

The (spectral harmonic) test-field method characterizing emergent phenomena in MRI turbulence

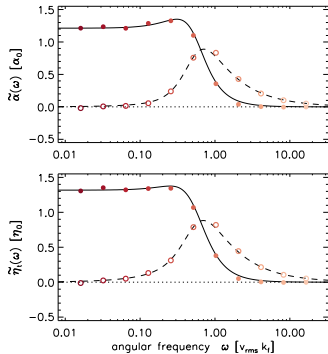
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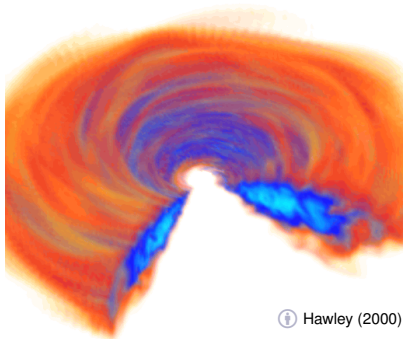
1 context / theory

- conceptual issues
- dynamo theory in a nutshell

2 methods / results

- the quasi-kinematic test-field method
- results pertaining to MRI turbulence

the engine: magnetorotational instability

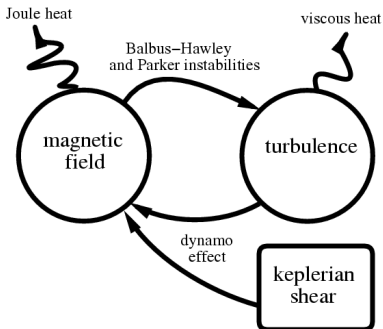


ⓘ Hawley (2000)

- accretion luminosity via turbulent viscosity
 - non-thermal coronal X-ray emission via B-fields
 - weak magnetic fields destabilise shear flow
MRI, Balbus & Hawley (1991)
 - robust linear instability – problem solved ...
-
- ... twenty-five years later, saturation mechanism remains enigmatic
 - attempts

?	linear theory	?	parasitic instabilities
?	direct simulations	?	mean-field dynamo
?	limit-cycle regime	?	transition into chaos

the key: survival of large-scale fields



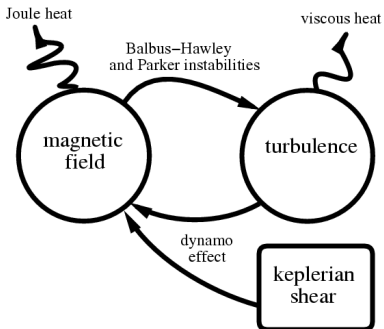
Brandenburg (1995)

- stratified shearing boxes have all ingredients for a classical strato-cyclonic dynamo
- large-scale dynamo is less likely Pm-dependent Brandenburg (2001)
- tall-enough (un-)stratified ZNF converged Davis, Stone & Pessah (2010), Shi, Stone & Huang (2016), Ryan et al. (2017), Coleman et al. (2017)
- (cyclic) dynamo already seen in unstratified ZNF case Lesur & Ogilvie (2008), Herault et al. (2013/15) Squire & Bhattacharjee (2015a/b)

complementary route:

study evolution of embedded **poloidal flux** (itself subject to turbulence)

the key: survival of large-scale fields



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■ complementary route:

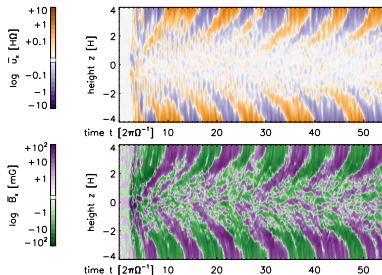
study evolution of embedded **poloidal flux** (itself subject to turbulence)

