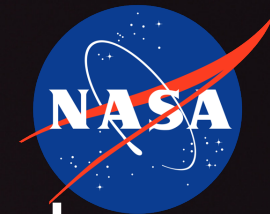




JSI JOINT SPACE-SCIENCE INSTITUTE



Discovering **rare transients** and
multi-messenger **counterparts** with the
Zwicky Transient Facility

Igor Andreoni
Neil Gehrels Fellow

Oskar Klein Center
October 17, 2023

Image credit: DESY

The transient sky



Walter Baade

“The modern era of transients with controlled cadence and a physics-based enquiry began with F. Zwicky and W. Baade.”



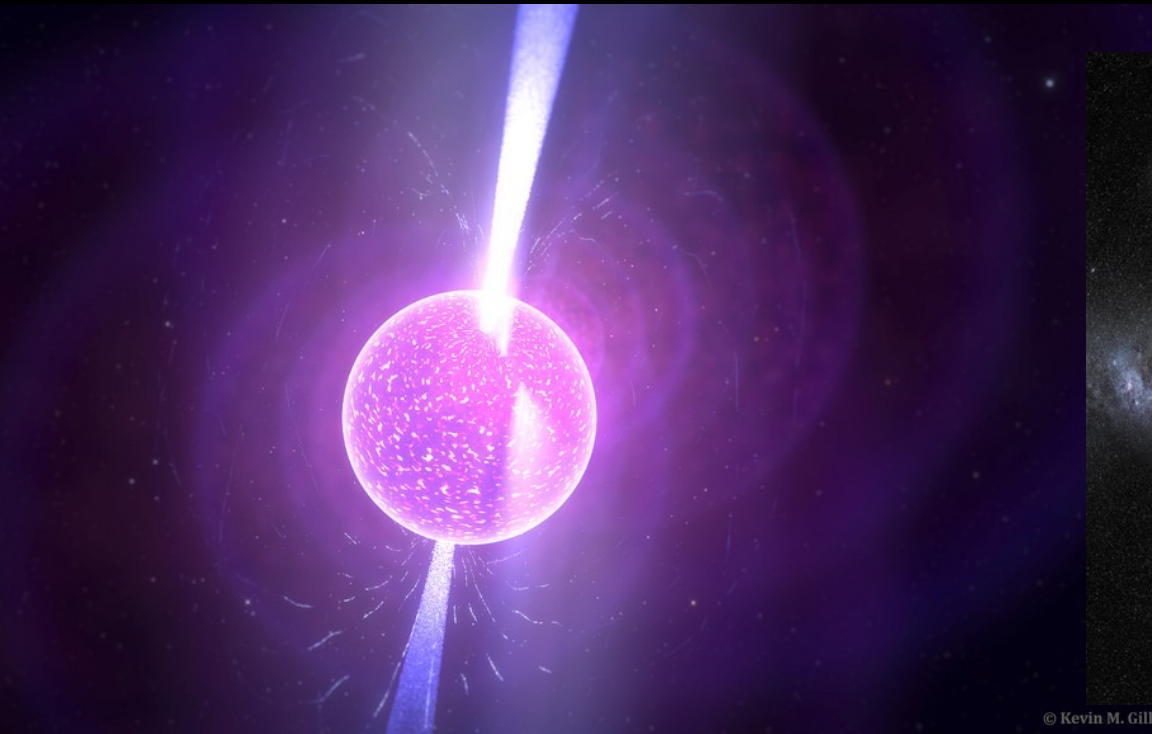
Fritz Zwicky

Kulkarni (2012), “*Cosmic Explosions*”

Death of massive stars

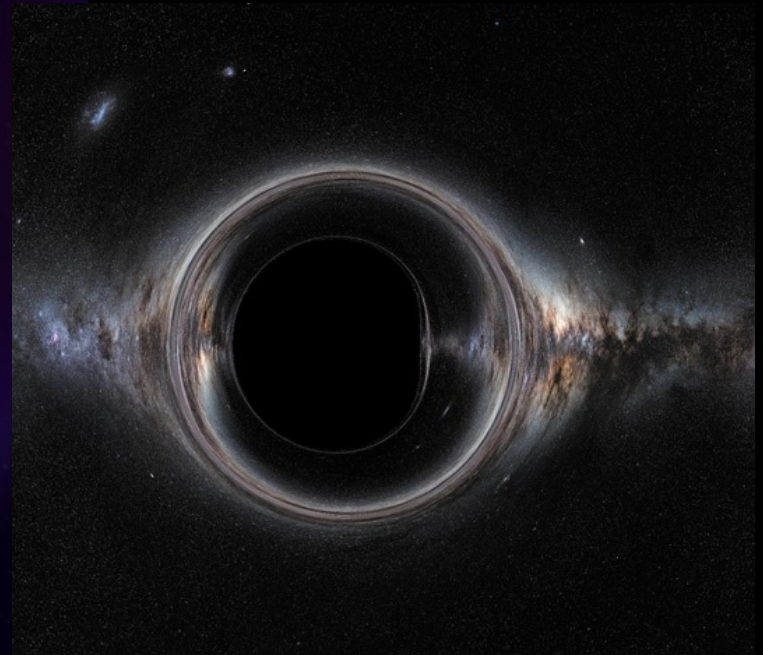
What is left **after the explosion?**

Neutron star



© Kevin M. Gill

Black hole



credit: ESO

Electromagnetic counterparts to Gravitational Waves

Gamma-ray burst

Binary neutron star and
neutron star–black hole systems
are GW multi-messenger sources

Tidally stripped ejecta
(low Y_e , high neutron content)

Post-merger ejecta
(broad range of neutron content)

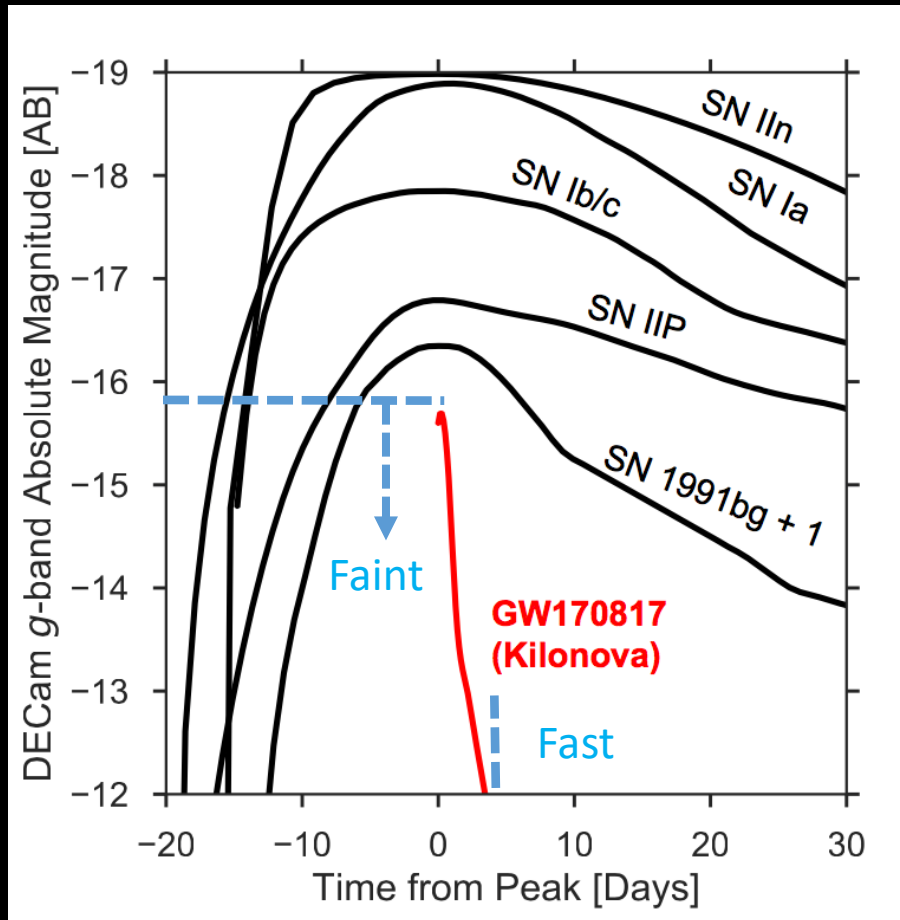
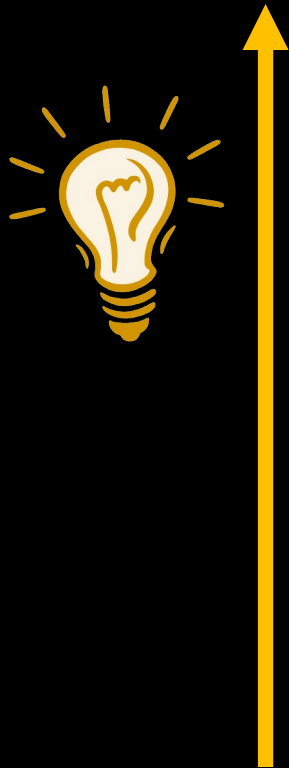
Kilonova
(optical/IR)

