

Towards gravitational scattering at fifth order in G

UCLA Mani L. Bhaumik Institute
for Theoretical Physics

Michael Ruf
From Amplitudes to Gravitational Waves, Nordita, July 24 2023

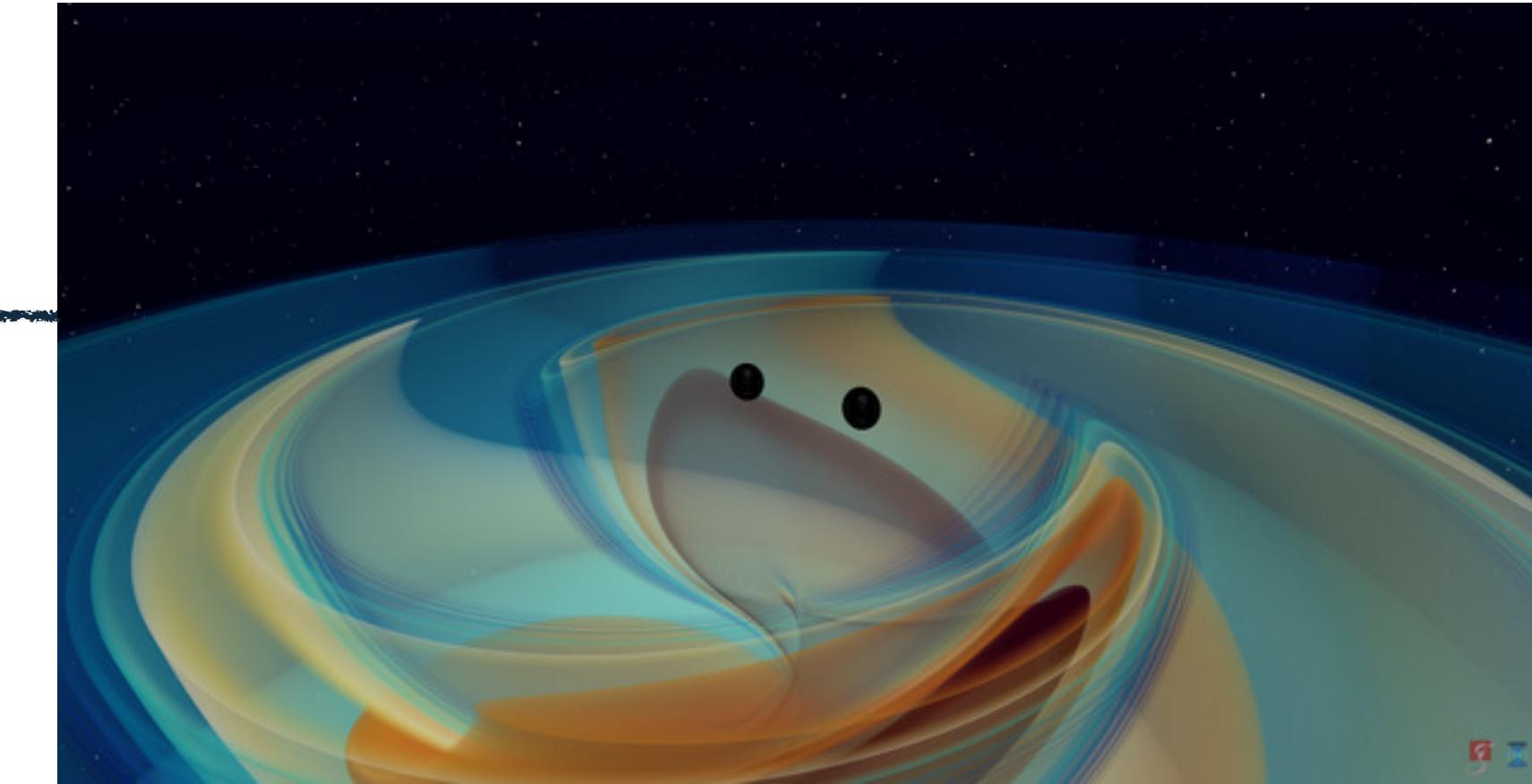
Based on work in collaboration w/ [Bern, Herrmann, Parra-Martinez, Roiban, A. Smirnov, V. Smirnov, Solon, Shen, Zeng]

Outline

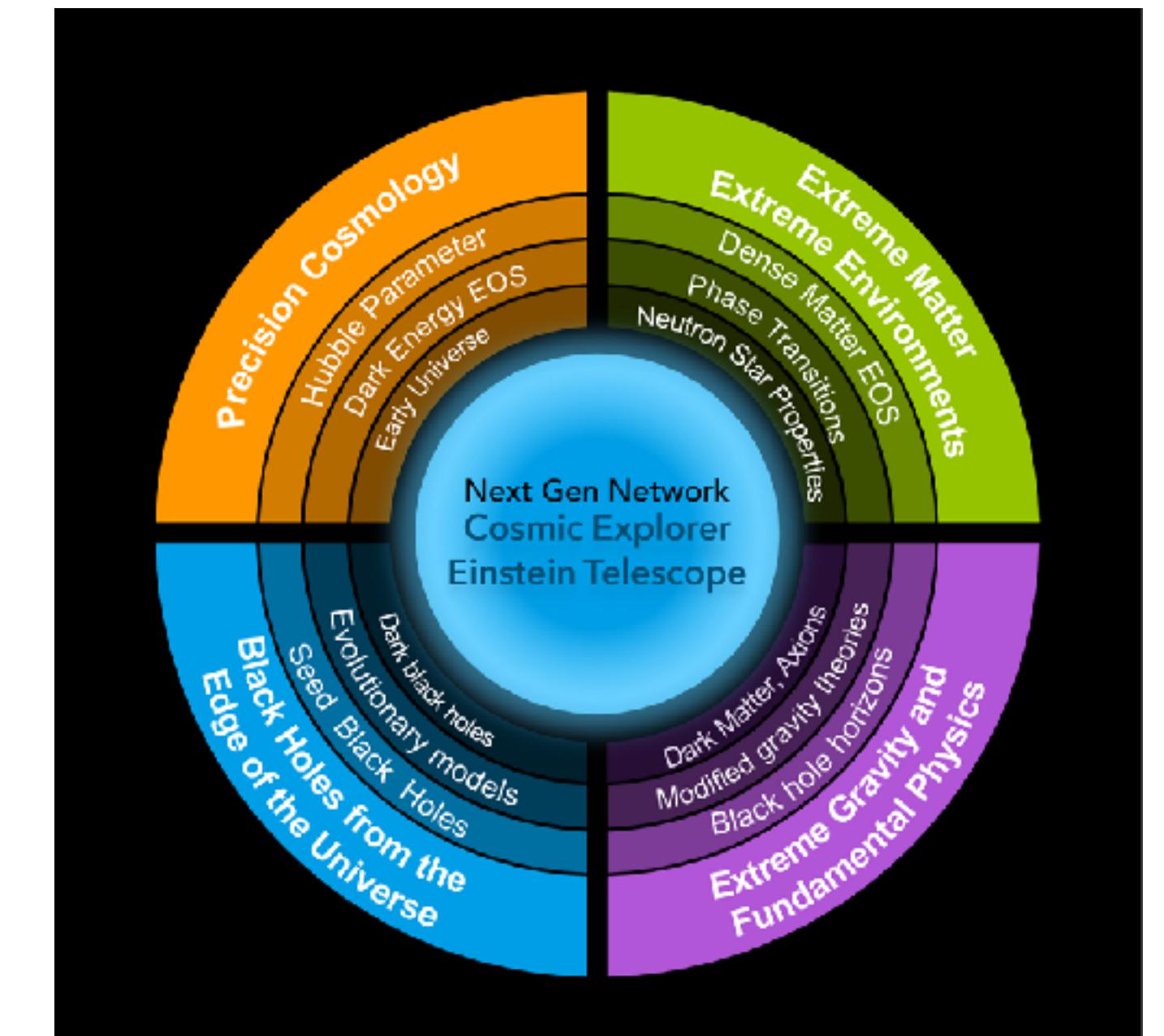
- Part 1: Recap scattering at order G^4
- Part 2: Towards scattering at order G^5

Motivation

- GW abundant, important source: compact binary systems
- Physics goals:
 - Strong-field tests of GR, new physics
 - BH properties, abundance etc.
 - Ultra-dense matter (neutron star equation of state)
 - Multi-messenger astronomy
 - ...

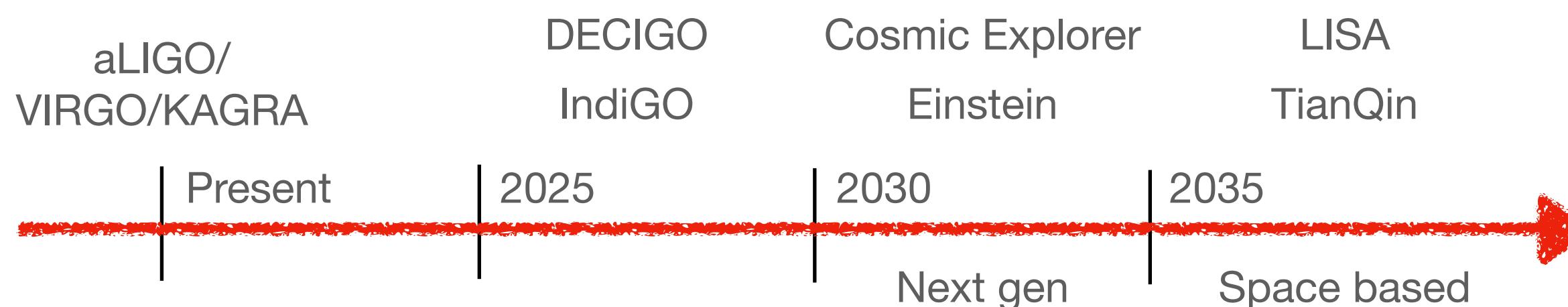
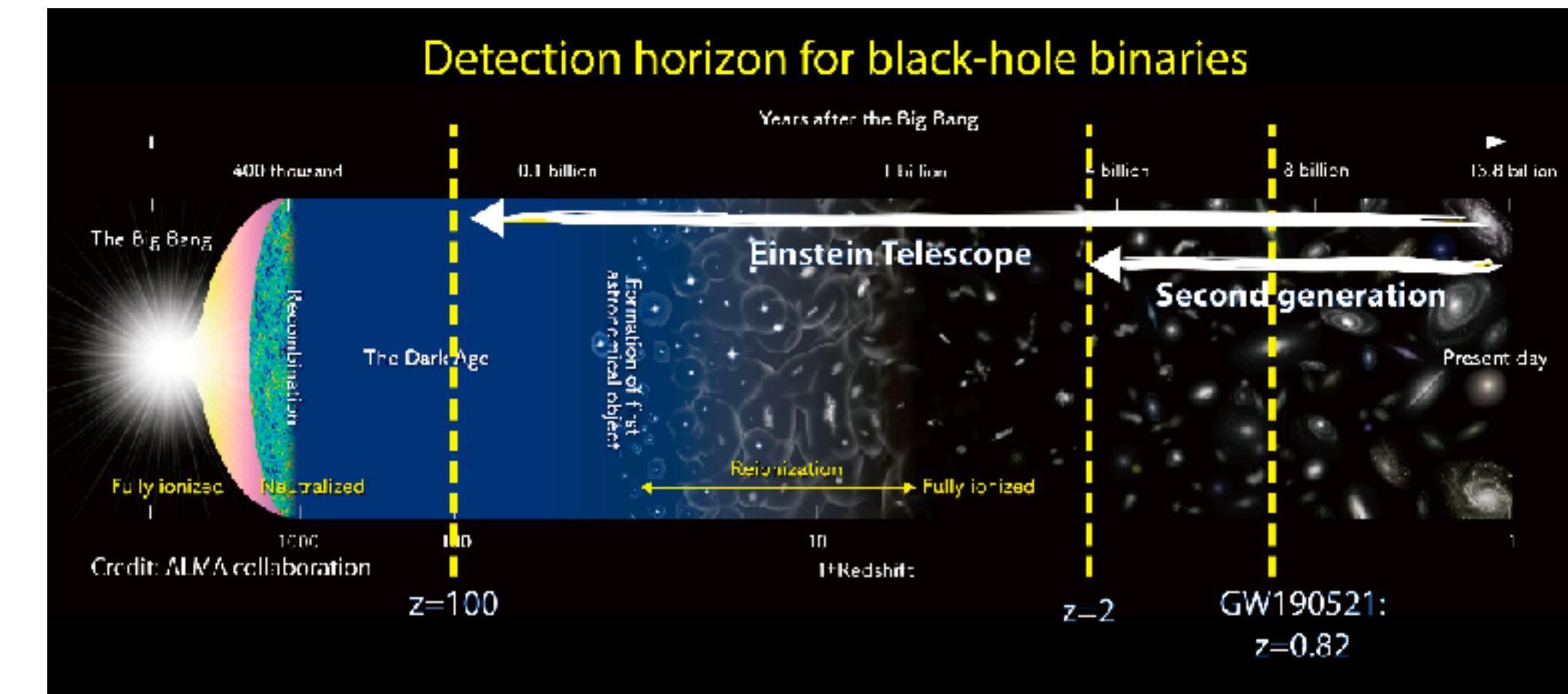


[GW190521, LIGO]



Motivation

- Present detectors (LIGO/VIRGO/KAGRA):
 - O(100) events. O4: 1 event/2.5 days
- Next-gen. experiments (ca. 2035):
 - More precision 10-100x improved S/N, wider frequency band
 - More data (bigger reach + sensitivity)
 - Extreme corners of parameter space, e.g. EMRI



