

NORDITA Winter School 2024 in Particle Physics and Cosmology

Neutrino physics II: Neutrino Mass

Thomas Schwetz-Mangold



Stockholm, 15-17 Jan 2024

Neutrinos oscillate...



... and have mass \Rightarrow physics beyond the Standard Model

- ▶ Part I: Neutrino Oscillations
- ▶ Part II: Neutrino mass - Dirac versus Majorana
- ▶ Part III: Neutrinos and physics beyond the Standard Model

Neutrinos oscillate...



... and have mass \Rightarrow physics beyond the Standard Model

- ▶ Part I: Neutrino Oscillations
- ▶ **Part II: Neutrino mass - Dirac versus Majorana**
- ▶ Part III: Neutrinos and physics beyond the Standard Model

Outline

Absolute neutrino mass

- Neutrino mass from cosmology

- Beta decay – the KATRIN experiment

- Neutrinoless double-beta decay

Fermion masses

- Dirac mass

- Majorana mass

- Dirac versus Majorana neutrinos in the SM

The Standard Model and neutrino mass

Outline

Absolute neutrino mass

Neutrino mass from cosmology

Beta decay – the KATRIN experiment

Neutrinoless double-beta decay

Fermion masses

Dirac mass

Majorana mass

Dirac versus Majorana neutrinos in the SM

The Standard Model and neutrino mass

3-flavour neutrino parameters

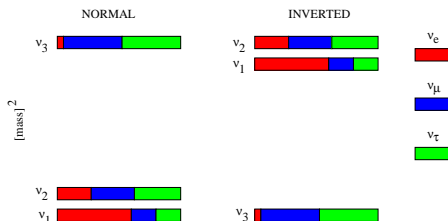
- ▶ 3 masses: Δm_{21}^2 , Δm_{31}^2 , m_0
- ▶ 3 mixing angles: θ_{12} , θ_{13} , θ_{23}
- ▶ 3 phases: 1 Dirac (δ), 2 Majorana (α_1, α_2)

neutrino oscillations

absolute mass observables

lepton-number violation

(neutrinoless double-beta decay)



Absolute neutrino mass

Three ways to measure absolute neutrino mass:

- ▶ Cosmology
(with caveats: cosmological model/data selection)
- ▶ Endpoint of beta spectrum: ${}^3\text{H} \rightarrow {}^3\text{He} + e^- + \bar{\nu}_e$
(experimentally challenging \rightarrow KATRIN)
- ▶ Neutrinoless double beta-decay: $(A, Z) \rightarrow (A, Z + 2) + 2e^-$
(with caveats: lepton number violation)

Absolute neutrino mass

Three ways to measure absolute neutrino mass:
sensitive to different quantities

- ▶ Cosmology

(with caveats: cosmological model/data selection)

$$\sum_i m_i$$

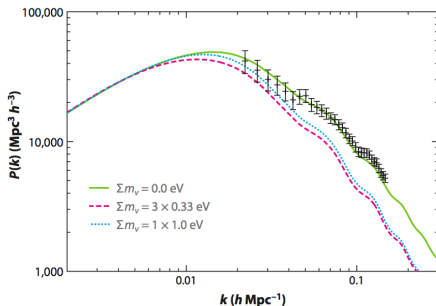
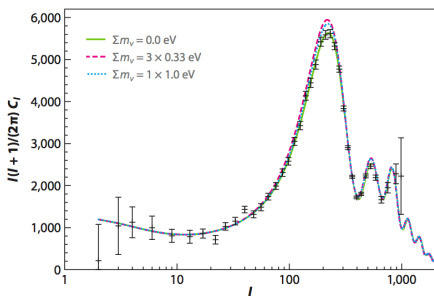
- ▶ Endpoint of beta spectrum: ${}^3\text{H} \rightarrow {}^3\text{He} + e^- + \bar{\nu}_e$
(experimentally challenging \rightarrow KATRIN)

$$m_\beta^2 = \sum_i |U_{ei}^2| m_i^2$$

- ▶ Neutrinoless double beta-decay: $(A, Z) \rightarrow (A, Z + 2) + 2e^-$
(with caveats: lepton number violation)

$$m_{ee} = \left| \sum_i U_{ei}^2 m_i \right|$$

Effect of neutrino mass on CMB and LSS



data points: WMAP 3yr and 2dF '05

Y.Y.Y. Wong, 1111.1436

- ▶ CMB: mainly height of 1st peak
- ▶ LSS: suppression of structure at scales smaller than 1–10 Mpc
- ▶ effects correlated with other parameters of the Λ CDM model

see Lesgourgues, Pastor, astro-ph/06034494 for a review

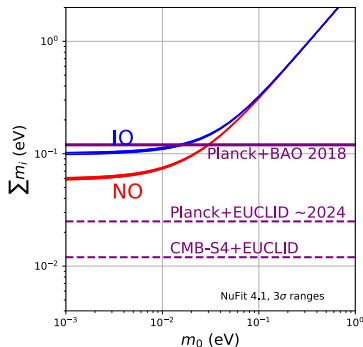
Neutrino mass from cosmology

$$\sum_i m_i \approx \begin{cases} m_0 + \sqrt{m_0 + \Delta m_{21}^2} + \sqrt{m_0 + \Delta m_{31}^2} & (NO) \\ m_0 + 2\sqrt{m_0 + |\Delta m_{31}^2|} & (IO) \end{cases}$$

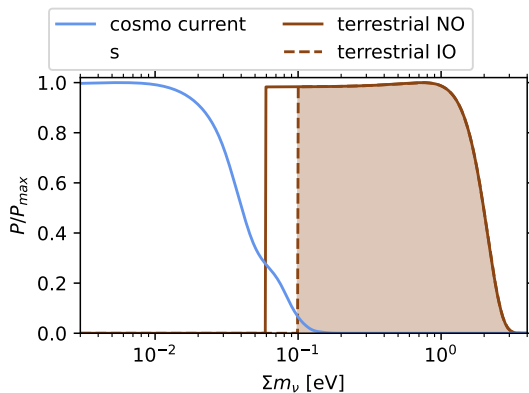
- ▶ minimal value predicted for $m_0 = 0$:

$$\sum m_i \Big|_{\min} \approx \begin{cases} 98.6 \pm 0.85 \text{ meV} & (IO) \\ 58.5 \pm 0.48 \text{ meV} & (NO) \end{cases}$$

