

Liam McAllister Cornell

Axions in Stockholm, July 1, 2025

Why does string theory have axions?

Antisymmetric tensor fields in six extra dimensions

Why many axions? How many?

Rich topology of extra dimensions. $N \sim 100$ to 100,000.

What are their properties?

Couplings scale with N.

Why does it matter what happens in string theory?

Axion couplings depend on nature of quantum gravity.

String theory has many axions because of the topological complexity of the six extra dimensions,

and the number N of axions matters: large-N theories are qualitatively different from small-N theories.

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and the number N of axions matters: large-N theories are qualitatively different from small-N theories.

We have, for the first time, constructed large-N theories in actual solutions of string theory.

We are in the process of discovering the properties of these theories.

Summary

Research program: computing the string axiverse, the landscape of many-axion effective theories in string theory.

We have made advances in computational geometry allowing study of theories with $N \gg 1$ axions.

We have constructed part of the axiverse, in type IIB string theory on Calabi-Yau threefolds.

We find geometric hierarchies involving powers of N, and corresponding hierarchies in the low-energy couplings.

Consequences for strong CP, DM abundance, superradiance, axion-photon couplings.

We can already exclude many string models.

Based on

Demirtas, Long, L.M., Stillman 2018 'Kreuzer-Skarke Axiverse': C_4 axions in type IIB on CY_3

topological complexity

⇒ hierarchies

Mehta, Demirtas, Long, Marsh, L.M., Stott 2021 Black hole superradiance

Demirtas, Gendler, Long, L.M., Moritz 2021 Strong CP problem

Gendler, Marsh, L.M., Moritz 2023 Axion-photon couplings

Sheridan, Carta, Gendler, Jain, Marsh, L.M., Righi, Rogers, Schachner 2024 Fuzzy dark matter

Bellas, Halverson, L.M., Vander Ploeg Fallon, Zhu 2025 Axions in F-theory $\begin{array}{l} \text{many more axions} \\ \Rightarrow \text{stronger effects} \end{array}$





Plan

- I. Axions in string theory
- II. Geometric hierarchies at $N \gg 1$
- III. Hierarchies in axion couplings

Setting

We start with superstring theory, for which the fundamental solution is $\mathbb{R}^{9,1}$.

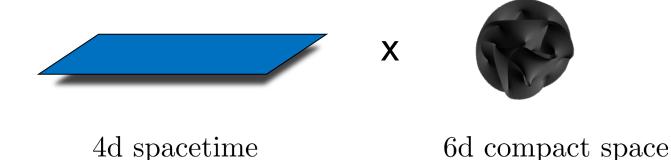
Setting

We start with superstring theory, for which the fundamental solution is $\mathbb{R}^{9,1}$.

We study compactifications in the geometric regime:

$$ds_{10}^2 = g_{\mu\nu}^{(3+1)} dx^{\mu} dx^{\nu} + ds_{X_6}^2$$

with X_6 a compact six-manifold that is large compared to a string.



Geometry of the extra dimensions

Einstein equations in vacuum: $\mathbf{Ricci} = 0$

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$$R_{\mu\nu} = 0$$

and

$$R_{mn}=0$$

$$\mu, \nu \in \{0, 1, 2, 3\}$$

$$m, n \in \{4, 5, 6, 7, 8, 9\}$$





4d spacetime

6d compact space

456789

Geometry of the extra dimensions

Einstein equations in vacuum: $\mathbf{Ricci} = 0$

$$R_{\mu\nu} = 0$$

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 and $R_{mn} = 0$

Vacuum solution of string theory = $\mathcal{M}^{3,1} \times [\text{compact Ricci-flat six-manifold}]$ [e.g. 'Calabi-Yau threefold']





4d spacetime

6d compact space

456789

Kaluza-Klein reduction

$$ds^{2} = \eta_{\mu\nu} dx^{\mu} dx^{\nu} + R^{2} d\vartheta^{2} \qquad \qquad \vartheta \cong \vartheta + 2\pi$$

Given a scalar Φ in 5d:

$$\Phi = \sum_{n} c_n \Phi_n(x^\mu, \vartheta)$$

$$\Phi_n(x^{\mu}, \vartheta) = \phi_n(x^{\mu}) \cdot \varphi_n(\vartheta) \qquad \varphi_n(\vartheta) = e^{in\vartheta}$$

$$0 = \Box_5 \Phi_n(x^\mu, \vartheta) \Leftrightarrow \left(\Box_4 + \frac{n^2}{R^2}\right) \phi(x^\mu) = 0$$

Massless 5d scalar Φ gives Kaluza-Klein tower of 4d scalars ϕ_n . Massless 4d scalar from zero-mode φ_0 .

In string theory it's the same, but with more extra dimensions.

