

Modelling black hole scattering using comoving hyperboloidal coordinates

Amplitudes, Strong-Field Gravity, and Resummation (Nordita 2026)

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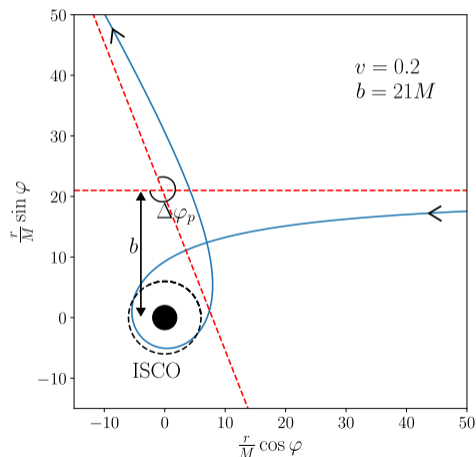
April 15, 2026



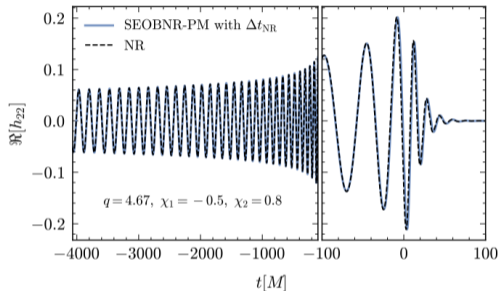
Motivation

High-precision probe of strong-field interaction

- ▶ GW astronomy: scattering angle calibrates EOB Hamiltonian
- ▶ Amplitudes and EFT ideas: scatter-to-bound relations, double copies, etc...

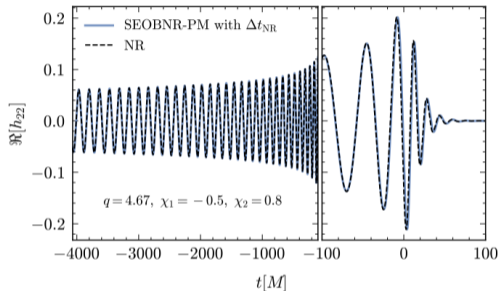


SF, EOB, PM



[Buonanno et al., 2024]

EOB informed by PM. But we can use SF as well.



[Buonanno et al., 2024]

EOB informed by PM. But we can use SF as well.

Complementary expansions of the same system

	1PM	2PM	3PM	4PM	5PM	...
0SF	G	G^2	G^3	G^4	G^5	...
1SF			νG^3	νG^4	νG^5	...
2SF					$\nu^2 G^5$...
					\vdots	\ddots

Credit: Oliver Long

$n\text{SF} \implies$ all orders in G ,

$n\text{PM} \implies$ all orders in ν

Lightning review of GSF

Systematic expansion of the field equations in the small mass ratio $q = \mu/M$

$$g_{\mu\nu} = g_{\mu\nu}^{\text{Sch./Kerr}} + qh_{\mu\nu}^{(1)} + \mathcal{O}(q^2) \quad \text{"+"} \quad T_{\mu\nu} = qT_{\mu\nu}^{(1)} + q^2 T_{\mu\nu}^{(2)} + \mathcal{O}(q^3),$$

↓ EFE

$$z^a(\tau) = z_{\text{geod}}^a + qz_{(1)}^a + \mathcal{O}(q^2)$$

0SF:

- ▶ geodesic motion on background BH spacetime
- ▶ bound/unbound orbits known analytically

1SF:

- ▶ secondary treated as a point-particle $T_{\mu\nu} \propto \delta^{(4)}(x - x_p)$ moving on a fixed geodesic of the background

Toy model: Scalar charge in Schwarzschild

Scalar charge Q , with inertial mass μ but no gravitational mass.

Small parameter $q_s = Q^2/(\mu M)$

Scalar field Φ satisfies Klein-Gordon equation

$$\square\Phi = -4\pi Q \int_{-\infty}^{\infty} d\tau \frac{\delta^{(4)}(y^\mu - y_p^\mu(\tau))}{\sqrt{-g}}$$

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Spherical harmonic decomposition: $\Phi(t, r, \theta, \varphi) = (Q/r) \sum_{\ell, m} \psi_{\ell m}(t, r) Y_{\ell m}(\theta, \varphi)$

