

Problem sheet / winter school 2024 / intro cosmology

Problem 1:

geodesic equation: $\frac{d^2 x^\mu}{d\lambda^2} = -\Gamma_{\kappa\rho}^\mu \frac{dx^\kappa}{d\lambda} \frac{dx^\rho}{d\lambda}$

(a) $ds^2 = -dt^2 + a^2 \delta_{ij} dx^i dx^j \Rightarrow g_{\mu\nu} = \begin{pmatrix} -1 & & & \\ & a^2 & & \\ & & a^2 & \\ & & & a^2 \end{pmatrix} \Rightarrow g^{\mu\nu} = \begin{pmatrix} -1 & & & \\ & \frac{1}{a^2} & & \\ & & \frac{1}{a^2} & \\ & & & \frac{1}{a^2} \end{pmatrix}$

$$\Gamma_{\alpha\beta}^\mu = \frac{1}{2} g^{\mu\lambda} (\partial_\alpha g_{\beta\lambda} + \partial_\beta g_{\alpha\lambda} - \partial_\lambda g_{\alpha\beta})$$

$$\Gamma_{\rho 0}^0 = \Gamma_{0\rho}^0 = \frac{1}{2} g^{00} (\partial_0 g_{\rho 0} + \partial_\rho g_{00} - \partial_0 g_{0\rho}) = 0$$

$$\Gamma_{00}^\mu = \frac{1}{2} g^{\mu\lambda} (\partial_0 g_{0\lambda} + \partial_0 g_{0\lambda} - \partial_\lambda g_{00}) = 0$$

$$\Gamma_{ij}^k = \frac{1}{2} g^{k\ell} (\partial_i g_{j\ell} + \partial_j g_{i\ell} - \partial_\ell g_{ij}) = 0$$

$$\Gamma_{ij}^0 = \frac{1}{2} g^{00} (\partial_i g_{0j} + \partial_j g_{0i} - \partial_0 g_{ij}) = +\frac{1}{2} \frac{1}{a^2} (a^2) \dot{g}_{ij} = +\frac{\dot{a}}{a} g_{ij} = H g_{ij}$$

$$\Gamma_{0i}^i = \frac{1}{2} g^{i\ell} (\partial_0 g_{j\ell} + \partial_j g_{0\ell} - \partial_\ell g_{0j}) = \frac{1}{2} \frac{1}{a^2} (a^2) \dot{g}^{i\ell} g_{j\ell} = \frac{\dot{a}}{a} \delta_\ell^i = H \delta_\ell^i$$

(b) $P^\mu = \frac{dx^\mu}{d\lambda} = (E, p^i)$

$$\frac{d^2 x^\mu}{d\lambda^2} = -\Gamma_{\kappa\rho}^\mu \frac{dx^\kappa}{d\lambda} \frac{dx^\rho}{d\lambda} \Rightarrow \frac{d}{d\lambda} P^\mu = -\Gamma_{\alpha\beta}^\mu P^\alpha P^\beta$$

$$\mu=0: \frac{dE}{d\lambda} = -a\dot{a} \delta_{ie} P^i P^e$$

$$\text{use } P_\mu P^\mu = g_{\mu\nu} P^\mu P^\nu = 0 \Rightarrow -(P^0)^2 + a^2 \delta_{ij} P^i P^j \Rightarrow E^2 = a^2 \delta_{ij} P^i P^j$$

$$\text{and } \frac{dE}{d\lambda} = \frac{dt}{d\lambda} \frac{dE}{dt} = E \frac{dE}{dt}$$

$$\Rightarrow E \frac{dE}{dt} = -\frac{\dot{a}}{a} E^2 \Rightarrow \frac{1}{E} \frac{dE}{dt} = -\frac{1}{a} \frac{da}{dt} \Rightarrow \frac{d}{dt} \ln(aE) = 0$$