Axions from extra dimensions

$$ds^{2} = \eta_{\mu\nu}dx^{\mu}dx^{\nu} + R^{2}d\vartheta^{2} \qquad \qquad \vartheta \cong \vartheta + 2\pi$$

Given a gauge field (1-form) A_M in 5d:

$$S_5 \supset -\int d^5x \, F_{MN} F^{MN} \,, \qquad M, N \in 0, \dots 4$$

$$\theta := \int_{0}^{2\pi} d\vartheta \, A_4$$

$$S_4 \supset -\frac{1}{2} \int d^4x \, f^2 \partial_\mu \theta \partial^\mu \theta \,, \qquad f^2 = \frac{2}{\pi R^2}$$

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$$ds^{2} = \eta_{\mu\nu} dx^{\mu} dx^{\nu} + ds_{\text{CY3}}^{2}$$

$$S_{10} \supset -\int d^{10}x F_{MNPQR} F^{MNPQR}, \qquad M, N \in 0, \dots 9$$

$$\theta_{i} := \int_{\Sigma_{i}} \left[dx^{4} dx^{5} dx^{6} dx^{7} \right] A_{4}$$

$$S_{4} \supset -\frac{1}{2} \int d^{4}x \sum_{i} f_{i}^{2} \partial_{\mu} \theta_{i} \partial^{\mu} \theta_{i}, \qquad f_{i}^{2} \propto \text{Vol}(\text{CY}_{3})^{-4/3}$$

A 1-form A_1 integrated over a 1-manifold yields an axion.

A 4-form A_4 integrated over a 4-manifold yields an axion.

In typical Calabi-Yau threefolds there are many submanifolds ⇒ many axion fields.

Setting

A **QGEFT** is an effective theory, of gravity and other fields, that results from an ultraviolet completion of gravity, such as a compactification of string theory.

Axion fields and couplings in QGEFT depend on topology, geometry, fields of extra dimensions.

We aim to understand axions in QGEFTs.

Strategy: enumeration of compactifications on Calabi-Yau threefolds (CY_3) .

Topologically simple compactifications yield simple QGEFTs.

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- No large dimensionless parameters. NDA works.
- Planck-suppressed operators have expected size:

$$\mathcal{L} \supset c_{\Delta} \frac{\mathcal{O}_{\Delta}}{M_{\rm pl}^{\Delta-4}} \text{ with } c_{\Delta} \gtrsim \mathcal{O}(1)$$

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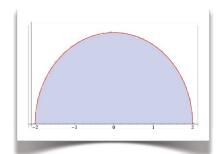
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• e.g.
$$\mathcal{O} = \sum_{i,j=1}^{N} M_{ij} \phi^i \phi^j$$
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Treating eigenvalues as \sim equal numbers is too simple.

M.C.D. Marsh, L.M., Wrase 2011



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Enumerate compactifications of string theory at large N.

Obstacle:

Until recently, topologically complex Calabi-Yau compactifications — those with $N\gg 1$ axions — were too complex to analyze.

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Kreuzer and Skarke, 2000

Cost of many steps is $\sim e^N$ in Sage/TOPCOM/Instanton/etc.

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We have overcome this problem.

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We found polynomial-time algorithms for 'everything', and implemented them in a software package, CYTools.

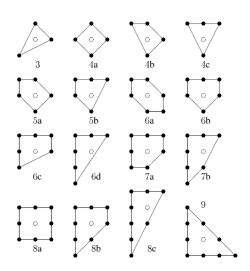
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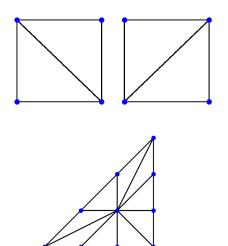
Combina-toric geometry

Key to unlocking large N:

Constructing CY_3 as hypersurfaces in toric varieties, and exploiting combinatoric structures.

Toric varieties correspond to triangulations of polytopes.