MRI dynamo vs. MHD winds

■ butterfly diagram also in the velocity

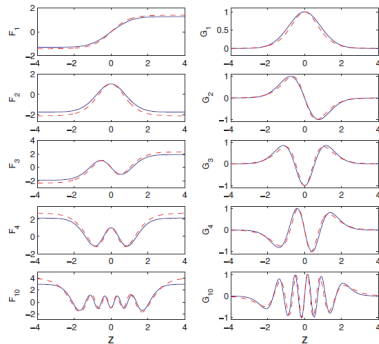
$$\mathbf{V}^{\text{ch}} = \mathbf{u}_0 F\left(\frac{z}{H}\right) e^{\text{st}} [\mathbf{e}_x \cos \theta + \mathbf{e}_y \sin \theta]$$

$$\mathbf{B}^{\text{ch}} = \mathbf{B}_0 G\left(\frac{z}{H}\right) e^{\text{st}} [\mathbf{e}_x \sin \theta - \mathbf{e}_y \cos \theta]$$




→ model 'A1' from Gressel, Nelson & Turner (2011)

strong fields → low wave-number



weak fields → high wave-number

 Latter, Fromang & Gressel (2010)

topological constraints

■ Dynamical quenching

(new notation: $a = A'$, $b = B'$, ...)

■ non-linear effects in the EMF

$$\partial_t \bar{\mathcal{E}} = \overline{u \times (\partial_t b)} + \overline{(\partial_t u) \times b}$$

$$\rightarrow \alpha = \alpha_K + \alpha_M = -\frac{1}{3} \tau_K \langle \omega \cdot u \rangle + \frac{1}{3} \tau_M \langle j \cdot b \rangle / \rho$$

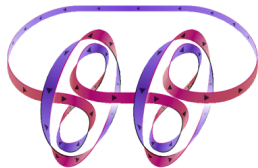
■ magnetic helicity evolution

$$\partial_t \langle \bar{A} \cdot \bar{B} \rangle = +2 \langle \bar{\mathcal{E}} \cdot \bar{B} \rangle - 2\eta \langle \bar{J} \cdot \bar{B} \rangle$$

$$\partial_t \langle a \cdot b \rangle = -2 \langle \bar{\mathcal{E}} \cdot \bar{B} \rangle - 2\eta \langle j \cdot b \rangle$$

■ time evolution for **effective** α effect

$$\partial_t \alpha = -2\eta_t k_f^2 \left(\frac{\alpha \langle \bar{B}^2 \rangle - \eta_t \langle \bar{J} \cdot \bar{B} \rangle + \text{fluxes}}{B_{\text{eq}}^2} + \frac{\alpha - \alpha_K}{\eta_t / \eta} \right)$$

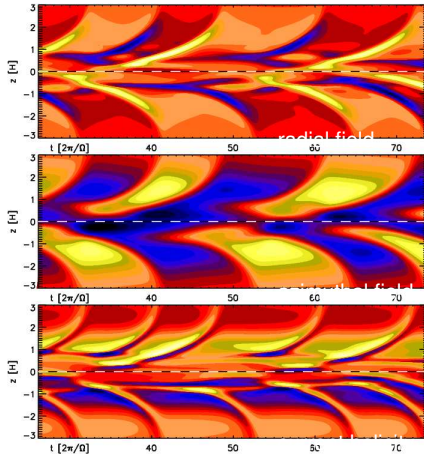


Blackman (2014)

using $\alpha_K = \text{const.}$, $\langle \bar{\mathcal{E}} \cdot \bar{B} \rangle = \langle \alpha \bar{B} \cdot \bar{B} \rangle - \langle \eta_t \bar{J} \cdot \bar{B} \rangle$ and $\langle a \cdot b \rangle \simeq k_f^{-2} \langle j \cdot b \rangle$

- ■ Blackman & Field (2000) ■ Vishniac & Cho (2001) ■ Blackman & Brandenburg (2002)
- Vishniac & Shapovalov (2014) ■ Squire & Bhattacharjee (2015a/b)