KG equation reduces to 1+1D system

$$\psi_{,uv} + V_\ell(r)\psi = S(\tau)\delta(r - r_p(\tau))$$

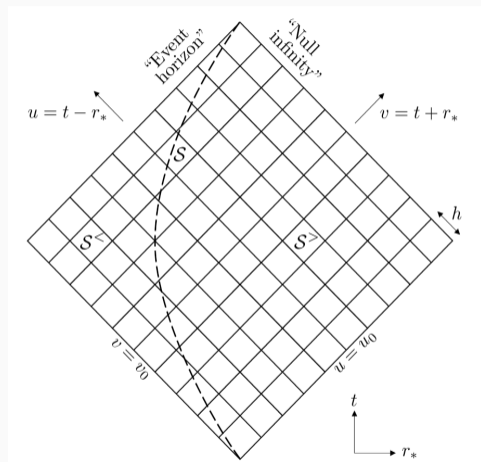
u, v : Eddington-Finkelstein coordinates

- ▶ Initial data (usually 0) on the characteristic rays $u = u_0$ and $v = v_0$.

Scalar field ansatz

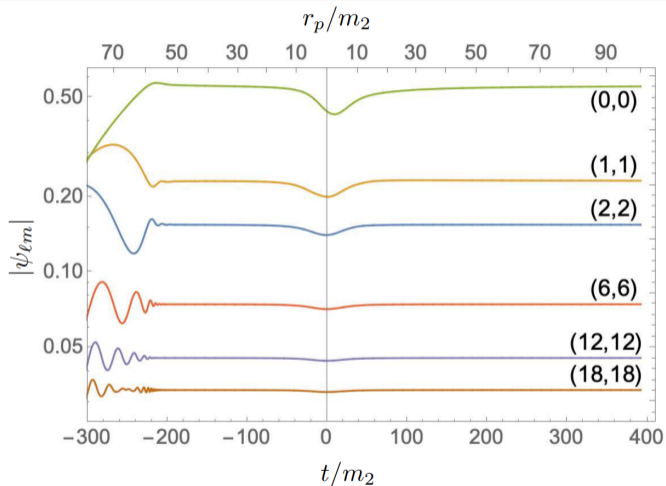
$$\psi(t, r) = \psi^>(t, r)\Theta(r - r_p(\tau)) + \psi^<(t, r)\Theta(r + r_p(\tau))$$

- ▶ Ansatz leads to jump conditions $J_\alpha = (\psi_{,\alpha}^+ - \psi_{,\alpha}^-)|_p$
- ▶ Discretise KG equation, evolve using jumps



Credit: Oliver Long

Scalar field results



16

Credit: Oliver Long

Self-force needs the “regular” part of the field

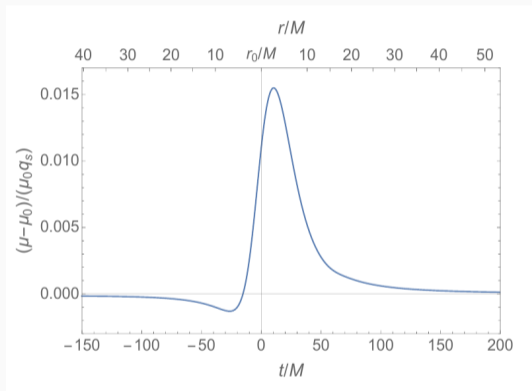
$$u^\beta \nabla_\beta (\mu u^\alpha) = Q \nabla^\alpha \Phi^R = Q^2 \nabla^\alpha \sum_\ell (\psi_\ell - B(t)) / r_p$$

Can be projected along u^α and orthogonal to it

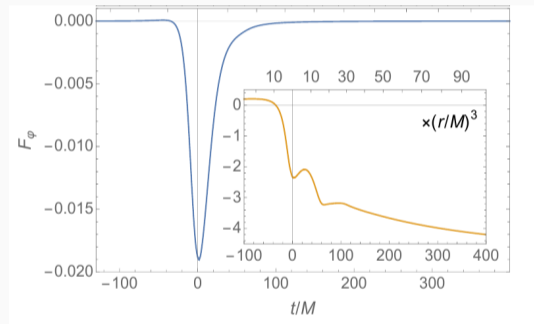
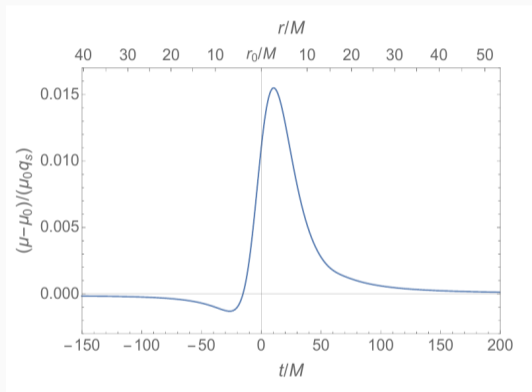
$$\text{Tangent: } \mu(t) = \mu_0 + Q\Phi^R$$

$$\text{Orthogonal: } \mu u^\beta \nabla_\beta u^\alpha = Q(\delta_\beta^\alpha + u^\alpha u_\beta) \nabla^\beta \Phi^R =: \mu q_s F^\alpha$$