473,800,776 4d reflexive polytopes

 $<10^{428}~\rm{CY}_3$ Demirtas, L.M., Rios-Tascon 2020

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analysis

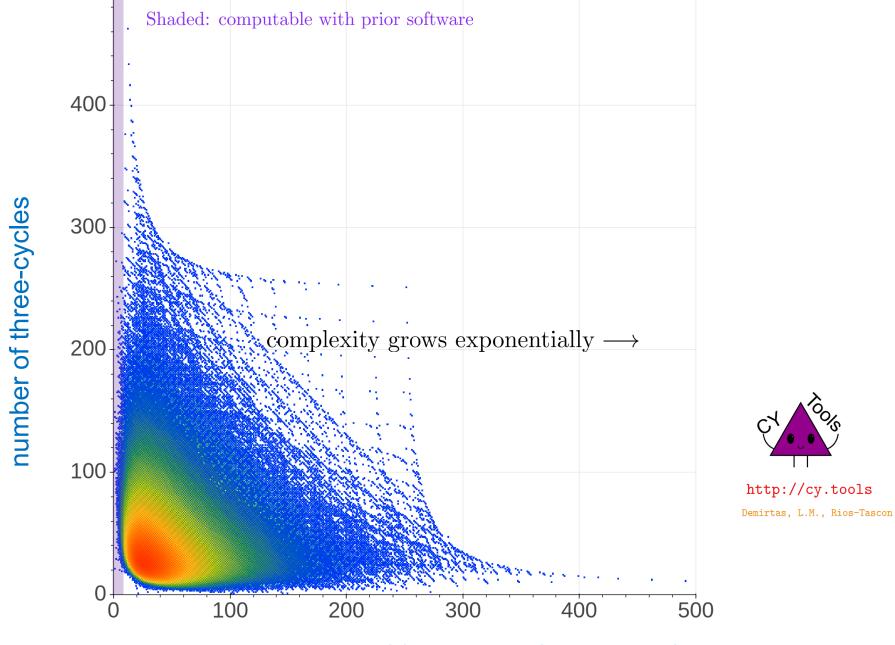


algebraic geometry



combinatorics

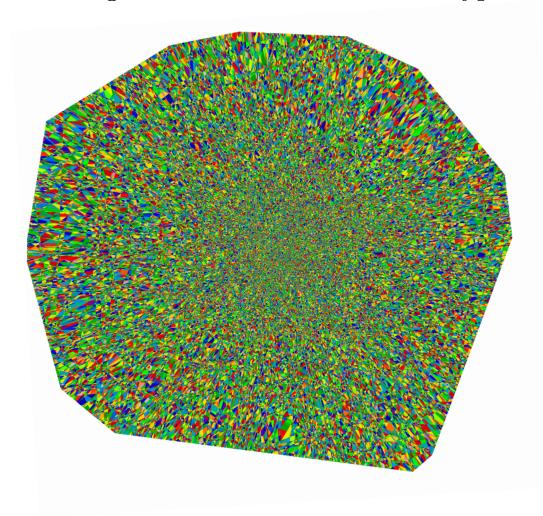
easier to automate



number of four-cycles (and axions)

The Kreuzer-Skarke Axiverse

Using CYTools we generated millions of CY_3 hypersurfaces.



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Compactifying type IIB string theory, we constructed an ensemble of N-axion QGEFTs, $1 \le N \le 491$.

These are incarnations of the string axiverse.

Arvanitaki, Dimopoulos, Dubovsky, Kaloper, March-Russell 09

Null hypothesis: isotropic internal space, one length scale L.

Our result: hierarchies of cycle sizes, polynomial in N.

Demirtas, Long, L.M., Stillman 2018

Topologically complex CY compactifications are anisotropic.

Consequence: hierarchical axion couplings.

Plan

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Cycle sizes

Curvature expansion: a fundamental expansion in string theory.

Coupling ℓ_s/L . ℓ_s : string length L: typical length

Can compute QGEFT when $\ell_s \ll L$.

Strategy: ensure that appropriate cycles are large in units of L.

Specifically: arrange that all 2-cycles have volume $> \ell_s^2$.

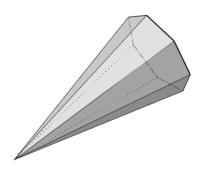
Let's see how to achieve this.

The Kähler cone

The number of 2-cycles is the number of axions, N.

So moduli space of 2-cycle sizes is $\subset \mathbb{R}^N$.

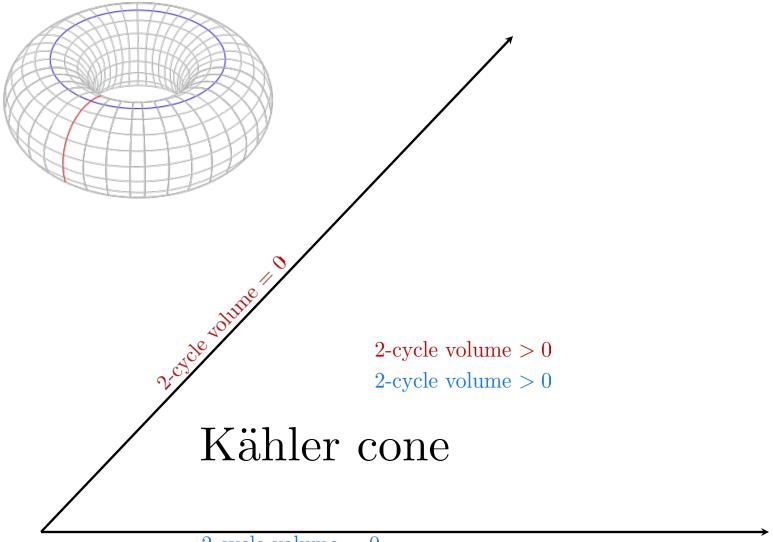
In fact it's a cone:



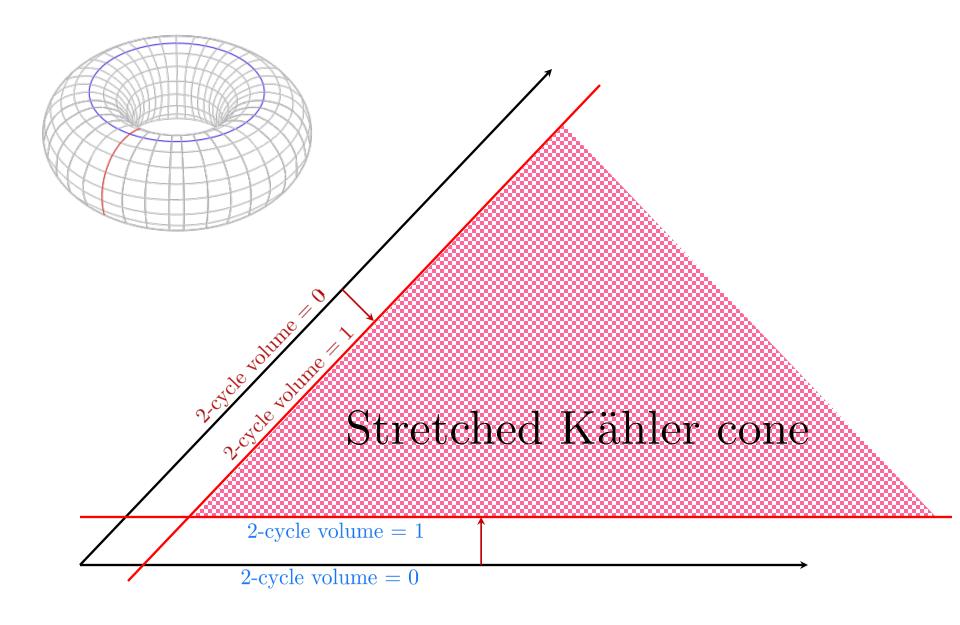
The Kähler cone $\mathcal{K} \subset \mathbb{R}^N$ is the region where all 2-cycles have volume ≥ 0 .

