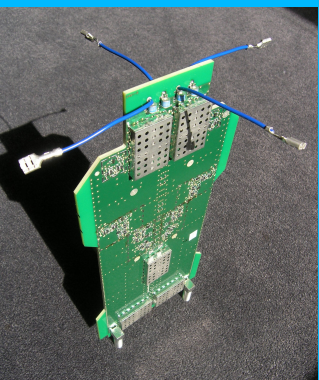
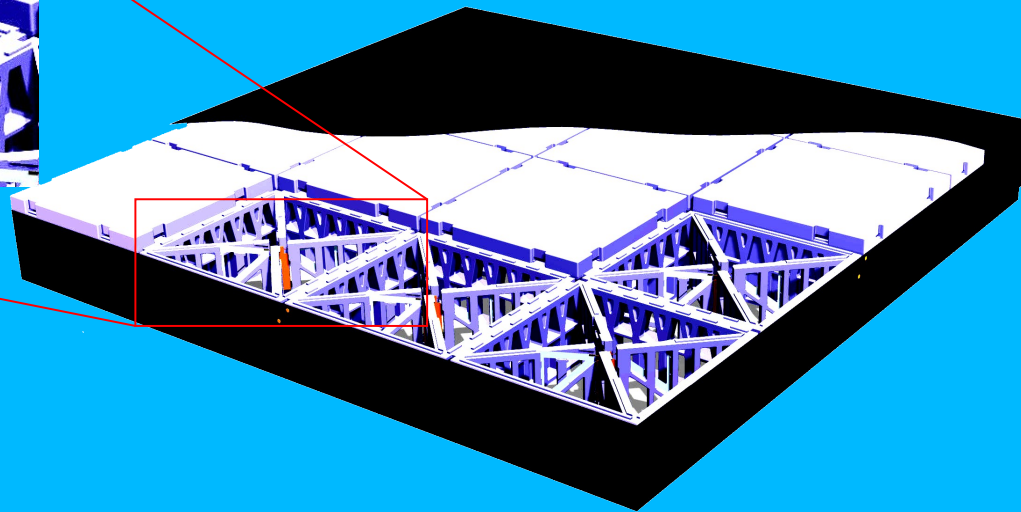


The LOFAR Observatory

Swedish LOFAR Meeting, 15 Jan 2009
René Vermeulen,
Director Radio Observatory





- Stations: electronically steered phased arrays of antennas
 - No moving parts
 - All-sky monitoring, transient buffer boards
 - Rapid electronic direction switching
 - Multiple simultaneous beams on sky

- Remotely operated from ASTRON Dwingeloo

- Aperture synthesis imaging, Tied Array, Transient Buffers

- 3 Gbit/s per station via fibre to BlueGene supercomputer

- Data model allows distributed archiving and post-processing

LOFAR on a European scale

ASTRON



LOFAR basics

Frequency ranges :	LBA 30 - 80 MHz Isolated dipoles	HBA 115 - 240 MHz Tiles of 4x4 dipoles	
18+ NL Core Stations	96 dipoles	2x24 dipoles	~2 km
18+ NL Remote Stations	96 dipoles	48 dipoles	~100 km
8+ International Stations	96 dipoles	96 dipoles	>1000 km

Sensitivity (after 4 h, 4 MHz bandwidth)

- @ 50 MHz ~ 3 mJy
- @ 150 MHz ~ 0.15 mJy

Up to 8 simultaneous 4MHz beams possible

LOFAR-NL

LOFAR

E-LOFAR: our ambitions for expanded capabilities in the next decade

LOFAR: In the beginning...

- Original science concepts:
 - George Miley & others in NL university community
 - Low-frequency radio surveys: under-explored terrain
- Original technology concepts:
 - Jaap Bregman & others at ASTRON
 - Phased arrays: great promise for the future
- NL subsidies/consortium: multi-disciplinary sensor network
 - Development of applications & operations left to communities
 - Geophysics, atmospheric, climatology, agriculture, ...
 - Astrophysics / radio telescope is dominant application
- International interest and collaboration growing rapidly
 - National LOFAR Astronomy Consortium in each country
 - International LOFAR Astronomy Consortium being formed

- Applications driving the design, called “key science projects” (KSP)
 - The epoch of re-ionisation
 - Extragalactic surveys
 - Transients and pulsars
 - Cosmic rays, particle astrophysics
 - Cosmic magnetism
 - Solar science, space weather
- For each KSP area, a team of astronomers is involved with ASTRON in realising the required technical capabilities
 - Will be recognised by means of privileged access to observing time
- Capabilities allow very wide range of projects that fall: centrally / periferally / not at all, within the general areas envisaged by: one / several overlapping, of the KSP teams