Mass evolution and SF [Barack and Long, 2022]



Mass evolution and SF [Barack and Long, 2022]

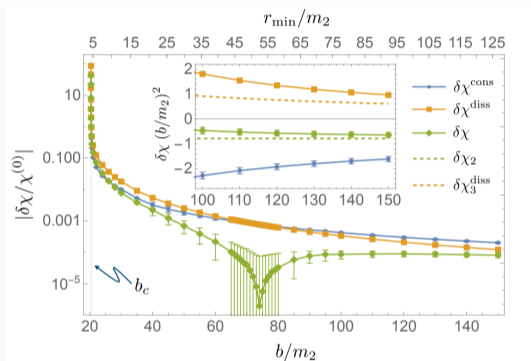


Self-force and scattering angle [Barack and Long, 2022]

Self-force leads to a change in the scattering angle $\chi = \chi^{(0)} + q_s \delta\chi$,

$$\delta\chi = \sum_{\pm} \int_{r_{\min}}^{\infty} \mathcal{G}_E^{\pm}(r) F_t^{\pm} + \mathcal{G}_L^{\pm}(r) F_{\varphi}^{\pm}$$

\mathcal{G} 's known analytically along the orbit



Credit: Oliver Long

More has been done comparing SF and PM, esp. in the context of resummation (See Olly's talk!)

For gravity, solve the $s = \pm 2$ component of adjoint Teukolsky equation.

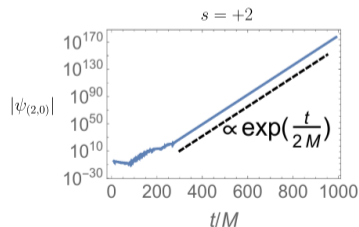
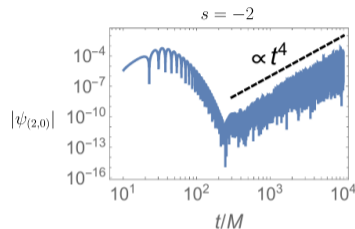
In Schwarzschild

$$\psi_{,uv}^{\pm} + U_s(r)\psi_{,u}^{\pm} + V_s(r)\psi_{,v}^{\pm} + W_s(r)\psi^{\pm} = 0, \quad (s = \mp 2 \text{ for } \psi^{\pm})$$

Jump conditions for the Hertz potential constructed from the source for the Weyl scalars [Long and Barack, 2021]

Divergent modes are solutions of the Teukolsky equation, but don't satisfy retarded BC's.

Lack of initial data!

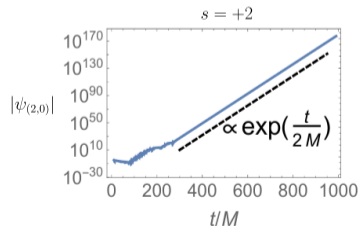
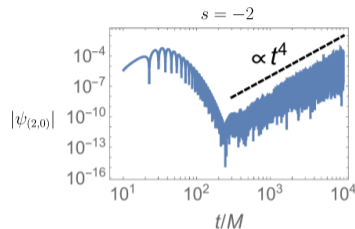


Things blow up [Long and Barack, 2021]

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Work-around in Schwarzschild: transform to Regge-Wheeler variable. But this can't be generalised to Kerr!

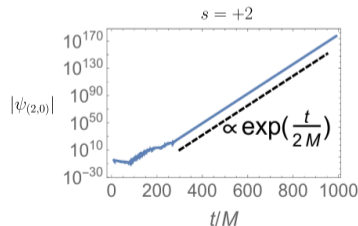
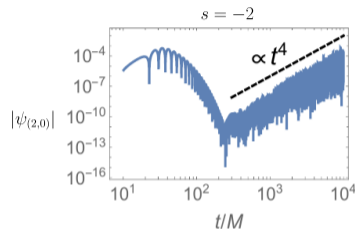


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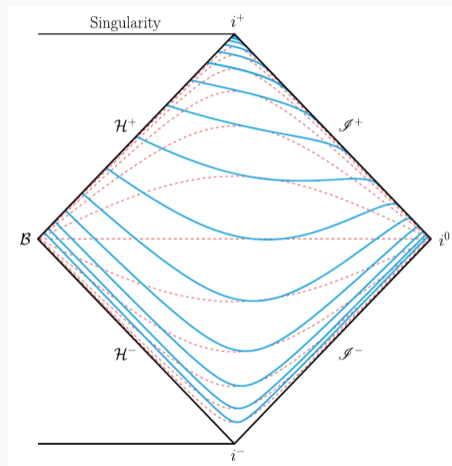
Work-around in Schwarzschild: transform to Regge-Wheeler variable. But this can't be generalised to Kerr!

