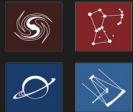


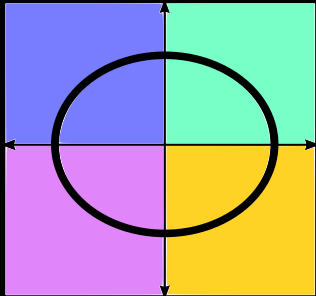
SGWB GENERATED BY SUPERCONDUCTING COSMIC STRINGS

Based on: I. Yu. Rybak & L. Sousa, JCAP 11 (2022) 024 (arXiv: 2209.01068)

I. Yu. Rybak & L. Sousa, Phys.Rev.D 111 (2025) 8, 083502 (arXiv: 2412.17154)

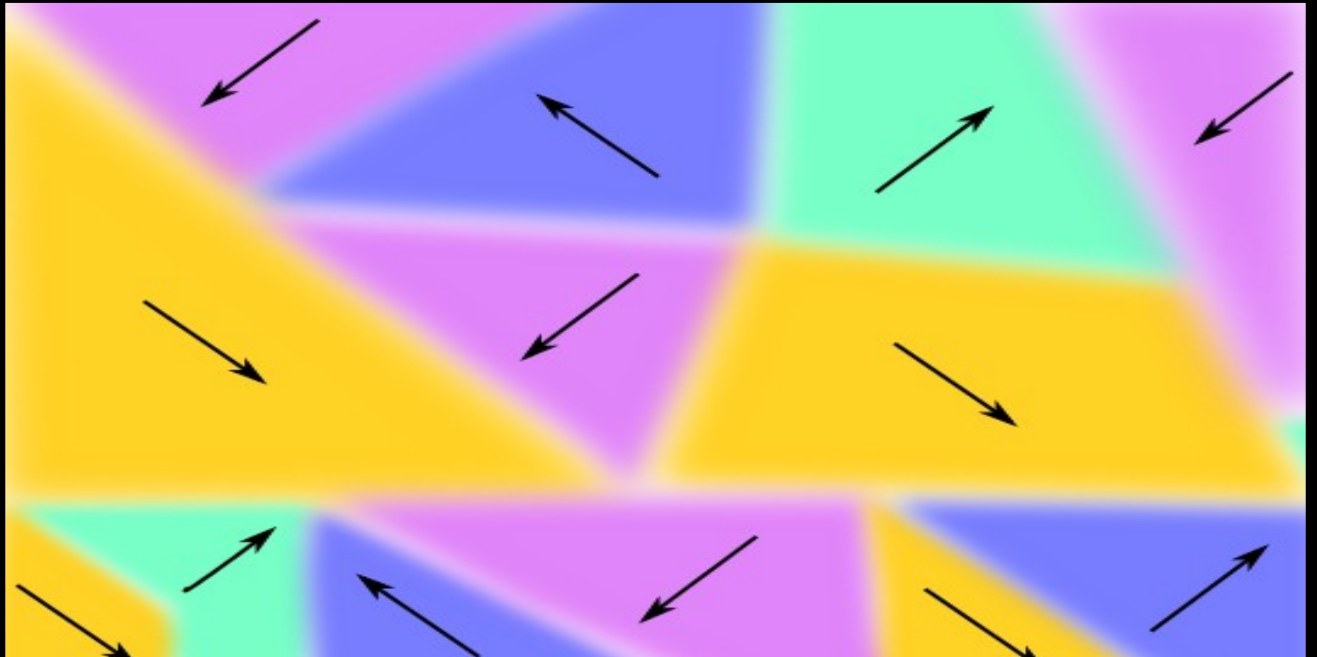
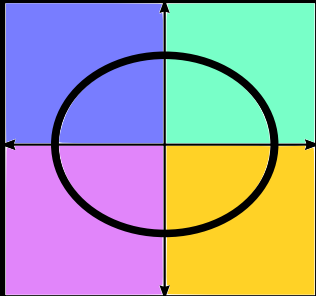


COSMIC STRING FORMATION



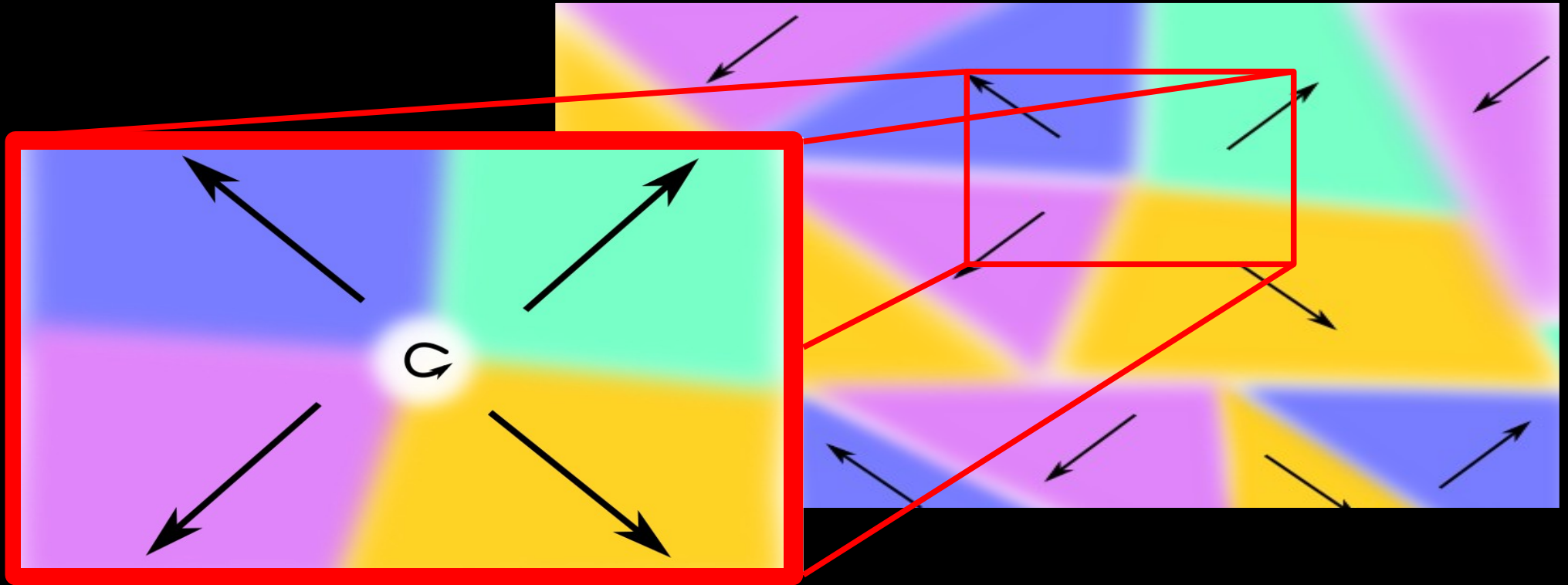
COSMIC STRING FORMATION

Different patches of the universe fall into different minima!



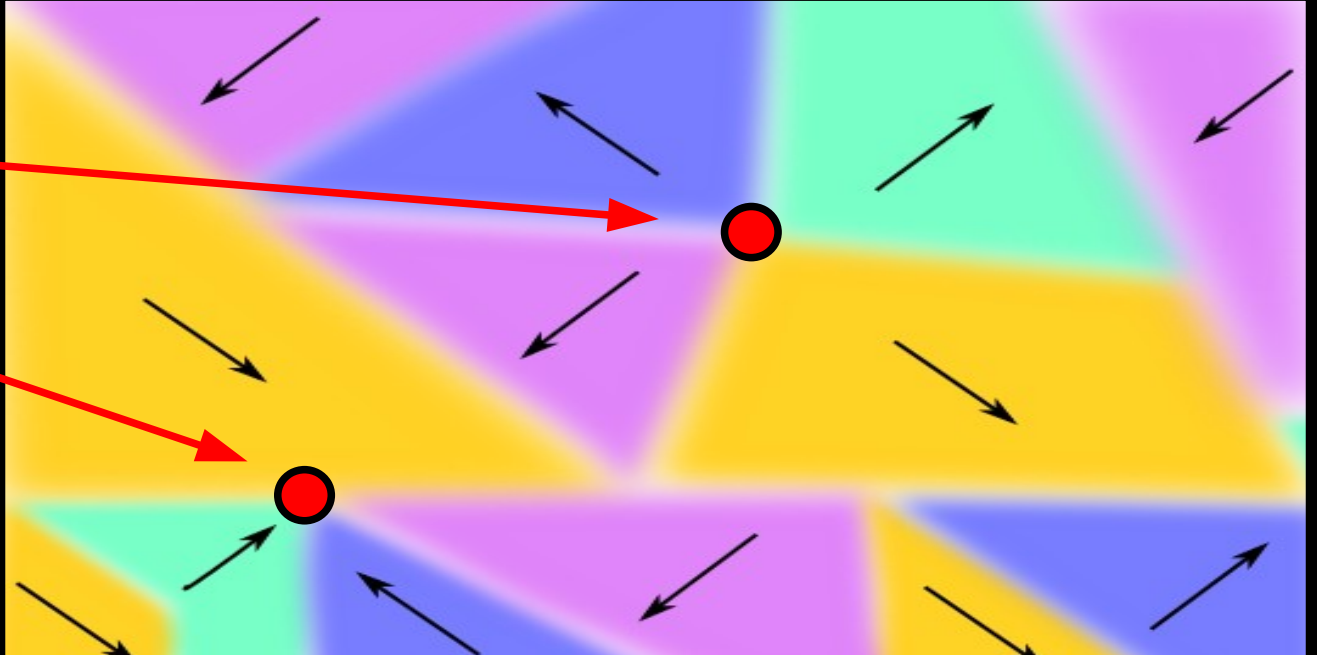
COSMIC STRING FORMATION

What lurks in the edges?



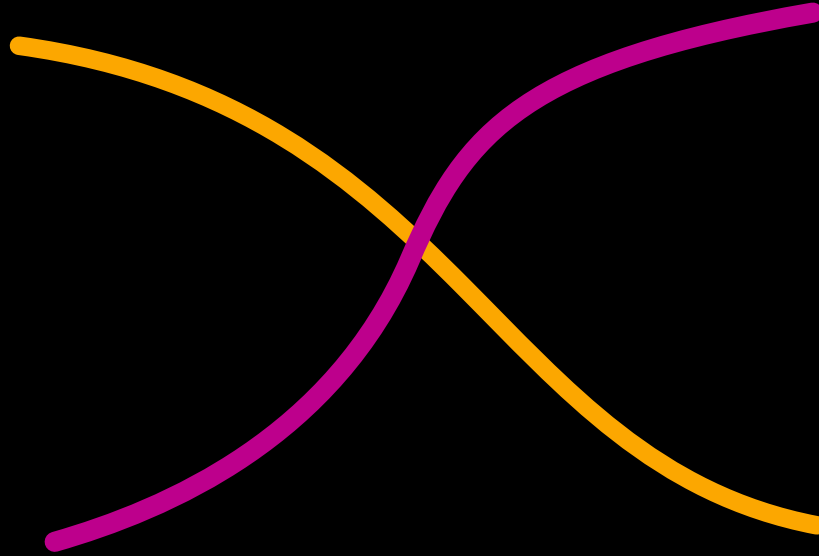
COSMIC STRING FORMATION

COSMIC
STRINGS!



