

# Testing the relic neutrino decay solution to the EDGES anomaly and ARCADE2 excess with neutrino oscillation experiments

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Based on: arXiv: 2311:XXXXX, coming soon!

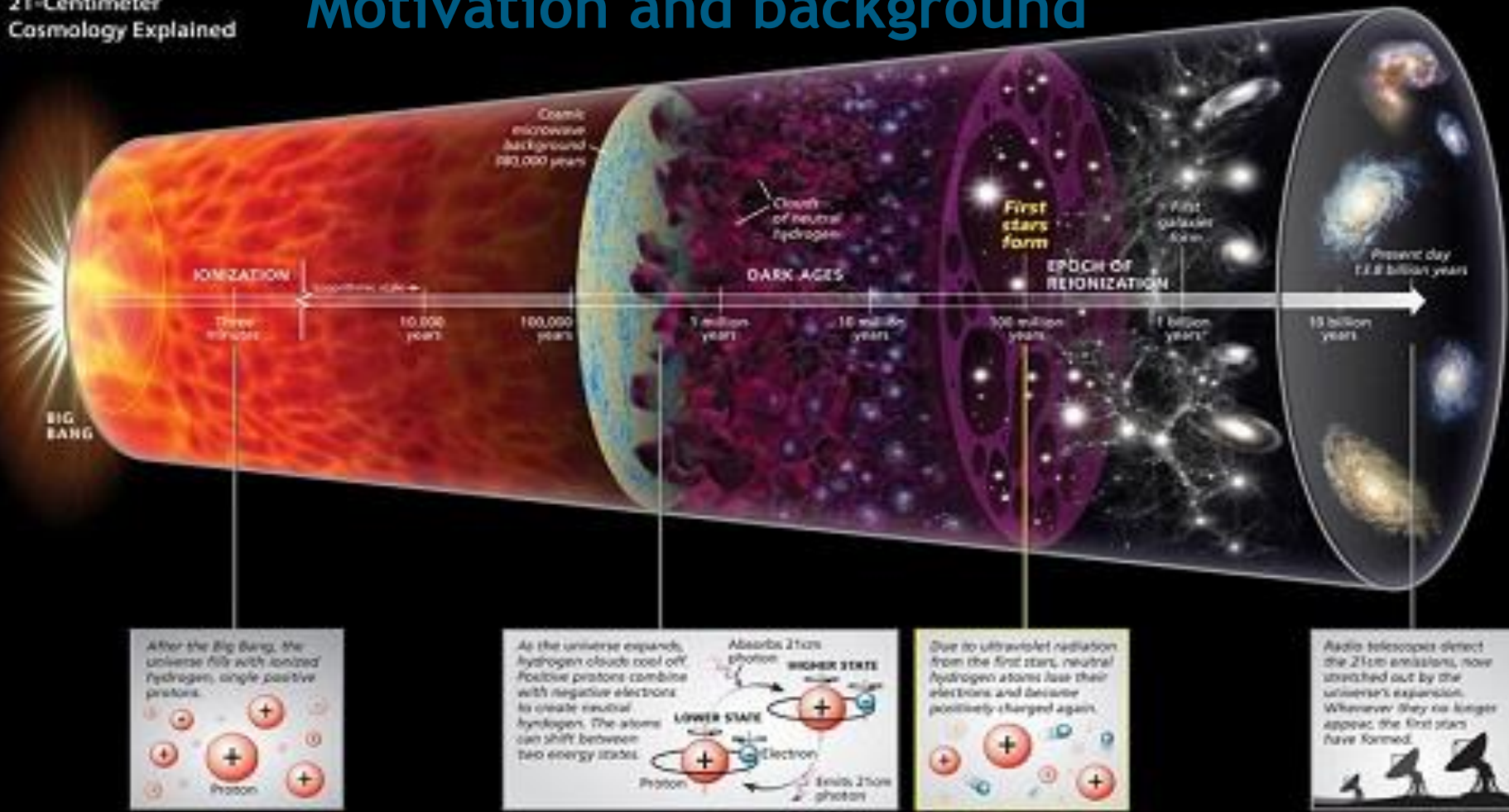
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University of  
**Southampton**