On the walls of the cone, one or more 2-cycles have zero size. Inside the cone, all (holomorphic) 2-cycles have positive size.

The Stretched Kähler cone is the subregion of \mathcal{K} where all 2-cycles have volume ≥ 1 .



-cycle volume = 0



We work in the stretched Kähler cone, where the curvature expansion is plausibly controlled.

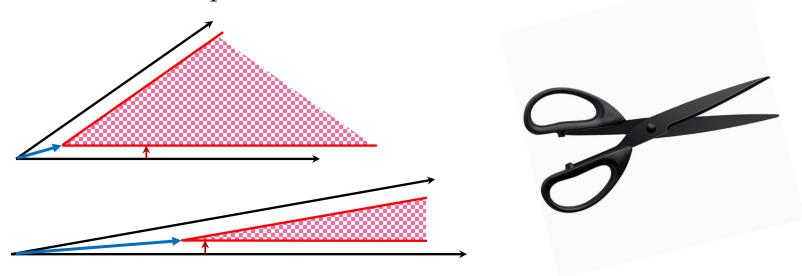
In a narrow cone with many walls, the subcone that is distance ≥ 1 from every wall has its apex far from the origin.

Fact: for $N \gg 1$, Kähler cones are narrow and have many walls.

So for $N \gg 1$, stretched Kähler cones have apex far from the origin.

This means that some cycles are large.

Their size is a power of N.



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So for $N \gg 1$, stretched Kähler cones have apex far from the origin.

This means that some cycles are large.

Their size is a power of N.

But other cycles have volumes of order unity.

A main result: hierarchies of cycle sizes, by powers of N.

'Topologically complex CY compactifications are anisotropic'.

Consider a spherical cow of radius R_{cow} .

Q: find minimum R_{cow} s.t. all its spots have size $R_{\text{spot}} > 1$ meter.



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A1: For an isotropic cow, $R_{\rm cow} \sim R_{\rm spot}$, so $R_{\rm cow,min} \sim 1$.

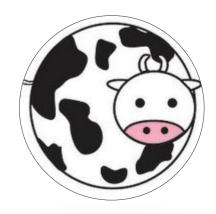


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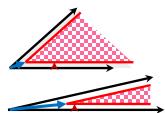
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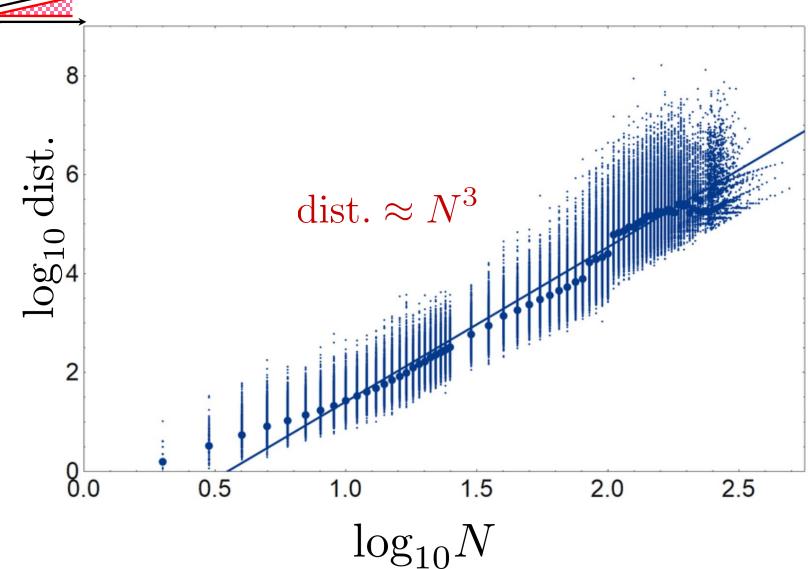
A2: For an anisotropic cow with many spots of widely varied size, one has $R_{\text{cow,min}} \gg 1$.



CY₃ are anisotropic, with many cycles of widely varied size.



Distance to tip of cone



Story so far

We constructed a landscape of string compactifications, and studied the resulting many-axion theories.

We found geometric hierarchies that are powers of the number of axions, N.