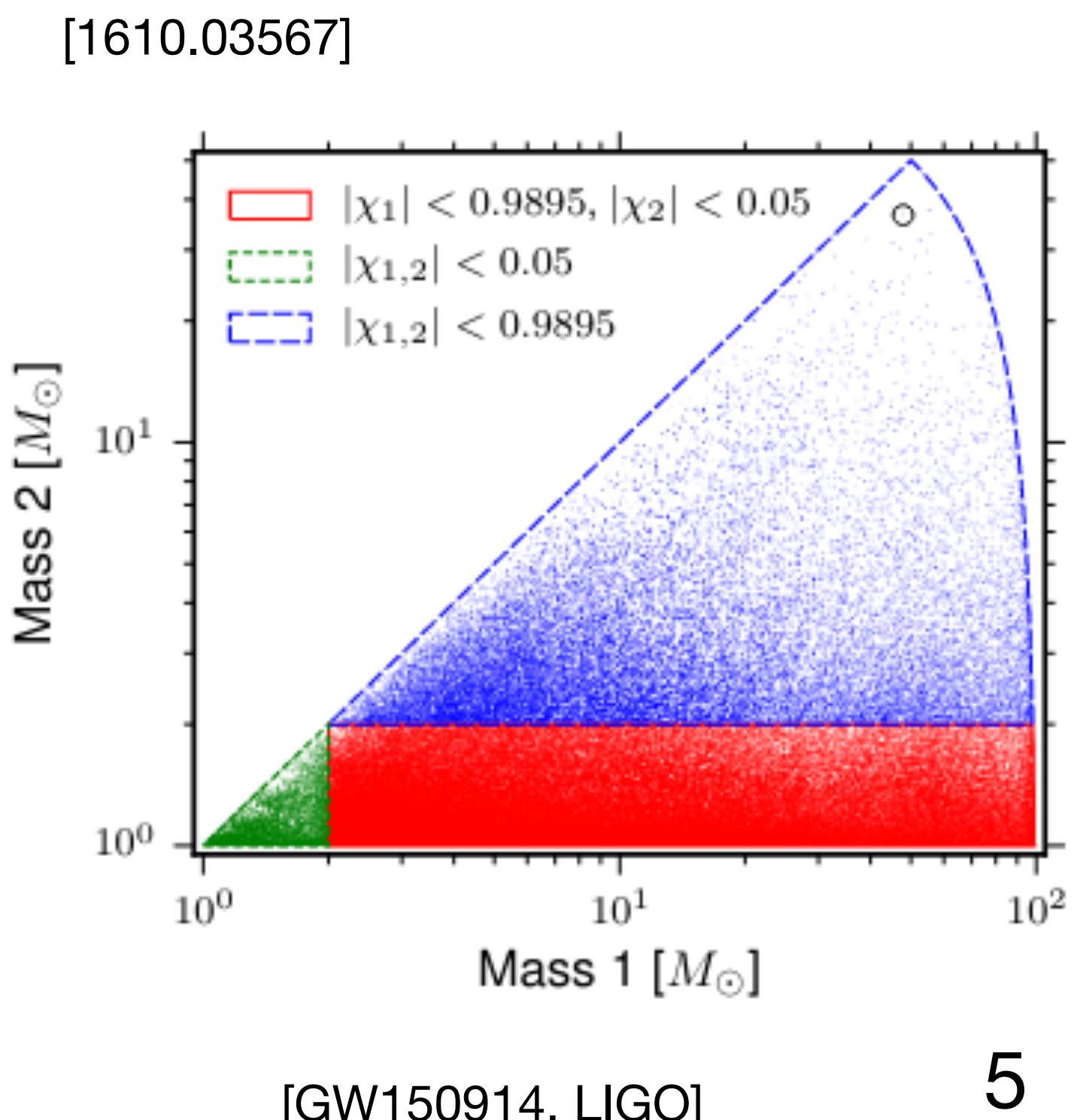
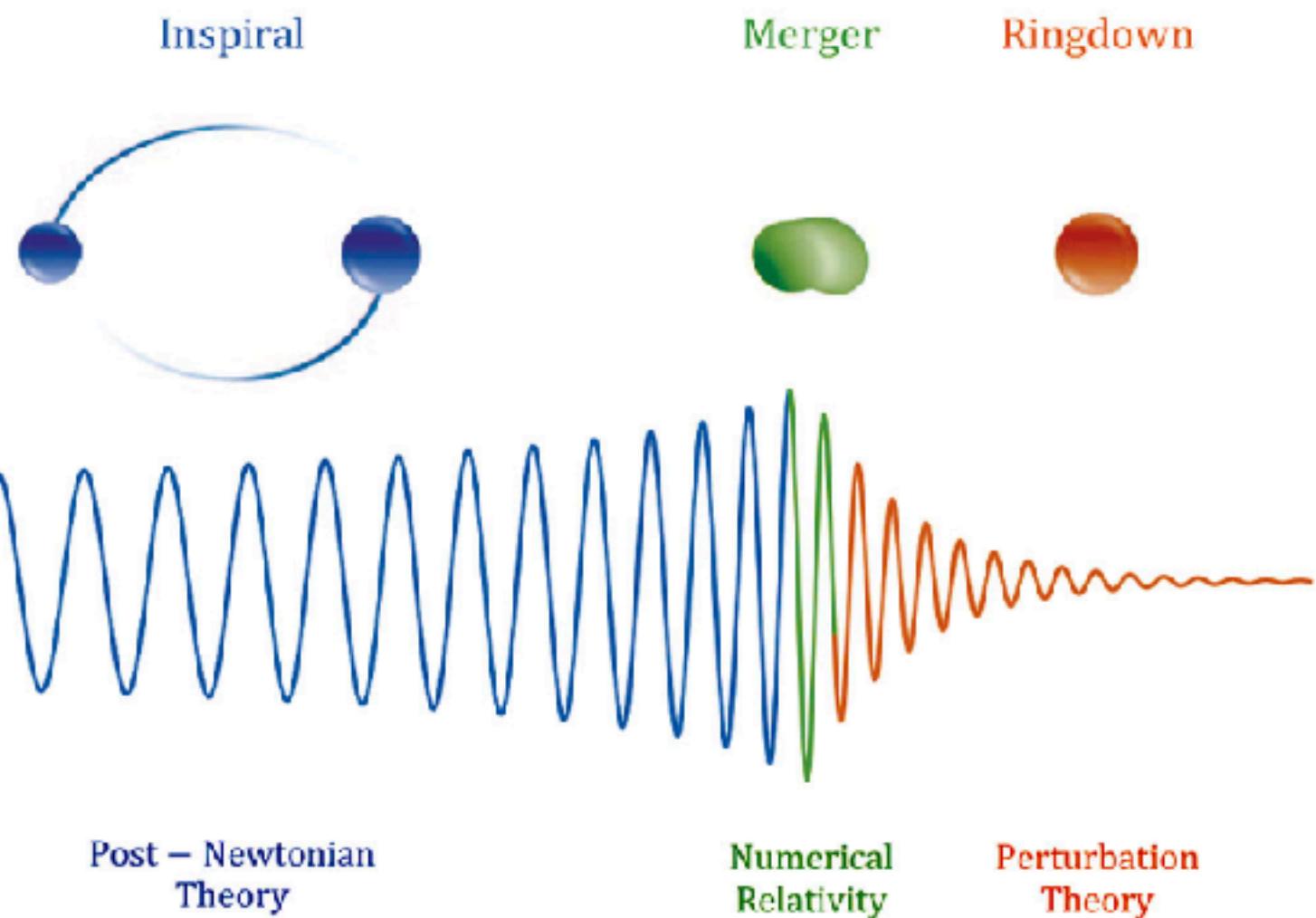
Need high-precision wave-form modelling!

Buonanno's talk

Motivation

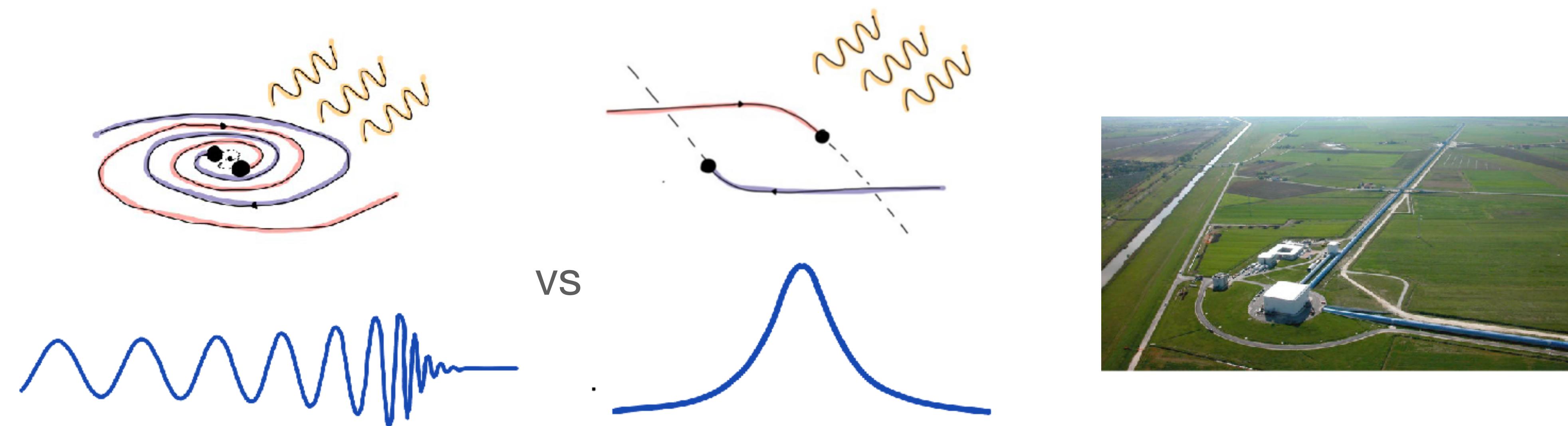
- Numerical waveform from Einstein eqn $G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$
- Significant resource requirements:
 - $O(10^5)$ CPU h/ NR template
 - GW150914: 250k templates
 - Challenging in PS-corners: $m_1 \ll m_2, v \rightarrow c, |\vec{L}|/m \rightarrow 1$
- Solution: analytic and hybrid models
(GW150914: post-Newtonian + effective-one-body)
- Corrections to Newton's potential to high orders

$$V(r) = -\frac{G\mu M}{r} + \frac{1}{c^2} \left[-\frac{3G\mu M v^2}{2r} + \frac{G^2 M^2}{r^2} \right] + \dots$$



Gravitational Scattering

Process of interest: scattering of compact massive objects



Hard to observe in GW observatories. Why bother?

Gravitational Scattering

Why bother?

1. Arguably simplest process, determined by initial data $\{p_1, p_2, b\}$

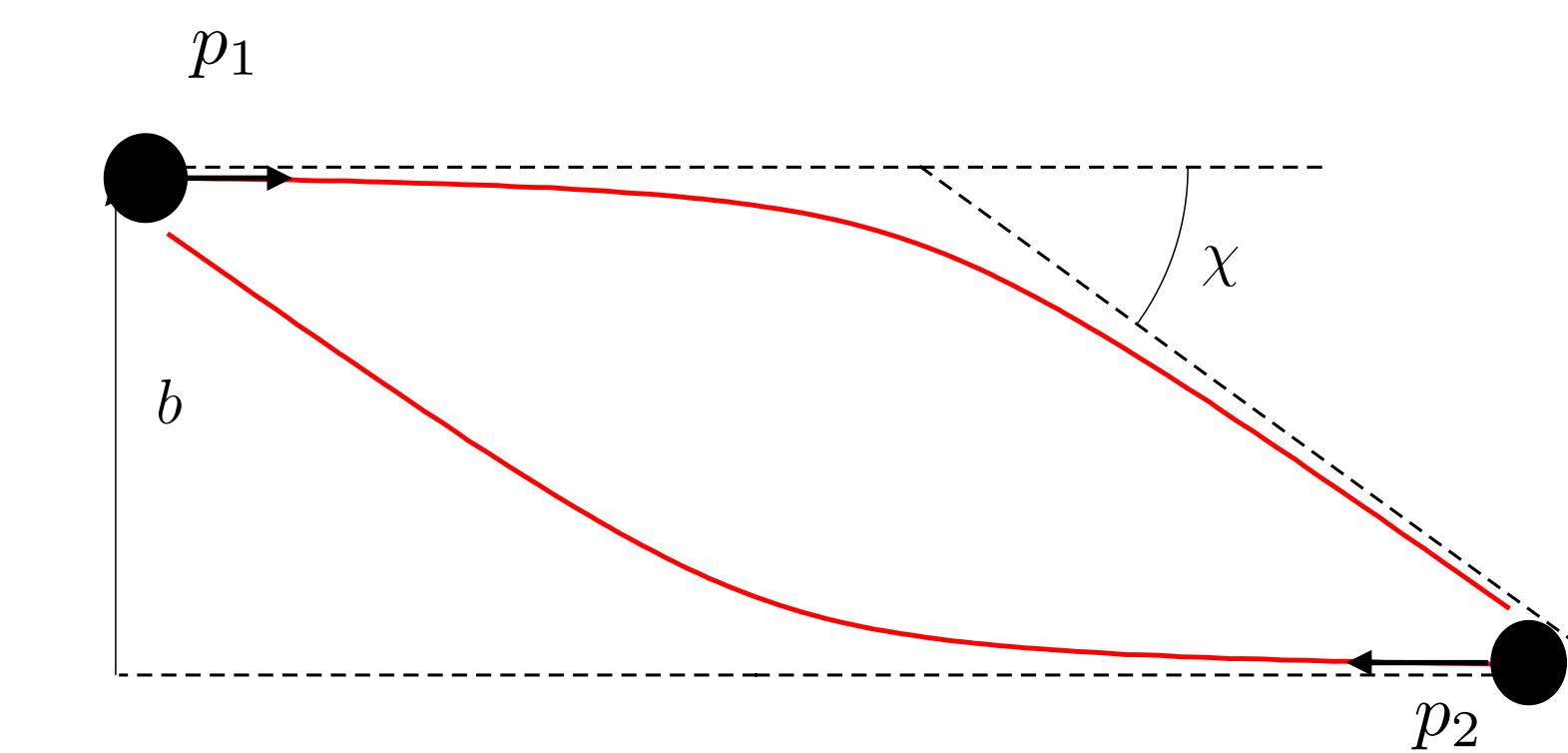
- Gauge/coordinate invariant approach
- Large separations: perturbative, no merger
- Benefits for numerical and analytic approach

2. Connection to the bound problem

- Key subtleties (e.g. hereditary effects) are present
- Universal information (e.g. instantaneous potentials)

3. Matches well with the amplitudes program

Relativistic treatment exposes additional structures, e.g. mass polynomiality [Damour]



