

Strongly Lensed Supernovae in the LSST era

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with: LSST DESC + IPTF/ZTF Cosmo WG

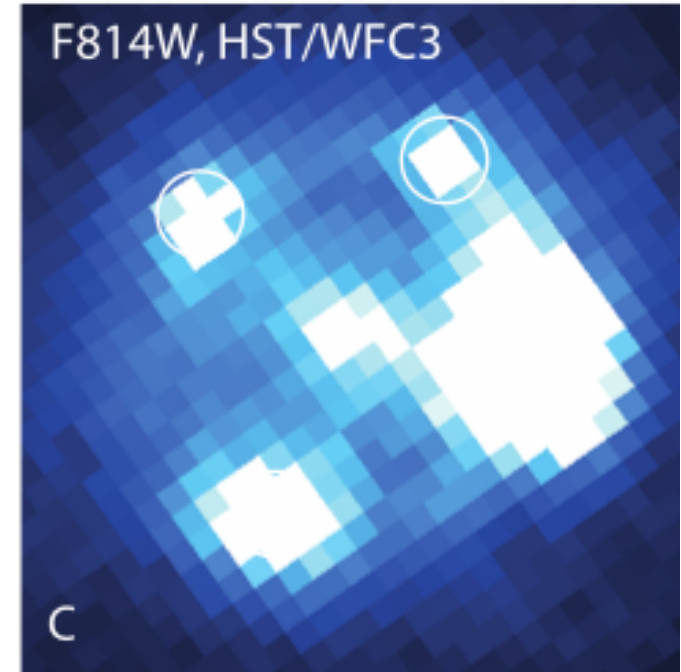


OKC@15, October 18 2023





Outline



Motivation

Recent discoveries

[SD+20b](#); [Mortsell, SD+ 2021](#); [Johansson, .. SD,+2021](#); [Goobar+, Nat. As, 2023](#)

Towards precision cosmology with LSST

[Birrer, SD, Shajib, 2022](#), [Dhawan + Pierel in prep.](#)

New time-delay inference method

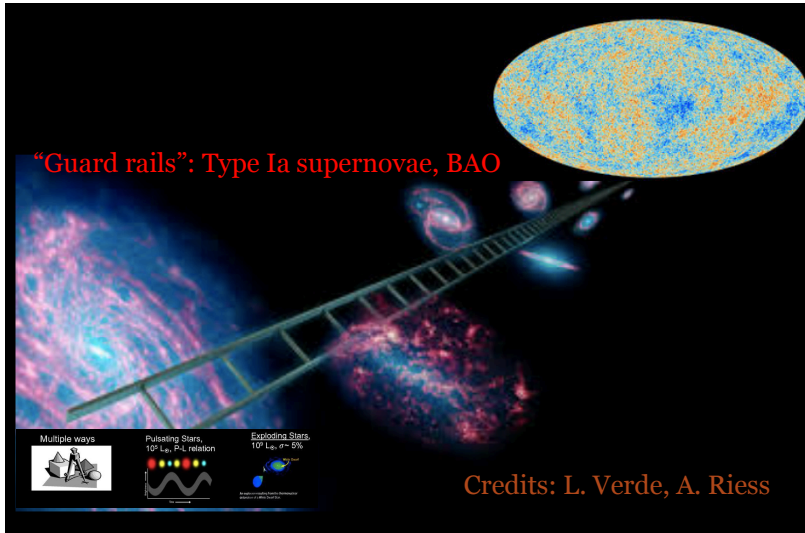
[Crascall-Kennedy,SD,+ in prep.](#)



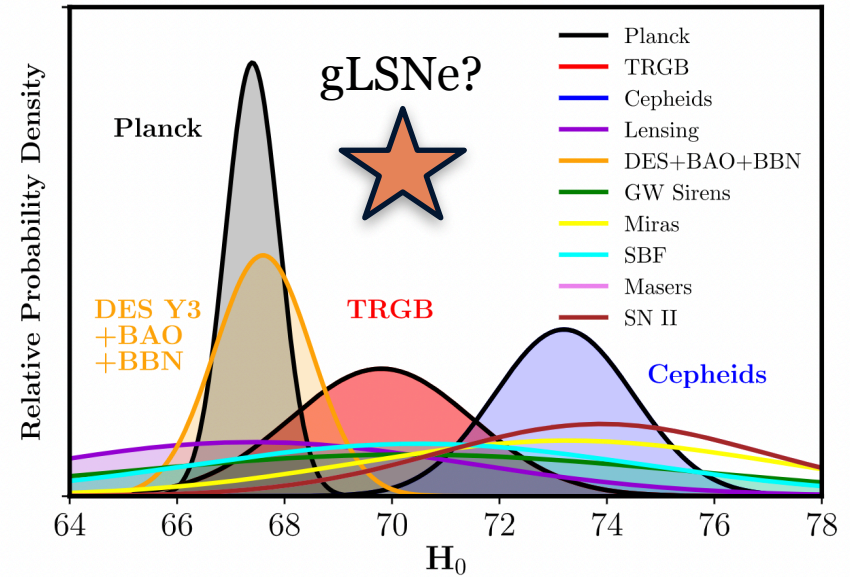
Motivation

- H_0 : Absolute scale of the universe
- End-to-end test of background expansion

Credits: Freedman 2021



Recent Published H_0 Values



- New physics?

- Unknown Systematics?

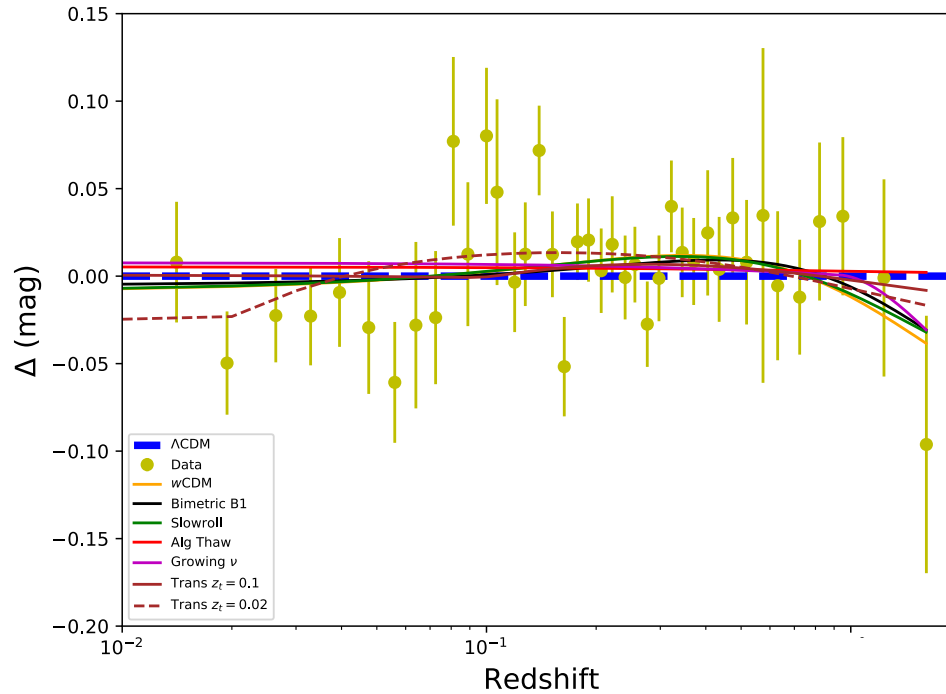


Need independent methods

- Novel absolute distance measurement (e.g. lensed transients)



Cosmological Solutions?