# Motivation and background



## Cosmological Puzzles

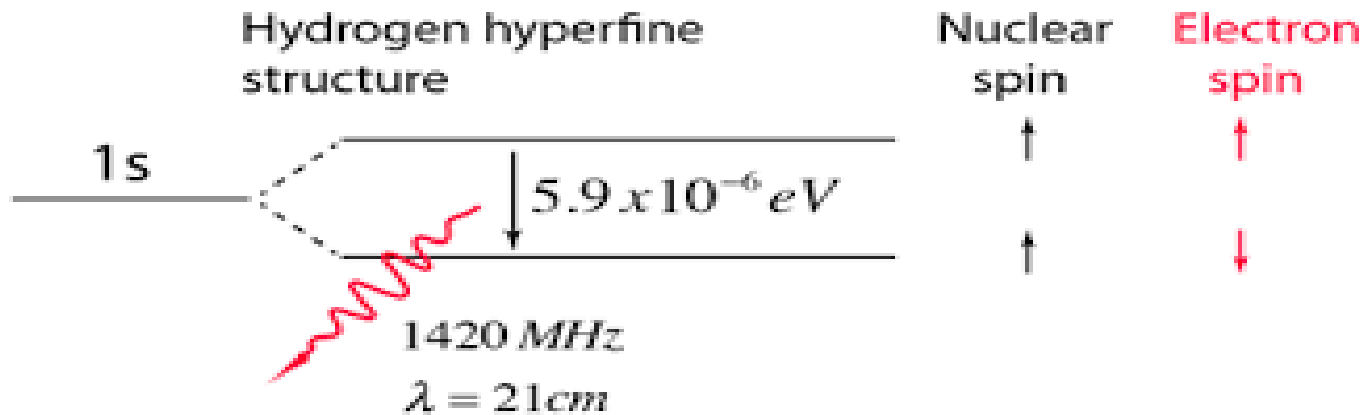
1. Matter-Antimatter asymmetry
2. Dark Matter
3. Neutrino masses



Motivation for BSM Physics

Cosmological Observations:  
a powerful investigative tool<sup>2</sup>

# 21 cm Radiation



**21 cm line of hydrogen arises from the hyperfine splitting of the 1S ground state of Hydrogen due to the spin flip transition of electron.**

$$\Delta E = 5.9 \times 10^{-6} \text{ eV} = 1420 \text{ MHz} = 21 \text{ cm}$$

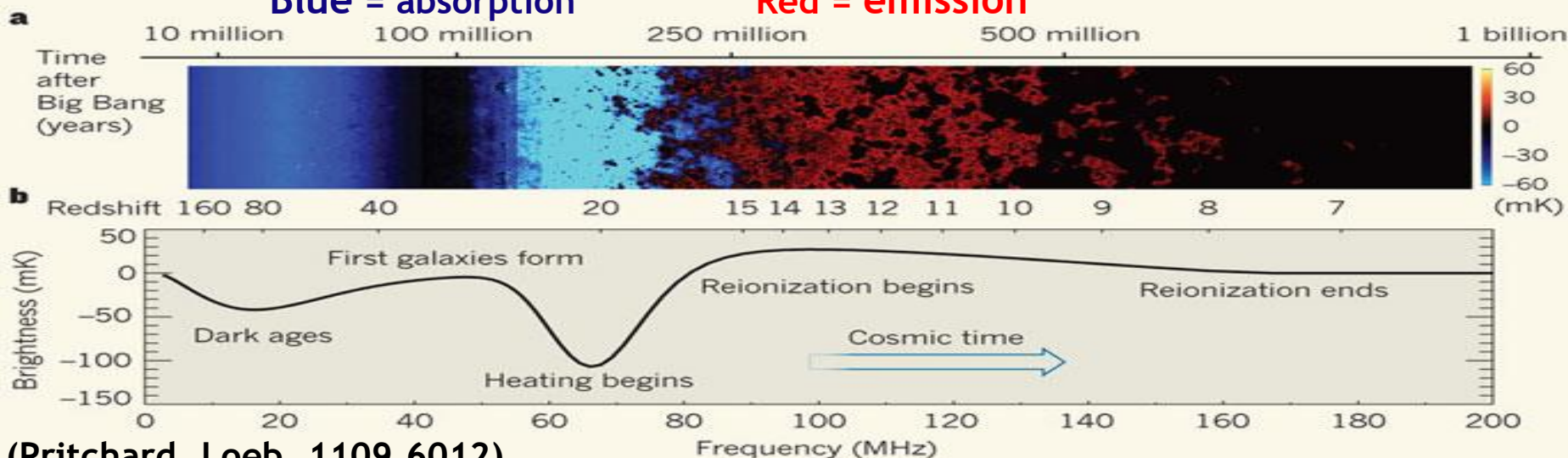
Predicted by Van De Hulst in 1944, Discovered by Ewen and Purcell in 1951

First time revealed that Milky Way has spiral structure.