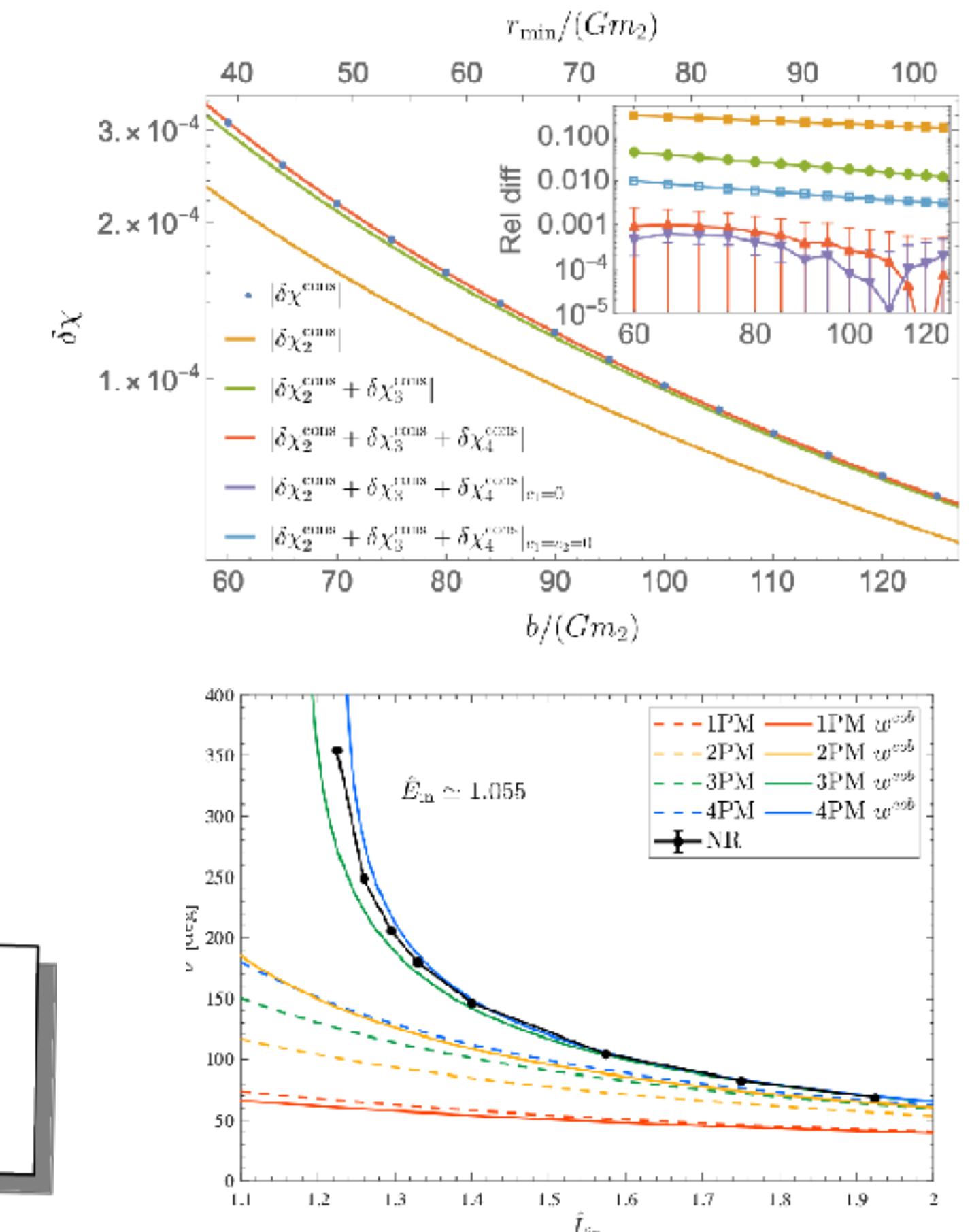
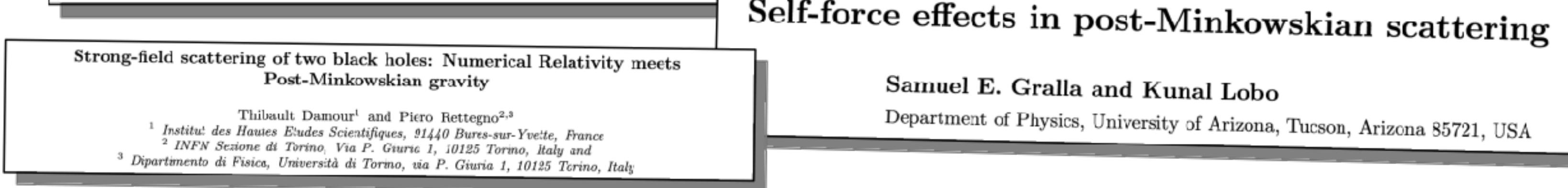
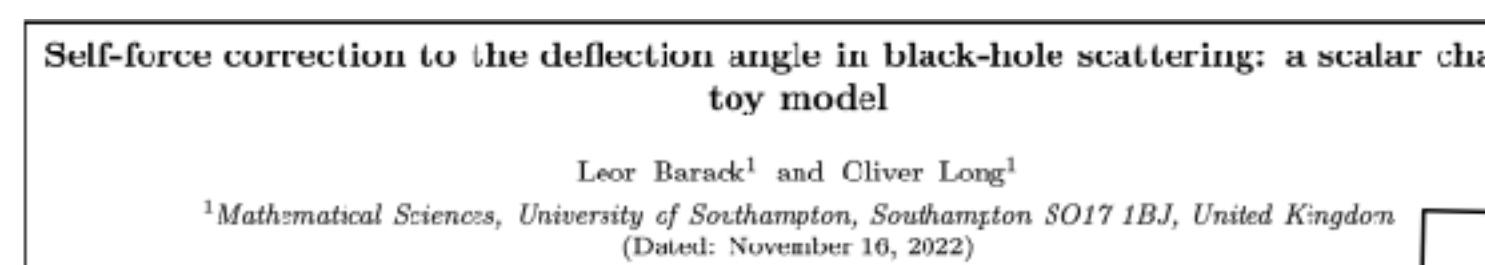
$$V(r) = -\frac{G\mu M}{r} + \frac{1}{c^2} \left[-\frac{3G\mu M \mathbf{v}^2}{2r} + \frac{G^2 M^2}{r^2} \right] + \dots$$

Gravitational Scattering

- The classical GR community is very interested in scattering

[Barack, Berti, Bini, Buonanno, Cardoso, Damour, East, Geralico, Gralla, Guercilena, Hinder, Hinderer, Hopper, Khalil, Lobo, Long, van de Meent, Nagar, Pfeiffer, Pratten Pretorius, Pretorius, Rettegno, Rezzolla, Schmidt, Sperhake, Steinhoff, Thomas, Vines, Whittall, Yunes, . . .]

- See e.g. Barack's workshop talk for collaborative effort
- Comparisons with percent-level agreement!
- Complementary approaches:
 - PM in weak field
 - SF/NR in strong field
 - ultimate goal: new hybrid models (e.g. SF-PM)



Gravitational Scattering

- Weak field perturbation theory $Gm/b \ll 1$ (PM-expansion) for large impact parameter scattering
- Two flavours:
 - Worldline based (classical or QFT [Mogull, Plefka, Steinhoff]) ← Talks by Kälin and Mogull + Workshop talk by Plefka
 - Amplitudes based
- Field is very mature, large number of new results in the past 4 years
 - High orders in perturbation theory
 - Spin ← Talk by Vines & tidal effects
 - Waveforms ← Talk by Travaglini
 - Radiative quantities ← Talk by Heissenberg
 - Self-force ← Talk by Cristofoli's
- And better understanding
 - High energy limit
 - Exponentiation (Eikonal, HEFT, Radial action,...)
 - Spin, angular momentum,...

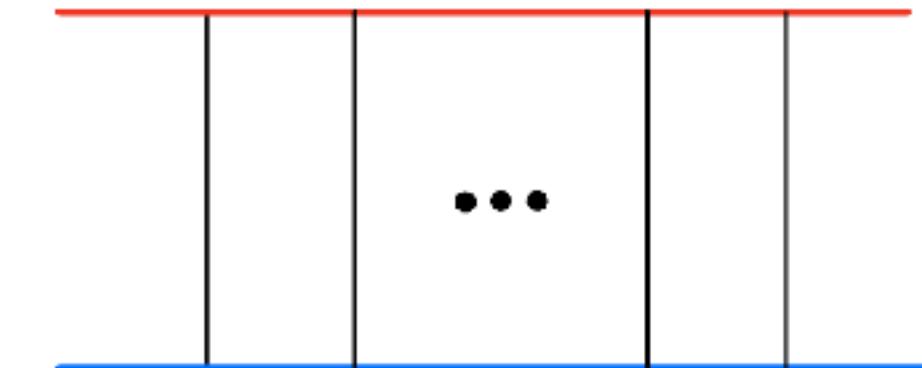
Classical physics from Amplitudes

- Basic idea:
 - Can evaluate scattering amplitudes very efficiently
 - S-Matrix has all information on classical scattering
- Old idea, dating back to '60 [Corinaldesi, Iwasaki, Feynman, Barker, Gupta, Kaskas]
- Revived by Damour's 2017 paper
- Modern program: state of the art results to enter GW template pipeline

Quantum viewpoint for classical physics extremely healthy!

Amplitudes to observables

- Amplitude not observable: $\mathcal{M}_{L\text{-loop}} \sim \frac{1}{\hbar^{L+1}}$, $\mathcal{M}_{L\text{-loop}} \sim \frac{1}{\epsilon^L}$
- Observables through
 - Direct computation [Kosower, Maybee, O'Connell; Damgaard, Hansen, Planté, Vanhove (4PM)]
 - Hamiltonian (Schrödinger eqn. or EFT matching [Rothstein,Neill; Cheung, Solon, Rothstein])
 - Stationary phase/generating functionals (eikonal, partial waves, heavy particle phase,...)
- Amplitude \leftrightarrow radial action [Bern, Parra-Martinez, Roiban, **MSR**, Shen, Solon, Zeng]



$$\mathcal{M} = i \int_J (e^{iI_r(J)/\hbar} - 1), \quad I_r(J, E) = \int_{\text{trajectory}} p_r(J, E) dr \quad \chi(J, E) = -\partial_J I_r(J, E)$$

Gauge invariants

Very efficient extraction that meshes with relativistic integration

Classical Limit

- Classical physics: Large number of soft exchanges $q = \hbar\bar{q}$

$$1 \ll J^2 \sim \frac{s}{q^2} \sim \frac{m_i^2}{q^2} \rightarrow q^2 \ll m_i^2 \sim s$$

$$l_{\text{compton}} \sim \frac{1}{M} \ll R_S \sim GM \ll b$$

- Relativistic regions: [Benecke, Smirnov]

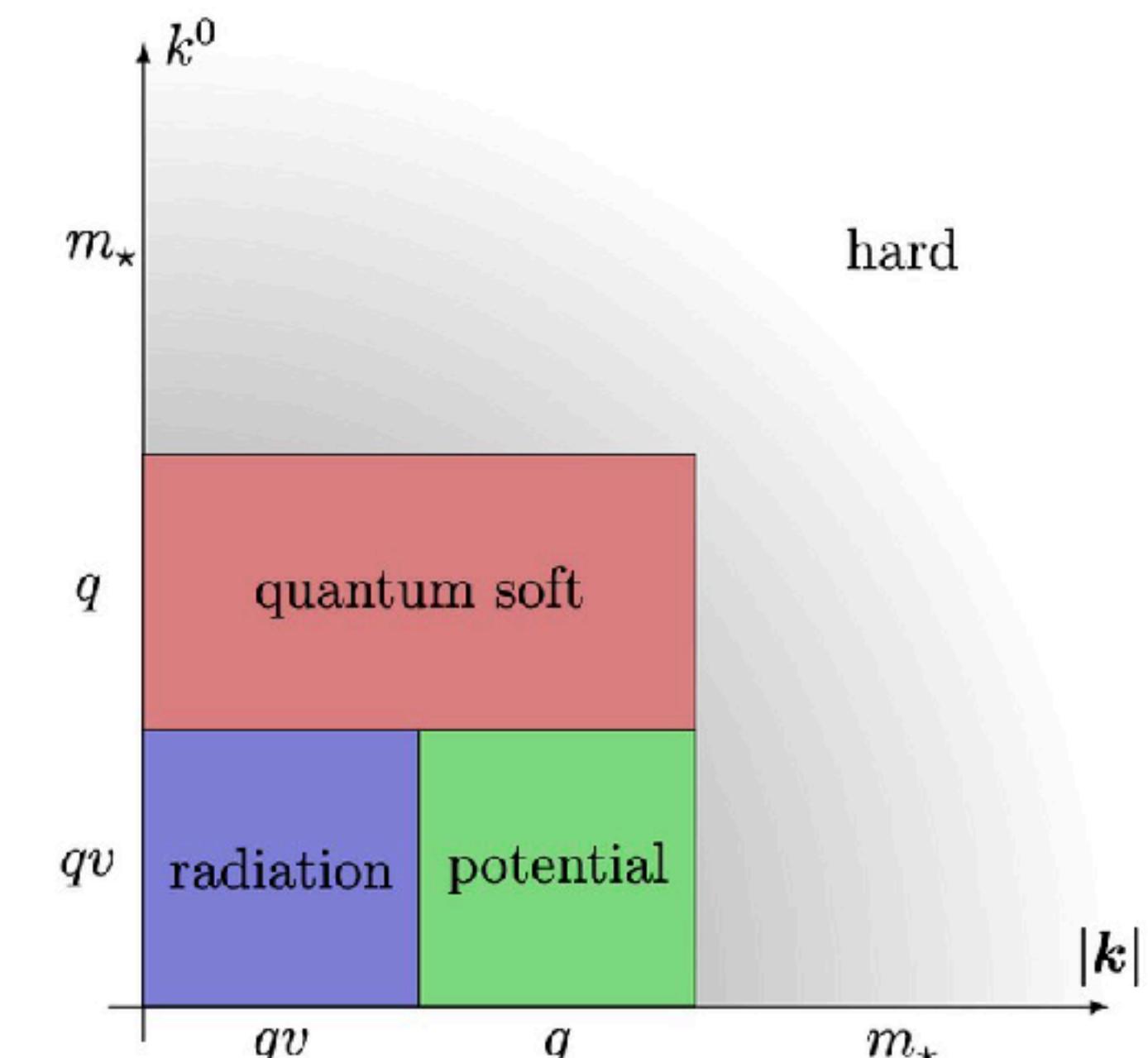
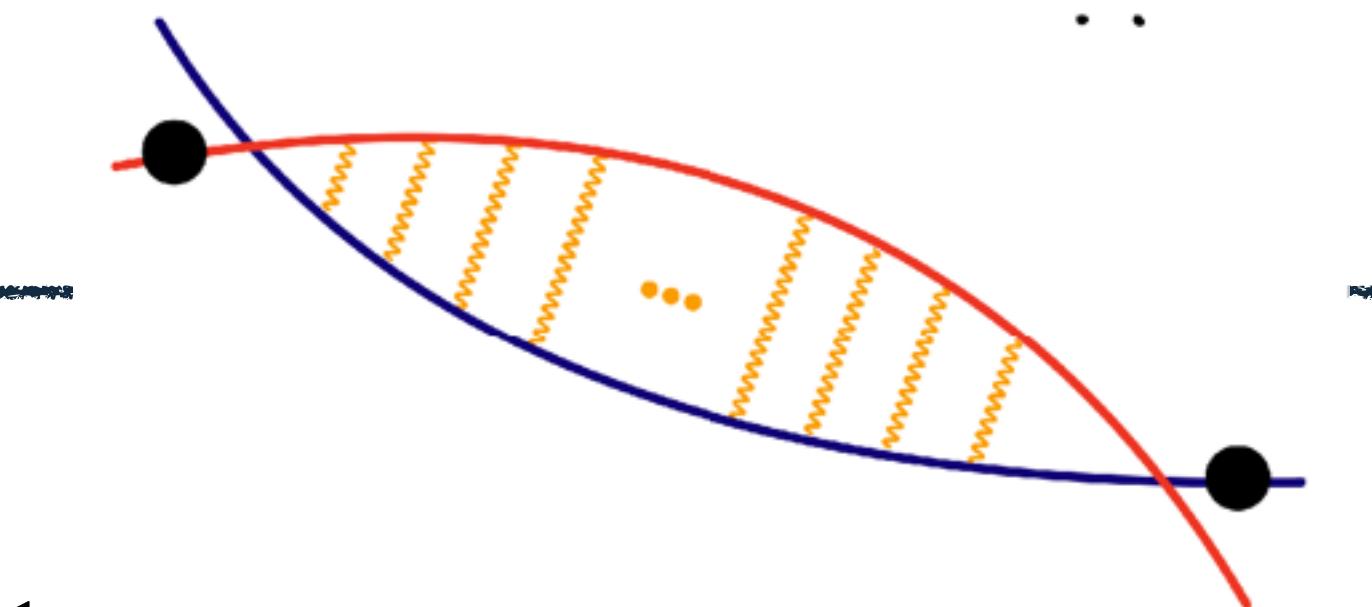
- Hard (h): $\ell \sim m \leftarrow \text{UV, quantum } \lambda_{\text{compton}} \sim b$
- Soft (s): $\ell \sim q \leftarrow \text{long range } \lambda_{\text{compton}} \ll b$

- Threshold expansion: $v \sim |\vec{p}_{\text{COM}}|/\sqrt{s}$
 - Potential (p): $(\omega, \vec{\ell}) \sim (|q|v, |q|) \leftarrow \text{instantaneous}$
 - Radiation (r): $(\omega, \vec{\ell}) \sim (|q|v, |q|v)$

- Classical physics (p)+(r), not well-defined separately

- Formally $v \ll 1$, resumption to $v \sim \mathcal{O}(1)$

$$v + \frac{v^3}{3} + \frac{v^5}{5} + \dots = \operatorname{arctanh}(v)$$



Amplitude computation

- Amplitude computation in parts
 - Constructing the integrand
 - Integral reduction
 - Evaluation of “master” integrals
- Separate problems each requiring some attention
- Integration identical in different approaches [Parra-Martinez, **MSR**, Zeng ‘20]
 - Amplitudes-based (traditional [Bern et. al.], heavy particle EFT [Damgaard, Haddad, Helset; Brandhuber et. al], velocity cuts [Bjerrum-Bohr, Plante, Vanhove], ...)
 - Worldline-based (traditional [Porto, Kalin], WQFT [Mogull, Plefka, Steinhoff])
- Improvements in integration beneficial for community
- More precision → more loops. $4\text{PM} = G^4 = 3 \text{ loops}$, $5\text{PM} = G^5 = 4 \text{ loops}$, ...