dynamically quenched mean-field model

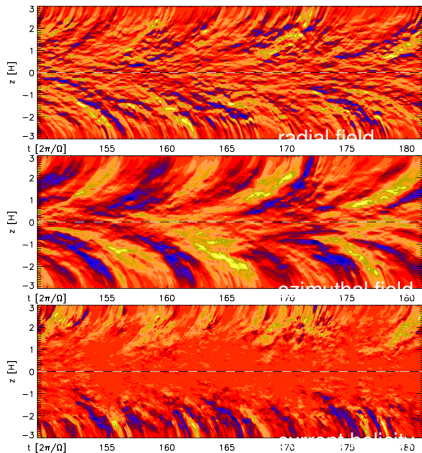


■ 1D “ α_M ” toy model / ■ direct MHD simulation

📄 Gressel (2010), MNRAS 405

- reproduces (at least decently) qualitative features:
 - asymmetry in B_R and B_ϕ
 - intermittent parity, chaotic features (Rm dependent)
 - frequency doubling in helicity (caused by phase shift)
- quantitative agreement difficult due to sensitive parameter dependencies
- note that MRI dynamo always appears quenched / dominated by fluctuations

dynamically quenched mean-field model



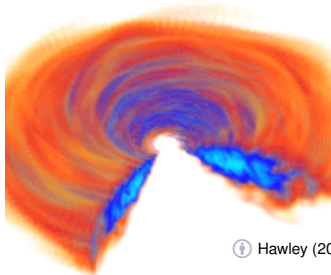
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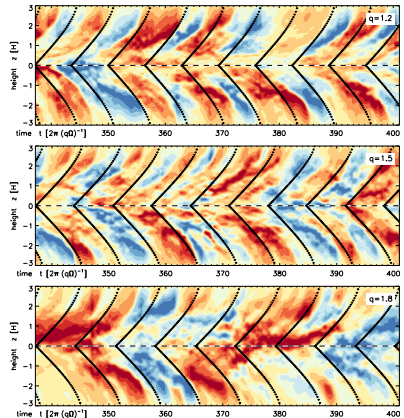


For the remainder of the talk, we will investigate the hypothesis that the coherent fields that are appearing in accretion disk turbulence are a consequence of a classical *stratocyclonic* mean-field dynamo.



 Hawley (2000)

 Gressel & Pessah (2015)





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emergent phenomena

■ The induction term:

$$\blacksquare \nabla \times (\mathbf{V} \times \mathbf{B}) \equiv -\mathbf{B}(\nabla \cdot \mathbf{V}) + (\mathbf{B} \cdot \nabla) \mathbf{V} - (\mathbf{V} \cdot \nabla) \mathbf{B}$$

■ Mean-field approach:

■ split into mean + fluctuation

$$\mathbf{V} = \bar{\mathbf{V}} + \mathbf{v}' \quad \text{and} \quad \mathbf{B} = \bar{\mathbf{B}} + \mathbf{b}'$$

■ mean-field equation with $\bar{\mathcal{E}} = \overline{\mathbf{v}' \times \mathbf{b}'}$

$$\partial_t \bar{\mathbf{B}} = \nabla \times (\bar{\mathbf{V}} \times \bar{\mathbf{B}} + \bar{\mathcal{E}} - \eta_m \nabla \times \bar{\mathbf{B}})$$

■ Parametrise small-scale effects $\bar{\mathcal{E}}$ as a functional of $\bar{\mathbf{V}}, \bar{\mathbf{B}}, \overline{f(\mathbf{v}')}$

■ typically $\bar{\mathcal{E}}_i = \alpha_{ij} \bar{B}_j - \tilde{\eta}_{ij} \epsilon_{jkl} \partial_k \bar{B}_l$



ⓘ A. Fletcher/R. Beck, SuW/HHT, STScl/AURA

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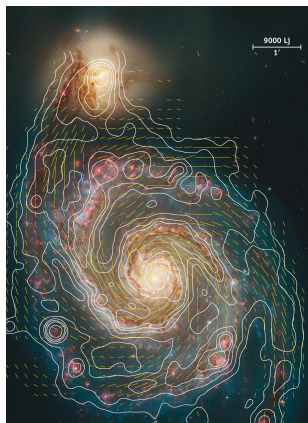
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Ⓐ A. Fletcher/R. Beck, SuW/HHT, STScl/AURA

the mean induction equation

- Splitting-up terms in the induction equation

$$\partial_t \overline{\mathbf{B} + \mathbf{b}'} = \nabla \times [(\overline{\mathbf{V}} + \mathbf{v}') \times (\overline{\mathbf{B}} + \mathbf{b}')] + \eta_m \nabla^2 \overline{\mathbf{B} + \mathbf{b}'}$$