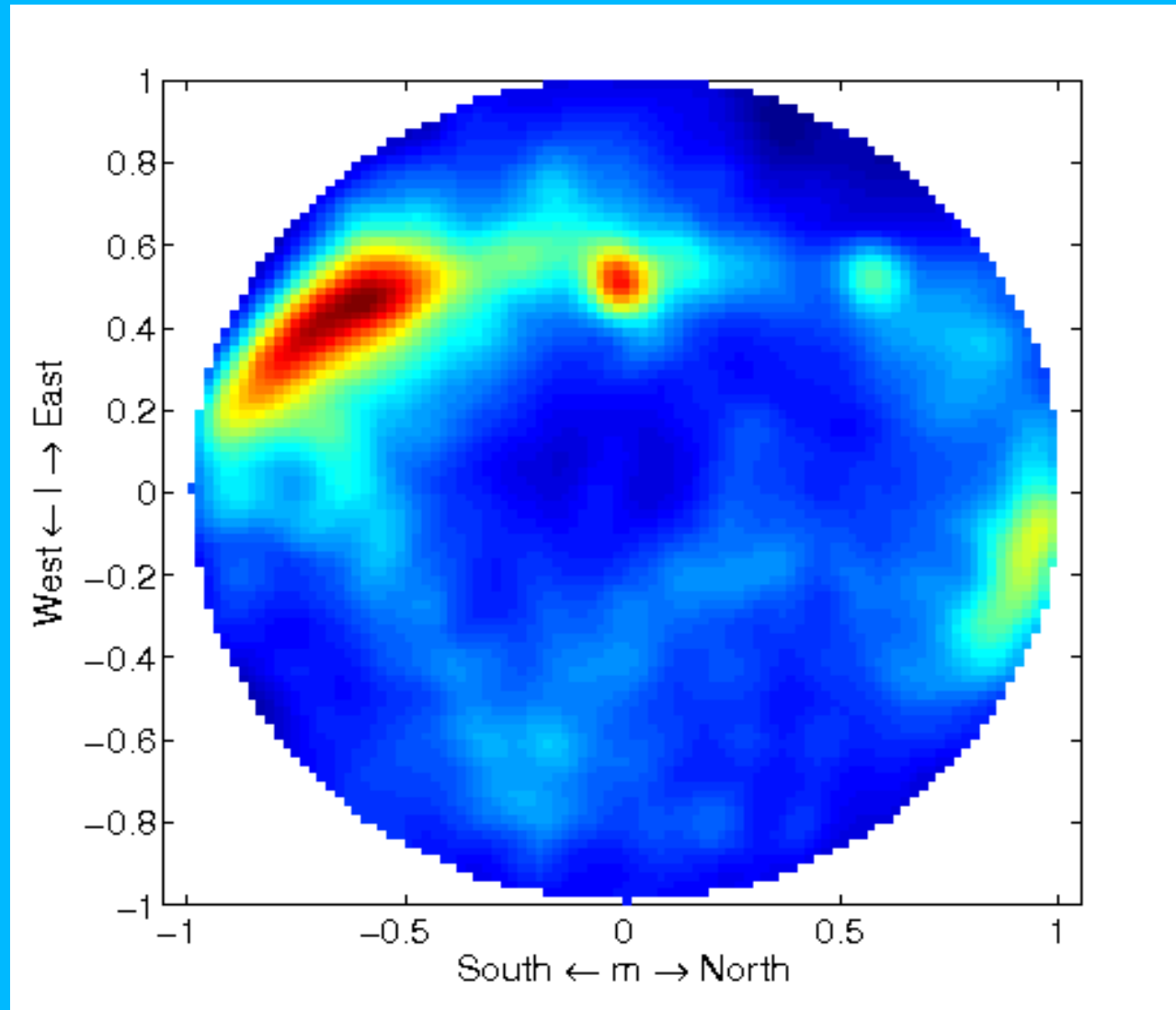
Wide range of astrophysical applications

A sample:

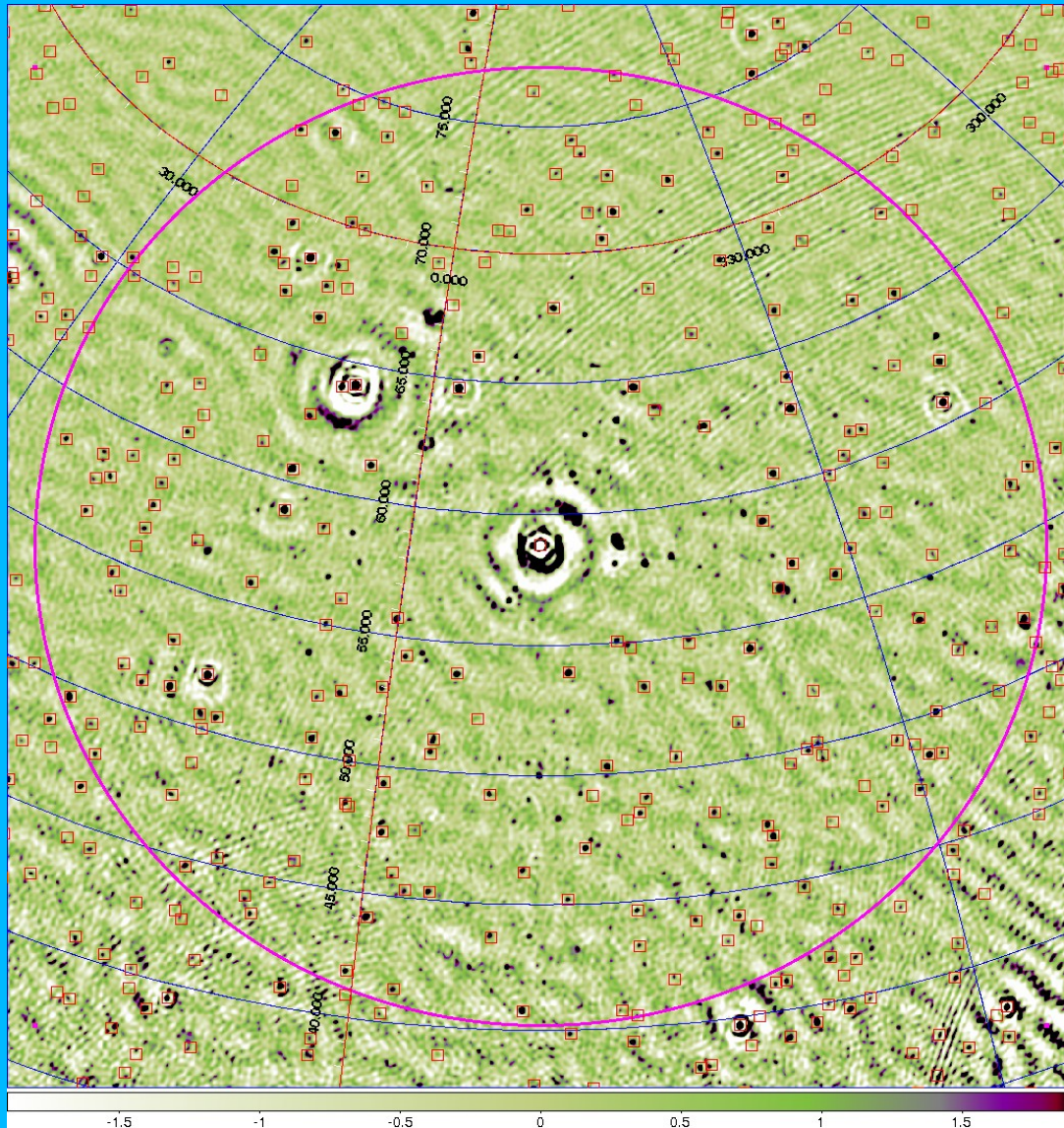
- The Sun and space weather
- Planets, exo-planets and SETI
- Single and multiple stellar transient/variable radio sources
- Pulsar science
- GRBs
- Galactic structure, ISM, SNR, starforming regions, magnetic properties
- Nearby Galaxies: structure, magnetic fields, dynamics
- Formation of (massive) galaxies and clusters
- Star formation in galaxies at high redshift
- Evolution of AGN activity / young and old radio galaxies
- Variable/transient extragalactic radio sources
- Properties of clusters and halos, large-scale magnetic structures
- Cosmic rays, particle astrophysics
- The Epoch-of-Reionisation
- Serendipitous discoveries
- **YOUR TOPIC HERE**

Test: Single Station All-Sky image

The Effelsberg sky at 14h on 29 October 2007



Test: Deep Image with HBA



Test: Confusion-limited image at 50 MHz

Sarod Yatawatta

16 dipoles (only ~ 70 baselines)

3 x 24h

38 - 59 MHz (B=6 MHz)

