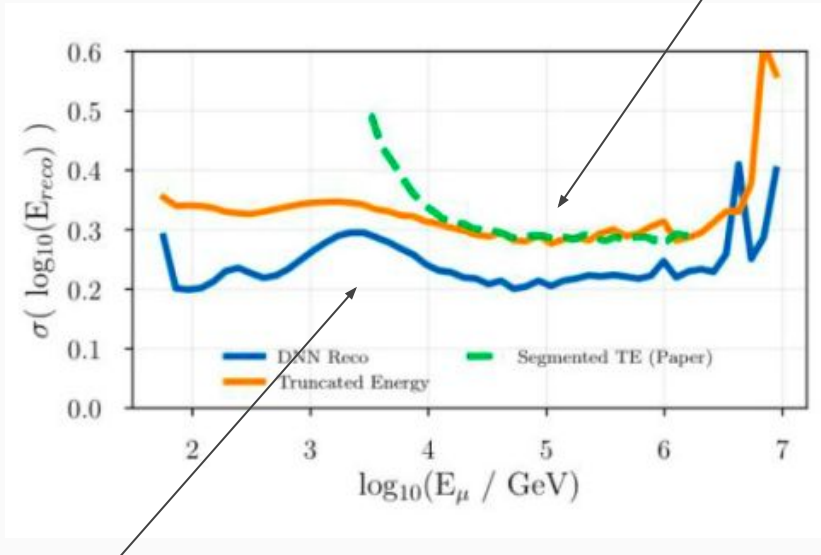
# New track processing in 2021 -> some changes in contours



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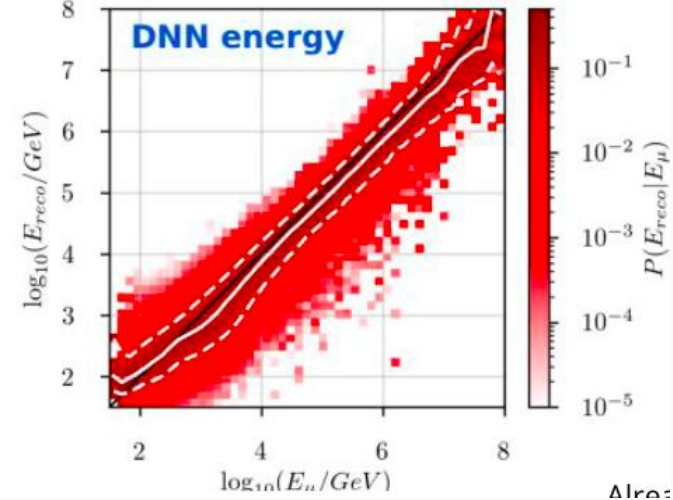
Neural-network based muon energy prediction  
was part of processing

