



# Progress in Flat Space Holography

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**From Amplitudes to Gravitational Waves - Nordita, July 2023**

# Motivation

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## Holography (most basic test)

- Asymptotic/large gauge symmetries of bulk (gravity) theory  $\leftrightarrow$  global symmetries of boundary (quantum) theory

[Brown, Henneaux '86]

## Gravity in 4D Asymptotically Flat Spacetimes

- Lorentz symmetries act like global conformal symmetries of the 2D sphere at infinity superrotations  $\leftrightarrow$  Virasoro

[Barnich, Troessaert '09; Kapec, Lysov, Pasterski, Strominger '14]

- Subleading soft graviton mode  $\leftrightarrow$  generator of Virasoro!

[Cachazo, Strominger '14; Kapec, Mitra, A.R., Strominger '16]

$$\lim_{\omega \rightarrow 0} \text{[diagram: shaded circle with four lines and a wavy line labeled } \omega \hat{q}] = \left[ \frac{1}{\omega} S^{(0)}(\hat{q}; p_i) + S^{(1)}(\hat{q}; p_i, J_i) + \mathcal{O}(\omega) \right] \times \text{[diagram: shaded circle with four lines]}$$

**Goal:** Look for 2D conformal field theory dual to gravity in 4D AFS:  $\Lambda = 0$  holography

# Outline



Celestial amplitudes from  
AdS Witten diagrams



Top-down celestial holography

Conformally soft sector of  $CFT_3$

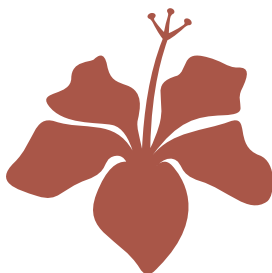
Soft gluon and graviton algebras



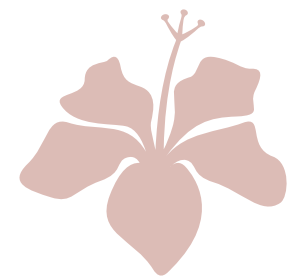
Operator product expansions  
and conformally soft symmetries



Generic features of celestial  
2, 3, 4-point functions



Massless scattering  
The boost basis and celestial amplitudes



# Spinor helicity variables

**Massless** Poincare representations: particle states  $|p_\mu, \sigma\rangle$  labelled by momentum  $p_\mu$  and helicity  $\sigma$

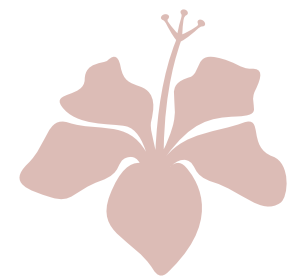
Lorentz group  $SO^+(1,3) \simeq SL(2,\mathbb{C})/\mathbb{Z}_2$  :  $p_\mu \rightarrow P \equiv p_\mu \sigma^\mu$ ,  $\sigma_\mu$  Pauli matrices

$$\Lambda_{\mu\nu} p^\nu \rightarrow XPX^\dagger, \quad X = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \in SL(2,\mathbb{C})$$

Spinor helicity variables:  $p^2 = 0 \iff \det P = 0 \implies P_{\alpha\dot{\alpha}} = \lambda_\alpha \tilde{\lambda}_{\dot{\alpha}}$

Let  $\lambda = \sqrt{2\omega} \begin{pmatrix} z \\ 1 \end{pmatrix}$ ,  $\tilde{\lambda} = \eta \lambda^*$  with  $\eta = \pm$  depending on whether particle is incoming or outgoing. The Lorentz group then acts as:

$$\lambda \rightarrow \lambda' = e^{i\theta} \sqrt{2\omega'} \begin{pmatrix} z' \\ 1 \end{pmatrix} \quad \text{where} \quad \omega' = |cz + d|^2 \omega, \quad z' = \frac{az + b}{cz + d}, \quad e^{2i\theta} = \frac{cz + d}{\bar{c}\bar{z} + \bar{d}}$$



# Massless scattering amplitudes

Lorentz transformations:  $\lambda \rightarrow \lambda' = e^{i\theta} \sqrt{2\omega'} \begin{pmatrix} z' \\ 1 \end{pmatrix}$  where  $\omega' = |cz + d|^2 \omega$ ,  $z' = \frac{az + b}{cz + d}$ ,  $e^{2i\theta} = \frac{cz + d}{\bar{c}\bar{z} + \bar{d}}$

• Little group = Lorentz subgroup that preserves momenta:  $\lambda' = e^{i\theta} \lambda \implies p' = p$

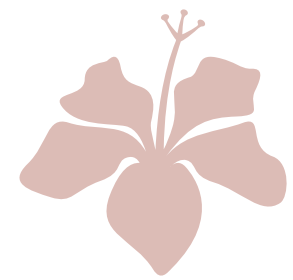
• Under Lorentz transformations:  $X(\Lambda) |p, \sigma\rangle = \sum_{\sigma'} D_{\sigma\sigma'} | \Lambda p, \sigma'\rangle$ , where  $D$  is some **representation** of the little group

in 4D little group is  $SO(2) \rightarrow$  reps. labelled by helicity  $\sigma$

Scattering amplitudes of  $n$  massless particles labelled by  $n$  pairs  $(p_a, \sigma_a)$ :

• Translation invariance  $\implies \mathcal{M}(p_a, \sigma_a) = M(p_a, \sigma_a) \delta^4 \left( \sum_{a=1}^n p_a \right)$

• Lorentz covariance  $\implies M^\Lambda(p_a, \sigma_a) = \prod_a D_{\sigma_a \sigma'_a} M((\Lambda p)_a; \sigma'_a)$



# Massless scattering amplitudes

- Little group = Lorentz subgroup that preserves momenta:  $\lambda' = e^{i\theta}\lambda \implies p' = p$

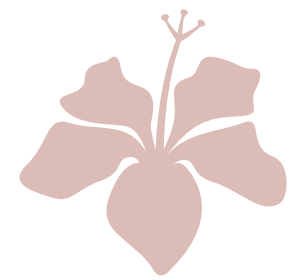
In terms of the **spinor-helicity variables**, the (stripped) amplitudes  $M$  are **homogeneous** functions of  $(\lambda_a, \tilde{\lambda}_a)$  with helicity-weights given by the little group:

$$M(e^{i\theta}\lambda, e^{-i\theta}\tilde{\lambda}) = e^{-2i\theta h} M(\lambda, \tilde{\lambda})$$

- This strongly restricts the scattering amplitudes. For example, for massless 3-point scattering:

$$\text{Momentum conservation + Lorentz invariance} \implies \begin{cases} \lambda_1 \sim \lambda_2 \sim \lambda_3 \implies M \sim [12]^a [23]^b [31]^c & [ij] \equiv \tilde{\lambda}_{i,\dot{\alpha}} \tilde{\lambda}_j^{\dot{\alpha}} \\ \tilde{\lambda}_1 \sim \tilde{\lambda}_2 \sim \tilde{\lambda}_3 \implies M \sim \langle 12 \rangle^a \langle 23 \rangle^b \langle 31 \rangle^c & \langle ij \rangle \equiv \lambda_{i,\alpha} \lambda_j^\alpha \end{cases}$$

$a, b, c$  fixed in terms of helicities by little group scaling!



# Boost eigenstates

Lorentz transformations:  $\lambda \rightarrow \lambda' = e^{i\theta} \sqrt{2\omega'} \begin{pmatrix} z' \\ 1 \end{pmatrix}$  where  $\omega' = |cz + d|^2 \omega$ ,  $z' = \frac{az + b}{cz + d}$ ,  $e^{2i\theta} = \frac{cz + d}{\bar{c}\bar{z} + \bar{d}}$

Lorentz group  $\sim$  global conformal group in 2D

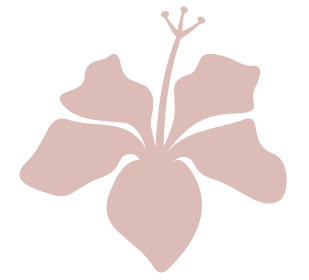
## Apply 2D CFT methods to scattering amplitudes in 4D?