(c) $\frac{d}{dt} \ln(aE) = 0 \Rightarrow E \propto a^{-1} \Rightarrow \lambda \propto a \Rightarrow z = \frac{\Delta\lambda}{\lambda} = \frac{\lambda_0 - \lambda_1}{\lambda_1} = \frac{a_0 - a_1}{a_1}$

w.l.o.g. $a_0 \equiv 1 \Rightarrow z = \frac{1 - a(t)}{a(t)}$

Problem 2:

$$T_{\nu}^{\mu} = \begin{pmatrix} -\dot{\xi} & & \\ & P & 0 \\ 0 & P & P \end{pmatrix}$$

$$\nabla_{\mu} T_{\nu}^{\mu} = \partial_{\mu} T_{\nu}^{\mu} + \Gamma_{\mu\lambda}^{\mu} T_{\nu}^{\lambda} - \Gamma_{\nu\mu}^{\lambda} T_{\lambda}^{\mu} \stackrel{!}{=} 0$$

$$\nu=0 \Rightarrow \partial_0 T_0^0 + \Gamma_{\rho 0}^0 T_0^{\rho} - \Gamma_{0\rho}^{\lambda} T_{\lambda}^{\rho} = 0$$

$$\Rightarrow -\dot{\xi} + \Gamma_{i0}^i (-\dot{\xi}) - \Gamma_{0i}^i T_i^i = 0$$

$$\Rightarrow -\dot{\xi} - 3H\dot{\xi} - 3HP = 0$$

$$\Rightarrow \underline{\dot{\xi} + 3H(\dot{\xi} + P) = 0}$$

(b) $\dot{\xi} + 3H(\dot{\xi} + P) = 0$

$$\Rightarrow \frac{d\xi}{dt} + 3\frac{\dot{a}}{a}\xi + 3\frac{\dot{a}}{a}P = 0$$

$$\stackrel{a \neq 0}{\Rightarrow} \frac{d}{dt}(a^3 \xi) + P \frac{d(a^3)}{dt} = 0$$

$$\Rightarrow \frac{d}{dt}(V\xi) + P \frac{dV}{dt} = 0$$

$$\Rightarrow \underline{dU + PdV = 0}$$

(c) $P(t) = w\xi(t)$

$$\Rightarrow \dot{\xi} + 3H(1+w)\xi = 0$$

$$\Rightarrow \dot{\xi} + 3\frac{\dot{a}}{a}(1+w)\xi = 0$$

$$\Rightarrow \frac{d\xi}{\xi} = -3(1+w) \frac{da}{a}$$

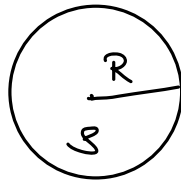
$$\Rightarrow d \ln \xi = -3(1+w) d \ln a$$

$$\Rightarrow \ln \xi = -3(1+w) \ln a + C_0$$

$$\Rightarrow \xi = \exp(C_0) \exp(-3(1+w) \ln a)$$

$$\Rightarrow \xi = \xi_0 \left(\frac{a}{a_0}\right)^{-3(1+w)}$$

Problem 3:



(a)

grav. force

$$\ddot{R} = -\frac{GM}{R^2} \quad \text{where} \quad M = \frac{4\pi}{3} R^3 \rho = \text{const}$$

$$\Rightarrow \dot{R}\ddot{R} = -GM \frac{\dot{R}}{R^2} \Rightarrow \frac{1}{2}\dot{R}^2 = +GM \frac{1}{R} + E$$

E is energy per unit mass of particle.

$$\Rightarrow \dot{R}^2 = +\frac{8\pi G}{3} \rho R^2 + 2E$$

identify: $R = R_0 a(t)$ & $E = -\frac{K}{2} \Rightarrow$ $\left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3} \rho - \frac{K}{a^2 R_0^2}$
Friedmann

(b) transformation $a dy = dt$

$$\Rightarrow \dot{a} = \frac{da}{dt} = \frac{da}{dy} \frac{dy}{dt} = a' \frac{1}{a} (= H)$$

$$\Rightarrow \ddot{a} = \frac{d}{dt} \left(\frac{a'}{a} \right) = \frac{d}{dy} \left(\frac{a'}{a} \right) \frac{dy}{dt} = \left(\frac{a''}{a} - \left(\frac{a'}{a} \right)^2 \right) \frac{1}{a}$$

1. Friedmann: $\left(\frac{a'}{a}\right)^2 = \frac{8\pi G}{3} a^2 \rho$

2. Friedmann: $\frac{a''}{a} - \left(\frac{a'}{a}\right)^2 = -\frac{4\pi G}{3} a^2 (\rho + 3P)$

$$\Rightarrow \frac{a''}{a} = -\frac{4\pi G}{3} a^2 (-\rho + 3P)$$

(c) ansatz: $a(t) = \left(\frac{t}{t_0}\right)^\alpha \Rightarrow \dot{a} = \alpha \left(\frac{t}{t_0}\right)^{\alpha-1} \frac{1}{t_0} \Rightarrow \ddot{a} = (\alpha-1)\alpha \left(\frac{t}{t_0}\right)^{\alpha-2} \frac{1}{t_0^2}$