Likelihood based



Neural network prediction

DNN encoding -> predict energy

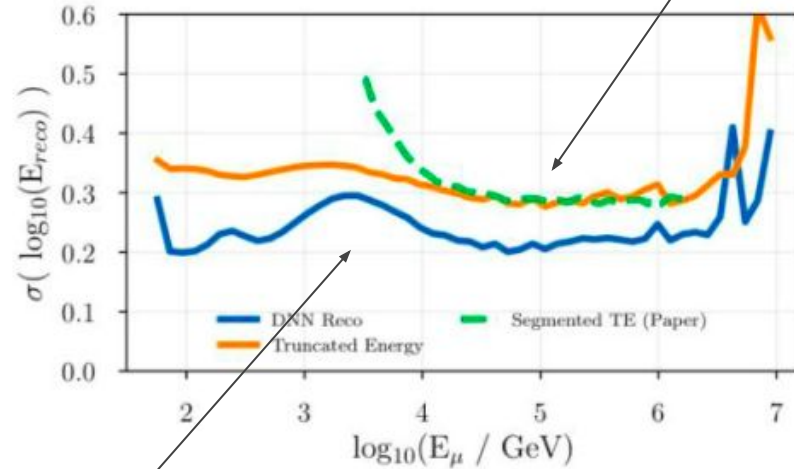


# New track processing in 2021 -> some changes in contours

Neural-network based muon energy prediction  
was part of processing

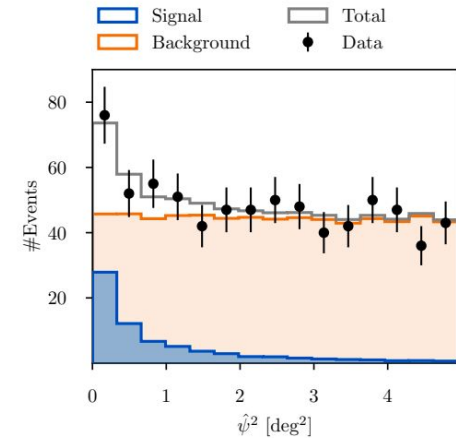
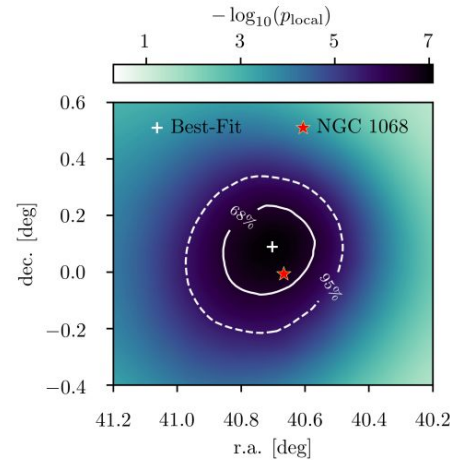
+ improved likelihood formalism

Likelihood based

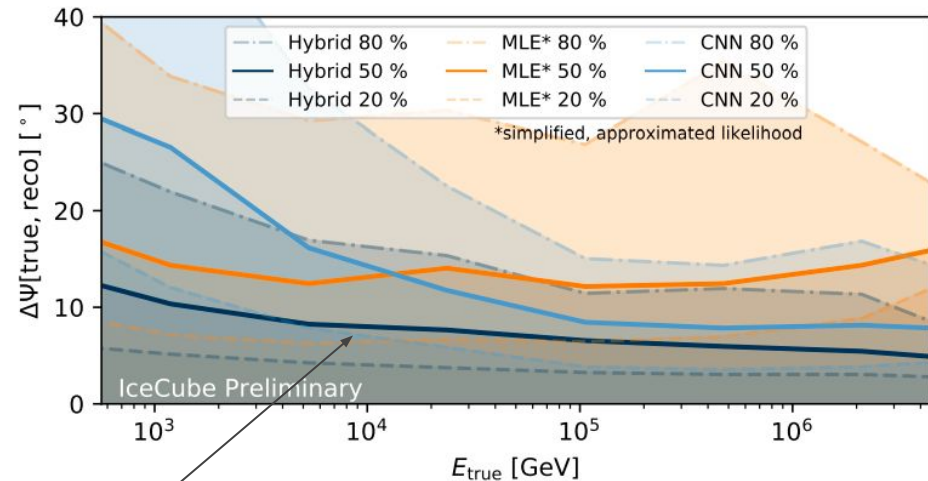
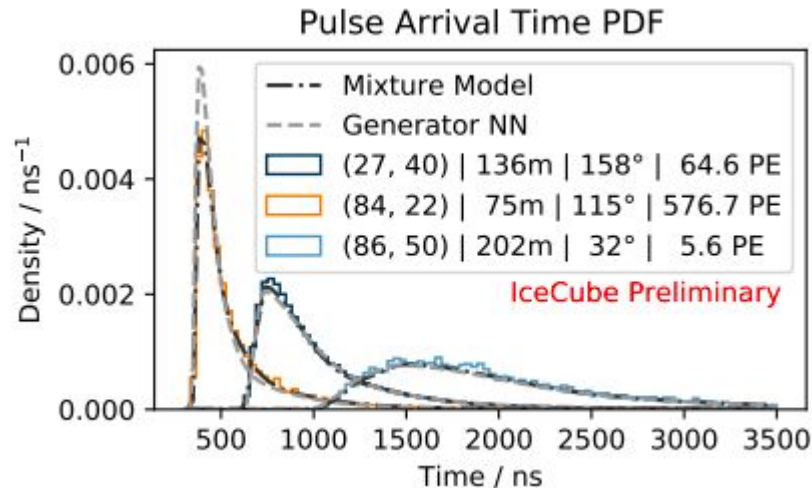


Neural network prediction

Frist time-independent point source: NGC 1068  
(Science, Nov. '22)