Solution: Hyperboloidal compactification!
This should under-resolve incoming waves from \mathcal{I} and outgoing waves from \mathcal{H} and suppress the divergent modes



Hyperboloidal compactification

Usual t -coordinate singular at spatial infinity and bifurcation sphere.



[Macedo et al., 2022]

Hyperboloidal compactification

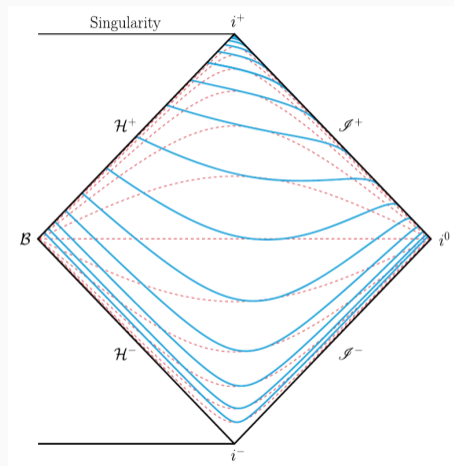
Usual t -coordinate singular at spatial infinity and bifurcation sphere.

Modify t -coordinate by removing divergence near spatial infinity and bifurcation sphere

$$\frac{t}{M} = \tau + \frac{1}{\sigma} - 2 \log [\sigma(1 - \sigma)]$$

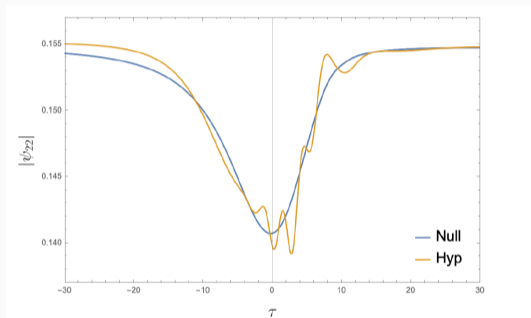
Compactify spatial coordinate $r = 2M/\sigma$

Exterior region $r \in (2M, \infty) \rightarrow \sigma \in (0, 1)$



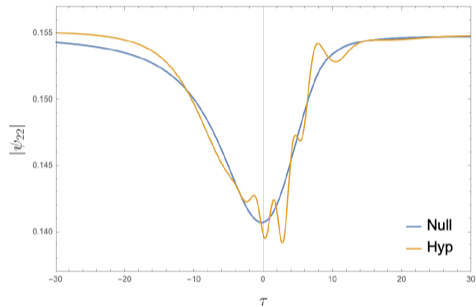
[Macedo et al., 2022]

Issues with minimal gauge



Credit: Oliver Long

Issues with minimal gauge



Credit: Oliver Long

Poor resolution of the field, even at periastron.

Comoving transformation

Keep $t \rightarrow \tau$.

Make the radial transformation comoving $\sigma = 1/2 \implies r = r_p$.

$$\frac{t}{M} = \tau + \frac{1}{\sigma} - 2 \log [\sigma(1 - \sigma)],$$
$$\frac{r^*}{M} = \frac{r_p^*(\tau)}{M} + \frac{1}{\sigma} - 2 + 2 \log \left[\frac{(1 - \sigma)}{\sigma} \right]$$

Motivation:

$$\sigma \rightarrow 0 \equiv \mathcal{I}^+ \implies u = \text{const.}, \quad v \rightarrow \infty,$$
$$\sigma \rightarrow 1 \equiv \mathcal{H}^+ \implies v = \text{const.}, \quad u \rightarrow \infty$$

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At large r_p the grid samples only the region near the particle.
No resolution of the wave zone or the strong field!