Gravitational
Waves

Reviews e.g.: Metzger 2019, Nakar 2019, Margutti & Chornock 2021

Image credit: NASA

Kilonova: faster & fainter than supernovae



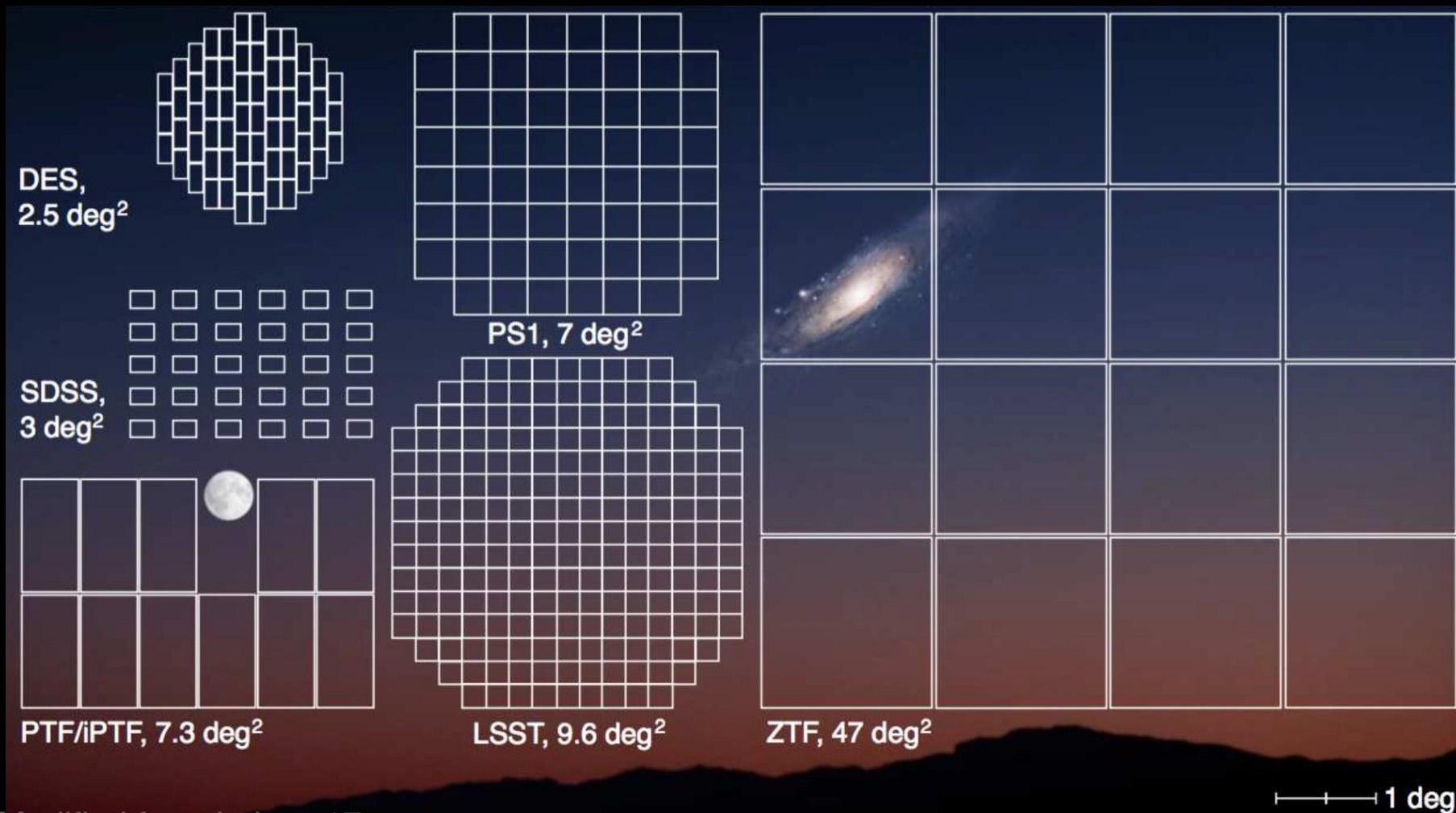
modified from
Andreoni+2018



$$\text{Magnitude} = -2.5 * \log_{10}(\text{flux}) + C$$

Faint + Fast = hard to catch!

Zwicky Transient Facility



Modified from Laher+17

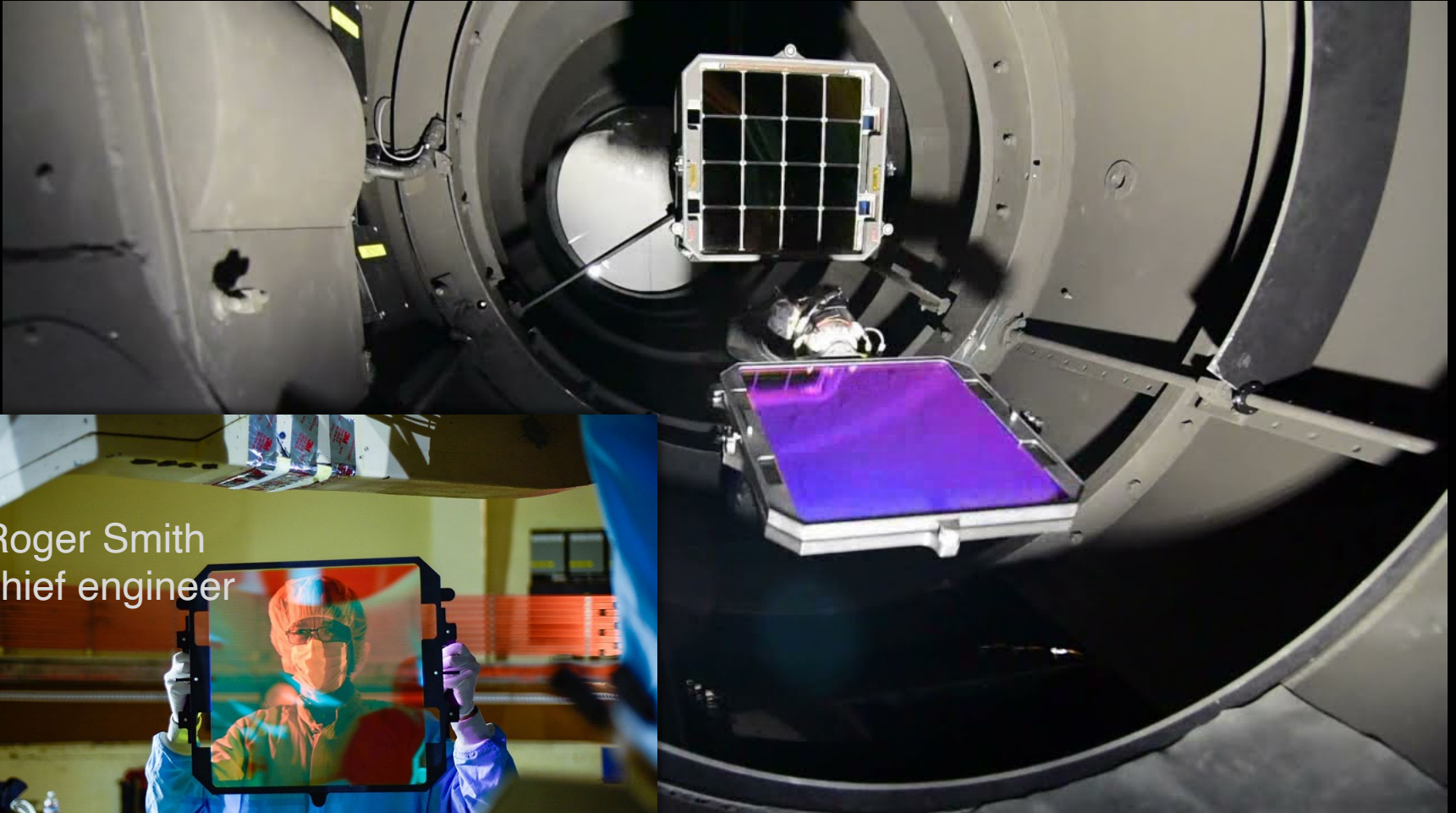
Zwicky Transient Facility

Zwicky Transient Facility $g < 20.5$ mag



Modified from Laher+17

Zwicky Transient Facility



48-inch Samuel Oschin telescope at Palomar

October 17, 2023



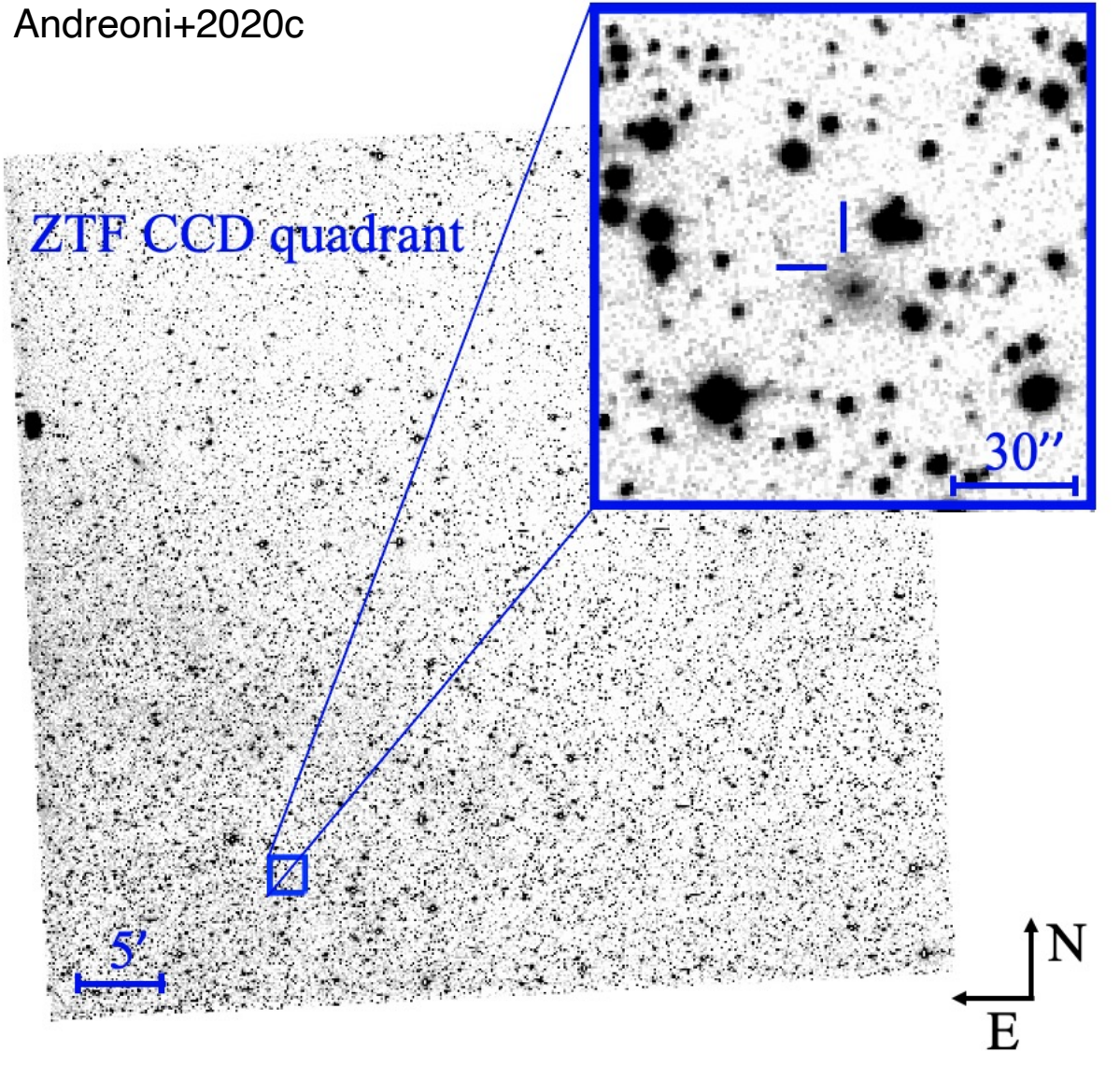
JSI



Igor Andreoni

Zwicky Transient Facility (ZTF)