- Using Reynold's averaging rules

- idempotence: $\overline{\overline{f}} = \overline{f}$
- symmetric perturbations: $\overline{f'c} = 0$
- summation: $\overline{f + g} = \overline{f} + \overline{g}$
- mixed product: $\overline{f' \times g'c} = \overline{f'} \times \overline{g'c} = 0$

- Leads to the mean-field induction equation

$$\partial_t \overline{\mathbf{B}} = \nabla \times (\overline{\mathbf{V}} \times \overline{\mathbf{B}}) + \nabla \times \overline{\mathcal{E}} + \eta_m \nabla^2 \overline{\mathbf{B}}$$

with turbulent **EMF** $\overline{\mathcal{E}} = \overline{\mathbf{v}' \times \mathbf{b}'}$ subsuming all small-scale effects

motivating α/η -type closures

- Aim: find an evolution equation for the mean EMF

$$\bar{\mathcal{E}}(\mathbf{z}, t) = \overline{\mathbf{v}'(\mathbf{z}, \mathbf{t}) \times \mathbf{b}'(\mathbf{z}, \mathbf{t})} = \overline{\mathbf{v}'(\mathbf{z}, \mathbf{t}) \times \int_{\tau=0}^{\mathbf{t}} \partial_{\tau} \mathbf{b}'(\mathbf{z}, \tau) d\tau} + \dots$$

- Compute magnetic field fluctuations:

$$\partial_t(\bar{\mathbf{B}} + \mathbf{b}') = \nabla \times [(\bar{\mathbf{V}} + \mathbf{v}') \times (\bar{\mathbf{B}} + \mathbf{b}')] + \eta \nabla^2 (\bar{\mathbf{B}} + \mathbf{b}')$$

$$- \partial_t \bar{\mathbf{B}} = \nabla \times [\bar{\mathbf{V}} \times \bar{\mathbf{B}} + \overline{\mathbf{v}' \times \mathbf{b}'}] + \eta \nabla^2 \bar{\mathbf{B}}$$

$$\partial_t \mathbf{b}' = \nabla \times [\bar{\mathbf{V}} \times \mathbf{b}' + \mathbf{v}' \times \mathbf{b}' - \overline{\mathbf{v}' \times \mathbf{b}'} + \mathbf{v}' \times \bar{\mathbf{B}} - \eta \nabla \times \mathbf{b}']$$

- third-order moments appear when substituted into $\bar{\mathcal{E}}(\mathbf{z}, t)$
- ad-hoc parametrisation $\bar{\mathcal{E}}_i = \alpha_{ij} \bar{B}_j + \eta_{ijk} \partial_k \bar{B}_j$



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spectral harmonic TF method

- Ansatz for test-field inhomogeneity:

$$\begin{aligned}\bar{\mathcal{B}}_{(0)} &= \cos(\omega t) \cos(k_z z) \hat{\mathbf{x}}, & \bar{\mathcal{B}}_{(1)} &= \cos(\omega t) \sin(k_z z) \hat{\mathbf{x}}, \\ \bar{\mathcal{B}}_{(2)} &= \cos(\omega t) \cos(k_z z) \hat{\mathbf{y}}, & \bar{\mathcal{B}}_{(3)} &= \cos(\omega t) \sin(k_z z) \hat{\mathbf{y}}.\end{aligned}$$

- Evolution equation for associated fluctuations:

$$\partial_t \mathcal{B}'_{(\nu)} = \nabla \times \left[\mathbf{v}' \times \bar{\mathcal{B}} + \bar{\mathbf{v}} \times \mathcal{B}'_{(\nu)} - \overline{\mathbf{v}' \times \mathcal{B}'_{(\nu)}} + \mathbf{v}' \times \mathcal{B}'_{(\nu)} - \eta_m \nabla \times \mathcal{B}'_{(\nu)} \right]$$

- Then, using $\bar{\mathcal{E}}^{(\nu)} \equiv \overline{\mathbf{v}' \times \mathcal{B}'_{(\nu)}}$, the matrix inversion is trivial:

$$\begin{pmatrix} \tilde{\alpha}_{ij}(k_z, \omega) \\ k_z \tilde{\beta}_{ijz}(k_z, \omega) \end{pmatrix} = e^{i\omega t} \begin{pmatrix} \cos(k_z z) & \sin(k_z z) \\ -\sin(k_z z) & \cos(k_z z) \end{pmatrix} \begin{pmatrix} \bar{\mathcal{E}}_i^{(2j-2)} \\ \bar{\mathcal{E}}_i^{(2j-1)} \end{pmatrix}$$

non-locality / finite “domain of dep.”

