

# Gravitational wave emission from a global string loop

Jorge Baeza-Ballesteros

In collaboration with E. Copeland, D. G. Figueroa and J. Lizarraga  
Based on arXiv/2308.08456

IFIC, University of Valencia-CSIC

First Nordic Cosmology Meeting - 25th October 2023



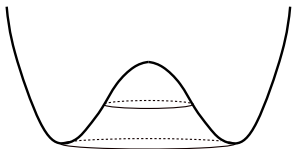
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# Cosmic strings in the early universe

Cosmic strings are **one-dimensional topological defects**

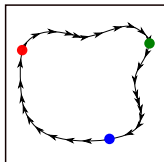
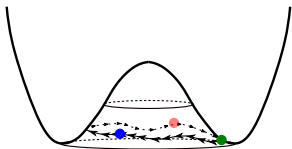
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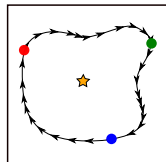
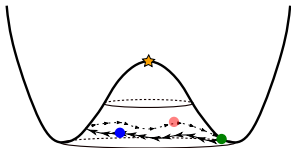
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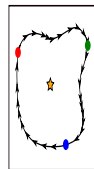
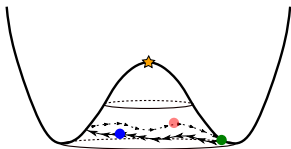
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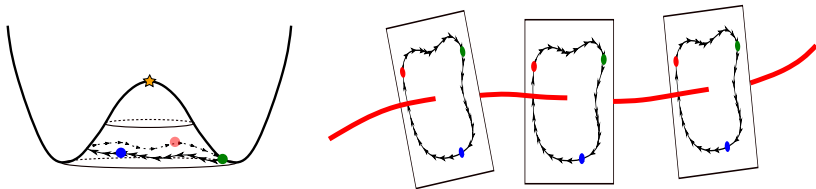
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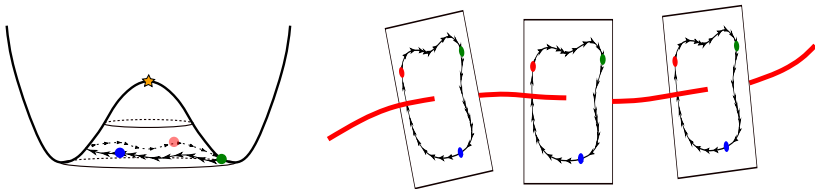
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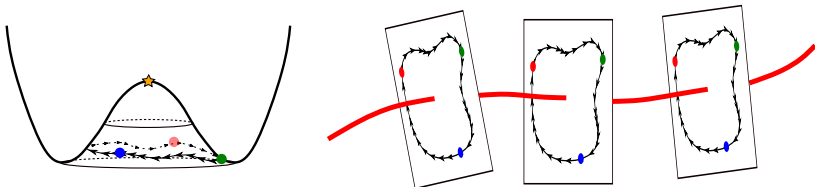


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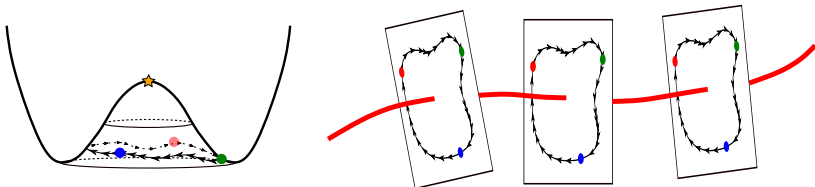


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**Detection:** CMB, gravitational wave (GW) background...

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Need to constrain the GW signature from cosmic strings

# The Nambu-Goto approximation

Traditional picture  $\rightarrow$  **Nambu-Goto approximation** (zero width)

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- ▶ Loops decay via GWs radiated in all harmonic frequencies  $\nu_j = j/T$ , with  $T = L_{\text{str}}/2$

$$P_j = \Gamma G\mu^2 \frac{j^{-q}}{\zeta(q)} \longrightarrow P_{\text{GW}} = \dot{E}_{\text{GW}} = \sum_{j=1}^{\infty} P_j = \Gamma G\mu^2$$

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**Goal:** Particle and GW emission using lattice simulations

# Global string loops in flat spacetime

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$$V(\varphi) = \lambda \left( |\varphi|^2 - \frac{v^2}{2} \right)^2 \quad \longrightarrow$$

Massive ( $m = \sqrt{2\lambda}v$ )  
and massless radiation

**Global strings** ( $r_c \sim m^{-1}$ )

# Global string loops in flat spacetime

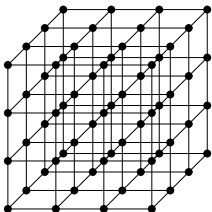
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Lattice simulations with **CosmoLattice** [Figueroa et al, 2020, 2021]



$$\equiv \frac{L}{\delta x}$$

**Dimensionless variables:**

$$\tilde{x} = \sqrt{\lambda} v x$$

$$\tilde{t} = \sqrt{\lambda} v t$$

$$\tilde{\varphi} = \varphi/v$$

# Initial conditions

We consider **two types of loops**:

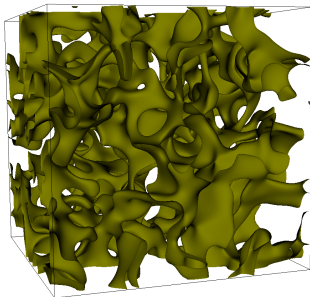
**Network loops**  
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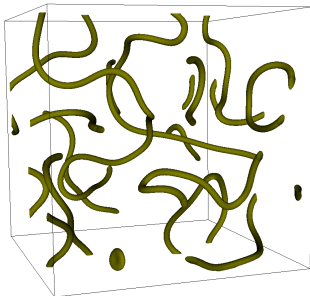
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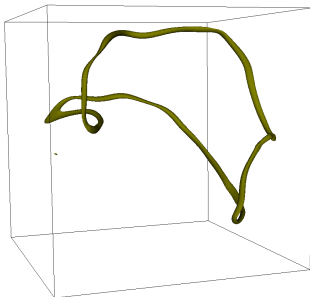
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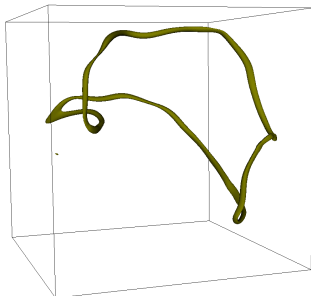
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## Artificial loops

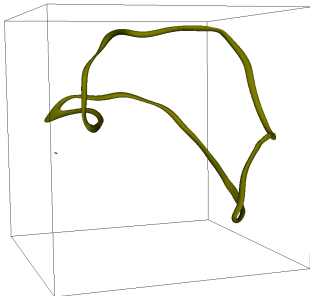
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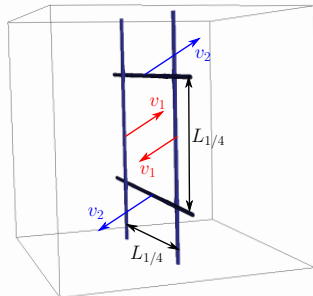
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Random initial conditions

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Control over initial conditions

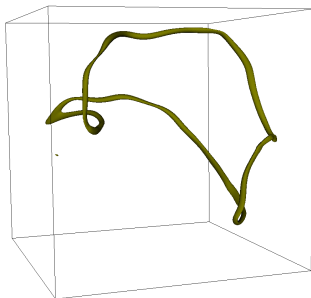


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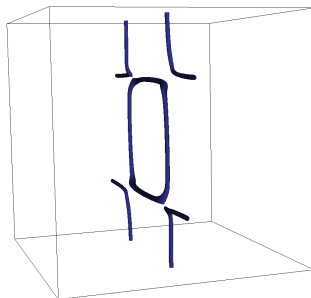
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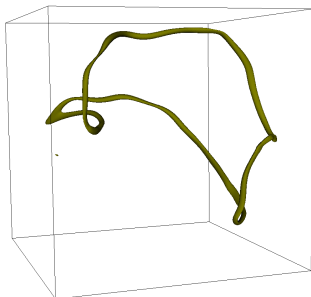
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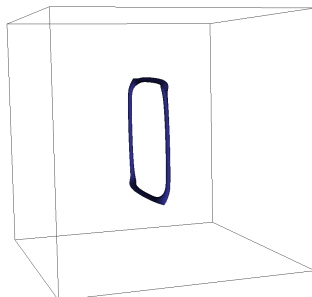
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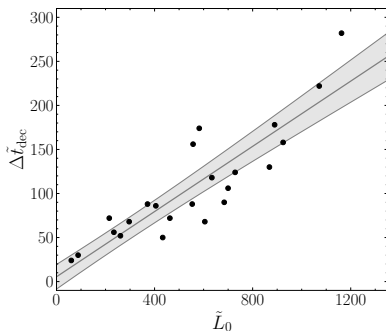


Control over initial conditions

# Loop decay time

Decay time vs **initial length** ( $L_0 \lesssim 1700r_c$ ) → linear relation

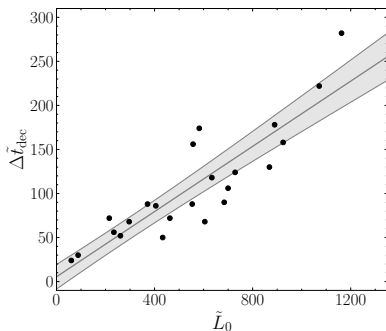
**Network**



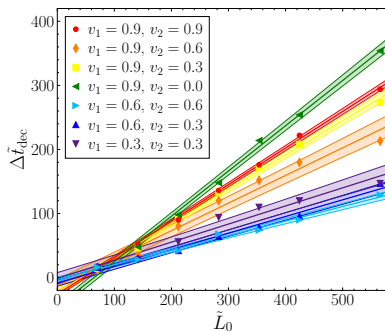
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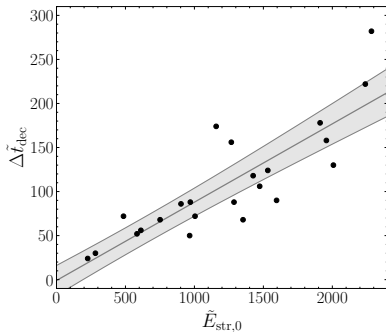