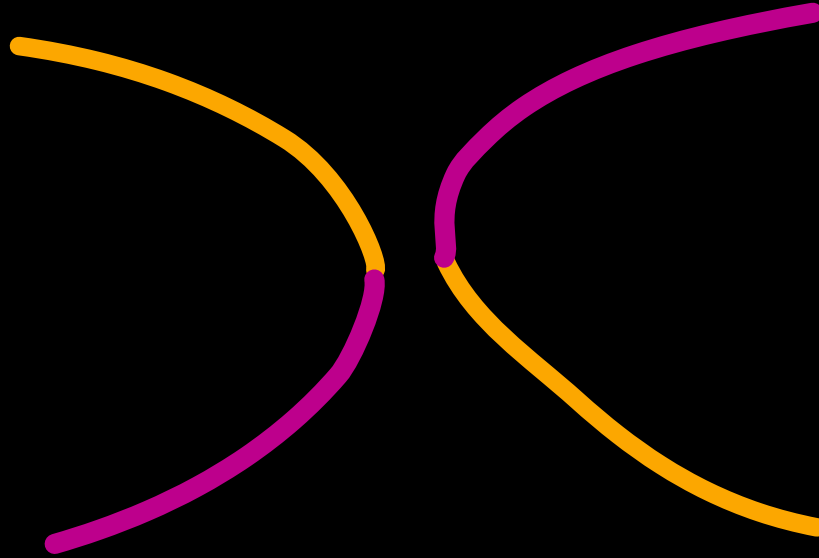
COSMIC STRING LOOPS & GWs

Upon collision



COSMIC STRING LOOPS & GWs

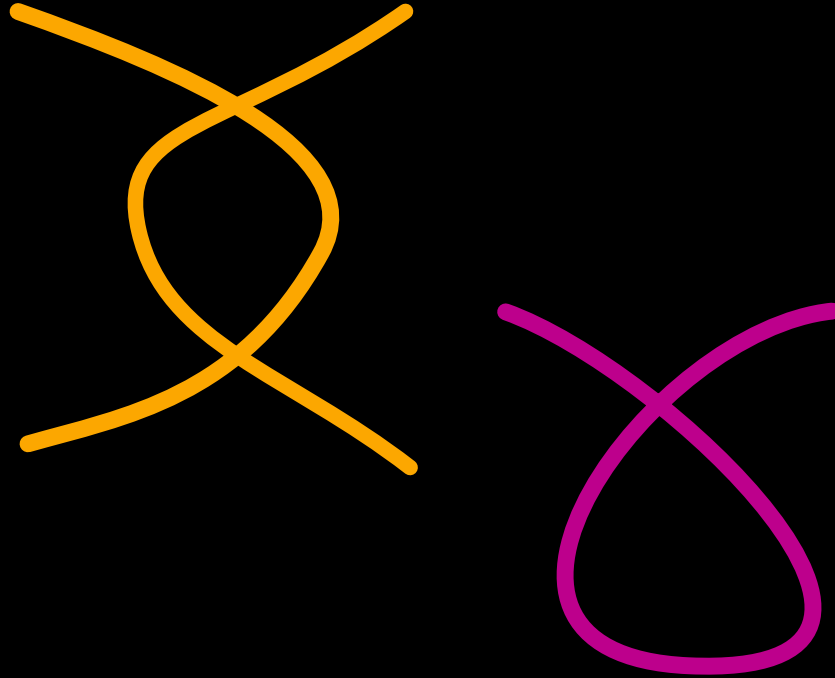
Upon collision



Strings exchange partners and intercommute.

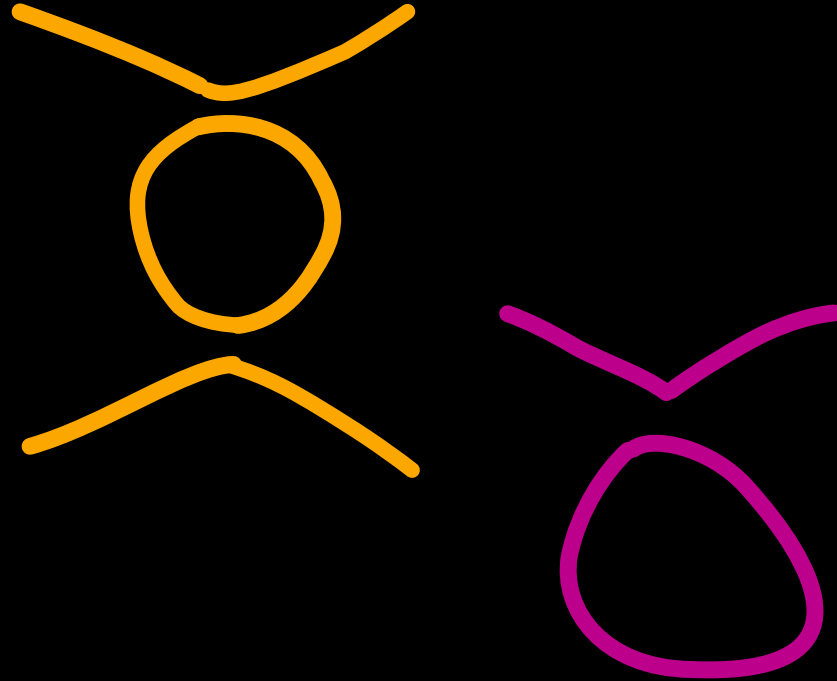
COSMIC STRING LOOPS & GWs

Some situations...



COSMIC STRING LOOPS & GWs

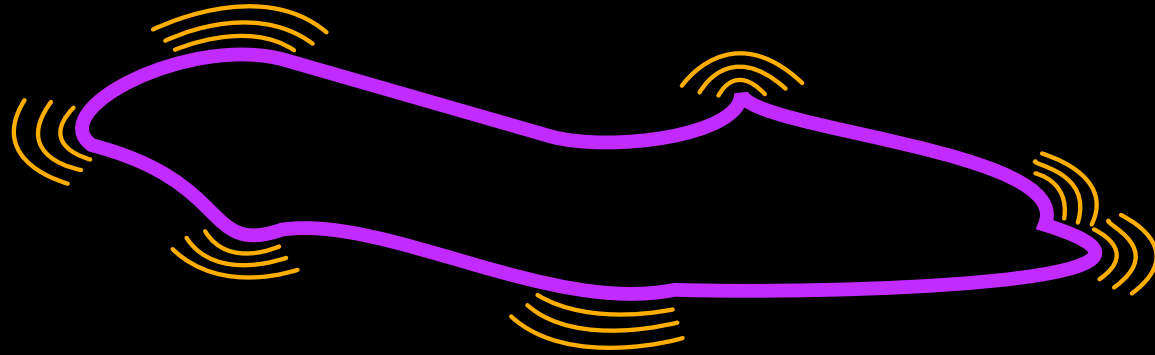
Some situations...



lead to the copious production of cosmic string loops!

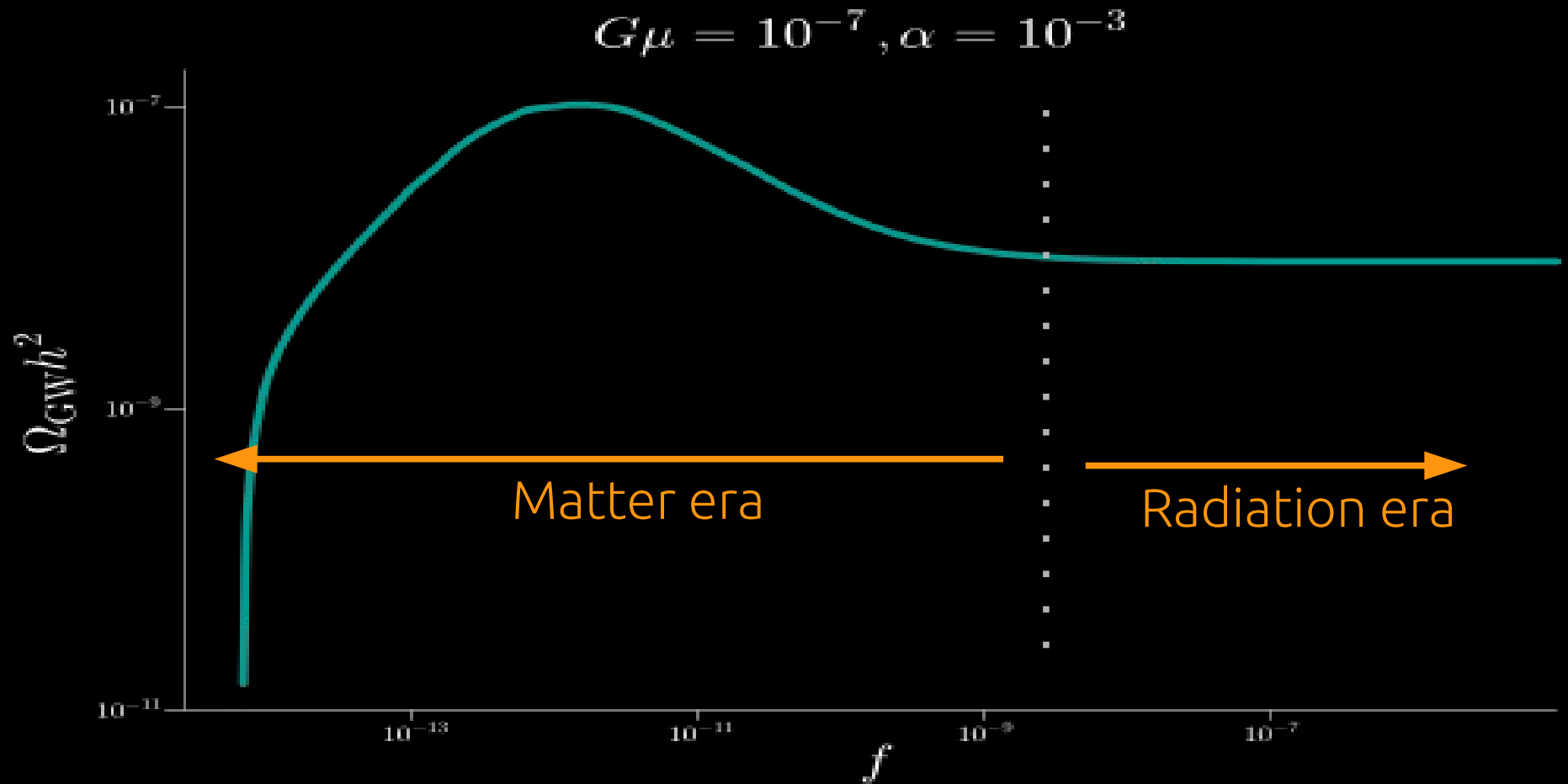
COSMIC STRING LOOPS & GWs

Loops detach from the network and start to oscillate



and generate transient gravitational wave signals.

THE TYPICAL COSMIC STRING SGWB



LOOPS AND GRAVITATIONAL WAVES

TO COMPUTE THE SGWB GENERATED
BY COSMIC STRINGS WE NEED TO KNOW THE
NUMBER DENSITY OF LOOPS $n(\ell(t), t)$

How many loops
are created?

How much of the
energy of the loops
goes into GWs?

What is the
emission
spectrum of
loops?

SUPERCONDUCTING STRINGS

In many GUTs, cosmic strings may carry currents and behave as thin superconductor wires (Witten 1985).

Spontaneous current generation may happen even in the standard GUT paradigm (Davis and Peter, 1995).

SUPERCONDUCTING STRINGS

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REALISTIC COSMIC STRINGS MAY
BE CURRENT-CARRYING AND THUS HAVE
INTERNAL DEGREES OF FREEDOM

SUPERCONDUCTING STRINGS

SUPERCONDUCTING STRINGS BEHAVE AS ELASTIC STRINGS which have a distinct tension and energy per unit length

There are two types of perturbations:

- * Wiggles: transverse displacements
- * Woggles: longitudinal displacements

SUPERCONDUCTING STRINGS

SUPERCONDUCTING STRINGS BEHAVE AS **ELASTIC STRINGS** which have a distinct tension and energy per unit length

There are two types of perturbations:

- * Wiggles: transverse displacements
- * Woggles: longitudinal displacements

We use **TRANSONIC ELASTIC STRINGS**, for which the speed of propagation of wiggles and woggles are equal, as a proxy to study the gravitational wave background generated by superconducting strings

SUPERCONDUCTING STRINGS

TRANSONIC STRING LOOPS

$$X^\mu = \frac{T_\ell}{2\pi} [X_+^\mu(\sigma_+) + X_-^\mu(\sigma_-)] ,$$

$$\phi = \frac{T_\ell}{2\pi} [F_+(\sigma_+) + F_-(\sigma_-)] ,$$

with

$$\mathbf{X}'^2_+(\sigma_+) = 1 - F'^2_+(\sigma_+), \quad \mathbf{X}'^2_-(\sigma_-) = 1 - F'^2_-(\sigma_-)$$

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GW EMISSION EFFICIENCY

Power emitted in GWs:

$$P = G\mu_0^2\Gamma, \quad \Gamma \equiv \sum_{j=0}^{\infty} \Gamma_j = \sum_{j=0}^{\infty} \int \omega_j^2 \frac{|I_j^{22} - I_j^{33}|^2 + |I_j^{23} + I_j^{32}|^2}{2^3 \pi T_\ell^2} d\Omega$$

GRAVITATIONAL WAVE
EMISSION EFFICIENCY

GW EMISSION EFFICIENCY

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GRAVITATIONAL WAVE
EMISSION EFFICIENCY

SPECTRUM OF
EMISSION:

Tells us how the power is
split into the different
harmonic modes

$$\omega_j = 2\pi j / T_\ell$$

GW EMISSION EFFICIENCY

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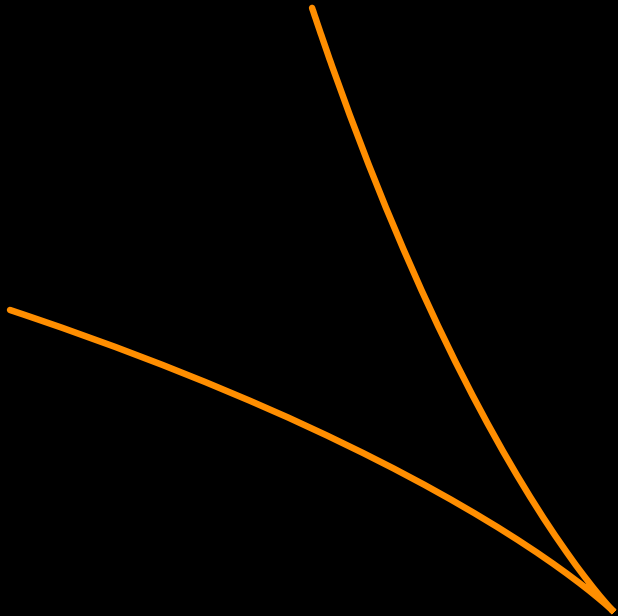
with

$$I_\pm^\mu = \frac{T_\ell}{\pi} \int_0^{2\pi} X_\pm'^\mu e^{-\frac{iT_\ell}{2\pi} k_\nu X_\pm^\nu} d\sigma_\pm$$

$$\tilde{T}^{\mu\nu}(\omega, \mathbf{k}) = \frac{1}{T_\ell} \int T^{\mu\nu} e^{-ik_\nu X^\nu} d^4x = \frac{\mu_0}{2T_\ell} I_+^{(\mu} I_-^{\nu)}$$

EMISSION FROM A CUSP-LIKE POINT

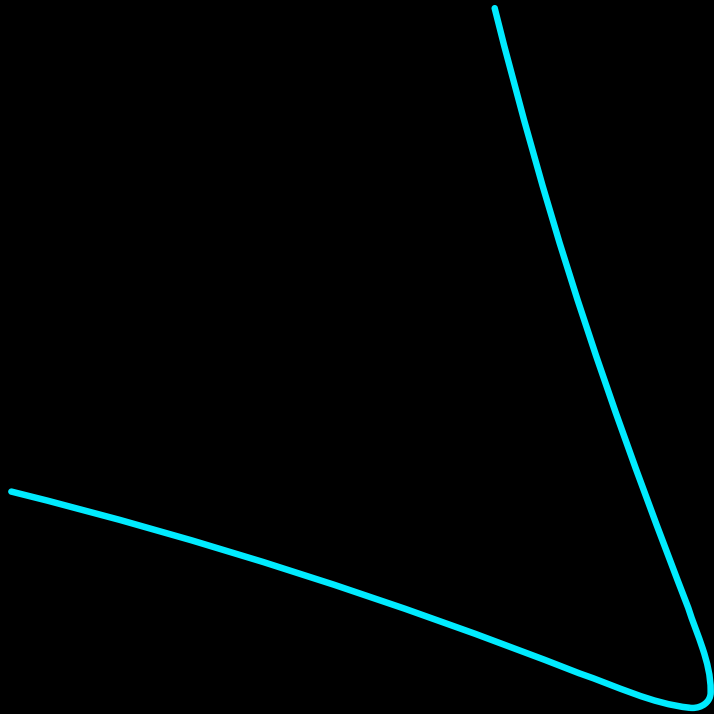
At some points on string loops, known as **CUSPS**, string velocity can equal the speed of light.



CUSPS EMIT
HIGHLY BEAMED
GRAVITATIONAL WAVE
BURSTS

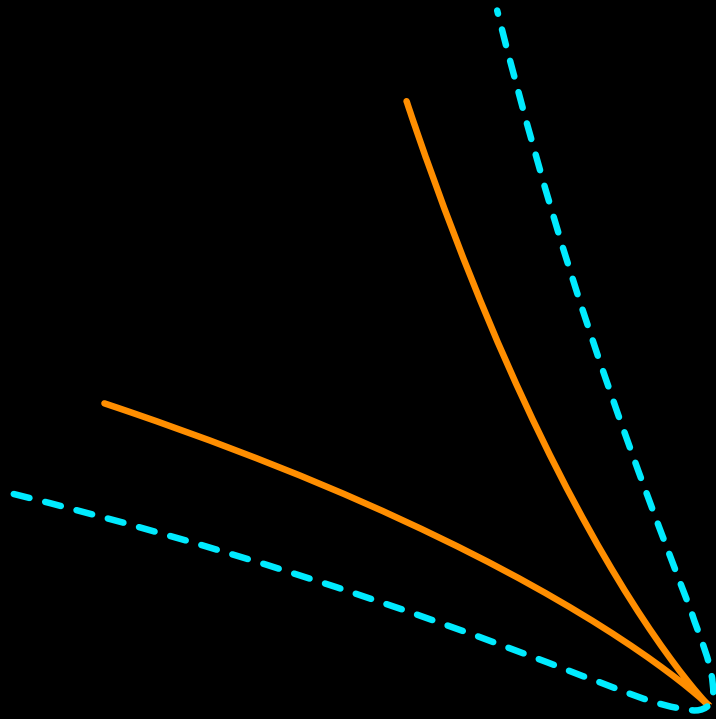
EMISSION FROM A CUSP-LIKE POINT

For current-carrying strings, CUSPS CAN NO LONGER FORM.



EMISSION FROM A CUSP-LIKE POINT

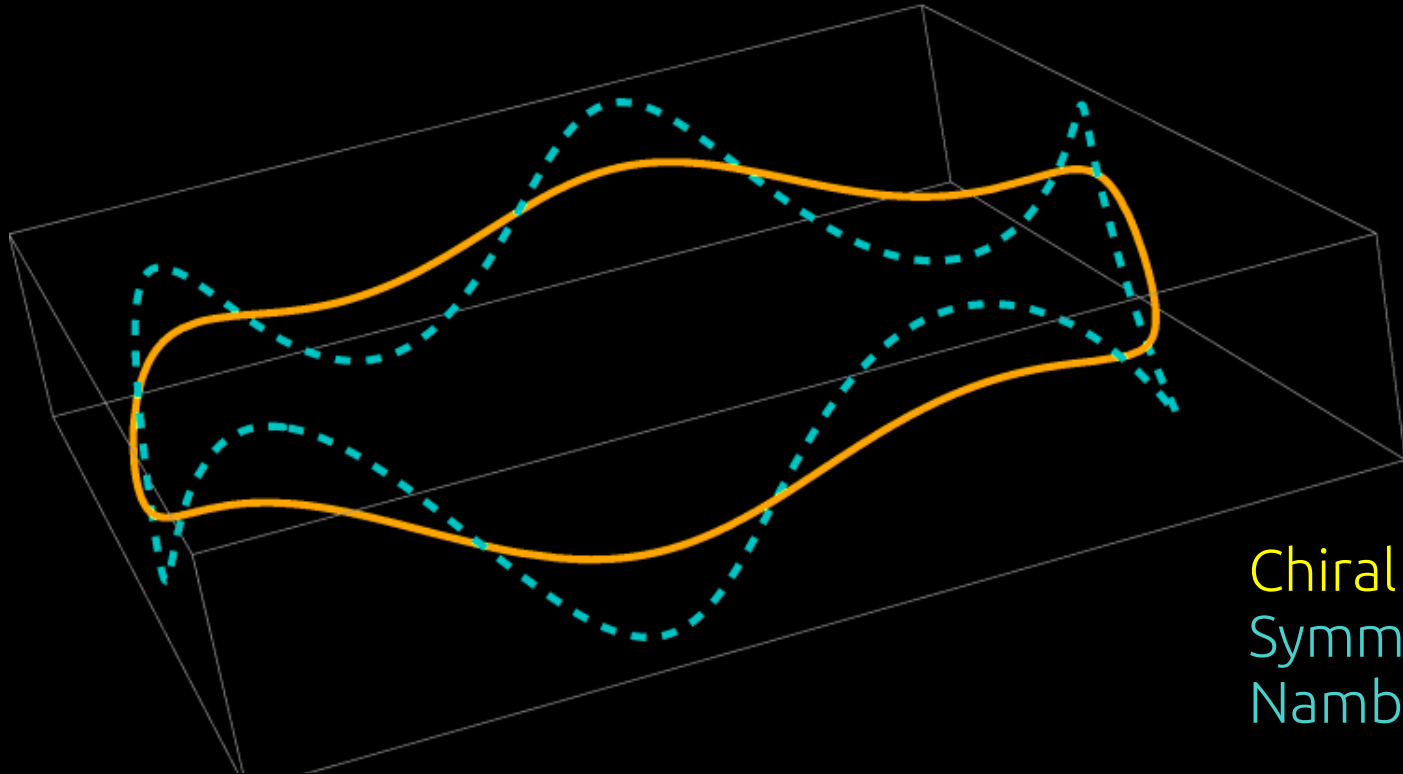
For current-carrying strings, CUSPS CAN NO LONGER FORM.



IF CURRENT
IS SYMMETRICAL,
HOWEVER,
CUSP-LIKE POINTS MAY
FORM, BUT NOW
THEIR VELOCITY IS
SUBLUMINAL

EMISSION FROM BURDEN LOOPS

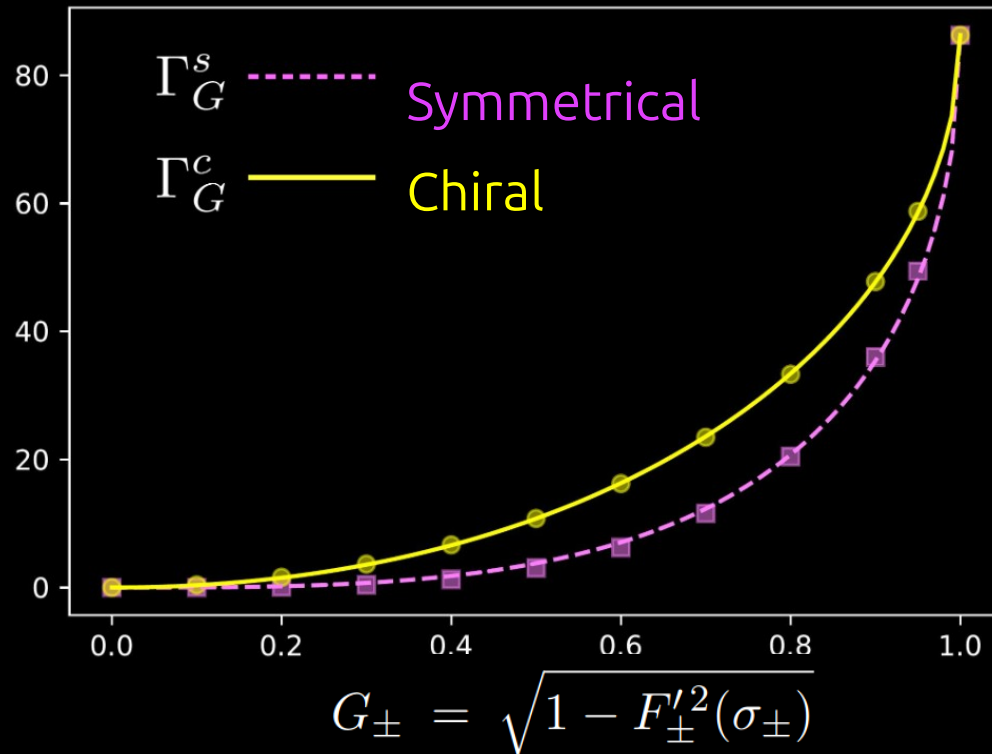
Burden Loops create cusps (or cusp-like points) at discrete instants of time