$H(t) = \frac{\dot{a}}{a} = \frac{\alpha}{t} \Rightarrow H_0 = \frac{\alpha}{t_0}$; moreover $S(t) = \rho_0 a^{-3(1+w)}$

Friedmann 1: $\frac{\dot{a}^2}{a^2} = \frac{8\pi G}{3} \rho_0 a^{-3(1+w)} \Rightarrow \alpha^2 \frac{1}{t^2} = \frac{8\pi G}{3} \rho_0 \left(\frac{t}{t_0}\right)^{-3(1+w)}$

$$\Rightarrow H_0^2 = \frac{8\pi G}{3} \rho_0 \quad \text{and} \quad -2 = -3(1+w)\alpha \Rightarrow \boxed{\alpha = \frac{2}{3(1+w)}}$$

Friedmann 2: $\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} \rho_0 a^{-3(1+w)} (1+3w) \Rightarrow \frac{\alpha(\alpha-1)}{t^2} = -\frac{4\pi G}{3} \rho_0 \left(\frac{t}{t_0}\right)^{-3(1+w) \frac{2}{3(1+w)}} (1+3w)$

$$\Rightarrow \alpha(\alpha-1) = -\frac{4\pi G}{3} \rho_0 t_0^2 (1+3w) \Rightarrow \alpha(\alpha-1) = -\frac{1}{2} H_0^2 t_0^2 (1+3w)$$

$$\Rightarrow \alpha(\alpha-1) = -\frac{1}{2} \alpha^2 (1+3w) \Rightarrow \alpha-1 = -\frac{1}{2} \alpha (1+3w) \Rightarrow \frac{3}{2} \alpha (1+w) = 1 \quad \checkmark$$

conformal time: $a(\eta) = \left(\frac{\eta}{\eta_0}\right)^\beta \Rightarrow a' = \beta \left(\frac{\eta}{\eta_0}\right)^{\beta-1} \frac{1}{\eta_0} \Rightarrow a'' = (\beta-1)\beta \left(\frac{\eta}{\eta_0}\right)^{\beta-2} \frac{1}{\eta_0^2}$

Friedmann 1: $\left(\frac{a'}{a}\right)^2 = \frac{8\pi G}{3} \rho_0 a^{-1-3w} \Rightarrow \frac{\beta^2}{\eta_0^2} = \frac{8\pi G}{3} \rho_0 \left(\frac{\eta}{\eta_0}\right)^{(-1-3w)\beta}$

$$\Rightarrow \frac{\beta^2}{\eta_0^2} = H_0^2 = \frac{8\pi G}{3} \rho_0 \quad \text{and} \quad -2 = (-1-3w)\beta \Rightarrow \boxed{\beta = \frac{2}{1+3w}}$$

Friedmann 2: $\frac{a''}{a} = +\frac{4\pi G}{3} (1-3w) \rho_0 a^{-3(1+w)+2}$

$$\Rightarrow \beta(\beta-1) \frac{1}{\eta_0^2} = \frac{4\pi G}{3} (1-3w) \rho_0 \left(\frac{\eta}{\eta_0}\right)^{[-3(1+w)+2]\beta}$$

$$\Rightarrow \frac{\beta(\beta-1)}{\eta_0^2} = \frac{4\pi G}{3} (1-3w) \rho_0 \quad \text{and} \quad [-3(1+w)+2]\beta = -2 \quad \checkmark$$

$$\Rightarrow \frac{\beta-1}{\beta} = \frac{1}{2} (1-3w) \Rightarrow 2(\beta-1) = \beta(1-3w) \Rightarrow \beta(1+3w) = 2 \quad \checkmark$$

Problem 4

$$ds^2 = a^2(\eta) \left[-(1+2A) dy^2 + 2\partial_i B dx^i dy + \left[(1+2C - \frac{2}{3}\nabla^2 E) \delta_{ij} + 2\partial_i \partial_j E \right] dx^i dx^j \right]$$

small coordinate refo $\tilde{x}^\mu(x) = x^\mu + \xi^\mu(x)$, where $\xi^\mu = (\xi^0, \partial_i \xi)$ and $x^\mu = (y, x^i)$