Dependence on  
initial conditions

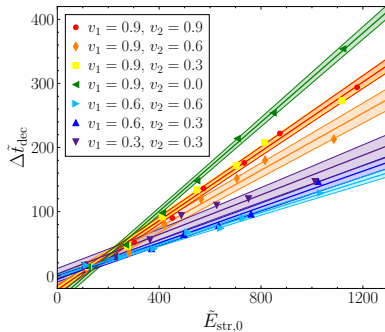
# Loop decay time

Decay time vs **initial string energy** → linear relation

**Network**



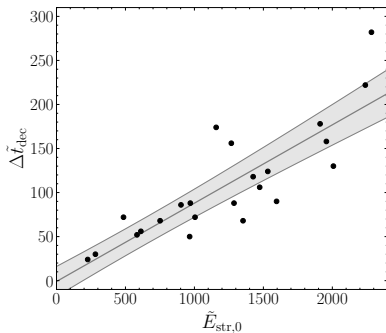
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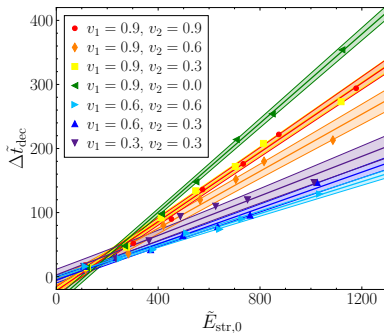
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$$P_{\varphi} = -\frac{dE_{\text{str}}}{dt} \approx (11.2 \pm 1.6)v^2$$

## Artificial



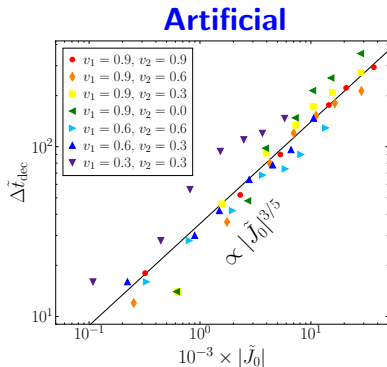
$$P_{\varphi} = -\frac{dE_{\text{str}}}{dt} \approx (3 - 8)v^2$$

# Loop decay time

Decay time vs **angular momentum** → **universal scaling**

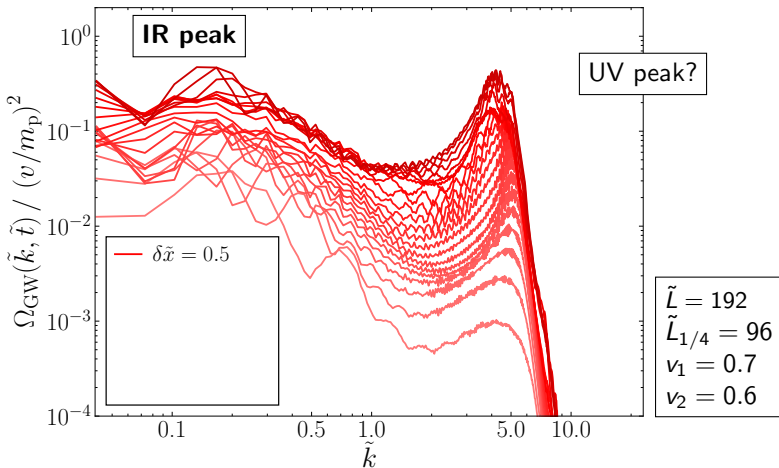
$$\vec{J} = -2 \int_{\text{str}} d^3x \operatorname{Re} \left[ \vec{x} \times \dot{\varphi} \vec{\nabla} \varphi^* \right]$$

↓  
Distance to loop's  
geometric center



# GW results: UV sensitivity of artificial loops

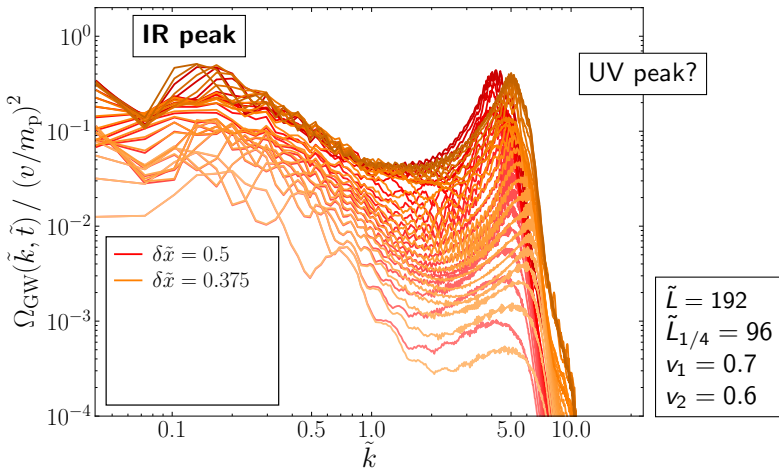
Fixed loop and lattice size + reduce lattice spacing