Chiral
Symmetrical and
Nambu-Goto

SPECTRUM OF EMISSION OF BURDEN LOOPS

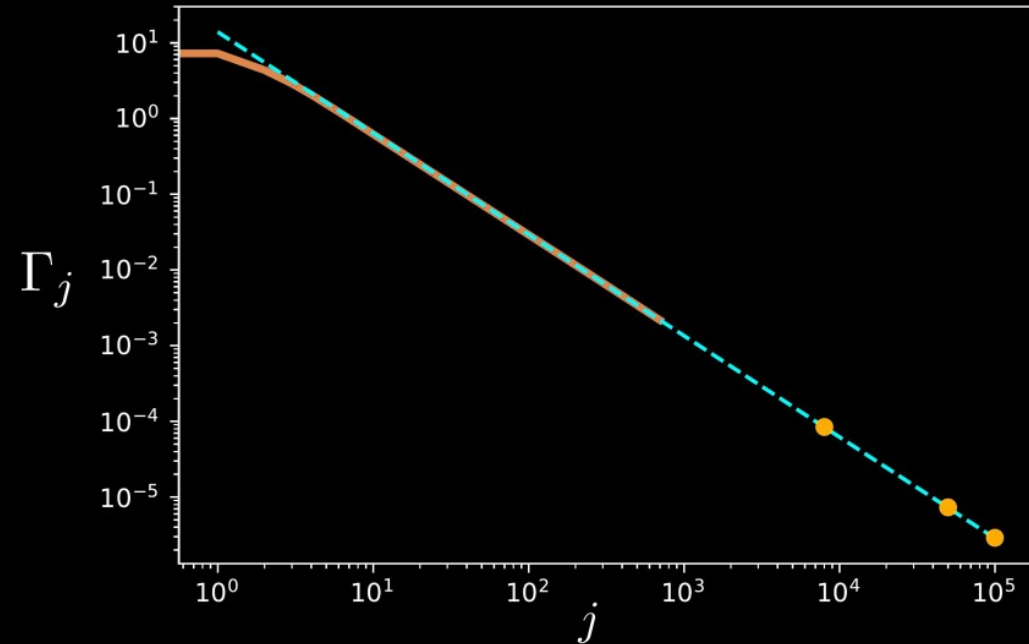
Gravitational wave emission efficiency decreases with increasing current



SPECTRUM OF EMISSION OF BURDEN LOOPS

THE PRESENCE OF CURRENT
LEADS TO A SUPPRESSION OF THE
EMISSION OF GRAVITATIONAL
RADIATION

SPECTRUM OF EMISSION OF BURDEN LOOPS



Without current, we have a power law spectrum

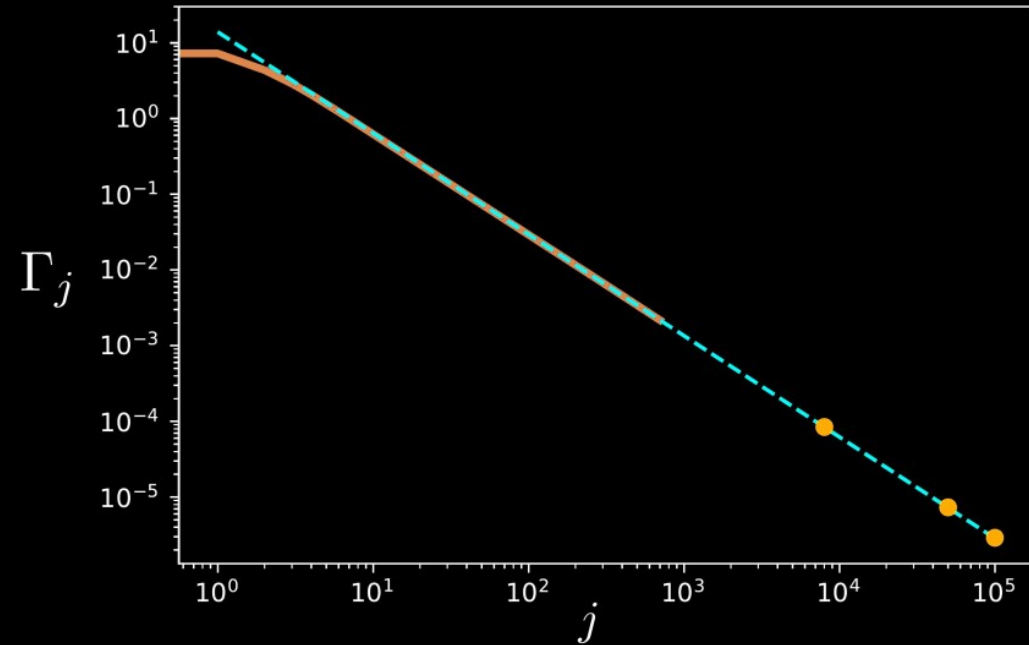
$$\Gamma_j \sim j^{-q}$$

For large enough j .

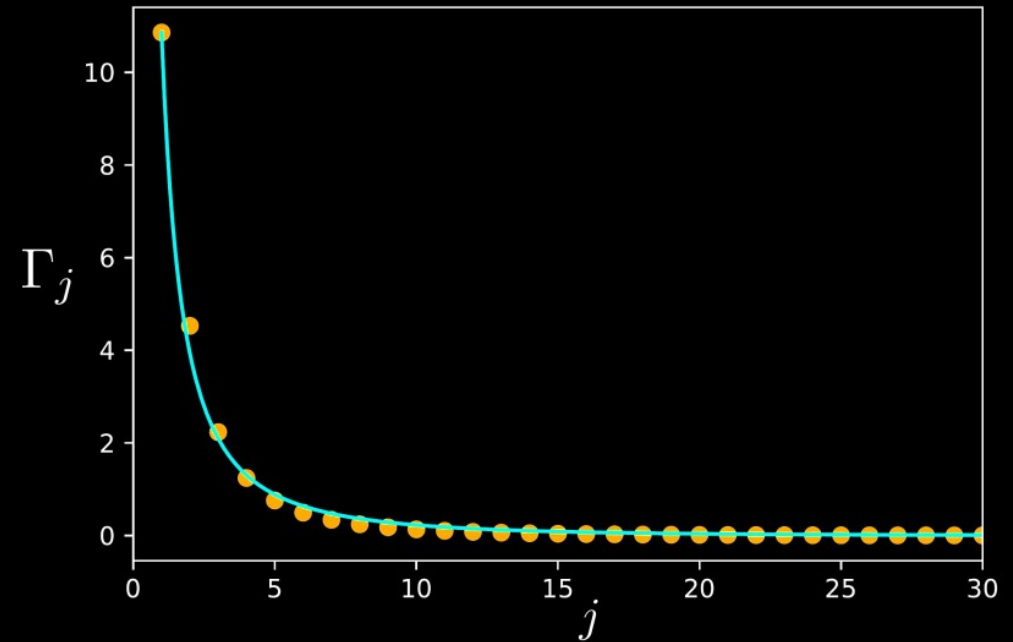
(Vaschaspati & Vilenkin, 1985)

SPECTRUM OF EMISSION OF BURDEN LOOPS

Without current



With current



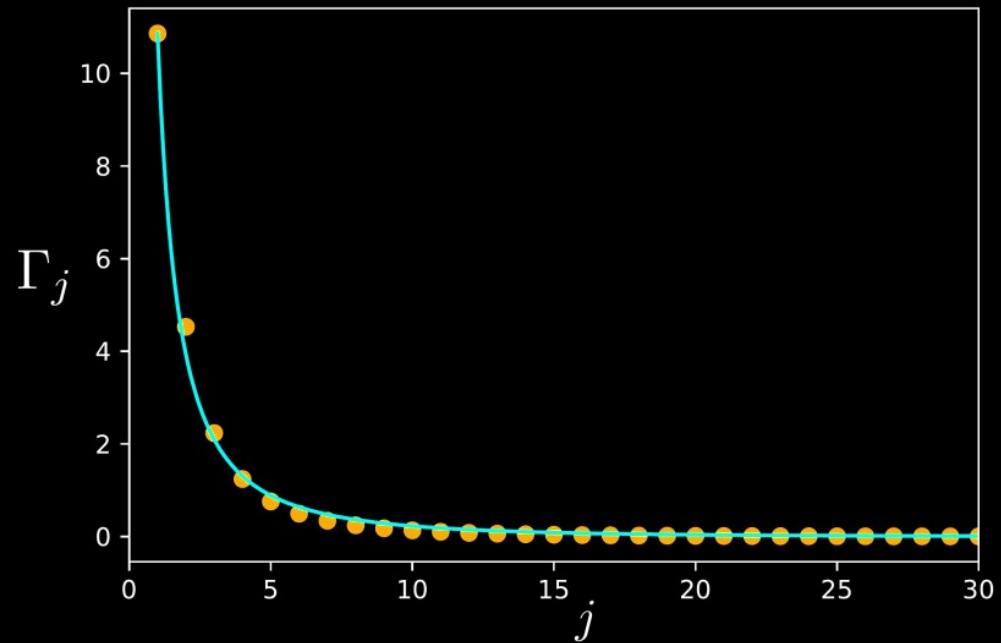
SPECTRUM OF EMISSION OF BURDEN LOOPS

With current, we have an
exponential decay

$$\Gamma_j \sim j^{-q} e^{-j f_m(G_{\pm})}$$

with

$$f_m(G_{\pm}) = a_m (1 - \sqrt{G_{\pm}})^{b_m}$$



SPECTRUM OF EMISSION OF BURDEN LOOPS

THE POWER EMITTED IN GWS BY
A LOOP WITH QUASI-CUSPS DECAYS
EXPONENTIALLY WITH INCREASING
HARMONIC MODE FOR
NULL AND SYMMETRICAL CURRENTS

EMISSION FROM A KINK

String collisions and intercommutation give rise to discontinuities in the string tangent, known as **KINKS**.

For Nambu-Goto strings, kinks travel along the string at the speed of light.

KINKS EMIT
“FAN-LIKE”
GRAVITATIONAL WAVE
BURSTS

EMISSION FROM A KINK

String collisions and intercommutation give rise to discontinuities in the string tangent, known as **KINKS**.

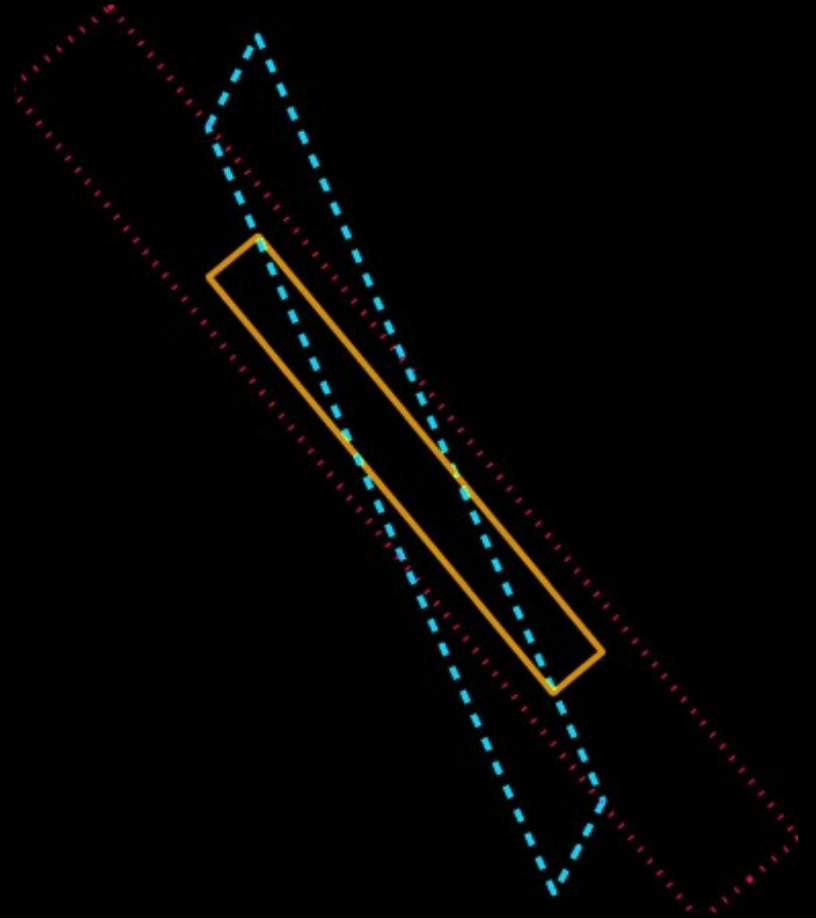
If there is current traveling along the string, kinks are necessarily slower