Final transformation

$$\frac{t}{M} = \tau + \frac{1}{\sigma^{N_\sigma}} - 2 \log [\sigma(1 - \sigma)],$$
$$\frac{r^*}{M} = \frac{r_p^*(\tau)}{M} + \frac{1}{\sigma^{N_\sigma}} - 2^{N_\sigma} + 2 \log \left[\frac{(1 - \sigma)}{\sigma} \right],$$

N_σ : controls sampling of wave zone, $N_\sigma = 1, 2, \dots$

Final transformation

$$\frac{t}{M} = \tau + \frac{1}{\sigma N_\sigma} - 2 \log [\sigma(1 - \sigma)],$$
$$\frac{r^*}{M} = \alpha(\sigma) \frac{r_p^*(\tau)}{M} + \frac{1}{\sigma N_\sigma} - 2^{N_\sigma} + 2 \log \left[\frac{(1 - \sigma)}{\sigma} \right]$$

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$\alpha(\sigma)$: controls sampling of strong field, $\alpha(1) = 0$

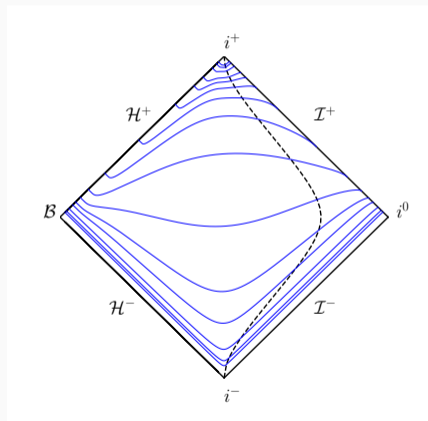
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N_σ : controls sampling of wave zone, $N_\sigma = 1, 2, \dots$

$\alpha(\sigma)$: controls sampling of strong field, $\alpha(1) = 0$

$$\alpha(\sigma) = \begin{cases} 1 - (2\sigma - 1)^{N_\sigma + 1}, & \sigma > 1/2 \\ 1, & \sigma \leq 1/2 \end{cases}$$



- Scalar field sourced by point charge in Schwarzschild

$$\square\Phi = -4\pi q \int d\tau \frac{\delta^4(x^a - x_p^a(\tau))}{\sqrt{-g}}.$$

- Multipole decomposition: $\Phi = \frac{Q}{r} \sum_{\ell,m} \psi_{\ell m}(\tau, \sigma) Y_{\ell m}(\theta, \phi)$

Modal KG equation:

$$\ddot{\psi} + C_{\tau\sigma}\dot{\psi}' + C_{\sigma\sigma}\psi'' + C_{\tau}\dot{\psi} + C_{\sigma}\psi' + C_0\psi = S(\tau)\delta(\sigma - 1/2)$$

Order reduction:

$$\begin{aligned}\dot{\Psi} &= \Pi, \\ \dot{\Pi} &= S(\tau)\delta(\sigma - 1/2) - (C_{\tau\sigma}\partial_\sigma + C_\tau)\Pi \\ &\quad - (C_{\sigma\sigma}\partial_\sigma^2 + C_\sigma\partial_\sigma + C_0)\Psi\end{aligned}$$

- ▶ Approximate spatial derivatives by finite differences
- ▶ Evolve in time with an RK4 integrator

But how do we incorporate the source into the scheme?

Jump conditions

- ▶ Scalar field ansatz satisfying KG equation

$$\Psi = \Psi^+ \Theta(\sigma - 1/2) + \Psi^- \Theta(\sigma + 1/2)$$

- ▶ Field kinked at $\sigma = 1/2$
- ▶ Jump conditions relate the field and derivatives on one side to corresponding quantities on the other side in terms of the source

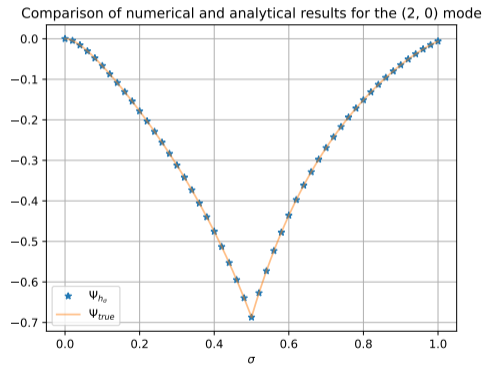
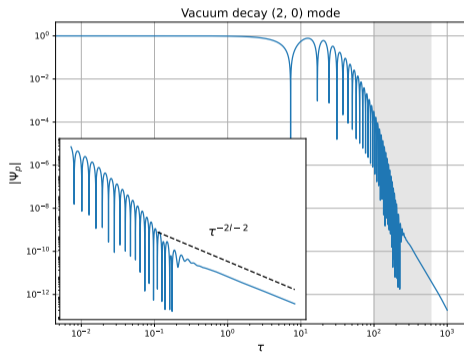
At the particle's location