# 21 cm Cosmology

Blue = absorption

Red = emission



(Pritchard, Loeb, 1109.6012)

$$T_{21}(z) \simeq 23 \text{ mK} (1 + \delta_B) x_{H_I}(z) \left( \frac{\Omega_B h^2}{0.02} \right) \left[ \left( \frac{0.15}{\Omega_m h^2} \right) \left( \frac{1+z}{10} \right) \right]^{1/2} \left[ 1 - \frac{T_\gamma(z)}{T_S(z)} \right]$$

- ❑ A powerful investigative astrophysical tool
- ❑ Maps Hydrogen gas intervening between a source and the observer.
- ❑ In the case of the global cosmological signal the source is the primordial radiation itself interacting with Hydrogen gas at all redshifts below  $z=1100$  (recombination)

# EDGES anomaly

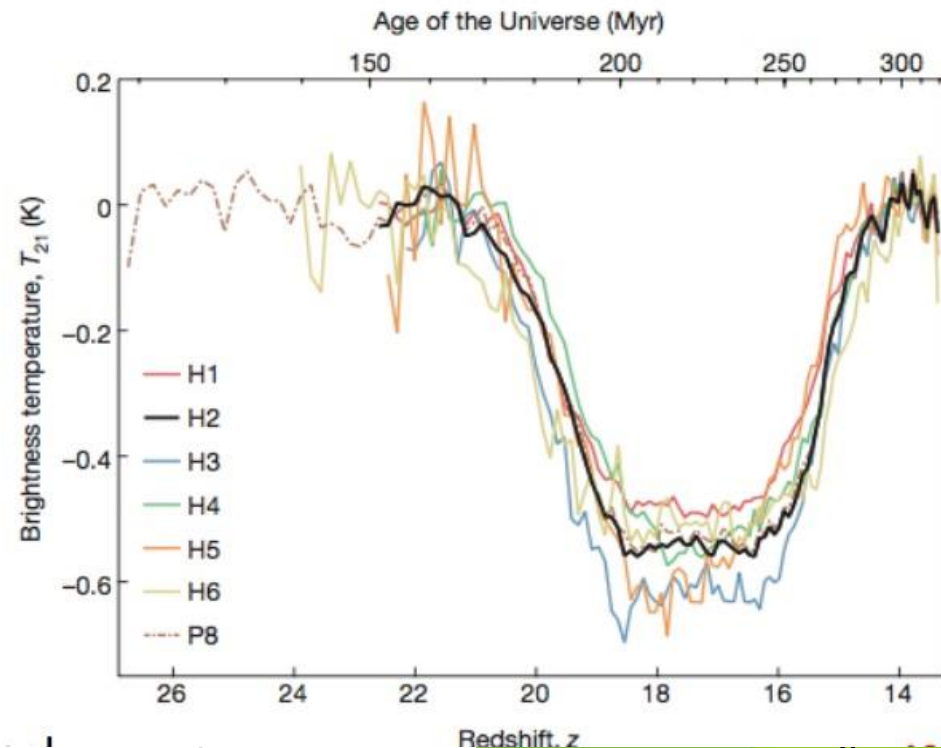


- ❑ EDGES measured a 21 cm (global) signal at a frequency  $\sim 70\text{MHz}$  and  $z_E \sim 17.2$
- ❑ It finds an absorption signal that is double compared to the expected one in a cosmological standard model

- ❑ A doubled signal can be explained:  
(1) in terms of a colder gas (earlier decoupling?)  
(2) additional non-thermal background that increases  $T_\gamma$

The spin temperature is related to the gas temperature:

$$\left(1 - \frac{T_\gamma}{T_S}\right) \simeq \frac{x_c + x_\alpha}{1 + x_c + x_\alpha} \left(1 - \frac{T_\gamma}{T_{\text{gas}}}\right)$$



(Bowman et al., Nature 555 (2018) 7694, 67-70)

# ARCADE anomaly



- ❑ ARCADE2 has measured the absolute temperature of the sky in the radio frequency range 3–90 GHz .
- ❑ It found an extragalactic excess, in addition to the CMB, at the low frequency range, approx. b/w 3 and 8 GHz.
- ❑ ARCADE 2 measurements are made at  $z=0$ .

❑ ARCADE 2 expt. found  $T_0 = 2.729$  K

Table 1. ARCADE 2 measurements of the effective temperature

$i$	$\nu_i$ (GHz)	$E_i$ ( $10^{-5}$ eV)	$T_{\gamma 0}^i$ (K)	$\bar{T}_{\text{rbe}}^i$ (mK)	$\delta \bar{T}_{\text{rbe}}^i$ (mK)
1	3.20	1.36	2.792	63	10
2	3.41	1.41	2.771	42	9
3	7.97	3.30	2.765	36	14
4	8.33	3.44	2.741	12	16
5	9.72	4.02	2.732	3	6
6	10.49	4.34	2.732	3	6
7	29.5	12.2	2.529	-200	155
8	31	12.82	2.573	-156	76
9	90	3.72	2.706	-23	19

(Astrophys.J. 734 (2011) 5)