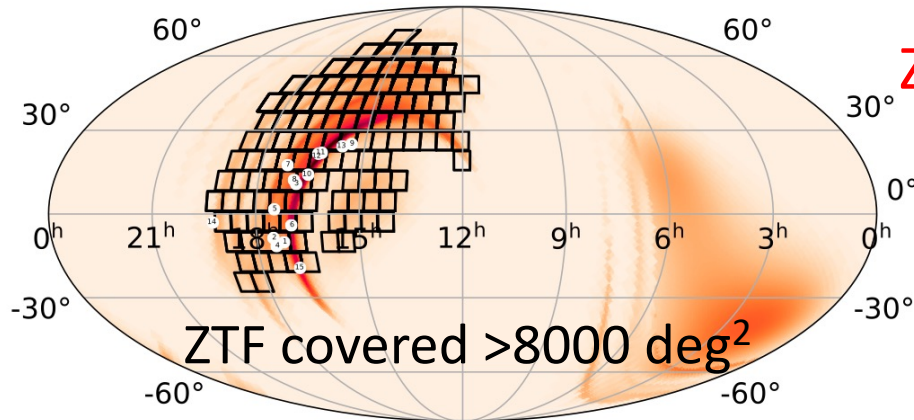
Andreoni+2020c



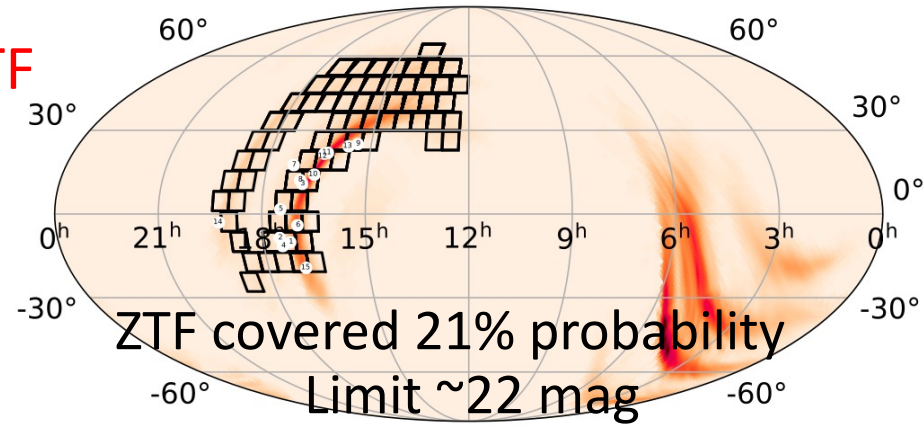
ZTF follow-up of the NS-NS merger GW190425

BAYESTAR

LaInference



ZTF

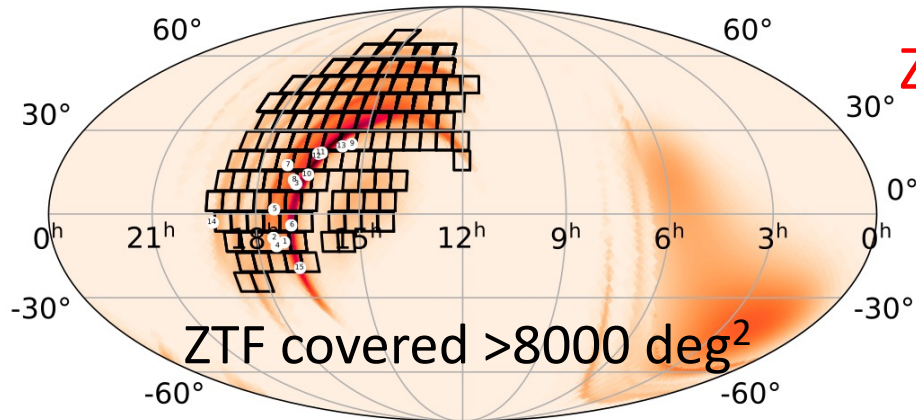


15 promising candidates
None was a kilonova

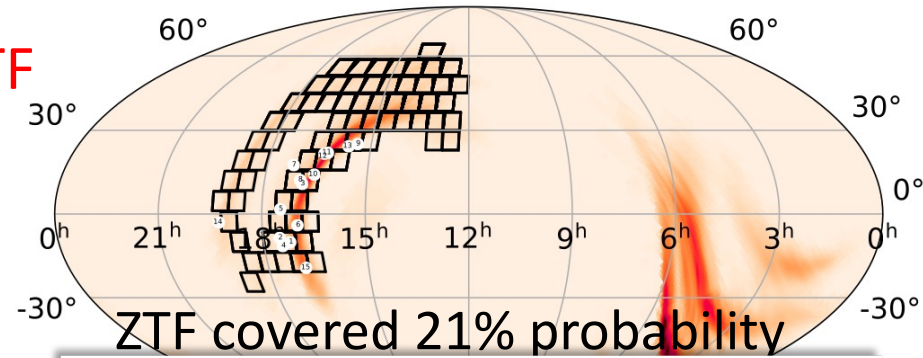
Coughlin, ..., IA et al. (2019), ApJL, 885, 1, L19

ZTF follow-up of the NS-NS merger GW190425

BAYESTAR

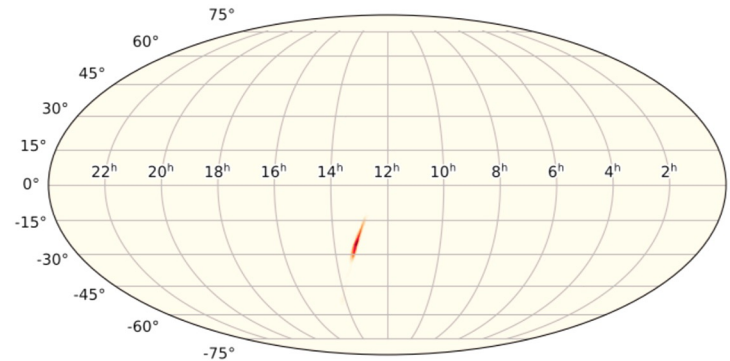


LaInference



15 promising candidates
None was a kilonova

GW170817



Coughlin, ..., IA et al. (2019), ApJL, 885, 1, L19

ZTF follow-up of 13 NS mergers during O3

For ZTF follow-up of neutron star mergers in O3 please see:

Second binary NS merger follow-up

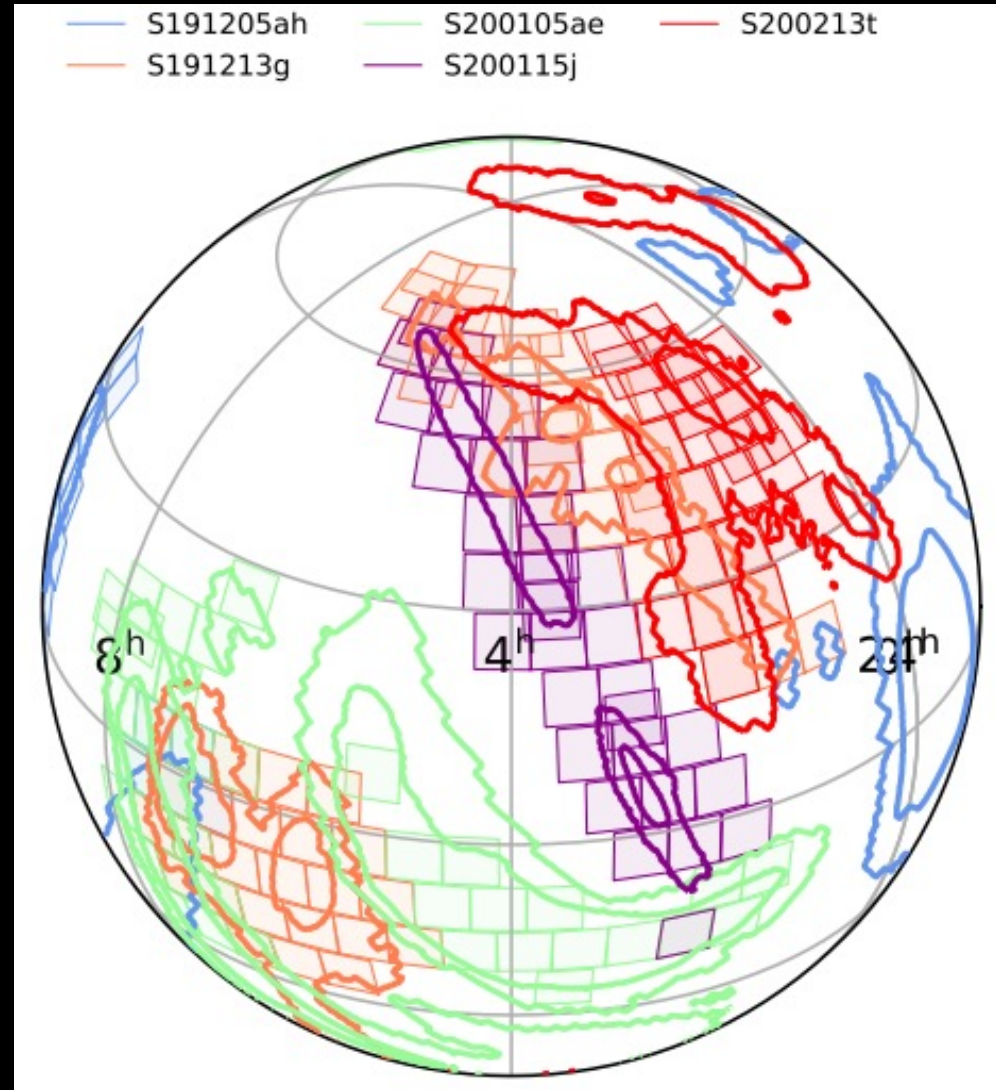
Coughlin et al. (2019),
ApJL, 885, 1, L19