- ▶ detection of non-zero neutrino mass expected soon!
- ▶ current limit close to IO minimum excluding IO with cosmology: ongoing discussion [Gariazzo et al., 2205.02195]

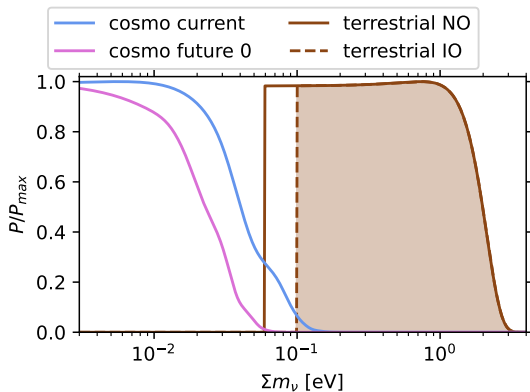


Neutrino mass from cosmology vs terrestrial



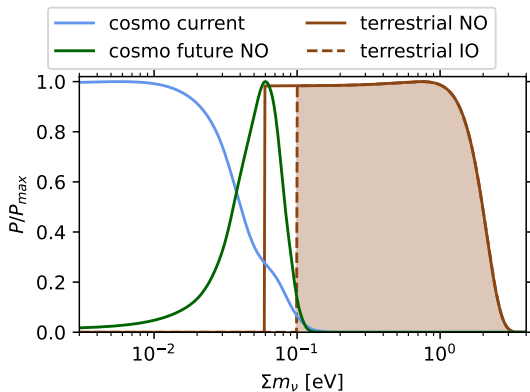
Gariazzo, Mena, TS, 2302.14159

Neutrino mass from cosmology vs terrestrial



Gariazzo, Mena, TS, 2302.14159

Neutrino mass from cosmology vs terrestrial



Gariazzo, Mena, TS, 2302.14159

Beta decay

$$N(A, Z) \rightarrow N(A, Z + 1) + e^- + \bar{\nu}_e$$

$$\frac{d\Gamma}{dE_e} = \frac{G_F^2 m_e^5}{2\pi^2} \cos^2 \theta_c |\mathcal{M}|^2 F(Z, E_e) \underbrace{E_e p_e E_\nu p_\nu}_{\text{phase space}}$$

Tritium decay: ${}^3\text{H} \rightarrow {}^3\text{He} + e^- + \bar{\nu}_e$

$$M_{{}^3\text{H}} = 2.808\,920\,8205 \times 10^6 \text{ keV}$$

$$M_{{}^3\text{He}} = 2.808\,391\,2193 \times 10^6 \text{ keV}$$

$$m_e = 510.9989 \text{ keV}$$

$$Q \equiv M_{{}^3\text{H}} - M_{{}^3\text{He}} - m_e = 18.6023 \text{ keV} \ll M_{{}^3\text{H}}, M_{{}^3\text{He}}$$

$$\kappa \equiv M_{{}^3\text{He}}/M_{{}^3\text{H}} = 1 - 1.89 \times 10^{-4}$$

Beta decay

$$N(A, Z) \rightarrow N(A, Z + 1) + e^- + \bar{\nu}_e$$

$$\frac{d\Gamma}{dE_e} = \frac{G_F^2 m_e^5}{2\pi^2} \cos^2 \theta_c |\mathcal{M}|^2 F(Z, E_e) \underbrace{E_e p_e E_\nu p_\nu}_{\text{phase space}}$$

Tritium decay: ${}^3\text{H} \rightarrow {}^3\text{He} + e^- + \bar{\nu}_e$

$$M_{{}^3\text{H}} = 2.808\,920\,8205 \times 10^6 \text{ keV}$$

$$M_{{}^3\text{He}} = 2.808\,391\,2193 \times 10^6 \text{ keV}$$

$$m_e = 510.9989 \text{ keV}$$

$$Q \equiv M_{{}^3\text{H}} - M_{{}^3\text{He}} - m_e = 18.6023 \text{ keV} \ll M_{{}^3\text{H}}, M_{{}^3\text{He}}$$

$$\kappa \equiv M_{{}^3\text{He}}/M_{{}^3\text{H}} = 1 - 1.89 \times 10^{-4}$$

Tritium beta decay

use E-momentum conservation, calculate electron kin. energy:

$$T \equiv E_e - m_e = \frac{1}{2M_{3\text{H}}} \left[(M_{3\text{H}} - m_e)^2 - M_{3\text{He}}^2 - 2M_{3\text{He}}E_\nu \right]$$

T has a maximum when E_ν has a minimum:

$$\begin{aligned} m_\nu = 0: \quad T_{\max,0} &= \frac{1}{2M_{3\text{H}}} \left[(M_{3\text{H}} - m_e)^2 - M_{3\text{He}}^2 \right] \\ &= Q - \frac{(M_{3\text{H}} - M_{3\text{He}})^2}{2M_{3\text{H}}} \approx Q - 3.4 \text{ eV} \end{aligned}$$

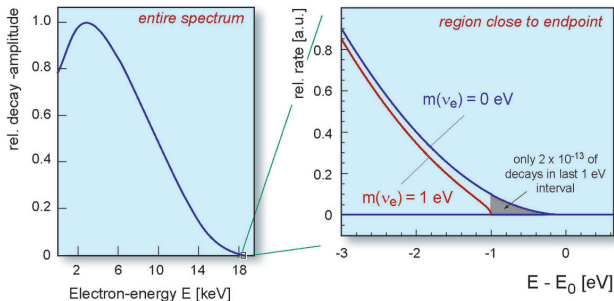
$$m_\nu > 0: \quad T_{\max} = T_{\max,0} - \kappa m_\nu$$

\Rightarrow finite neutrino mass leads to a shift in electron spectrum endpoint

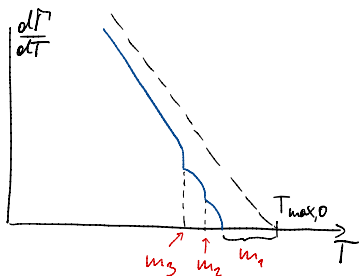
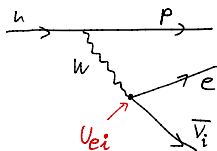
Tritium decay spectrum close to the endpoint

phase space factor: $E_\nu p_\nu = E_\nu \sqrt{E_\nu^2 - m_\nu^2}$, use $E_\nu \approx \frac{M_{3\text{H}}}{M_{3\text{He}}}(T_{\text{max},0} - T)$:

$$\frac{d\Gamma}{dT} \propto (T_{\text{max},0} - T) \sqrt{(T_{\text{max},0} - T)^2 - \kappa^2 m_\nu^2}$$