**Obstacle:** asymptotic states are not in highest weight representations of  $SL(2, \mathbb{C})$

Under  $p(\omega, z, \bar{z}) \rightarrow \lambda p(\omega, z, \bar{z})$  (4D boosts towards  $(z, \bar{z}) = 2D$  dilatations),  $|\omega, z, \bar{z}\rangle \rightarrow |\lambda\omega, z, \bar{z}\rangle \neq \lambda |\omega, z, \bar{z}\rangle$  (since eg.  $[K, P] \neq 0$ )

**Diagonalize boosts:**  $|\Delta, z, \bar{z}\rangle \equiv \int_0^\infty d\omega \omega^{\Delta-1} |\omega, z, \bar{z}\rangle \rightarrow \int_0^\infty d\omega \omega^{\Delta-1} |\lambda\omega, z, \bar{z}\rangle = \lambda^{-\Delta} |\Delta, z, \bar{z}\rangle$

# Celestial amplitudes



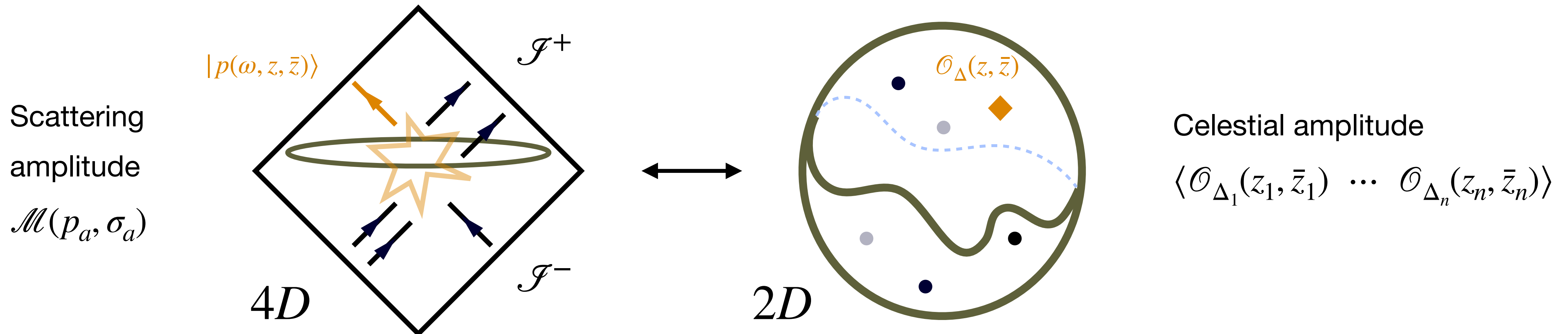
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Scatter boost eigenstates  $|\Delta, z, \bar{z}\rangle$  instead of energy-momentum-eigenstates  $|\omega, z, \bar{z}\rangle$

**Conformal primary basis:**

$$\mathcal{M}(p_a, \sigma_a) \longrightarrow \widetilde{\mathcal{M}}(\Delta_a, z_a, \bar{z}_a) = \prod_{i=1}^n \left( \int_0^\infty d\omega_i \omega_i^{\Delta_i-1} \right) M(p_a, \sigma_a) \delta^4 \left( \sum_{a=1}^n p_a \right)$$







# Examples of celestial amplitudes

Two-point functions:  $\widetilde{\mathcal{M}}(z_1, \bar{z}_1; z_2, \bar{z}_2) \propto \delta(\Delta_1 + \Delta_2 - 2) \delta^{(2)}(z_1 - z_2)$  [Pasterski, Shao, Strominger '17; Pasterski, Shao '17; Stieberger, Taylor '18]

Three-point functions:  $\widetilde{\mathcal{M}} \propto z_{21}^{-h_1-h_2+h_3} z_{23}^{-h_2-h_3+h_1} z_{13}^{h_2-h_1-h_3} \delta(\bar{z}_{12}) \delta(\bar{z}_{23}) \int_0^\infty d\omega \omega^{\Delta_1+\Delta_2+\Delta_3-3-s_1-s_2-s_3-2}$ ,  $h = \frac{\Delta + s}{2}$

Four-point functions:  $\widetilde{\mathcal{M}}(z_i, \bar{z}_i; \beta) = \underbrace{K(z_i, \bar{z}_i)}_{\text{kinematics}} X(z, \beta) \underbrace{\int_0^\infty d\omega \omega^{\beta-1} \mathcal{M}(\omega^2, -z\omega^2)}_{\equiv \mathcal{A}(\beta, z), \text{ dynamics}}$ ,  $\beta \equiv \sum_{i=1}^4 \Delta_i - 4$ ,  $X \propto \delta(z - \bar{z})$   
[momentum conservation]



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[momentum conservation]

**Analytic properties** Poor UV behavior  $\mathcal{M} \propto \omega^p \implies \widetilde{\mathcal{M}} \propto \int_0^\infty d\omega \omega^{\beta+p-1} \propto \delta(\beta + p), \beta + p \in i\mathbb{R}$

Good UV behavior  $\mathcal{M} = \lambda \frac{M^2}{\omega^2 - M^2} \implies \widetilde{\mathcal{M}} \propto \frac{\lambda M^\beta}{\sin \pi\beta/2}$  "low-" and "high-energy" poles in  $\beta$



# Properties of celestial amplitudes

Two-point functions:  $\widetilde{\mathcal{M}}(z_1, \bar{z}_1; z_2, \bar{z}_2) \propto \delta(\Delta_1 + \Delta_2 - 2) \delta^{(2)}(z_1 - z_2)$

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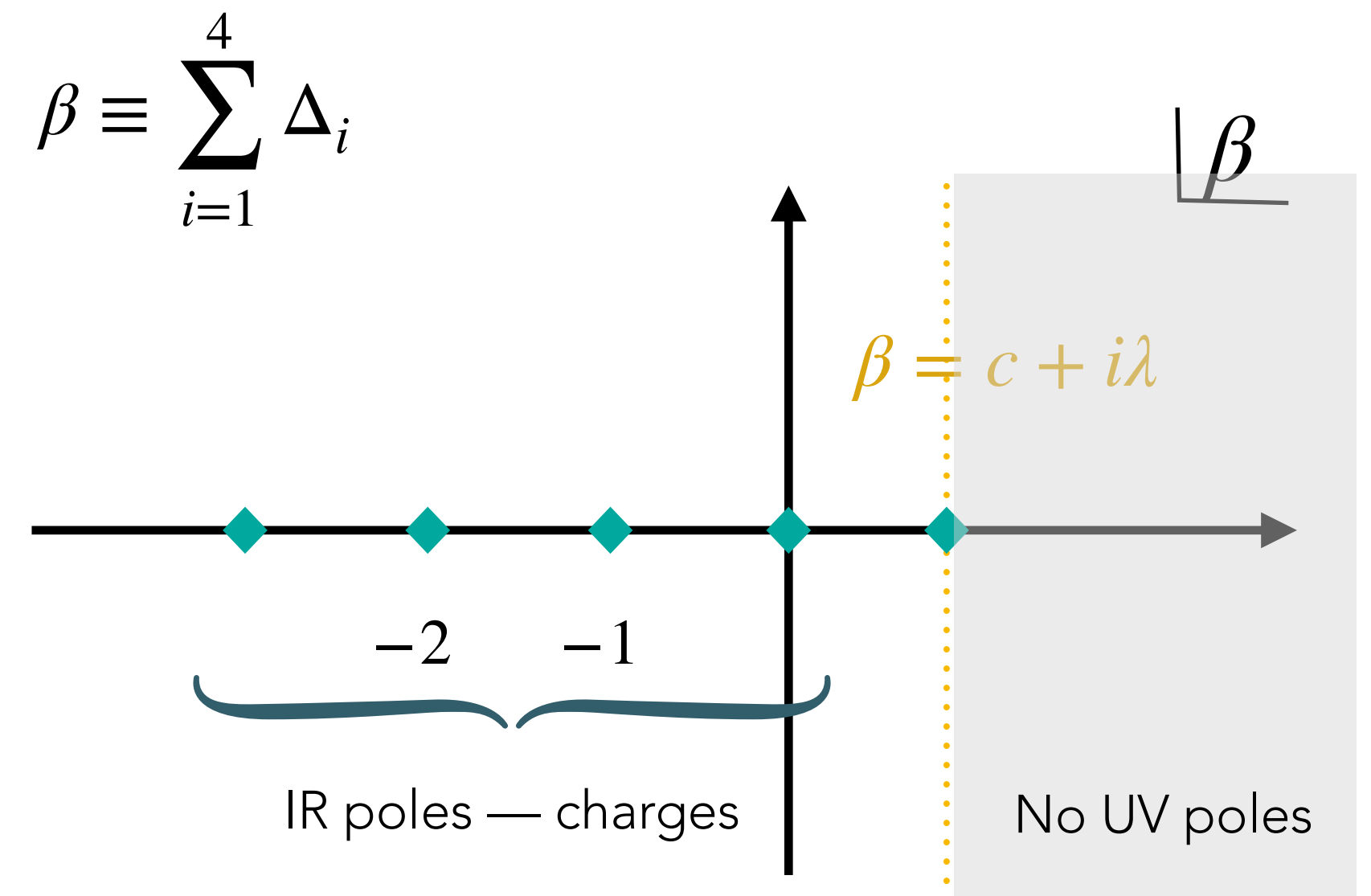
**Analytic properties**