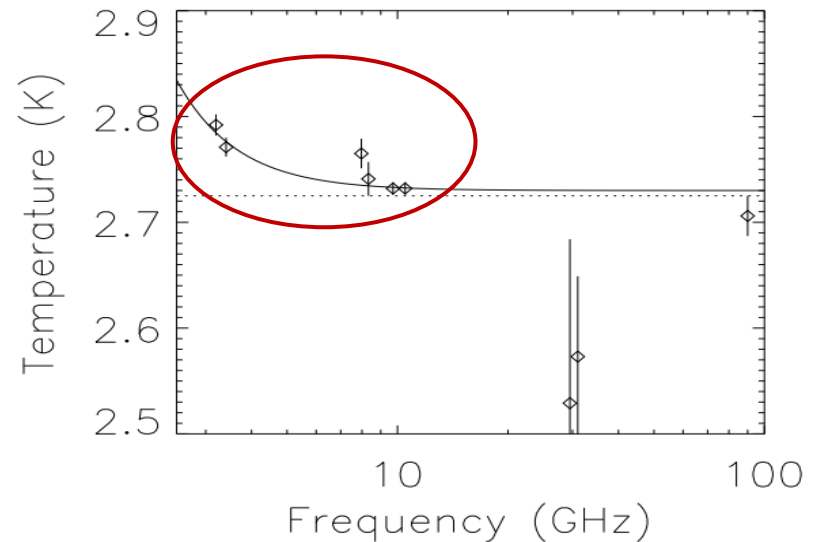


FIG. 4.— The thermodynamic temperature as a function of frequency. The solid line is the best fit to the ARCADE 2 data with a constant CMB temperature plus a synchrotron like component with an assumed  $-2.62$  index. The vertical lines are  $\pm 1\sigma$ . The dotted line is the FIRAS CMB temperature.

# Can neutrino decay solve these anomalies?

Phys.Lett.B 790 (2019) 64-70

- ❑ Considered the possibility that  $\nu_i \rightarrow \nu_s + \gamma$  with  $m_i - m_s = E_{21} z_{\text{decay}} / z_E$
- ❑ Active neutrinos have to decay non-relativistically.
- ❑ The specific intensity produced by the decays is found to be

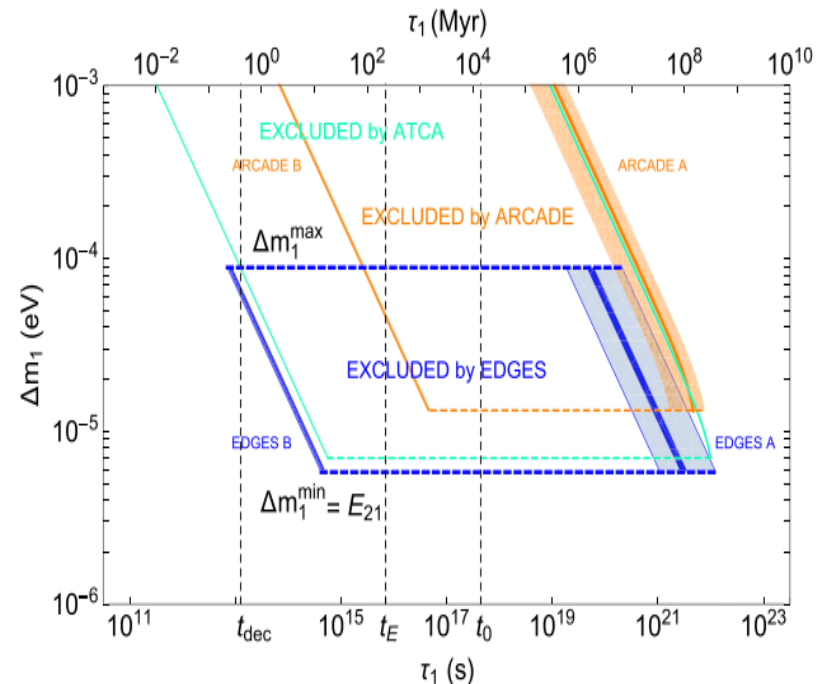
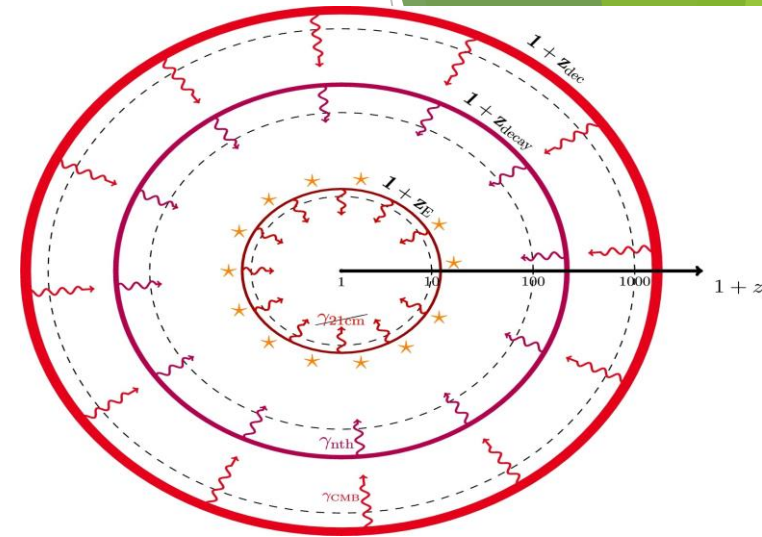
$$I_{\text{nth}}(E_{21}, z_E) = \frac{1}{4\pi} \frac{d\varepsilon_{\gamma_{\text{nth}}}}{dE} = \frac{n_{\nu_1}^{\infty}(z_E)}{4\pi} \left( \frac{E_{21}}{\Delta m_1} \right)^{3/2} \frac{e^{-\frac{t_E}{\tau_1} \left( \frac{E_{21}}{\Delta m_1} \right)^{3/2}}}{H_E \tau_1}$$