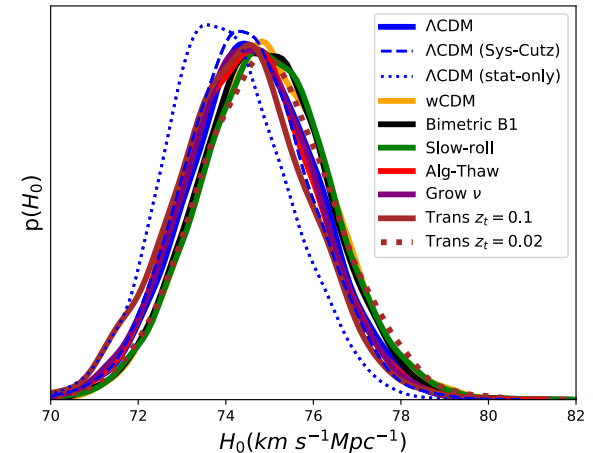
SD+20c

Simultaneous dark energy + H_0 constraint

Exotic models, including late-time transition

No model deviates from smooth- Λ

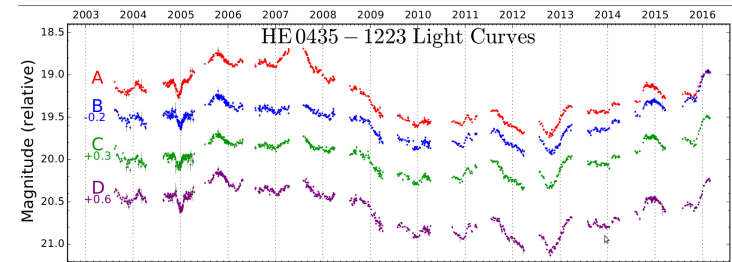
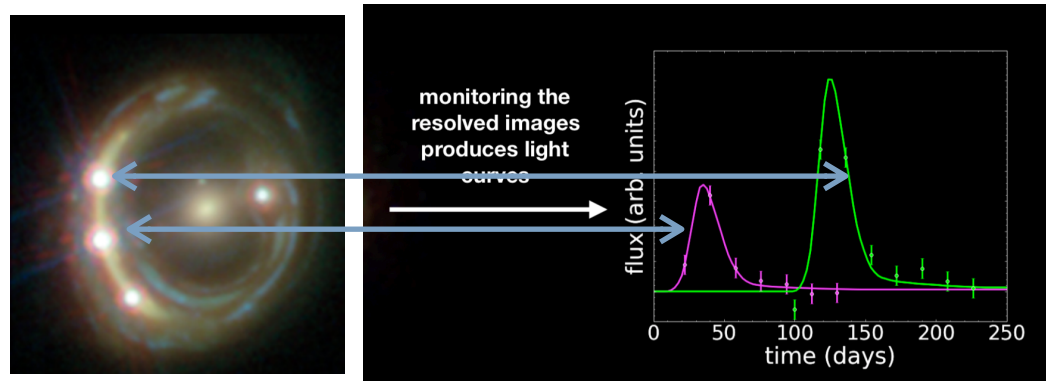
No transition seen in M_B (Lovick, SD, Handley in prep.)





Time-delay cosmography

Typical lensed SN and QSO light curves



$$\Delta t \propto D_{\Delta t} \times \phi_{\text{lens}} \Rightarrow D_{\Delta t} \propto \frac{1}{H_0}$$

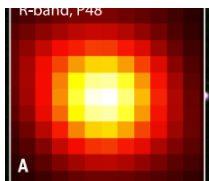
Time delay Time-delay distance Lens potential (from mass model)

Independent of SN distance scale!

- Independent discovery method to lensed quasars
 - gISNe => “standardisable candle”
 - Much less monitoring required
- First proposed in Refsdal 1964 (for SNe, used for QSOs)



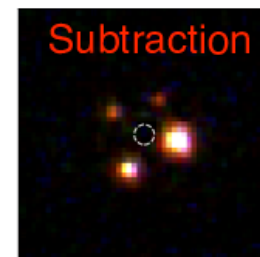
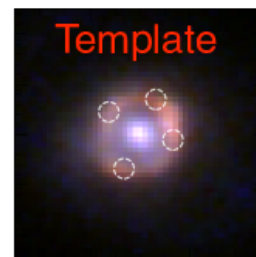
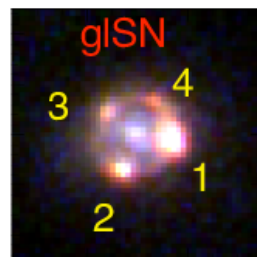
First Resolved lensed SN Ia



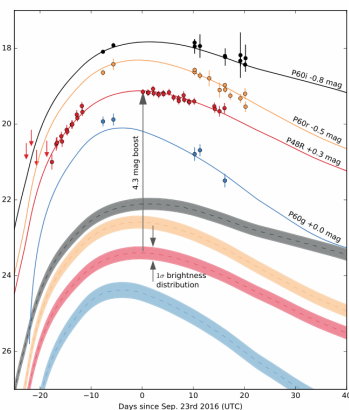
Discovery in unresolved data



Follow-up: HST / AO



HST/WFC resolved image, template and subtraction => not possible for QSOs!



