Constraining the axiverse with reionization

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Based on 2506.19096 and paper to appear soon

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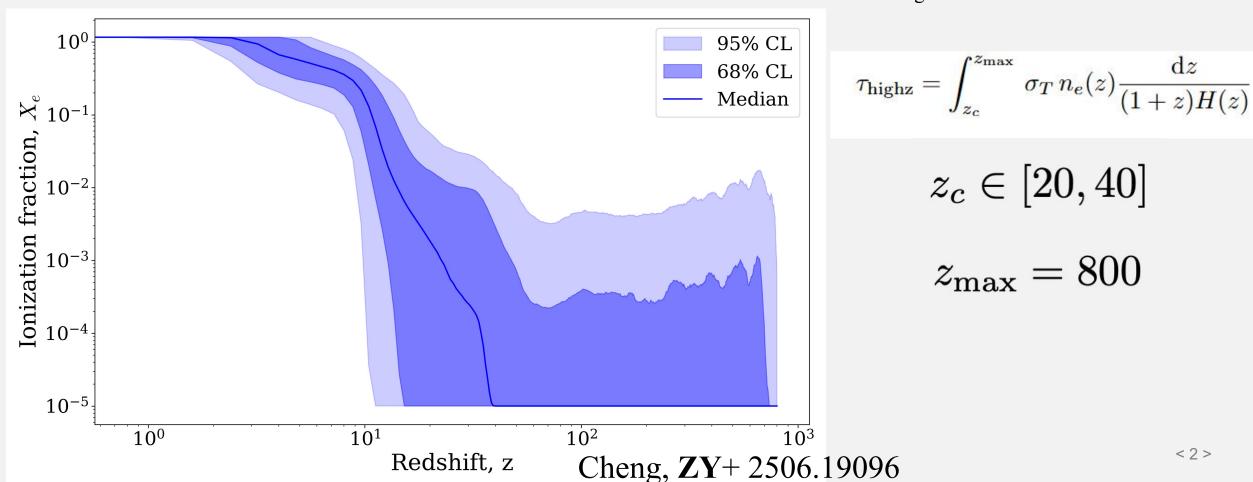
Model independent method for reionization history reconstruction



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We develop a Gaussian Process Regression to constraint the ionization fraction

We use *Planck* low-*l* EE polarization to compute the optical depth τ_{highz}



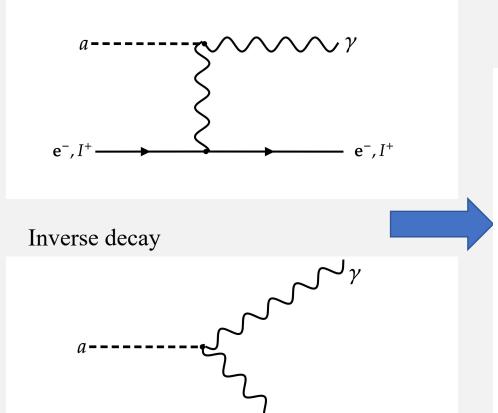
Axion production and decay



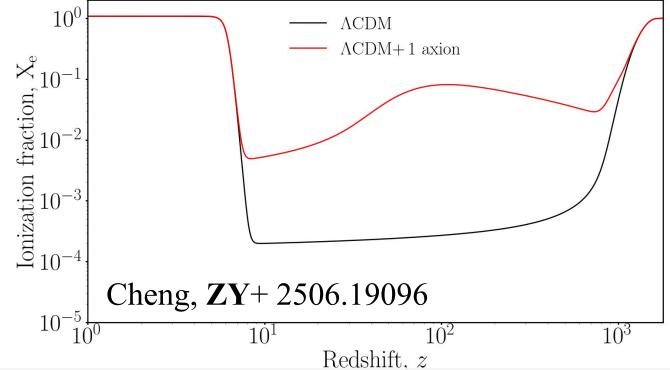
Axion freeze-in via axion photon coupling

$$\mathcal{L}_{a\gamma\gamma} = -\frac{1}{4} g_{a\gamma\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu} \,,$$