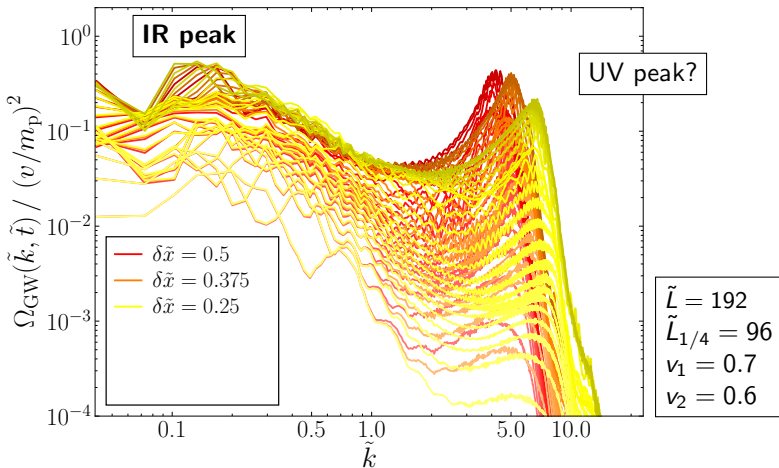
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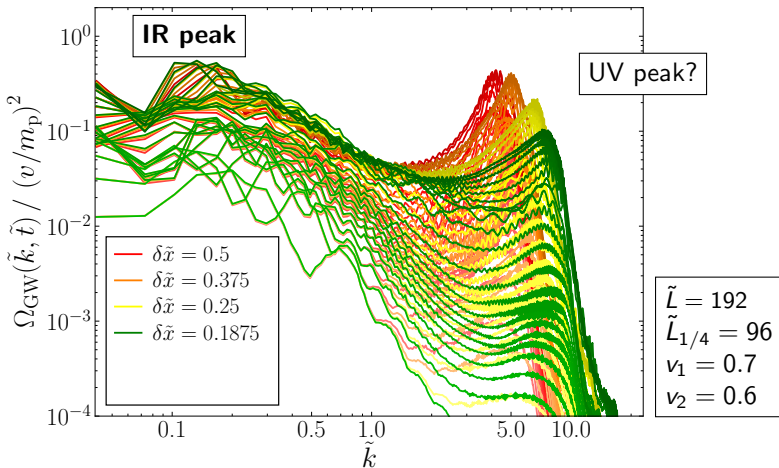
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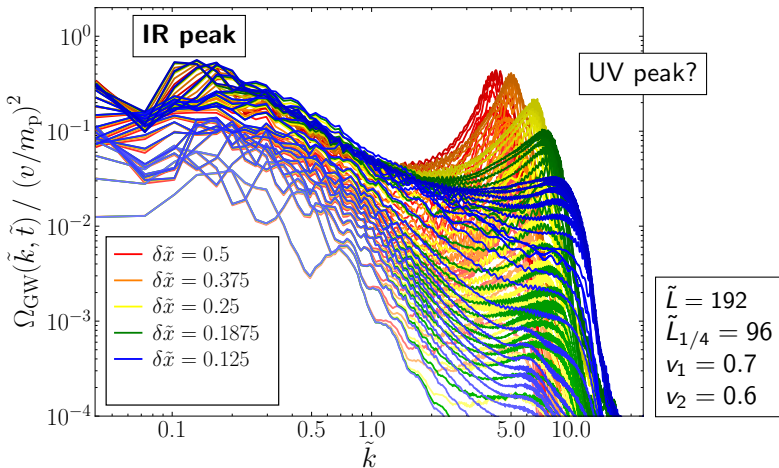
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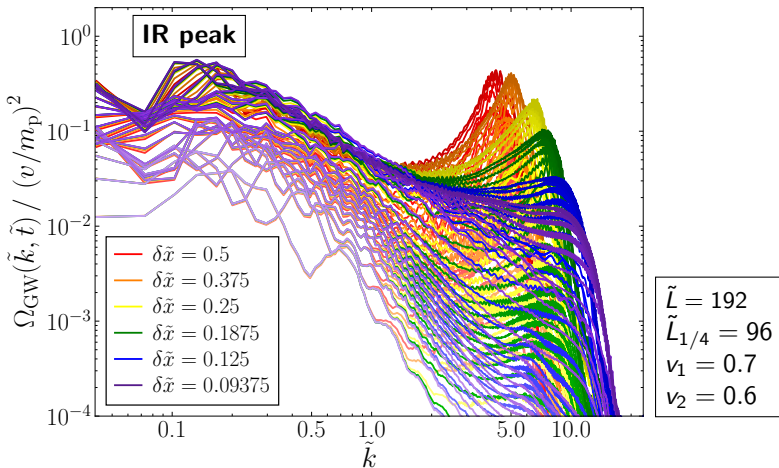
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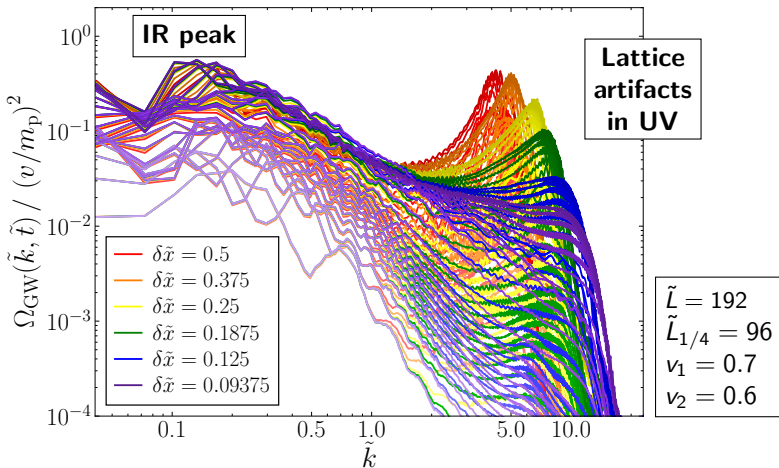