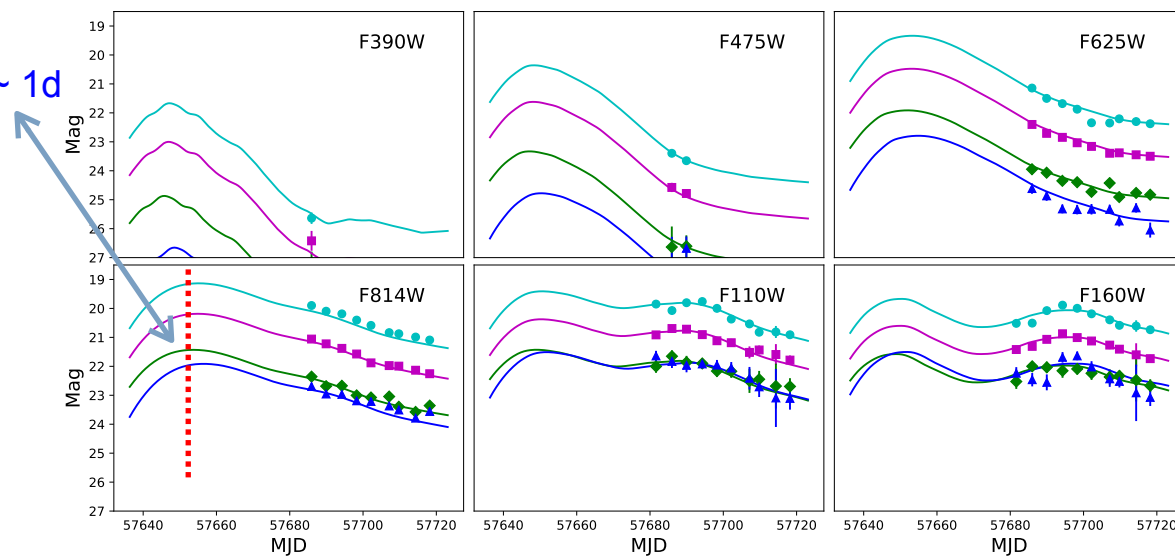
>50 times brighter than normal SNIa at $z \sim 0.4$: a 30σ outlier! (Goobar+2017)

Just a few orbits: errors $\sim 1d$

Very small time-delays (~ 1 day):
Not ideal for measuring H_0

Max. light simulations
=> five times smaller error

Long wavelength lever arm for
extinction constraints



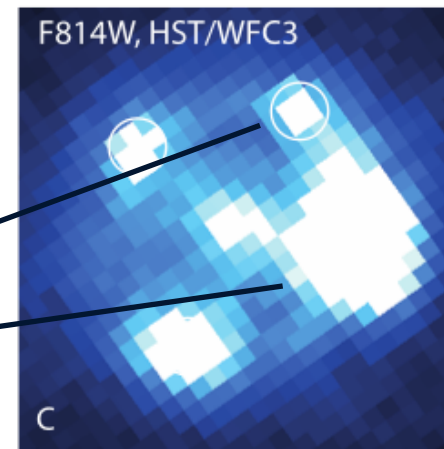
SD+20b



iPTF16geu: Magnification + extinction

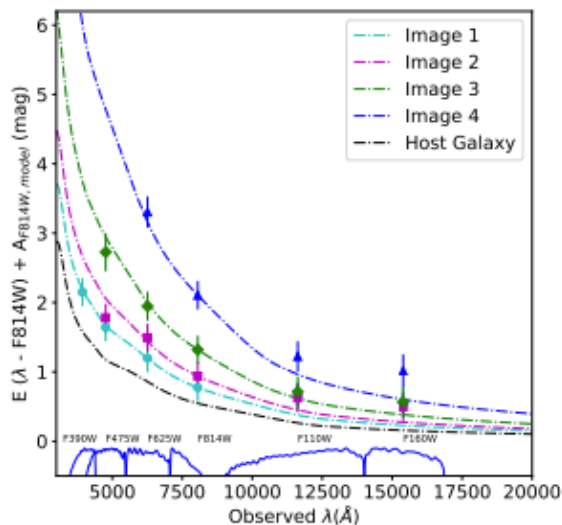
Important probe of dust in lens galaxy
LoS $R_v <$ Milky Way values

Model independent lensing magnification

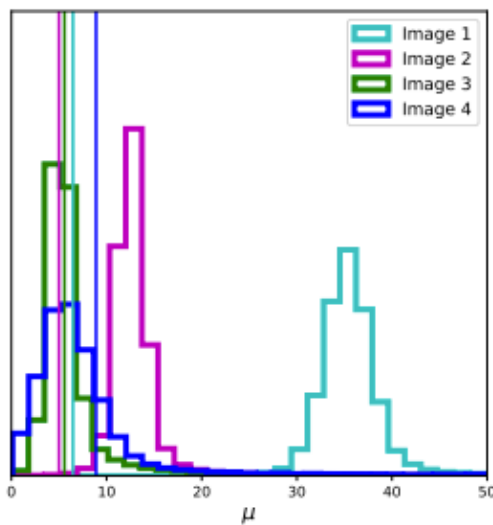


Surprisingly different brightness?

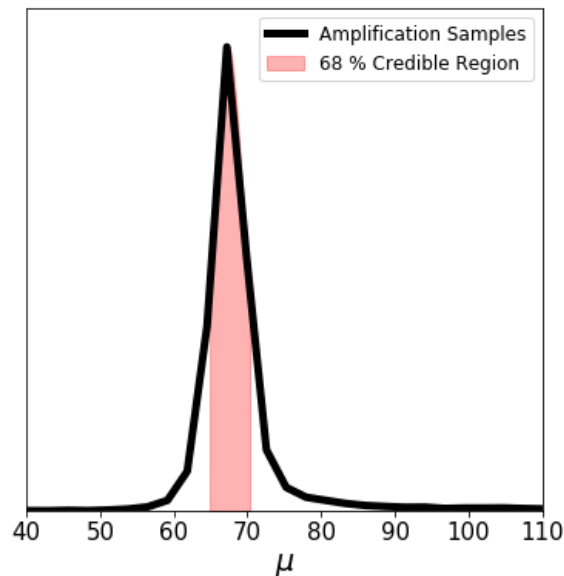
Preliminary magnification (μ) \sim 52
With extinction correction 67 \pm 3



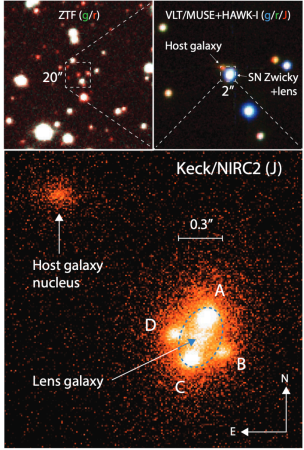
Spectroscopy in Johansson, ..., SD, + '21



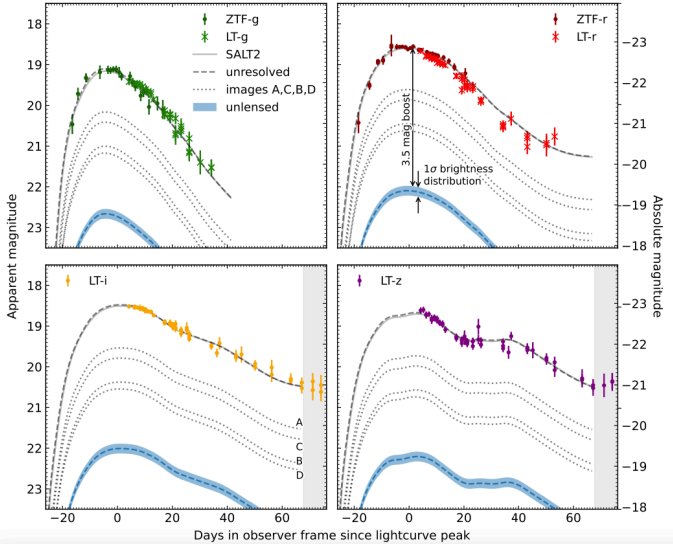
Modelling details in Mortzell, ..., SD, + '21