The EDGES signal is explained imposing

$$R \equiv \frac{I_{\text{nth}}(E_{21}, z_E)}{I_{\text{CMB}}(E_{21}, z_E)} = \frac{T_{\gamma_{\text{nth}}}(E_{21}, z_E)}{T_{\text{CMB}}(z_E)} = R_E \equiv 1.15^{+2.15}_{-0.8}$$

or alternatively one can always interpret the EDGES results as an upper bound  $R < R_E$  resulting in an excluded region.

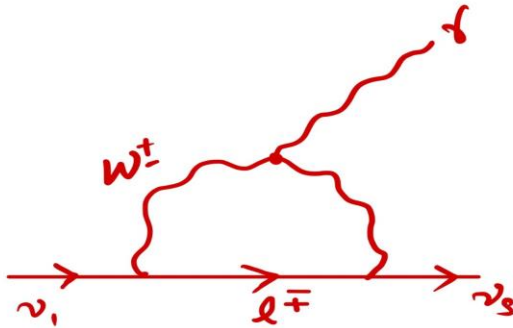
- ❑ Intriguingly the same mechanism can also explain the ARCADE excess in RB and the two allowed regions marginally overlap!



# Failure of minimal model

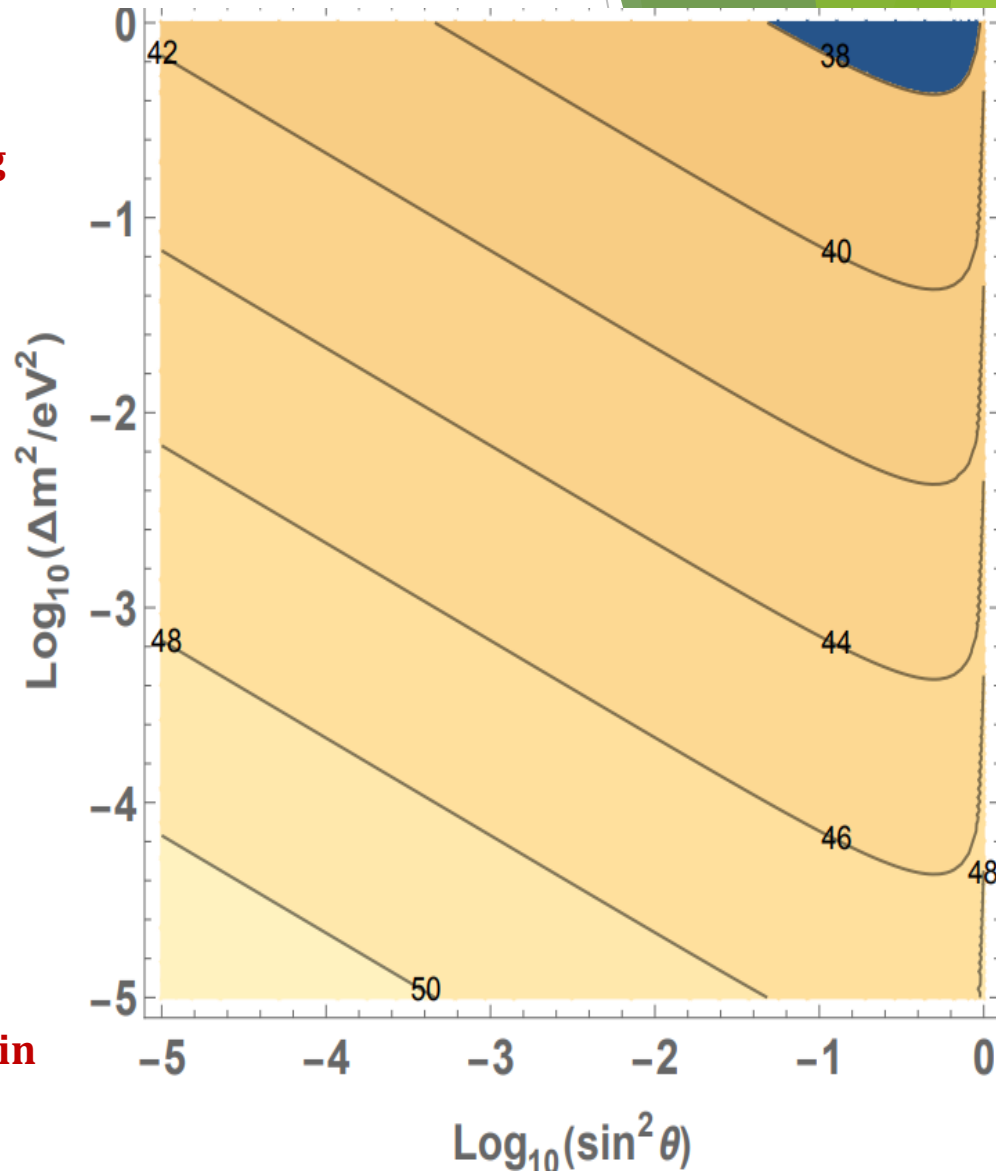
- Minimal Type-I seesaw
- As a result of active-sterile mixing