Constraints on NS-BH merger

Anand & Coughlin et al. (2021),
Nature Astronomy, 5, 46

Kilonova luminosity function

Kasliwal et al. (2020),
ApJ, 905, 2, 145

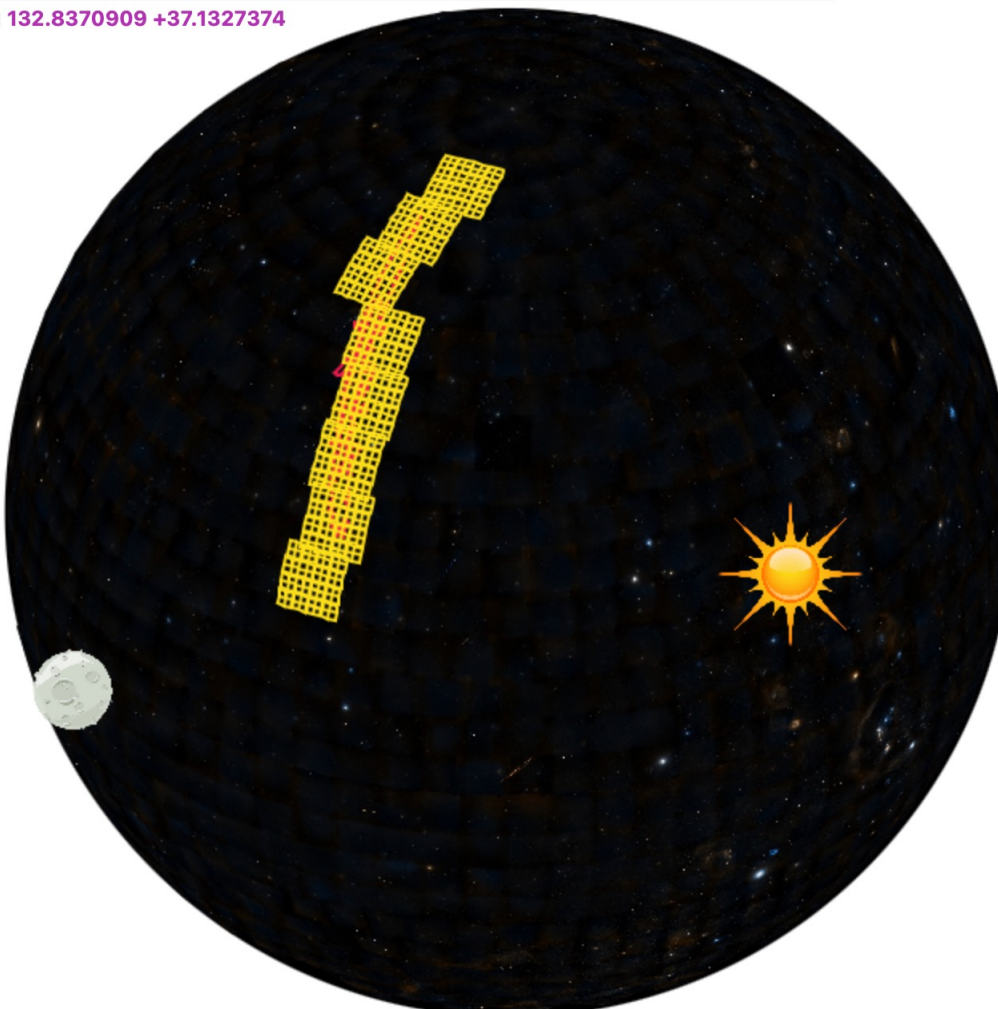


Modified from Laher+17

ZTF follow-up of GW events during O4

Preliminary 2023-06-27 01:53:52 Preliminary 1 2023-06-27 01:54:00 Initial 2023-06-27 01:54:00 Update 2023-06-27 03:55:07.176219

J2000d



Follow-Up

- Instruments
- ZTF
- GRB Coverage

Sources

- Galaxies
- XRT Sources

Visit:
treasuremap.space

Keep an eye on
[NASA GCN circulars](#)

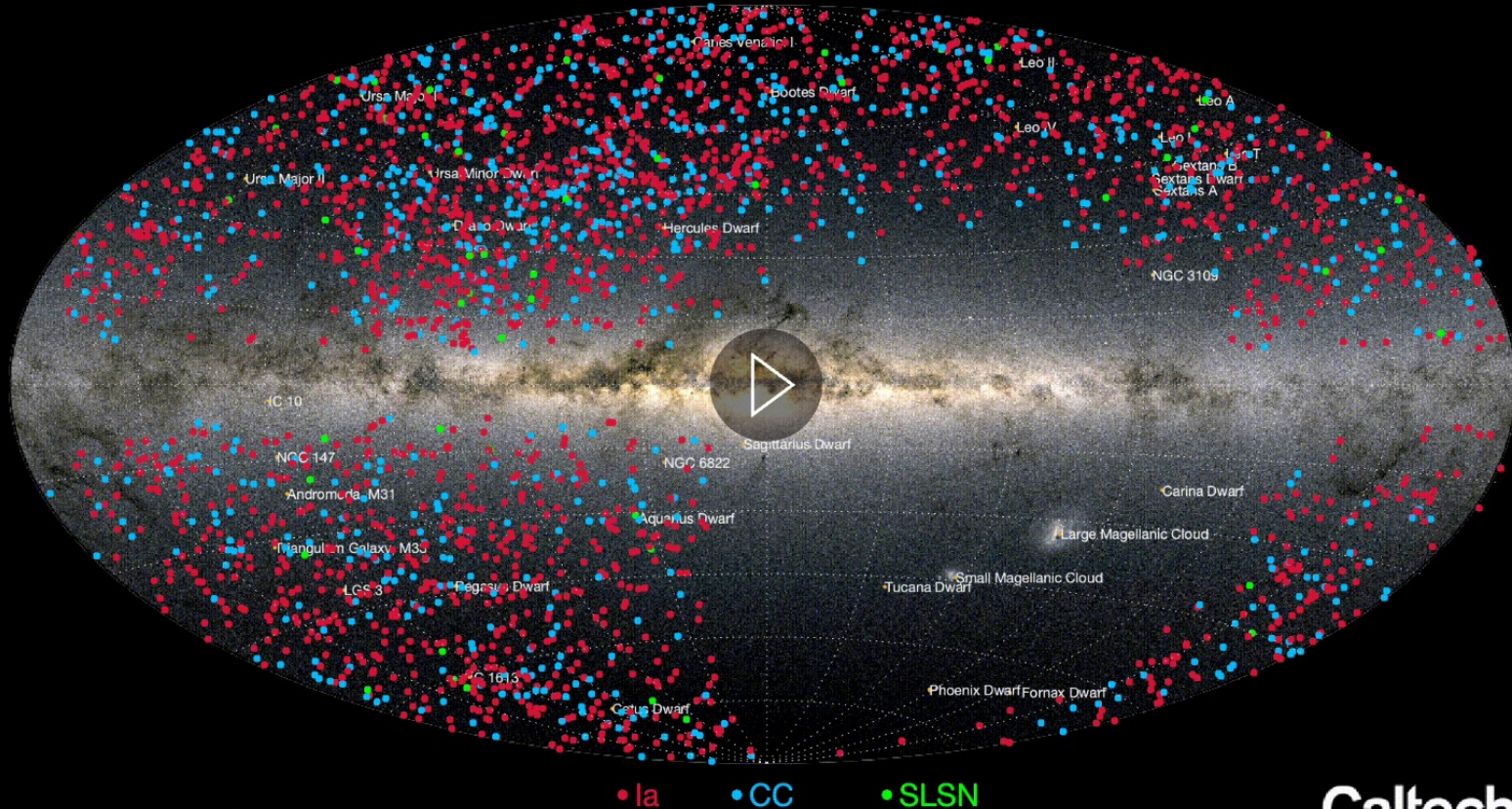
FoV: 180°

Un-triggered ZTF transient searches

Supernovae Classified:

2836

Date: 23-Mar-2020



Underlying image credit: ESA/Gaia/DPAC

Caltech

DR 19: **51 million** images, **787 billion** source detections

October 17, 2023

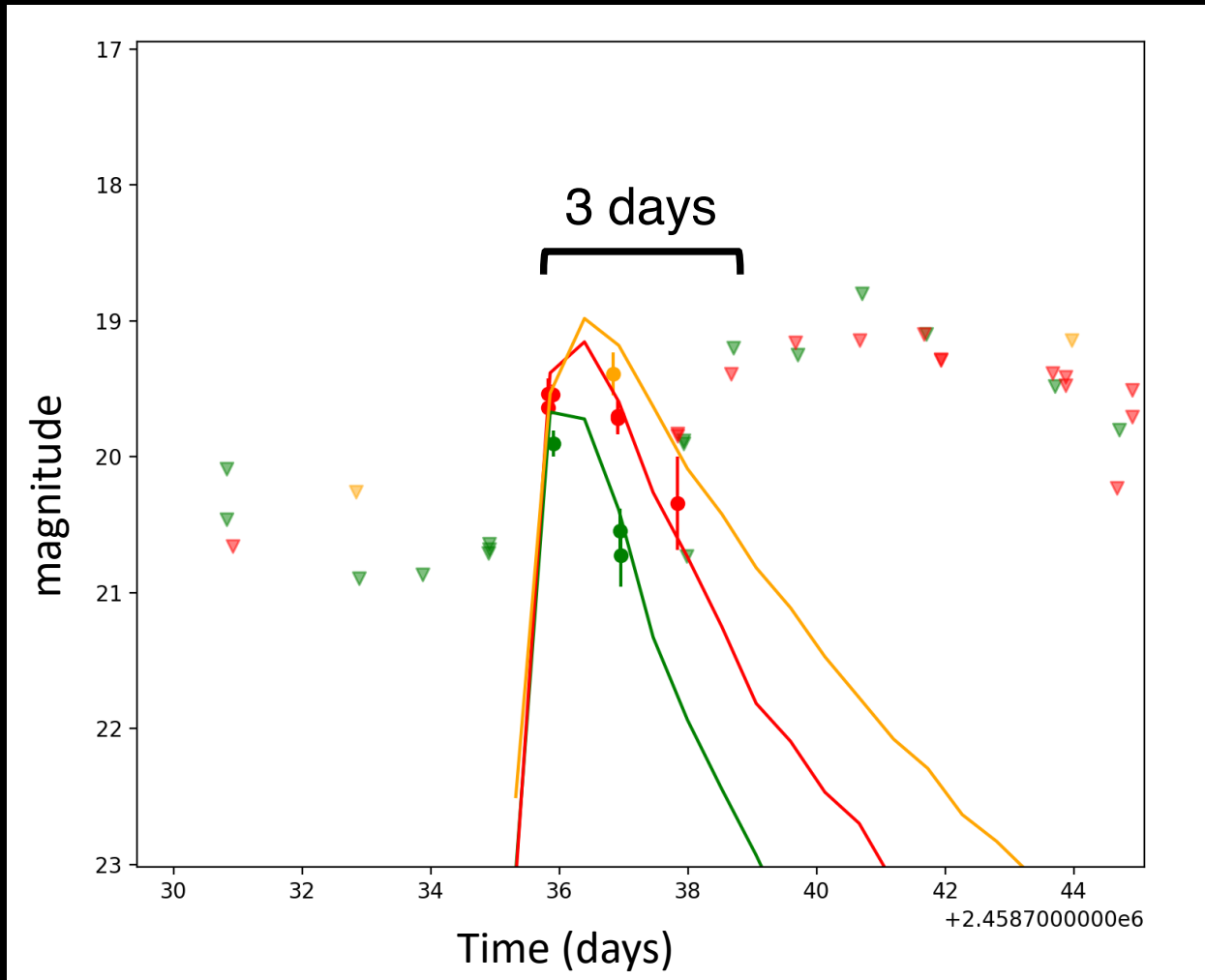


JSI



Igor Andreoni

Are there Kilonovae in ZTF data?



Simulated distant kilonova in ZTF data (plot by: Sagués-Carracedo @OKC)

ZTF Realtime Searching and Triggering

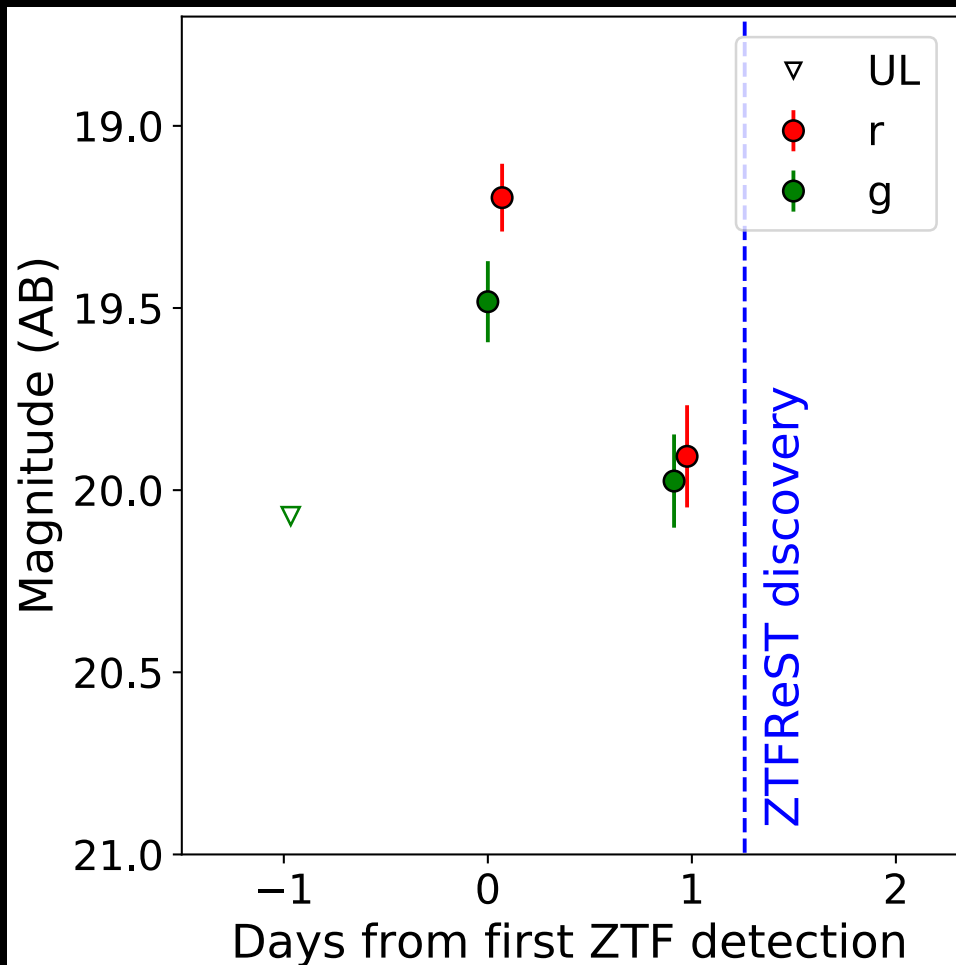


growth-astro/ztfrest



Michael Coughlin

Andreoni & Coughlin et al. (2021), ApJ, 918, 2, 63



Near **real-time** implementation of the search methods used in Andreoni et al. (2020d)

Supernova shock cooling
~ a dozen

Serendipitous GRB afterglows

- 5 with GRB association (long)
- **5 confirmed, un-triggered** afterglows w/o GRB association

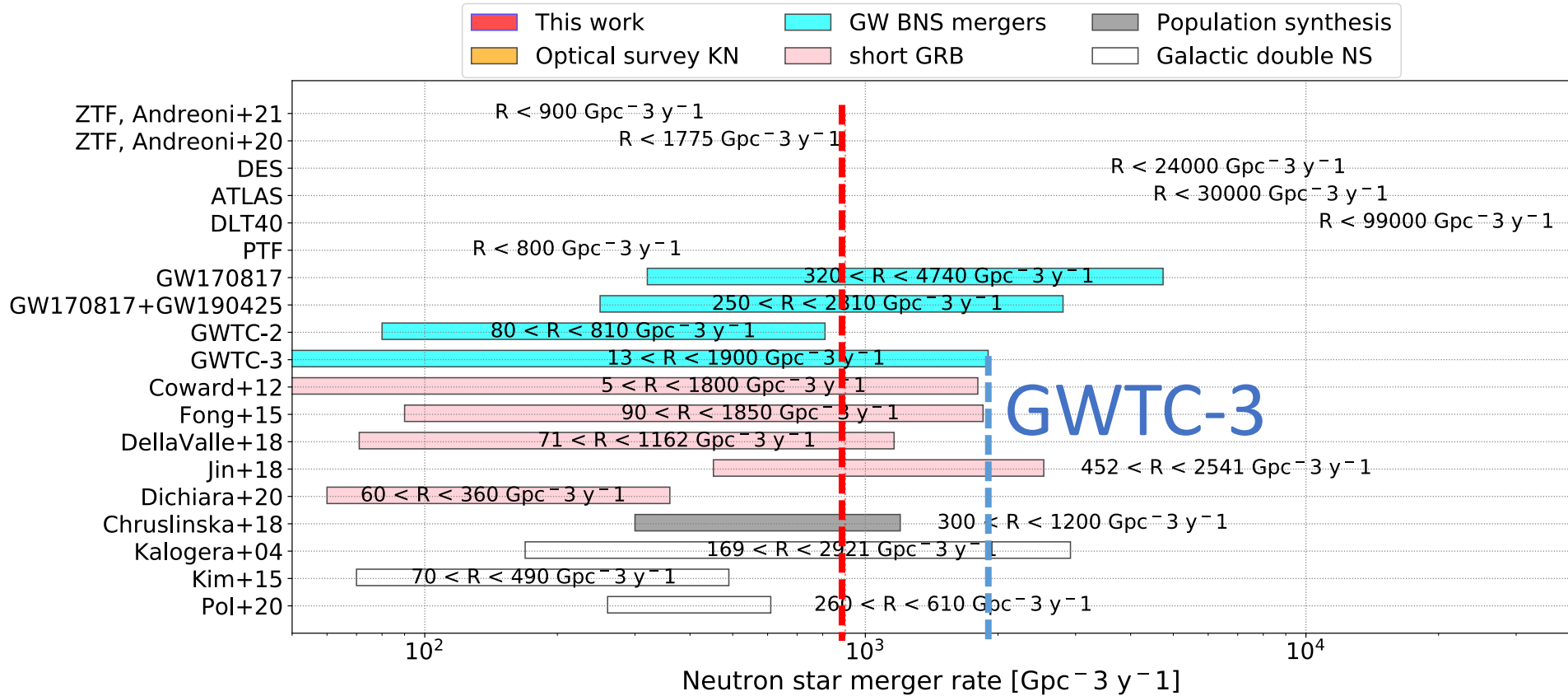
see also Ho+2020,2021,2022;
Andreoni+2020d

Kilonovae
Still waiting...