GRAVITATIONAL WAVE
BURST FROM KINKS
ON SUPERCONDUCTING
STRINGS ARE
WEAKER

EMISSION FROM CUSPLESS LOOPS

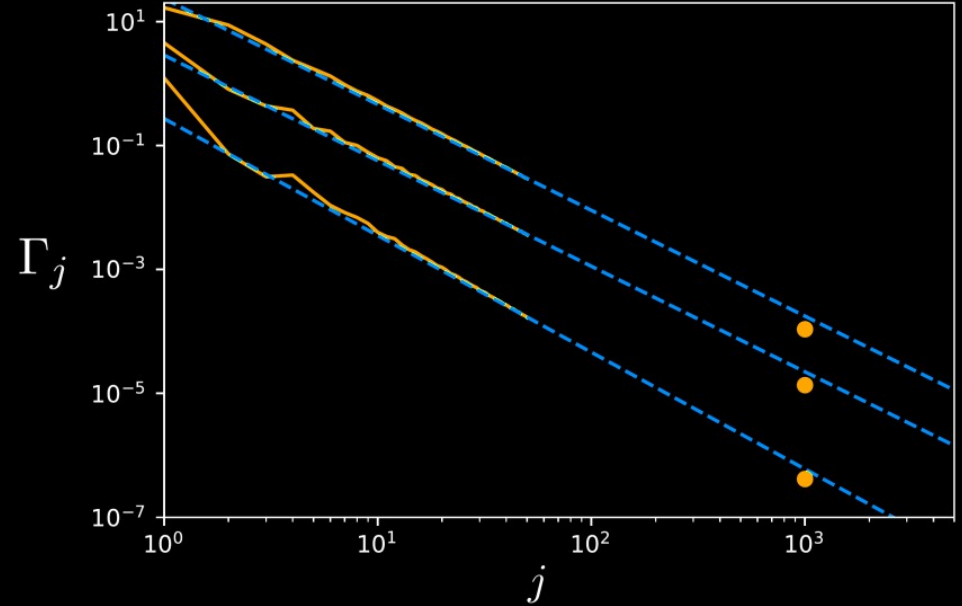
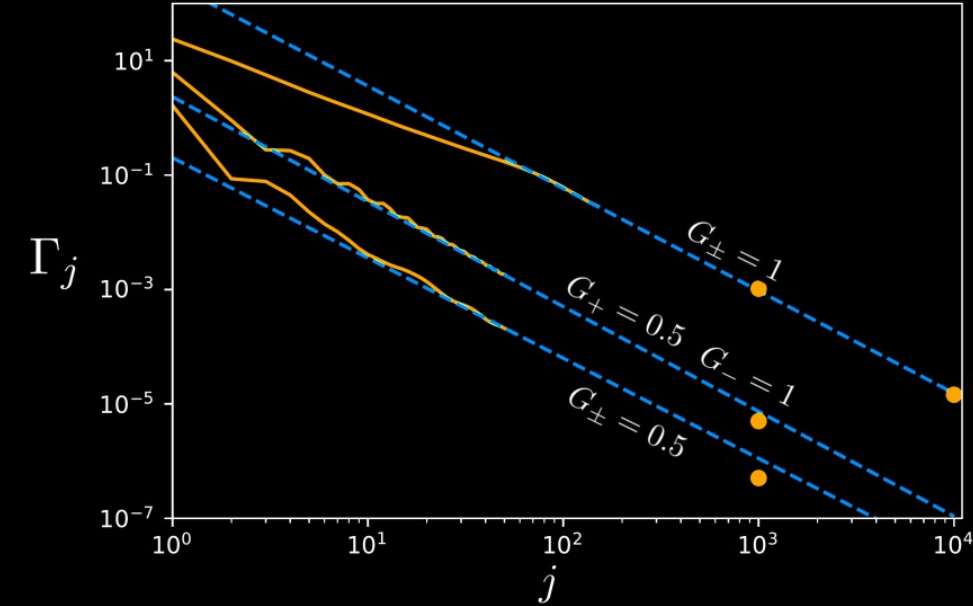
Loops that are
Piece-wise straight
And have no cusps

Chiral
Symmetrical
Nambu-Goto



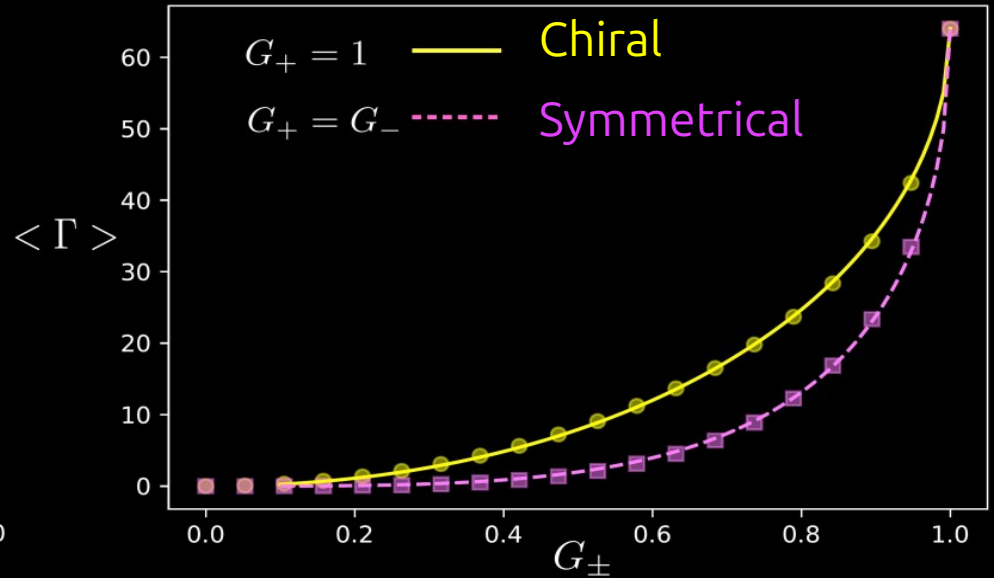
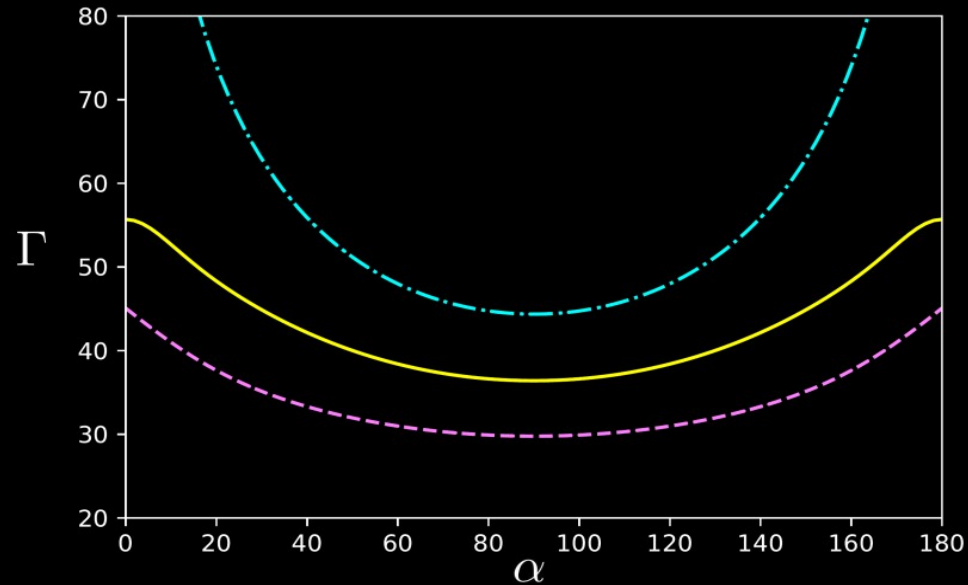
EMISSION FROM CUSPLESS LOOPS

In this case, the spectrum of emission is not affected by the inclusion of current: **it is still a power law**



EMISSION FROM CUSPLESS LOOPS

The gravitational wave efficiency, however, is affected precisely in the same way as before



IMPACT OF CURRENT ON GW EFFICIENCY

GRAVITATIONAL WAVE EMISSION EFFICIENCY FOR CURRENT-CARRYING STRINGS IS GIVEN BY

$$\Gamma_G^m = \Gamma_0 (1 - |F'_\pm|) B_\Gamma^m ,$$

Where, for symmetrical currents,

$$B_\Gamma^s \approx 2 \quad \text{For loops with cusps}$$

$$B_\Gamma^s \approx 1.5 \quad \text{For loops with kinks}$$

And $B_\Gamma^s \approx \sqrt{2} B_\Gamma^c$ for chiral superconducting loops.

LOOPS AND GRAVITATIONAL WAVES

TO COMPUTE THE SGWB GENERATED
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NUMBER DENSITY OF LOOPS $n(\ell(t), t)$

How many loops
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How much of the
energy of the loops
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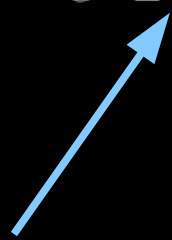
What is the
emission
spectrum of
loops?

LOOPS AND VECTOR RADIATION

Power emitted in Vector radiation:

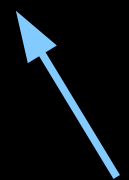
$$P = \sum_j P_j = e^2 \Gamma^{\text{em}},$$

VECTOR RADIATION
EMISSION EFFICIENCY



$$\frac{dP_j}{d\Omega} = -\frac{\omega_j^2}{2\pi} j^\mu j_\mu^* = e^2 \frac{d\Gamma^{\text{em}}}{d\Omega}$$

SPECTRUM OF
EMISSION:



Tells us how the power is
split into the different
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LOOPS AND VECTOR RADIATION

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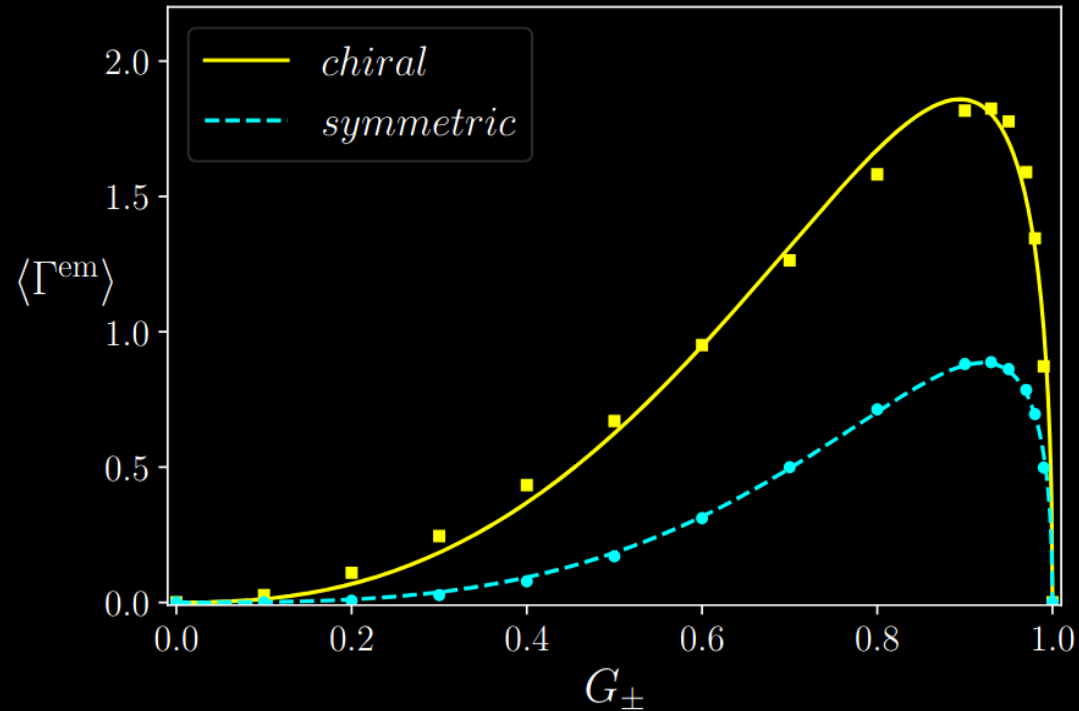
$$P = \sum_j P_j = e^2 \Gamma^{\text{em}}, \quad \frac{dP_j}{d\Omega} = -\frac{\omega_j^2}{2\pi} j^\mu j_\mu^* = e^2 \frac{d\Gamma^{\text{em}}}{d\Omega}$$

with

$$j^\mu = e \frac{L}{8\pi^2} (I_- J_+^\mu - I_+ J_-^\mu),$$
$$I_\pm = \int_0^{2\pi} F'_\pm e^{ij(\sigma_\pm - \frac{2\pi}{L} \mathbf{n} \cdot \mathbf{X}_\pm)} d\sigma_\pm,$$
$$J_\pm^\mu = \int_0^{2\pi} X'^\mu_\pm e^{ij(\sigma_\pm - \frac{2\pi}{L} \mathbf{n} \cdot \mathbf{X}_\pm)} d\sigma_\pm.$$