Very good UV behavior

$$\lim_{\omega \rightarrow \infty} \mathcal{M}(\omega^2, -z\omega^2) \propto e^{-\omega^2/M^2}$$



$$\lim_{\beta \rightarrow \infty} \widetilde{\mathcal{M}}(\beta, z) \rightarrow \frac{M^\beta}{2} \Gamma(\beta/2)$$





# Other examples

**Loop corrections**  $\widetilde{\mathcal{M}}(\beta, z) \supset \int_0^{\omega_*} d\omega \omega^{\beta-1} \log^r \omega \propto \frac{\partial^r}{\partial \beta^r} \frac{1}{\beta} \propto \frac{1}{\beta^{r+1}}$  (loop order  $\sim$  higher order pole in  $\beta$ )

All loop formula in planar  $\mathcal{N} = 4$  SYM: UV finite, IR divergences exponentiate

$$M = \exp \left[ \sum_{\ell=1}^{\infty} a^{\ell} \left( f_{\epsilon}^{(\ell)} M_{\ell\epsilon}^{(1)} + C_{\epsilon}^{(\ell)} + E_{\epsilon}^{(\ell)} \right) \right] M_{\text{tree}} \longrightarrow \widetilde{\mathcal{M}} = \exp \left[ \sum_{L=1}^{\infty} a^L \left( f_{\epsilon}^{(L)} \mathcal{F}_1(z, L\epsilon) + C_{\epsilon}^{(L)} + \mathcal{E}^{(L)}(z, L\epsilon) \right) \hat{P}^{L\epsilon} \right] \widetilde{M}_{\text{tree}}$$

$$a = \frac{g^2 N}{8\pi^2} (4\pi e^{-\gamma_E})^{\epsilon}$$

- $\hat{P}$  is a conformally invariant operator  $\propto e^{\frac{i}{2} \sum_{i=1}^4 \partial_{\Delta_i}}$
- $z$  is a conformally invariant cross-ratio [Bern, Dixon, Smirnov '05]  
[Gonzales, Puhm, Rojas '20]
- Recent work celestial open string amplitudes [above + Donnay, Giribet '23]

- IR divergences in QED and gravity may be removed in a basis of eigenstates of large gauge charge

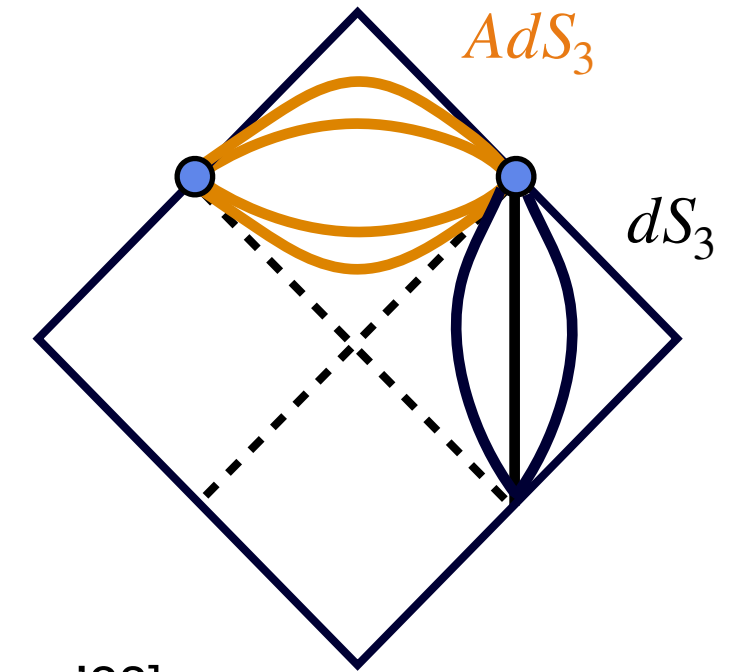


# Non-perturbative backgrounds

Translation breaking backgrounds appear to smoothen out singularities

$$\widetilde{\mathcal{M}}_B(1^-, 2^-, 3^+, \dots, n^+) \sim \frac{z_{12}^3}{z_{23}z_{34}\dots z_{n1}} \int \widetilde{d^3Q} g(Q) \int d\omega_1 \omega_1^{\Delta_1} \int d\omega_2 \omega_2^{\Delta_2} \int \prod_{j=3}^n d\omega_j \omega_j^{\Delta_j-2} \delta^{(4)}\left(Q + \sum_i \eta_i \omega_i \hat{q}_i\right)$$

conformal primary massive scalar ( $\Delta = 2$ )



- 3-point function  $\propto$  standard CFT 3-point function [Casali, Melton, Strominger '22; Stieberger, Taylor, Zhu '22; Sleight, Taronna '23]
- Celestial two-point functions in various different backgrounds recently computed to leading order in the coupling [Gonzo, McLaughlin, Puhm '22]

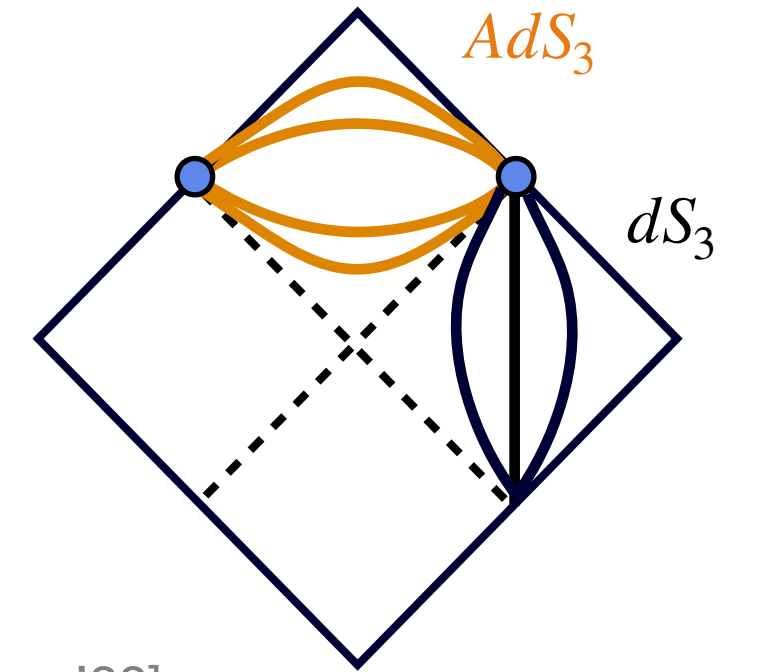


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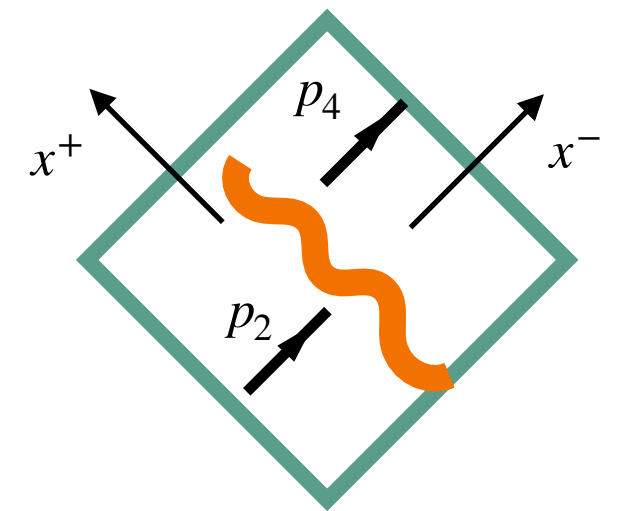
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[Gonzo, McLaughlin, Puhm '22]



- Celestial 2-point function

$$\widetilde{A}_{\text{shock}}(\Delta_2, z_2, \bar{z}_2; \Delta_4, z_4, \bar{z}_4) = 4\pi \int d^2x_{\perp} \frac{i^{\Delta_2 + \Delta_4} \Gamma(\Delta_2 + \Delta_4)}{[-q_{24,\perp} \cdot x_{\perp} - h(x_{\perp}) + i\epsilon]^{\Delta_2 + \Delta_4}}$$