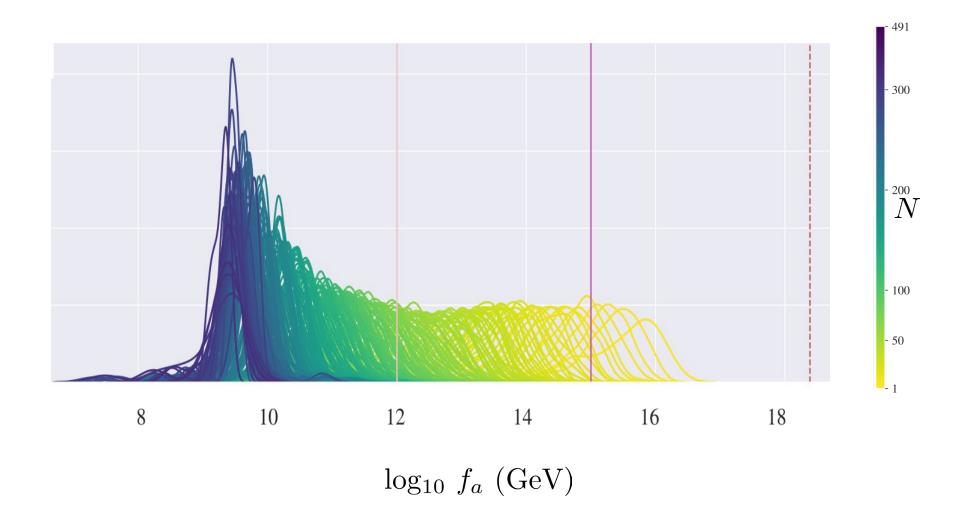
Next: the geometric hierarchies lead to hierarchies in the low-energy couplings, with consequences for phenomenology.

Warning: not clear that these extra-dimensional axions can form axion strings.

Consequence #1:

Decay constants spread over wide range, and diminish with N.

Decay constants f_a



Consequence #2:

Fuzzy axion dark matter is extremely rare for $N \gtrsim 15$.

Fuzzy axion dark matter

Could dark matter consist of ultralight axions?

For $m_a \lesssim 10^{-18} \,\mathrm{eV}$, chance of future tests.

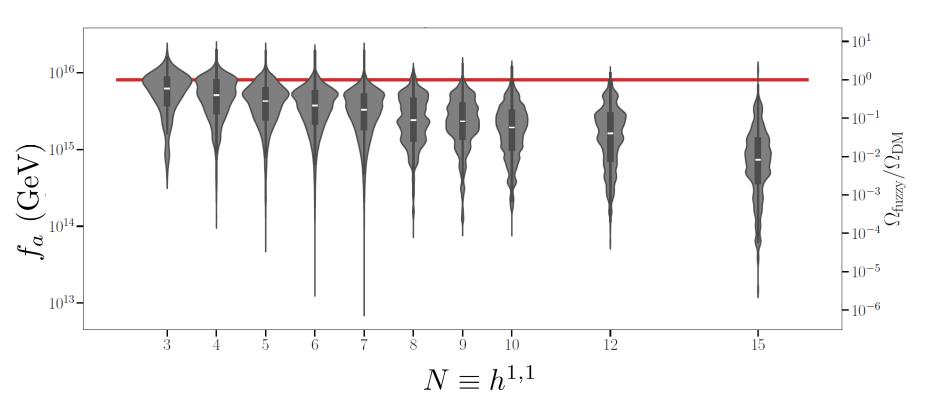
For $m_a \sim 10^{-21} \, \text{eV}$, significant constraints.

Misalignment abundance $\Omega_a \propto f_a^2 \sqrt{m_a}$.

 \Rightarrow need $f_a \gtrsim 10^{16} \, \text{eV}$ for abundant fuzzy axion DM.

We find f_a decreases with N, and is typically too small for fuzzy DM for $N \gtrsim 15$.

Fuzzy axions at small N



 f_a typically too small for fuzzy DM for $N \gtrsim 15$.

Fuzzy DM generally accompanied by overabundance of heavier-axion DM. Exceptions: fibration, or tuned cosmology.

Consequence #3:

Strong CP problem solved by PQ mechanism, without a PQ quality problem, for $N \gtrsim 15$.

PQ quality problem

The Peccei-Quinn solution of the strong CP problem is sensitive to Planck-scale CP-breaking, and so requires UV completion. This is the PQ quality problem.

High-scale physics, even Planck-scale physics, can easily break PQ symmetry badly enough to ruin mechanism.

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If there is another term that breaks the PQ shift symmetry,

e.g.
$$V_{\rm breaking} = \Lambda^4 \cos(\bar{\theta} + \psi)$$
 $\psi \in [0, 2\pi)$
then $\langle \bar{\theta} \rangle \sim \frac{\Lambda^4}{\Lambda_{\rm QCD}^4} \lesssim 10^{-10} \Leftrightarrow \Lambda^4 \lesssim 6 \times 10^{-14} \text{ GeV}^4$

Write in terms of $\Lambda^4 \equiv M_{\rm pl}^4 e^{-S}$ as $S \gtrsim 200.$

So every instanton carrying PQ charge must have $S \gtrsim 200$.

PQ quality problem

Every instanton carrying PQ charge must have $S \gtrsim 200$.