LOOPS AND VECTOR RADIATION

Vector radiation emission efficiency also decreases with increasing current, but is also suppressed for very low currents



LOOPS AND VECTOR RADIATION

VECTOR RADIATION EMISSION EFFICIENCY FOR CURRENT-CARRYING STRINGS IS GIVEN BY

$$\langle \Gamma^{\text{em}} \rangle = \Gamma_0^{\text{em}} |F'_{\pm}| (1 - |F'_{\pm}|)^D$$

With

	Γ_0^{em}	D
Symmetric loop with quasi-cusps	4.9	1.6
Symmetric loop with kinks	10.5	1.8
Chiral loop with quasi-cusps	8.6	1.1
Chiral loop with kinks	8.6	1.2

SUPERCONDUCTING LOOPS EVOLUTION

Loop length: $\dot{\ell} = -G\mu_0\Gamma^{\text{gr}}(\mathcal{V}) - \tilde{e}^2\Gamma^{\text{em}}(\mathcal{V}),$

SUPERCONDUCTING LOOPS EVOLUTION

Loop length: $\dot{\ell} = -G\mu_0\Gamma^{\text{gr}}(\mathcal{Y}) - \tilde{e}^2\Gamma^{\text{em}}(\mathcal{Y}),$

Current: $\dot{\mathcal{Y}}\Big|_{\text{rad}} = -\frac{\mathcal{Y}}{\ell} \dot{\ell}\Big|_{\text{rad}}$

SUPERCONDUCTING LOOPS EVOLUTION

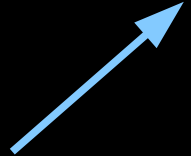
Loop length:

$$\dot{\ell} = -G\mu_0\Gamma^{\text{gr}}(\mathcal{Y}) - \tilde{e}^2\Gamma^{\text{em}}(\mathcal{Y}),$$

Current:

$$\dot{\mathcal{Y}} = \frac{\mathcal{Y}}{\ell} [G\mu_0\Gamma^{\text{gr}}(\mathcal{Y}) + \tilde{e}^2\Gamma^{\text{em}}(\mathcal{Y}) - A(\mathcal{Y})]$$

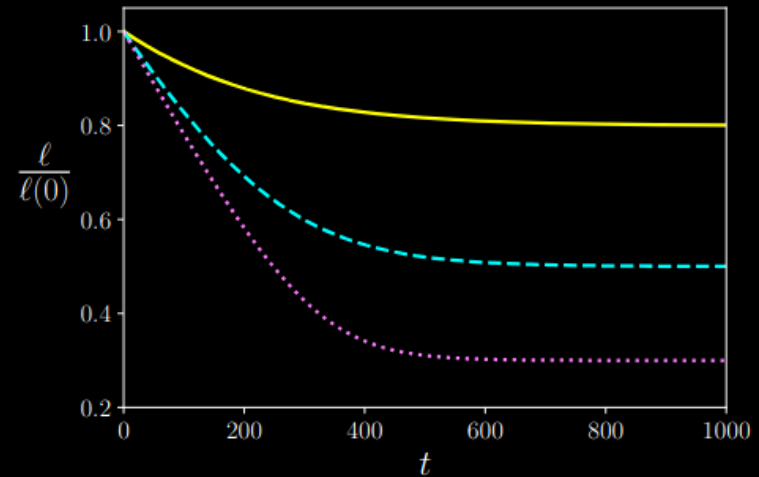
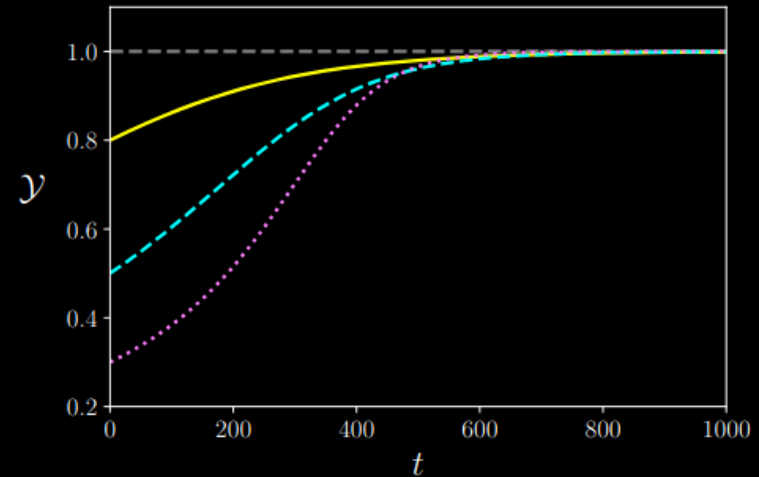
Charge leakage



SUPERCONDUCTING LOOPS EVOLUTION

In the absence of charge leakage,
VORTONS FORM

NO SIGNIFICANT
GW EMISSION
IS
EXPECTED

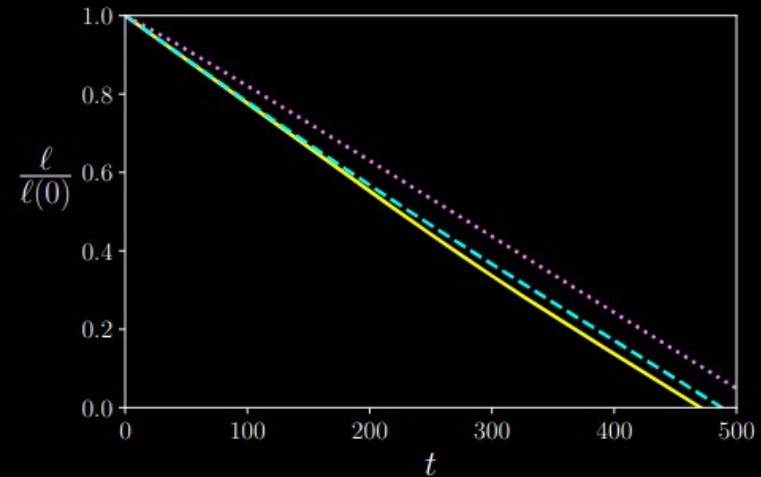
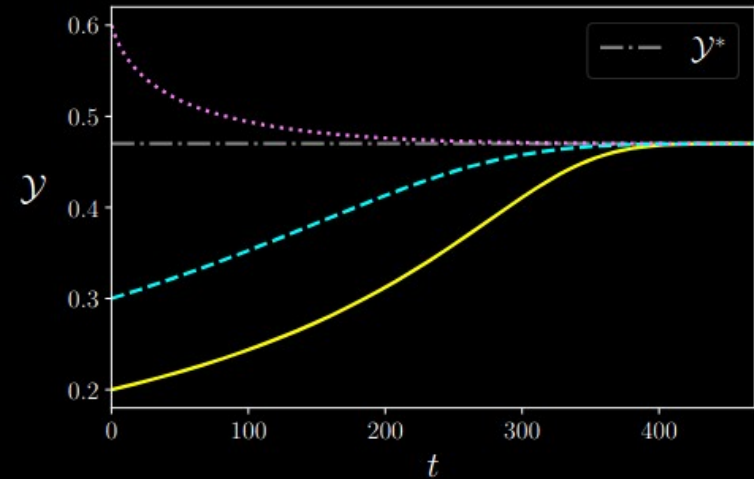


SUPERCONDUCTING LOOPS EVOLUTION

But in the presence of charge leakage, A SCALING SOLUTION EMERGES:

$$\dot{\ell} = \text{constant} \quad \text{and} \quad \dot{\mathcal{Y}} = 0$$

$$G\mu_0\Gamma^{\text{gr}}(\mathcal{Y}^*) + \tilde{e}^2\Gamma^{\text{em}}(\mathcal{Y}^*) = A(\mathcal{Y}^*)$$



SUPERCONDUCTING LOOPS EVOLUTION

WE MAY HAVE A SIGNIFICANT
SGWB FROM SUPERCONDUCTING
LOOPS WITH CONSTANT CURRENT

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(SEMI-)ANALYTICAL APPROACH

Sousa & Avelino 2013 (arXiv: 1304.2445); Sousa, Avelino & Guedes 2020 (arXiv: 2002.01709)

PREMISE: the large-scale evolution of the string network determines how much energy goes into loops:

$$\dot{\rho}_\ell = \left. \frac{d\rho}{dt} \right|_{loops} = \tilde{c}v \frac{\rho}{L} \sim \frac{v}{L^3}$$

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**NUMBER DENSITY
OF LOOPS
PRODUCED:**

$$\dot{n}_l = \frac{\dot{\rho}_\ell}{\alpha L}$$

EVOLUTION OF STRINGS WITH CURRENT

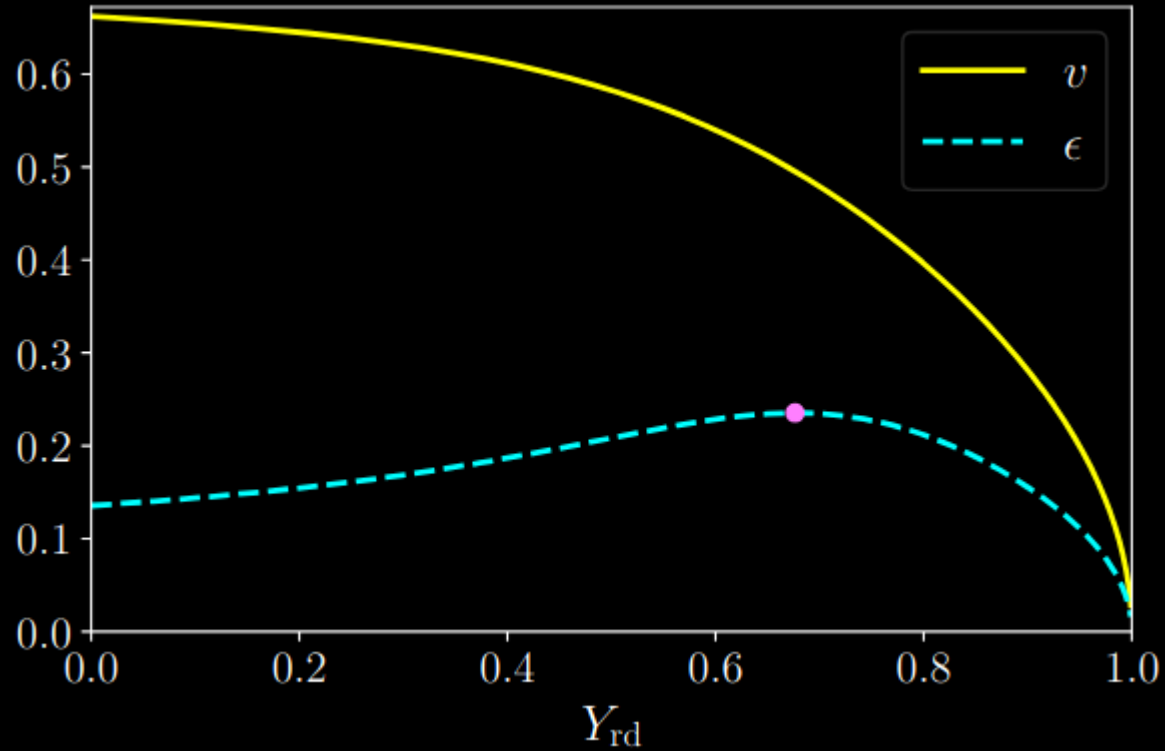
Martins, Peter, Rybak & Shellard 2021 (arXiv: 2011.09700 and arXiv: 2108.03147)