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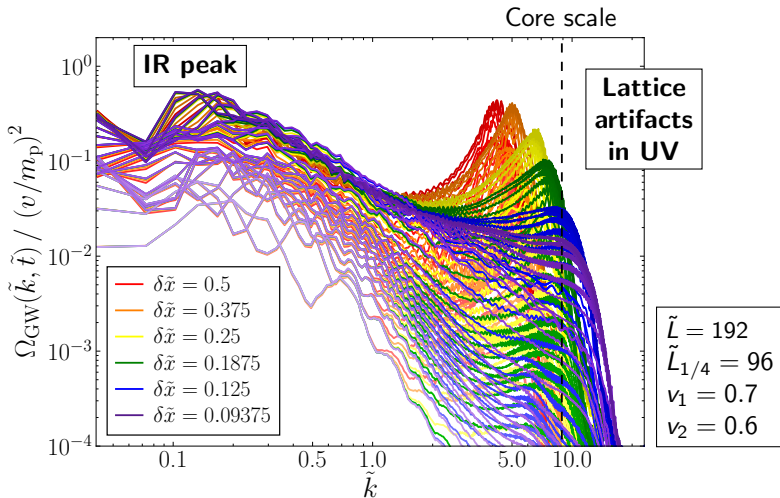
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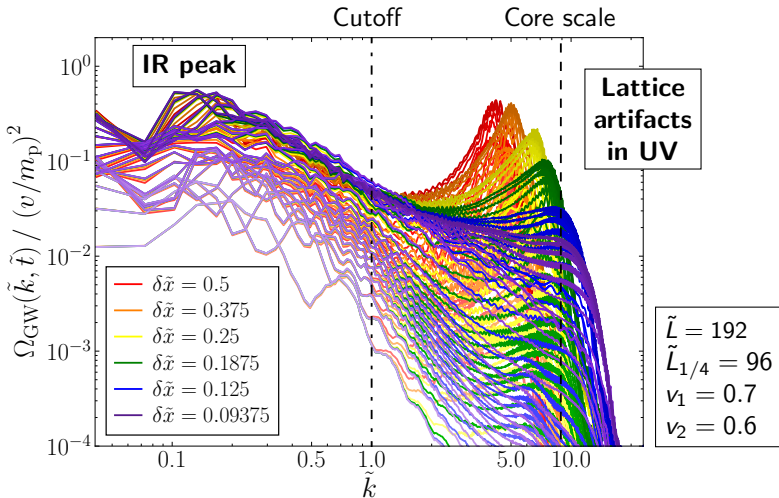
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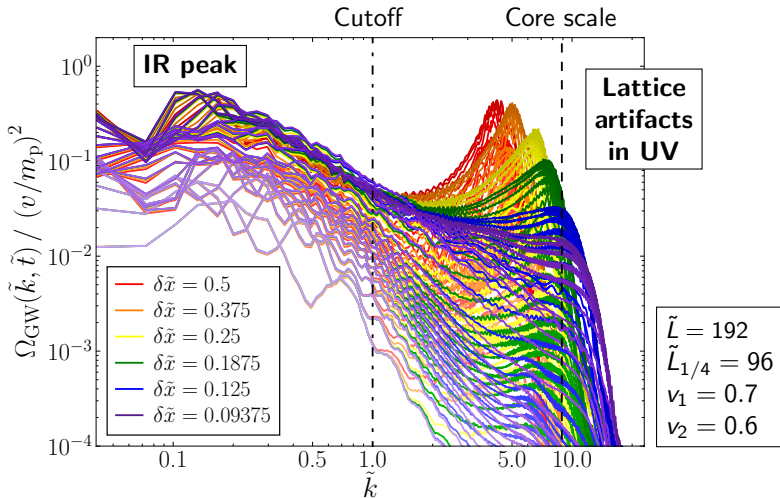
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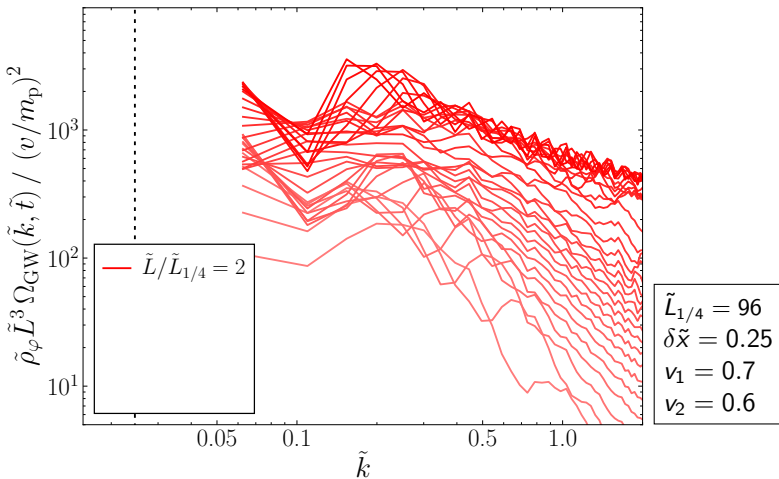