Take into account neutrino mixing



incoherent sum of individual mass states:

$$\frac{d\Gamma}{dT} = \sum_i |U_{ei}|^2 \frac{d\Gamma_i}{dT}$$

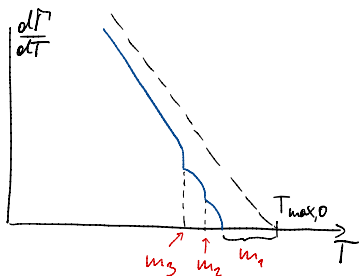
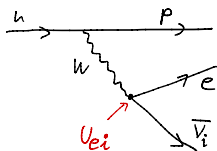
$$\propto (T_{max,0} - T) \sum_i |U_{ei}|^2 \sqrt{(T_{max,0} - T)^2 - \kappa^2 m_i^2}$$

for $T_{max,0} - T \gg \Delta m$:

$$\frac{d\Gamma}{dT} \approx (T_{max,0} - T) \sqrt{(T_{max,0} - T)^2 - \kappa^2 m_\beta^2}$$

$$m_\beta^2 \equiv \sum_i |U_{ei}|^2 m_i^2$$

Take into account neutrino mixing



incoherent sum of individual mass states:

$$\frac{d\Gamma}{dT} = \sum_i |U_{ei}|^2 \frac{d\Gamma_i}{dT}$$

$$\propto (T_{max,0} - T) \sum_i |U_{ei}|^2 \sqrt{(T_{max,0} - T)^2 - \kappa^2 m_i^2}$$

for $T_{max,0} - T \gg \Delta m$:

$$\frac{d\Gamma}{dT} \approx (T_{max,0} - T) \sqrt{(T_{max,0} - T)^2 - \kappa^2 m_\beta^2}$$

$$m_\beta^2 \equiv \sum_i |U_{ei}|^2 m_i^2$$

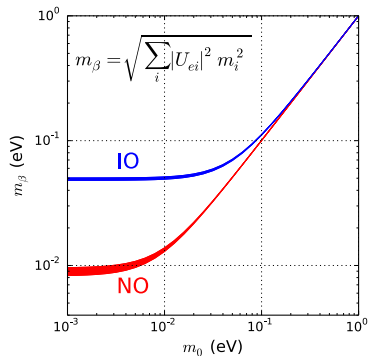
The effective mass

$$m_\beta^2 \equiv \sum_i |U_{ei}|^2 m_i^2 \approx \begin{cases} m_0^2 + |U_{e2}|^2 \Delta m_{21}^2 + |U_{e3}|^2 \Delta m_{31}^2 & \text{(NO)} \\ m_0^2 + (1 - |U_{e3}|^2) |\Delta m_{31}^2| & \text{(IO)} \end{cases}$$

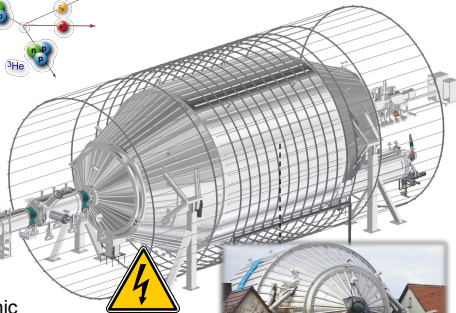
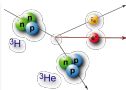
minimum values for $m_0 = 0$:

$$m_\beta^{\min} \approx \begin{cases} 9 \text{ meV} & \text{(NO)} \\ 50 \text{ meV} & \text{(IO)} \end{cases}$$

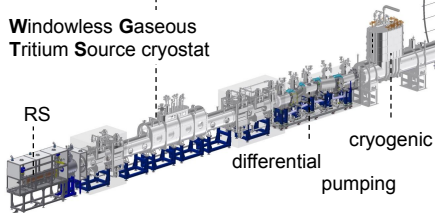
for $m_0 \gg |\Delta m_{31}^2|$: $m_\beta \approx m_0$



KATRIN overview: 70 m long beamline



Windowless Gaseous
Tritium Source cryostat

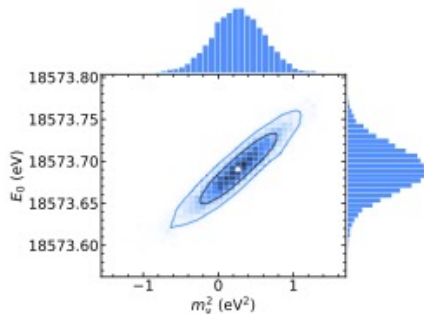
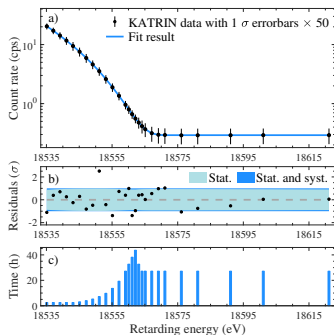


Main Spectrometer



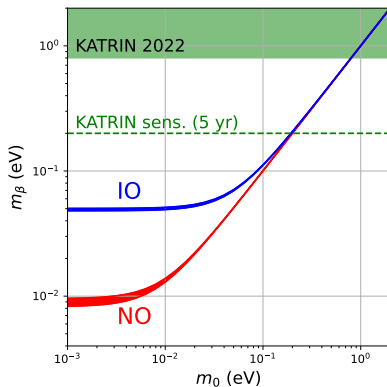
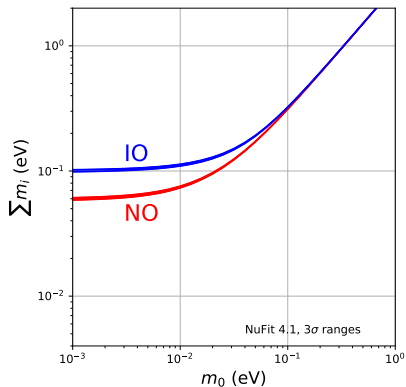
KATRIN results