[de Gioia, A.R. '22]



# Chiral algebras

- 2D QFT with  $SL(2, \mathbb{C})$  symmetry generated by 
$$\left\{ \begin{array}{l} L_{-1} = -\partial_z, \quad L_0 = -z\partial_z, \quad L_1 = -z^2\partial_z \\ \bar{L}_{-1} = -\partial_{\bar{z}}, \quad \bar{L}_0 = -\bar{z}\partial_{\bar{z}}, \quad \bar{L}_1 = -\bar{z}^2\partial_{\bar{z}} \end{array} \right\}$$

$$sl(2)_L \times sl(2)_R \quad \text{commutation relations} \quad [L_m, L_n] = (m-n)L_{m+n}, \quad [\bar{L}_m, \bar{L}_n] = (m-n)\bar{L}_{m+n}$$

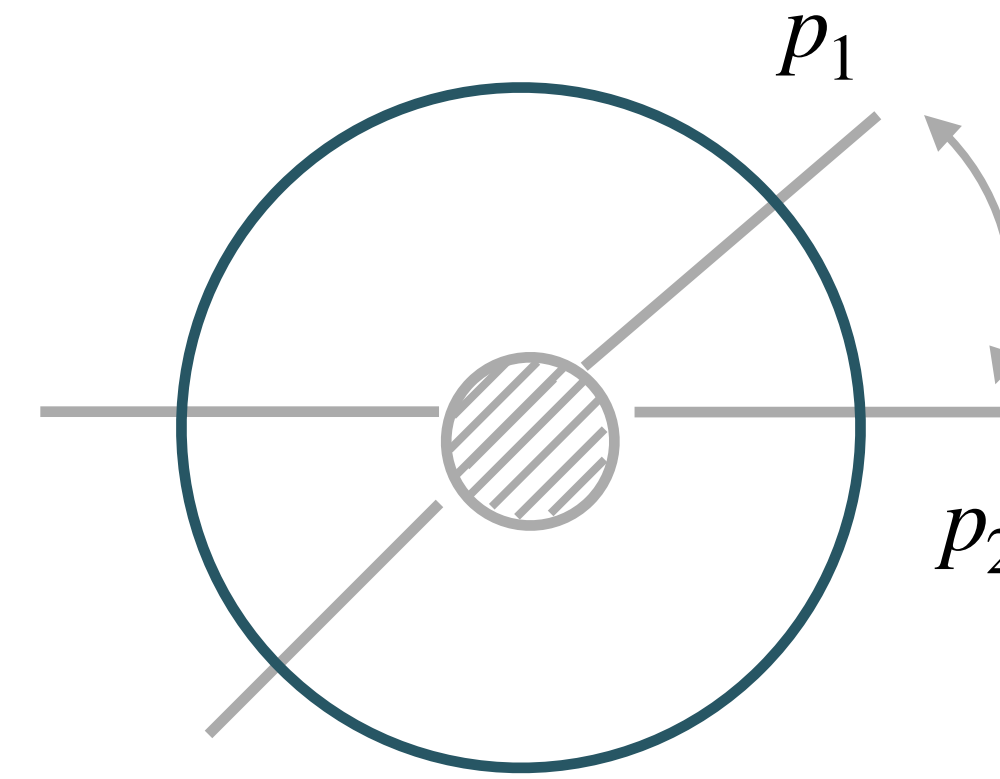
- Meromorphicity condition  $\bar{\partial}\mathcal{O}_\Delta^s(z, \bar{z}) = 0 \implies \mathcal{O}_\Delta^s(z, \bar{z}) = \mathcal{O}_h(z)$  of dimension/weight  $\Delta = h = s \in \mathbb{N}/2$

$$\implies \text{infinity of conserved charges} \quad O_n \equiv \oint dz z^{n+h-1} \mathcal{O}(z) \quad \mathcal{O}_h(z) = \sum_n \frac{O_n}{z^{n+h}}$$

- Global subalgebra = modes that annihilate the vacuum at both 0 and  $\infty \implies 1-h \leq n \leq h-1$



# Celestial operator products



Leading OPE in CCFT = collinear factorization in 4D

## Gluons in Yang-Mills theory:

Global conformal invariance in (Lorentzian) 2D CCFT  $\implies \mathcal{O}_{\Delta_1}^a(z_1, \bar{z}_1) \mathcal{O}_{\Delta_2}^b(z_2, \bar{z}_2) \sim if_c^{ab} \frac{C(\Delta_1, \Delta_2)}{z_{12}} \mathcal{O}_{\Delta_1+\Delta_2-1}^c(z_2, \bar{z}_2) + \dots$

Subleading soft gluon theorem  $\sim$  symmetry action in 2D CCFT

- Invariance of OPE under these transformations  $\implies C(\Delta_1, \Delta_1) = B(\Delta_1 - 1, \Delta_2 - 1)$ ,  $B(x, y) = \int_0^1 dt t^{x-1} (1-t)^{y-1}$
- Also follows from associativity or Poincare symmetry upon including  $SL(2, \mathbb{R})$  descendants





# Conformally soft gluon algebras

- Conformally soft gluons of **positive helicity** are operators with  $s = 1$ ,  $\Delta = k \in \mathbb{Z}$ ,  $k \leq 1 \iff h = \frac{k+1}{2}$ ,  $\bar{h} = \frac{k-1}{2}$

(Similar story can be told for negative helicity gluons  $s = -1$ )

- Note:  $\bar{h} \leq 0 \implies$  finite dimensional  $sl(2)_R$  representations:  $[\bar{L}_1, \bar{\partial}^m \mathcal{O}(z, \bar{z})] = m(2\bar{h} + m - 1)\bar{\partial}^{m-1} \mathcal{O}(z, \bar{z})$

$$\bar{\partial}^{2-k} \mathcal{O}(z, \bar{z}) = 0 \implies \mathcal{O}_k^a(z, \bar{z}) = \sum_n \frac{O_{k,n}^a(z)}{\bar{z}^{n+\frac{k-1}{2}}}, \quad \frac{k-1}{2} \leq n \leq \frac{1-k}{2}$$

- Similar to global symmetry algebras with respect to  $sl(2)_R$  upon taking **light transform!**



# Conformally soft gluon algebras

$$\mathcal{O}_k^a(z, \bar{z}) = \sum_n \frac{O_{k,n}^a(z)}{\bar{z}^{n+\frac{k-1}{2}}}, \quad \frac{k-1}{2} \leq n \leq \frac{1-k}{2}$$

$$k \leq 1, \quad k \in \mathbb{Z}$$

$$s \equiv 1 - k \geq 0$$

• **Light transform**  $L[\mathcal{O}_{h,\bar{h}}](z, \bar{z}) \equiv \int_{\mathbb{R}} \frac{d\bar{w}}{2\pi i} \frac{\mathcal{O}_{h,\bar{h}}(z, \bar{w})}{(\bar{z} - \bar{w})^{2-2\bar{h}}} \quad (h, \bar{h}) \rightarrow (h, 1 - \bar{h})$

• Take  $\mathcal{O}_{h,\bar{h}} = \mathcal{O}_k^a(z, \bar{z}) \implies L[\mathcal{O}_k^a](z, \bar{z}) = \sum_n O_{k,n}^a(z) \int \frac{d\bar{w}}{2\pi i} \frac{1}{\bar{w}^{n+\frac{k-1}{2}}} \frac{1}{(\bar{z} - \bar{w})^{2-(k-1)}}$

$$\propto \bar{z}^{\frac{k-1}{2}-n-1} \text{ upon changing variables } \bar{w} \rightarrow \bar{w}\bar{z}$$

$\implies$  The light-transformed operator has  $\bar{h}_L = \frac{3-k}{2}$  and  $1 - \bar{h}_L \leq n \leq \bar{h}_L - 1$  cf. global modes of  $sl(2)_R$  chiral algebra



# Celestial symmetry algebras

- The conformally soft gluon modes (as well as their light-transforms) form an algebra

$$k \leq 1, \quad k \in \mathbb{Z}$$

$$s \equiv 1 - k \geq 0$$

## Strategy to compute it

$$\begin{aligned} \mathcal{O}_{\Delta_1}^a(z_1, \bar{z}_1) \mathcal{O}_{\Delta_2}^b(z_2, \bar{z}_2) &\sim if_c^{ab} \frac{B(\Delta_1 - 1, \Delta_2 - 1)}{z_{12}} \mathcal{O}_{\Delta_1 + \Delta_2 - 1}^c(z_2, \bar{z}_2) + \dots \\ &\sim if_c^{ab} \frac{1}{z_{12}} \sum_{n=0}^{\infty} B(\Delta_1 - 1 + n, \Delta_2 - 1) \frac{\bar{z}_{12}^n}{n!} \bar{\partial}^n \mathcal{O}_{\Delta_1 + \Delta_2 - 1}^c(z_j, \bar{z}_j) + O(z_{ij}^0) \end{aligned}$$