# Neural-network based likelihood prediction: $\nu_e$ direction reconstruction

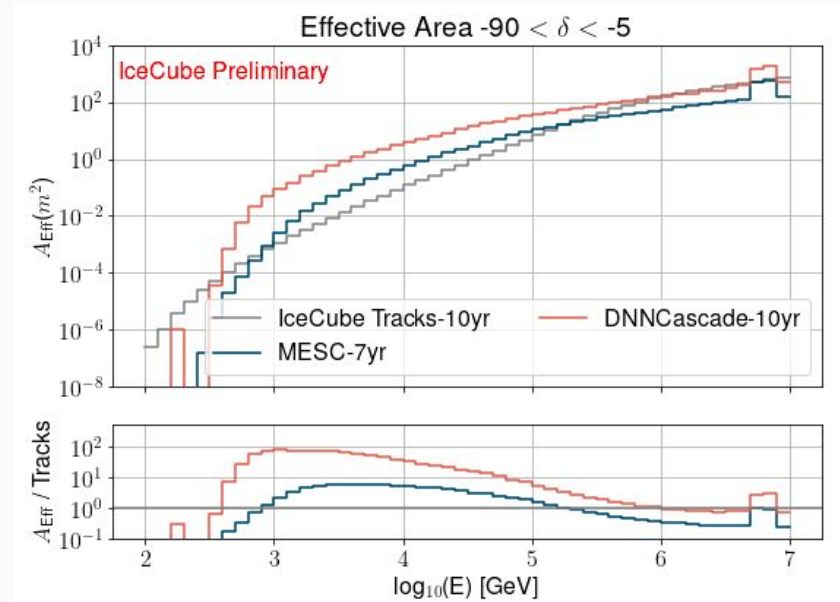
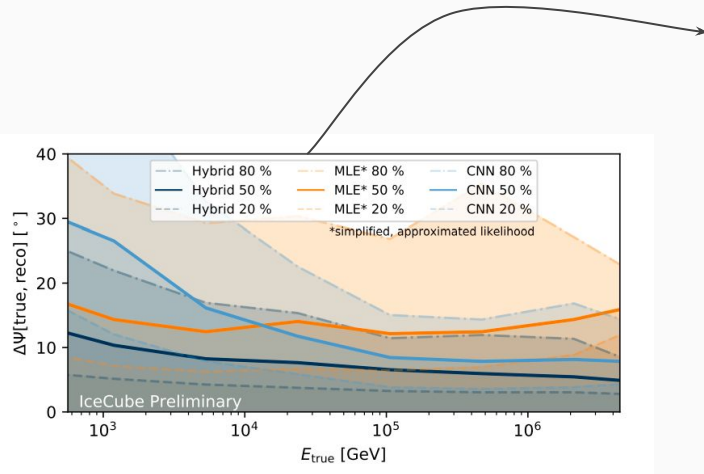


$$\mathcal{L}_{\text{event}}(\vec{x} = \{\vec{c}, \vec{t}\} | \vec{\xi}) = \prod_d^D \text{Poisson}\left(\sum_i^{N_d} c_{d,i} | \lambda_d(\vec{\xi})\right) \cdot \prod_i^{N_d} P_d(t_{d,i} | \vec{\xi})^{c_{d,i}}$$

PoS-ICRC2021-1065



Deep Learning **crucial** in a new analysis pipeline (**see press conference on June 30 for more details on result**)



Multiple deep learning  
Reconstructions in event selection  
→ **Factor X 10-100** more signal events than  
preceeding event selections at low energies

Per-Event contours often systematically wrong...  
Is there a better way?