$$\mathcal{M}(p_1, p_2, p_3, p_4) = \int d^d \ell M(\ell) = f(\sigma) \log(\sigma) + \dots$$

Integrand – Generalized unitarity $M(\ell)$

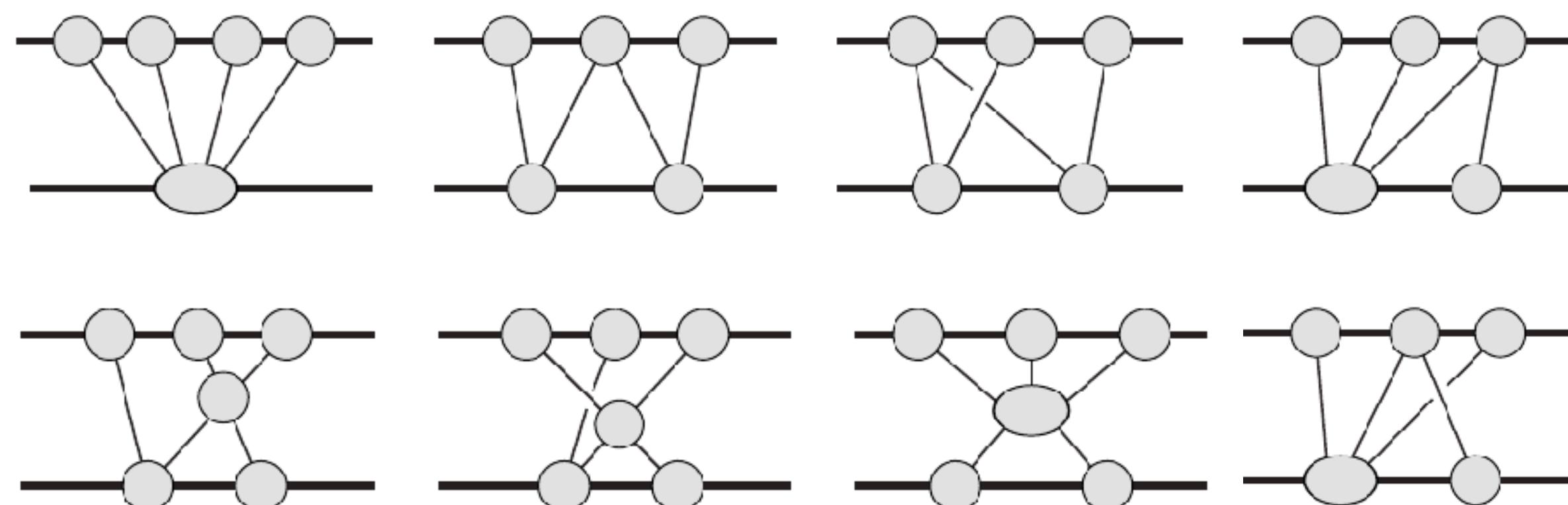
- Integrands from different methods
 - Feynman rules from Einstein-Hilbert action → straightforward
 - Generalized unitarity [Bern, Dixon, Dunbar, Kosower] → clean, more efficient
- Building blocks: on-shell tree amplitudes
 - Recursion from simpler trees (BCFW, Berends-Giele)
 - Double-copy → QCD trees
 - Sometimes even closed-form expressions!
- Recurring theme: on-shell amplitudes have structure (symmetries, analytic properties) – good idea to use this

Surprisingly, gravity interactions complicated but no bottleneck!

Integrand – Generalized unitarity

- Construct integrand from unitary cuts, at $\mathcal{O}(G^4)$:

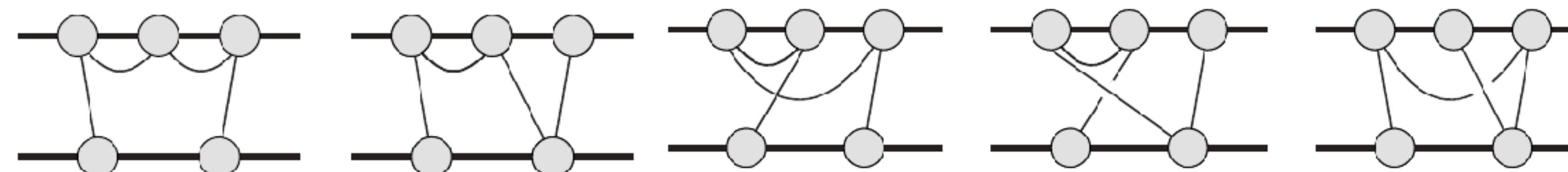
Potential/
Instantaneous



Blobs = tree amplitudes
Lines = on-shell states
 $E^2 = \vec{p}^2 + m^2$

$$\Rightarrow M(\ell)$$

Tail



- Drastically simplified in classical limit
 - No graviton loops, self energies, matter contacts
 - 1 matter line per loop

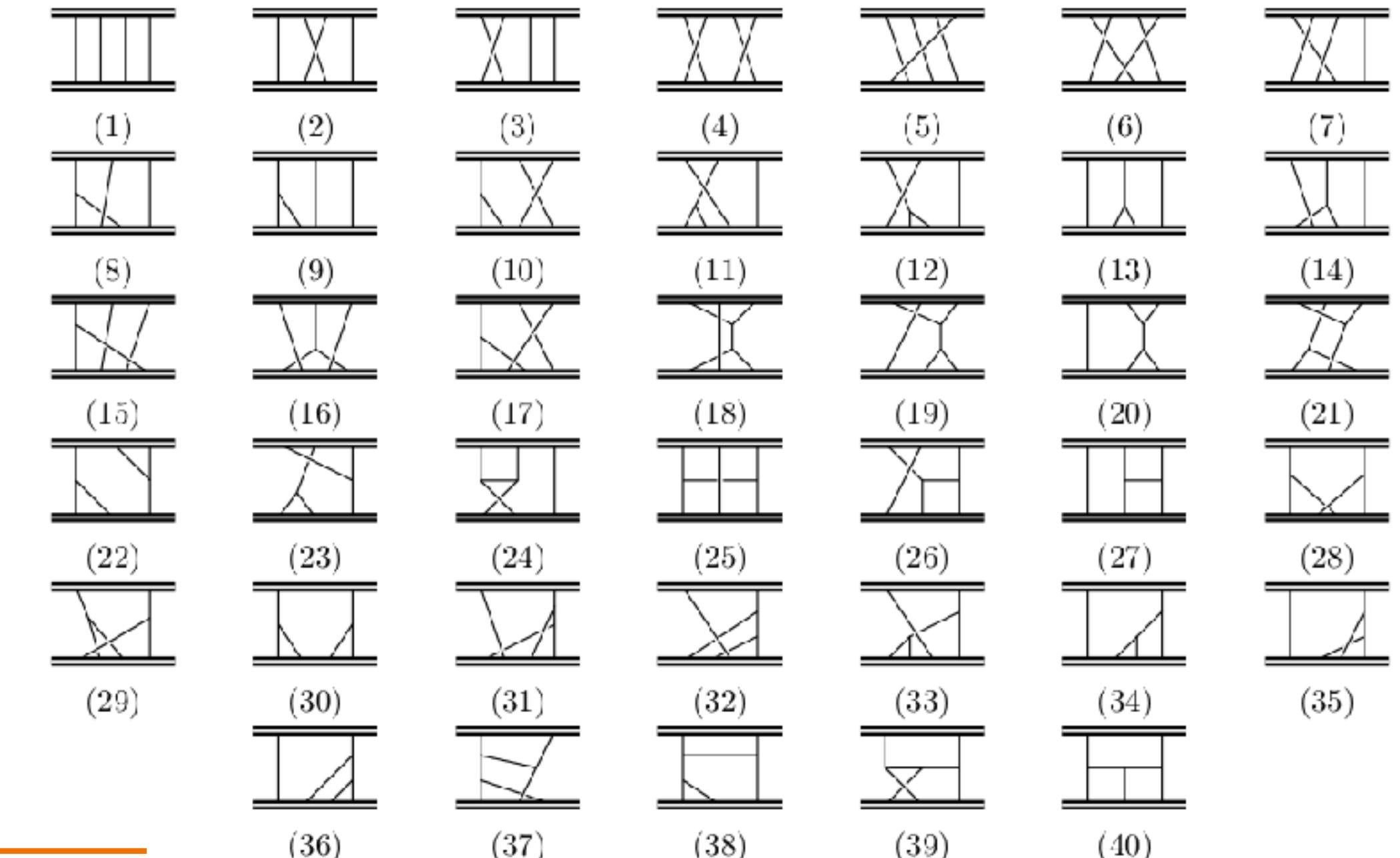
Integration

- Amplitude organized in 40 integral families

- Integrals as in heavy quark EFT



$$= \int d^D \ell \frac{1}{\ell^2} \frac{1}{(\ell - q)^2} \frac{1}{2u_1 \cdot \ell} \frac{1}{2u_2 \cdot \ell}$$

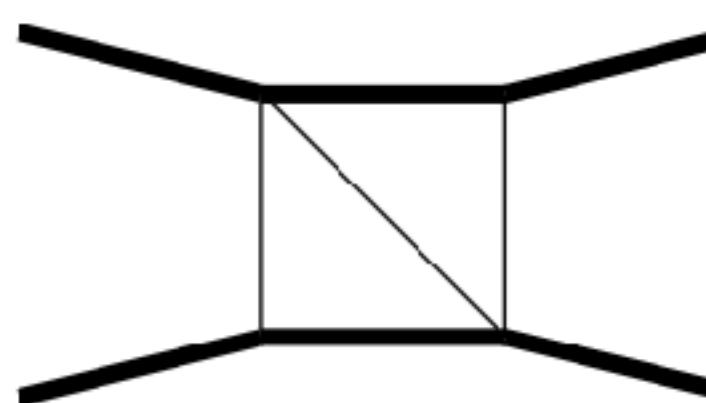


- Classical limit crucial:

- Fewer independent integrals

- Single-variable problem! $y = u_1 \cdot u_2 = 1/\sqrt{1 - v^2}$

- Simpler functions



\sim

$$\begin{aligned}
 &= 16 \log \frac{-t}{m^2} \left[\mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1+y}, \frac{1}{2}; \bar{x}, \vec{a}\right) - \mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1+y}, \frac{1}{2}; \bar{x}, \vec{a}\right) \right] \\
 &+ \mathcal{E}_4\left(-\frac{1}{\infty}, \frac{1}{1}, \frac{1}{1}; \bar{x}, \vec{a}\right) + \mathcal{E}_4\left(-\frac{1}{1}, \frac{1}{1}, \frac{1}{1}; \bar{x}, \vec{a}\right) + \zeta_2 \mathcal{E}_4\left(-\frac{1}{\infty}, \bar{x}, \vec{a}\right) + \zeta_2 \mathcal{E}_4\left(-\frac{1}{1}, \bar{x}, \vec{a}\right) \\
 &- 8 \left(8\zeta_2 + 4\text{Li}_2(y) + \log^2 y \right) \left[\mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1+y}, \frac{1}{2}; \bar{x}, \vec{a}\right) + \mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1+y}, \frac{1}{2}; \bar{x}, \vec{a}\right) - \mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1+y}, \frac{1}{2}; \bar{x}, \vec{a}\right) \right] \\
 &- 32 \zeta_2 \left[\mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1}; \bar{x}, \vec{a}\right) - \mathcal{E}_4\left(-\frac{1}{1}, \bar{x}, \vec{a}\right) \right] + 16 \mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1+y}, \frac{1}{1}; \bar{x}, \vec{a}\right) \\
 &- 32 \mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1+y}, \frac{1}{2}; \bar{x}, \vec{a}\right) - 16 \mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1+y}, \frac{1}{1}; \bar{x}, \vec{a}\right) + 32 \mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1+y}, \frac{1}{2}; \bar{x}, \vec{a}\right) \\
 &+ 16 \mathcal{E}_4\left(-\frac{1}{\infty}, \frac{1}{1}, \frac{1}{1}; \bar{x}, \vec{a}\right) - 24 \mathcal{E}_4\left(-\frac{1}{\infty}, \frac{1}{1}, \frac{1}{1}; \bar{x}, \vec{a}\right) - 32 \mathcal{E}_4\left(-\frac{1}{0}, \frac{1}{1+y}, \frac{1}{2}; \bar{x}, \vec{a}\right) \\
 &+ 16 \mathcal{E}_4\left(-\frac{1}{1}, \frac{1}{1}, \frac{1}{1}; \bar{x}, \vec{a}\right) + 40 \mathcal{E}_4\left(-\frac{1}{1}, \frac{1}{1}, \frac{1}{1}; \bar{x}, \vec{a}\right) - 32 \mathcal{E}_4\left(-\frac{1}{1}, \frac{1}{1}, \frac{1}{1}; \bar{x}, \vec{a}\right) \\
 &+ \frac{4}{3} (12\text{Li}_3(y) + 24\zeta_2 \log y + \log^3 y) [\mathcal{E}_4\left(-\frac{1}{\infty}, \bar{x}, \vec{a}\right) + \mathcal{E}_4\left(-\frac{1}{1}, \bar{x}, \vec{a}\right)] \\
 &+ 64\zeta_4 - 32\zeta_2 \text{Li}_2(y) + 16\text{Li}_4(y) + 8\zeta_2 \log^2 y + \frac{1}{3} \log^4 y.
 \end{aligned} \tag{5.8}$$



$\arccosh(y) \log(-t)$

Integration

- Thousands of integrals

$$I_{\text{1-loop}}[a_1, a_2, a_3, a_4] = \int d^D \ell \frac{1}{[\ell^2]^{a_1}} \frac{1}{[(\ell - q)^2]^{a_2}} \frac{1}{[2u_1 \cdot \ell]^{a_3}} \frac{1}{[2u_2 \cdot \ell]^{a_4}}, \quad I_{\text{3-loop}}[a_1, \dots, a_{15}]$$

- Integrals satisfy integration-by-parts (IBP) relations [Tkachov, Chetyrkin 1981]

$$0 = \int d^D \ell u_2 \cdot \frac{\partial}{\partial \ell} \frac{1}{[\ell^2]^{a_1}} \frac{1}{[(\ell - q)^2]^{a_2}} \frac{1}{[u_1 \cdot \ell]^{a_3}} \frac{1}{[u_2 \cdot \ell]^{a_4}} = -a_1 I[a_1 + 1, a_2, a_3, a_4 - 1] - a_2 I[a_1, a_2 + 1, a_3, a_4 - 1] + \dots$$

- Linear algebra problem using Laporta's algorithm [Laporta '01] [FIRE, Kira, Reduze...]

$$0 = -I[2,1,1,0] - I[1,2,1,0] + \dots$$

$$0 = -I[2,1,1,1] - I[1,2,0,1] + \dots$$

⋮

- Scales with number and range of a_i

Integration [Parra-Martinez, MSR, Zeng '20]

- Integral reduction: 137 irreducible “master” integrals
- Master integrals satisfy Fuchsian differential equations (DE) [Kotikov '91]

$$\frac{d}{dx} \vec{I} = \sum_k \frac{d \log(w_k)}{dx} A_k(\epsilon) \vec{I}, \quad w_k \in \{x, 1 \pm x, 1 + x^2\}, \quad |\vec{I}| = 137, \quad A_k(\epsilon) \in M_{137}(\mathbb{Q})[\epsilon]$$