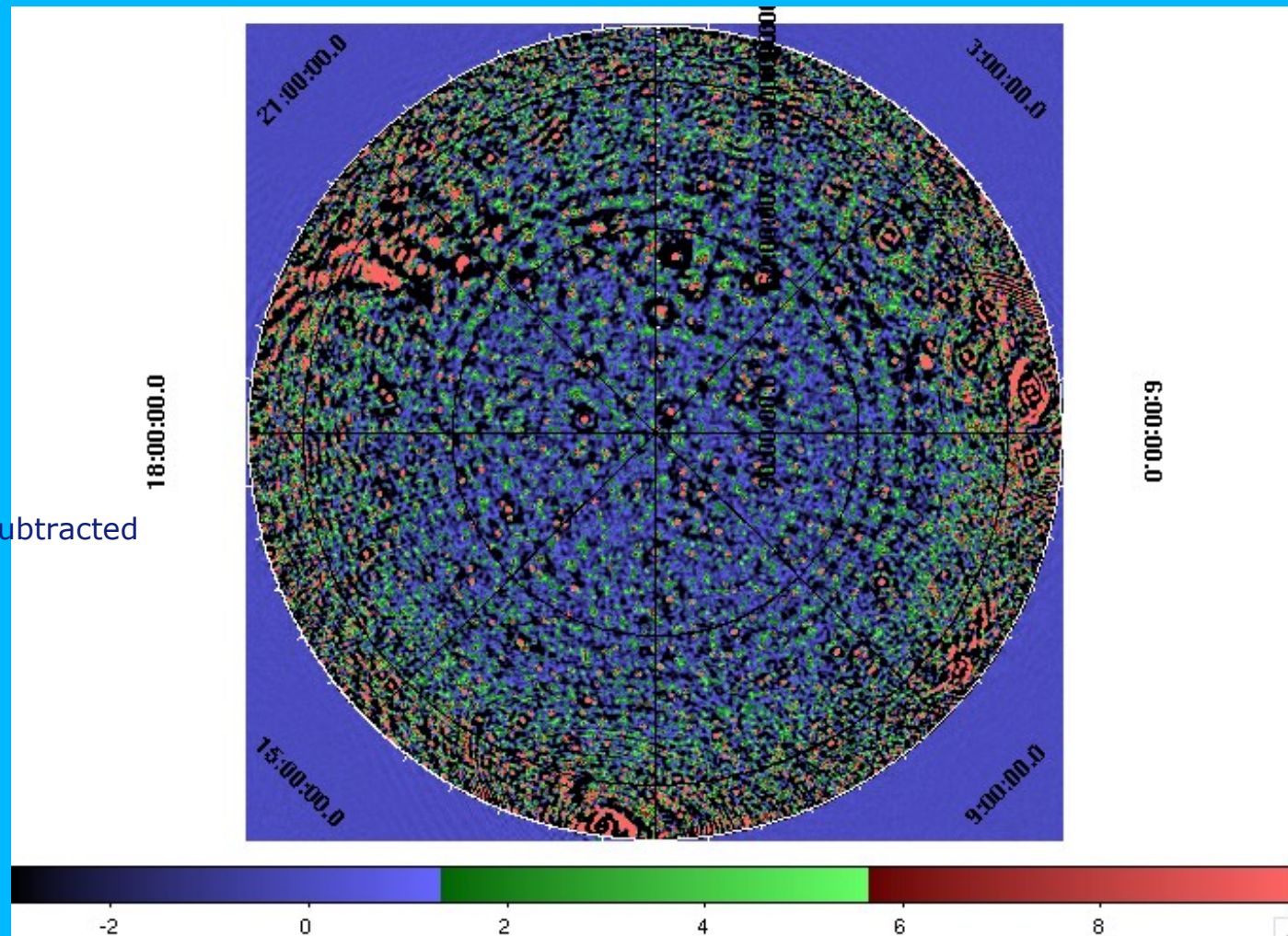
~ 800 sources !