- classical local closure relation:

$$\bar{\mathcal{E}}_i(z, t) = \alpha_{ij}(z, t) \bar{B}_j(z, t) - \eta_{ij}(z, t) \varepsilon_{jzl} \partial_z \bar{B}_l(z, t).$$

- non-local generalisation:


$$\bar{\mathcal{E}}_i(z) = \int \hat{\alpha}_{ij}(z, \zeta) \bar{B}_j(z - \zeta) - \hat{\eta}_{ij}(z, \zeta) \varepsilon_{jzl} \partial_z \bar{B}_l(z - \zeta) d\zeta,$$

$$\tilde{\mathcal{E}}_i(k_z) = \tilde{\alpha}_{ij}(k_z) \tilde{B}_j(k_z) - \tilde{\eta}_{ij}(k_z) i k_z \varepsilon_{jzl} \tilde{B}_l(k_z)$$

- empiric *ad hoc* model for kernel function:

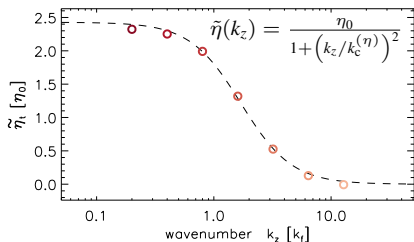
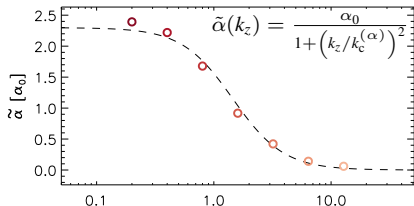
$$\tilde{\alpha}(k_z) = \frac{\alpha_0}{1 + \left(k_z/k_c^{(\alpha)}\right)^2}, \quad \tilde{\eta}(k_z) = \frac{\eta_0}{1 + \left(k_z/k_c^{(\eta)}\right)^2}$$

 Brandenburg, Rädler & Schinnerer (2008), A&A 482, 739

 Gressel & Pessah (2015), ApJ 810, 59


helical forcing (k -dependence)

- benchmark: \rightarrow helically-forced hydro
(i.e., strict kinematic limit)



$$\begin{aligned} \partial_t \rho + \nabla \cdot (\rho \mathbf{v}) &= 0, \\ \partial_t (\rho \mathbf{v}) + \nabla \cdot (\rho \mathbf{v} \mathbf{v}) &= -\nabla p + \nabla \cdot \boldsymbol{\tau} \\ &\quad + \mathbf{f}_{\text{hel}}(\mathbf{k}, t) \end{aligned}$$

- amplitudes for the α effect (top) and turbulent diffusion, η_T (bottom)
- helically-forced isothermal turbulence with $k_f = 5$, $\text{Ma} \simeq 0.1$, and $\text{Re} = \text{Rm} \simeq 10$.
- dashed lines indicate Lorentzian fits with $k_c^{(\alpha)} = 1.45$ and $k_c^{(\eta)} = 1.70$

 Gressel & Elstner (2020), MNRAS 494, 1180

 Bbg, Rädler & Schinnerer (2008), A&A 482, 739

time-lag / “memory” effects

- classical local closure relation:

$$\bar{\mathcal{E}}_i(z, t) = \alpha_{ij}(z, t) \bar{B}_j(z, t) - \eta_{ij}(z, t) \varepsilon_{jzl} \partial_z \bar{B}_l(z, t).$$

- non-instantaneous generalisation:

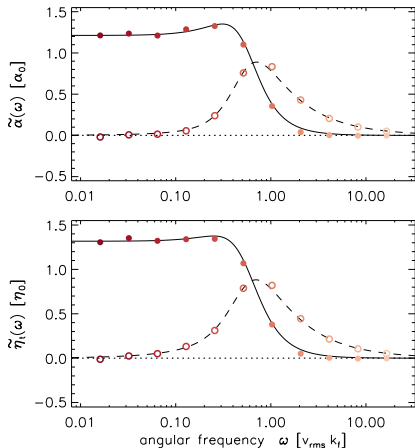
$$\bar{\mathcal{E}}_i(z, t) = \int \hat{\alpha}_{ij}(z, t') \bar{B}_j(z, t - t') - \hat{\eta}_{ij}(z, t') \varepsilon_{jzl} \partial_z \bar{B}_l(z, t - t') dt',$$

$$\tilde{\mathcal{E}}_i(\omega) = \tilde{\alpha}_{ij}(\omega) \tilde{B}_j(\omega) - \tilde{\eta}_{ij}(\omega) \mathbf{ik}_z \varepsilon_{jzl} \tilde{B}_l(\omega)$$

- empiric *ad hoc* model for kernel function:

$$\tilde{\alpha}(\omega) = \alpha_0 \frac{1 - \mathbf{i} \omega \tau_c^{(\alpha)}}{\left(1 - \mathbf{i} \omega \tau_c^{(\alpha)}\right)^2 + \left(\omega_0^{(\alpha)} \tau_c^{(\alpha)}\right)^2},$$

helical forcing (frequency dependence)





- for $k_f \simeq 3$, $Ma \simeq 0.1$, $Re = Rm \simeq 22$.
- solid and dashed lines show a simultaneous fit to the real and imaginary part, respectively

$$\Re = \frac{1 + (\omega^2 + \omega_0^2) \tau_c^2}{4 \omega^2 \tau_c^2 + (1 - (\omega^2 - \omega_0^2) \tau_c^2)^2}$$

$$\Im = \frac{1 + (\omega^2 - \omega_0^2) \tau_c^2}{4 \omega^2 \tau_c^2 + (1 - (\omega^2 - \omega_0^2) \tau_c^2)^2} \omega \tau_c$$

- slight tension with previous results...
(but also different codes, Pencil vs. NIRVANA-III)

 Gressel & Elstner (2020), MNRAS 494, 1180

 Hubbard & Brandenburg (2009), ApJ 706, 712



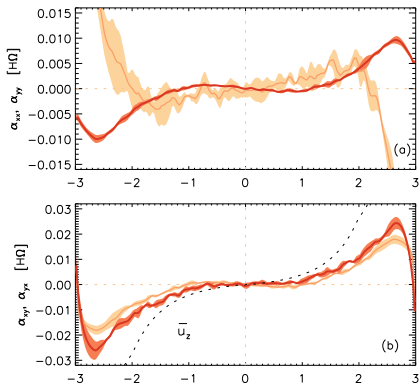
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test-field α effect for MRI turbulence

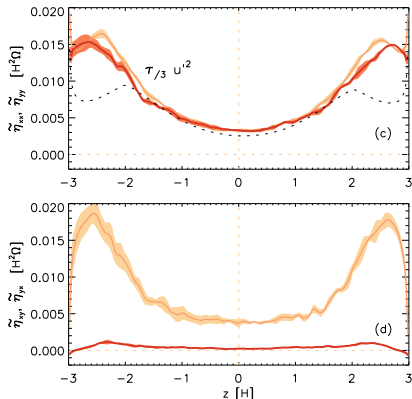


 Gressel & Pessah (2015), ApJ 810, 59

- new test-field results for weaker shear of $q = 1.2$
- pronounced negative α effect near midplane
Brandenburg (1998),
Rüdiger & Pipin (2000)
- as previously: off-diagonal tensor elements both positive (anisotropic pumping)