$$\begin{aligned} \Rightarrow g_{\mu\nu}(x) &= \frac{\partial \tilde{x}^\alpha}{\partial x^\mu} \frac{\partial \tilde{x}^\beta}{\partial x^\nu} \tilde{g}_{\alpha\beta}(\tilde{x}) \\ &= \frac{\partial \tilde{x}^\alpha}{\partial x^\mu} \frac{\partial \tilde{x}^\beta}{\partial x^\nu} \tilde{g}_{\alpha\beta}(x^\mu + \xi^\mu) \\ &= \frac{\partial \tilde{x}^\alpha}{\partial x^\mu} \frac{\partial \tilde{x}^\beta}{\partial x^\nu} \left(\tilde{g}_{\alpha\beta}(x) + \xi^\nu \frac{\partial}{\partial x^\nu} \tilde{g}_{\alpha\beta} \right) \\ &= \frac{\partial \tilde{x}^\alpha}{\partial x^\mu} \frac{\partial \tilde{x}^\beta}{\partial x^\nu} \left(\tilde{g}_{\alpha\beta} + 2\mathcal{H} \xi^0 \tilde{g}_{\alpha\beta} \right) \end{aligned}$$

$$\begin{aligned} \square: g_{00}(x) &= \left(1 + \frac{d}{dy} \xi^0\right) \left(1 + \frac{d}{dy} \xi^0\right) \tilde{g}_{00} (1 + 2\mathcal{H} \xi^0) + \frac{d}{dy} \xi^i \frac{d}{dy} \xi^i (\tilde{g}_{ij} + 2\mathcal{H} \xi^0 \tilde{g}_{ij}) \\ &= \mathcal{O}(\xi^2) \approx 0 \\ &= \left(1 + 2(\xi^0)'\right) \tilde{g}_{00} (1 + 2\mathcal{H} \xi^0) \end{aligned}$$

$$\Rightarrow 1 + 2A = \left(1 + 2(\xi^0)'\right) (1 + 2\tilde{A}) (1 + 2\mathcal{H} \xi^0) \quad \Rightarrow 1 + 2A = 1 + 2(\xi^0)' + 2\tilde{A} + 2\mathcal{H} \xi^0$$

$$\Rightarrow \tilde{A} = \underline{\underline{A - (\xi^0)' - \mathcal{H} \xi^0}}$$

Newtonian gauge: $\tilde{B} = \tilde{E} = 0 \Rightarrow B + \xi^0 - \xi^1 = 0 \quad \& \quad E - \xi = 0$
 \Rightarrow choose $\underline{\underline{\xi = E}}$ and $\underline{\underline{\xi^0 = E' - B}}$

Enough freedom because $\xi(y, \vec{x})$ and $\xi^0(y, \vec{x})$ are independent.

Newtonian \rightarrow spatially flat gauge ($\tilde{C} = \tilde{E} = 0$)

start with $B = E = 0$ & $A = \mathcal{H}$ & $C = -\Phi$ (Newtonian gauge)

$$\text{now demand } \tilde{C} = 0 \text{ \& } \tilde{E} = 0 \Rightarrow C - \mathcal{H} \xi^0 - \frac{1}{3} \nabla^2 \xi = 0 \text{ \& } \underline{\underline{E - \xi = 0}}$$

$$\Rightarrow -\Phi - \mathcal{H} \xi^0 = 0 \text{ \& } \xi = 0 \Rightarrow \underline{\underline{\xi^0 = -\frac{1}{\mathcal{H}} \Phi}} \text{ \& } \underline{\underline{\xi = 0}}$$

We see that we can go to spatially flat gauge by shifting the time coordinate.

Problem 5

$$(a) \mathcal{R} = -C - \frac{1}{3} K^2 E - \mathcal{H}(v+B)$$

$$\begin{aligned} \tilde{\mathcal{R}} &= -C - \frac{1}{3} K^2 E - \mathcal{H}(v+B) - \mathcal{H} \check{f}^0 + \frac{1}{3} K^2 \check{f} - \frac{1}{3} K^2 \check{f} + \mathcal{H}(\check{f}^1 + \check{f}^0 - \check{f}^1) \\ &= \mathcal{R} \end{aligned}$$

$$\mathcal{R} = \underline{\Phi} - \mathcal{H}v \quad (\text{in Newtonian gauge: } B=E=0)$$

$$(b) \mathcal{R}' = \underline{\Phi}' - \mathcal{H}'v - \mathcal{H}v'$$

$$= \underline{\Phi}' - \mathcal{H}'v + (\mathcal{H}^2 v + \frac{\mathcal{H}\bar{P}'}{\bar{S}+\bar{P}} v) + \frac{\mathcal{H}}{\bar{S}+\bar{P}} \delta\mathcal{P} + \underline{\mathcal{H}\Phi}$$