$$\nu_i \rightarrow \nu_s + \gamma$$



$$\Gamma(\nu_i \rightarrow \nu_j \gamma) \simeq \frac{9\alpha G_F^2 m_i^5}{256\pi^4} \sin^2(2\theta) \left( \frac{\Delta m^2}{m_i^2} \right)^2$$

- Life-time very large: cannot explain these anomalies!





# Next-to-minimal model

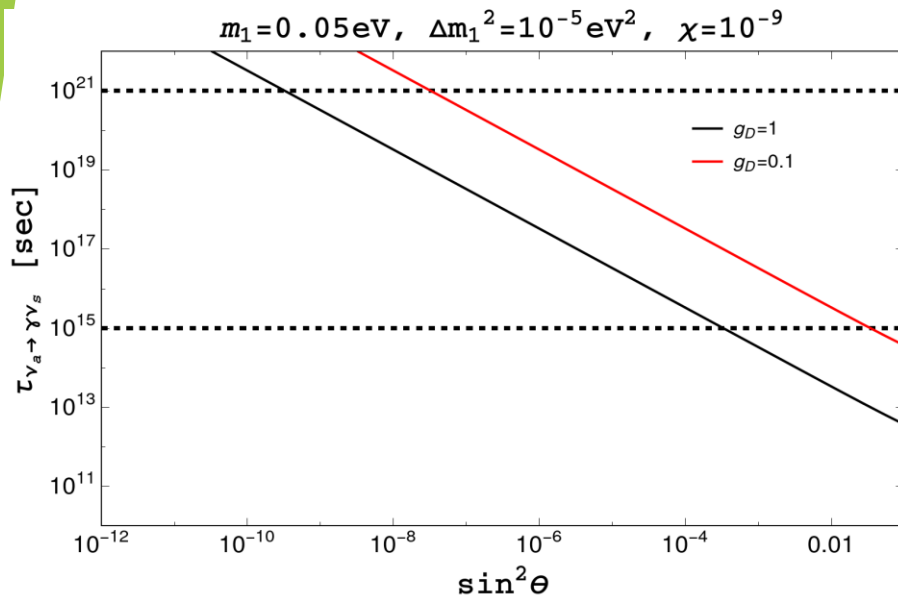
$$\mathcal{L}^D = \frac{m_D^2}{2} X_{D\mu} X_D^\mu + \bar{\nu}_D \gamma_\mu D^\mu \nu_D + \frac{\sin \alpha}{2} F^{\mu\nu} X_{\mu\nu} + h \bar{L}_\alpha H \nu_D + \frac{1}{2} M_{\nu D} \nu_D \nu_D$$