Aker et al., 1909.06048 (PRL19), 2105.08533 (Nature Phys. 22)

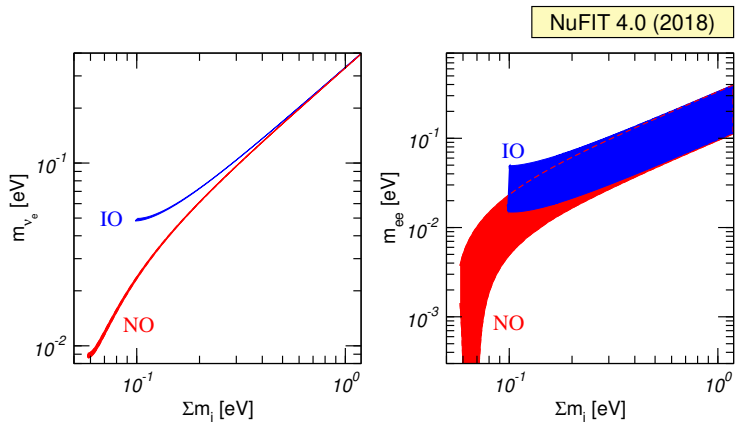


$$m_\beta^2 = 0.26 \pm 0.34 \text{ eV}^2$$

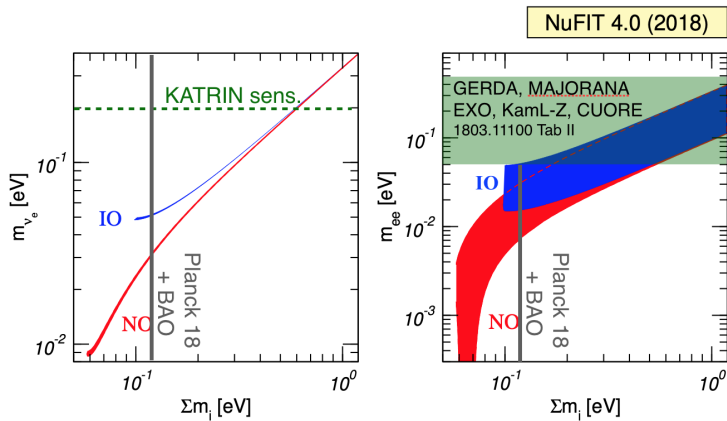
$$m_\beta < 0.8 \text{ eV (90\% CL)}$$

Cosmology and β decay observables

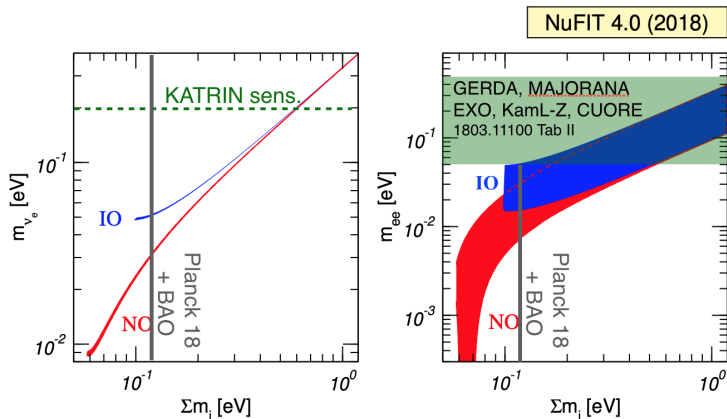
Absolute neutrino mass



Absolute neutrino mass



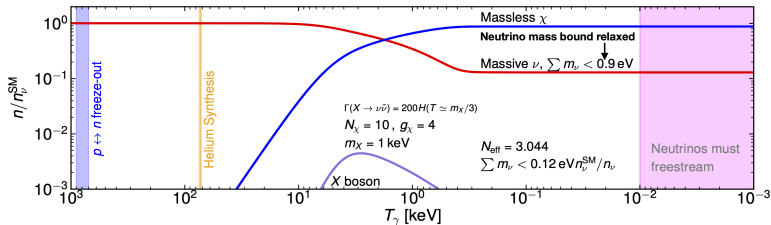
Absolute neutrino mass



relies on standard three-flavour scenario and standard cosmology
 Any inconsistency would indicate new physics beyond 3 flavour neutrino mass!
 discussion of non-standard neutrino cosmology: [Alvey et al., 2111.14870](#)

One example how to relax the cosmological bound

Farzan, Hannestad, 1510.02201; Escudero, TS, Terol-Calvo, 2211.01729

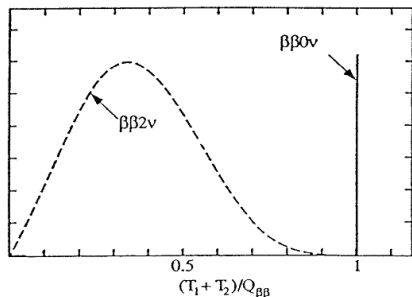


- ▶ introduce $N_\chi \gtrsim 10$ generations of massless sterile neutrinos with $\theta_{\nu_\chi} \sim 10^{-3}$
- ▶ a vector mediator X , $m_X \sim 10$ keV, $U(1)_{\text{dark}}$ breaking around ~ 1 GeV
- ▶ convert active neutrinos into sterile neutrinos between BBN and recombination
- ▶ mass bound gets relaxed: $\sum m_\nu < 0.12 \text{ eV} (1 + 2N_\chi/3)$

Neutrinoless double-beta decay

2-neutrino double-beta decay: $(A, Z) \rightarrow (A, Z + 2) + 2e^- + 2\bar{\nu}_e$

neutrinoless double-beta decay: $(A, Z) \rightarrow (A, Z + 2) + 2e^-$



Example ^{76}Ge (GERDA experiment):

$$2\beta 2\nu : T_{1/2} = (1.8 \pm 0.1) \times 10^{21} \text{ yr}$$

$$2\beta 0\nu : T_{1/2} > 2.1 \times 10^{25} \text{ yr}$$

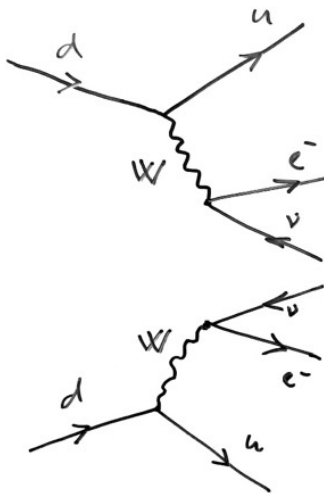
(importance of energy resolution and background suppression)