$$\Psi^+ = \Psi^-$$

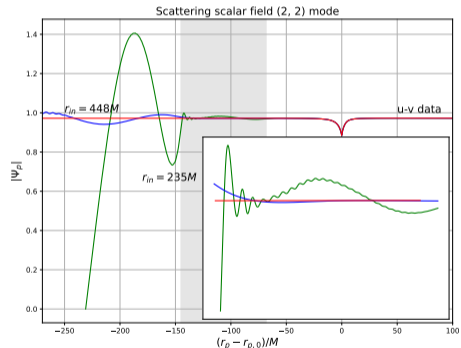
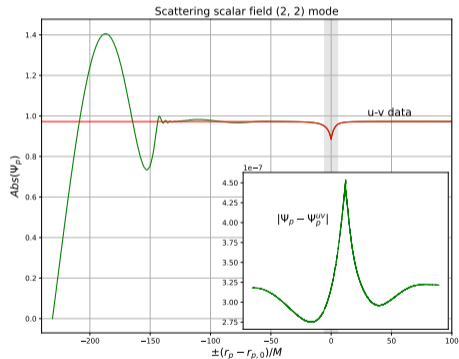
$$\partial_\sigma \Psi^+ = \partial_\sigma \Psi^- + J_\sigma, \quad J_\sigma = S/C_{\sigma\sigma}|_p$$

$$\partial_\sigma^2 \Psi^+ = \partial_\sigma^2 \Psi^- + J_{\sigma\sigma}, \quad J_{\sigma\sigma} = -(C_{\tau\sigma} \dot{J}_\sigma + C_\sigma J_\sigma)/C_{\sigma\sigma}|_p$$

Results: Vacuum + Circular



Scattering: comparison to null code

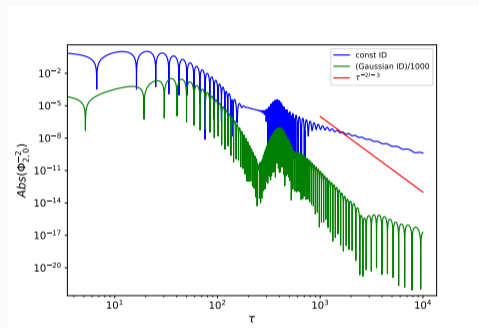


Teukolsky with hyperboloidal slices

Hertz potential ψ^\pm solves the adjoint Teukolsky equation with $s = \mp 2$

In vacuum

$$\begin{aligned} \ddot{\psi}_\pm + D_{\tau\sigma}^s \dot{\psi}'_\pm + D_{\sigma\sigma}^s \psi''_\pm \\ + D_\tau^s \dot{\psi}_\pm + D_\sigma^s \psi'_\pm + D_0^s \psi_\pm = 0, \end{aligned}$$

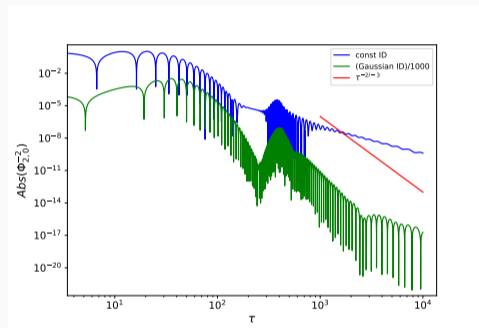


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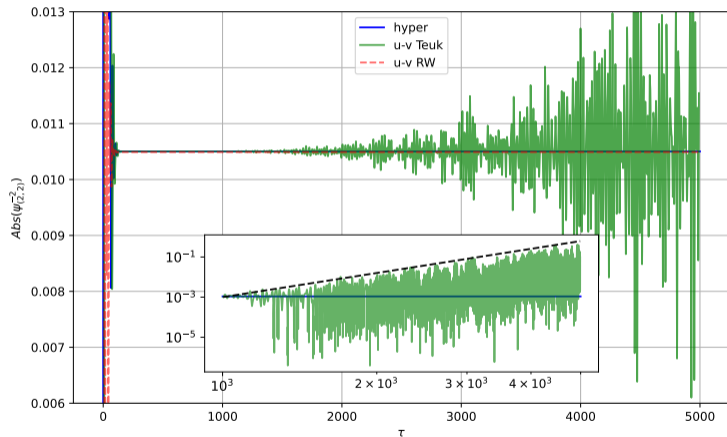
$$\begin{aligned} \ddot{\psi}_\pm + D_{\tau\sigma}^s \dot{\psi}'_\pm + D_{\sigma\sigma}^s \psi''_\pm \\ + D_\tau^s \dot{\psi}_\pm + D_\sigma^s \psi'_\pm + D_0^s \psi_\pm = 0, \end{aligned}$$



The field doesn't blow up!