Constraints on Kilonova and neutron star merger rates



Constraints on Kilonova and neutron star merger rates

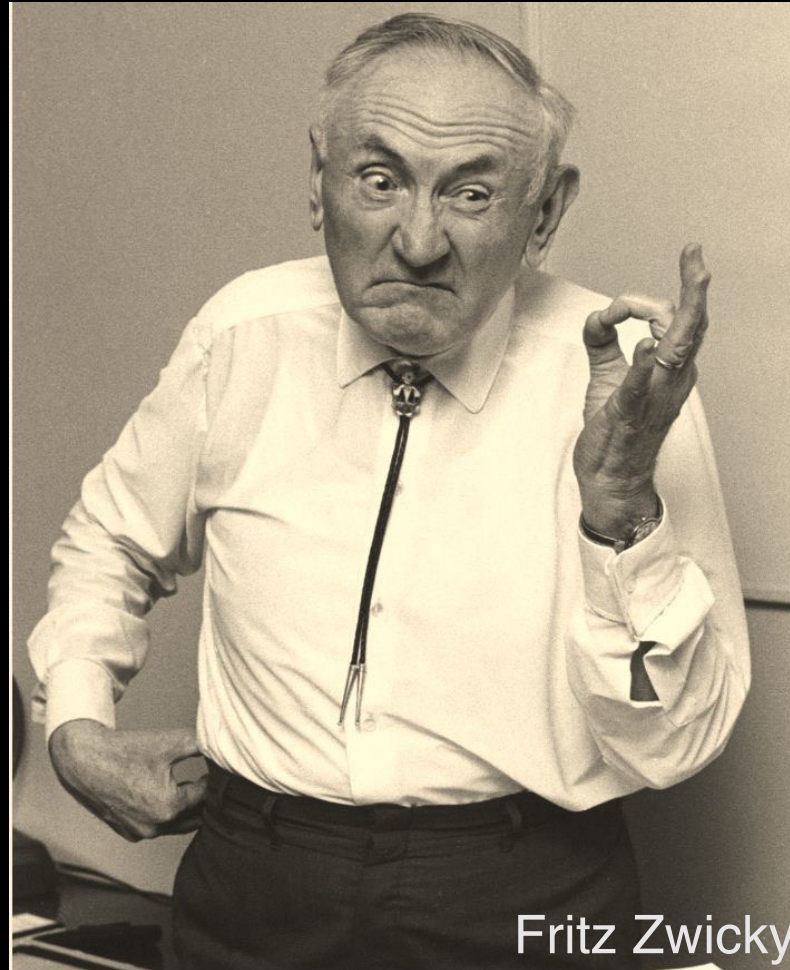


Andreoni & Coughlin et al. (2021), ApJ, 918, 2, 63

Model grid in Andreoni et al. (2020d), ApJ, 904, 2, 155

ZTF constrained the rate of GW170817-like kilonovae to be $R < 900 \text{ Gpc}^{-3} \text{ y}^{-1}$