IN THE RADIATION ERA:

$$L_{\text{ph}} = \xi_r t \quad \text{and} \quad v = v_r ,$$

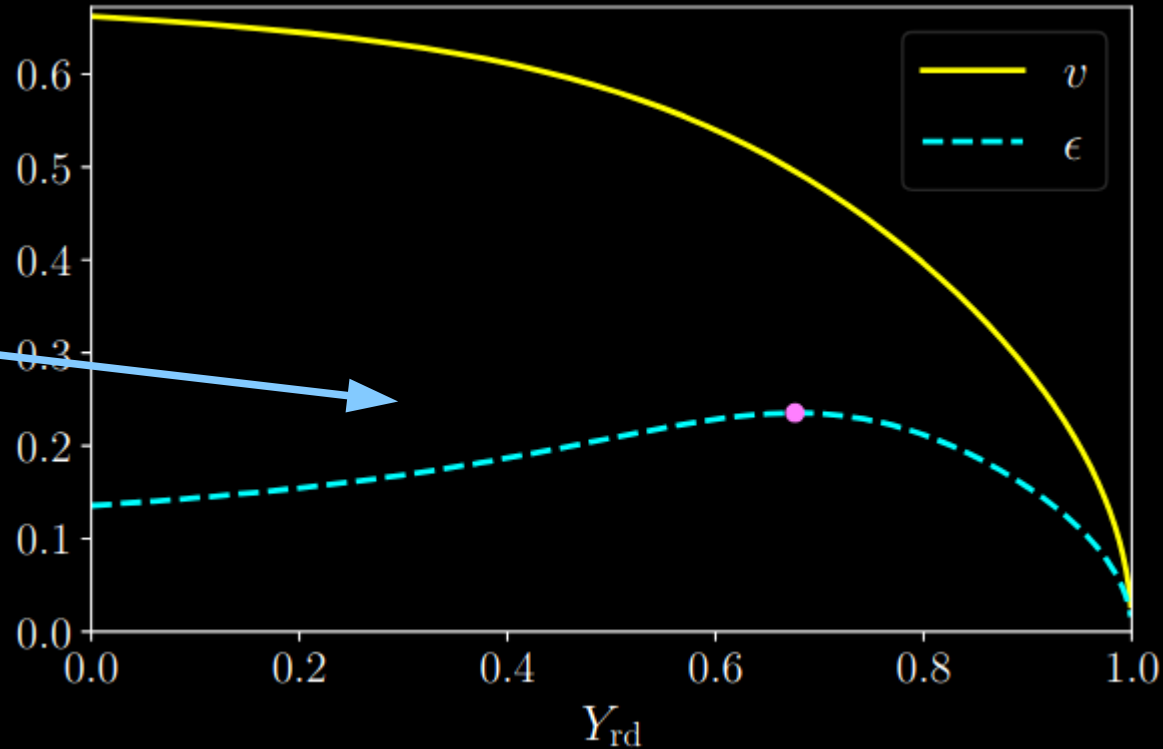
$$\xi_r^2 = (1 - Y)k(k + \tilde{c}) \quad \text{and} \quad v_r^2 = (1 - Y)\frac{k}{k + \tilde{c}}$$

EVOLUTION OF STRINGS WITH CURRENT



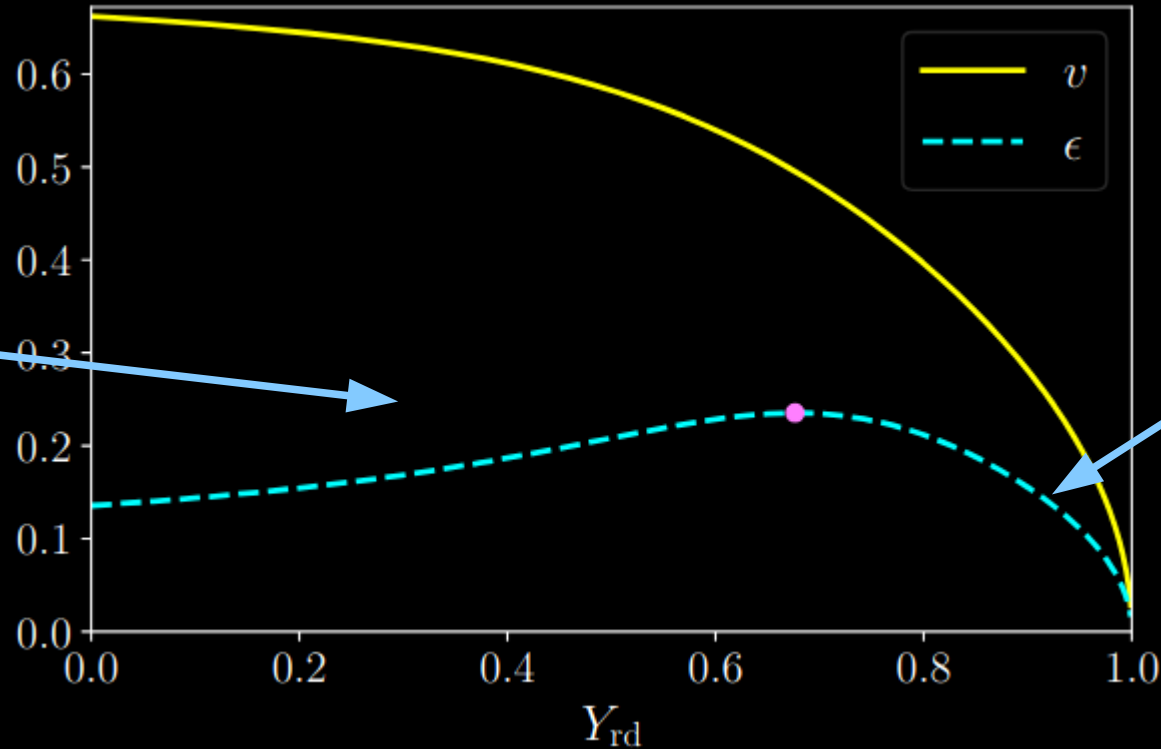
EVOLUTION OF STRINGS WITH CURRENT

Number of
loops
produced is
suppressed
for small
currents



EVOLUTION OF STRINGS WITH CURRENT

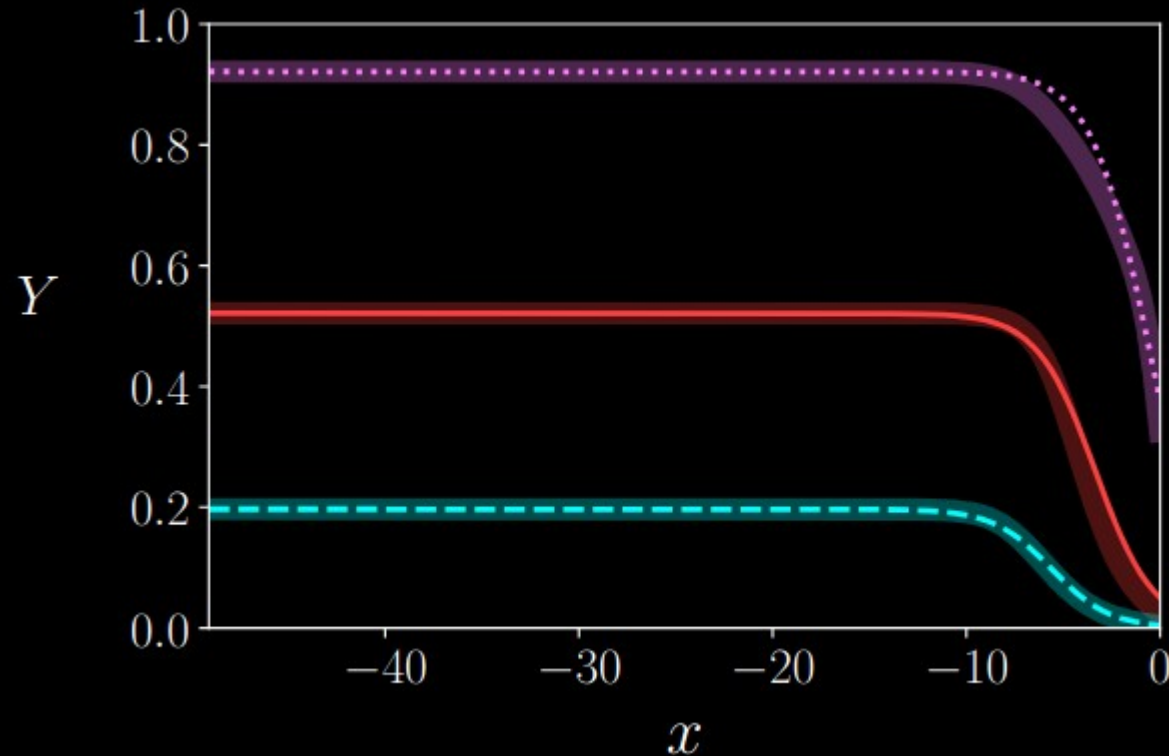
Number of loops produced is suppressed for small currents



Number of loops produced is enhanced for large currents

EVOLUTION OF STRINGS WITH CURRENT

Current is quickly dissipated in the matter era



IMPACT OF CURRENT ON THE SGWB

IN THE RADIATION ERA:

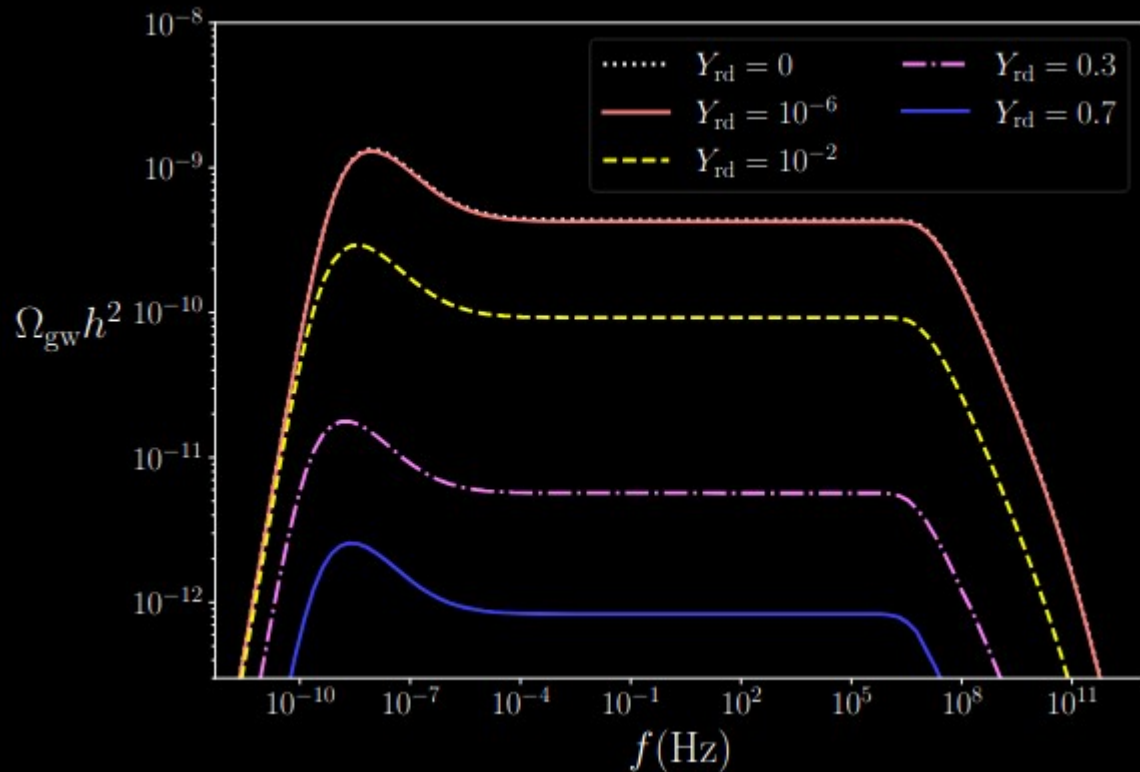
THE AMPLITUDE OF THE PLATEAU
MAY BE SIGNIFICANTLY AFFECTED AND
DEPENDS ON THE VALUE OF CURRENT

IN THE MATTER ERA:

THE CONTRIBUTION OF MATTER-ERA
LOOPS IS (MOSTLY) UNNAFFECTED

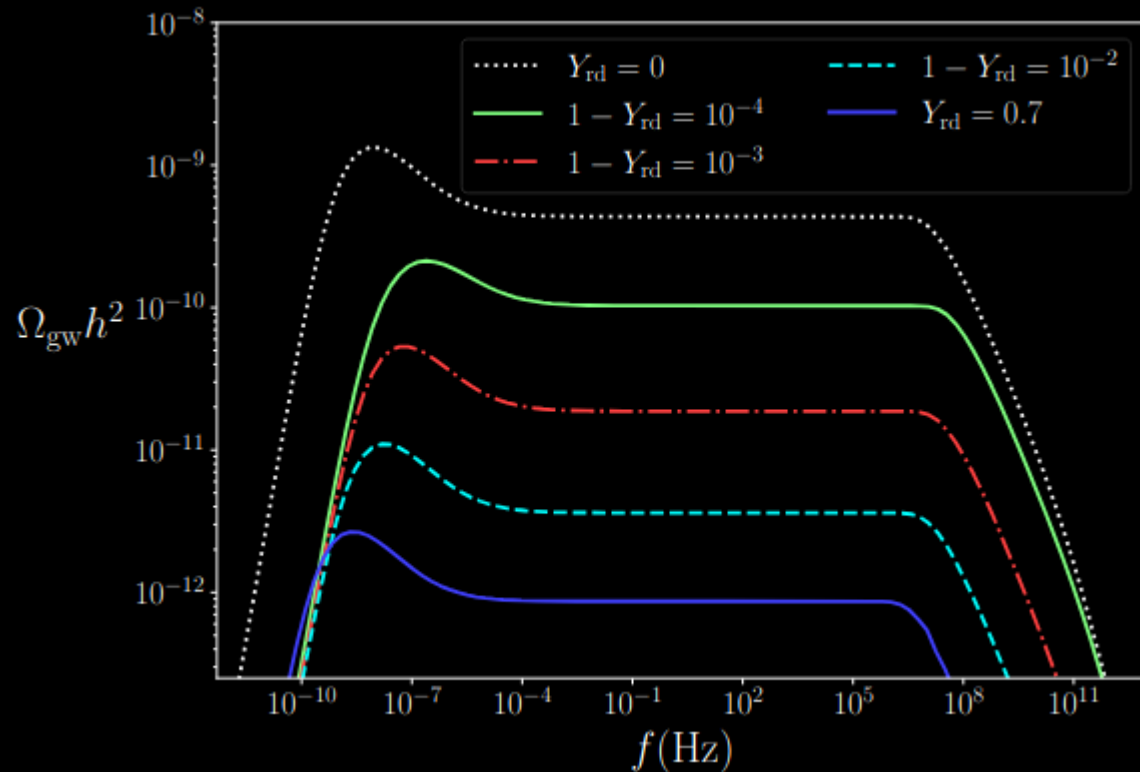
IMPACT OF CURRENT ON THE SGWB

For small enough currents:



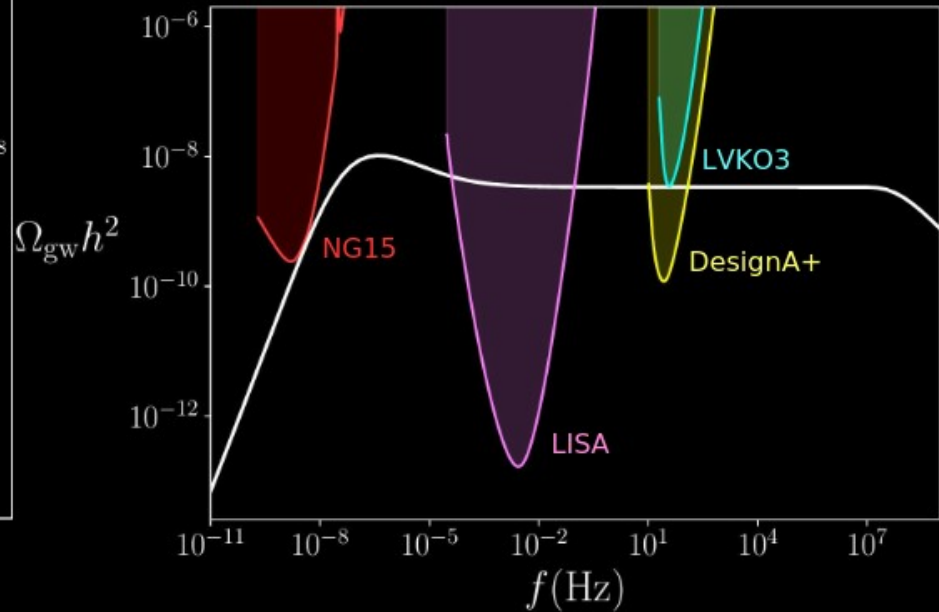
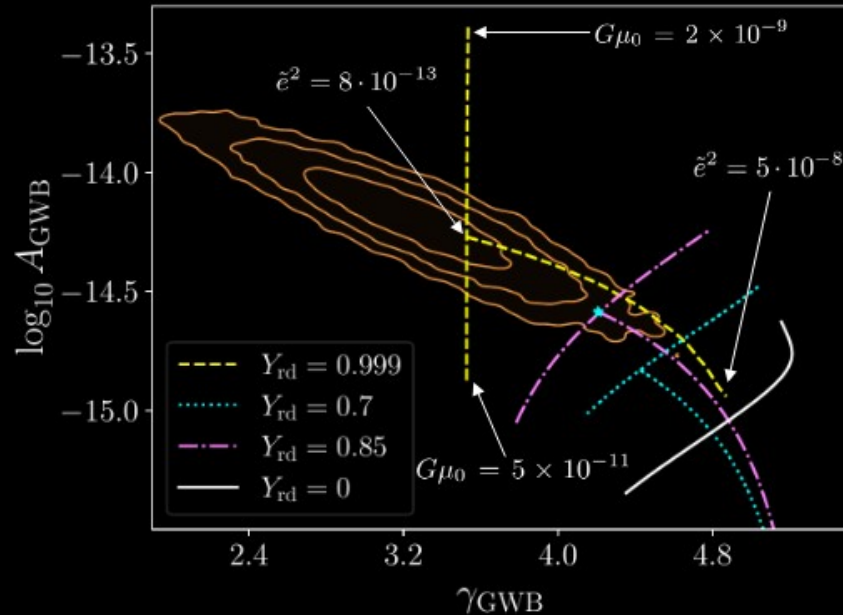
IMPACT OF CURRENT ON THE SGWB

For large enough currents:



IMPACT OF CURRENT ON THE SGWB

Superconducting strings with high current fit better NANOGrav data than standard cosmic strings

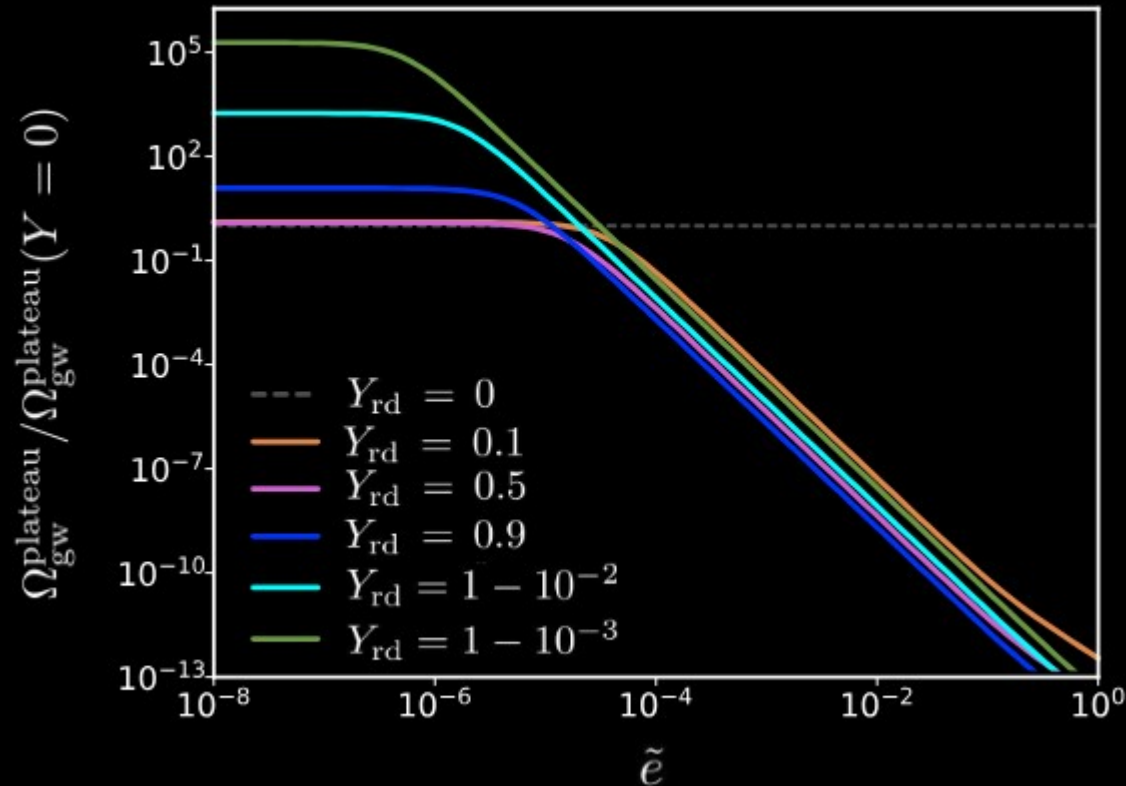


IMPACT OF CURRENT ON THE SGWB

CURRENT MAY ACTUALLY HELP
RECONCILE COSMIC STRINGS WITH
NANOGRAB DATA!

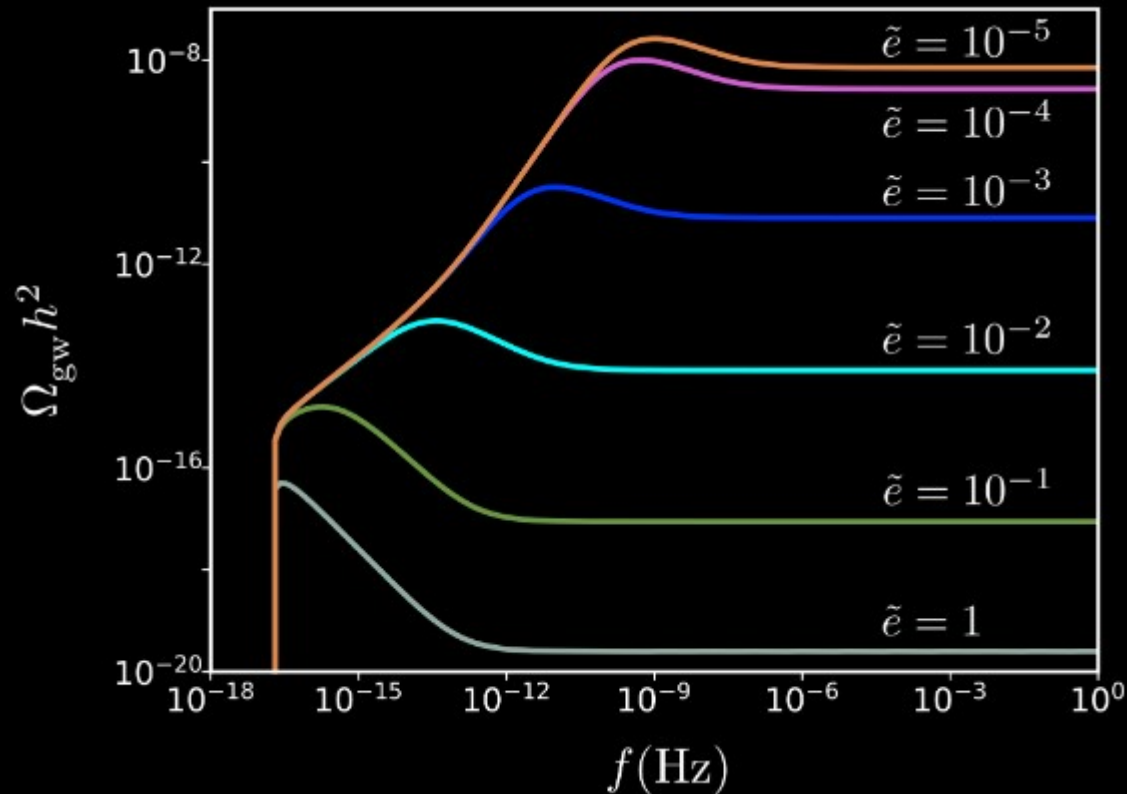
IMPACT OF CHARGE ON THE SGWB

As the charge of the current carriers increases, the amplitude is suppressed:



IMPACT OF CHARGE ON THE SGWB

But the shape of the spectrum is also affected



IMPACT OF CURRENT ON THE SGWB

THERE MAY BE DISTINCTIVE SIGNATURES
ON THE SGWB SPECTRUM GENERATED
BY STRINGS, WHICH MAY ALLOW US
TO PROBE THE UNDERLYING
PARTICLE PHYSICS SCENARIO

TO SUM UP

- *Current may significantly decrease the efficiency of emission of gravitational radiation by cosmic strings and affect the spectrum of emission;
- *The amplitude of the SGWB for chiral superconducting strings is highly dependent on the value of current and on the charge of current carriers;
- *There may also be significant changes to the shape of the spectrum if the charge of current carriers is large enough;

THANKS!