PSF $\sim 0.5^\circ$

noise $\sim 0.5 - 1$ Jy

CasA/CygA ($\sim 20,000$ Jy) subtracted

- dipole beam corrected
- no deconvolution



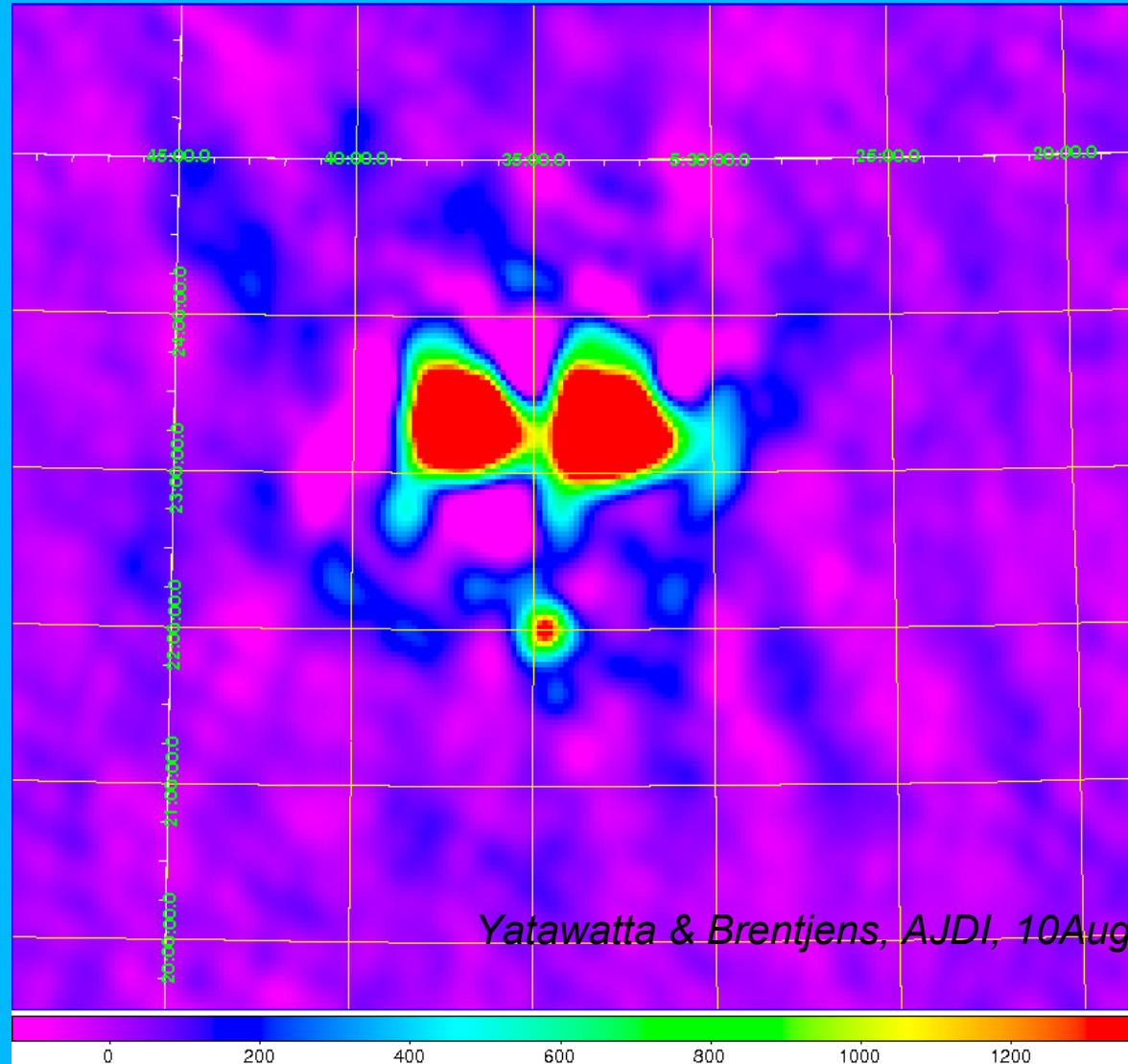
Test: Imaging Sun-Taurus conjunction

~12h syntheses
on 14+15 Jun 2008

~ 150 MHz

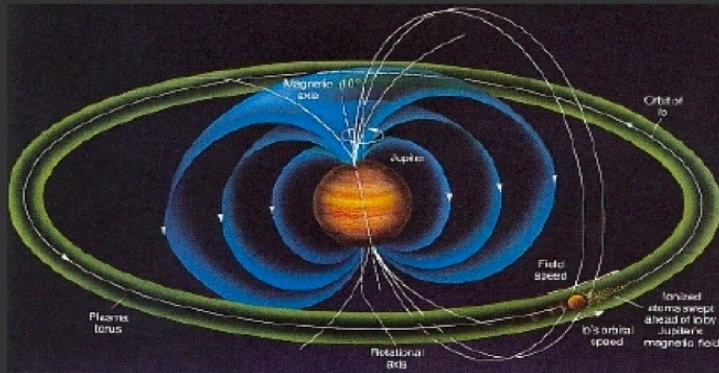
~ 10' PSF

Sun moves,
Taurus stays put



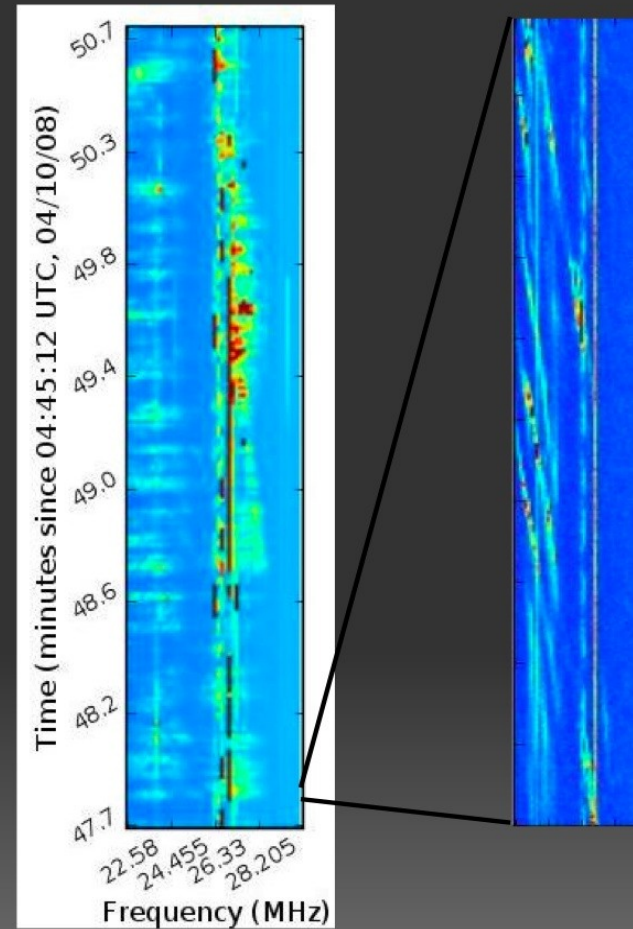
Yatawatta & Brentjens, AJDI, 10Aug08

Test: Jupiter dynamic spectrum



Jupiter/Io magnetosphere radiates cyclotron maser

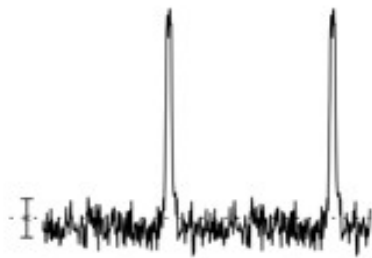
To Earth and CS1



Test: Pulsar Detection

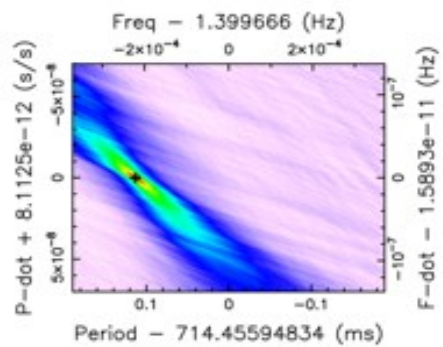
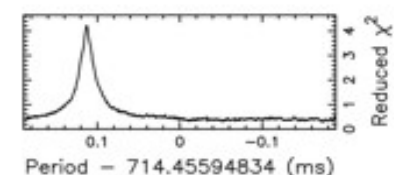
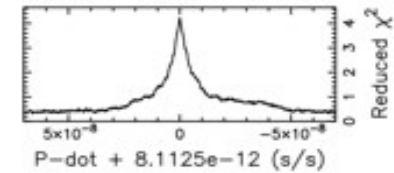
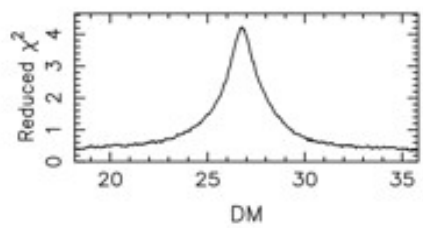
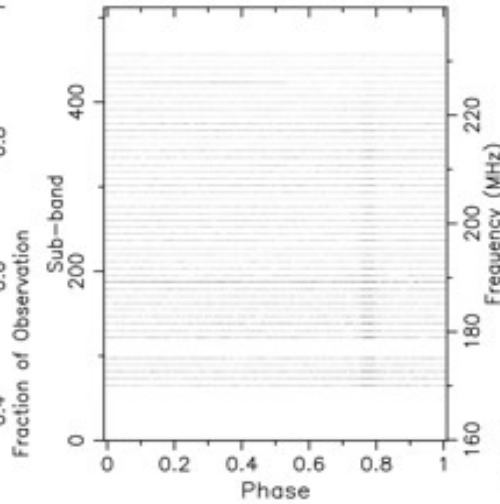
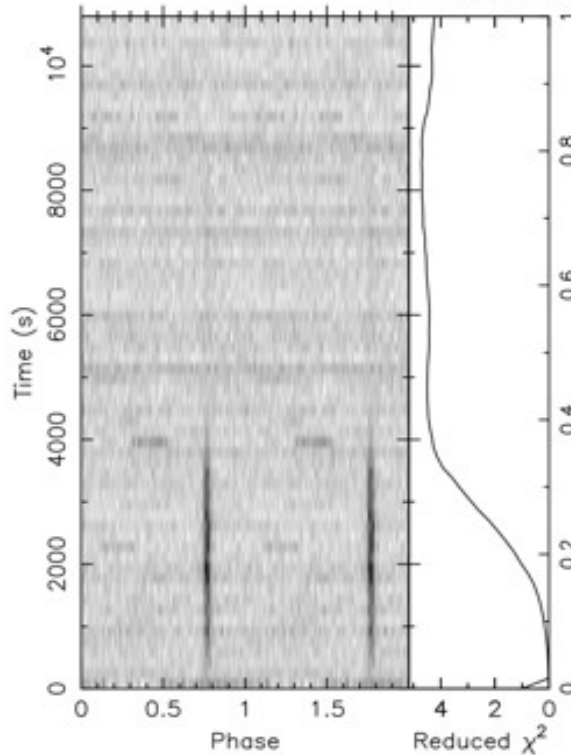
2 Pulses of Best Profile B0329+54_080411.sub000

Search Information



Candidate: 714.50ms_Cand
 Telescope: GBT
 Epoch_{topo} = 50000.000000000000
 Epoch_{bary} = 50000.00384567267
 T_{sample} = 0.0008192
 Data Folded = 13172736
 Data Avg = 29.47
 Data StdDev = 457.2
 Profile Bins = 256
 Profile Avg = 1.52e+06