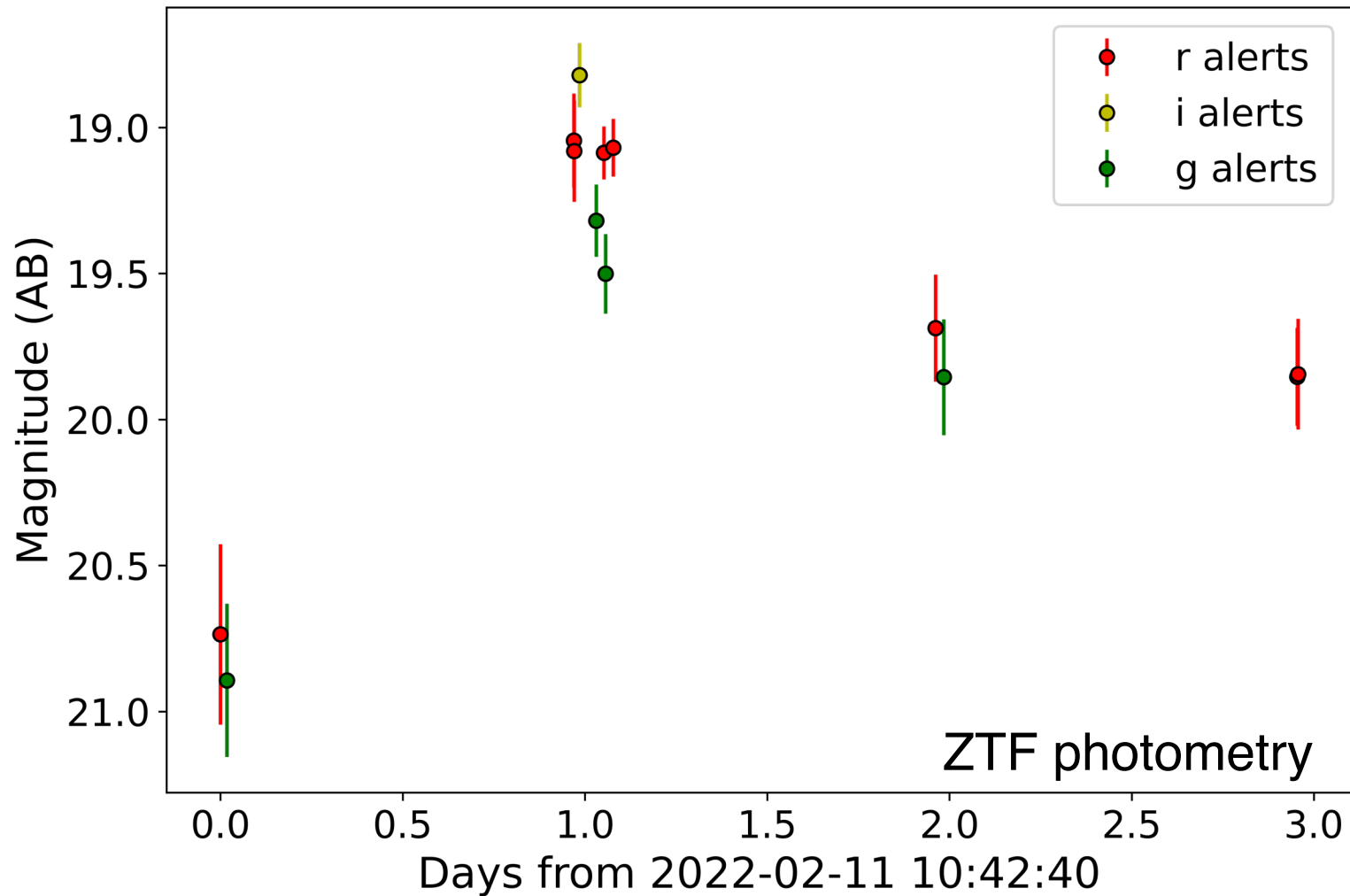
Ready for the unexpected



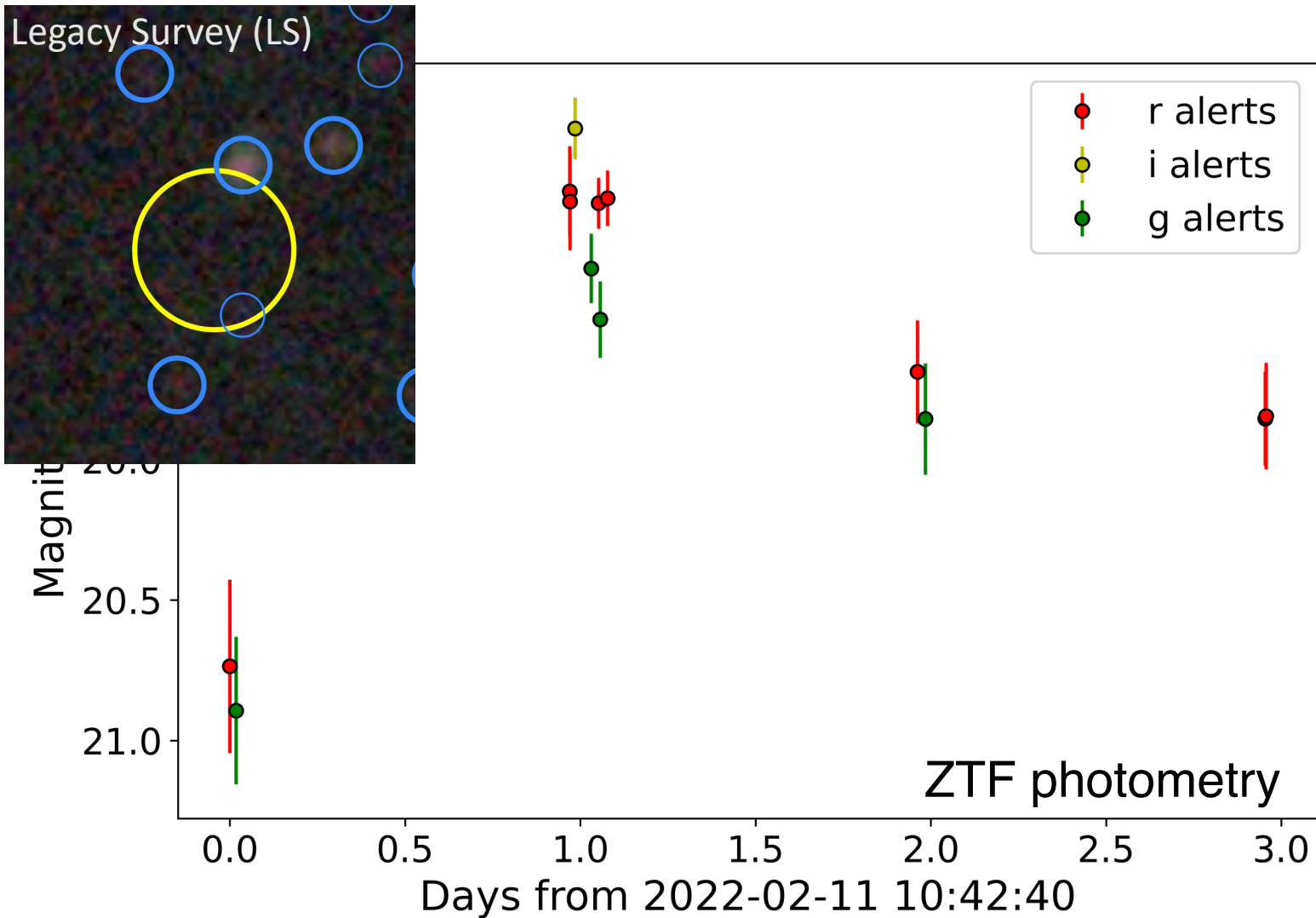
Fritz Zwicky

Andreoni & Coughlin et al. (2022), *Nature*, 612, 7940

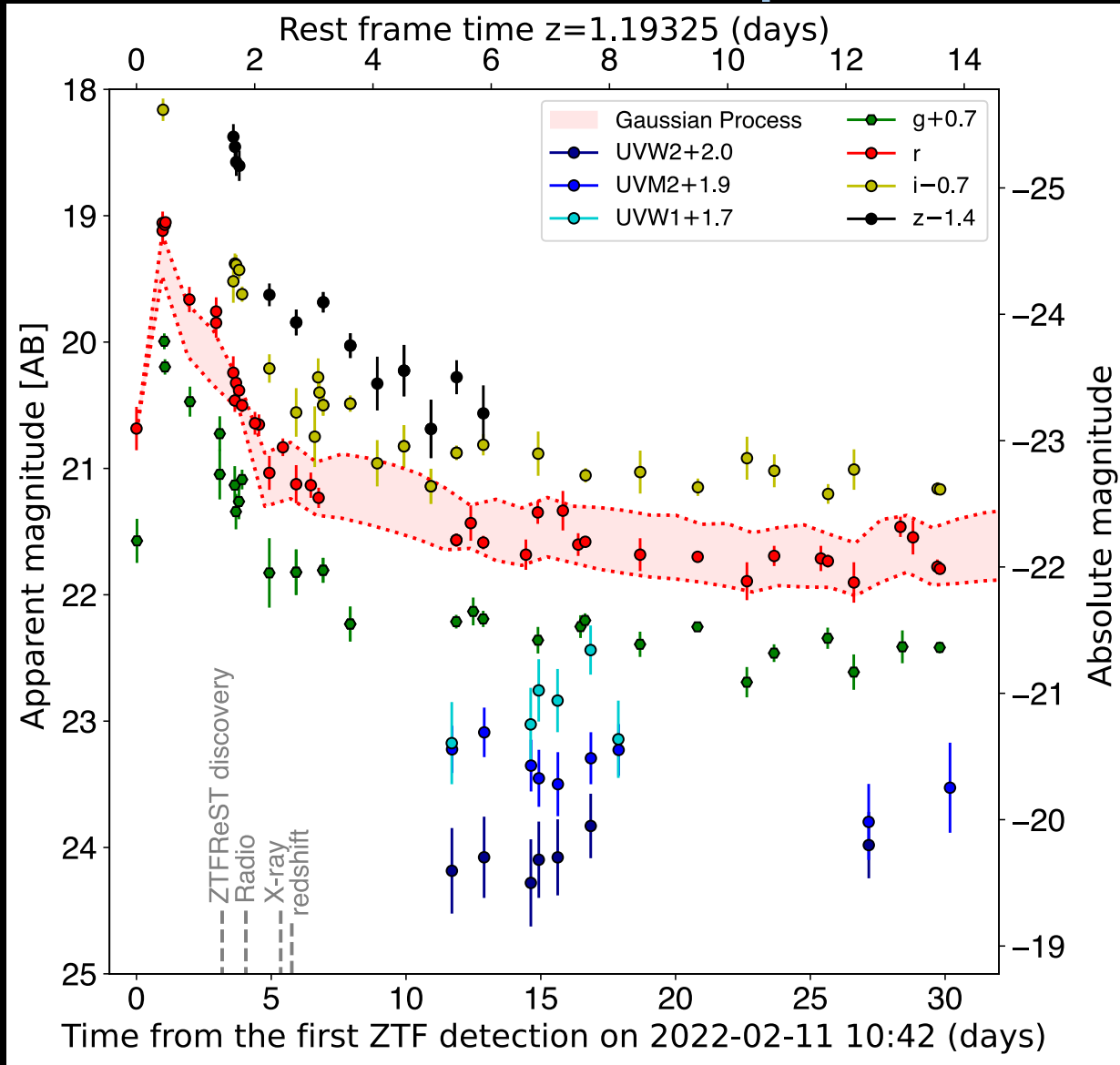
AT2022cmc: Discovery



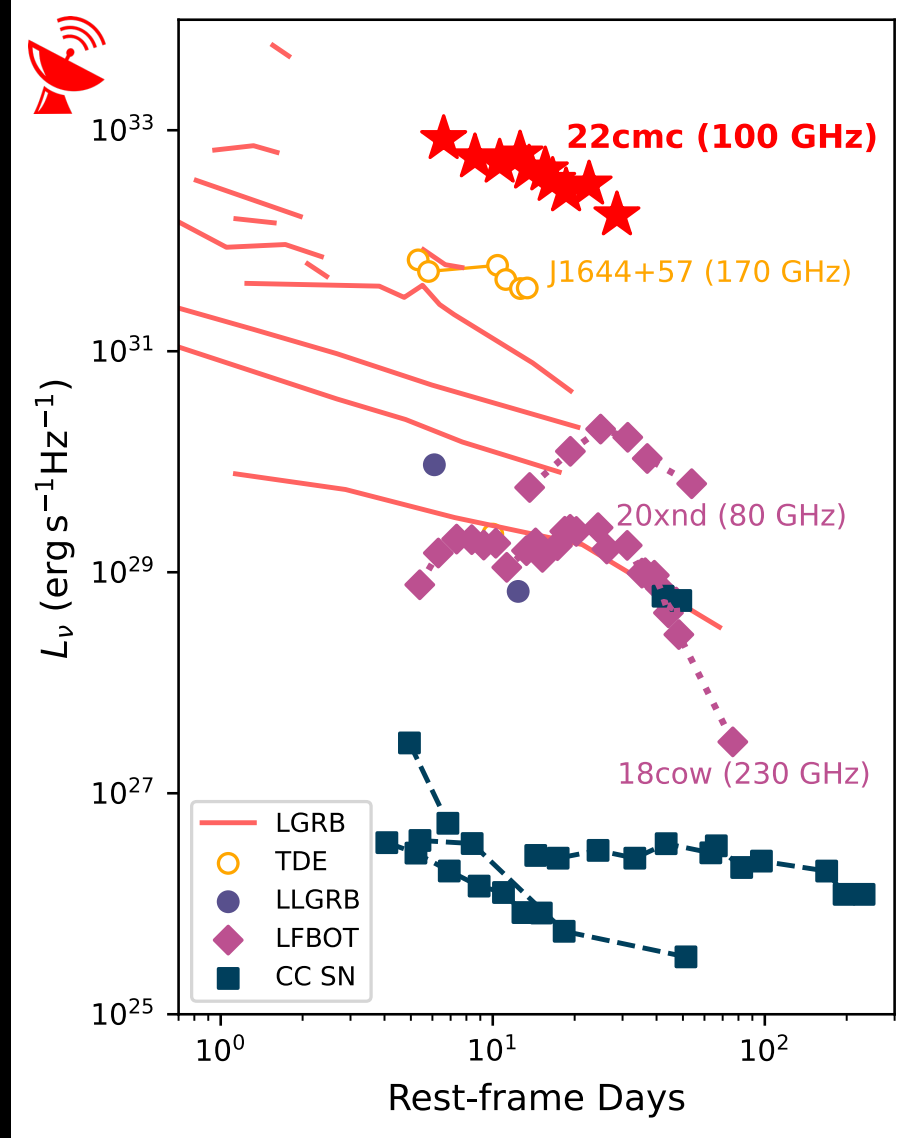
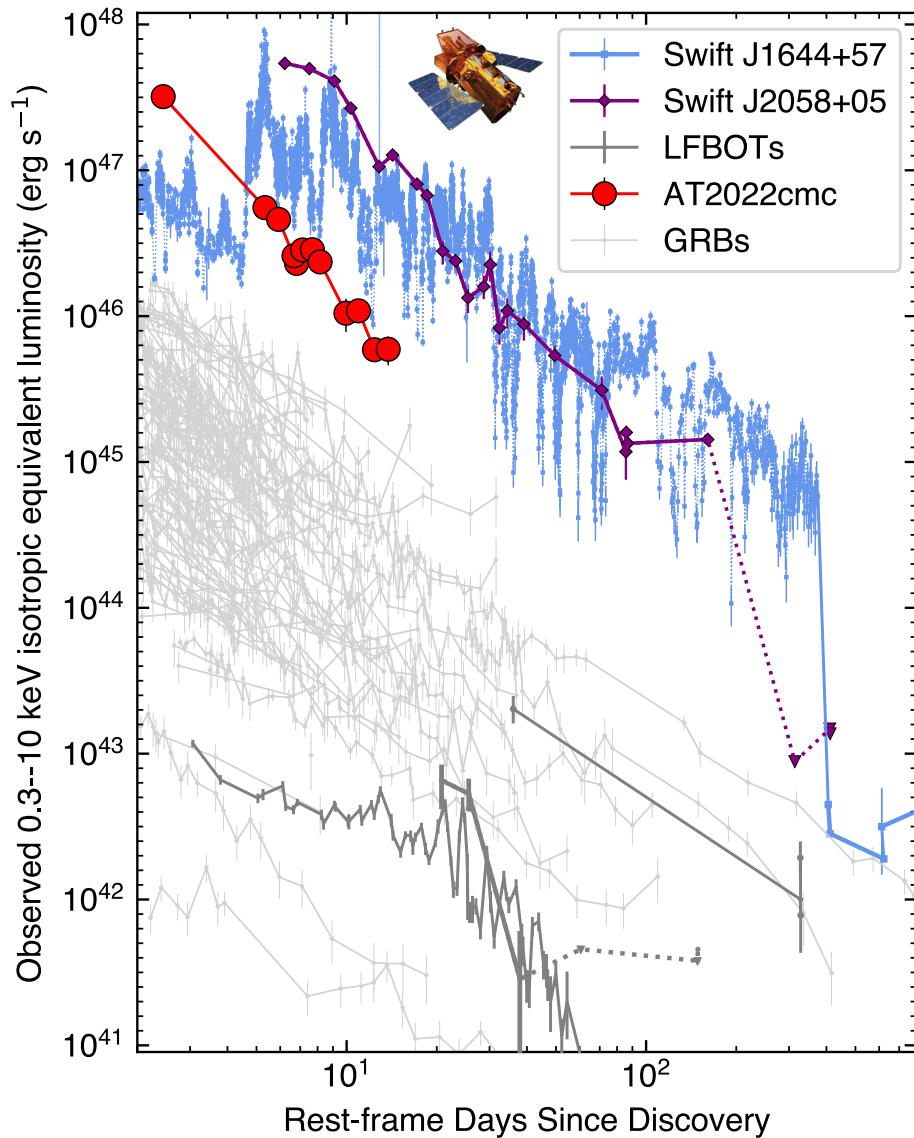
AT2022cmc: Discovery