Similar results for network loops

# GW results: IR sensitivity of artificial loops

Fixed loop and lattice spacing + Increase box size

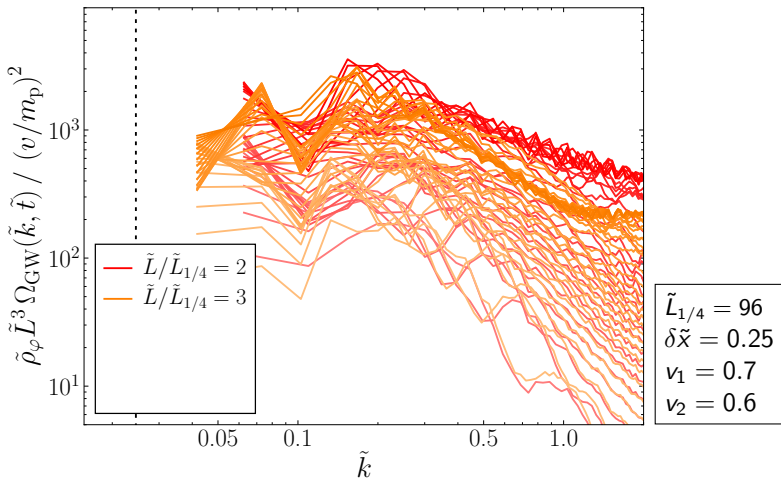
Loop length  
scale



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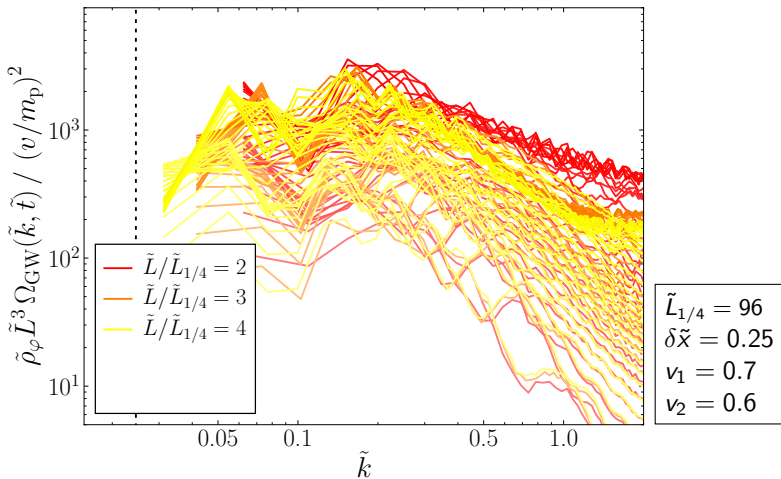
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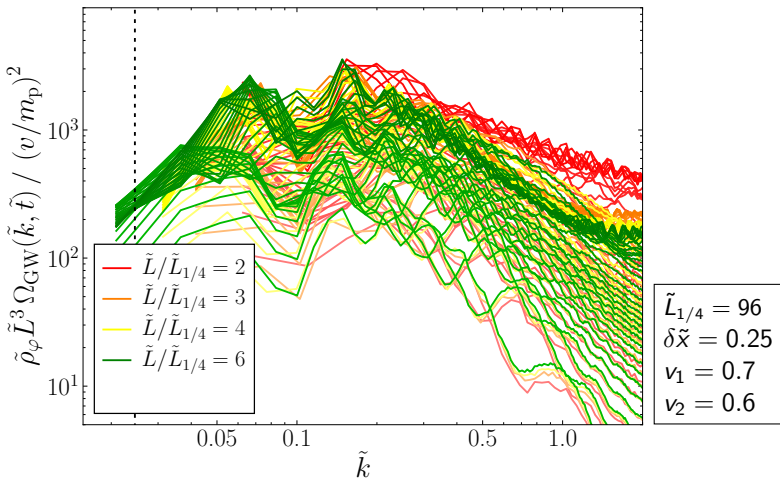
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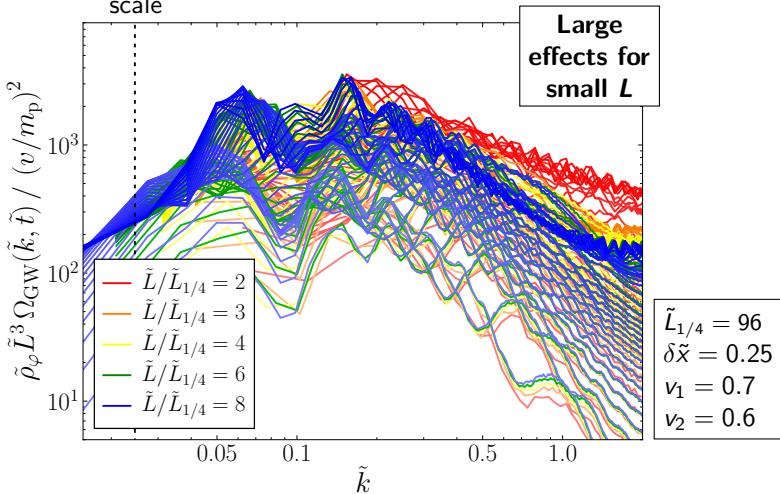