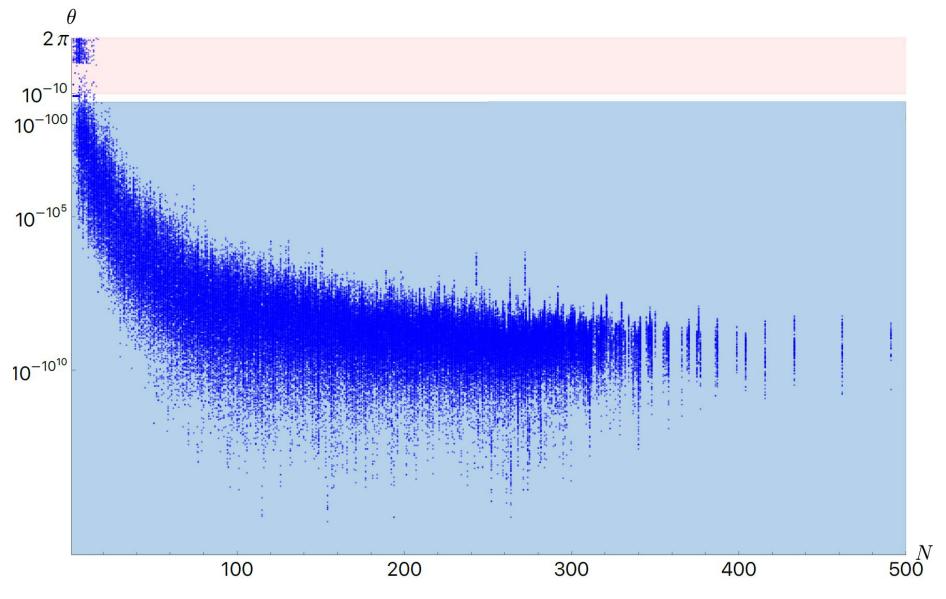
This dramatic sensitivity to Planck-scale physics is an opportunity for string theory: the effects in question can be computed.

In our ensemble this breaking comes from certain instantons.

We explicitly compute the leading such instantons.

When $N \gtrsim 15$, the Planck-scale breaking is negligible, and the quality problem is solved.

PQ quality

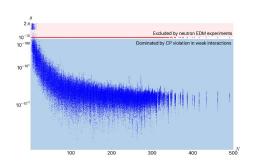


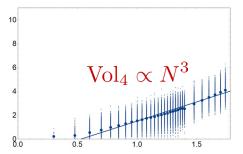
Demirtas, Gendler, Long, L.M., Moritz, 2112.04503

PQ quality problem in string theory

What's going on:

 $\Delta\theta \propto e^{-2\pi \text{Vol}_4}$, for leading 4-cycle.





Anisotropy of topologically complex space

- \Rightarrow large 4-cycles
- \Rightarrow small θ -angle.

Consequence #4:

Photon couplings to ultralight axions are hierarchically suppressed, if there is a stringy instanton on the QED cycle.

Axion-photon couplings

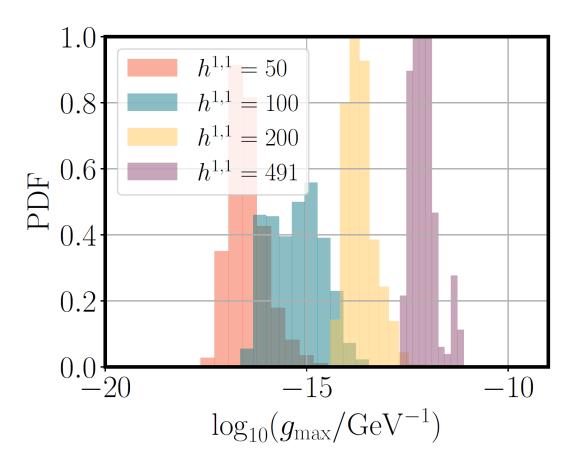
$$\mathcal{L}_{\text{QED}} = -\frac{1}{4} g_{a\gamma\gamma} \, \varphi^a \, F_{\mu\nu} \tilde{F}^{\mu\nu}$$

We have not built actual SM.

QED proxy: 4-cycle $\Sigma_{\rm QED}$ that 'could' host QED.

 $g_{a\gamma\gamma} \sim 1/f \Rightarrow \text{coupling increases}$ with N

Axion-photon couplings



 $g_{a\gamma\gamma} \sim 1/f \Rightarrow \text{coupling increases}$ with N

Light threshold

Gauge group with θ -angle

$$\mathcal{L}_{\text{gauge}} = \theta_a \, F_{\mu\nu} \tilde{F}^{\mu\nu}$$

and potential $\mathcal{L} \supset \Lambda^4 \cos(\theta_a)$ $\varphi_a = \theta_a/f_a, \quad a = 1, \dots, N$

$$\varphi_a = \theta_a/f_a, \quad a = 1, \dots, N$$

has negligible coupling to axions φ_j with $m_j \ll m_{\text{gauge}} \equiv \Lambda^2/f_a$.

QED does not have ordinary instantons of its own.

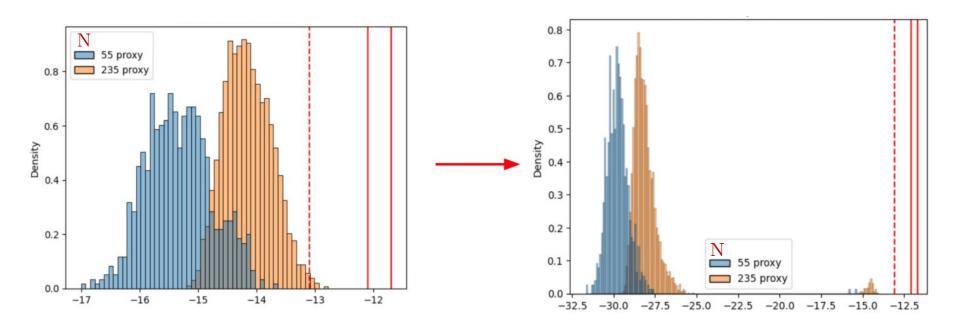
But there can be stringy instantons!