$$D^\mu = \partial^\mu + i g_D X_D^\mu$$

$$\nu_1 = \cos \theta \nu_a - \sin \theta \nu_D$$

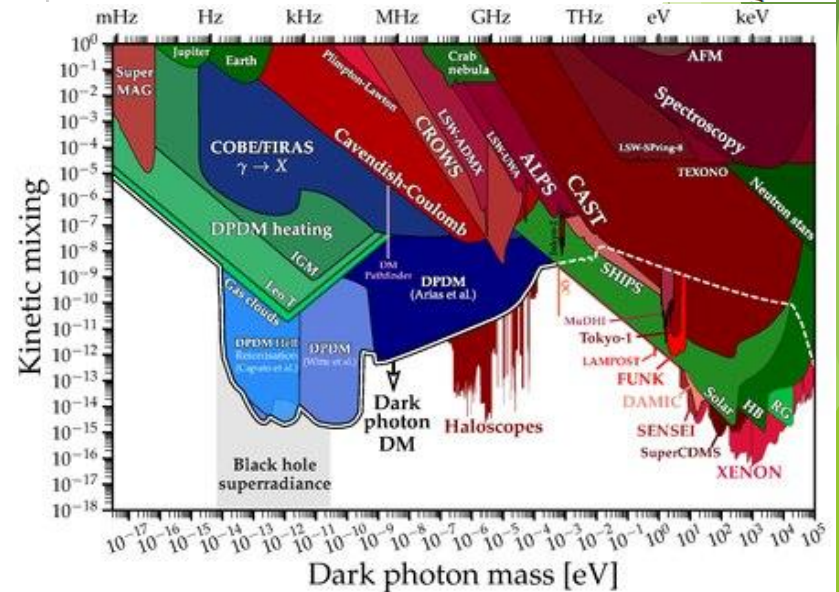
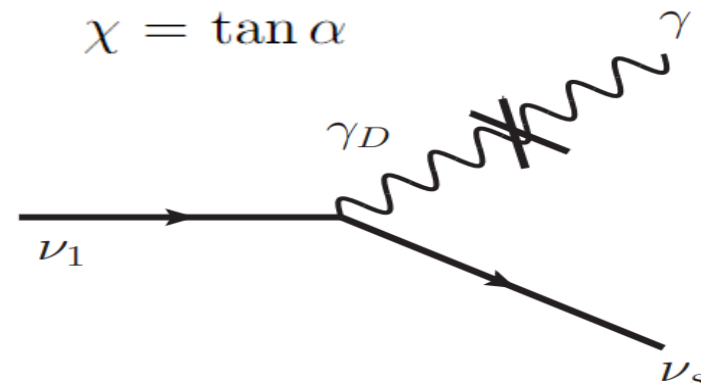
$$\nu_s = \sin \theta \nu_a + \cos \theta \nu_D.$$

$$\Gamma_{\nu_1 \rightarrow \nu_s \gamma} \simeq \frac{g_D^2 \sin^2(2\theta)}{256\pi} \chi^2 \frac{m_1}{2} \left(1 - \frac{m_{\nu_s}^2}{m_1^2}\right)^2$$

