The Role of Galaxies and AGN during Reionization Insights from JWST ASPIRE QSO fields and IGM Tomography —

Koki Kakiichi Cosmic Dawn Center, University of Copenhagen

DAWN COSMIC DAWN CENTER

Nordita, Cosmic Dawn at High Latitudes, 24 June 2024







JWST pushes the frontier of distant galaxies & reionization ...

Credit: NIRCam GOODS-South field

NASA, ESA, CSA, B. Robertson, B. Johnson, S. Tacchella, M. Rieke, D. Eisenstein, A. Pagan



JWST pushes the frontier of distant galaxies & reionization ... When do we think that we have 'solved' cosmic reionization?

Credit: NIRCam GOODS-South field NASA, ESA, CSA, B. Robertson, B. Johnson, S. Tacchella, M. Rieke, D. Eisenstein, A. Pagan



Lya forest is the cleanest probe of the Intergalactic Medium (IGM)



$$\frac{1}{2} \frac{1}{s^{-1}} \int_{0}^{-1} \left(\frac{T}{10^4 \,\mathrm{K}} \right)^{-0.72} \left(\frac{1+z}{7} \right)^{9/2}$$

Physics of the IGM

Meiksin+09, Becker+15, McQuinn+16 for reviews





JWST spectroscopic survey of galaxies in QSO fields



& scope of JWST EIGER (Kashino+23) and ASPIRE (Wang+23) surveys with NIRCam/WFSS

JWST ASPIRE Cycle 1 (PI: Wang) NIRCam/F356W Grism Spectroscopic Galaxy Survey of 25 QSO Fields at z>6



Wang et al incl. KK 2023 z>6 QSO

"JWST as Redshift Machine"



JWST ASPIRE Cycle 1 (PI: Wang) NIRCam/F356W Grism Spectroscopic Galaxy Survey of 25 QSO Fields at z>6

'Shallow & wide' survey to reveals ~300 5.3<z<7 galaxies with [OIII] lines in the Lya forest regions of 25 background QSOs

cf. EIGER survey targets 6 QSO fields with deep exposure

See also Daichi Kashino's Talk

Wang et al incl. KK 2023



"JWST as Redshift Machine"





JWST map of galaxy-IGM spatial correlations at z~6



[OIII] emitters reside in the highly ionized IGM at z~6

See also Xiangyu Jin's Talk

Kakiichi & Jin + in prep





JWST map of galaxy-IGM spatial correlations at z~6



[OIII] emitters reside in the highly ionized IGM at z~6

See also Xiangyu Jin's Talk

Kakiichi & Jin + in prep





JWST map of galaxy-IGM

Kakiichi & Jin + in prep







Kakiichi+ in prep





Observed [OIII] emitters (M_{UV}≈-20) are sub-dominant contributor to the ionising background in this ionised region of the IGM

$$\frac{\Gamma_{\rm HI}^{\rm OIII}}{\Gamma_{\rm HI}} < 0.003 \left(\frac{f_{\rm esc}^{\rm OIII}}{0.10}\right) \left(\frac{\xi_{\rm ion}^{\rm OIII}}{10^{25.5} \, {\rm erg}^{-1} {\rm Hz}}\right)$$

Faint AGN?



Fainter or unseen galaxies are likely the main source of reionization during its final stage



Galaxy-Lya forest cross-correlation Evidence for galaxies residing in the highly ionized IGM environment at z~5.8



Kakiichi+ in prep

Galaxy-Lya forest cross-correlation Comparison with halo model + fluctuating UV background



Based on Kakiichi+18 & Meyer+20 conditional luminosity function + RT model framework Source model matched to Bouwens+21 luminosity function & Harikane+22 galaxy clustering

Kakiichi+ in prep

Galaxy-Lya forest cross-correlation THESAN full cosmological radiation hydrodynamic simulations



Universal prediction of all reionization theories = Galaxies in the ionized IGM / bubbles.

Kakiichi+ in prep

Underlying spatial distribution of galaxies & reionization morphology



Rapid redshift evolution of galaxy-Lyα forest cross-correlation at the final stages of reionziation



Kakiichi+ in prep



Do we understand the error? Yes, sort of.