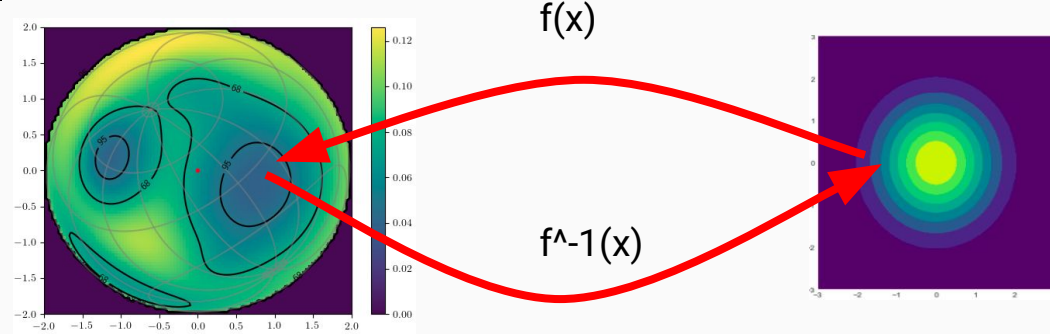
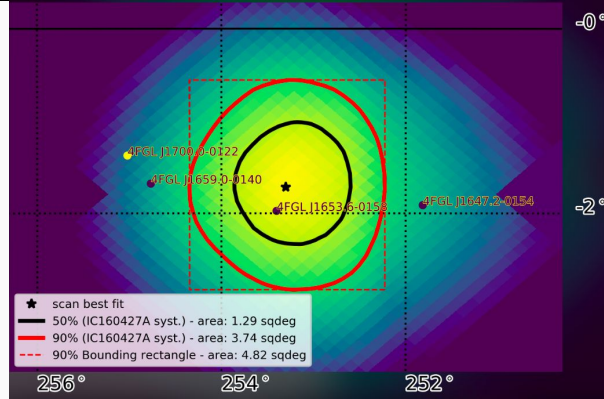
$$p_{\theta}(x|y) = p_0(f_{g_{\theta}(y)}^{-1}(x)) \cdot |\det J_{g_{\theta}(y)}^{-1}(x)|$$