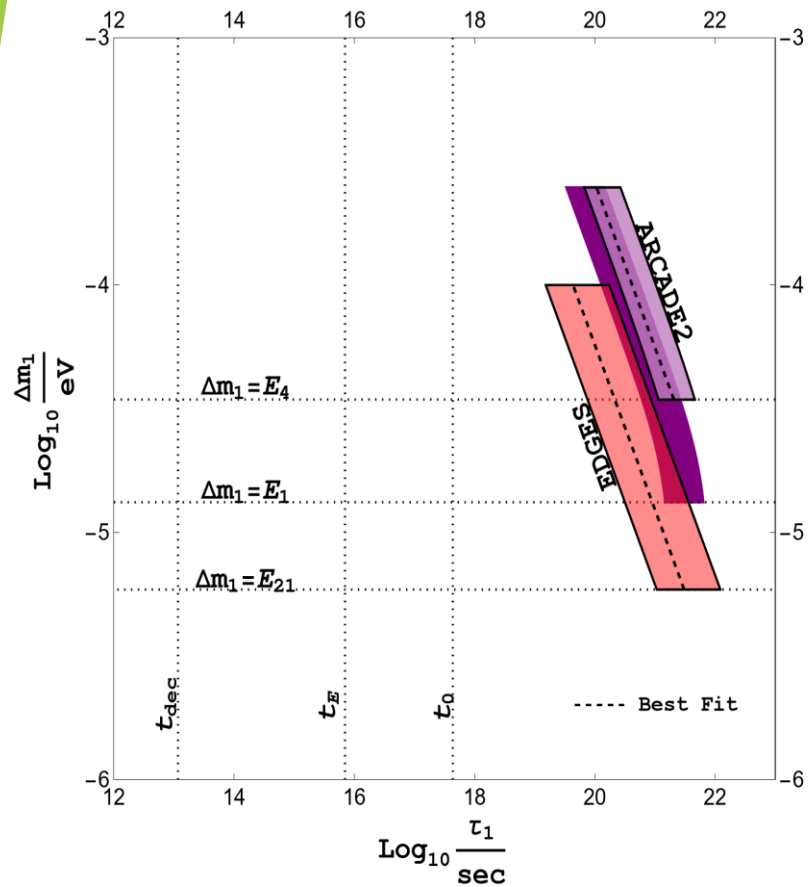


$$\gamma = A \cos \alpha, \gamma_D = X_D - \sin \alpha A$$

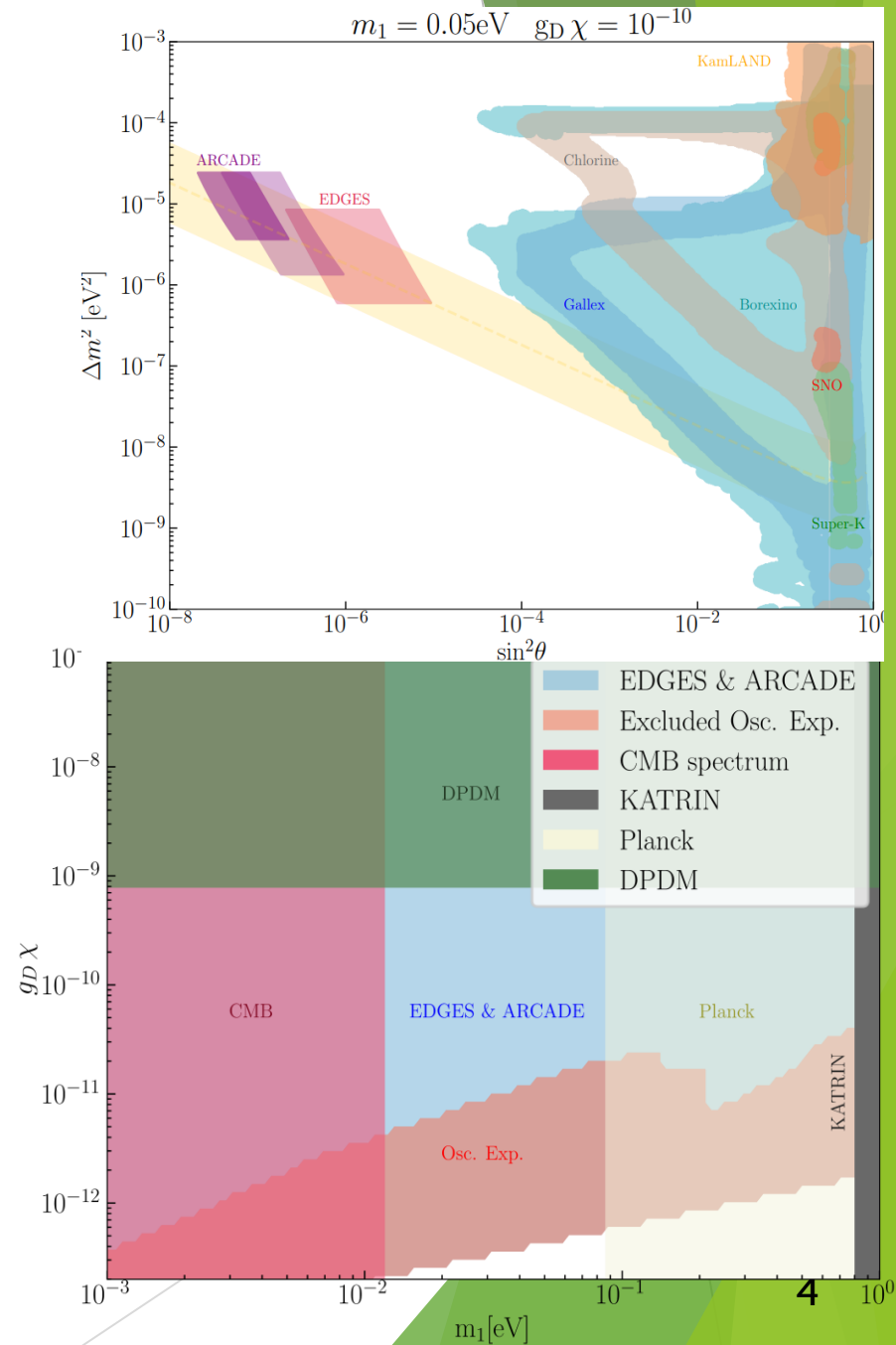
$$\chi = \tan \alpha$$



# Results



$$T_{\text{nth}}(E, 0) \simeq \frac{6 \zeta(3)}{11 \sqrt{\Omega_{\text{M}0}}} \frac{T_0^3}{E^{1/2} \Delta m_1^{3/2}} \frac{t_0}{\tau_1} \left( 1 + \frac{a^3}{a_{\text{eq}}^3} \right)^{-\frac{1}{2}}$$



# Conclusions

- ❑ **The anomalies observed by ARCADE-2 and EDGES are not explained within the standard cosmological model and there is currently no clear astrophysical solution so they might indicate the presence of new physics.**
- ❑ **The radiative decays of relic neutrinos into sterile neutrinos provide a simultaneous solution to both puzzles**
- ❑ **We show that, in our model, the decay of relic neutrinos into sterile neutrinos also implies active-sterile neutrino mixing and how neutrino oscillation experiments place interesting constraints on the parameters of the model.**
- ❑ **Neutrinos in Cosmology is not just a topic with important historical results, but it is still one of the best motivated routes to understand the cosmological puzzles**



THANK YOU!