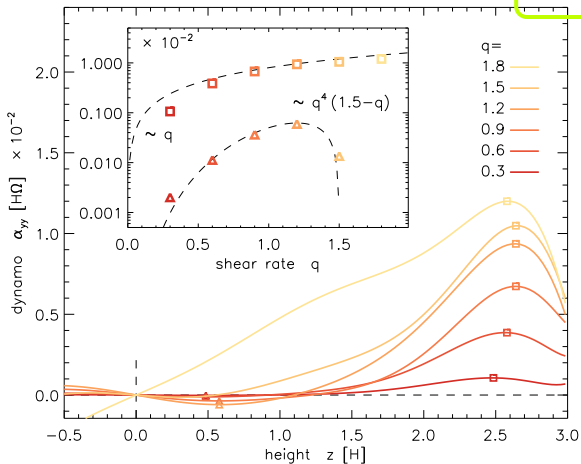
test-field turbulent η for MRI turbulence


- turbulent diffusion consistent with theory (i.e., for $z < 2H$ / high β_P)
- off-diagonals both positive (i.e., no independent Rädler effect–dynamo)
- weak $\tilde{\eta}_{yx}$ responsible for butterfly diagram?!





Gressel & Pessah (2015), ApJ 810, 59

shear-rate dependence of coefficients

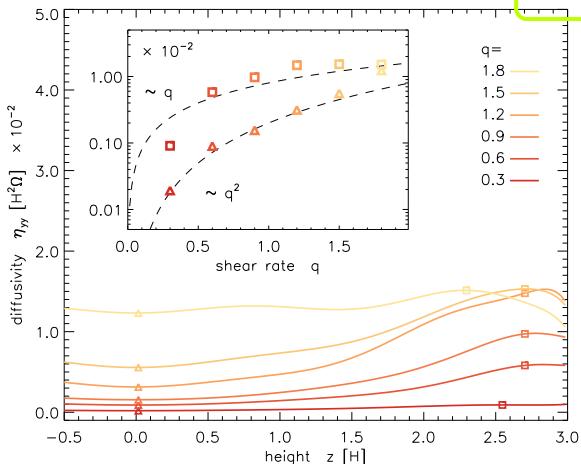



 Rüdiger & Pipin (2000)


 Ziegler & Rüdiger (2001)


 Gressel & Pessah (2015), ApJ 810, 59

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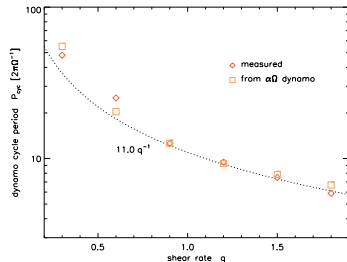
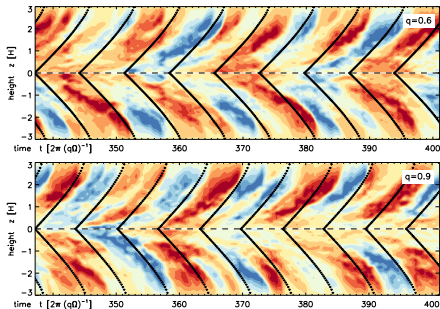


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 Gressel & Pessah (2015), ApJ 810, 59

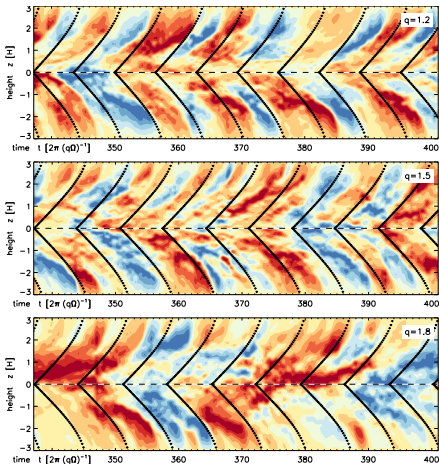
the dynamo cycle period



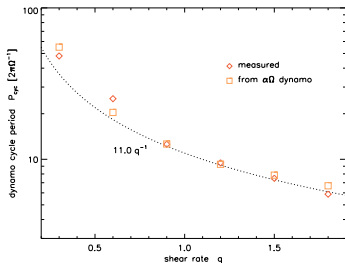
- $\omega_{\text{cyc}} \simeq \left| \frac{1}{2} \alpha_{yy} q \Omega k_z \right|^{1/2}$
- shear-rate dependence predicted well by $\alpha\Omega$ dispersion relation
- propagation direction still “wrong”