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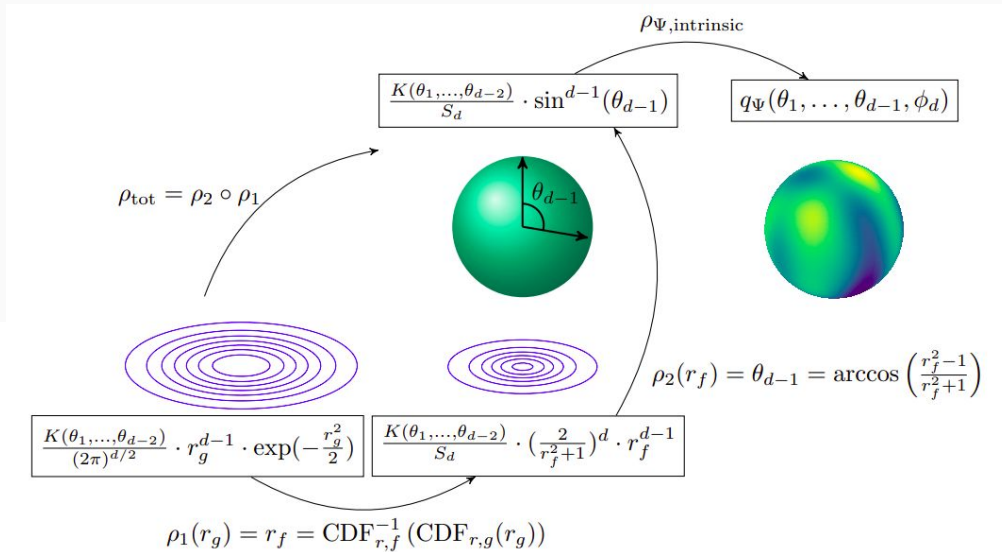
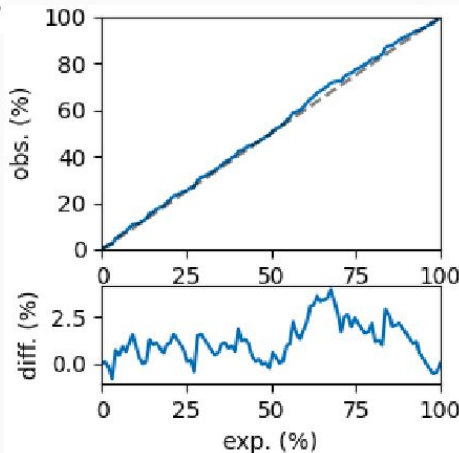
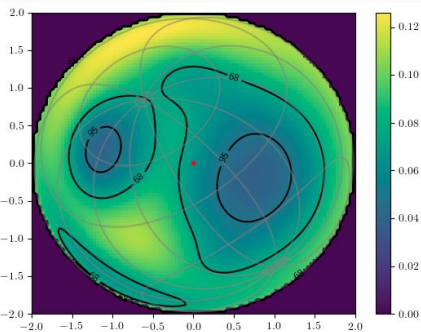
ATel #10791; *Yasuyuki T. Tanaka (Hiroshima University), Sara Buson (NASA/GSFC),  
Daniel Kocevski (NASA/MSFC) on behalf of the Fermi-LAT collaboration*  
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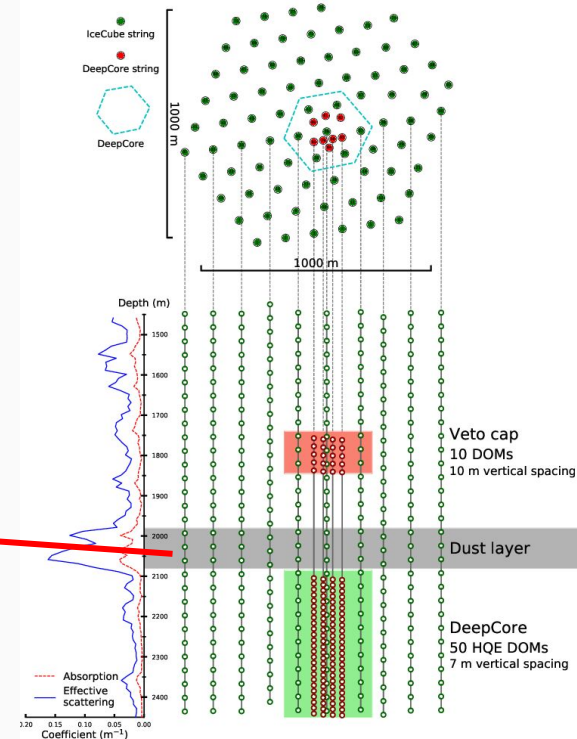
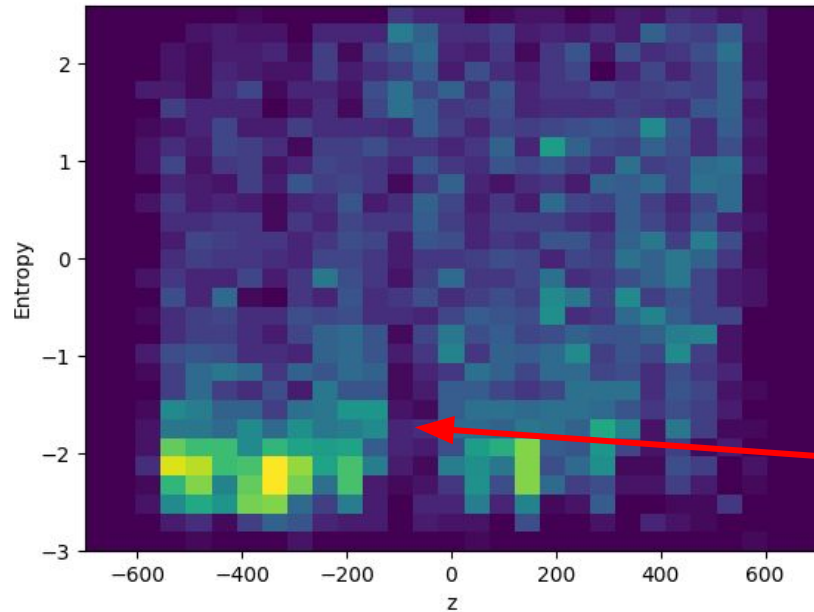


- Coverage can be checked **without numerical integration**  
(for arbitrary shaped contours, including distributions on the sphere)