Include  $sl(2)_R$  descendants



# Celestial symmetry algebras

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Include  $sl(2)_R$  descendants

Key observation: only a finite number of terms have poles at  $\Delta_1 = 1 - s, s \in \mathbb{Z}_+$

$$R_s^a(z_1, \bar{z}_1) \mathcal{O}_{\Delta_2}^b(z_2, \bar{z}_2) \sim if_c^{ab} \frac{1}{z_{12}} \sum_{n=0}^s \binom{s + 1 - \Delta_2 - n}{s - n} \frac{\bar{z}_{12}^n}{n!} \bar{\partial}^n \mathcal{O}_{\Delta_2 - s}^c(z_2, \bar{z}_2) + O(z_{12}^0)$$

residue of conformal primary gluon at  $\Delta_i = 1 - s$



# Celestial symmetry algebras

- The conformally soft gluon modes (as well as their light-transforms) form an algebra

Taking  $\mathcal{O}_{\Delta_2}^b$  soft too (ie. take residue at  $\Delta_2 = 1 - s$ )

$$R_s^a(z_1, \bar{z}_1) R_{s'}^b(z_2, \bar{z}_2) \sim if_c^{ab} \frac{1}{z_{12}} \sum_{n=0}^s \binom{s+s'-n}{s'} \frac{\bar{z}_{12}^n}{n!} \bar{\partial}^n R_{s+s'}^c(z_2, \bar{z}_2) + O(z_{12}^0)$$

The  $sl(2)_R$  modes  $R_n^{s,a}(z) \equiv \oint \frac{d\bar{z}}{2\pi i} \bar{z}^{n+\frac{k-3}{2}} R_s^a(z, \bar{z})$  form an algebra:  $[R_n^{s,a}, R_m^{s',b}] = C(s, n; s', m) if_c^{ab} R_{m+n}^{s+s',c}$

Much simpler algebra of modes of the (right) light transforms of  $R_s^a(z)$ :  $[S_m^{s,a}, S_n^{s',b}] = if_c^{ab} S_{m+n}^{s+s',c}$

$S_s^a$  has weights  $(h, \bar{h}) = \left( \frac{2-s}{2}, \frac{2+s}{2} \right)$

[Guevara, Himwich, Pate, Strominger '21; Jiang '21; Guevara '21]



# Celestial symmetry algebras

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Light transforms of (conformally soft gluons)  $R_s^a(z)$ :  $[S_m^{s,a}, S_n^{s',b}] = if_c^{ab} S_{m+n}^{s+s',c}$

Similar story in gravity: light transforms of (conformally) soft gravitons form a w-infinity algebra

$$[w_m^s, w_n^{s'}] = (m(s+1) - n(s'+1)) w_{m+n}^{s+s'}$$

**Spacetime interpretation:** algebra also follow from hierarchy of differential equations extracted from the Einstein equations

$$\dot{\mathcal{Q}}_s = D\mathcal{Q}_{s-1} + \frac{(1+s)}{2} C\mathcal{Q}_{s-2}, \quad s \in \mathbb{Z}_+$$

tower of celestial Ward identities  $\longleftrightarrow$  tower of soft theorems

$\mathcal{Q}_s$  encode multipole moments of the gravitational field



# Chiral algebras from higher dimensions

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[Beem, Lemos, Liendo, Peelaers, Rastelli, van Rees '13]

- Chiral algebras are special, rigid structures; theories that possess them are vastly constrained
- Unexpected to encounter beyond 2D CFT
- Nevertheless they may appear as “sectors” of higher-dimensional SCFT, eg.  $4D \mathcal{N} = 2$  SYM!

Consider  $\mathbb{R}^2 \subset \mathbb{R}^4$  preserving  $sl(2)_L \times sl(2)_R \subset so(6)$  & look for operators that transform trivially under an  $sl(2)$  copy

Naive obstruction: Trivial under  $sl(2) \implies$  trivial under full  $so(6)$

Bypass by looking for  $\widehat{sl}(2)$  that is exact with respect to some operator  $\mathbb{Q}$  such that  $\mathbb{Q}^2 = 0$  and take cohomology wrt.  $\mathbb{Q}$



# Chiral algebras from higher dimensions

[Beem, Lemos, Liendo, Peelaers, Rastelli, van Rees '13]

Taking cohomology wrt.  $\mathbb{Q}$  = twisting

- Schematically  $\mathbb{Q} \sim \mathcal{Q} + \mathcal{S} \implies [\mathcal{O}]_{\mathbb{Q}}(z)$  where  $[\mathcal{O}]_{\mathbb{Q}} = \{ \mathcal{O} \mid \{ \mathbb{Q}, \mathcal{O} \} = 0, \mathcal{O} \neq \{ \mathbb{Q}, \mathcal{O}' \} \}$

**Examples of chiral algebras for  $\mathcal{N} = 2$  SCFT:**

Free hypermultiplet  $\rightarrow$  free symplectic boson algebra  $q_I(z)q_J(w) \sim \frac{\epsilon_{IJ}}{z-w}$

Free vector multiplet  $\rightarrow$  (b, c) system  $b(z)\partial c(w) \sim \frac{1}{(z-w)^2}, \quad \partial c(z)b(w) \sim -\frac{1}{(z-w)^2}$



# Top down construction of flat space holography

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Celestial gluon and graviton algebras related to 2D chiral algebras arising from SCFT/twisted holography/twistor theory

[Gaiotto, Costello '18; Adamo, Bu, Costello, Mason, Paquette, Sharma '21, '22, '23]

**Goal:** Look for 4D asymptotically flat bulk theory dual to celestial 2D chiral algebras



# Top down construction of flat space holography

[Costello, Paquette, Sharma '22, '23]

Consider  $\mathbb{R}^4 \simeq \mathbb{C}^2$  with coordinates  $x^\mu \rightarrow x = x_\mu \sigma^\mu$

$$u^{\dot{\alpha}} \equiv x^{1\dot{\alpha}}, \quad \hat{u}^{\dot{\alpha}} \equiv x^{2\dot{\alpha}}$$

Equip  $\mathbb{C}^2 \setminus \{0\}$  with metric associated to Kähler form  $\omega = \partial \bar{\partial} K$ ,

$$\partial \equiv du^{\dot{\alpha}} \partial_{u^{\dot{\alpha}}}, \quad \bar{\partial} \equiv d\hat{u}^{\dot{\alpha}} \partial_{\hat{u}^{\dot{\alpha}}}, \quad \|u\|^2 = u^{\dot{\alpha}} \hat{u}_{\dot{\alpha}}$$

$$K = \|u\|^2 + \log \|u\|^2$$

- also known as the Burns metric
- self-dual,  $R = 0$ ,  $R_{\mu\nu} \neq 0$ , asymptotically flat:

$$g_{\mu\nu} = \delta_{\mu\nu} + \mathcal{O}(\|u\|^{-2}), \quad \|u\|^2 \propto \delta_{\mu\nu} x^\mu x^\nu \rightarrow \infty$$

- bulk theory:  $WZW_4$  on Burns space



# Top down construction of flat space holography

[Costello, Paquette, Sharma '22, '23]

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- bulk theory:  $WZW_4$  on Burns space

$$\widetilde{\mathbb{C}}^2 = \mathbb{C}^2 \text{ with origin replaced by } \mathbb{CP}^1$$

$$\mathcal{S} = \frac{N}{8\pi^2} \int_{\widetilde{\mathbb{C}}^2} \partial\bar{\partial}K \wedge \text{tr} (g\partial g^{-1} \wedge g\bar{\partial}g^{-1})$$

$$- \frac{N}{24\pi^2} \int_{\widetilde{\mathbb{C}}^2 \times [0,1]} \partial\bar{\partial}K \wedge \text{tr} (\tilde{g}d\tilde{g}^{-1})^3$$

$$g : \widetilde{\mathbb{C}}^2 \rightarrow SO(8), \quad \frac{iN}{2\pi} \int_{\mathbb{CP}^1} \partial\bar{\partial}K = N \in \mathbb{Z}_+$$