Neutrinoless double-beta decay



- ▶ an observation of this process would prove that **lepton number is violated**
- ▶ proves Majorana nature of neutrinos
- ▶ BUT no direct prove of neutrino mass (a different mechanism could be responsible)

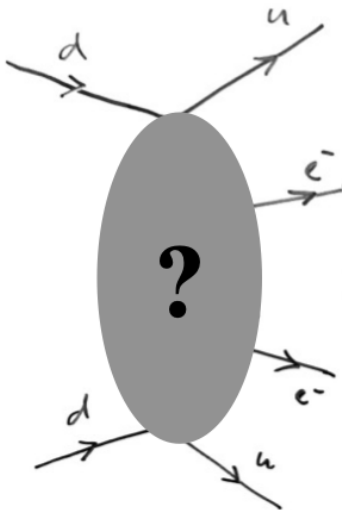
The neutrino-mass mechanism



The neutrino-mass mechanism



The neutrino-mass mechanism



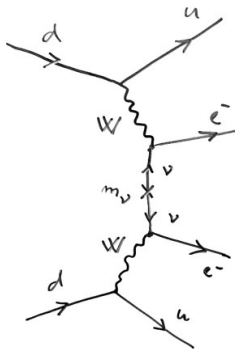
BUT: what we observe is just $\Delta L = 2$

The neutrino-mass mechanism

assuming that light neutrino exchange is responsible for the decay:

$$m_{\beta\beta} = |\mathcal{M}_{ee}| \quad (\text{in basis where ch. lepton mass matrix is diag.})$$

$$= \left| \sum_i U_{ei}^2 m_i \right| = |c_{13}^2 c_{12}^2 m_1 + c_{13}^2 s_{12}^2 e^{i\alpha_1} m_2 + s_{13}^2 e^{i\alpha_2} m_3|$$



coherent sum of individual neutrino masses

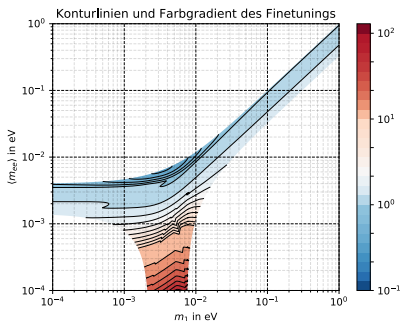
The neutrino-mass mechanism

assuming that light neutrino exchange is responsible for the decay:

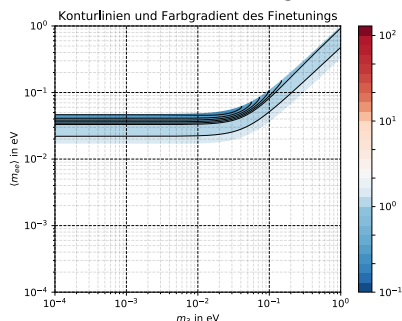
$$m_{\beta\beta} = |\mathcal{M}_{ee}| \quad (\text{in basis where ch. lepton mass matrix is diag.})$$

$$= \left| \sum_i U_{ei}^2 m_i \right| = |c_{13}^2 c_{12}^2 m_1 + c_{13}^2 s_{12}^2 e^{i\alpha_1} m_2 + s_{13}^2 e^{i\alpha_2} m_3|$$

normal ordering

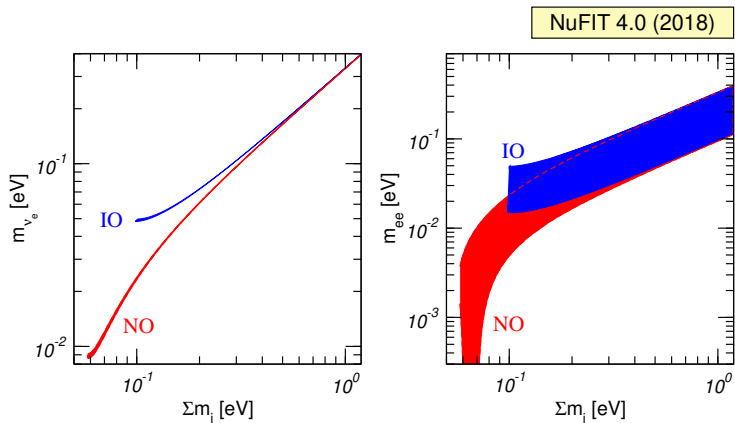


inverted ordering

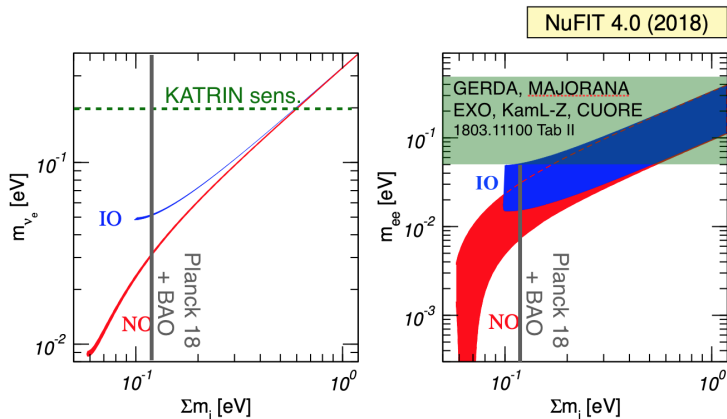


M. Eichhorn, BSc thesis, KIT 2018

Absolute neutrino mass



Absolute neutrino mass



uncertainties on $m_{\beta\beta}$ due to Majorana phases and relating $T_{1/2}$ to $m_{\beta\beta}$ due to nuclear matrix elements
 recent review: [Agostini et al., 2202.01787](#)