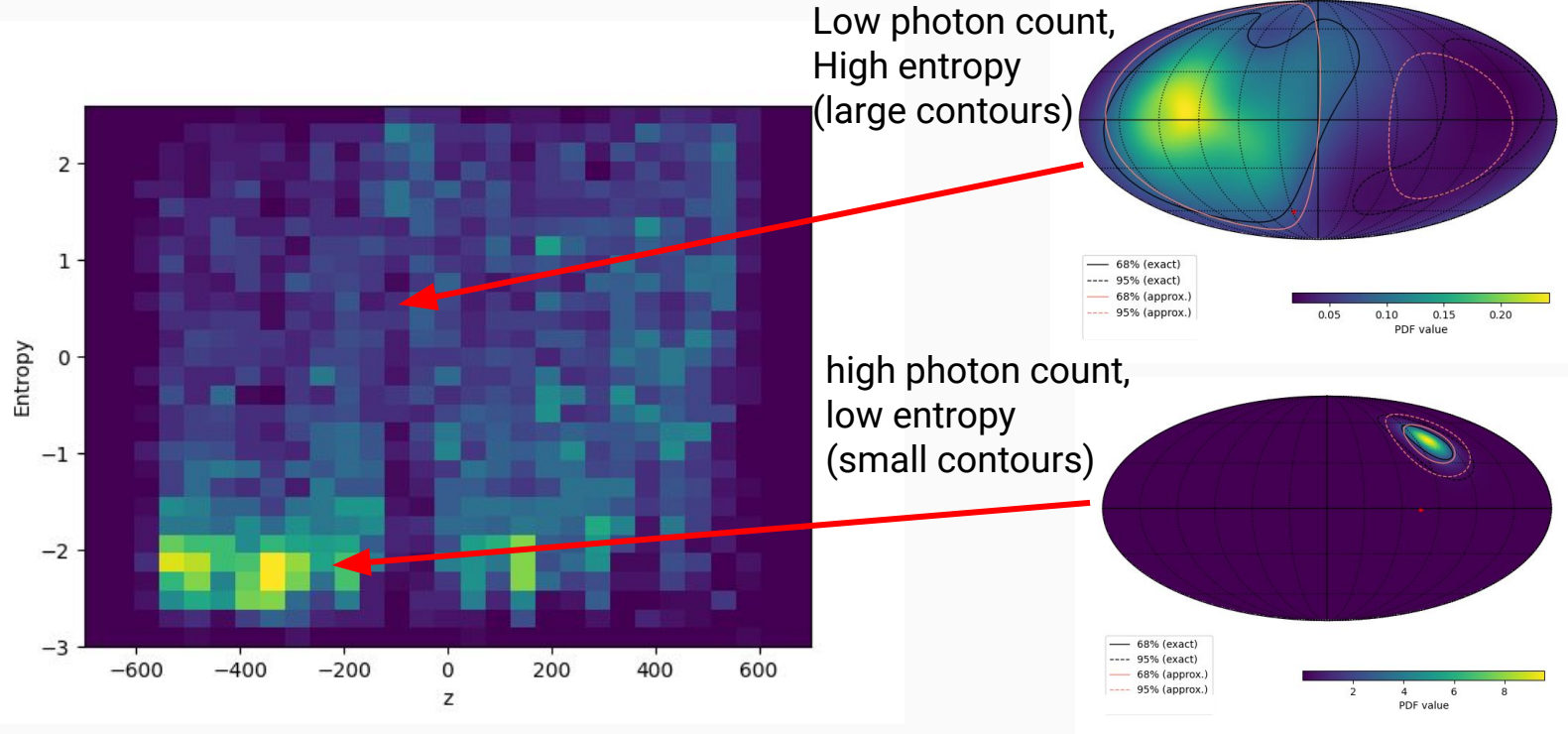
arXiv:2008.05825

- The normalizing flow **picks up ice structure** (we have dust layers in the Antarctic ice)





- The normalizing flow **picks up ice structure** (we have dust layers in the Antarctic ice)

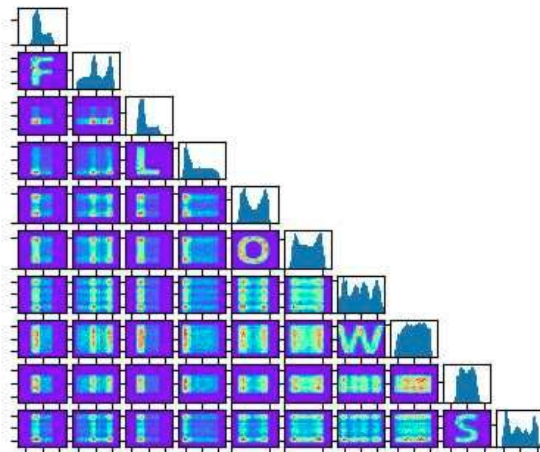
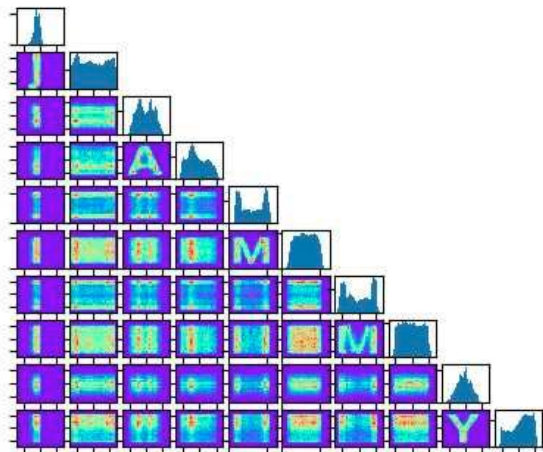