Poor resolution of \mathcal{I}^+ \implies no clear tail. But we don't need this for GSF calculations.

Gravity: circular geodesics ($r_p^* = 9M$)



Gravity: scattering geodesics

We don't explicitly know the jumps analytically. But it satisfies

$$\frac{dJ}{dt_p} = F(\Psi_0, \Psi_4, \partial\Psi_0, \partial\Psi_4, \dots)$$

Algorithm:

- ▶ numerically solve $dJ/dt_p = \dots \rightarrow J$
- ▶ Compute $J_\sigma, J_{\sigma\sigma}, \dots$ [Long and Barack, 2021]
- ▶ Evolve adjoint Teukolsky for Φ
- ▶ Reconstruct the metric perturbation and extract GSF

$$\Phi \xrightarrow{\nabla^2} h^{rec} \rightarrow \text{GSF}$$





Conclusion and future work

- ▶ Compute the jumps and evolve the full gravity problem.
- ▶ Metric reconstruction of sourced case to compute scattering angle, energy- and angular momentum losses
- ▶ Strong-field PM resummation

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- ▶ Strong-field PM resummation

Thank You!

References

-  Barack, L., & Long, O. (2022). Self-force correction to the deflection angle in black-hole scattering: A scalar charge toy model. *Phys. Rev. D*, *106*(10), 104031. <https://doi.org/10.1103/PhysRevD.106.104031>
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-  Long, O., & Barack, L. (2021). Time-domain metric reconstruction for hyperbolic scattering. *Phys. Rev. D*, *104*(2), 024014. <https://doi.org/10.1103/PhysRevD.104.024014>
-  Macedo, R. P., Leather, B., Warburton, N., Wardell, B., & Zengino ğlu, A. ı. (2022). Hyperboloidal method for frequency-domain self-force calculations. *Phys. Rev. D*, *105*, 104033. <https://doi.org/10.1103/PhysRevD.105.104033>

Extra material

Finite difference stencils: vac + particle

Away from boundaries and next-to-boundary points

$$(\Psi'_i)^{(\text{VFD})} = \frac{1}{12h_\sigma} [(\Psi_{i-2} - \Psi_{i+2}) - 8(\Psi_{i-1} - \Psi_{i+1})] + \mathcal{O}(h_\sigma^4),$$

$$(\Psi''_i)^{(\text{VFD})} = \frac{1}{12h_\sigma^2} [- (\Psi_{i-2} + \Psi_{i+2}) + 16(\Psi_{i-1} + \Psi_{i+1}) - 30\Psi_i] + \mathcal{O}(h_\sigma^4).$$

At the particle

$$(\Psi_p^+)' = (\Psi'_p)^{(\text{VFD})} + \frac{1}{36} (18J_\sigma - 6h_\sigma J_{\sigma\sigma} + h_\sigma^3 J_{\sigma\sigma\sigma\sigma}) + \mathcal{O}(h_\sigma^4),$$

$$(\Psi_p^+)'' = (\Psi''_p)^{(\text{VFD})} + \frac{1}{18h_\sigma} (-21J_\sigma + 9h_\sigma J_{\sigma\sigma} - 2h_\sigma^2 J_{\sigma\sigma\sigma}) + \mathcal{O}(h_\sigma^3).$$

FD stencil at $p \pm 1$

Next to the particle

$$(\Psi_{p\pm 1}^+)' = (\Psi'_{p\pm 1})^{(\text{VFD})} + \frac{1}{288} (\mp 24J_\sigma + 12h_\sigma J_{\sigma\sigma} \mp 4h_\sigma^2 J_{\sigma\sigma\sigma} + h_\sigma^3 J_{\sigma\sigma\sigma\sigma}) + \mathcal{O}(h_\sigma^4),$$

$$(\Psi_{p\pm 1}^+)'' = (\Psi''_{p\pm 1})^{(\text{VFD})} + \frac{1}{288h_\sigma} (24J_\sigma \mp 12h_\sigma J_{\sigma\sigma} + 4h_\sigma^2 J_{\sigma\sigma\sigma} \mp h_\sigma^3 J_{\sigma\sigma\sigma\sigma}) + \mathcal{O}(h_\sigma^3).$$