- Change of basis $\vec{I} \rightarrow \vec{J} = T\vec{I}$ to canonical form [Henn '13]

$$\frac{d}{dx} \vec{J} = \epsilon \sum_k \frac{d \log(w_k)}{dx} B_k \vec{J}$$

$$\sigma = \frac{1+x^2}{2x} \quad 1 < \sigma \leftrightarrow x \in (0,1)$$

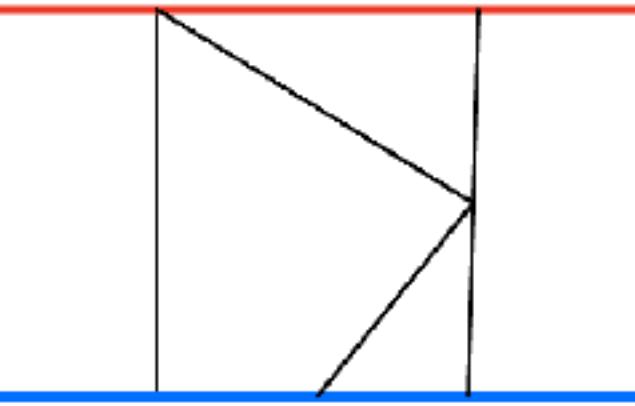
- Order-by-order solution in terms of (generalized) polylogarithms

$$\vec{J} = \sum_n \epsilon^n \vec{J}_n \quad \vec{J}_{n+1} = \sum_k A_k \int_0^x dz \left[\frac{d}{dz} \log(w_k(z)) \right] \vec{J}_n$$

- Boundary conditions: Regularity/scaling fixes most. Rest computed in the static limit $\sigma \rightarrow 1$

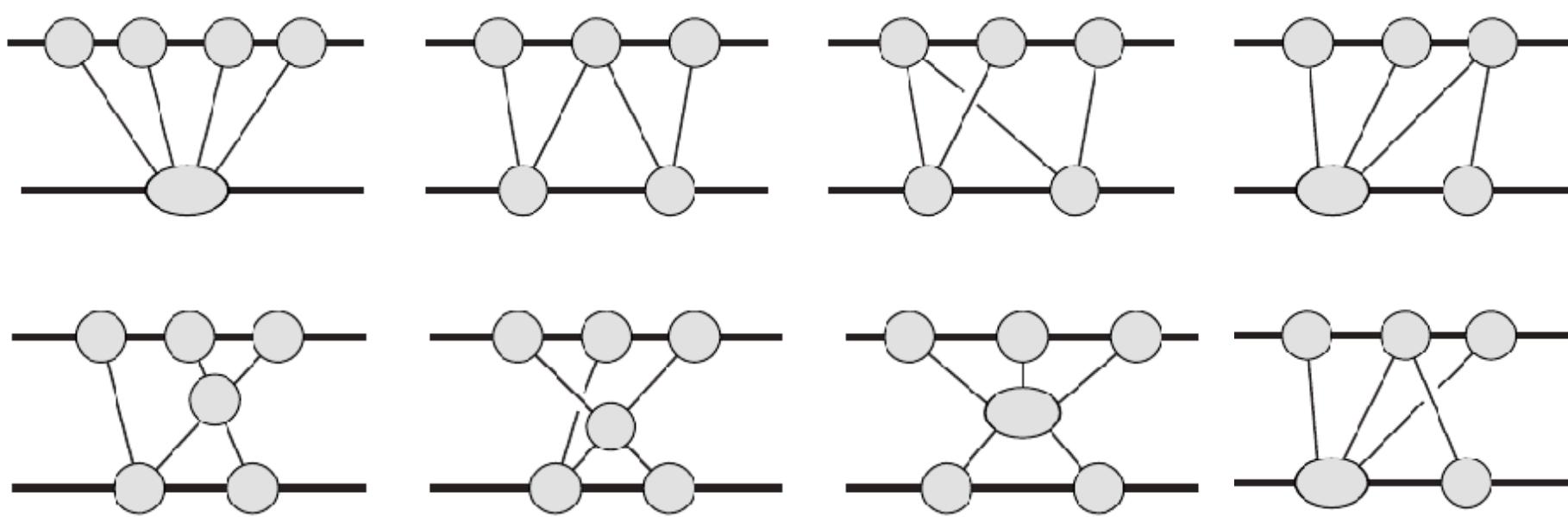
Integration – Elliptic sector

- Some cases: no canonical form


$$= \frac{1}{\epsilon^2} \frac{8}{\sigma + 1} K^2 \left(\frac{1 - \sigma}{1 + \sigma} \right) + \mathcal{O}(\epsilon^{-1})$$

- Elliptics only in the potential region \rightarrow hardest for integration
- Strategy 1:
 - Split amplitude $\mathcal{M} = \mathcal{M}_{\text{poly}} + \mathcal{M}_{\text{elliptic}}$
 - Solve $\mathcal{M}_{\text{poly}}$ through DE
 - For $\mathcal{M}_{\text{elliptic}}$ compute series to high orders and match to ansatz
- Strategy 2: Generalized kernels instead of $\frac{d \log(w_k)}{dx}$, [Dlapa et al.]

Integration – Elliptic sector



- Ansatz from contact integrals

$$\sim \left\{ K^2 \left(\frac{1-\sigma}{1+\sigma} \right), E \left(\frac{1-\sigma}{1+\sigma} \right) K \left(\frac{1-\sigma}{1+\sigma} \right), E^2 \left(\frac{1-\sigma}{1+\sigma} \right) \right\}$$

Not allowed to collapse any line!

- Expand integrals/amplitude in ν using DE and match

$$\mathcal{M}_{4,\text{elliptic}} \sim -\pi^2 \left(\frac{41}{16} + \frac{33601\nu^2}{3072} + \dots \# \nu^{400} \right) = r_4 \pi^2 + r_5 K \left(\frac{1-\sigma}{1+\sigma} \right)^2 + r_6 E \left(\frac{1-\sigma}{1+\sigma} \right) K \left(\frac{1-\sigma}{1+\sigma} \right) + r_7 E \left(\frac{1-\sigma}{1+\sigma} \right)^2$$

- 60 orders to fix r_i , 400 to check
- Avoids complicated integrals (elliptic polylogs) in intermediate steps
- (Pre-)Canonical form useful to make series expansion efficient

Classical scattering at $O(G^4)$ – Results

$$\mathcal{M}_4^{\text{cons}} = G^4 M^7 \nu^2 |\vec{q}| \pi^2 \left[\mathcal{M}_4^{\text{p}} + \nu \left(4 \mathcal{M}_4^{\text{t}} \log \left(\frac{\nu}{2} \right) + \mathcal{M}_4^{\pi^2} + \mathcal{M}_4^{\text{rem}} \right) \right] + \text{iterations}$$

$$\mathcal{M}_4^{\text{p}} = -\frac{35 (1 - 18\sigma^2 + 33\sigma^4)}{8 (\sigma^2 - 1)}$$

$$\mathcal{M}_4^{\text{t}} = r_1 + r_2 \log \left(\frac{\sigma+1}{2} \right) + r_3 \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}}$$

$$\mathcal{M}_4^{\pi^2} = r_4 \pi^2 + r_5 K\left(\frac{\sigma-1}{\sigma+1}\right) E\left(\frac{\sigma-1}{\sigma+1}\right) + r_6 K^2\left(\frac{\sigma-1}{\sigma+1}\right) + r_7 E^2\left(\frac{\sigma-1}{\sigma+1}\right)$$

$$\mathcal{M}_4^{\text{rem}} = r_8 + r_9 \log \left(\frac{\sigma+1}{2} \right) + r_{10} \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}} + r_{11} \log(\sigma)$$

$$+ r_{13} \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}} \log \left(\frac{\sigma+1}{2} \right) + r_{14} \frac{\operatorname{arccosh}^2(\sigma)}{\sigma^2 - 1}$$

$$+ r_{15} \operatorname{Li}_2\left(\frac{1-\sigma}{2}\right) + r_{16} \operatorname{Li}_2\left(\frac{1-\sigma}{1+\sigma}\right) + r_{17} \frac{1}{\sqrt{\sigma^2 - 1}} \left[\operatorname{Li}_2\left(-\sqrt{\frac{\sigma-1}{\sigma+1}}\right) - \operatorname{Li}_2\left(\sqrt{\frac{\sigma-1}{\sigma+1}}\right) \right]$$

Classical scattering at $O(G^4)$ – Results

$$\mathcal{M}_4^{\text{cons}} = G^4 M^7 \nu^2 |\vec{q}| \pi^2 \left[\mathcal{M}_4^{\text{p}} + \nu \left(4 \mathcal{M}_4^{\text{t}} \log\left(\frac{\nu}{2}\right) + \mathcal{M}_4^{\pi^2} + \mathcal{M}_4^{\text{rem}} \right) \right] + \boxed{\text{iterations}}$$

$$\mathcal{M}_4^{\text{p}} = -\frac{35 (1 - 18\sigma^2 + 33\sigma^4)}{8 (\sigma^2 - 1)}$$

$$\mathcal{M}_4^{\text{t}} = r_1 + r_2 \log\left(\frac{\sigma+1}{2}\right) + r_3 \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}}$$

$$\mathcal{M}_4^{\pi^2} = r_4 \pi^2 + r_5 K\left(\frac{\sigma-1}{\sigma+1}\right) E\left(\frac{\sigma-1}{\sigma+1}\right) + r_6 K^2\left(\frac{\sigma-1}{\sigma+1}\right) + r_7 E^2\left(\frac{\sigma-1}{\sigma+1}\right)$$

$$\mathcal{M}_4^{\text{rem}} = r_8 + r_9 \log\left(\frac{\sigma+1}{2}\right) + r_{10} \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}} + r_{11} \log(\sigma)$$

$$+ r_{13} \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}} \log\left(\frac{\sigma+1}{2}\right) + r_{14} \frac{\operatorname{arccosh}^2(\sigma)}{\sigma^2 - 1}$$

$$+ r_{15} \operatorname{Li}_2\left(\frac{1-\sigma}{2}\right) + r_{16} \operatorname{Li}_2\left(\frac{1-\sigma}{1+\sigma}\right) + r_{17} \frac{1}{\sqrt{\sigma^2 - 1}} \left[\operatorname{Li}_2\left(-\sqrt{\frac{\sigma-1}{\sigma+1}}\right) - \operatorname{Li}_2\left(\sqrt{\frac{\sigma-1}{\sigma+1}}\right) \right]$$

$$\int_{\vec{\ell}} \frac{\tilde{I}_{r,1}^4}{Z_1 Z_2 Z_3} + \int_{\vec{\ell}} \frac{\tilde{I}_{r,1}^2 \tilde{I}_{r,2}}{Z_1 Z_2} + \int_{\vec{\ell}} \frac{\tilde{I}_{r,1} \tilde{I}_{r,3}}{Z_1} + \int_{\vec{\ell}} \frac{\tilde{I}_{r,2}^2}{Z_1}$$

As predicted by amplitude-action relation
Highly non-trivial check!

Classical scattering at $O(G^4)$ – Results

Radial action at order G^4 $I_r^4(\vec{q}) \sim \chi^4$

$$\mathcal{M}_4^{\text{cons}} = G^4 M^7 \nu^2 |\vec{q}| \pi^2 \left[\mathcal{M}_4^{\text{p}} + \nu \left(4 \mathcal{M}_4^{\text{t}} \log \left(\frac{\nu}{2} \right) + \mathcal{M}_4^{\pi^2} + \mathcal{M}_4^{\text{rem}} \right) \right] + \text{iterations}$$

$$\mathcal{M}_4^{\text{p}} = -\frac{35 (1 - 18\sigma^2 + 33\sigma^4)}{8 (\sigma^2 - 1)}$$

$$\mathcal{M}_4^{\text{t}} = r_1 + r_2 \log \left(\frac{\sigma+1}{2} \right) + r_3 \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}}$$

$$\mathcal{M}_4^{\pi^2} = r_4 \pi^2 + r_5 K\left(\frac{\sigma-1}{\sigma+1}\right) E\left(\frac{\sigma-1}{\sigma+1}\right) + r_6 K^2\left(\frac{\sigma-1}{\sigma+1}\right) + r_7 E^2\left(\frac{\sigma-1}{\sigma+1}\right)$$

$$\mathcal{M}_4^{\text{rem}} = r_8 + r_9 \log \left(\frac{\sigma+1}{2} \right) + r_{10} \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}} + r_{11} \log(\sigma)$$

$$+ r_{13} \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}} \log \left(\frac{\sigma+1}{2} \right) + r_{14} \frac{\operatorname{arccosh}^2(\sigma)}{\sigma^2 - 1}$$

$$+ r_{15} \operatorname{Li}_2\left(\frac{1-\sigma}{2}\right) + r_{16} \operatorname{Li}_2\left(\frac{1-\sigma}{1+\sigma}\right) + r_{17} \frac{1}{\sqrt{\sigma^2 - 1}} \left[\operatorname{Li}_2\left(-\sqrt{\frac{\sigma-1}{\sigma+1}}\right) - \operatorname{Li}_2\left(\sqrt{\frac{\sigma-1}{\sigma+1}}\right) \right]$$

Classical scattering at $O(G^4)$ – Results

$$\mathcal{M}_4^{\text{cons}} = G^4 M^7 \nu^2 |\vec{q}| \pi^2 \left[\mathcal{M}_4^{\text{p}} + \nu \left(4 \mathcal{M}_4^{\text{t}} \log\left(\frac{\nu}{2}\right) + \mathcal{M}_4^{\pi^2} + \mathcal{M}_4^{\text{rem}} \right) \right] + \text{iterations}$$

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Elliptics, from ansatz

$$\mathcal{M}_4^{\text{rem}} = r_8 + r_9 \log\left(\frac{\sigma+1}{2}\right) + r_{10} \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}} + r_{11} \log(\sigma)$$

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Classical scattering at $O(G^4)$ – Results

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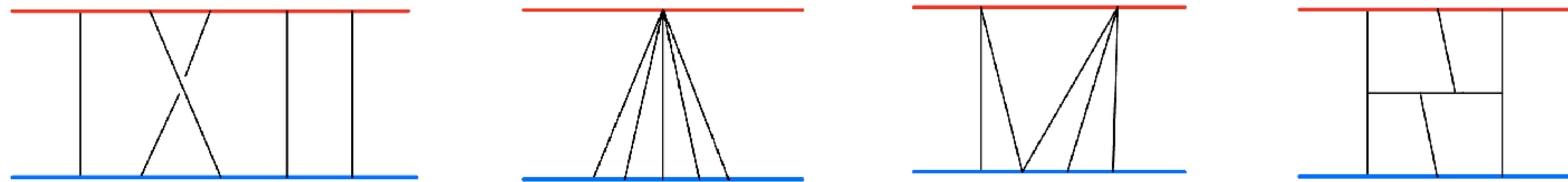
$$+ r_{13} \frac{\operatorname{arccosh}(\sigma)}{\sqrt{\sigma^2 - 1}} \log \left(\frac{\sigma+1}{2} \right) + r_{14} \frac{\operatorname{arccosh}^2(\sigma)}{\sigma^2 - 1}$$

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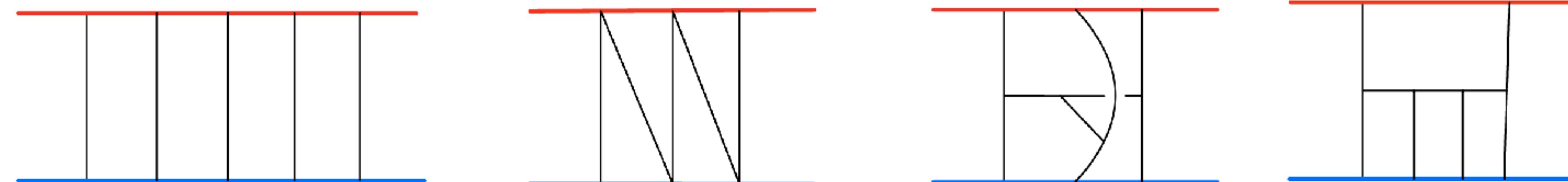
From explicit solution of DE