1" ~ 0.3"



0.01"



Small telescopes also vital for bright lensed SNe

JWST NIRCcam obs (Compared to HST live + template)

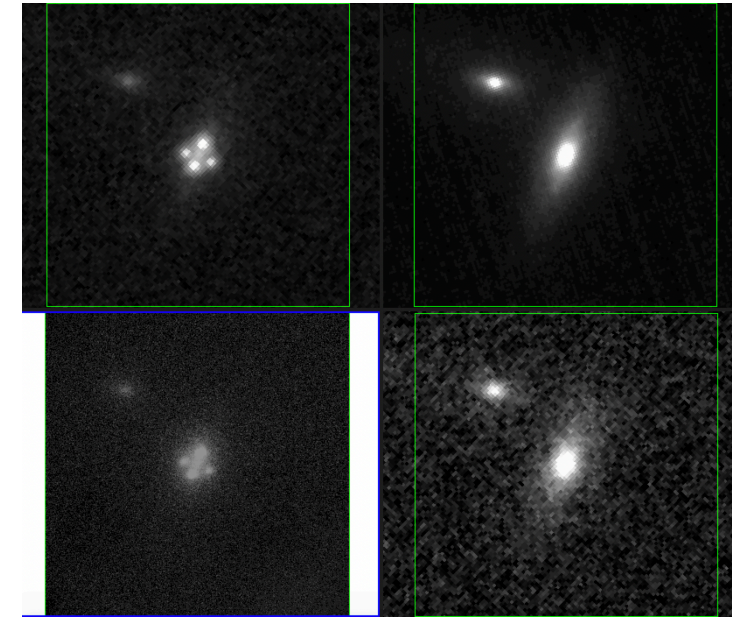
(Goobar+2023)

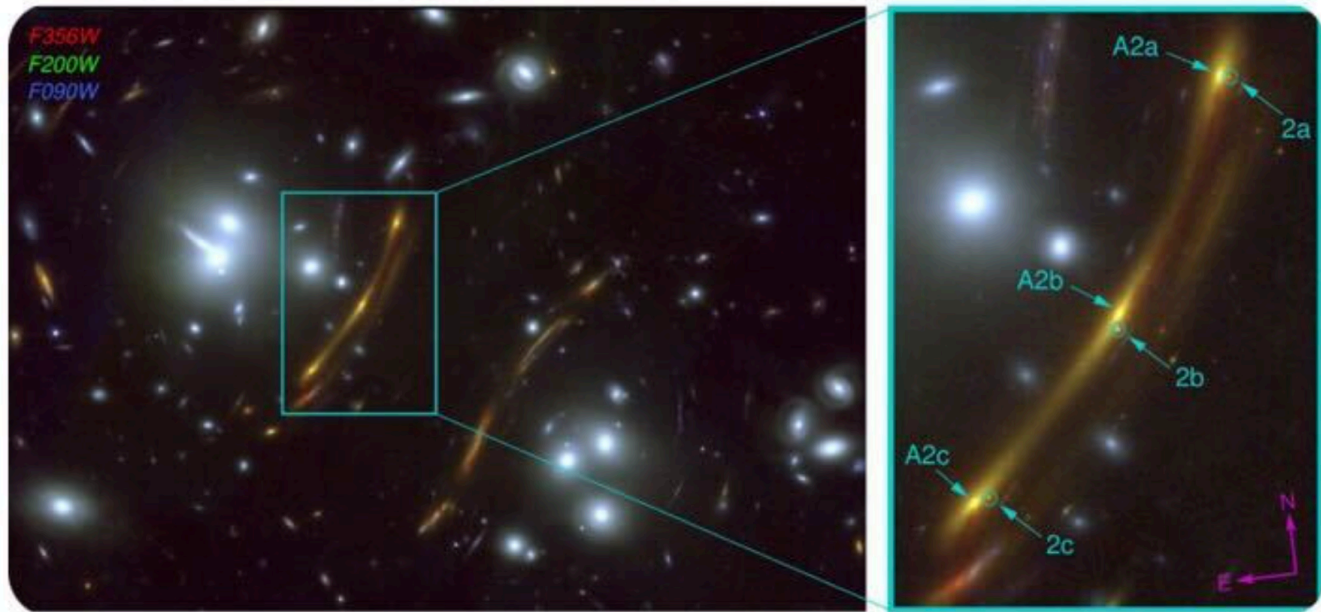
Multiband P48+ LT data
Accurate extinction constraints
PI: Perley, SD

~ 3.5 mag > SN Ia at $z_s = 0.354$

Low extinction in host + lens

Compact system $\theta_E < 0.2''$: study central stellar IMF





Discovery in NIRCAM obs of PLCK G165.7+67

Triply imaged SN Ia at $z = 1.78$

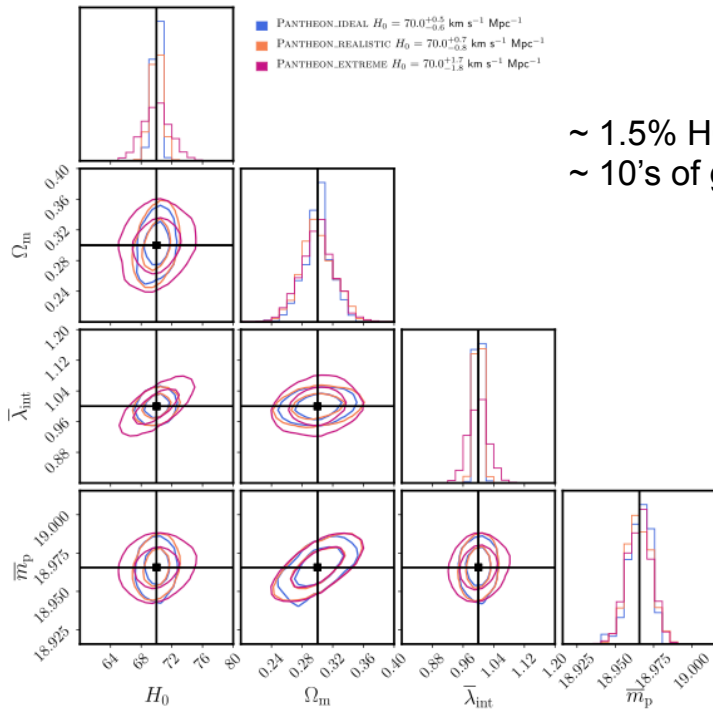
Cluster lens, long expected time-delay

Multiband follow-up with NIRCcam (F090W- \rightarrow F444W)