# Top down construction of flat space holography

[Costello, Paquette, Sharma '22]

$$\begin{aligned} \mathcal{S} &= \frac{N}{8\pi^2} \int_{\widetilde{\mathbb{C}^2}} \partial\bar{\partial}K \wedge \text{tr} (g\partial g^{-1} \wedge g\bar{\partial}g^{-1}) \\ &\quad - \frac{N}{24\pi^2} \int_{\widetilde{\mathbb{C}^2} \times [0,1]} \partial\bar{\partial}K \wedge \text{tr} (\tilde{g}d\tilde{g}^{-1})^3 \\ &\rightarrow \frac{N}{8\pi^2} \int_{\mathbb{C}^2} \partial\bar{\partial}K \wedge \text{tr} \left( \partial\phi \wedge \bar{\partial}\phi - \frac{1}{3}\phi[\partial\phi, \bar{\partial}\phi] \right) + \mathcal{O}(\phi^4) \end{aligned}$$

- look for perturbative solutions by setting  $g = e^\phi$ ,  $\tilde{g} = e^{t\phi}$
- “asymptotic” states obey the wave equation on Burns space admitting a family of solutions:

$$\phi_a(z, \tilde{\lambda}) = \sum_{k,\ell} \frac{1}{k!\ell!} \tilde{\lambda}_1^k \tilde{\lambda}_2^\ell \phi_a[k, \ell](z)$$



# Top down construction of flat space holography

[Costello, Paquette, Sharma '22]

$$\mathcal{S} = \frac{N}{8\pi^2} \int_{\widetilde{\mathbb{C}^2}} \partial\bar{\partial}K \wedge \text{tr} (g\partial g^{-1} \wedge g\bar{\partial}g^{-1}) - \frac{N}{24\pi^2} \int_{\widetilde{\mathbb{C}^2} \times [0,1]} \partial\bar{\partial}K \wedge \text{tr} (\tilde{g}d\tilde{g}^{-1})^3$$

$$\rightarrow \frac{N}{8\pi^2} \int_{\mathbb{C}^2} \partial\bar{\partial}K \wedge \text{tr} \left( \partial\phi \wedge \bar{\partial}\phi - \frac{1}{3}\phi[\partial\phi, \bar{\partial}\phi] \right) + \mathcal{O}(\phi^4)$$

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## Holographic dictionary

$$\phi_a[k, \ell](z) \leftrightarrow J_a[k, \ell](z)$$

$J_a[k, \ell](z)$  2D chiral algebra generators



# Top down construction of flat space holography

[Costello, Paquette, Sharma '22]

$$\mathcal{S} = \frac{N}{8\pi^2} \int_{\widetilde{\mathbb{C}^2}} \partial\bar{\partial}K \wedge \text{tr} (g\partial g^{-1} \wedge g\bar{\partial}g^{-1}) - \frac{N}{24\pi^2} \int_{\widetilde{\mathbb{C}^2 \times [0,1]}} \partial\bar{\partial}K \wedge \text{tr} (\tilde{g}d\tilde{g}^{-1})^3$$

$$\rightarrow \frac{N}{8\pi^2} \int_{\mathbb{C}^2} \partial\bar{\partial}K \wedge \text{tr} \left( \partial\phi \wedge \bar{\partial}\phi - \frac{1}{3}\phi[\partial\phi, \bar{\partial}\phi] \right) + \mathcal{O}(\phi^4)$$

4D perturbative gluon amplitudes on Burns space

- look for perturbative solutions by setting  $g = e^\phi$ ,  $\tilde{g} = e^{t\phi}$
- “asymptotic” states obey the wave equation on Burns space admitting a family of solutions:

$$\phi_a(z, \tilde{\lambda}) = \sum_{k,\ell} \frac{1}{k!\ell!} \tilde{\lambda}_i^k \tilde{\lambda}_j^\ell \phi_a[k, \ell](z)$$

$ij \rightarrow a$

2D OPE of  $J_{ij}[\tilde{\lambda}](z)$

matches identity contribution to OPE

- two-point  $A(1,2) = -\frac{N}{z_{12}^2} J_0 \left( 2\sqrt{\frac{[12]}{z_{12}}} \right) \text{tr} (T_{a_1} T_{a_2})$

- Matching also established for the three-point amplitudes/correlators

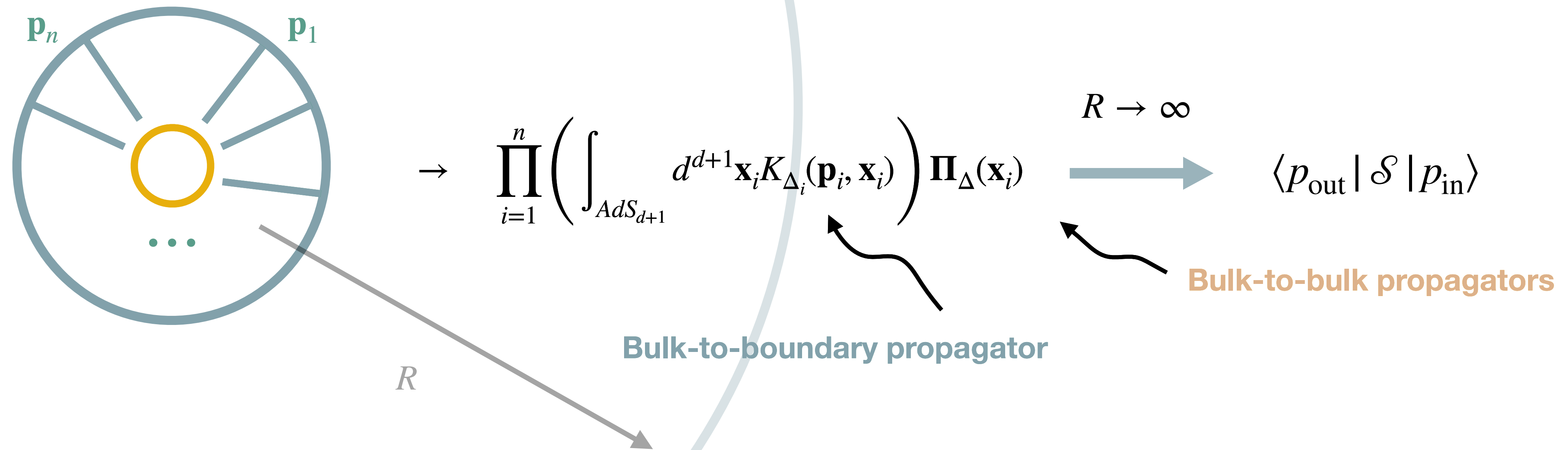


# AdS/CFT in the flat space limit

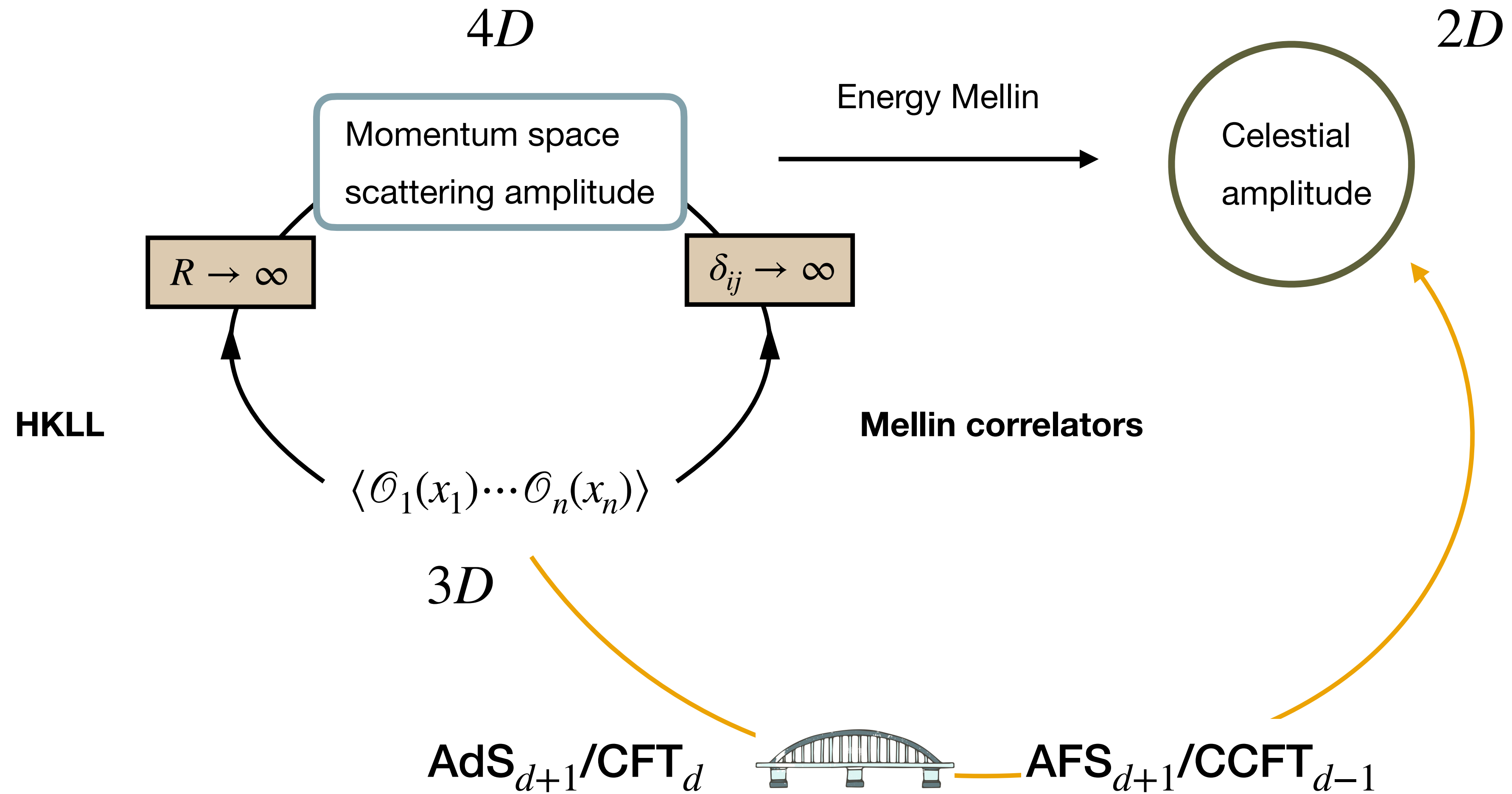
$\sum$  Amplitudes in  $AdS_{d+1}$  (Witten diagrams)  $\leftrightarrow$  Correlation functions in  $CFT_d$