Classical scattering at $O(G^4)$ – Summary

- Classical scattering at G^4 from amplitudes approach very efficient
- Main improvement from importing integration techniques from collider physics (IBP, DE, Generalised unitarity)
- Results confirmed and extended to include radiation and spin [Dlaps et al., Jacobsen et al., Damgaard et. al.]
- Main bottleneck integration
 - Integral reduction
 - Evaluation of master integrals

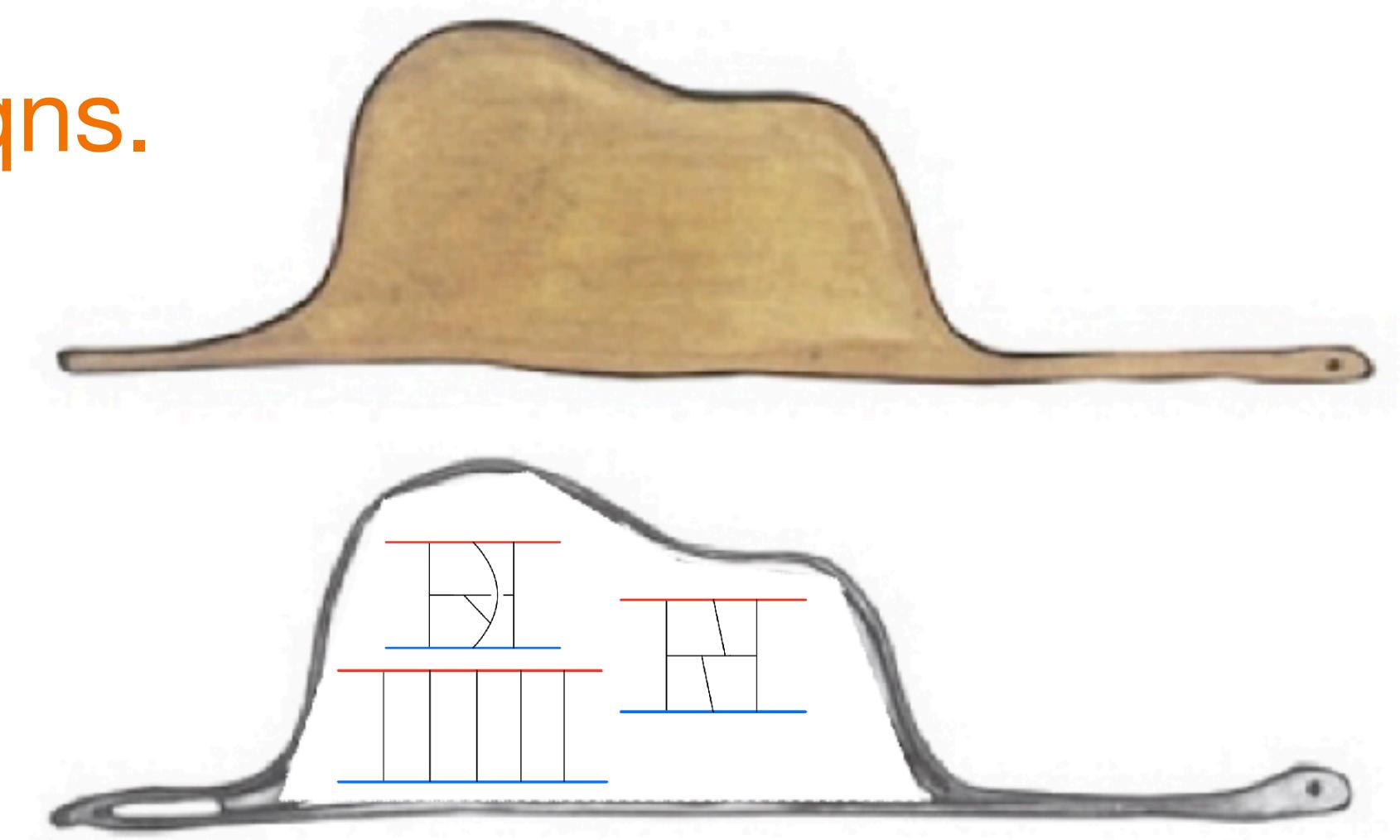


Towards Gravitational scattering at G^5



Amplitude computation

- Amplitude computation comes in parts
 - Constructing the integrand ✓ 51 cuts for 5PM, 6PM straightforward
 - Integral reduction ?
 - Evaluation of master integrals ?
- Integral reduction:
 - linear algebra problem, exponential growth in # eqns.
 - Intractable without major improvements
- Evaluating integrals:
 - DE's for almost all families
 - Series soon, analytic results harder



The 5PM problem is hard, don't try to swallow it whole!

Slicing the 5PM problem

- Can we eat our cake one piece at a time? Expansions?



$$\mathcal{M}_{\text{5PM}}(\nu, m_1, m_2)$$

- PN expansion $\nu \rightarrow 0$
 - Complicated topologies suppressed
 - Important for phenomenology
- High-energy expansion $\nu \rightarrow 1$
 - Less well understood, important conceptual questions
- Expanding loses information on functional structure

Slicing the 5PM problem

- Can we eat our cake one piece at a time? Expansions?



$$\mathcal{M}_{\text{5PM}}(\nu, m_1, m_2)$$

- Hierarchical limit (SF) $m_1 \ll m_2$ $\nu = \frac{\mu}{M} = m_1 m_2 / (m_1 + m_2)^2 \leq 1/4$

- Organization into gauge-invariant objects

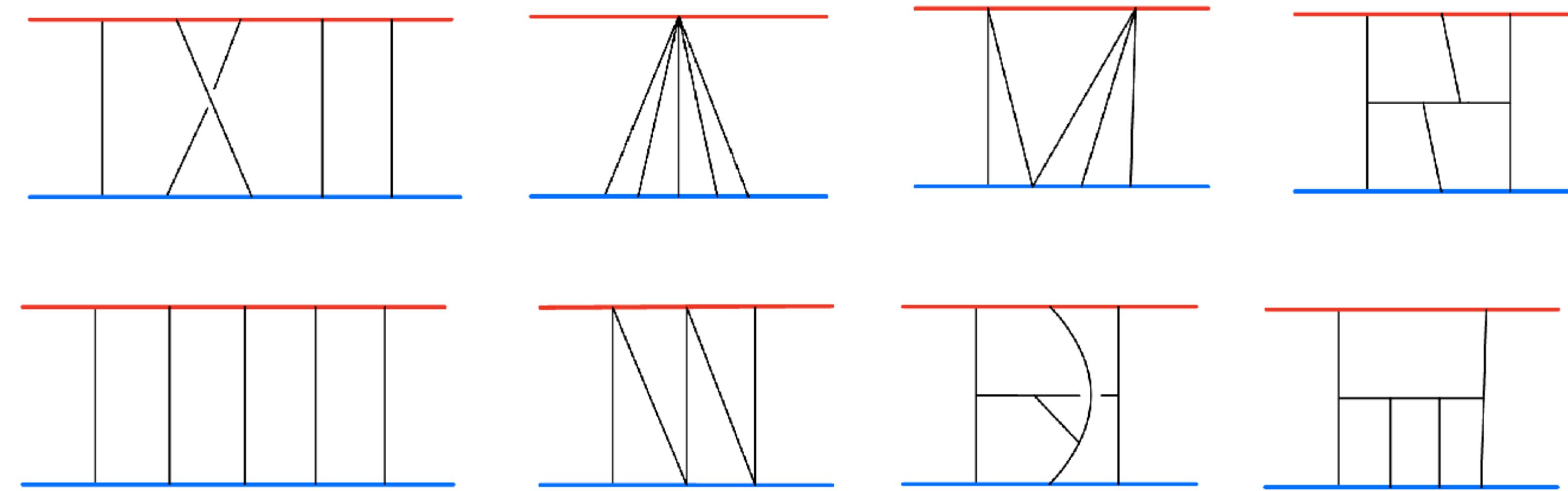
$$\mathcal{M}_{\text{5PM}} = \mathcal{M}_{\text{5PM}}^{\text{0SF}} + \nu \mathcal{M}_{\text{5PM}}^{\text{1SF}} + \nu^2 \mathcal{M}_{\text{5PM}}^{\text{2SF}} \leftarrow \text{trivial from amplitudes}$$

- Useful expansion for equal-mass case! Similar to QCD $\frac{1}{N_c} = \frac{1}{3}$
 - Complicated integrals suppressed

Slicing the 5PM problem

- Can we eat our cake one piece at a time? Expansions?

$\mathcal{M}_{5PM} =$



+ $\mathcal{O}(10^4)$ more

Slicing the 5PM problem

- Can we eat our cake one piece at a time? Expansions?

$$\mathcal{M}_{5\text{PM}}^{0\text{SF}} =$$

+ $\mathcal{O}(10^4)$ more

Slicing the 5PM problem

- Can we eat our cake one piece at a time? Expansions?

$$\mathcal{M}_{\text{5PM}}^{\text{1SF}} =$$

+ $\mathcal{O}(10^4)$ more

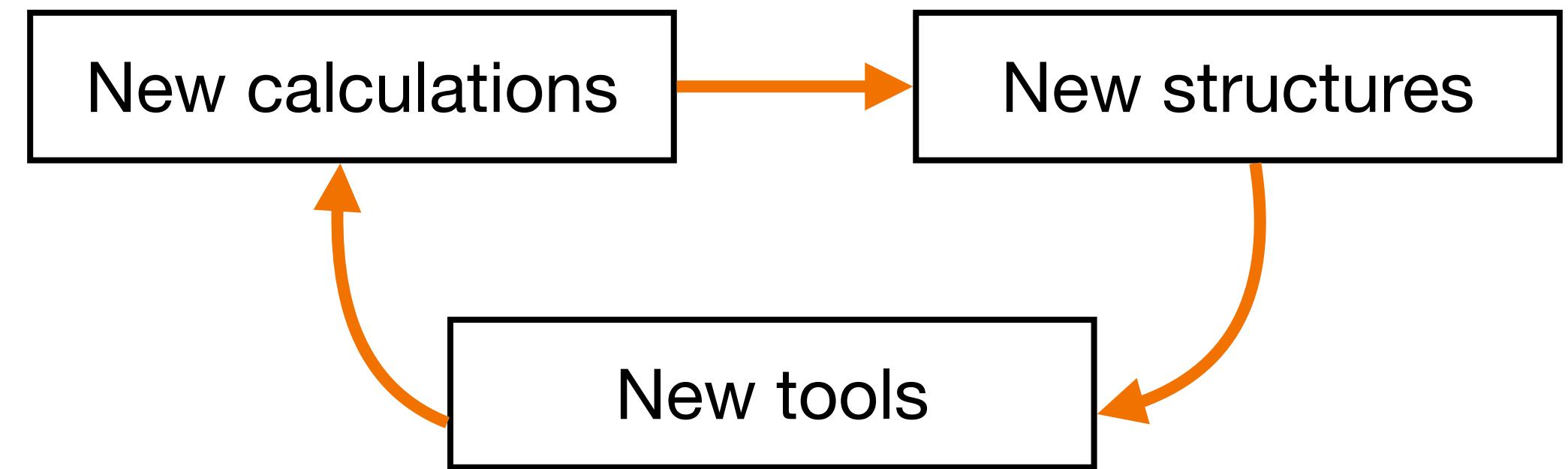
Slicing the 5PM problem

- Can we eat our cake one piece at a time? Expansions?

$$\mathcal{M}_{\text{5PM}}^{\text{2SF}} =$$
$$+ \mathcal{O}(10^4) \text{ more}$$

Slicing the 5PM problem

- Work on 1SF in progress
- Use a simpler model without approximations:
 - Maximal SUSY, scalar toy models,...
 - Electrodynamics
- Main criteria:
 - Sizeable overlap with GR
 - Significantly more complicated than 4PM
 - Real world system, applications to phenomenology



Electrodynamics checks all the boxes

Slicing the 5PM problem

- Can we eat our cake one piece at a time?

$$\mathcal{M}_{5\text{PL}}^{\text{QED}} =$$

As complicated as in GR! 2SF graphs + $\mathcal{O}(10^4)$ more

Classical scattering at $O(\alpha^5)$

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}^2 - \frac{1}{2\xi}(\partial_\mu A^\mu)^2 + \sum_{i=1}^2 [|D_\mu \phi_i|^2 - m_i^2 |\phi_i|^2]$$

- QED integrand trivial: ~ 1000 Feynman diagrams
- Deep expansion in the classical limit $\mathcal{M}_4 \sim \frac{1}{\hbar^5} + \dots + \frac{1}{\hbar}$
- Integral reduction: 10^6 integrals $\rightarrow 1107$ masters, 23 families
- For the integration integral reduction use FIRE+LiteRed:
 - Choosing better basis of master integrals [Smirnov; Usovitch]
 - Removing redundant equations

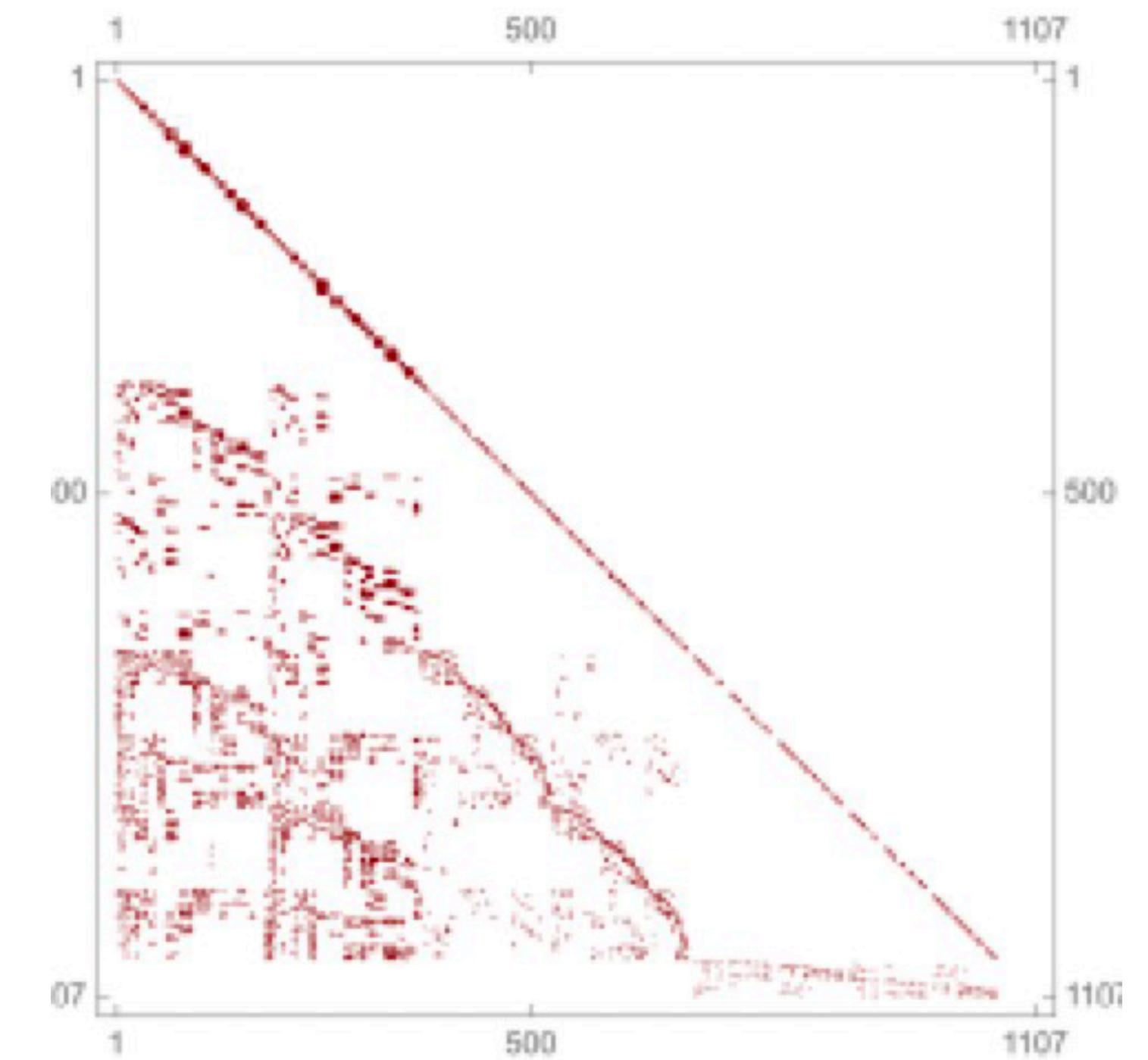
Classical scattering at $O(\alpha^5)$ – Integration