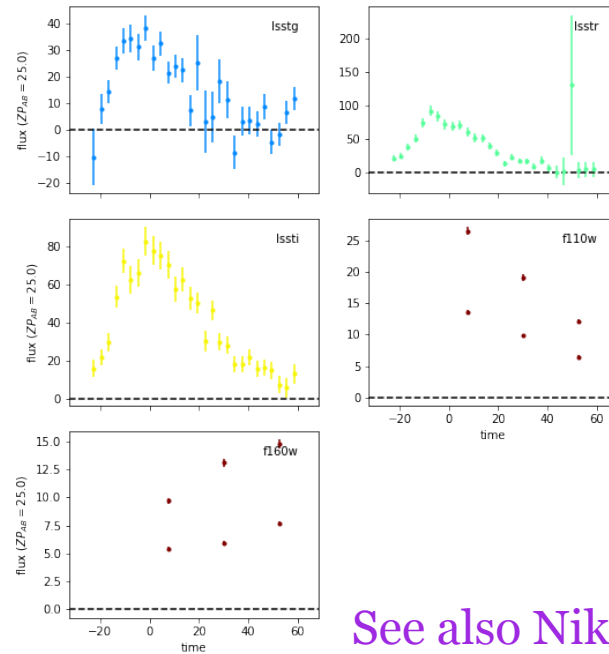
Expected H_0 at 10% \rightarrow 7% with reference epochs



Gearing up for LSST



~ 1.5% H_0 with LSST
 ~ 10's of gLSNe



Birrer, SD, Shajib, 2022

See also Nikki's talk for rates and parameters

High-Cadence LSST + sparse space-based follow-up (SD + Pierel in prep.)



Spectroscopic Time-Delay 11

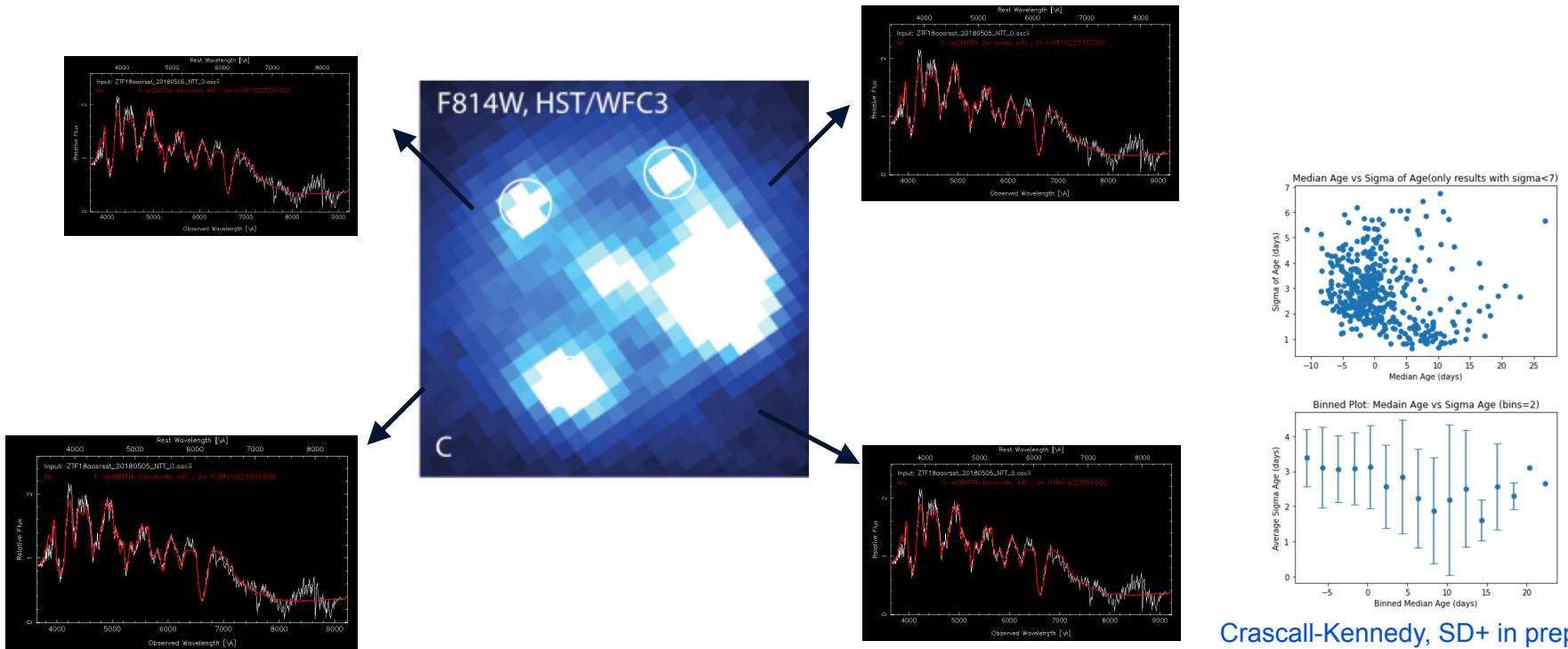
Spectroscopic time-delay: “one-shot” method (Johansson+2021 for 16geu)

Cross-correlate against template spectra

Small errors ($< 2d$) near maximum light

Works similarly well for core-collapse SNe

Example SNID fit for a typical low-z SNIa spectrum





Conclusions

- Current survey discoveries
 - 16geu: short time delays, extreme magnification
 - SN Zwicky: Most compact galaxy lens to date
 - Excellent laboratory for spectroscopic studies
- Forecasts for cosmology
 - Independent H_0 at 1.5% with LSST
 - Detect a large sample with feasible spectroscopy
- New inference methods
 - Spectroscopy is a promising alternative