 Gresell & Pessah (2015), ApJ 810, 59

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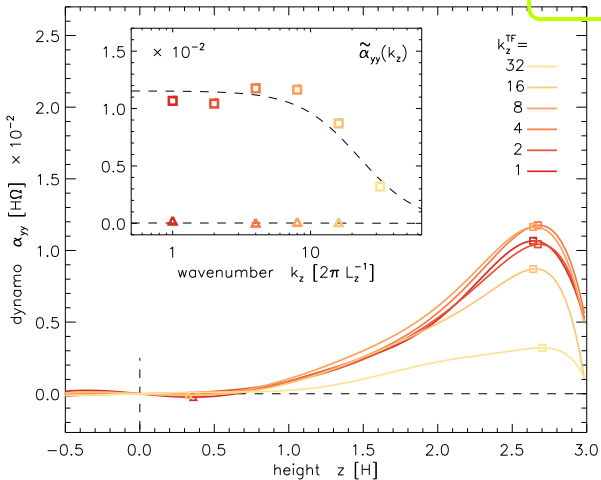



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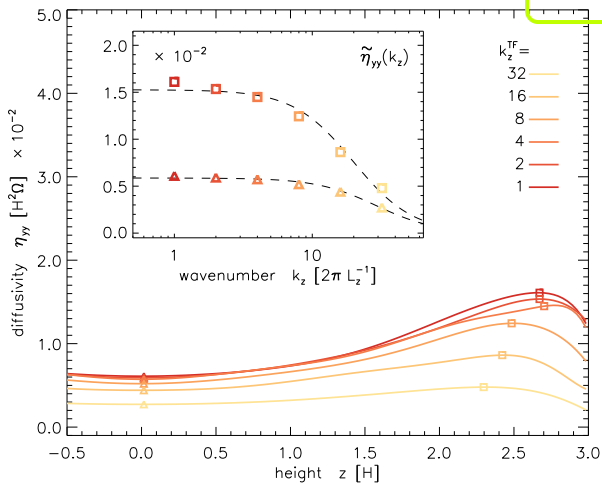
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scale-dependence of coefficients




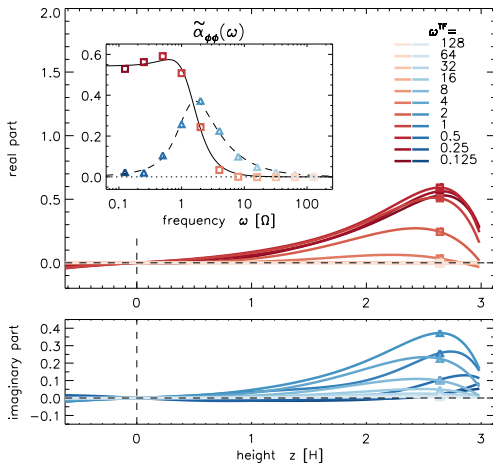
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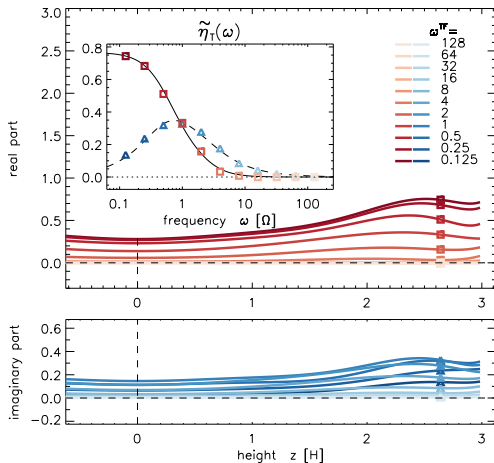
 Gresel & Pessah (2015), ApJ 810, 59

time-lag-dependence of coefficients

dynamo $\alpha_{\phi\phi} [\text{H}\Omega \times 10^{-2}]$  Gressel & Pessah (2022), ApJ 928, 118

time-lag-dependence of coefficients

diffusivity η_T [$H^2\Omega \times 10^{-2}$]



Gressel & Pessah (2022), ApJ 928, 118

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Thank you for your attention.