- No elliptic integrals for QED
- Differential equations in **canonical form**

$$\partial_x \vec{I} = \epsilon \sum_{k,n} f_k^n(x) A_{k,n} \vec{I}, \quad r = |\vec{I}| = 1107,$$

- In terms of **cyclotomic** kernels

$$f_n^k(x) = \frac{x^k}{\Phi_n(x)}, \quad \Phi_{1,2} = x \pm 1, \Phi_4 = 1 + x^2, \Phi_{3,6} = 1 \pm x + x^2$$



Classical scattering at $O(\alpha^5)$

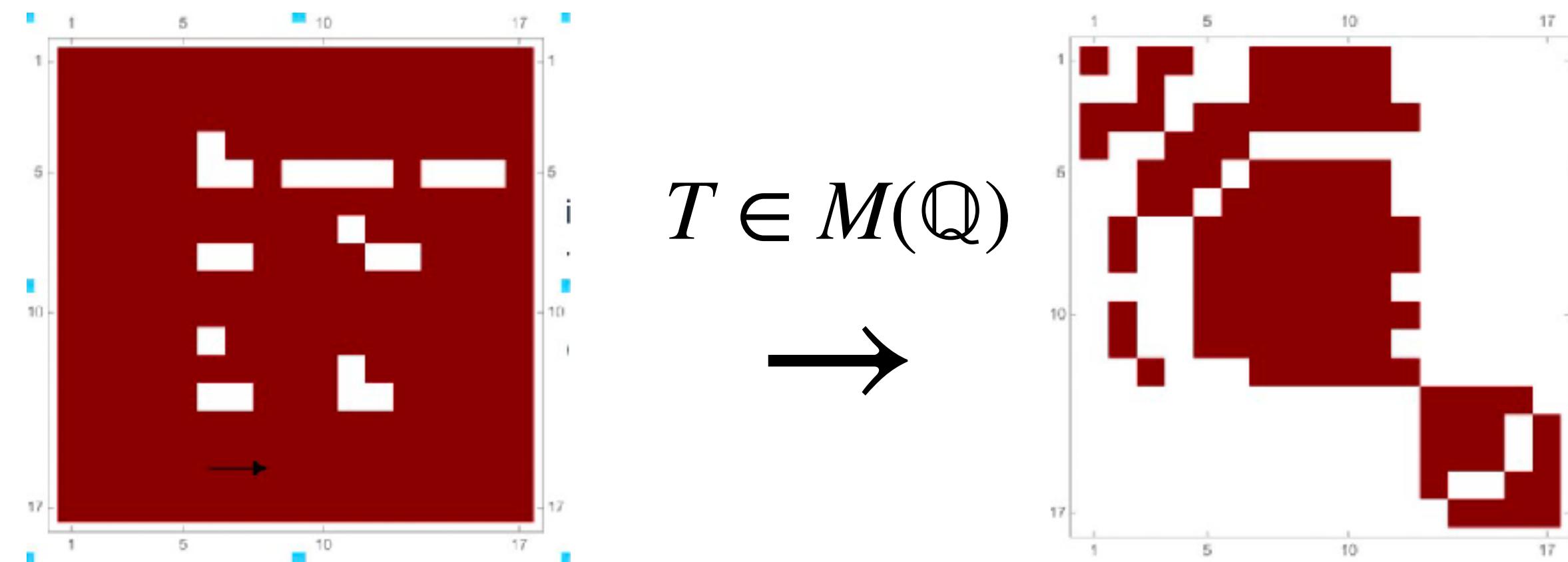
- Cyclotomic harmonic polylogs introduced in [Ablinger, Blumlein, Schneider 2011]
 - Manifestly real $C_4^0(x) = \int \frac{dx}{1+x^2} = \frac{i}{2} \int dx \left[\frac{1}{x+i} - \frac{1}{x-i} \right] = \frac{i}{2}(\log(x+i) - \log(x-i)) = \arctan(x)$
 - Shuffle algebra (minimal basis)
 - Integration and differential rules
 - Series expansion & numerics
- Special combinations (amplitude) in terms of real Li's

$$C_{400}^{100}(x) - C_{000}^{200}(x) = -\frac{\text{Li}_3(x^6)}{18} + \frac{1}{6} \text{Li}_2(x^6) \log(x) + \frac{4 \text{Li}_3(x^3)}{9} - \frac{2}{3} \text{Li}_2(x^3) \log(x) + \frac{\text{Li}_3(x^2)}{2} - \frac{1}{2} \text{Li}_2(1-x^2) \log(x) - 4 \text{Li}_3(x) - 2 \text{Li}_2(-x) \log(x) - \log(1-x^2) \log^2(x) + \frac{1}{4} \pi^2 \log(x) + \frac{707}{32} \zeta_3$$

Classical scattering at $O(\alpha^5)$ – Integration

- DE has additional structure: sparse, top sectors don't talk to bottom
- Canonical DE invariant under rational transformation: factorize

$$\epsilon \sum_{k,n} f_k^n(x) A_{k,n} \rightarrow \epsilon \sum_{k,n} f_k^n(x) A'_{k,n}$$



- Bonus relations between integrals, linked to special structure of eikonal integrals

Classical scattering at $O(\alpha^5)$ – Results

$$\chi_{\text{pot}}^{\text{5PL}} = \frac{\alpha^5(m_1 + m_2)^4}{30J^5E^4(\sigma^2 - 1)^{5/2}} \times \left[r_0^{(0)} + \sum_{k=1}^{12} \left(\nu r_k^{(1)} + \nu^2 r_k^{(2)} \right) f_k \right]$$

Classical scattering at $O(\alpha^5)$ – Results

Fifth post-Lorentzian order

$$\chi_{\text{pot}}^{\text{5PL}} = \frac{\alpha^5 (m_1 + m_2)^4}{30 J^5 E^4 (\sigma^2 - 1)^{5/2}} \times \left[r_0^{(0)} + \sum_{k=1}^{12} \left(\nu r_k^{(1)} + \nu^2 r_k^{(2)} \right) f_k \right]$$

Classical scattering at $O(\alpha^5)$ – Results

$$\chi_{\text{pot}}^{\text{5PL}} = \frac{\alpha^5(m_1 + m_2)^4}{30J^5E^4(\sigma^2 - 1)^{5/2}} \times \left[r_0^{(0)} + \sum_{k=1}^{12} \left(\nu r_k^{(1)} + \nu^2 r_k^{(2)} \right) f_k \right]$$

Contribution from potential
modes well-defined! No tail!

Classical scattering at $O(\alpha^5)$ – Results

$$\chi_{\text{pot}}^{\text{5PL}} = \frac{\alpha^5(m_1 + m_2)^4}{30J^5E^4(\sigma^2 - 1)^{5/2}} \times \left[r_0^{(0)} + \sum_{k=1}^{12} \left(\nu r_k^{(1)} + \nu^2 r_k^{(2)} \right) f_k \right]$$

Mass polynomiality

Second order E&M self-force!

Classical scattering at $O(\alpha^5)$ – Results

$$\begin{aligned}
r_9^{(1)} &= r_{12}^{(1)} = 240(\sigma^2 - 1)^2, \\
r_{11}^{(1)} &= 120(\sigma^2 - 1)(\sigma^2 + 2\sigma - 1), \\
r_6^{(1)} &= r_7^{(1)} = r_{10}^{(1)} = 0, \\
r_1^{(2)} &= \frac{405\sigma(15-44\sigma^2)}{16(1-4\sigma^2)^2} - \frac{15(10\sigma^2+2\sigma-3)}{\sigma^3} \\
&\quad + \frac{-2048\sigma^7+6656\sigma^6+17872\sigma^5+20000\sigma^4}{16} \\
&\quad + \frac{-7740\sigma^3-22560\sigma^2-6635\sigma-2080}{16}, \\
r_2^{(2)} &= \sqrt{\sigma^2-1} \left[\frac{45(1232\sigma^4-1168\sigma^2+287)}{16(4\sigma^2-1)^3} \right. \\
&\quad \left. + \frac{30(20\sigma^3-9\sigma^2-4\sigma+3)}{\sigma^4} \right. \\
&\quad \left. + \frac{5}{16}(1776\sigma^4+8192\sigma^3+10820\sigma^2+11776\sigma+3223) \right], \\
r_3^{(2)} &= -\frac{30(16\sigma^4+36\sigma^3-11\sigma^2-6\sigma+3)}{\sigma^5} \\
&\quad + 20(212\sigma^3+350\sigma^2+328\sigma+319), \\
r_4^{(2)} &= \frac{2880(\sigma+1)(3\sigma+1)}{\sqrt{\sigma^2-1}}, \\
r_6^{(2)} &= 480(\sigma^2-1)^{3/2}(2\sigma^2-1), \\
r_7^{(2)} &= 45\sigma(\sigma^2-1)^{5/2}, \\
r_9^{(2)} &= -480(\sigma^2-1)(\sigma^2-\sigma-1), \\
r_{10}^{(2)} &= -135(\sigma^2-1)^2, \\
r_{12}^{(2)} &= -480(\sigma^2-1)(\sigma^2-2\sigma-1), \\
r_5^{(2)} &= r_8^{(2)} = r_{11}^{(2)} = 0.
\end{aligned}
\tag{14}$$

Poles at $\sigma = 0, \pm 1/2, \pm 1$

Outside of the scattering region $1 < \sigma$
Implications for bound state?

$$\left[r_0^{(0)} + \sum_{k=1}^{12} \left(\nu r_k^{(1)} + \nu^2 r_k^{(2)} \right) f_k \right]$$

Rational coefficients

Classical scattering at $O(\alpha^5)$ – Results

$$\begin{aligned}
 f_1 &= 1, \quad f_2 = C_0^0(x), \quad f_3 = C_{0,0}^{0,0}(x), \quad f_4 = C_{0,0,0}^{0,0,0}(x), \\
 f_5 &= -C_{1,0}^{0,0}(x) + C_{2,0}^{0,0}(x) + \frac{\pi^2}{4}, \\
 f_6 &= -C_{2,0}^{0,0}(x) + C_{4,0}^{1,0}(x) - \frac{\pi^2}{16}, \\
 f_7 &= C_{3,0}^{0,0}(x) + 2C_{3,0}^{1,0}(x) + C_{6,0}^{0,0}(x) - 2C_{6,0}^{1,0}(x) + \frac{\pi^2}{6}, \\
 f_8 &= -C_{0,1,0}^{0,0,0}(x) + C_{0,2,0}^{0,0,0}(x) + \frac{\pi^2}{4}C_0^0(x) + \frac{7\zeta_3}{2}, \\
 f_9 &= -C_{0,2,0}^{0,0,0}(x) + C_{0,4,0}^{0,1,0}(x) - \frac{\pi^2}{16}C_0^0(x) - \frac{21\zeta_3}{16}, \\
 f_{10} &= C_{0,3,0}^{0,0,0}(x) + 2C_{0,3,0}^{0,1,0}(x) + C_{0,6,0}^{0,0,0}(x) - 2C_{0,6,0}^{0,1,0}(x) \\
 &\quad + \frac{1}{6}\pi^2C_0^0(x) + \frac{28\zeta_3}{9}, \\
 f_{11} &= -C_{1,0,0}^{0,0,0}(x) + C_{2,0,0}^{0,0,0}(x) - \frac{7\zeta_3}{4}, \\
 f_{12} &= -C_{2,0,0}^{0,0,0}(x) + C_{4,0,0}^{1,0,0}(x) + \frac{21\zeta_3}{32}. \tag{15}
 \end{aligned}$$

Transcendental functions

$$\left[r_0^{(0)} + \sum_{k=1}^{12} \left(\nu r_k^{(1)} + \nu^2 r_k^{(2)} \right) f_k \right]$$

Functions are special:

$$\text{No } \zeta\text{-values } f_k = \sum_{n,r} \log^r(1-x) a_n^r (1-x)^n, \quad a_n^r \in \mathbb{Q}$$

Only specific contributions of indices (symbology)