- CFT (Mellin) correlators related to flat space scattering amplitudes at **large AdS radius**

[Polchinski '99; Susskind '99; Giddings '99; Penedones '10;...; Hijano, Neuenfeld '20]

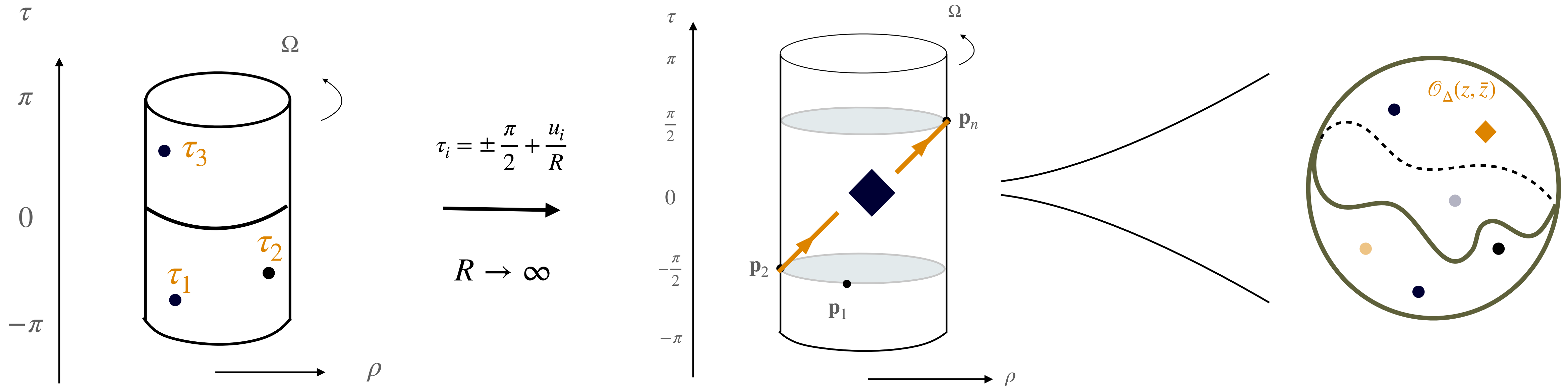


# Celestial amplitudes from AdS/CFT





# Celestial amplitudes from AdS/CFT



AdS<sub>4</sub> boundary observables

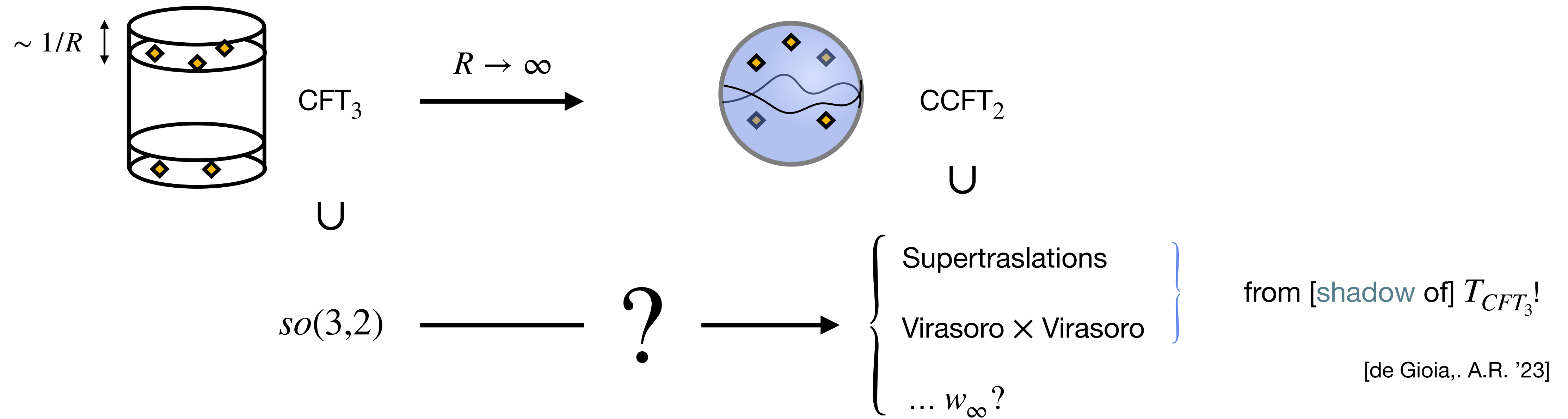


3+1D flat space celestial observables

# Celestial sector in CFT



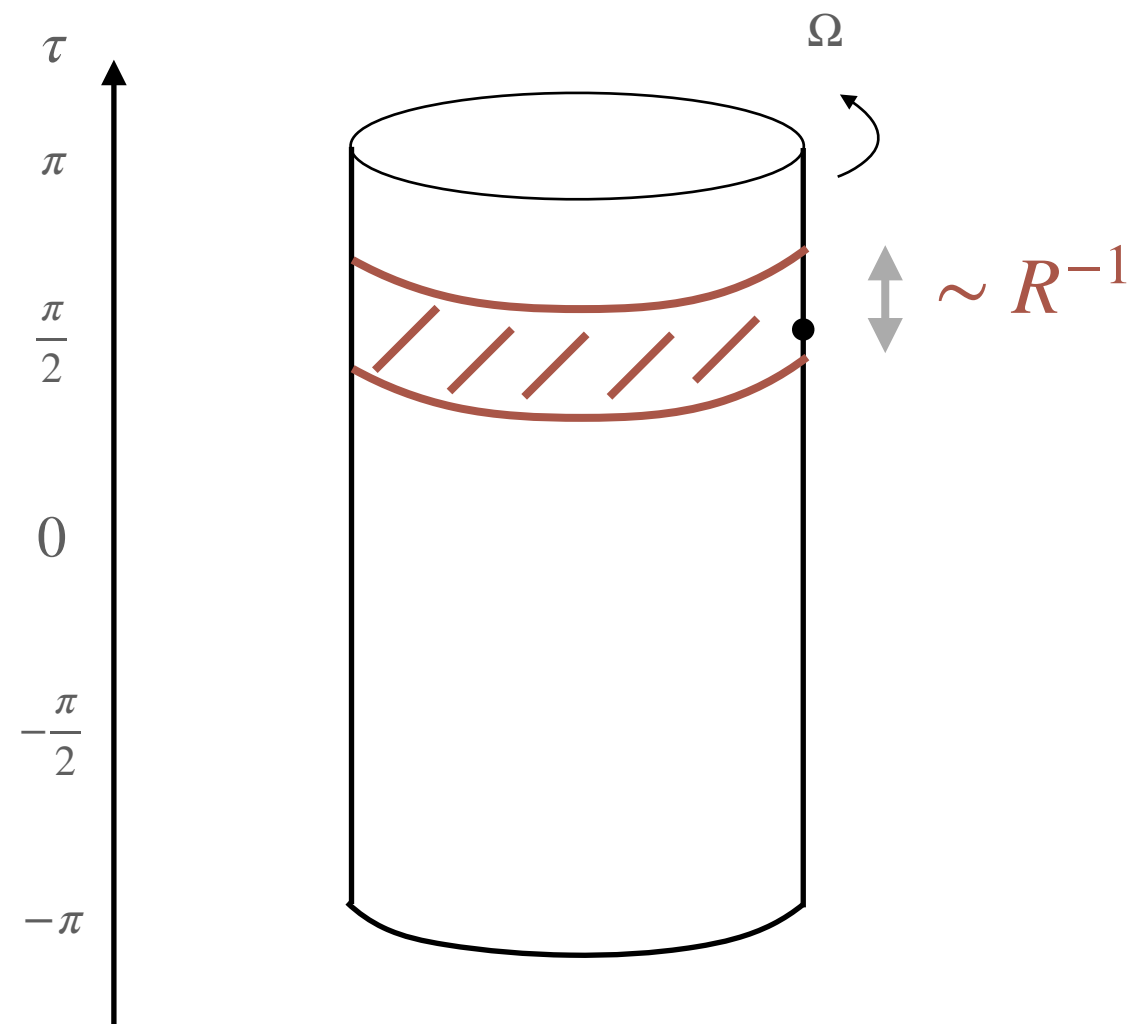
Infinity of symmetries from  $CFT_d$





# Symmetries of infinitesimal time intervals in 3D CFT

Analyze conformal symmetries in infinitesimal interval on the Lorentzian cylinder:



$\sim R^{-1}$ 
 $ds^2 = -d\tau^2 + 2\gamma_{z\bar{z}}dzd\bar{z}, \quad \gamma_{z\bar{z}} = \frac{2}{(1+z\bar{z})^2}$ 
 $\xrightarrow{\tau = \tau_0 + \frac{u}{R}}$ 
 $ds^2 = -\boxed{R^{-2}du^2} + 2\gamma_{z\bar{z}}dzd\bar{z}$ 
  
 $\rightarrow 0 \text{ as } R \rightarrow \infty$

Conformal Killing vectors in the interval:  $\nabla_\mu \epsilon_\nu + \nabla_\nu \epsilon_\mu = \frac{2}{d} \nabla \cdot \epsilon(x) g_{\mu\nu}$

As  $R \rightarrow \infty$ , solutions parameterized by a function  $f$  and a vector field  $Y^A$  on the sphere:

$$\epsilon^\pm = \left[ \mp \frac{iR}{2} F_\pm(u) D \cdot Y(z, \bar{z}) + f(z, \bar{z}) \right] \partial_u + F_\pm(u) Y^A(z, \bar{z}) \partial_A$$



# BMS<sub>4</sub> algebra in the strip

- For constant  $f$  and  $Y$  global CKV,  $\epsilon^\pm$  reorganize into generators of  $so(3,2)$  - Lorentz generators  $M^{\mu\nu}$  in 5d embedding space
- Inonu-Wigner contraction  $\mathcal{P}^\mu = \frac{1}{R}M^{4\mu}$ ,  $\mu = 0, \dots, 3$  with  $\mathcal{P}^\mu$ ,  $M_{\mu\nu}$  fixed as  $R \rightarrow \infty$  yields Poincare algebra

- For  $Y^A(z, \bar{z})$  arbitrary CKV, contraction yields 
$$\begin{cases} L_Y = iY^A \partial_A + i\frac{u}{2} D \cdot Y \partial_u + O(R^{-2}) \\ T_f \equiv i\epsilon_f = if(z, \bar{z}) \partial_u + O(R^{-2}) \end{cases}$$

which generate **ebms<sub>4</sub>**  $[T_{f_1}, T_{f_2}] = O(R^{-2})$ ,  $[L_{Y_1}, L_{Y_2}] = iL_{[Y_1, Y_2]} + O(R^{-2})$ ,  $[T_f, L_Y] = iT_{f'=\frac{1}{2}(D \cdot Y)f - Y(f)} + O(R^{-2})$

→ asymptotic symmetry algebra of 4D AFS!



# CCFT operators from 3D CFT operators

Conformal transformations in the strip  $\sim$  celestial symmetries in the “flat space” limit

$$\delta_\epsilon \mathcal{O}_\Delta(x) = - \left[ (\nabla \cdot \epsilon) \frac{\Delta}{3} + \epsilon^\mu \nabla_\mu + \frac{i}{2} \nabla_\mu \epsilon_\nu S^{\mu\nu} \right] \mathcal{O}_\Delta(x) \longrightarrow \text{transformation of CCFT}_2 \text{ primary operator}$$