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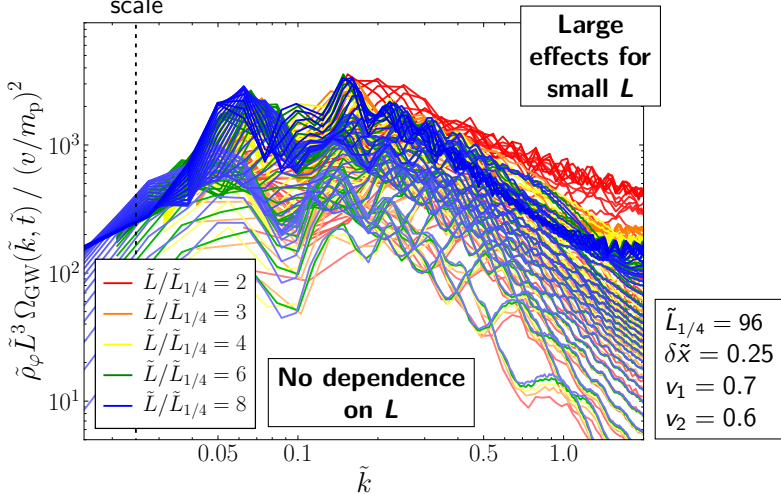


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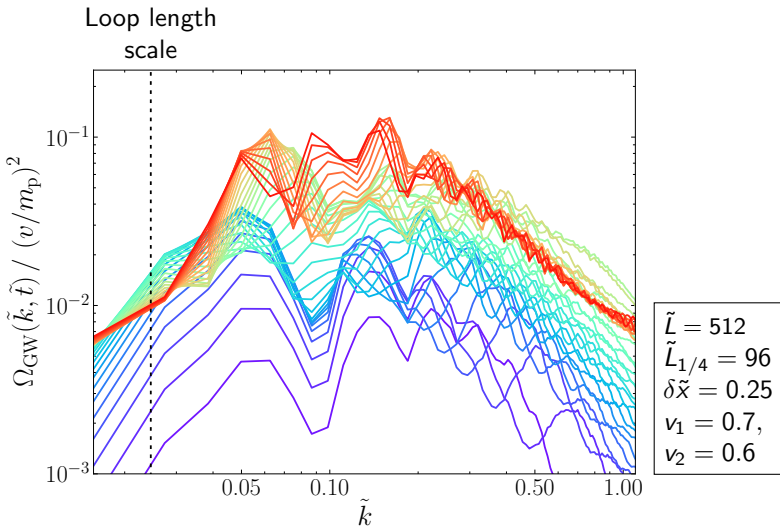
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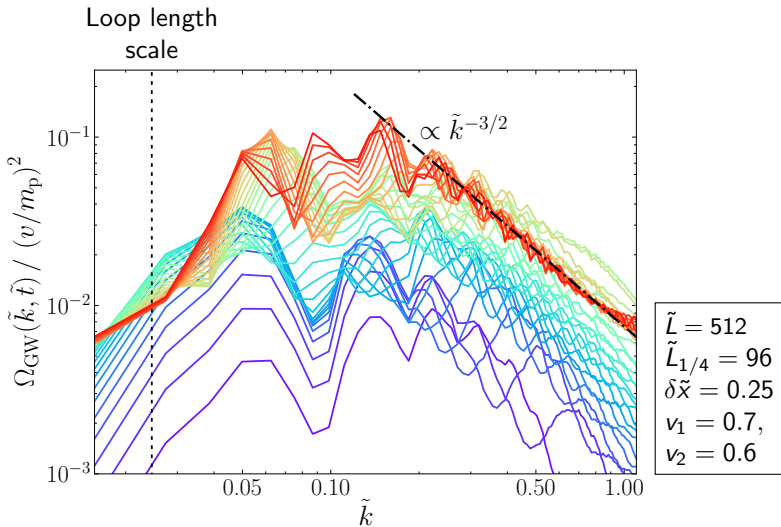
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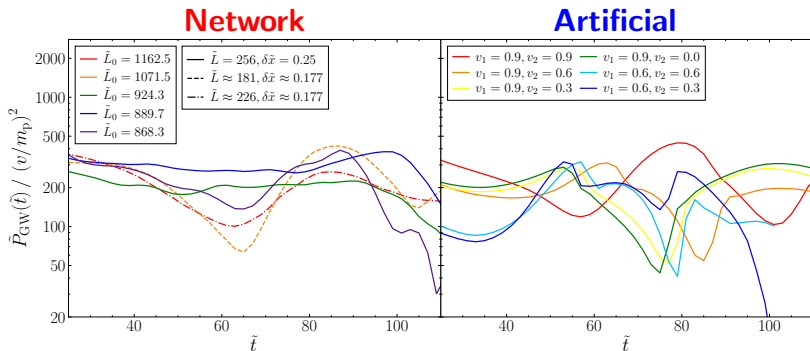