[https://github.com/thoglu/jammy\\_flows](https://github.com/thoglu/jammy_flows)

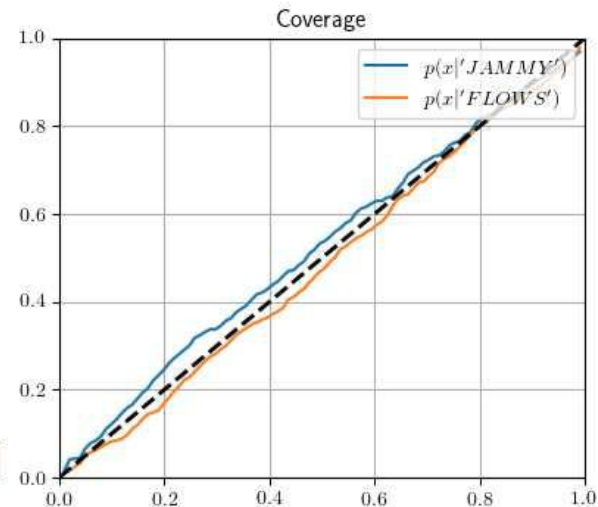
```
import jammy_flows
```

```
pdf=jammy_flows.pdf("e4+s2+e4", "gggg+n+gggg")
```

pdf structure: e4+s2+e4



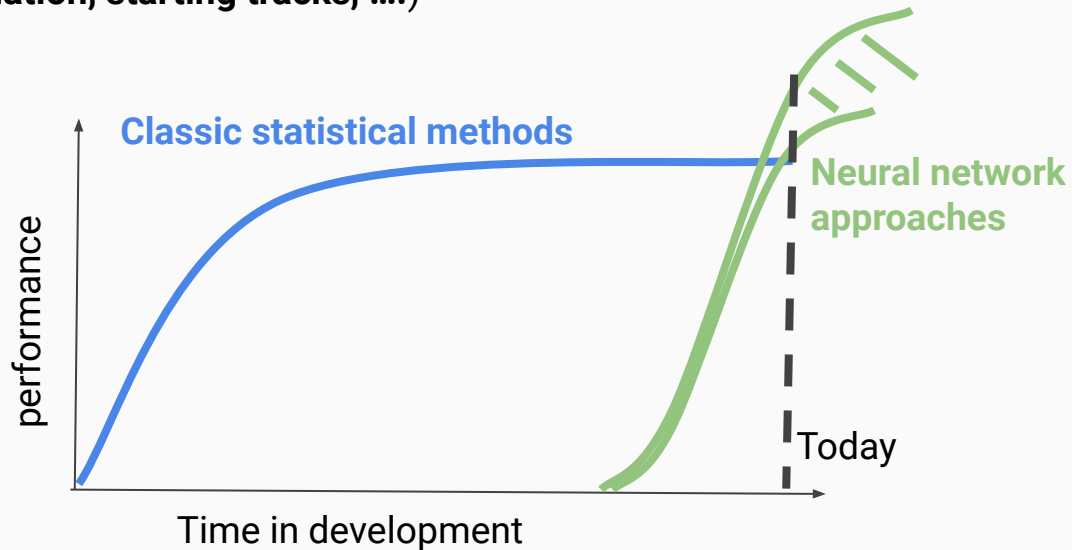
“A generic pdf on a tensor product of manifolds”



- More than 80% of technical effort in IceCube by now involves neural network and deep learning
- Already offers faster and more precise reconstructions for a large number of use cases (**cascade-type topologies, energy estimation, starting tracks, ....**)

Classical llh approaches run into  
bottleneck  
(**20+ years** of history in  
IceCube/Amanda)

New Machine learning methods  
(only **~2-3 years** of investigation)  
already better in many areas



**New physics results and improvements of existing results guaranteed in the next years !**