Outline

Absolute neutrino mass

Neutrino mass from cosmology

Beta decay – the KATRIN experiment

Neutrinoless double-beta decay

Fermion masses

Dirac mass

Majorana mass

Dirac versus Majorana neutrinos in the SM

The Standard Model and neutrino mass

Dirac fermion

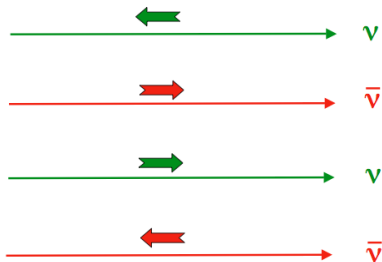
$$\mathcal{L}_D = i\bar{\psi}\gamma^\mu\partial_\mu\psi - m\bar{\psi}\psi$$

Dirac equation:

$$(i\gamma^\mu\partial_\mu - m)\psi = 0$$

ψ is a 4-component object: 2 helicity states for particle and anti-particle

4 mass-degenerate states:



Representations of SM are chiral fields

left- and right-chirality projection operators:

$$P_L = \frac{1}{2}(1 - \gamma_5), \quad P_R = \frac{1}{2}(1 + \gamma_5)$$

left and right chiral fields (irreducible representations of Lorentz group):

$$P_L \psi_L = \psi_L, \quad P_R \psi_R = \psi_R, \quad \psi = \psi_L + \psi_R$$

Dirac Lagrangian:

$$\begin{aligned} \mathcal{L}_D &= i\bar{\psi}\gamma^\mu\partial_\mu\psi - m\bar{\psi}\psi \\ &= i\bar{\psi}_L\gamma^\mu\partial_\mu\psi_L + i\bar{\psi}_R\gamma^\mu\partial_\mu\psi_R - m\bar{\psi}_L\psi_R - m\bar{\psi}_R\psi_L \end{aligned}$$

Dirac equation (mass term mixes chiralities):

$$\begin{aligned} i\gamma^\mu\partial_\mu\psi_L - m\psi_R &= 0 \\ i\gamma^\mu\partial_\mu\psi_R - m\psi_L &= 0 \end{aligned}$$

Representations of SM are chiral fields

left- and right-chirality projection operators:

$$P_L = \frac{1}{2}(1 - \gamma_5), \quad P_R = \frac{1}{2}(1 + \gamma_5)$$

left and right chiral fields (irreducible representations of Lorentz group):

$$P_L \psi_L = \psi_L, \quad P_R \psi_R = \psi_R, \quad \psi = \psi_L + \psi_R$$

Dirac Lagrangian:

$$\begin{aligned} \mathcal{L}_D &= i\bar{\psi}\gamma^\mu\partial_\mu\psi - m\bar{\psi}\psi \\ &= i\bar{\psi}_L\gamma^\mu\partial_\mu\psi_L + i\bar{\psi}_R\gamma^\mu\partial_\mu\psi_R - m\bar{\psi}_L\psi_R - m\bar{\psi}_R\psi_L \end{aligned}$$

Dirac equation (mass term mixes chiralities):

$$\begin{aligned} i\gamma^\mu\partial_\mu\psi_L - m\psi_R &= 0 \\ i\gamma^\mu\partial_\mu\psi_R - m\psi_L &= 0 \end{aligned}$$

Dirac Lagrangian:

$$\begin{aligned}\mathcal{L}_D &= i\bar{\psi}\gamma^\mu\partial_\mu\psi - m\bar{\psi}\psi \\ &= i\bar{\psi}_L\gamma^\mu\partial_\mu\psi_L + i\bar{\psi}_R\gamma^\mu\partial_\mu\psi_R - m\bar{\psi}_L\psi_R - m\bar{\psi}_R\psi_L\end{aligned}$$

invariant under a $U(1)$ symmetry

$$\psi_L \rightarrow e^{i\alpha}\psi_L, \quad \psi_R \rightarrow e^{i\alpha}\psi_R$$

conserved quantum number (charge, lepton number, ...)

particle is different from anti-particle

\Rightarrow any charged Fermion has to be a Dirac particle

Majorana field: replace ψ_R by ψ_L^c :

$$\psi = \psi_L + \psi_L^c$$

with particle- antiparticle conjugation \hat{C} :

$$\hat{C}: \quad \psi \rightarrow \psi^c \equiv C\bar{\psi}^T \equiv C\gamma_0^T\psi^*$$

$$C^{-1}\gamma^\mu C = -\gamma^{\mu T}, \quad C^\dagger = C^{-1} = -C^*$$

\hat{C} changes chirality:

$$\psi_L \rightarrow (\psi_L)^c \equiv \psi_L^c \quad \text{with} \quad P_R\psi_L^c = \psi_L^c, \quad P_L\psi_L^c = 0$$

Majorana fermion

the Majorana field $\psi = \psi_L + \psi_L^c$ fulfills the Majorana condition

$$\psi = \psi^c$$

“is its own anti-particle”

only 2 independent
(mass-degenerate) states:



Majorana fermion

$$\mathcal{L}_M = i\bar{\psi}_L \gamma^\mu \partial_\mu \psi_L + \frac{m}{2} \left[\psi_L^T C^{-1} \psi_L + \text{h.c.} \right]$$

- ▶ explicitly built out of only ψ_L (2 dof)
- ▶ this Lagrangian is not invariant under $\psi_L \rightarrow e^{i\alpha} \psi_L$
- ▶ Majorana mass term breaks all $U(1)$ charges by 2 units
- ▶ cannot define “particle” and “anti-particle”
- ▶ any (electrically) charged particle cannot be a Majorana particle

In weak interactions we speak about
“neutrinos” and “antineutrinos”

How can the neutrino be a Majorana particle,
being its own antiparticle?