# GW results: radiation power

$$P_{\text{GW}}(t) \equiv \frac{L^3 \rho_\varphi}{2T} \int_{t-T}^{t+T} dt' \frac{d}{dt'} \int_0^{k_{\text{cut}}} \Omega_{\text{GW}}(k, t') d \log k$$

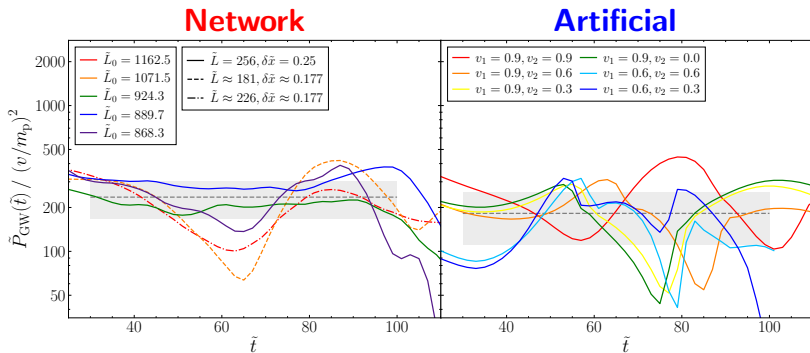
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No dependence on string length

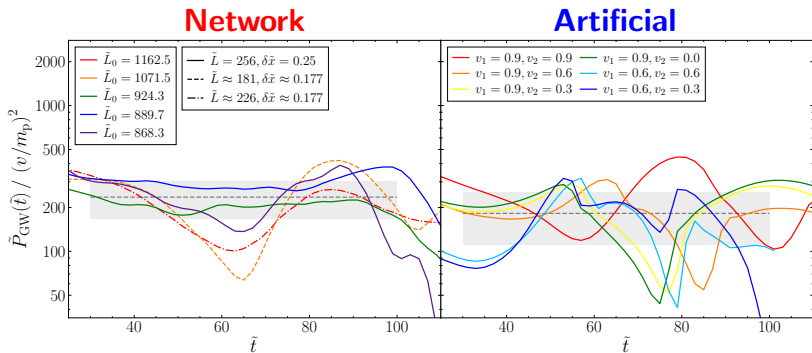
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$$P_{\text{GW}} = (240 \pm 80)v^2(v/m_p)^2$$

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$$\frac{P_{\text{GW}}}{P_\varphi} \approx \mathcal{O}(10) \left( \frac{v}{m_p} \right)^2, \quad v/m_p \lesssim 10^{-6} - 10^{-3}$$

[Hindmarsh et al. (2019),  
Benabou et al. (2023)]

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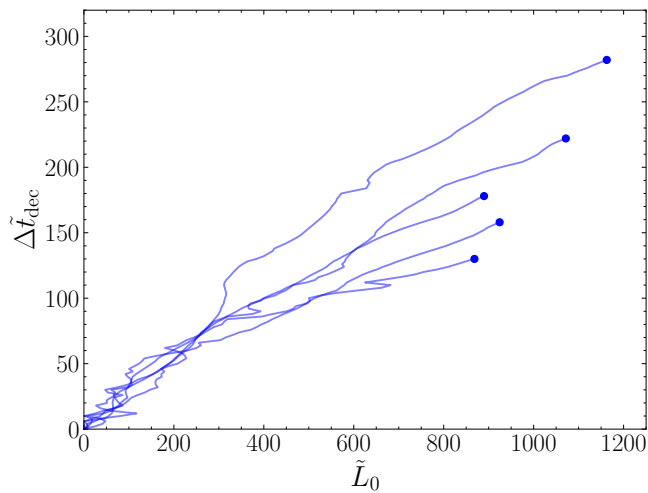
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# Thank you for your attention!

# Time-evolution of the string length



# GW results: UV sensitivity of network loops

Fixed loop and lattice size + reduce lattice spacing  
Coarse graining