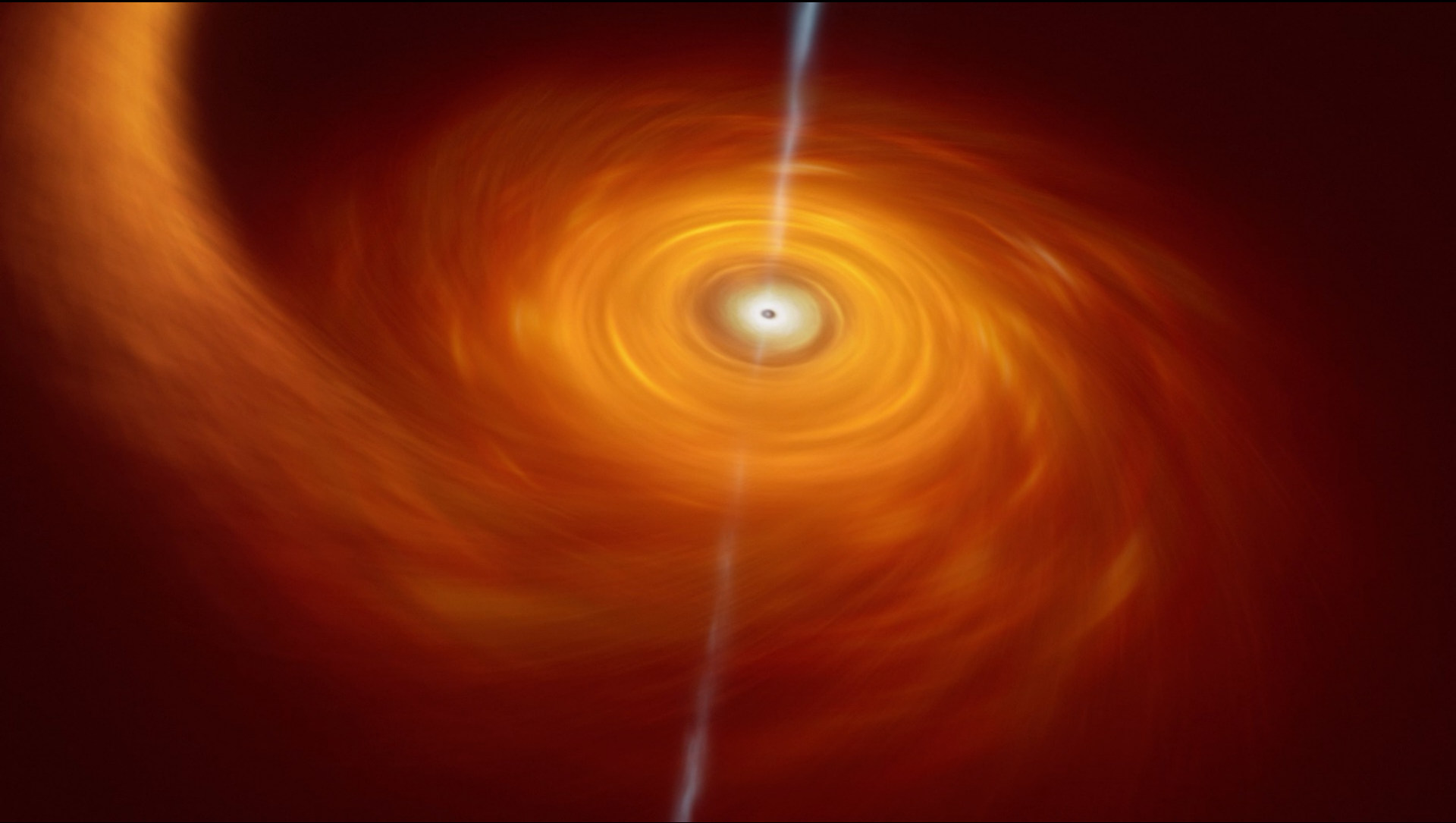
AT2022cmc: UV/Optical/nIR



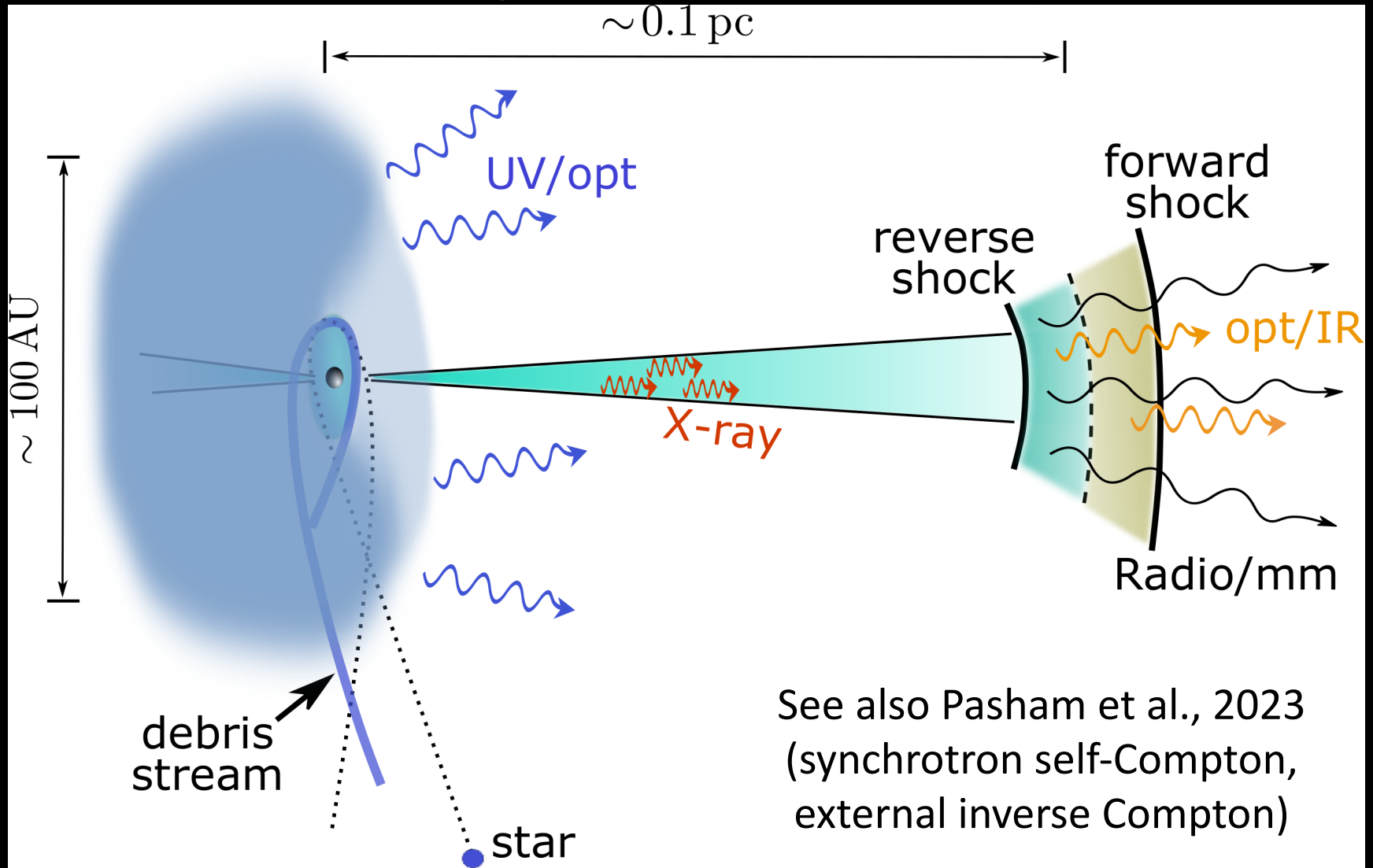
AT2022cmc: X-rays and radio/mm



Jetted Tidal Disruption Events



AT2022cmc: a jetted tidal disruption event



AT2022cmc: The **furthest** TDE ever observed - the **first** jetted TDE identified by an **optical** survey

Summary/Take-away

Many **open questions** remain about kilonovae and neutron star mergers

Systematic ZTF follow-up of LIGO-Virgo-KAGRA triggers **science** (individual events, luminosity function) even without a positive counterpart detection – waiting for a kilonova in O4!

Systematic searches in ZTF constrained the rate of GW170817-like kilonovae to be $R < 900 \text{ Gpc}^{-3} \text{ y}^{-1}$, which can tell us something about the neutron star merger rate

Un-triggered searches for fast transients unveiled the first optically-discovered **jetted TDE**, which ZTF confirmed to be $\sim 1\%$ of the TDE population