In the SM neutrinos only left-chiral fields participate in weak interactions:

the left-handed field ν_L acts as “neutrino”

the right-handed field $\bar{\nu}_L$ acts as “antineutrino”

- ▶ we need a “L” and a “R” neutrino state for weak interactions (to describe “neutrino” and “antineutrino”)
- ▶ we need a “L” and a “R” neutrino state to form a mass term

Majorana:

- ▶ those states are identical (there are only two independent states, ν_L, ν_L^c)

Dirac:

- ▶ the R state to form the mass term is different than the one acting as “antineutrino” in weak interactions (4 independent states) → “right-handed neutrino”: does not participate in weak interactions

In the SM neutrinos only left-chiral fields participate in weak interactions:

the left-handed field ν_L acts as “neutrino”

the right-handed field $\bar{\nu}_L$ acts as “antineutrino”

- ▶ we need a “L” and a “R” neutrino state for weak interactions (to describe “neutrino” and “antineutrino”)
- ▶ we need a “L” and a “R” neutrino state to form a mass term

Majorana:

- ▶ those states are identical (there are only two independent states, ν_L, ν_L^c)

Dirac:

- ▶ the R state to form the mass term is different than the one acting as “antineutrino” in weak interactions (4 independent states) → “right-handed neutrino”: does not participate in weak interactions

In the SM neutrinos only left-chiral fields participate in weak interactions:

the left-handed field ν_L acts as “neutrino”

the right-handed field $\bar{\nu}_L$ acts as “antineutrino”

- ▶ we need a “L” and a “R” neutrino state for weak interactions (to describe “neutrino” and “antineutrino”)
- ▶ we need a “L” and a “R” neutrino state to form a mass term

Majorana:

- ▶ those states are identical (there are only two independent states, ν_L, ν_L^c)

Dirac:

- ▶ the R state to form the mass term is different than the one acting as “antineutrino” in weak interactions (4 independent states) \rightarrow “right-handed neutrino”: does not participate in weak interactions

Chirality versus helicity

physical states are helicity eigenstates:

$$\frac{\vec{\sigma}\vec{p}}{|\vec{p}|}\psi_{\pm} = \pm\psi_{\pm}$$

for massless fermions helicity and chirality coincides:

$$\psi_{-} = \psi_L, \quad \psi_{+} = \psi_R \quad (\text{massless})$$

for relativistic massive fermions ($m \ll E$) we have:

$$\psi_{-} \approx \psi_L + \frac{m}{2E}\psi_R, \quad \psi_{+} \approx \psi_R + \frac{m}{2E}\psi_L$$

OBS: here “ ψ_R ” denotes the right-chiral field in the mass term, which corresponds to ψ^c in the Majorana case

Chirality versus helicity

physical states are helicity eigenstates:

$$\frac{\vec{\sigma}\vec{p}}{|\vec{p}|}\psi_{\pm} = \pm\psi_{\pm}$$

for massless fermions helicity and chirality coincides:

$$\psi_{-} = \psi_L, \quad \psi_{+} = \psi_R \quad (\text{massless})$$

for relativistic massive fermions ($m \ll E$) we have:

$$\psi_{-} \approx \psi_L + \frac{m}{2E}\psi_R, \quad \psi_{+} \approx \psi_R + \frac{m}{2E}\psi_L$$

OBS: here “ ψ_R ” denotes the right-chiral field in the mass term, which corresponds to ψ^c in the Majorana case

Chirality versus helicity

physical states are helicity eigenstates:

$$\frac{\vec{\sigma}\vec{p}}{|\vec{p}|}\psi_{\pm} = \pm\psi_{\pm}$$

for massless fermions helicity and chirality coincides:

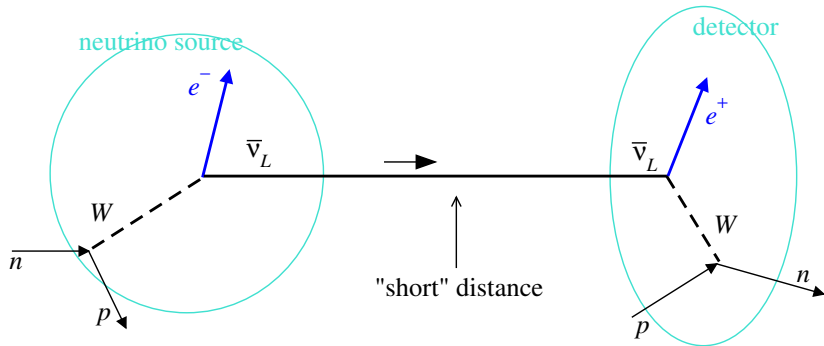
$$\psi_{-} = \psi_L, \quad \psi_{+} = \psi_R \quad (\text{massless})$$

for relativistic massive fermions ($m \ll E$) we have:

$$\psi_{-} \approx \psi_L + \frac{m}{2E}\psi_R, \quad \psi_{+} \approx \psi_R + \frac{m}{2E}\psi_L$$

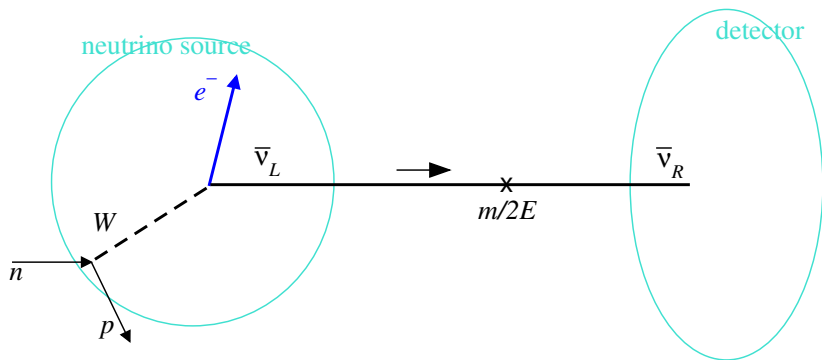
OBS: here “ ψ_R ” denotes the right-chiral field in the mass term, which corresponds to ψ^c in the Majorana case

A typical neutrino experiment (massless neutrinos)



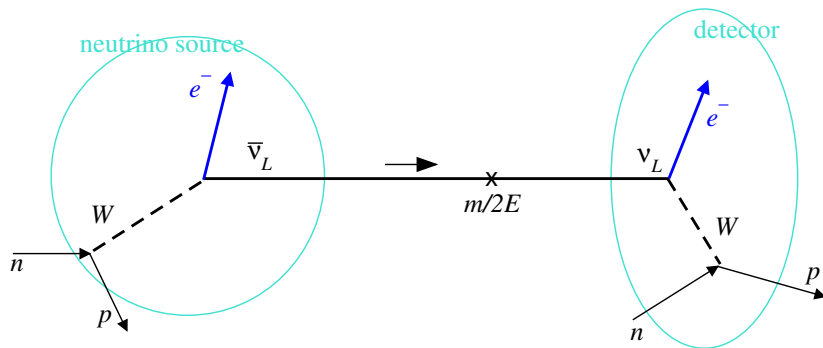
Mass induced chirality flip - Dirac

with a probability suppressed wrt leading diagram by $(m/2E)^2 \lesssim 10^{-12}$