Theory

 $\overline{T}_{\text{IGM}}^2$ $\operatorname{Cov}[\langle \hat{T}_{\mathrm{IGM}}(r_i) \rangle, \langle \hat{T}_{\mathrm{IGM}}(r_j) \rangle] = 2\pi$ $\operatorname{Var}[\hat{P}_{g\alpha}(k)] = P_{g\alpha}(k)^2 + \left(P_g(k) + n_{g,3D}^{-1}\right)$

Cosmic variance

~1/(Survey Volume)~1/(# of quasar fields)

Noise in Lya forest. ~1/(signal-to-noise of quasar spectra)

$$\frac{k}{2}\bar{j}_0(k \mid r_i)\bar{j}_0(k \mid r_j)\operatorname{Var}[\hat{P}_{g\alpha}(k)]$$

$$\left(P_{\alpha}(k) + P_{\alpha}^{1\mathrm{D}}(k_{\parallel})n_{\mathrm{eff},\mathrm{2D}}^{-1}\right)$$

McQuinn & White 2011

Linear power spectra with UV background fluctuations (Pontzen 2014, Gontcho-a-Gontcho+2014)





Towards Mapping the Cosmic Frontier in "3D"





For $z \approx 3 - 6$, we need 30-m class telescopes...



gas in foreground

CLAMATO: Lee+14,18 LATIS: Newman+20

Japeli+19







Mapping the Cosmic Web: "Photometric" IGM Tomography



Photometric measurement of the IGM transmission along background galaxies

Kakiichi+22,23

Mapping the Cosmic Web: "Photometric" IGM Tomography



Photometric measurement of the IGM

is similar to weak lensing tomography





Pilot Subaru/HSC study of Photometric IGM Tomography at z~5







Pilot Subaru/HSC study of Photometric IGM Tomography at z~5







JWST: Mapping the Universe in 3D JWST Wide-Field Grism Spectroscopic Survey of the COSMOS-Web Field

"COSMOS-3D" Large 265 hour public treasury program z~0-9 spectroscopic surve NIRCam/F444W WFSS **Complete full NIRCam imaging** F115W (2nd epoch), F200W & F356W Parallel MIRI imaging F1000W & F2100W

This is built on legacy of FRESCO (Oesch+23) & 3D-HST (Brammer+12) & many years of the community effort on the COSMOS field, e.g. COSMOS-Web (Casey & Kartaltepe+23), COSMOS2020 (Weaver+21) ... Scoville+07





PI: Kakiichi, Co-PIs: Wang, Fan, Liu, Yang, Egami



JWST: Mapping the Universe in 3D 3D Pano-Chromatic View of the Universe from sub-kpc to tens of Mpc over z~0-9

NIRCam/F444W grism mode + Panchromatic multi-wavelength data

z~0-9

Large-scale clustering and formation of early galaxies

Crisis of ΛCDM ? or Efficient build-up of stellar mass? i.e. Galaxy formation vs Cosmological interpretations



Large-scale clustering and demographics of SMBHs

Mysteries surrounding Little Red Dots & Formation of early SMBHs Map all foreground & background galaxies across cosmic time from z~0 to 9

Galaxies in the Cosmic Web JWST galaxies × IGM tomography **Epoch of Reionization**







JWST-Mapping the Universe in 3D 3D Pano-Chrom dic View of the Universe from sub-kpc to tens of Mpc over z~0-9

NIRCam/F444W g

z~0-9

ism mode + Panchromatic n ulti-wavelength data

Large-scale clustering and formation of early galaxies $\Lambda CDM? or$ Crisis c Efficient build-up of stellar mass? i.e. Galaxy formation vs Cosmological interpretations

Subaru HSC & PFS & Euclid

Large-scale clustering and demographics of SMBHs

Mysteries surrounding Little Red Dots & Formation of early SMBHs

Map all foreground & bac ground galaxies across cosmic time from z to 9

Galaxies in th e Cosmic Web **IGM tomography** JWST galaxies > Epoch of R ionization



COSMOS-3D COSMOS (Subaru/HSC NB816)

3x10⁷ Mpc³ 1200 arcmin²

COSMOS-Web (F115W/F150W/F27 W/F444W)





F Mapping the Universe in 30 3D Pano-Synergy/iew of the Universe $z \simeq 5.7$ IGM tomography

COSMOS (Subaru

COS

Euclid Subaru

Z ~/

Ef

i.e. Galax

Dream of galaxy × 21cm tomography may come true? Morphological analysis of ionised bubbles and the spatial correlation with galaxies Kakiichi+17, Busch et al 20, Giri & Mellema 21, Elbers & van de Weygaert 19,23





Summary

Spatial correlation between galaxy and Lyα forest

Evidence for [OIII] emitters residing in highly ionized IGM at the final stage of 1. reionization from statistical [OIII] emitter-Lyα forest cross-correlation

2. Evidence for faint or unseen galaxies completing reionization from individual [OIII] emitter - transmission spike associations

• Photometric IGM tomography is one of the key driver of a Cycle-3 public treasury program "COSMOS-3D"

JWST Wide-Field Grism Spectroscopic Survey of the COSMOS-Web Field with NIRCam/F444W + parallel MIRI imaging (F1000W & F2100W) + time-domain F115W

from JWST ASPIRE QSO field survey



Back-up