High-resolution observations

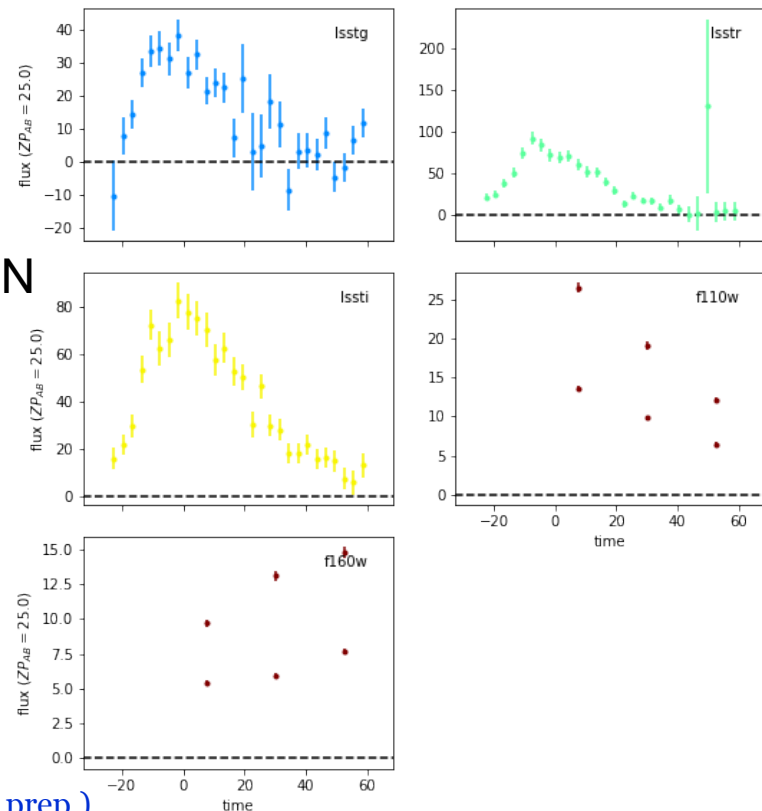
LSST lightcurves (mostly) unresolved

Lensing parameter constraints require follow-up

Combine with space based photometry

- Phase for triggering
- Number of observations

Preliminary results suggest 5 observations per gLSN

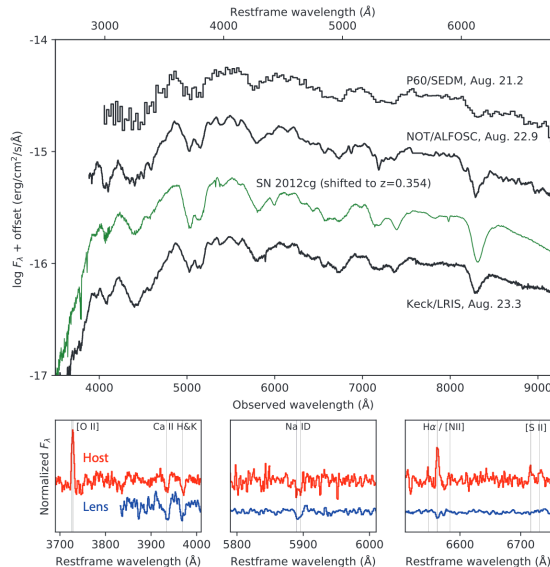
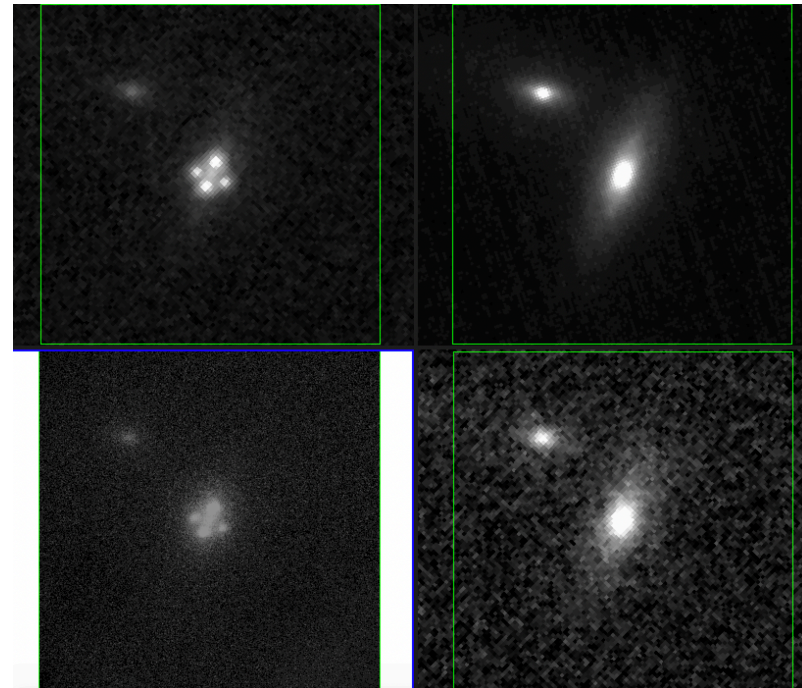




Spatial Curvature Constraints

- Nebular phase (~ 1 year post explosion) observations
- Four filter photometry + spectra (NIRCam + NIRSpec IFU)
- Faint SN, important constraints on lens light
- Important for post-explosion host + lens modelling
- Nebular spectra -> shed light on progenitors
- Photospheric spectra from the ground