from (16b) and (16c):

$$4\pi G a^2 \delta\mathcal{S} = -K^2 \Phi + 3\mathcal{H} 4\pi G a^2 (\bar{S} + \bar{P}) v$$

$$\Rightarrow v = \frac{1}{3\mathcal{H}} \frac{\delta\mathcal{S}}{(\bar{S} + \bar{P})} + \frac{K^2 \Phi}{12\mathcal{H}\pi G a^2 (\bar{S} + \bar{P})}$$

$$\Rightarrow \mathcal{R}' = \underline{\Phi}' + \mathcal{H}\Phi + (\mathcal{H}' + \mathcal{H}^2) v + \frac{\mathcal{H}}{\bar{S} + \bar{P}} \delta\mathcal{P} + \frac{1}{3} \frac{\bar{P}'}{(\bar{S} + \bar{P})^2} \delta\mathcal{S} + \frac{\bar{P}' K^2 \Phi}{(\bar{S} + \bar{P})^2 12\pi G a^2}$$

from (11)

$$\begin{aligned} \mathcal{H}^2 - \mathcal{H}' &= \left(\frac{a'}{a}\right)^2 - \frac{a''}{a} + \left(\frac{a'}{a}\right)^2 = 2\mathcal{H}^2 - \frac{a''}{a} = \frac{8\pi G a^2}{3} \left(2\bar{S} - \frac{\bar{S}}{2} + \frac{3}{2}\bar{P}\right) \\ &= 4\pi G a^2 (\bar{S} + \bar{P}) \end{aligned}$$

$$\begin{aligned} \Phi' + \mathcal{H}\Phi + (-\mathcal{H}' + \mathcal{H}^2) v &= -4\pi G a^2 (\bar{S} + \bar{P}) v + (\mathcal{H}^2 - \mathcal{H}') v \\ &= 0 \end{aligned}$$

$$\Rightarrow (\bar{S} + \bar{P}) \frac{\mathcal{R}'}{\mathcal{H}} = \left(\delta\mathcal{P} + \frac{1}{3} \frac{\bar{P}'}{\mathcal{H}(\bar{S} + \bar{P})} \delta\mathcal{S} \right) + \frac{\bar{P}' K^2 \Phi}{(\bar{S} + \bar{P}) 3\mathcal{H} 4\pi G a^2}$$

$$\Rightarrow (\bar{S} + \bar{P}) \frac{\mathcal{R}'}{\mathcal{H}} = \left(\delta\mathcal{P} - \frac{\bar{P}'}{\bar{S}'} \delta\mathcal{S} \right) - \frac{\bar{P}' K^2 \Phi}{\bar{S}' 4\pi G a^2}$$

We have seen that $\Phi \approx \text{const}$ for $k \ll \mathcal{H}$

$$\begin{aligned} \frac{\mathcal{R}'}{\mathcal{H}} &= \frac{1}{(\bar{s} + \bar{p})} (\delta P - \delta p) - \frac{\bar{p}'}{\bar{s}'} \left(\frac{k^2}{\mathcal{H}^2 - 3\mathcal{E}'} \right) \Phi \\ &= \mathcal{O}\left(\frac{k^2}{\mathcal{H}^2}\right) \approx 0 \end{aligned}$$

$$(c) \mathcal{R} = \Phi - \mathcal{H}v \stackrel{(16c)}{=} \Phi + \mathcal{H} \frac{\Phi' + \mathcal{H}\Phi}{4\pi G a^2 (\bar{s} + \bar{p})}$$

$$= \Phi + \frac{\mathcal{H}\Phi' + \mathcal{H}^2\Phi}{(1+w)\mathcal{H}^2 \frac{3}{2}}$$

$$\stackrel{\Phi' \ll 0}{=} \Phi \left(1 + \frac{2}{3} \frac{1}{1+w} \right) = \frac{5+w}{3} \Phi = \frac{5+3w}{3+3w} \Phi$$

transition from radiation to matter era:

$$\frac{5+3\frac{1}{3}}{3+3\frac{1}{3}} \Phi_{\text{rad}} = \frac{5}{3} \Phi_{\text{matter}} \Rightarrow \underline{\underline{\Phi_{\text{matter}} = \frac{9}{10} \Phi_{\text{rad}}}}$$