Then the photon has negligible couplings to axions j with

$$m_j \ll m_{\text{light}} \equiv f_{\text{QED}}^{-1} e^{-2\pi \text{Vol(QED)}}$$

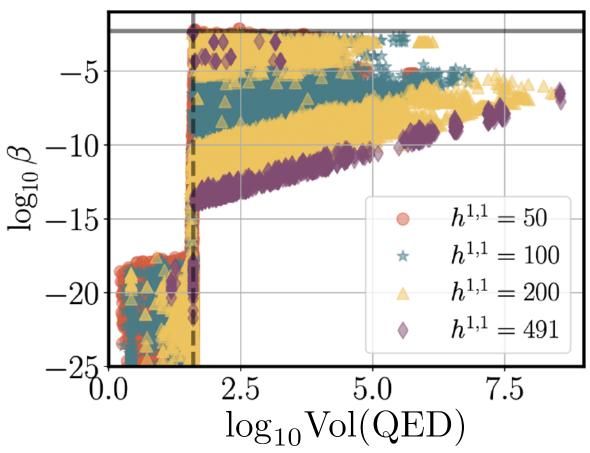
Light threshold: a mass scale set by stringy instantons on $\Sigma_{\rm QED}$. Axions with $m \ll m_{\text{light}}$ have negligible photon couplings. Makes ultralight axions invisible.

Light threshold



Birefringence prospects

Maximum birefringence



 $\beta \approx 0.3^{\circ}$ possible for Vol(QED) ≈ 40 .

Minami, Komatsu 20

Coincidentally, $\alpha^{-1} \approx \frac{1}{40}$ at GUT scale in SM.

So far, the string axiverse. Now, a work in progress: F-theory.

F-theory = type IIB string theory with strong + varying coupling

Compactify F-theory on a CY_4 that is a fibration over a base B_3 .

In general B_3 is not a CY_3 .

Consequence: much wider range of possible topologies.

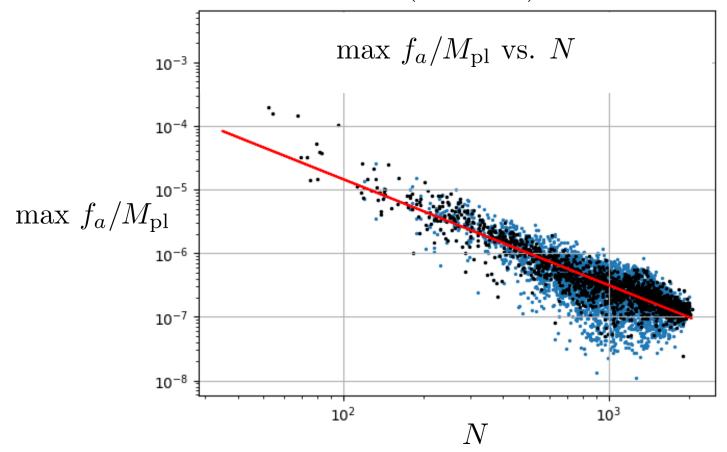
String theory on CY₃: $N \leq 491$.

F-theory on CY₄: $N \le 181,820$. Wang 2020

best-understood ensemble: $N \leq 2{,}591$ Halverson, Long, Sung 17

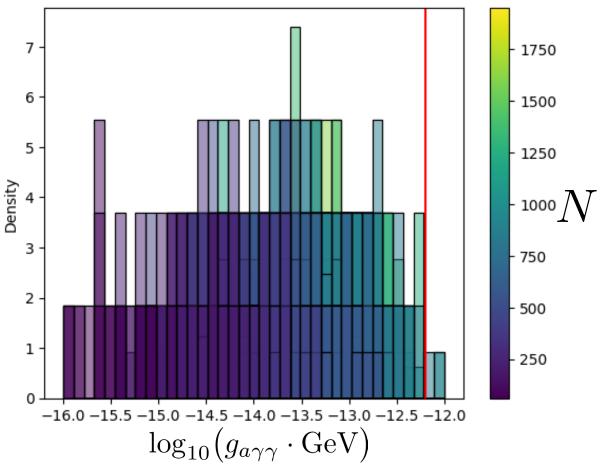
'Tree' ensemble, of toric threefold bases. $N \leq 2{,}591$ cf. Halverson, Long, Sung 17

Result: trends seen in CY₃ ($N \le 491$) continue for $N \le 2,591$.



Result: $g_{a\gamma\gamma}$ increases with N.

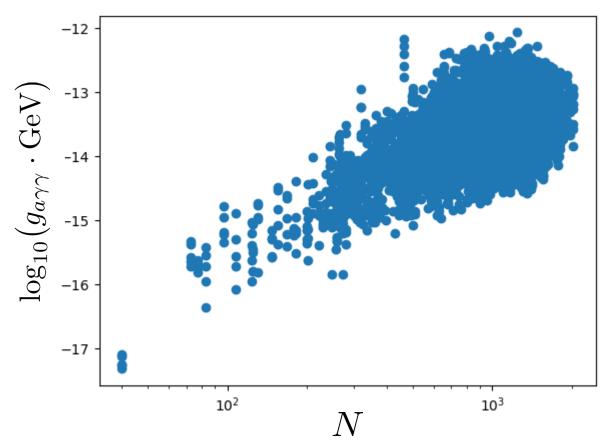
Here: preliminary first sample up to $N \leq 2{,}591$.