 Profile StdDev = 1.037e+05

RA_{J2000} = 03:32:59.0000 DEC_{J2000} = 54:34:43.0000
 Best Fit Parameters
 Reduced χ^2 = 4.225 P(Noise) < 1.77e-101 ($\approx 21.4\sigma$)
 Dispersion Measure (DM) = 26.759
 P_{topo} (ms) = 714.52499(51) P_{bary} (ms) = 714.56905(51)
 P'_{topo} (s/s) = -0.2(3.7)x10⁻¹⁰ P'_{bary} (s/s) = -0.1(3.7)x10⁻¹⁰
 P''_{topo} (s/s²) = 0.0(2.2)x10⁻¹³ P''_{bary} (s/s²) = 0.0(2.2)x10⁻¹³
 Binary Parameters
 P_{orb} (s) = N/A e = N/A
 a_{sin(i)/c} (s) = N/A ω (rad) = N/A
 T_{peri} = N/A



LOFAR: a broad common-user Observatory

- LOFAR scope has gone from “project” to “observatory”
- Observatory aims at maximal, long-term overall scientific output
 - Integrated operations via ASTRON/Radio Observatory
 - National consortia participate: International Astronomy Consortium
- Many individual user groups
 - Each focusing on maximal yield in their specific research topic(s)
 - In individual collaborations that fit specific science and style
- Substantial contributions are rewarded with privileged access
 - Development, commissioning of common, integrated infrastructure
 - Participation in integrated operations
- Rapidly growing Open Skies portion