NIRCam image of host + lens



SN Zwicky spectra compared to local SN 2012cg

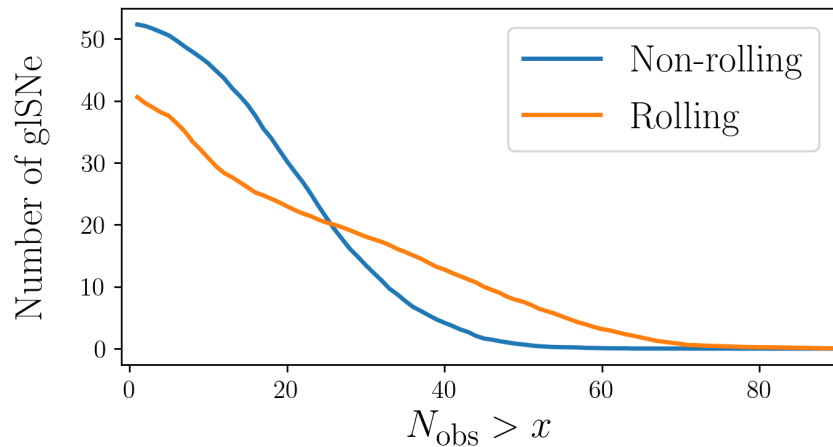


Detectability in LSST

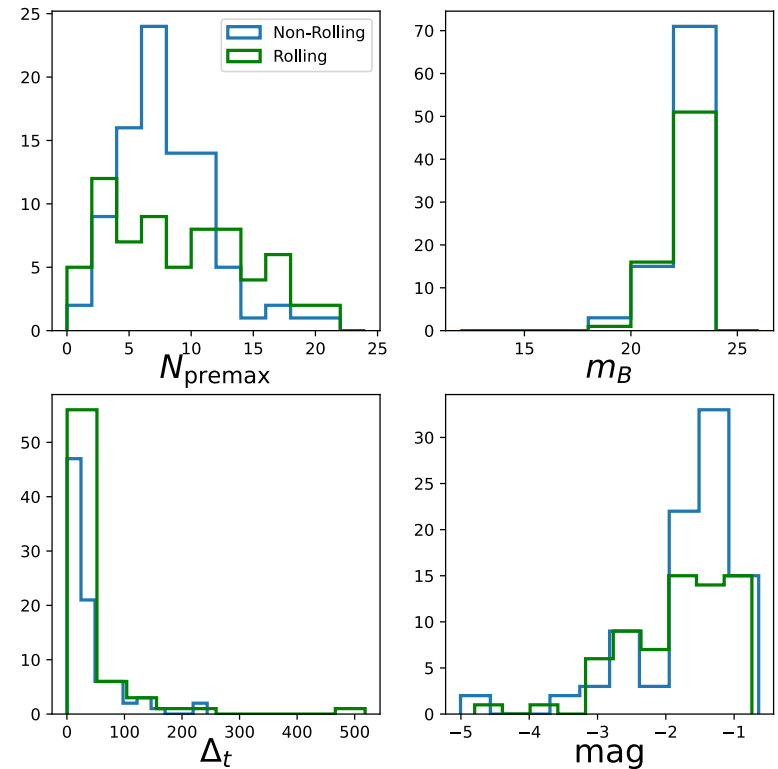
Baseline v3.0 cadence

Many 10's expected per year: ~ 20 with large Δt

Rolling has fewer \rightarrow denser sampling



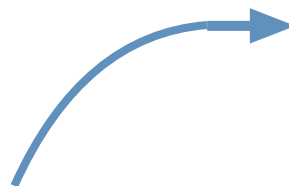
Several with early sampling for discovery
Bright for 4m spec classification
Long time delays for cosmology



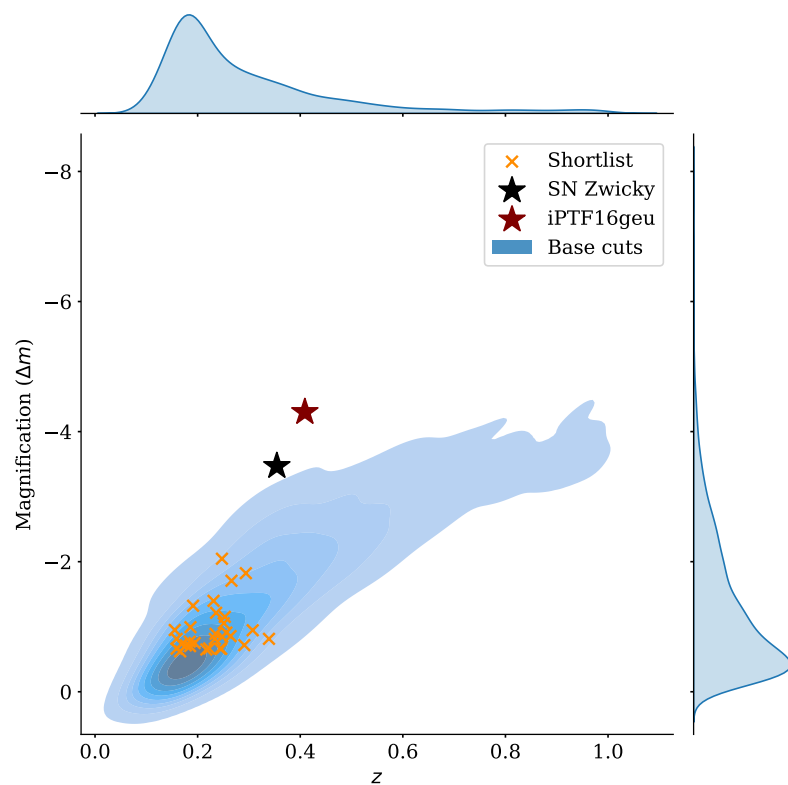


ZTF archive search for lensed SNe

- Systematic search for strongly lensed SNe Ia in the ZTF archive.
- Applying cuts based on simulations.
 - Cuts on photo-z, distance to host, peak absolute magnitude, SN and host colours.
- 31 930 alerts \rightarrow 30 candidates.
- Paper out this year!



Magnification vs. z for the best 7000 candidates (base cuts), and the shortlist candidates.

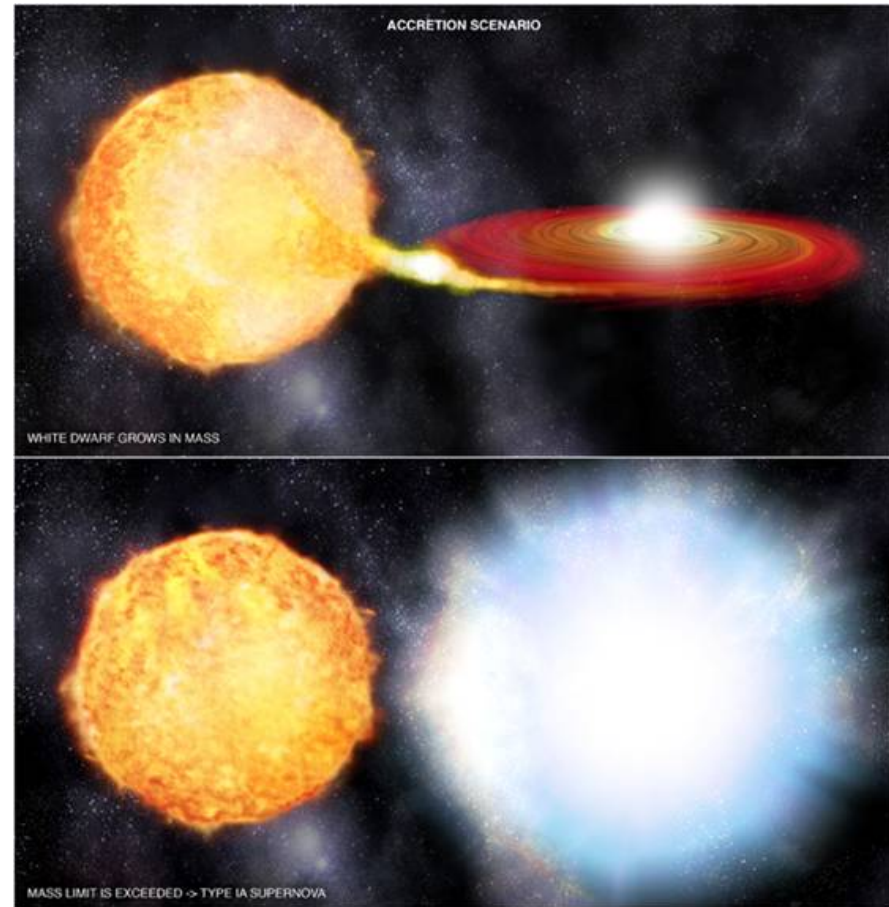




What are Type Ia supernovae?