$$\mathfrak{h} \equiv \frac{\hat{\Delta} + s}{2}, \quad \bar{\mathfrak{h}} \equiv \frac{\hat{\Delta} - s}{2}, \quad \hat{\Delta} \equiv \Delta + u \partial_u$$

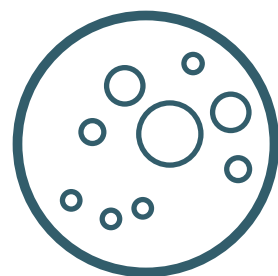
- Diagonalize weights via  $\widehat{\mathcal{O}}_\Delta(z, \bar{z}; \Delta_0) \equiv N(\Delta, \Delta_0) \int_{-\infty}^{\infty} du u^{-\Delta_0} \mathcal{O}_\Delta(u, z, \bar{z})$

- Same as transform relating Carrollian and celestial conformal field theories

[Donnay, Fiorucci, Herfray, Ruzziconi '22]

- Shadow stress tensor Ward identity in 3D CFT lead to **leading and subleading conformally soft graviton theorems in 2D CCFT**

[Kapec, Mitra '18; de Gioia, A.R. '23]



# Flat space holography summary



- some of the novelties recovered in flat limit of (bottom up) AdS/CFT

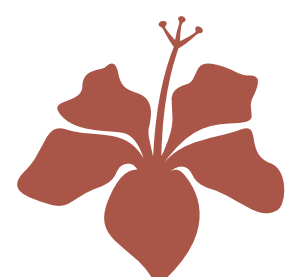
- from OPE to conformally soft currents

- top down realization of flat space holography

- high- and low-energy behavior of 4D scattering amplitudes reflected in analytic structure of celestial amplitudes in  $\Delta$

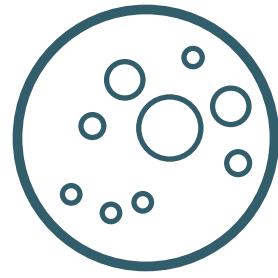
- low point amplitudes are distributions

- singularities disappear in special backgrounds



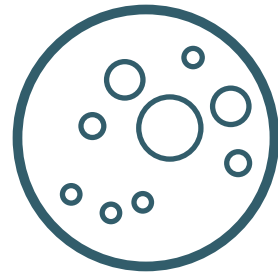
**Boost eigenbasis**

$$|\Delta, z, \bar{z}\rangle \equiv \int_0^\infty d\omega \omega^{\Delta-1} |\omega, z, \bar{z}\rangle \rightarrow \int_0^\infty d\omega \omega^{\Delta-1} |\lambda\omega, z, \bar{z}\rangle = \lambda^{-\Delta} |\Delta, z, \bar{z}\rangle$$



- Bulk/geometric interpretation of in flat space limit
- AdS boundary conditions -  $\bar{T}T$  deformations?
- Central extensions
- Entanglement entropy; black holes?
- Spectrum of CCFT; conformal block decompositions
- Connections to all  $\Lambda$  via AdS/dS slicing
- Top down constructions

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## Connections to scattering amplitudes

- OPE from EFT
- Higher-derivative & loop corrections to OPE

[He, Jiang, Ren, Spradlin, Taylor, Volovich, Zhu....]

- Double copy constructions [Casali, Puhm, Sharma,...]

- IR divergences + Dan's talk

- Self-dual amplitudes and black holes

- Discrete basis

## Asymptotic symmetries and Carrollian FT

String theory, BFFS, ...

[Adamo, Ball, Cotler, Crawley, Donnay, Fiorucci, Guevara, He, Kapec, Mason, Mitra, Narayana, Ruzziconi, Salzer, Strominger, Sharma, Tropper, Wang...]

...





Thank you!