- Scarce/distinct resources to be allocated:
 - Telescope time, i.e. opportunity to determine setup/configuration
 - Data storage capacity and throughput
 - (Post)processing capacity and throughput
 - Access privileges to data / archive access, for specific science purposes, piggyback use

- 3-tiered allocation mechanism with pre-arranged fractions to various communities and “rights holders”
 - Operators (rewards sustained operational involvement)
 - Contributors (rewards building of common infrastructure)
 - Major shares initially, allowing some large, well-focused projects
 - Open Skies (fosters broad user base)
 - Limited, shared-risks commissioning opportunities for first call
 - Fraction grows rapidly in later years

Programme Committee assessment for all

Single, independent Programme Committee

- Composed of independent international experts
- Uniform assessment of all projects (new & progress reports)
 - Scientific merit
 - Technical feasibility
 - Project plan: timelines, people, data analysis resources
 - Scope, focus, demarcation
 - Cannot “reserve” technique, sky area, “umbrella science”,...
 - Piggy-back & archive use allowed/encouraged
 - Suitability for privileged (“large programme”) support

National Consortia can sponsor projects

- Rights accumulated by National Consortia:
 - Contributions of all adhering members added

- National Consortia then distribute their acquired rights:
 - To any/all of their favourite (large) projects

- Scientists are free to set up projects and collaborations for maximal efficiency:
 - Propose per project, with relevant group of collaborators
 - Seek collaborations for greater observing efficiency
 - Obtain (multiple) national sponsorships
 - Or use Open Time

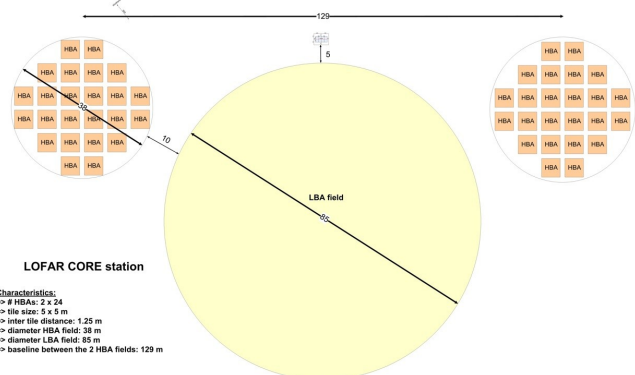
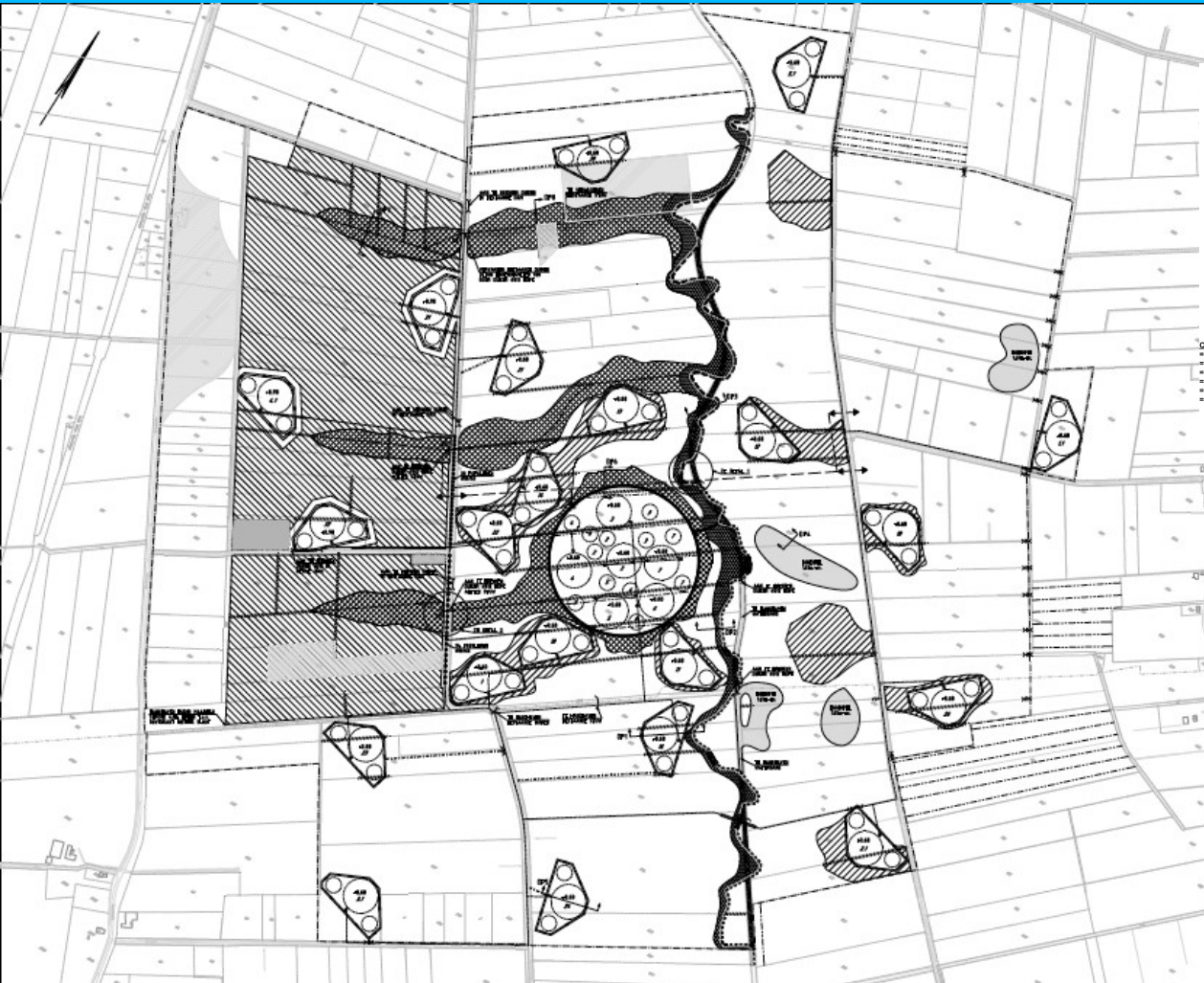
LOFAR commissioning phase