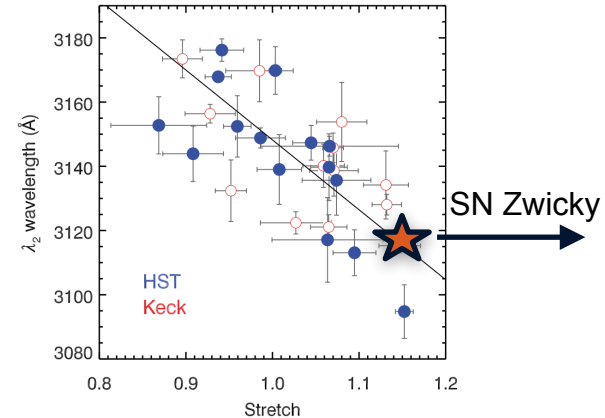
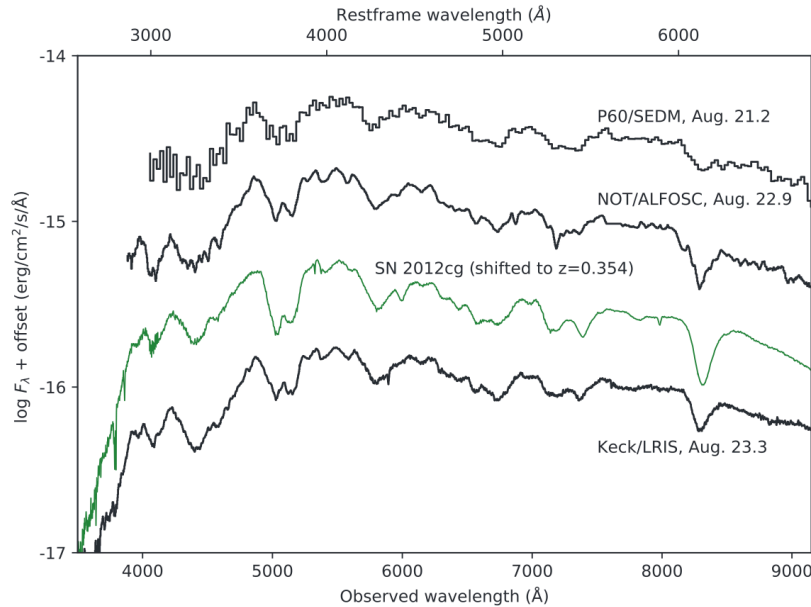
18

NOT standard but calibratable -> small scatter, reduce lensing uncertainties



dark energy systematics -> lensing helps study high-z SN physics

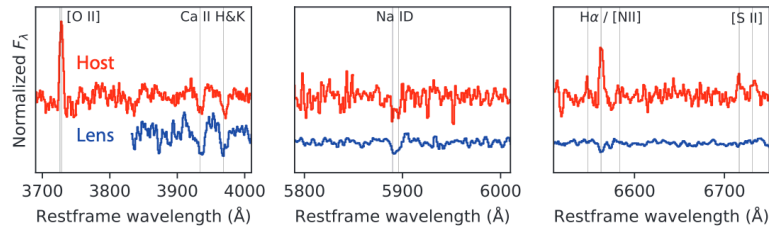
Spectroscopy of lensed SNe



Comparison to low- z SNe sample from Maguire +2012

SN Zwicky spectra compared to local SN 2012cg

Johansson, SD+, in prep



No signs of cosmic spectroscopic evolution!

JWST Cycle 2 NIRCam + NIRSpec proposal
Nebular observations of SN Zwicky

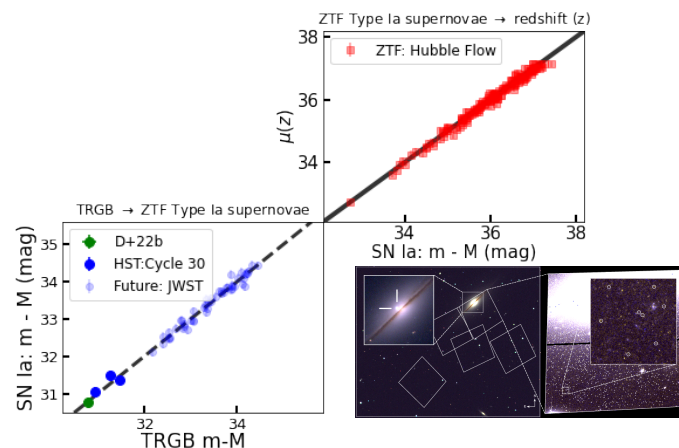
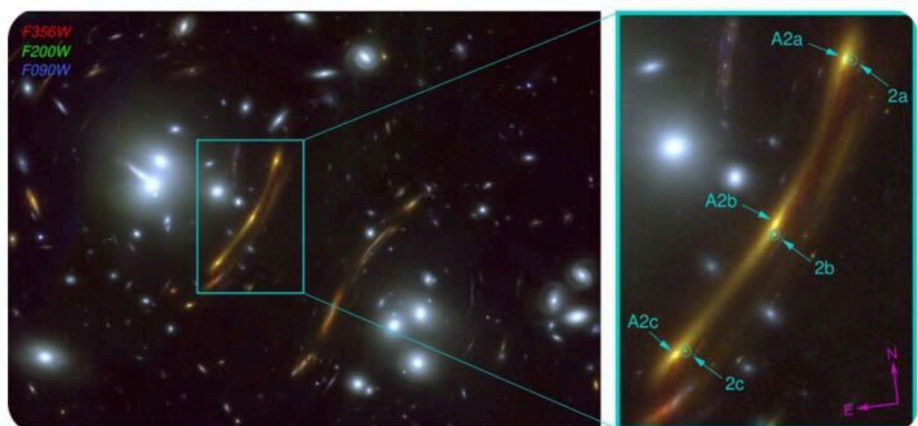
SN Hope



Discovery in NIRCAM obs of PLCK G165.7+67 ; PEARLS program

Cluster lens, long expected time-delay

Triply imaged SN Ia at $z \sim 2$; Follow-up with DDT ongoing

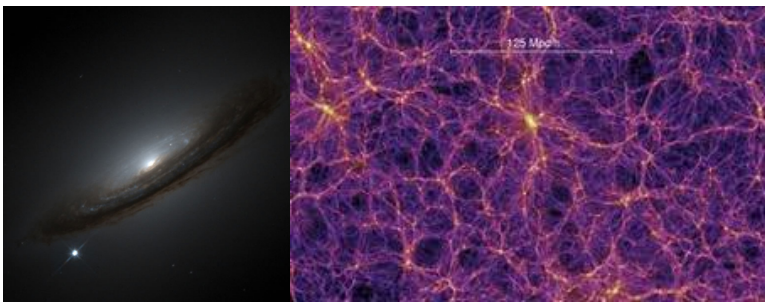


Strongly Lensed Supernovae

- Discovery with Vera C. Rubin Observatory
- Analyses with JWST IFU data
- Novel methods with current surveys

Local Hubble Constant

- Uniform, large Type Ia supernova datasets
- New distance indicators



Structure growth with SNe

- Bulk flows from local superstructures
- σ_8 with Type Ia supernovae