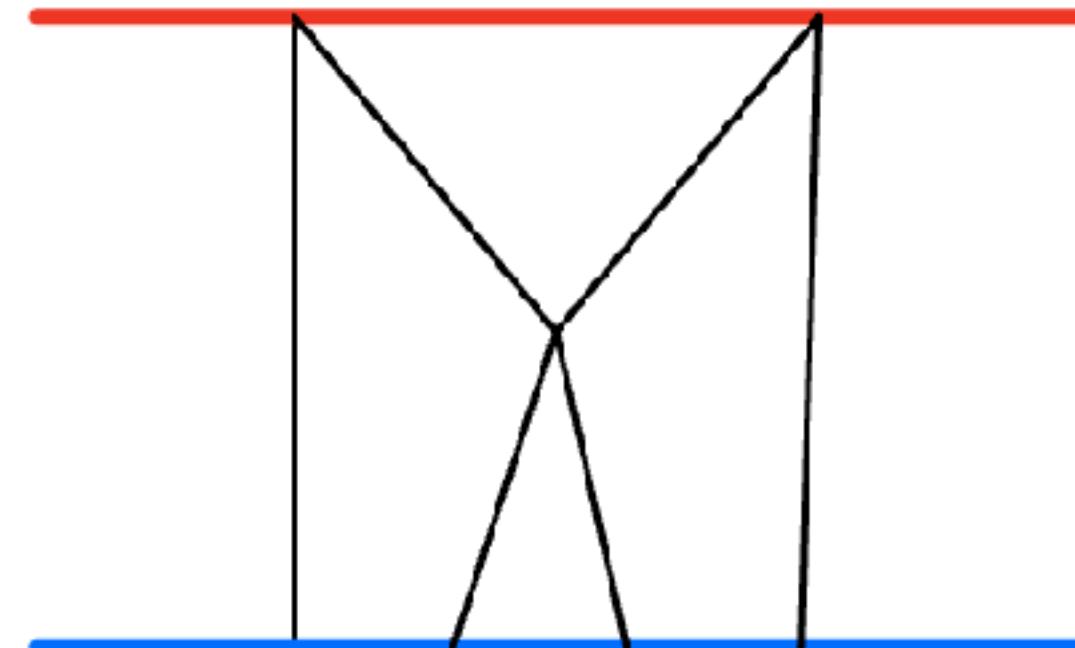
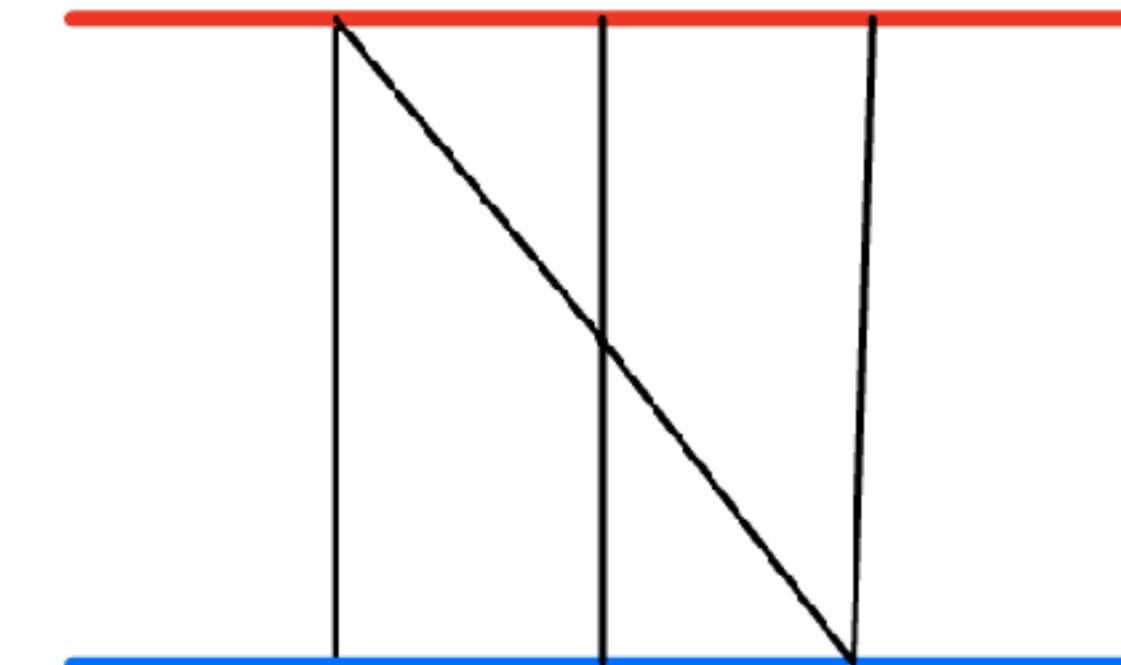
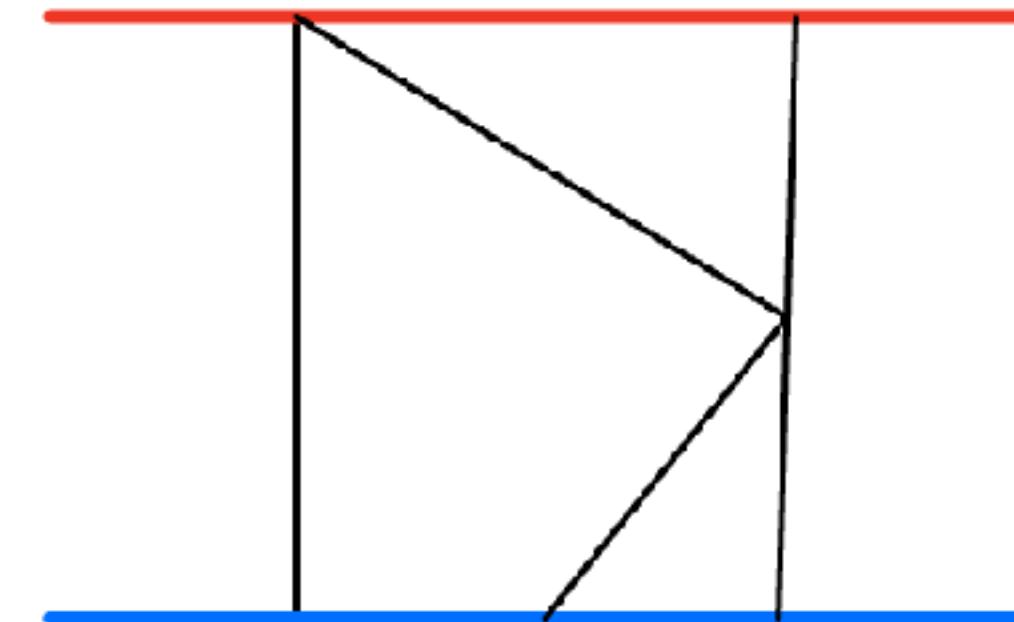
Alternative form: polylogs with real arguments

Classical scattering at $O(\alpha^5)$ – Lessons

- Computations at fifth order in perturbation theory are possible!
- E&M scattering at large impact parameter – heavy ion scattering
- First glimpse at the function space, organization in terms of CHPL useful
- Understand which parts of the computation are hardest → 2SF graphs
- QED is an useful playground for bootstrap ideas
 - High post-Coulombian orders from Fokker action
 - Simple space of transcendental function
 - Additional relation to energy loss
- Performed lots of checks of the result

Classical scattering at $O(G^5)$ – Status

- QED was surprisingly simple – only CHPL's
- For GR likely functions beyond polylogs and E, K
- Integrals more complicated than amplitude \rightarrow series + ansatz



$$\sim \left\{ K^2 \left(\frac{1-\sigma}{1+\sigma} \right), E \left(\frac{1-\sigma}{1+\sigma} \right) K \left(\frac{1-\sigma}{1+\sigma} \right), E^2 \left(\frac{1-\sigma}{1+\sigma} \right) \right\}$$

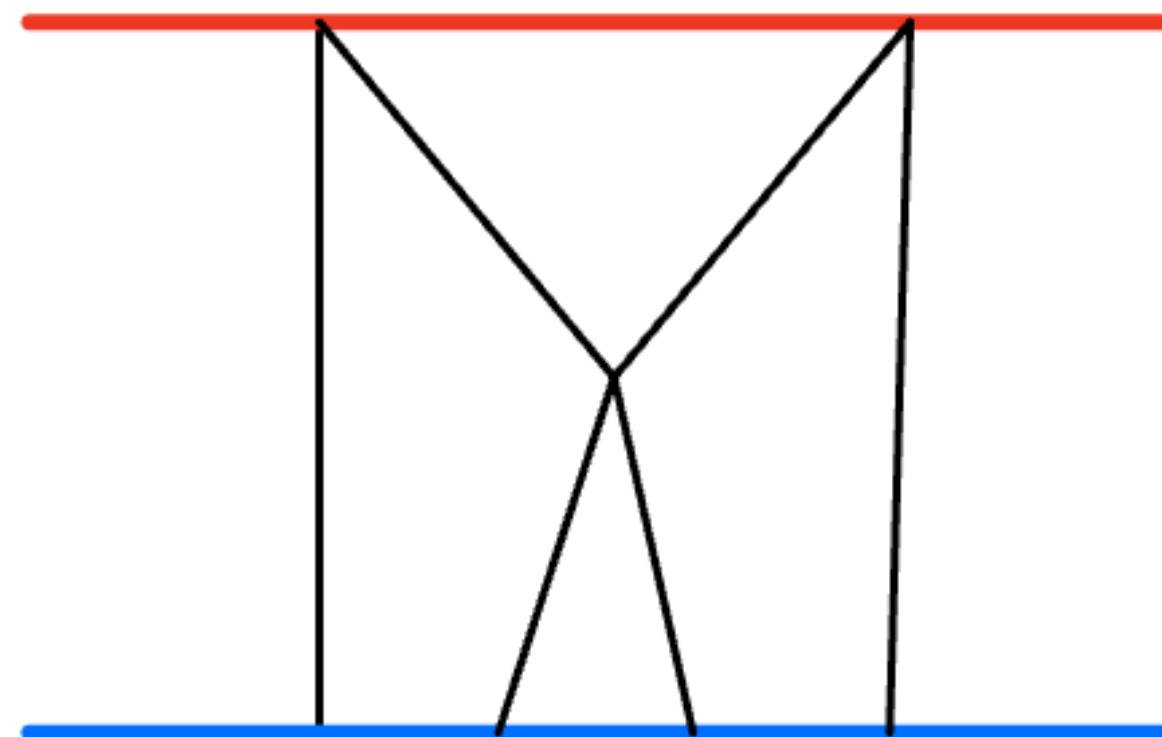
$$\sim \left\{ ? \right\}$$

Classical scattering at $O(G^5)$ – Status

Example 1 (1SF)

$$d \begin{pmatrix} f_1 \\ \vdots \\ f_5 \end{pmatrix} = \epsilon \sum_k d \log(w_k) A_k \begin{pmatrix} f_1 \\ \vdots \\ f_5 \end{pmatrix}, \quad w_k \in \{x, 1 \pm x, 1 + x^2\}$$

Same as at 3 loop

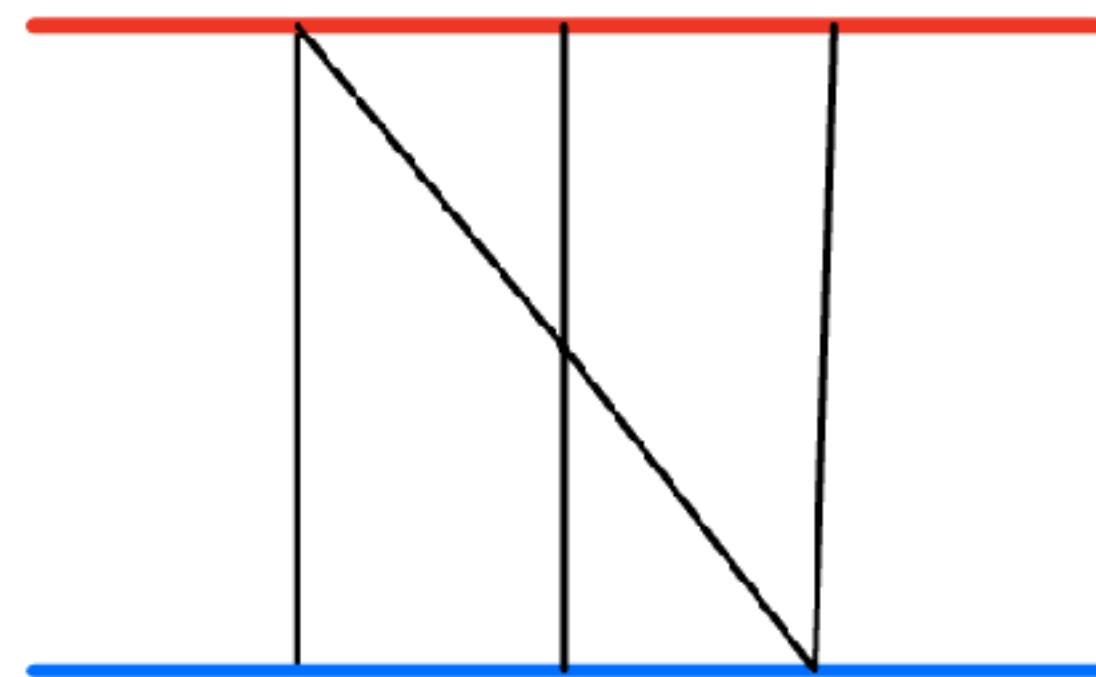


evaluates to CHPL's !

Classical scattering at $O(G^5)$ – Status

Example 2 (2SF)

$$d \begin{pmatrix} f_1 \\ \vdots \\ f_9 \end{pmatrix} = \sum_k d \log(w_k) A_k \begin{pmatrix} f_1 \\ \vdots \\ f_9 \end{pmatrix}, \quad w_k = \{x, 1 \pm x, 1 \pm x + x^2, 1 \pm 6x + x^2, 1 + 6x^2 + x^4\}$$



Beyond cyclotomic alphabet

$$x = 3 - 2\sqrt{2} = 0.171\dots \in [0,1]$$

$$x = -i\sqrt{3 - 2\sqrt{2}} \rightarrow \gamma = i$$

Inside the scattering region $x \in [0,1]!$

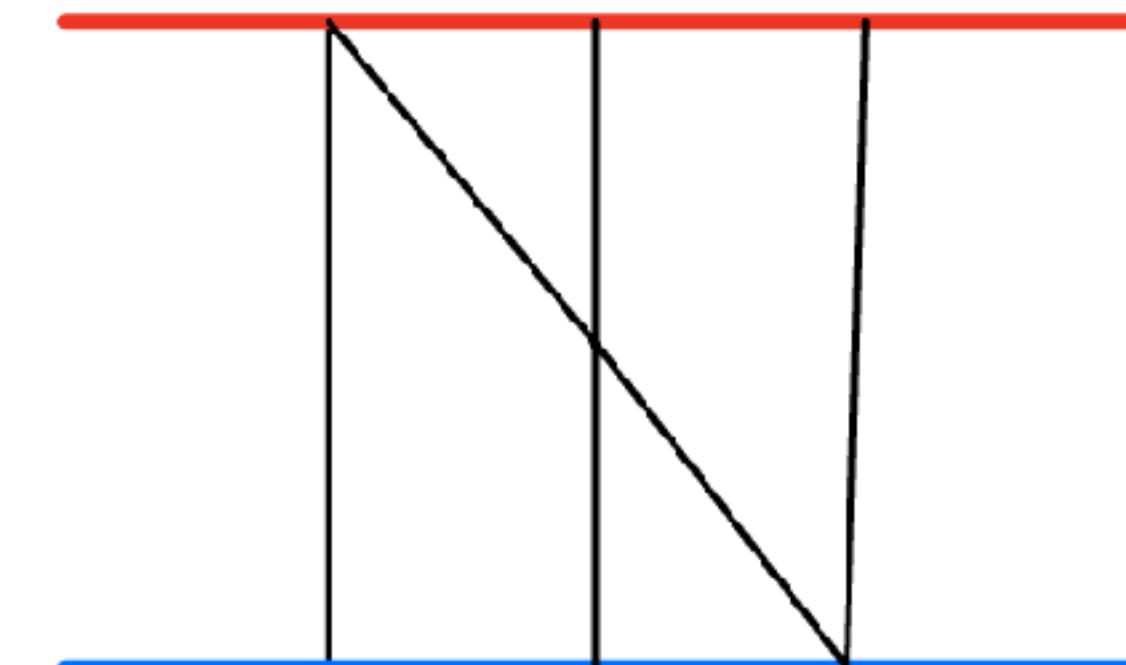
A_k eigenvalues $1/2 + a\epsilon, a \in \mathbb{Z}$
suggest elliptic (or more complicated)
integrals

Classical scattering at $O(G^5)$ – Todo

- Equivalently solve 9th order ODE – hard

$$d \begin{pmatrix} f_1 \\ \vdots \\ f_9 \end{pmatrix} = \epsilon \sum_k d \log(w_k) A_k \begin{pmatrix} f_1 \\ \vdots \\ f_9 \end{pmatrix}, \quad 0 = \sum_{k=0}^9 p_k(x, \epsilon) \frac{d^k}{dx^k} f_1$$

- Needs detailed study of the geometry involved
- Might need to resort to series expansion



Conclusion

- Amplitudes-based program to extract GR observables to high orders in perturbation theory
- Higher precision, more loops
- Scattering in QED at 4 loops important case study
 - Proof of principle computation
 - Large overlap with GR computation ~25% of the master integrals, including 2SF integrals
 - Identified bottlenecks and improved setup in integral reduction
- Progress towards scattering at G^5
 - Integrand constructed ✓
 - Differential equations for all but a few families (hopefully completed soon) ✓
 - Integral reduction challenging ? New ideas from collider physics could help
- Eventually reconnect with the bound problem

Optimistic on near-term progress – 5PM hard but doable!