Commissioning phase requires intensive partnerships, coordination

- With users and user groups (**YOU!**)
 - Technical design/development issues: the TWG
 - Software development: the LAD
 - Users Committee will be set up

- With national consortia (the Internationalisation Working Group)
 - Handle management/policy issues as they arise
 - Prepare appropriate consortium for operational phase

Rollout: Building the core area



The Superterp

ASTRON



Living on the edge ?



Roll out “the Information Super Highway”

ASTRON



Fighting the mud

ASTRON



Flattening of RS307

- Superterp needs to settle, dry for a few months.
- Remote fields (RS302, RS503, RS307) being prepared
 - Fields were very wet in Oct/Nov
 - Then dryer weather set in, but also colder, everything frozen now
 - antenna installation starts as soon as thawed/dried
- Effelsberg LBA in place
- Unterweilenbach, Tautenburg, Potsdam work in progress
- Hardware delivery issues largely solved
 - Hardware stockpiled/warehoused in Emmen for rapid rollout later
 - TBB delivery still delayed
- Rollout procedures are getting tested
 - Practical experience with field preparations
 - Cabinet electronic preparation in warehouse

LOFAR Rollout timeline

- Winter 2009: Limited test observing with first hardware
- Spring 2009: First processing software pipelines delivered
- Summer 2009: Intensive commissioning observing
- End of 2009: aiming for 36+8 stations in the field
- 2010: Ramp-up of science operations

Commissioning in 2009

- LOFAR observing resources will be limited in 2009:
 - Gradual increase of available hardware & software
 - Operational control tools also still in delivery/commissioning
 - Simultaneous debugging of many components
 - Still on high part of maintenance “bathtub curve”
- Scheduling and priorities set by observatory
 - Based on rollout & commissioning needs
 - Changes on short timescale
 - No guarantees for any time
- But there will be exciting opportunities !
 - The many technical needs MUST predominate
 - Some results will show-case LOFAR capabilities
 - Some resultant data will be useful for early science
 - Some tweaks or modest add-on observing may help this along

Astronomical commissioning observing plan

- Requirements for general, common LOFAR capabilities coordinated via LAD
 - A million-source-shallow-survey (MSSS)
 - to obtain initial Global Sky Model
 - for imaging pipeline and calibration purposes
 - Tied Array commissioning
 - Long baseline issues
 - Polarisation calibration
 - Etc.

New participants with time/manpower very welcome

LOFAR data processing workshop in February (M. Wise)

- Your own science projects: technical tests / pilot runs
- Bright ideas for novel initial science based on short observations of uncertain quality

The first call for LOFAR proposals

First call for proposals will be circulated soon,
deadline in spring 2009

- Validation/consolidation of commissioning period
- Very limited opportunities for Open Time in 2009
 - All on shared-risks basis
 - Either making novel use of ongoing commissioning observing
 - Or small new, stand-alone, or pilot observing projects
- If applicable, plans for 2009 to be described w.r.t. preparations for projects in 2010 and later
 - No long-term allocations made yet
 - But allows iteration: equitable balance between all interests
 - Specific science can start to be tackled by specific groups
 - Range of planned science becomes known

First call for proposals

- PC will provide scientific assessments, give feedback on longer term plans:
 - Oversubscriptions
 - Appropriateness, scope, focus
 - Overlaps, different groups allowed to use same data
 - Timeline, processing & analysis resources required
 - Growing Open Skies share to be defined
- All groups can then tune and focus their projects without losing any observing opportunities
 - Scoping out parameter space for 2010+
- Actual “full-scale science” proposals will be due later

- Overall science output from the LOFAR Observatory will benefit from:
 - breadth of the user community
 - healthy competition
 - Focused research carried out by dedicated research groups
- YOU:
 - Should set up your project scope and collaborations in the way you think best suits your science interests, your style of work, and the finite amount of your time and of telescope time
 - Within a large conglomerate or with a small group or on your own
 - A grand-design programme or a very focused project
 - Requesting observing time or data access rights
- LOFAR SCIENCE WILL BE EXCITING AND MULTI-FACETED !!

ASTRON's 2007 Christmas card