Bellas, Halverson, L.M., Vander Ploeg Fallon, Zhu WIP

Result: $g_{a\gamma\gamma}$ increases with N.

Here: preliminary first sample up to $N \leq 2,591$.



Models exist up to N = 181,820.

Conclusions and Outlook

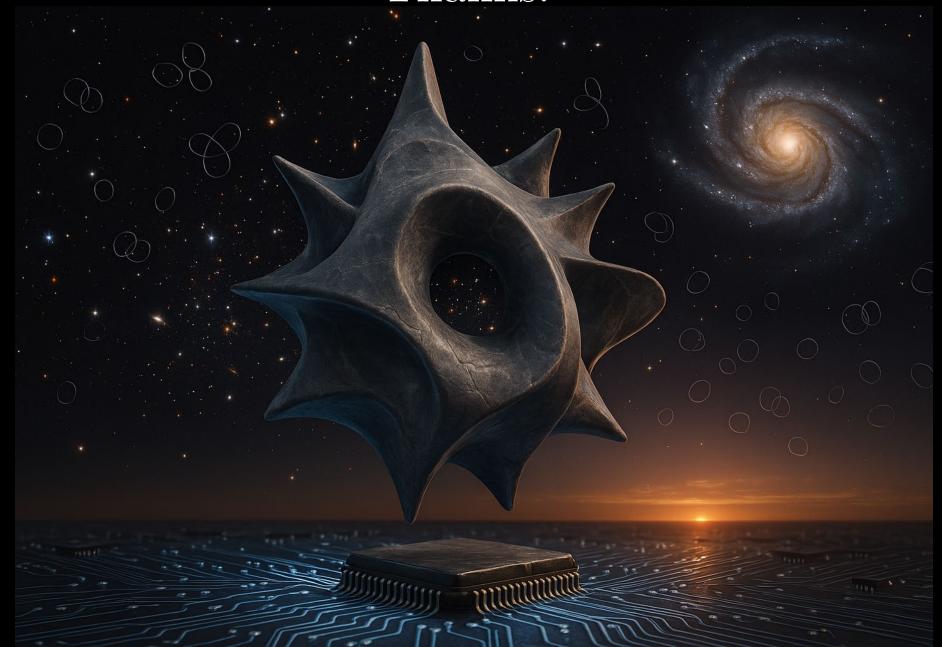
Advances in computational geometry have allowed us to construct parts of the string axiverse, with $N \gg 1$.

Geometric hierarchies from control of curvature expansion lead to hierarchies of couplings in the axion EFT.

Expect progress in:

- explicit SM constructions and couplings
- other string theories, other compactifications
- \bullet F-theory, with N up to 180,000
- moduli stabilization
- models of inflation and reheating
- detailed constraints

Thanks!



Axion couplings

$$N \text{ axions:} \qquad \theta^a := \int\limits_{a^{th} \text{ 4-cycle, } \Sigma_a} C_4 \qquad a = 1, \dots, N$$

$$\mathcal{L}_{\mathrm{axion}} = \mathcal{L}_{\mathrm{kin}} + \mathcal{L}_{\mathrm{pot}} + \mathcal{L}_{\mathrm{gauge}}$$

$$\mathcal{L}_{kin} = -\frac{1}{2} \, \mathcal{K}_{ab} \partial \theta^a \partial \theta^b$$

$$\mathcal{L}_{\text{pot}} = -\sum_{\text{instantons, } I} \Lambda_I^4 \left[1 - \cos(Q^I_a \theta^a) \right]$$

$$\mathcal{L}_{\text{gauge}} = \sum_{\text{gauge groups, A}} C_a^{\ A} \theta^a (F_{\mu\nu} \tilde{F}^{\mu\nu})_A$$

metric \mathcal{K}_{ab} : volumes, intersection numbers

scales Λ_I : 4-cycle volumes charges Q_a^I : topology

gauge groups: D-branes couplings C_a^A : topology

Computable in terms of topology and point in moduli space.

Turn the crank

- 1. Construct ensemble of CY₃ hypersurface topologies.
- 2. Sample moduli space inside stretched Kähler cone.
- 3. Specify (or model) D-brane configuration.
- 4. Identify contributing instantons.
- 5. Compute axion couplings.
- 6. Express in terms of mass+kinetic eigenstates.

Canonical axion couplings

$$\mathcal{L}_{\mathrm{axion}} = \mathcal{L}_{\mathrm{kin}} + \mathcal{L}_{\mathrm{mass}} + \mathcal{L}_{\mathrm{gauge}} + \dots$$

$$\mathcal{L}_{\text{kin}} = -\frac{1}{2} \sum_{a} (\partial \varphi^{a})^{2}$$

$$\mathcal{L}_{\text{mass}} = -\frac{1}{2} \sum_{a} m_{a}^{2} (\varphi^{a})^{2}$$

$$\mathcal{L}_{\text{gauge}} = -\frac{1}{4} \sum_{\text{groups, A}} g_{a\gamma^{A}\gamma^{A}} \varphi^{a} (F_{\mu\nu} \tilde{F}^{\mu\nu})_{A}$$

$$m_{a}^{2} = \frac{\Lambda_{a}^{4}}{f_{a}^{2}}$$

Fuzzy axion abundance