At $\sigma = 0$,

$$\Psi'_0 = \frac{1}{12h_\sigma} (-25\Psi_0 + 48\Psi_1 - 36\Psi_2 + 16\Psi_3 - 3\Psi_4) + \mathcal{O}(h_\sigma^4),$$

$$\begin{aligned} \Psi''_0 = \frac{1}{12h_\sigma^2} (45\Psi_0 - 154\Psi_1 + 214\Psi_2 - 156\Psi_3 \\ + 61\Psi_4 - 10\Psi_5) + \mathcal{O}(h_\sigma^4). \end{aligned}$$

For the stencil at the other boundary, replace index $0 \rightarrow N, 1 \rightarrow N-1, \dots$

For the next-to-boundary points, use an asymmetric stencil

$$\Psi'_1 = \frac{1}{12h_\sigma} (-3\Psi_0 - 10\Psi_1 + 18\Psi_2 - 6\Psi_3 + \Psi_4) + \mathcal{O}(h_\sigma^4)$$

$$\Psi''_1 = \frac{1}{12h_\sigma^2} (11\Psi_0 - 20\Psi_1 + 6\Psi_2 + 4\Psi_3 - \Psi_4) + \mathcal{O}(h_\sigma^3)$$

For the corresponding stencil at the $N-1$ point, replace index $0 \rightarrow N, 1 \rightarrow N-1, \dots$

Field ansatz:

$$\psi(\tau, \sigma) = \psi^+(\tau, \sigma)\delta(\sigma - 1/2) + \psi^-(\tau, \sigma)\delta(\sigma + 1/2)$$

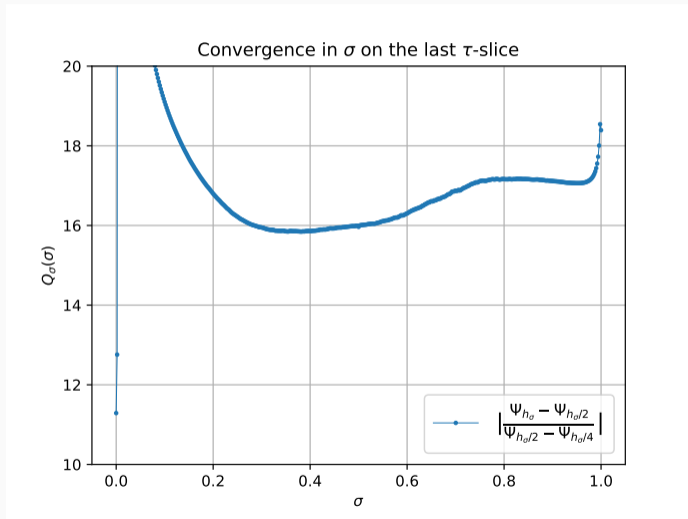
Leads to

$$S_{\ell m}(\tau) = \frac{\pi M f_p (3 - \dot{x}_p)}{8 E r_p} Y_{\ell m}^*(\pi/2, \varphi_p)$$

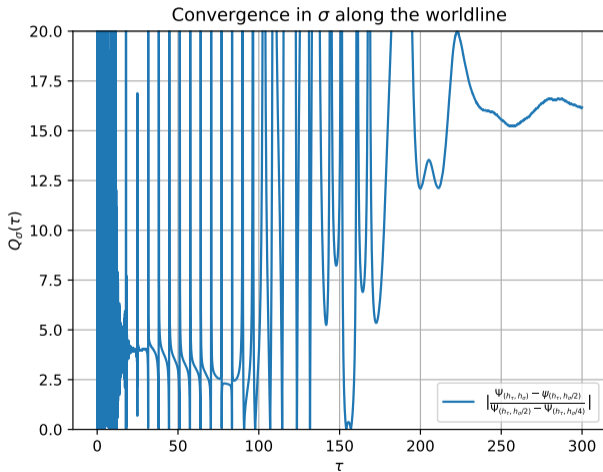
$$J_\sigma(\tau) = S(\tau)/C_{\sigma\sigma}|_p = -\frac{16\pi M f_p (3 - \dot{x}_p)}{E r_p (1 - \dot{x}_p^2)} Y_{\ell m}^*(\pi/2, \varphi_p)$$

$$J_{\sigma\sigma}(\tau) = -(C_{\tau\sigma}\dot{J}_\sigma + C_\sigma J_\sigma)/C_{\sigma\sigma}|_p$$

Circular orbit: convergence test in σ



Circular orbit: convergence test in σ (contd.)



Junk comparison