Mass induced chirality flip - Majorana

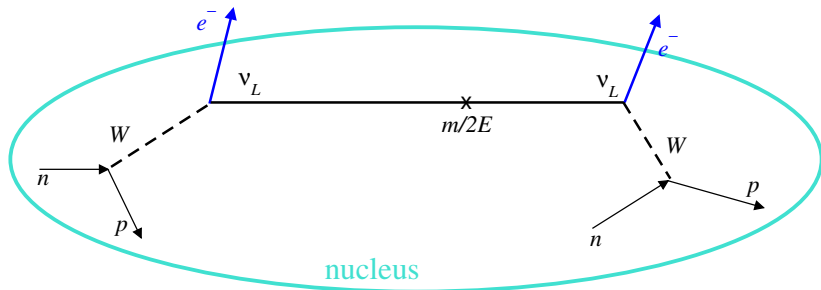
with a probability suppressed wrt leading diagram by $(m/2E)^2 \lesssim 10^{-12}$



Schechter, Valle, PRD 1981

Mass induced chirality flip - Majorana

Neutrinoless double-beta decay $(A, Z) \rightarrow (A, Z + 2) + 2e^-$



Outline

Absolute neutrino mass

Neutrino mass from cosmology

Beta decay – the KATRIN experiment

Neutrinoless double-beta decay

Fermion masses

Dirac mass

Majorana mass

Dirac versus Majorana neutrinos in the SM

The Standard Model and neutrino mass

Masses in the Standard Model

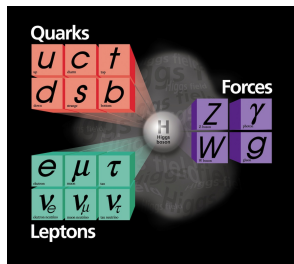
- ▶ The Standard Model has only one dimension full parameter: the vacuum expectation value of the Higgs:

$$\langle \phi \rangle \approx 174 \text{ GeV}$$

- ▶ All masses in the Standard Model are set by this single scale:

$$m_i = y_i \langle \phi \rangle$$

top quark: $y_t \approx 1$
 electron: $y_e \approx 10^{-6}$



Fermion masses in the Standard Model

fermions of one generation:

$$\text{quarks: } Q_L = \begin{pmatrix} u_L \\ d_L \end{pmatrix}, u_R, d_R \quad \text{leptons: } L_L = \begin{pmatrix} \nu_L \\ e_L \end{pmatrix}, e_R$$

mass terms from Yukawa coupling to Higgs ϕ

$$\mathcal{L}_Y = -\lambda_d \bar{Q}_L \phi d_R - \lambda_u \bar{Q}_L \tilde{\phi} u_R + \text{h.c.} \quad -\lambda_e \bar{L}_L \phi e_R + \text{h.c.}$$

$$\text{EWSB} \rightarrow -m_d \bar{d}_L d_R - m_u \bar{u}_L u_R + \text{h.c.} \quad -m_e \bar{e}_L e_R + \text{h.c.}$$

$$\tilde{\phi} \equiv i\sigma_2 \phi^*, \quad m_d = \lambda_d \frac{v}{\sqrt{2}}, \quad m_u = \lambda_u \frac{v}{\sqrt{2}}, \quad m_e = \lambda_e \frac{v}{\sqrt{2}}, \quad \langle \phi \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v \end{pmatrix}$$

Dirac mass terms for charged fermions

Fermion masses in the Standard Model

fermions of one generation:

$$\text{quarks: } Q_L = \begin{pmatrix} u_L \\ d_L \end{pmatrix}, u_R, d_R \quad \text{leptons: } L_L = \begin{pmatrix} \nu_L \\ e_L \end{pmatrix}, e_R$$

mass terms from Yukawa coupling to Higgs ϕ

$$\mathcal{L}_Y = -\lambda_d \bar{Q}_L \phi d_R - \lambda_u \bar{Q}_L \tilde{\phi} u_R + \text{h.c.} \quad -\lambda_e \bar{L}_L \phi e_R + \text{h.c.}$$

$$\text{EWSB} \rightarrow -m_d \bar{d}_L d_R - m_u \bar{u}_L u_R + \text{h.c.} \quad -m_e \bar{e}_L e_R + \text{h.c.}$$

$$\tilde{\phi} \equiv i\sigma_2 \phi^*, \quad m_d = \lambda_d \frac{v}{\sqrt{2}}, \quad m_u = \lambda_u \frac{v}{\sqrt{2}}, \quad m_e = \lambda_e \frac{v}{\sqrt{2}}, \quad \langle \phi \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v \end{pmatrix}$$

Dirac mass terms for charged fermions

- ▶ “right-handed” neutrinos would be complete gauge singlets in the SM
- ▶ no gauge interactions
- ▶ left out in the original formulation of the SM
⇒ no Dirac mass term for neutrinos

- ▶ Why is there no Majorana mass term?

- ▶ Lepton-number is an accidental symmetry in the SM → given the gauge symmetry and the field content of the SM we cannot construct a Majorana mass term for neutrinos (true at any loop order)

- ▶ “right-handed” neutrinos would be complete gauge singlets in the SM
- ▶ no gauge interactions
- ▶ left out in the original formulation of the SM
⇒ no Dirac mass term for neutrinos

- ▶ Why is there no Majorana mass term?

- ▶ Lepton-number is an accidental symmetry in the SM → given the gauge symmetry and the field content of the SM we cannot construct a Majorana mass term for neutrinos (true at any loop order)

In the SM neutrinos are massless because. . .

1. there are no right-handed neutrinos to form a Dirac mass term
2. because of the field content (scalar sector) and gauge symmetry lepton number¹ is an accidental global symmetry of the SM and therefore no Majorana mass term can be induced.
3. restriction to renormalizable terms in the Lagrangian

¹B-L at the quantum level

In the SM neutrinos are massless because. . .

1. there are no right-handed neutrinos to form a Dirac mass term
2. because of the field content (scalar sector) and gauge symmetry lepton number¹ is an accidental global symmetry of the SM and therefore no Majorana mass term can be induced.
3. restriction to renormalizable terms in the Lagrangian

Neutrino mass implies physics beyond the Standard Model

At least one of the above items needs to be violated

¹B-L at the quantum level