Primakoff process

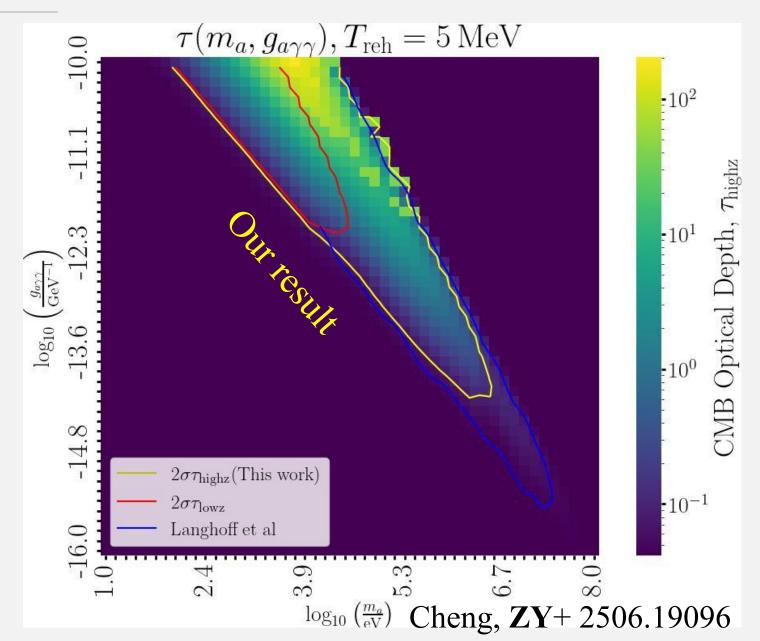


$$m_a = 10^5 \, \text{eV}, \; g_{a\gamma\gamma} = 6 \times 10^{-13} \, \text{GeV}^{-1}$$
 $T_{\text{reh}} = 5 \, \text{MeV}$



Single axion decay parameter space



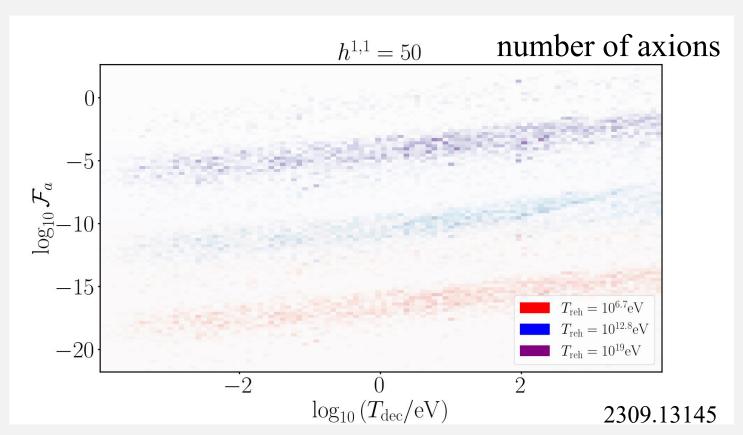


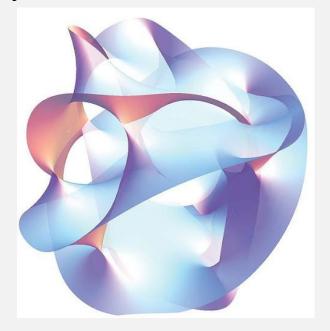
String axion abundance



Masses&couplings for multi-axions are predicted within string theory see Gendler+2309.13145

We obtain abundance and lifetime for each axion in a given theory

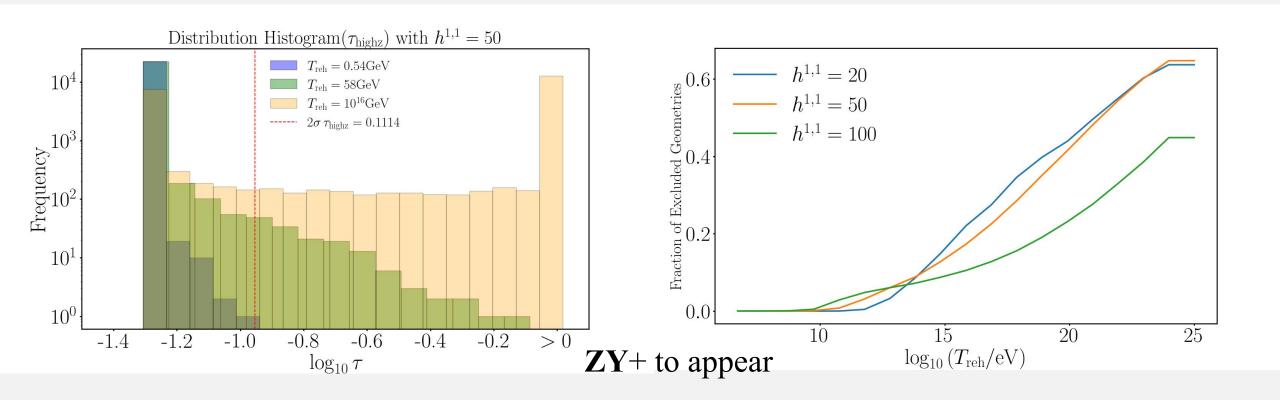




We want to constrain ensembles from string theory

Main preliminary results





We find that 50% of the models in ensembles with $h^{1,1} = 20$ and 50 prefer low reheating temperature $T_{reh} \le 10^{11} \text{ GeV} (10^{13} \text{ GeV})$ at 68% C.L.(95% C.L.).



谢谢!

